## Orbital magnetism in the half-metallic Heusler alloys

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Using the fully relativistic screened Korringa-Kohn-Rostoker method I study the orbital magnetism in the half-metallic Heusler alloys. Orbital moments are almost completely quenched and they are negligible with respect to the spin moments. The change in the atomic-resolved orbital moments can be easily explained in terms of the spin-orbit strength and hybridization effects. Finally, I discuss the orbital and spin moments derived from x-ray magnetic circular dichroism experiments.

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*Introduction*. Half-metallic ferromagnets are a new class of materials which attracted a lot of attention due to their possible applications in spintronics.<sup>1</sup> In these materials the two spin bands have a completely different behavior. While the majority spin band (referred to also as spin-up band) shows the typical metallic behavior, the minority spin band (spin-down band) is semiconducting. The spin-polarization at the Fermi level is 100% and these compounds could maximize the efficiency of the magnetoelectronic devices.<sup>2</sup>

de Groot and collaborators were the first to predict the existence of half-metallicity in the case of NiMnSb.<sup>3</sup> Since then a lot of materials have been predicted to be half-metals: other half-Heusler alloys (e.g., PtMnSb),<sup>4,5</sup> a large number of the full-Heusler alloys (e.g., Co<sub>2</sub>MnGe),<sup>6,7</sup> the quaternary Heusler alloys,<sup>8,9</sup> some oxides (e.g., CrO<sub>2</sub> and Fe<sub>3</sub>O<sub>4</sub>),<sup>10</sup> the manganites (e.g., La<sub>0.7</sub>Sr<sub>0.3</sub>MnO<sub>3</sub>),<sup>10</sup> the double perovskites (e.g., Sr<sub>2</sub>FeReO<sub>6</sub>),<sup>11</sup> the pyrites (e.g., CoS<sub>2</sub>),<sup>12</sup> the transition metal chalcogenides (e.g., CrSe) and pnictides (e.g., CrAs) in the zinc-blende or wurtzite structures,<sup>13–15</sup> and the diluted magnetic semiconductors (e.g., Mn impurities in Si or GaAs).<sup>16,17</sup> Heusler alloys are particularly interesting due to their very high Curie temperature and the similarity between their crystal structure and the zinc-blende structure adopted by the III-V and II-VI binary semiconductors like GaAs or ZnS.

Several papers have been devoted to the calculation of the electronic structure of the half-metallic Heusler alloys. All these studies produced a similar description of their magnetic properties.<sup>7,18,19</sup> In 2002 Galanakis *et al.* have shown that the appearance of the gaps in these alloys is directly connected to the magnetic spin moments and moreover that the total spin magnetic moment  $M_t$  scales linearly with the total number of valence electrons  $Z_t$  following the relation:  $M_t=Z_t$ –18 for the half-Heusler alloys like NiMnSb and  $M_t=Z_t$ –24 for the full Heusler alloys like Co<sub>2</sub>MnGe.<sup>4,6</sup> The orbital magnetic moments of these alloys on the other hand have attracted much less attention and results are scarce. Also experimentally only in few cases the orbital magnetic circular dichroic (XMCD) spectra of thin films.<sup>20,21</sup>

In this contribution I will present a study of the atomicresolved orbital magnetic moments of several Heusler alloys using first-principles calculations. This investigation was motivated by the experiments in Ref. 21 where the orbital moments derived from the XMCD spectra were one order of magnitude larger than the ones normally found in materials with cubic symmetry like the Heusler materials. In this case the spin-orbit coupling appears only in the fourth order of the perturbation theory contrary to materials crystallizing in highly anisotropic structures like the  $L1_0$ , e.g., CoPt, where the spin-orbit coupling appears in the second order and orbital moments are much higher.<sup>22</sup> My results show that orbital moments are much smaller than the experimental values, as expected by symmetry reasons. In the last section I discuss the discrepancy between the experimental and the theoretical results.

Calculations details. To calculate the orbital and spin magnetic moments I used the fully relativistic (FR) version of the Korringa-Kohn-Rostoker (KKR) multiple-scattering Green function method where the Dirac equation for the cellcentered potentials in the atomic spheres (ASA) is solved.<sup>23</sup> The Vosko, Wilk, and Nusair parametrization<sup>24</sup> of the local density approximation (LDA) is used for the exchange and correlation potential. This method has been already employed to calculate the effect of the spin-orbit coupling on the minority band gap in the case of half-metallic ferromagnets.<sup>25</sup> In the case of NiMnSb and similar half-Heusler alloys it was shown that the spin-orbit induces states within the gap but the effect is very weak and the alloys show a region of very high spin-polarization ( $\sim 99\%$ ) instead of a gap; defects have a much more pronounced effect on the destruction of the gap.<sup>26</sup>

If I compare the results obtained in this contribution by using the FR-KKR-ASA with the results obtained in Refs. 4 and 6 using the full-potential (FP) KKR method where the scalar-relativistic approximation is employed (the spin-orbit coupling is not taken into account), both versions of the KKR method reproduce a similar description of the spin magnetic moments; the differences are restricted to small deviations in the absolute values of the spin magnetic moments. Both  $C1_b$  and  $L2_1$  structures of the half- and full-Heusler alloys, respectively, are close-packed structures and ASA is expected to give a good description of their electronic structure with respect to FP. Moreover, spin-orbit is a weak effect and only marginally changes the spin moments. I should also note that LDA is known to underestimate the orbital moments by as much as 50% but reproduces the correct trends.22,27

Half-Heusler alloys containing Mn-Sb. The first family I

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	$m_{\rm spin}^X$	$m_{\rm orbit}^X$	$m_{ m spin}^{Mn}$	$m_{ m orbit}^{Mn}$	$m^{Sb}_{ m spin}$	$m_{ m orbit}^{Sb}$	$m_{ m spin}^{total}$	$m_{ m orbit}^{total}$	m <sup>total</sup>
FeMnSb	-0.973	-0.060	2.943	0.034	-0.040	-0.002	1.958	-0.028	1.930
CoMnSb	-0.159	-0.041	3.201	0.032	-0.101	-0.001	2.959	-0.010	2.949
NiMnSb	0.245	0.015	3.720	0.027	-0.071	-0.001	3.951	0.040	3.991
CuMnSb	0.132	0.006	4.106	0.032	0.028	-0.006	4.335	0.032	4.367
RhMnSb	-0.136	-0.033	3.627	0.035	-0.141	-0	3.360	0.001	3.361
PdMnSb	0.067	0.007	4.036	0.028	-0.117	-0	4.027	0.035	4.062
AgMnSb	0.106	0.006	4.334	0.031	0.040	-0.007	4.556	0.029	4.585
IrMnSb	-0.201	-0.094	3.431	0.092	-0.109	-0.001	3.130	-0.004	3.126
PtMnSb	0.066	0.006	3.911	0.057	-0.086	0	3.934	0.063	3.997
AuMnSb	0.134	0.021	4.335	0.027	0.056	-0.006	4.606	0.044	4.650

TABLE I. Spin  $(m_{spin})$  and orbital  $(m_{orbit})$  magnetic moments in  $\mu_B$  for the XMnSb half-Heusler compounds. The last three columns are the total spin and orbital magnetic moment and their sum, respectively.

will study is the MnSb-based half-Heusler alloys and in Table I I have gathered their magnetic moments. To this family belong the Fe-, Co-, Ni-, and PtMnSb which are halfmetallic (HM). RhMnSb and IrMnSb are isoelectronic to CoMnSb but the Fermi level falls within the minority valence band and the HM is lost (the total spin moments are slightly above the ideal value of  $3\mu_B$ ). The Cu-, Ag-, and AuMnSb have 23 valence electrons and if they were HM they should have a total spin moment of  $5\mu_B$ , but as it was shown in Ref. 4 this value is practically impossible to obtain; it is energetically more favorable to lose the HM. As a result also the spin moments of the Sb atoms are now parallel to the spin moments of the Mn atoms contrary to the other compounds.

The orbital moments are small with respect to the spin moments and only in the case of IrMnSb the  $m_{orbit}^{Ir}$  approaches the  $0.1\mu_B$ . In the case of the Sb atoms, the *sp*-bands lay low in energy and are almost completely filled for both spin directions.<sup>4</sup> There is only a very small majority spin *p*-weight around the Fermi level due to the antibonding *p*-*d* hybrids. As a result the antimonium orbital moment is practically zero for all compounds.

Mn atoms possess a large spin-magnetic moment in all Heusler alloys. The Mn spin-up states are practically completely occupied while Mn admixture in the occupied minority d states is limited; it is mainly the X atom which dominates the minority bonding d states.<sup>4</sup> Mn orbital moment is less than  $0.1\mu_B$  in all cases and remains parallel to the spin moment following the 3rd Hund rule. The latter rule, although derived for atoms, stands also for solids with few exceptions.<sup>28</sup> It states that if the d band is more than halffilled (Mn has 7*d*-electrons) then the spin and orbital moments should be parallel. Increasing the valence of the X atom by one electron either following the 3d series (Fe  $\rightarrow$  Co $\rightarrow$ Ni $\rightarrow$ Cu) or the 4d series (Rh $\rightarrow$ Pd $\rightarrow$ Ag) only scarcely changes the Mn orbital moment while there are significant variations in the value of the Mn spin moment. If now the X-atom changes along the 5d-elements series (Ir  $\rightarrow$  Pt $\rightarrow$ Au), the increase of the Mn spin moment by  $\sim 0.5 \mu_B$ at every step is accompanied by a large decrease of the Mn orbital moment which is practically halved. The increase of the spin moment is expected since the hybridization between Mn and a d atom decreases as the valence of the d atom increases leading to a more atomiclike electronic structure around the Mn site. The large effect on the Mn orbital moment in the case of the 5d atoms has been already discussed in Ref. 28, where using perturbation theory it was shown that the large spin-orbit coupling of the 5d elements has a large effect on the orbital moment of the 3d neighboring atoms in the case of alloys.

Finally for the X atom the orbital moment follows the Hund's rules and is always parallel to the spin magnetic moment. Note that the Fe, Co, Rh, and Ir spin moments are antiparallel with respect to the Mn atom. The orbital moment follows the changes of the spin moment and it increases as the number of valence electrons increase. As I substitute Co for Fe the orbital moment increases from  $-0.06\mu_B$  to  $-0.04\mu_B$  and then to  $0.015\mu_B$  for Ni in NiMnSb. The absolute value of the orbital moment depends strongly also on the spin-orbit coupling. This is clearly seen if I compare Ir with Co. Both atoms have similar spin moments;  $-0.16\mu_B$  for Co and  $-0.20\mu_B$  for Ir. On the other hand, cobalt's orbital moment is  $-0.04\mu_B$  while the Ir orbital moment is double as much  $(-0.09\mu_B)$ . Also hybridization plays an important role on the value of the orbital moment, e.g., in FePt Fe has a spin moment of  $2.9\mu_B$  instead of  $-1.0\mu_B$  in FeMnSb but the Fe orbital moment is similar in both cases; its absolute value is  $0.07\mu_B$  for FePt and  $0.06\mu_B$  for FeMnSb.<sup>29</sup>

Orbital moments from first-principle calculations exist for the Ni-, Pd-, and PtMnSb compounds obtained using the full-potential linear muffin-tin orbitals method (FPLMTO).<sup>30</sup> While results for NiMnSb are similar to the present calculations this is not the case for the Pd and Pt atoms in PdMnSb and PtMnSb compounds. FPLMTO predicts that their orbital moment is antiparallel to the spin moment contrary to the present calculations. This difference can arise from the treatment of the spin-orbit coupling. Whilst in the present calculations the Dirac equations are solved, in the case of the FPLMTO study the spin-orbit coupling is treated as a perturbation and since orbital moments are very small this can lead to such small deviations. Safe conclusions could me made only if the KKR calculations were performed in the same way as the FPLMTO ones.

Finally it was shown in Ref. 31 that the orbital moment is proportional to the difference between the number of states

	$m_{\rm spin}^X$	$m_{\rm orbit}^X$	$m_{ m spin}^Y$	$m_{ m orbit}^Y$	$m_{ m spin}^Z$	$m_{\rm orbit}^Z$	$m_{ m spin}^{total}$	$m_{ m orbit}^{total}$	m <sup>total</sup>
Co <sub>2</sub> MnAl	0.745	0.012	2.599	0.013	-0.091	-0	3.998	0.038	4.036
Co <sub>2</sub> MnSi	0.994	0.029	3.022	0.017	-0.078	0.001	4.932	0.076	5.008
Co <sub>2</sub> MnGe	0.950	0.030	3.095	0.020	-0.065	0.001	4.931	0.081	5.012
Co <sub>2</sub> MnSn	0.905	0.038	3.257	0.025	-0.079	0	4.988	0.101	5.089
Co <sub>2</sub> CrAl	0.702	0.012	1.644	0.008	-0.082	0	2.966	0.033	2.999
Co <sub>2</sub> FeAl	1.094	0.045	2.753	0.060	-0.095	-0	4.847	0.149	4.996
Fe <sub>2</sub> MnAl	-0.311	-0.015	2.633	0.014	-0.016	0.001	1.994	-0.014	1.980
Mn <sub>2</sub> VAl	-1.398	-0.034	0.785	-0.009	0.013	0.005	-1.998	-0.073	-2.071
Rh <sub>2</sub> MnAl	0.304	-0.011	3.431	0.034	-0.037	-0.001	4.002	0.011	4.013

TABLE II. Same as Table I for the  $X_2YZ$  full-Heusler compounds.

of majority and minority spin at the Fermi level:  $m_{orbit} \propto n^{\uparrow}(E_{\rm F}) - n^{\downarrow}(E_{\rm F})$ . In the case of the half-metallic systems  $n^{\downarrow}(E_{\rm F}) = 0$  and thus the total orbital moment should be parallel to the total spin moment. This is not the case always as can be seen in Table I. In Ref. 31 it was assumed that the  $t_{2g}$  and  $e_g$  states are degenerate and the local DOS of all atoms is a Lorentzian; thus the applicability of this relation is restricted.

Half-metallic full-Heusler alloys. In the second part of my study I will concentrate on the half-metallic full-Heusler alloys and in Table II I have gathered my results. The orbital moments are quite small like the half-Heuslers. In all cases with the exception of Rh atom in Rh<sub>2</sub>MnAl the Hunds rules are obeyed; note that for V in Mn<sub>2</sub>VAl the spin and orbital moments are antiparallel since V d valence shell is less than half-filled. The orbital moments of the sp atoms (Z sites) are almost zero for all cases as in the half-Heuslers.

The Co<sub>2</sub>Mn-Z type compounds are the most interesting since they present the highest Curie temperature among the known half-metals.<sup>?</sup> The comparison between the Al and Si compounds, which have one valence electrons difference, reveals large changes in their magnetic properties. The Co spin moment increases by nearly  $0.25\mu_B$  and the Co orbital moment follows this change since it is more than double for the Si compound. The increase in the Mn spin moments is proportionally smaller and so do the orbital moments. Substituting now Ge or Sn for Si, which are isovalent systems, has only a weak effect on the spin moments. Co spin moment slightly decreases while the Mn spin moment slightly increases. For both atoms the orbital moments show a small increase with the atomic number.

The next step is to substitute Cr for Mn in Co<sub>2</sub>MnAl. Co spin moment is not affected by this substitution and so does its orbital moment. Thus the Co orbital moment is mostly induced by the spin-orbit coupling at the Co moment and is insensitive to hybridization with the neighboring sites. Cr moments on the other hand have to account for the missing electron and are considerably smaller than the Mn ones. Substituting Fe for Mn in Co<sub>2</sub>MnAl has a more pronounced effect. Co spin moment increases by  $0.35\mu_B$  while its orbital moment is more than tripled. Its also interesting to compare Co<sub>2</sub>FeAl to the isoelectronic Co<sub>2</sub>MnSi. Co spin moment in the case of Co<sub>2</sub>FeAl is slightly larger while the Co orbital moment is increased by ~50%. Comparing Co<sub>2</sub>MnAl with Fe<sub>2</sub>MnAl reveals only small changes at the Mn site and the decrease in the total number for valence electrons is taken care by Fe atoms. Substituting now Rh for Co in the same compound leads to an increase of both the spin and orbital moments of Mn since the hybridization between Mn and Rh *d* states is considerably smaller than between the Mn and Co *d* states. Finally I also calculated the properties of Mn<sub>2</sub>VAl. The increased hybridization between the Mn and its neighboring Mn and V atoms leads to a large orbital moment at the Mn site although its spin moment is halved with respect to the cases above where Mn occupied the Y site.

To my knowledge, calculations of the orbital moment exist only by Picozzi *et al.*<sup>19</sup> for the Co<sub>2</sub>Mn-Si, -Ge, and -Sn compounds. The orbital moment at the Co site was found to be around  $0.02\mu_B$  and at the Mn site around  $0.008\mu_B$ . These moments are slightly smaller than my values. The differences can arise from the treatment of the spin-orbit coupling as perturbation in their calculations.

Experiments. Few experiments dedicated to the orbital magnetism exist on these compounds. These experiments involve the obtaining of the XMCD spectra of thin films. XMCD is the difference between the absorption spectra for left- and right-circular polarized light involving 2p core states excitations towards unoccupied d states. Elmers and collaborators<sup>21</sup> derived orbital moments of  $0.12\mu_B$  for Co,  $0.04\mu_B$  for Cr and  $0.33\mu_B$  for Fe in the case of a Co<sub>2</sub>Cr<sub>0.6</sub>Fe<sub>0.4</sub>Al thin film. If I compare these values with my calculations for the Co<sub>2</sub>CrAl and Co<sub>2</sub>FeAl compounds they are one order of magnitude larger. LDA usually gives orbital moments only halved with respect to experiments.<sup>22</sup> Also the XMCD derived spin moments are half of the theoretical predicted values. On the other hand, Kimura et al.<sup>20</sup> studied the NiMnSb and PtMnSb films and found that  $m_{orbit}^{total}/m_{spin}^{total}$ < 0.05 while in my calculations this ratio is around 0.01. The spin moments derived by Kimura et al. experiments are also comparable to the theoretical results. Thus the deviation between the present theoretical results and the experiments in Ref. 20 is considerably smaller than when compared to the ones in Ref. 21.

In both sets of experiments the orbital and spin moments are derived by applying the sum rules to the XMCD spectra. The sum rules have been derived using an ionic model<sup>32</sup> and their application to itinerant systems, in particular to low symmetry systems, is strongly debated<sup>33</sup> since XMCD

probes mainly the region near the surface of a film. Thus their application to experimental spectra is not straightforward. Elmer's and collaborators sum-rule derived total spin moment is halved not only with respect to the theoretical results but most importantly also with respect to the value derived from the SQUID measurements. This inconsistency even between XMCD and SQUID measurements on the same sample shows that the application of sum rules to derive the moments in the case of XMCD experiments on films is not really adequate.

Summary. I have studied the orbital magnetism in the half-metallic half- and full-Heusler alloys using the Dirac

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- <sup>1</sup>I. Žutić, J. Fabian, and S. Das Sarma, Rev. Mod. Phys. **76**, 323 (2004).
- <sup>2</sup>J. de Boeck, W. van Roy, J. Das, V. Motsnyi, Z. Liu, L. Lagae, H. Boeve, K. Dessein, and G. Borghs, Semicond. Sci. Technol. 17, 342 (2002).
- <sup>3</sup>R. A. de Groot, F. M. Mueller, P. G. van Engen, and K. H. J. Buschow, Phys. Rev. Lett. **50**, 2024 (1983).
- <sup>4</sup>I. Galanakis, P. H. Dederichs, and N. Papanikolaou, Phys. Rev. B 66, 134428 (2002).
- <sup>5</sup>M. Zhang, X. Dai, H. Hu, G. Liu, Y. Cui, Z. Liu, J. Chen, J. Wang, and G. Wu, J. Phys.: Condens. Matter **15**, 7891 (2003); M. Zhang, Z. Liu, H. Hu, G. Liu, Y. Cui, G. Wu, E. Brück, F. R. de Boer, and Y. Li, J. Appl. Phys. **95**, 7219 (2004).
- <sup>6</sup>I. Galanakis, P. H. Dederichs, and N. Papanikolaou, Phys. Rev. B **66**, 174429 (2002).
- <sup>7</sup>S. Fujii, S. Ishida, and S. Asano, J. Phys. Soc. Jpn. **64**, 185 (1995); S. Ishida, S. Fujii, S. Kashiwagi, and S. Asano, *ibid*. **64**, 2152 (1995).
- <sup>8</sup>Y. Miura, K. Nagao, and M. Shirai, Phys. Rev. B **69**, 144413 (2004).
- <sup>9</sup>I. Galanakis, J. Phys.: Condens. Matter 16, 3089 (2004).
- <sup>10</sup>R. J. Soulen, Jr., J. M. Byers, M. S. Osofsky, B. Nadgorny, T. Ambrose, S. F. Cheng, P. R. Broussard, C. T. Tanaka, J. Nowak, J. S. Moodera, A. Barry, and J. M. D. Coey, Science **282**, 85 (1998).
- <sup>11</sup>H. Kato, T. Okuda, Y. Okimoto, Y. Tomioka, K. Oikawa, T. Kamiyama, and Y. Tokura, Phys. Rev. B **69**, 184412 (2004).
- <sup>12</sup>T. Shishidou, A. J. Freeman, and R. Asahi, Phys. Rev. B 64, 180401 (2001).
- <sup>13</sup>I. Galanakis, Phys. Rev. B **66**, 012406 (2002); I. Galanakis and Ph. Mavropoulos, *ibid.* **67**, 104417 (2003).
- <sup>14</sup>S. Sanvito and N. A. Hill, Phys. Rev. B **62**, 15 553 (2000); B. Sanyal, L. Bergqvist, and O. Eriksson, *ibid*. **68**, 054417 (2003);
  W.-H. Xie, B.-G. Liu, and D. G. Pettifor, *ibid*. **68**, 134407 (2003); M. Zhang, H. Hu, G. Liu, Y. Cui, Z. Liu, J. Wang, G. Wu, X. Zhang, L. Yan, H. Liu, F. Meng, J. Qu, and Y. Li, J. Phys.: Condens. Matter **15**, 5017 (2003); J.-C. Zheng and J. W. Davenport, Phys. Rev. B **69**, 144415 (2004).
- <sup>15</sup>H. Akinaga, T. Manago, and M. Shirai, Jpn. J. Appl. Phys., Part 2 39, L1118 (2000); M. Shirai, J. Appl. Phys. 93, 6844 (2003).
- <sup>16</sup>A. Stroppa, S. Picozzi, A. Continenza, and A. J. Freeman, Phys. Rev. B 68, 155203 (2003).
- <sup>17</sup>H. Akai, Phys. Rev. Lett. **81**, 3002 (1998).

formalism within the framework of the Korringa-Kohn-Rostoker Green's function method. The quenching of the orbital moments is pretty complete and their values are very small with respect to the spin moments. The change in the atomic-resolved orbital moments can be easily explained in terms of the spin-orbit strength and hybridization effects. Moments derived by applying the sum rules to the experimental x-ray dichroic spectra of thin films should be treated with caution.

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- <sup>18</sup>E. Kulatov and I. I. Mazin, J. Phys.: Condens. Matter 2, 343 (1990); S. V. Halilov and E. T. Kulatov, *ibid.* 3, 6363 (1991); S. J. Youn and B. I. Min, Phys. Rev. B 51, 10 436 (1995); V. N. Antonov, P. M. Oppeneer, A. N. Yaresko, A. Ya. Perlov, and T. Kraft, *ibid.* 56, 13 012 (1997).
- <sup>19</sup>S. Picozzi, A. Continenza, and A. J. Freeman, Phys. Rev. B 66, 094421 (2002).
- <sup>20</sup>A. Kimura, S. Suga, T. Shishidou, S. Imada, T. Muro, S. Y. Park, T. Miyahara, T. Kaneko, and T. Kanomata, Phys. Rev. B 56, 6021 (1997).
- <sup>21</sup>H. J. Elmers, G. H. Fecher, D. Valdaitsev, S. A. Nepijko, A. Gloskovskii, G. Jakob, G. Schonhense, S. Wurmehl, T. Block, C. Felser, P.-C. Hsu, W.-L. Tsai, and S. Cramm, Phys. Rev. B 67, 104412 (2003).
- <sup>22</sup>W. Grange, I. Galanakis, M. Alouani, M. Maret, J.-P. Kappler, and A. Rogalev, Phys. Rev. B **62**, 1157 (2000).
- <sup>23</sup>N. Papanikolaou, R. Zeller, and P. H. Dederichs, J. Phys.: Condens. Matter 14, 2799 (2002).
- <sup>24</sup>S. H. Vosko, L. Wilk, and N. Nusair, Can. J. Phys. 58, 1200 (1980).
- <sup>25</sup> Ph. Mavropoulos, K. Sato, R. Zeller, P. H. Dederichs, V. Popescu, and H. Ebert, Phys. Rev. B **69**, 054424 (2004); Ph. Mavropoulos, I. Galanakis, V. Popescu, and P. H. Dederichs, J. Phys.: Condens. Matter **16**, 4261 (2004).
- <sup>26</sup>D. Orgassa, H. Fujiwara, T. C. Schulthess, and W. H. Butler, Phys. Rev. B **60**, 13 237 (1999); S. Picozzi, A. Continenza, and A. J. Freeman, *ibid*. **69**, 094423 (2004).
- <sup>27</sup>H. Ebert, Rep. Prog. Phys. **59**, 1665 (1996).
- <sup>28</sup>I. Galanakis, P. Ravindran, P. M. Oppeneer, L. Nordström, P. James, M. Alouani, H. Dreyssé, and O. Eriksson, Phys. Rev. B 63, 172405 (2001).
- <sup>29</sup>I. Galanakis, M. Alouani, and H. Dreyssé, Phys. Rev. B **62**, 6475 (2000).
- <sup>30</sup>I. Galanakis, S. Ostanin, M. Alouani, H. Dreyssé, and J. M. Wills, Phys. Rev. B **61**, 4093 (2000).
- <sup>31</sup> H. Ebert, R. Zeller, B. Drittler, and P. H. Dederichs, J. Appl. Phys. 67, 4576 (1990).
- <sup>32</sup> P. Carra, B. T. Thole, M. Altarelli, and X. Wang, Phys. Rev. Lett. **70**, 694 (1993); B. T. Thole, P. Carra, F. Sette, and G. van der Laan, *ibid.* **68**, 1943 (1992).
- <sup>33</sup>C. T. Chen, Y. U. Idzerda, H.-J. Lin, N. V. Smith, G. Meigs, E. Chaban, G. H. Ho, E. Pellegrin, and F. Sette, Phys. Rev. Lett. **75**, 152 (1995); R. Wu, D. Wang, and A. J. Freeman, *ibid.* **71**, 3581 (1993); R. Wu and A. J. Freeman, *ibid.* **73**, 1994 (1994).