

Modulational instability and energy localization in anharmonic lattices at finite energy density

Yuriy A. Kosevich*

Max Planck Institute for Physics of Complex Systems, Noethnitzer Str. 38, D-01187 Dresden, Germany

Stefano Lepri[†]

Dipartimento di Energetica "S. Stecco," Via Santa Marta 3, I-50139 Florence, Italy

(Received 7 June 1999)

The localization of vibrational energy, induced by the modulational instability of the Brillouin-zone-boundary mode in a chain of classical anharmonic oscillators with finite initial energy density, is studied within a continuum theory. We describe the initial localization stage as a gas of envelope solitons and explain their merging, eventually leading to a single localized object containing a macroscopic fraction of the total energy of the lattice. The initial-energy-density dependences of all characteristic time scales of the soliton formation and merging are described analytically. Spatial power spectra are computed and used for the quantitative explanation of the numerical results.

I. INTRODUCTION

Despite a large number of studies on localization of energy in pure anharmonic lattices,¹⁻⁴ the mechanism of their thermal generation is still under debate. This is clearly a central issue being directly related with the possibility of observing signature of such a class of excitations in experiments on real crystals. The problem of nanoscale energy localization has turned out to be rather important also in view of its possible application for different physical and biological systems. The theoretical studies at finite energy per particle dealt mainly with one-dimensional models belonging both to the Fermi-Pasta-Ulam (FPU) class of models^{5,6} (with acoustic spectrum in the harmonic limit) and to the Klein-Gordon one⁷ (with optical-phonon-type spectrum in the harmonic limit). Some attention has been also recently paid to the discretized version of the nonlinear Schrödinger equation in one dimension.^{8,9}

Of special interest is the numerically observed fact that excitation of short-wavelength extended modes of the chain and their instability with respect to perturbations with non-vanishing wave number leads generically to creation of localized states. The phenomenology emerging from the numerical studies^{5,6} can be summarized in three main stages. In the first one, after the development of the modulational instability, energy localizes along the chain in the form of an array of almost standing "bumps." After a given time (that increase upon decreasing the energy) those begin to merge diminishing their number and localization lengths. This second stage (which is reminiscent of the phenomenon of "soliton turbulence"¹⁰) ends with the emergence of a single localized and moving object (a "breather"), carrying along a macroscopic fraction of the total energy. In the third stage, the breather continues to slow down radiating its energy towards the background until it eventually disappears and the equipartition (thermal equilibrium) state is attained.

The aim of this paper is to give a consistent quantitative explanation of such behavior, in particular of the apparent paradoxical tendency of the system to increase localization as time goes by. Namely, we would like to clarify the role of

the finite vibrational energy density on the formation of strongly localized envelope solitons and breathers and their temporal evolution in the anharmonic FPU chain. This is accomplished by means of a suitable continuum approximation for the lattice dynamics. The approach in terms of the envelope-function equation including higher-order nonlinear terms has been already introduced to describe short-wavelength excitations in the FPU chain¹¹ and is reviewed in Sec. II. It is a useful framework both for computing localized solution in the form of envelope soliton (ES) and breather (Sec. III), and for describing the modulational instability (Sec. IV). Furthermore, going beyond the usual integrable approximation¹² is of crucial importance to explain the localization process (Secs. V and VI). We described analytically the initial-energy-density dependence of all characteristic time scales of the envelope soliton formation and merging. Particular emphasis has been also given to the analytic computation and numerical measurements of the spatial power spectra ("structure functions") that, at least in principle, are of importance for real experiments. Indeed, they turn out to be extremely useful for the physical interpretation of the simulation results (see, e.g., Secs. V and VII). Finally, a brief discussion of the finite-temperature effects is reported in Sec. VIII.

II. ENVELOPE-FUNCTION EQUATION FOR THE FPU MODEL

We start from the classical FPU Lagrangian of a monatomic chain of N particles of mass m with anharmonic nearest-neighbor potential of order $r \geq 3$:

$$\mathcal{L} = \sum_n \left[\frac{1}{2} m \dot{u}_n^2 - \sum_{\gamma=2}^r \frac{K_\gamma}{\gamma} (u_{n+1} - u_n)^\gamma \right], \quad (1)$$

where u_n is the (real scalar) displacement of the n th particle from its equilibrium position and K_γ are the harmonic ($\gamma = 2$) and anharmonic force constants. The lowest order potential describing the anharmonicity of the longitudinal or pure transverse motion in a centrosymmetric 1D lattice cor-

responds, respectively, to $r=3$ or $r=4$. From Eq. (1) we obtain the following equations of motion:

$$m\ddot{u}_n = \sum_{\gamma=2}^r K_\gamma [(u_{n+1} - u_n)^{\gamma-1} - (u_n - u_{n-1})^{\gamma-1}]. \quad (2)$$

In the following, periodic boundary conditions are always assumed.

In order to describe the short-wavelength excitations of the chain with wave number $k \sim \pi/a$ (a being the lattice spacing), i.e., near the Brillouin-zone boundary, it is convenient to introduce an envelope-function of the displacements $f_n = (-1)^n u_n$, $f_n \equiv f(x/a)$. The latter is assumed to be slowly varying on the interatomic scale: $af_x \ll f$. Substituting in Eq. (2) the expansion of the differences $u_{n\pm 1} - u_n$ up to the fourth order, one can obtain a nonlinear partial differential equation for f .¹¹ For instance, in the case of the purely quartic anharmonicity (i.e., the so-called β -FPU model with only K_2 and K_4 different from zero) such an equation reads

$$m\ddot{f} + 4K_2 f + K_2 a^2 f_{xx} + \frac{1}{12} K_2 a^4 f_{xxxx} + 16K_4 f^3 + 6K_4 a^2 f(f^2)_{xx} = 0, \quad (3)$$

where the subscript x stands for the spatial derivative of the corresponding order. A equation, similar to Eq. (3), can be also obtained for the cubic anharmonicity (α -FPU model with $\gamma=2$ and $\gamma=3$) and, more generally, for the FPU chain (1) with the anharmonic potential of the arbitrary (even or odd) order r .¹³

Due to the last term in left-hand side of Eq. (3), the latter substantially differs from the conventional nonlinear Klein-Gordon and its nonrelativistic version—the nonlinear Schrödinger equation. Actually, Eq. (3) can be reduced to the latter only in the limit of relatively small oscillation amplitudes (see Sec. IV below). Equation (3) differs also from the well-known Korteweg–de Vries equation which describes the evolution of the long-wavelength (i.e., the ‘‘acoustic-phonon-type’’) excitations and can be derived from Eq. (1) in the limit $au_x \ll 1$ (see, e.g., Ref. 13).

In what follows, we will use Eq. (3) for the adequate *analytic* description of all the main stages of modulation-instability-induced dynamic evolution of the hard (with $K_4 > 0$) quartic FPU chain. For simplicity of notation, we set K_2 , K_4 , a , and m to unity. With such a choice, the (adimensional) energy density (i.e., the energy per particle) ϵ is the only relevant physical parameter if one does not account for the finite-temperature effects (see Sec. VIII below). Moreover, we will always deal with the (classical) low-temperature limit that in our units corresponds to $\epsilon \ll 1$.

III. LOCALIZED SOLUTIONS: ENVELOPE SOLITONS AND BREATHERS

As a first step, we determine the form of the single ES (or periodic array of solitons) by looking for solutions of Eq. (3) in the monochromatic form $f(x, t) = \phi(x) \cos \omega t$. The frequency ω can be computed by first substituting such an ansatz in Eq. (3) and invoking then the so-called rotating wave approximation (RWA), namely, to replace $\cos^3 \omega t$ by $\frac{3}{4} \cos \omega t$. It is worth mentioning that such an approximation

is expected to hold for the considered short-wavelength vibrations due to the weakness of nonresonant interaction between the mode with fundamental frequency and its third harmonic. By further neglecting the fourth-order dispersive term in Eq. (3) we find its first integral as¹¹

$$(4 - \omega^2) \phi^2 + (1 + 9\phi^2)(\phi_x)^2 + 6\phi^4 = \text{const}. \quad (4)$$

The generic solution of Eq. (4) for a nonvanishing constant corresponds to a periodic array of ESs, and can be conveniently parametrized by the two extreme values of the amplitude ϕ_{\min} and ϕ_{\max} (so that $\phi_{\min} \leq \phi \leq \phi_{\max}$). Those are simply determined by the conditions that ϕ_x vanishes for $\phi = \phi_{\min}$ and $\phi = \phi_{\max}$. From those two conditions we derive immediately the frequency of the envelope-soliton array as a function of the extreme values of its amplitude:

$$\omega^2 = 4 + 6(\phi_{\min}^2 + \phi_{\max}^2). \quad (5)$$

Then from Eqs. (4),(5) we can determine the desired solution as

$$\left(\frac{d\phi}{dx}\right)^2 = 6 \frac{(\phi_{\max}^2 - \phi^2)(\phi^2 - \phi_{\min}^2)}{1 + 9\phi^2}. \quad (6)$$

Equation (6) can be solved by quadratures yielding the following implicit form for $\phi(x)$:

$$|x - x_0| = \frac{1}{\sqrt{6}} \int_{\phi}^{\phi_{\max}} \frac{\sqrt{1 + 9\phi'^2}}{\sqrt{(\phi_{\max}^2 - \phi'^2)(\phi'^2 - \phi_{\min}^2)}} d\phi',$$

$$\phi(x_0) = \phi_{\max}. \quad (7)$$

The single ES corresponds to $\text{const}=0$ in Eq. (4) and to $\phi_{\min}=0$ in Eq. (7). Although a general explicit form of Eq. (7) is not feasible, we can at least discuss the two limiting cases of small and large amplitudes. In the small-amplitude limit ($9\phi_{\max}^2 \ll 1$), the solution is given by (we include explicitly the time dependence)

$$f_s(x, t) = \phi_{\max} \frac{\cos(\omega_s t + \varphi)}{\cosh[(x - x_0) \phi_{\max} \sqrt{6}]}, \quad (8)$$

where the frequency as given by Eq. (5) with $\phi_{\min}=0$ and is approximately $\omega_s = 2 + 3\phi_{\max}^2/2$, whereas φ is the initial phase.

In the large-amplitude limit ($9\phi_{\max}^2 \gg 1$), it follows from Eq. (7) that the ES acquires a sinusoidal shape with a short amplitude-independent localization length¹¹

$$f_b(x, t) = \phi_{\max} \cos(\omega_b t + \varphi) \cos[k_b(x - x_0)]$$

$$\text{for } k_b |x - x_0| < \frac{\pi}{2} \quad (9)$$

and with $f_b \approx 0$ outside of the above specified domain. The effective wave number is $k_b = \sqrt{2/3}$ and ω_b given by Eq. (5) with $\phi_{\min}=0$. This is rather accurate approximation for the exact breather solution of the discrete FPU model (see Ref. 13 for a detailed comparison). Discrepancies have to be expected because, strictly speaking, the continuum approach is not completely justified in this short-wavelength limit. In-

deed, the analysis of the corresponding discrete-lattice equations predicts¹⁴ the existence of intrinsic localized modes (discrete breathers) with (approximately) sinusoidal envelope (9) but with $k_b = \pi/3$ and $\omega_b^2 = 3 + (81/16)\phi_{\max}^2$.

Having at hand the explicit solution (8), we can compute the spatial Fourier transform of the displacement field u associated to the single ES obtaining (notice that, by definition, the zero wave number for f corresponds to the wave numbers $\pm\pi$ for u and therefore $u(k) = \int u(x)e^{ikx}dx = \int f(x)e^{ikx} \cos \pi x dx = \frac{1}{2}[f(k-\pi) + f(k+\pi)]$)

$$u_s(k, t) = \frac{1}{2} \cos(\omega_s t + \varphi) [F_s(k - \pi, x_0) + F_s(k + \pi, x_0)], \quad (10)$$

where we have defined the form factor

$$F_s(k, x_0) = \frac{\pi}{\sqrt{6}} \frac{e^{ikx_0}}{\cosh[\pi k / \phi_{\max} 2\sqrt{6}]}. \quad (11)$$

Accordingly, the power spectrum [i.e., the square modulus of $u_s(k, t)$] is exponentially localized around $k = \pm\pi$ and its maximal value is independent of the amplitude ϕ_{\max} .

For later purposes, we compute also the total energy E_s of the ES. To this aim, one can use the expression for the Hamiltonian density h by the variation of which Eq. (3) can be derived¹¹

$$h = \frac{1}{2}(\dot{f})^2 + 2f^2 + 4f^4 - \frac{1}{2}(1 + 12f^2)(f_x)^2 + \frac{1}{24}(f_{xx})^2 \quad (12)$$

(remember that we are working in dimensionless units). Neglecting again the higher-order dispersive term and averaging h over time (in the RWA) and by means of Eq. (4) with $\text{const}=0$, we find that the energy density h_s associated with the ES is $2\phi^2 + 3\phi^4$. Using this expression together with Eq. (8), we finally find

$$\begin{aligned} E_s &= \int_{-\infty}^{+\infty} h_s dx = \int_{-\infty}^{+\infty} [2\phi^2 + 3\phi^4] dx \\ &= \frac{4}{\sqrt{6}} \phi_{\max} (1 + \phi_{\max}^2). \end{aligned} \quad (13)$$

In the leading approximation, the energy scales linearly with the amplitude.¹⁰

The Fourier transform of the breather solution (9) is

$$u_b(k, t) = \frac{1}{2} \cos(\omega_b t + \varphi) [F_b(k - \pi, x_0) + F_b(k + \pi, x_0)], \quad (14)$$

where we have defined

$$F_b(k, x_0) = \phi_{\max} e^{ikx_0} \frac{2k_b}{k_b^2 - k^2} \cos\left(\frac{\pi k}{2k_b}\right). \quad (15)$$

Notice that this function is finite for $k = k_b$. In particular, for $k_b = \pi/3$ the Fourier transform of the breather solution has the following characteristic form:

$$\begin{aligned} u_b(k, t) &= \frac{\pi}{3} \phi_{\max} \cos(\omega_b t + \varphi) \sin\left(\frac{3}{2}k\right) e^{ikx_0} \\ &\times \left[\frac{e^{-i\pi x_0}}{(k - 4\pi/3)(k - 2\pi/3)} \right. \\ &\left. - \frac{e^{i\pi x_0}}{(k + 4\pi/3)(k + 2\pi/3)} \right]. \end{aligned} \quad (16)$$

The power spectrum has basically a sinusoidal shape with two absolute maxima at $k = \pm\pi$ whose height is proportional to the square of the breather amplitude.

IV. MODULATIONAL INSTABILITY

In this section, we describe the modulational instability of the Brillouin-zone-boundary mode corresponding to the uniform envelope $f(x, t) \equiv f_0(t) = A \cos \omega t$. To compute the frequency ω one can resort again to the RWA mentioned in the previous section. In the small-amplitude limit $3A^2 \ll 1$ one obtains $\omega = 2 + 3A^2$, which amounts to a correction to the zone-boundary frequency $\omega_{\max} = 2\sqrt{K_2/m} \equiv 2$.

Let us consider a spatially inhomogeneous small perturbation $\delta f \propto \cos k_m x$, where the wave number k_m of the envelope-function modulation is assumed to be much smaller than the Brillouin-zone width π , an hypothesis that will be confirmed *a posteriori*. Then, from the linearization of Eq. (3) we find the equation for δf to be

$$\delta \ddot{f} + [4 - k_m^2 + 24A^2(1 + \cos 2\omega t)] \delta f = 0. \quad (17)$$

To obtain the growth rate s at the parametric resonance, described by Eq. (17), we take a solution of this equation as (see, e.g., Ref. 15),

$$\delta f = [a \cos \omega t + b \sin \omega t] e^{st} \cos k_m x, \quad (18)$$

where the unknown amplitudes a and b are assumed to be small with respect to A . Therefore, from Eqs. (17) and (18) we obtain the dispersion equation for the growth rate s :

$$s^4 + 2s^2[8 + 36A^2 - k_m^2] = k_m^2(24A^2 - k_m^2). \quad (19)$$

The maximal growth rate s_{\max} and the corresponding wave number k_m^* of the envelope-function modulational instability are thus given in the small-amplitude limit $3A^2 \ll 1$ by

$$k_m^* = A\sqrt{12}, \quad s_{\max} = 3A^2, \quad b = -a, \quad (20)$$

which, consistently, describes a long wavelength modulation with $k_m^*/2 \ll 1$.¹⁶ In the considered limit, such a value of s_{\max} coincides with the one computed from the stability analysis of the discrete-lattice equations.¹⁷ It is also worth to mention that there is a minimal (threshold) amplitude A_c for the modulational instability of the Brillouin-zone-boundary mode in a finite FPU chain of N particles, and $A_c \propto 1/N$ for $N \gg 1$, see, e.g. Ref. 6. We neglect such finite-size effects in our continuum approach which implies, in particular, that the amplitude A in Eq. (20) is well above the threshold one for the considered (finite) chain.

By means of Eq. (3) we can also describe the interesting phenomenon of the further appearance of spatial harmonics

of the ‘‘primary’’ perturbation, i.e., the components with wave numbers nk_m^* and $n=0,2,3,\dots$. To describe, for instance, the growth of the zeroth and second spatial harmonics, we assume δf to be of the following form [see Eqs. (18) and (20)]:

$$\begin{aligned}\delta f &= a[\cos \omega t - \sin \omega t]e^{s_{\max}t} \cos k_m^* x \\ &\quad + [c \cos \omega t + d \sin \omega t]e^{2s_{\max}t} \cos 2k_m^* x \\ &\quad + [e \cos \omega t + g \sin \omega t]e^{2s_{\max}t} \\ &\equiv \delta f_{k_m^*} + \delta f_{2k_m^*} + \delta f_0,\end{aligned}\quad (21)$$

where the amplitudes c , d , e , and g are assumed to be much smaller than both the condensate amplitude A and the amplitude a of the primary perturbation. Using Eq. (3), we obtain the following evolution equations for $\delta f_{2k_m^*}$ and δf_0 similar to Eq. (17) with a driving term $-48f_0(\delta f_{k_m^*})^2$ on the right-hand side:

$$\begin{aligned}\delta \dot{f}_{2k_m^*} + [4 - 4k_m^{*2} + 24A^2(1 + \cos 2\omega t)]\delta f_{2k_m^*} \\ = -12a^2A[2 \cos \omega t - \sin \omega t]e^{2s_{\max}t} \cos 2k_m^* x,\end{aligned}\quad (22)$$

$$\begin{aligned}\delta \dot{f}_0 + [4 + 24A^2(1 + \cos 2\omega t)]\delta f_0 \\ = -12a^2A[2 \cos \omega t - \sin \omega t]e^{2s_{\max}t}.\end{aligned}\quad (23)$$

From Eqs. (21)–(23) we find the amplitudes c , d , e , and g :

$$c = \frac{a^2}{2A}, \quad d = e = -\frac{a^2}{2A}, \quad g = -\frac{3a^2}{4A}.\quad (24)$$

Therefore the zeroth δf_0 and second $\delta f_{2k_m^*}$ spatial harmonics grow exponentially with a rate which is two times larger than that of the primary instability $\delta f_{k_m^*}$. Nevertheless, as their initial amplitudes are of second order with respect to the amplitude a of the primary instability, one expects that they will come into play only at a later stage.

More generally, we can show that the growth rate of the n th spatial harmonic $\delta f_{nk_m^*}$ of the primary instability is equal to ns_{\max} while its initial amplitude is proportional to the n th power of the amplitude a . Furthermore, it can be shown that small contributions to the growth of the zeroth harmonic δf_0 are induced by the higher-order harmonics. These arise from terms such as $(\delta f_{2k_m^*})^2 f_0$, $(\delta f_{k_m^*})^2 \delta f_{2k_m^*}$, etc., to the right-hand side of Eq. (23).

The above scenario is illustrated in Fig. 1, where the outcomes of a numerical simulation of the quartic FPU model are reported. The equations of motion were integrated with a third-order symplectic algorithm¹⁸ (microcanonical simulation) starting from the initial conditions

$$u_n(0) = 0, \quad \dot{u}_n(0) = (-1)^n \sqrt{2\epsilon}.\quad (25)$$

Actually, a small Gaussian white noise (with amplitude $\ll \sqrt{2\epsilon}$) has also been added to the initial velocities in order to speed up the instability. The left panels of Fig. 1 show the evolution of the energy density

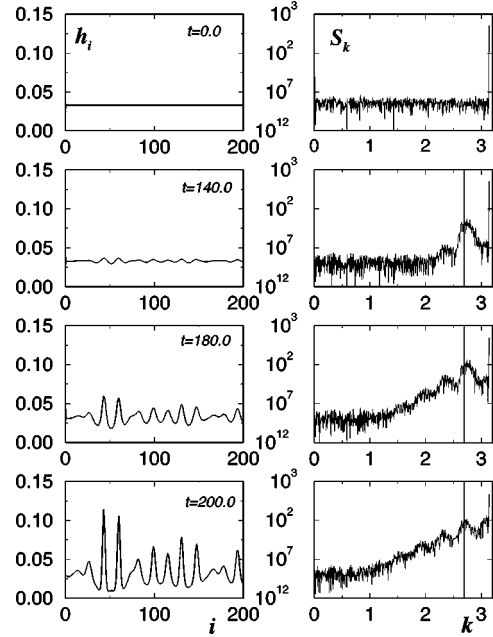


FIG. 1. Modulational instability of the boundary mode for the quartic FPU model with $\epsilon = 0.0330$, $N = 1024$. The left panels are snapshots of the energy density, the right ones the spatial Fourier spectrum of the velocities. The vertical lines correspond to the theoretical value of k_m^* given by Eq. (20).

$$h_n = \frac{1}{2} \dot{u}_n^2 + \frac{1}{2} \sum_{\gamma=2,4} \frac{1}{\gamma} [(u_{n+1} - u_n)^\gamma + (u_n - u_{n-1})^\gamma],\quad (26)$$

at subsequent times. The spatial power spectrum of the velocities

$$S_k = \frac{1}{N} \left| \sum_{n=1}^N \dot{u}_n e^{ikn} \right|^2, \quad k = 0, \pm \frac{2\pi}{N}, \pm \frac{4\pi}{N}, \dots, \pm \pi\quad (27)$$

shows the growth of the perturbation with the wave numbers $\pi - k_m^*$ (vertical line). Obviously a symmetric peak at $-\pi + k_m^*$ is also present as well as the further transfer of the energy towards the spatial harmonics with wave numbers nk_m^* , $n=0,2,3,\dots$, according to the above described mechanism. We also checked that the typical time for developing the instability scales as $A^{-2} \epsilon^{-1}$ as prescribed by Eq. (20).

V. THE FIRST LOCALIZATION STAGE: A SOLITON GAS

As already mentioned, the modulational instability produces a strong localization of the vibrational energy along the chain (see Fig. 2). We wish now to describe such a set of localized objects. More precisely we will show that it is basically a gas of ESs and study some of its properties.

Let us consider a set of N_s nonoverlapping ESs of the form (8) centered at the points x_l along the line and with initial phases φ_l . For the sake of simplicity, let us also assume them to have approximately the same amplitude. According to Eq. (10), the spatial Fourier transform of the corresponding displacement field will be

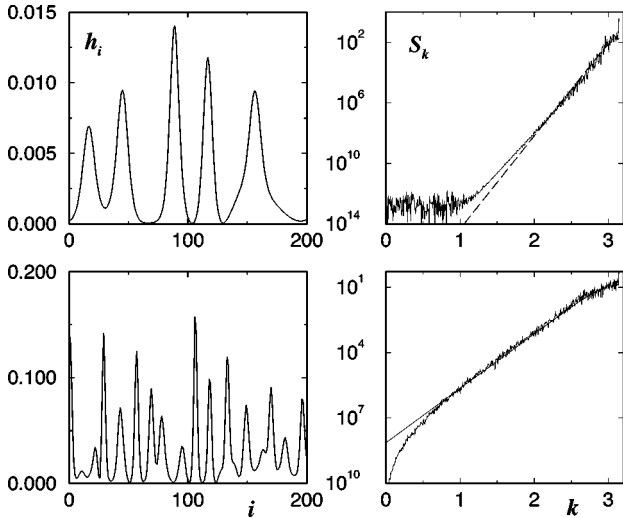


FIG. 2. The soliton gas generated after the instability for $N = 1024$ for $\epsilon = 0.0033$, $t = 8000$ (upper panels) and $\epsilon = 0.0330$, $t = 800$ (lower panels). The corresponding power spectra S_k of the particles' velocities (right panels) averaged over a time interval ≈ 100 . The dashed lines are the best fit with the formula (34).

$$u(k, t) = \frac{1}{2} \sum_{l=1}^{N_s} \cos(\omega_s t + \varphi_l) [F_s(k - \pi, x_l) + F_s(k + \pi, x_l)]. \quad (28)$$

The velocity power spectrum S_k as defined by Eq. (27) in the discrete case, can thus be approximated as (we use the explicit form of F_s)

$$S_k = \frac{1}{N} \langle |u(k, t)|^2 \rangle = \frac{\pi^2 \omega_s^2}{48N} \times \left\{ \frac{1}{\cosh^2[\pi(\pi - k)/\phi_{\max} 2\sqrt{6}]} \left| \sum_l e^{i\varphi_l + i(k - \pi)x_l} \right|^2 + \frac{1}{\cosh^2[\pi(k + \pi)/\phi_{\max} 2\sqrt{6}]} \left| \sum_l e^{i\varphi_l - i(k + \pi)x_l} \right|^2 \right\}, \quad (29)$$

where the angle brackets denote the averaging over a time interval larger than $2\pi/\omega_s$ (that is performed to remove the oscillations as well as statistical fluctuations). In the last expression we neglected exponentially small overlap between the components of Eq. (10) localized, respectively, around $k = \pi$ and $k = -\pi$.

Finding a more explicit expression for S_k requires of course some hypothesis on the statistical properties of the positions x_l and phases φ_l . For instance, in the limit of completely uncorrelated ESs one can replace the square modulus of each sum in Eq. (29) by N_s . To take into account weak partial correlations, we introduce a phenomenological parameter p ($p \sim 1$) and replace the mentioned square modulus by pN_s . Obviously, $p = 1$ corresponds to a completely random case while the cases $p > 1$ or $p < 1$ describe, respectively, the arrays of the partially correlated or anticorrelated ESs. With such a definition

$$S_k = p n_s \frac{\pi^2 \omega_s^2}{48} \left\{ \frac{1}{\cosh^2[\pi(\pi - k)/\phi_{\max} 2\sqrt{6}]} + \frac{1}{\cosh^2[\pi(k + \pi)/\phi_{\max} 2\sqrt{6}]} \right\}, \quad (30)$$

where $n_s = N_s/N$ is a density of the envelope solitons.

To compare the latter expression with the numerically measured S_k , one has to relate ω_s , n_s , and ϕ_{\max} in Eq. (30) with the energy density ϵ . As we deal with the limit $\epsilon \ll 1$, one has that $\omega_s \approx 2$ since $A \approx \sqrt{2\epsilon/2}$ [see Eq. (25) and the definition of the amplitude A]. Moreover, the density n_s will be given by the inverse wavelength of the modulational instability, namely,

$$n_s = \frac{k_m^*}{2\pi} = \frac{1}{\pi} \sqrt{\frac{3\epsilon}{2}}. \quad (31)$$

This amounts to say that a single localized object emerges from each wavelength. A similar reasoning allows us to relate the total energy E_s [which is basically proportional to ϕ_{\max} , see again Eq. (13)] to ϵ as

$$E_s = q \frac{\epsilon}{n_s} = q 2\pi \sqrt{\frac{\epsilon}{6}}. \quad (32)$$

The phenomenological parameter q , $0 \leq q \leq 1$, gives the fraction of the initial vibrational energy which is absorbed by the set of ESs (and actually accounts for the partial overlap between them). In particular, for $\epsilon \rightarrow 0$ we expect the soliton gas to be more and more rarified and to contain almost all of the initial energy in the lattice so that $q \rightarrow 1$.

A consequence of Eqs. (31) and (32) is that the ratio between the total energy of the ESs and their density is (almost) independent on ϵ :

$$\frac{E_s}{n_s} = q \frac{2\pi^2}{3}. \quad (33)$$

Before proceeding further, we numerically checked that the above picture is correct by evaluating the number and energy of solitons as a function of the initial energy density. The results reported in Fig. 3 show a very good agreement with both Eqs. (31) and (33). Notice that, as expected, the data are very well fitted with $q = 1$ at least for $\epsilon < 0.01$.

Finally, using Eqs. (30), (31), (32), and (13), we obtain the desired expression for the velocity power spectrum as a function of ϵ and the (unknown) parameters p and q only:

$$S_k = p \frac{\pi}{16} \sqrt{\frac{\epsilon}{6}} \omega_s^2 \left\{ \frac{1}{\cosh^2[(\pi - k)/q\sqrt{6\epsilon}]} + \frac{1}{\cosh^2[(\pi + k)/q\sqrt{6\epsilon}]} \right\}, \quad (34)$$

where $\omega_s = 2$. We obtained a very good fit of our numerical data, for energy density $\epsilon = 0.0033$, using Eq. (34) with $p = 0.50$ and $q = 1.0$ (see the right panels of Fig. 2). We interpret this by saying that the system is assimilated to a diluted gas of nonoverlapping and partially anticorrelated (i.e., out of phase) ESs. For larger initial energy density $\epsilon = 0.033$ the

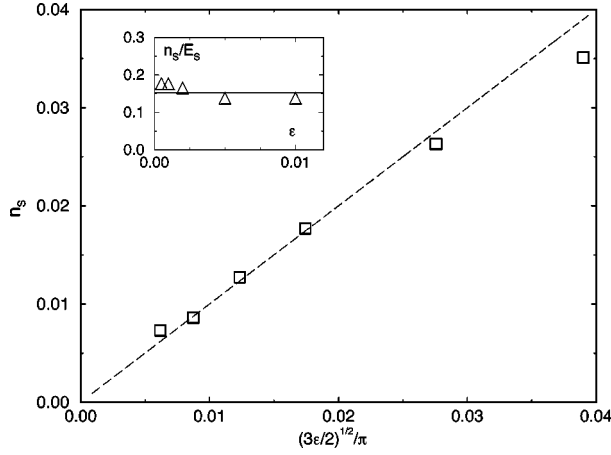


FIG. 3. The maximal initial density n_s of ESs as a function of the energy density for $N=1024$. The solid line corresponds to formula (31). The inset shows the ratio between the number and energy of the solitons: the horizontal line corresponds to Eq. (33) with $q=1$.

data are fitted by $p=1.12$ and $q=0.78$ (lower panels of Fig. 2). This latter case correspond, as expected, to a more dense gas of partially overlapping ESs, which is characterized, in comparison with the previous one, mainly by pair correlations between neighbors.

VI. THE SECOND LOCALIZATION STAGE

To understand why a set of almost nonoverlapping small-amplitude ES finally merge into a large-amplitude one, it is first of all convenient to recast the second order (in time) equation (3) for the real envelope-function in two first order equations. For small-amplitude oscillations with a frequency close to the bandedge $\omega_{\max}=2$ this is accomplished introducing the complex function $\psi(x,t)$ such that

$$f(x,t) = \frac{1}{2} [\psi(x,t) e^{-i\omega_{\max}t} + \psi^*(x,t) e^{i\omega_{\max}t}]. \quad (35)$$

Substituting this expression into Eq. (3) and assuming that $\dot{\psi} \ll \omega_{\max}\psi$, after some algebra one can obtain (in the RWA) the following coupled nonlinear-Schrödinger-type equations for ψ and ψ^*

$$\begin{aligned} -i\omega_{\max}\dot{\psi} + \frac{1}{2}\psi_{xx} + \frac{1}{24}\psi_{xxxx} + 6|\psi|^2\psi \\ + \frac{3}{2}[2\psi(\psi^*\psi_{xx} + |\psi_x|^2) + \psi^2\psi_{xx}^* + \psi^*\psi_x^2] = 0, \\ i\omega_{\max}\dot{\psi}^* + \frac{1}{2}\psi_{xx}^* + \frac{1}{24}\psi_{xxxx}^* + 6|\psi|^2\psi^* \\ + \frac{3}{2}[2\psi^*(\psi\psi_{xx}^* + |\psi_x|^2) + \psi^{*2}\psi_{xx} + \psi\psi_x^{*2}] = 0. \end{aligned} \quad (36)$$

Equations (36) can be expressed in canonical form as

$$i\omega_{\max}\dot{\psi} = \frac{\delta\mathcal{H}}{\delta\psi^*}, \quad -i\omega_{\max}\dot{\psi}^* = \frac{\delta\mathcal{H}}{\delta\psi},$$

where the Hamiltonian $\mathcal{H} = \int H dx$ has the following density H :

$$\begin{aligned} H = -\frac{1}{2} \left[|\psi_x|^2 - \frac{1}{12} |\psi_{xx}|^2 - 6|\psi|^4 + 6 \left(|\psi|^2 |\psi_x|^2 \right. \right. \\ \left. \left. + \frac{1}{4} (\psi^2 \psi_x^{*2} + \psi_x^2 \psi^{*2}) \right) \right]. \end{aligned} \quad (37)$$

One can ascertain that Eq. (36) admit the integrals of motion

$$\begin{aligned} \mathcal{H} = \int H dx, \quad \mathcal{P} = -\frac{i}{2} \int [\psi\psi_x^* - \psi^*\psi_x] dx, \\ \mathcal{N} = \int |\psi|^2 dx, \end{aligned} \quad (38)$$

which are the energy, momentum and number of quanta, respectively.

The inclusion of higher-order and nonlinear dispersive terms make equations (36) nonintegrable (i.e., they do not have an infinite number of constants of motion). This simple fact, together with the conservation laws (38) has deep consequences on the dynamics: it is in fact known¹⁹ that in this case the total number of solitons is not conserved. Actually, they can merge together enhancing their amplitudes and decreasing their number due to an exchange of both energy and number of quanta. Indeed, the Hamiltonian (37) describes an effective *repulsive* interaction between quasiparticles with *negative* (and amplitude-dependent) effective mass and it is therefore equivalent to a system of quasiparticles with *positive* effective mass and effective *attractive* interaction. As the solitons merge, the released binding energy is carried away by free waves and the ‘‘phononic’’ part of the spectrum is gradually populated.

More precisely, the kinematics of the soliton-soliton interaction process requires an increase of the number of quanta in the stronger soliton and a decrease in the weaker one accompanied by the release of some binding energy by means of free plane waves that give also the recoil momentum to the solitons.¹⁹ Accordingly, the reverse process should include a triple collision (two solitons and a free wave) and therefore is much less probable than the direct one. It is therefore clear that eventually only one large-amplitude soliton survives on the background of small-amplitude free waves with a continuous frequency spectrum. This argument explains qualitatively why the purely dynamical process of soliton merging is time irreversible.

The above reasoning can be further carried on to give a more quantitative description of the phenomenology. For instance, it is possible to explain the observed fact that the characteristic time for complete soliton merging in the quartic FPU chain scales with the energy density as ϵ^{-2} (Ref. 6, see also Ref. 5). To this aim, let us estimate the ES lifetime (or relaxation time) τ_{ss} due to the scattering mentioned above. It can be written as $\tau_{ss} = l/\bar{v}_s$, where l and \bar{v}_s are respectively the mean free path and average velocity of the ES during the relaxation process. The mean free path will be

in turn given by $l \sim 1/\sigma_{ss}n_s$, where n_s is a density of ESs and σ_{ss} is the scattering cross section.

If we assume the lifetime to be much larger than the inverse frequency of the soliton ($\tau_{ss} \gg 1/\omega_s$), we can resort to the Born approximation to estimate σ_{ss} . As time goes by and the soliton gas rarifies due to the process of merging, the approximation becomes more and more justified. Within this perturbative treatment and in the low velocity limit, we thus find that σ_{ss} is simply proportional to the square modulus of $\int U_{\text{int}} dx$, where U_{int} is the interaction potential. According to Eq. (12), the latter is proportional to the product of the envelope functions of two small-amplitude ESs: $U_{\text{int}} \propto f_1 f_2$. Assuming for simplicity that both of them have approximately the same amplitude ϕ_{max} , it is easy to ascertain from (8) that the overlap integral scales as ϕ_{max} , similar to the energy of a single small-amplitude ES (13). Therefore, the cross-section σ_{ss} is proportional to ϕ_{max}^2 and, according to Eqs. (32) and (13), this implies that $\sigma_{ss} \propto \epsilon$. Finally, as $n_s \propto \sqrt{\epsilon}$, see Eq. (31), we obtain that $l \propto \epsilon^{-1.5}$.

To estimate \bar{v}_s , we compute first the energy of the moving soliton of Eqs. (36). In the small-amplitude limit, when one can neglect the higher-order and nonlinear dispersive terms, Eqs. (36) admit the exact solution^{1,19}

$$\psi(x,t) = \psi_{\text{max}} \frac{\exp[i(kx - \Omega t)]}{\cosh[(x - V_g t) \psi_{\text{max}} \sqrt{6}]}, \quad (39)$$

where V_g is a group velocity of the soliton, $k = -V_g \omega_{\text{max}}$ is a soliton quasimomentum and $\Omega = (3\psi_{\text{max}}^2 - k^2/2)/\omega_{\text{max}}$. Using Eqs. (37), (38), and (39), we find the soliton number of quanta \mathcal{N}_s , momentum \mathcal{P}_s , and energy \mathcal{E}_s (for $\omega_{\text{max}} = 2$):

$$\mathcal{N}_s = \psi_{\text{max}} \sqrt{\frac{2}{3}}, \quad \mathcal{P}_s = -k\mathcal{N}_s = 2V_g\mathcal{N}_s, \quad (40)$$

$$\mathcal{E}_s = \frac{3}{2}\mathcal{N}_s^3 - 2\mathcal{N}_s V_g^2 = \frac{3}{2}\mathcal{N}_s^3 - \frac{\mathcal{P}_s^2}{2\mathcal{N}_s}.$$

These relations show that small-amplitude moving soliton can be indeed considered as a bound state of \mathcal{N}_s negative-effective-mass quasiparticles (harmonic modes) with a correspondingly *positive* binding energy equal to $\frac{3}{2}\mathcal{N}_s^3$ (see Ref. 20).

If we compare the definition of \mathcal{N}_s in Eq. (40) with expressions (32) and (13) for the energy of the ES (8), we conclude that $\mathcal{N}_s \propto \sqrt{\epsilon}$. Let us now assume that a kind of virial theorem holds for the set of solitons moving along the chain with periodic (or fixed) boundary conditions. This amounts to say that (time) averaged kinetic and potential energies of such finite dynamical system should be proportional (see, e.g., Ref. 15). It then follows from the definition (40) of \mathcal{E}_s that $\langle V_g^2 \rangle \sim \langle \mathcal{N}_s^2 \rangle \propto \epsilon$. The validity of such a relation can be physically justified by observing that, as a result of the exchange of the number of quanta during the scattering processes, the solitons receive the recoil quasimomentum k from free waves which in turn changes the soliton velocities thus providing some form of ‘‘thermalization.’’ From all the above follows that $\bar{v}_s \sim \sqrt{\langle V_g^2 \rangle} \propto \sqrt{\epsilon}$.

Finally, we obtain the estimate $\tau_{ss} = l/\bar{v}_s \propto \epsilon^{-2}$ and correspondingly $\tau_{ss} \gg 1/\omega_s \approx 1/2$, that is fully consistent with the

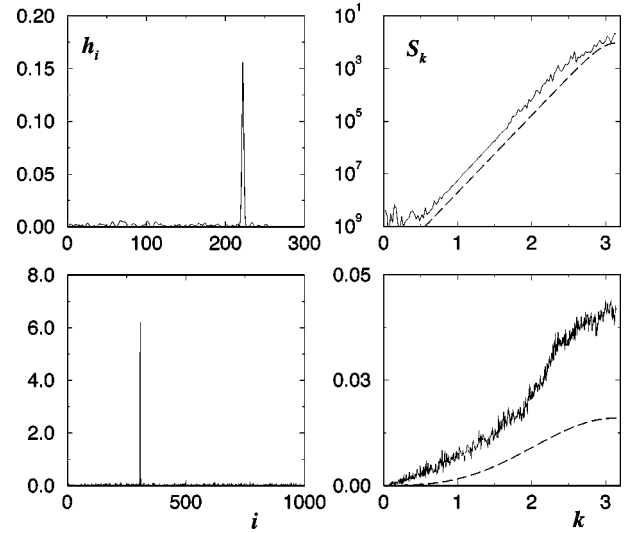


FIG. 4. Final breather for different energies $\epsilon = 0.0033$, $N = 256$, $t = 1.2 \cdot 10^6$ (upper panels) and $\epsilon = 0.0330$, $N = 1024$, $t = 8 \cdot 10^4$ (lower panels). The last spectrum has been averaged over a time 10^4 . Notice the different vertical scales. Dashed lines correspond to Eqs. (10), (11), and (16) with the numerically measured value of ϕ_{max} and $k_b = \pi/3$.

initial assumptions. As already mentioned, this scaling is in agreement with the numerical observation.^{5,6}

VII. THE BREATHER

In the previous section we explained why the scattering favors the creation of a single localized object. We now give a brief account of some of its properties. The h_n patterns reported in the left panels of Figs. 4 and 5 illustrate the remarkable fact that a macroscopic fraction of the total energy accumulates on a few sites. For instance, for the case $\epsilon = 0.033$ and $N = 1024$, only *three* sites bear an energy 10.9 corresponding to about 30% of the total. As expected from the previous section, the remaining is distributed along the

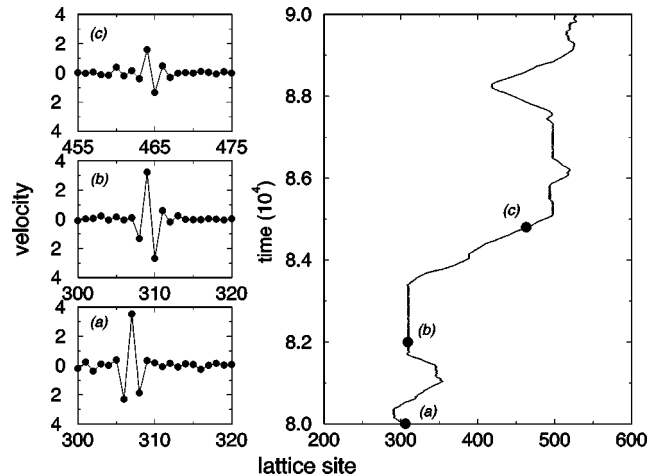


FIG. 5. The position of the maxima and three successive snapshots of the final breather as a function of time for $\epsilon = 0.0330$ and $N = 1024$. During this observation time the value of the breather energy is practically constant.

chain in the form of small amplitude (harmonic) oscillations (see again Fig. 4)

If we observe the breather on a time scale during which its amplitude does not change considerably, we see that it randomly alternates between a standing and moving state. Figure 5 clearly shows how each motion occurs with an almost constant velocity and is accompanied by changes of the shape. For a given amplitude, the modulus of such a velocity is always the same. We checked that its value is consistent with previous studies on exact breather solutions: for example, from Fig. 5 we found indeed the velocity to be 0.12(8) in excellent agreement with the value 0.120 obtained interpolating the data of Ref. 21.

The main issue concerning the velocity is the one of its selection. Although we will not enter into details here, we just anticipate that some analytic clue can be achieved by studying the moving solutions of Eq. (3)¹³. In particular, we expect the velocity to increase with the amplitude (almost linearly in the large-amplitude limit $9\phi_{\max}^2 \gg 1$), a fact that is actually consistent with ours as well as with other simulations.^{6,22}

Let us now discuss the power spectra of those strongly localized states. The right panels of Fig. 4 show how they differ depending on the amplitude. At smaller energies (upper panels) they resemble more closely an ES and the spectrum would have the exponentially decaying prescribed by Eq. (10). Upon increasing the energy, S_k acquires a sinusoidal component similar to Eq. (16) characteristic of a breather solution. This is simply the consequence of the fact that in the continuum approach the two are the only limiting cases of the general solution (7) with the amplitude ϕ_{\max} being determined only by the total energy of the lattice.

Obviously, the existence of the vibrational background affects the structure of the spectra. As the background contains a non-negligible energy (the remaining 70% in the case mentioned above), a complete description of the spectra requires some information on its features. One consequence that we can draw from Fig. 4 is that the naive picture of a single breather moving through a gas of small waves is oversimplified. In fact this would imply the spectrum to be a sum of a part similar to Eq. (14) plus an almost flat part. The form of S_k seems rather to indicate that only the high-frequency domain of the harmonic spectrum is populated.

To give a complete description of the relaxation process, we should now consider the final stage of the breather destruction (see also Ref. 6 for a discussion). This requires understanding how energy is transferred from large to small wave numbers. This issue involves in turn the interaction of a large-amplitude ES with small-amplitude plane waves. The latter problem has been studied in Ref. 23. Here we limit ourselves to observe that it is still not clear how to explain why the typical time scale for such a process should scale again as ϵ^{-2} . An analytic argument for that can be actually given only in the very last stage, when the system is sufficiently close to equilibrium and relaxation of Fourier modes can be described by usual linear-response theory.²⁴

VIII. FINITE-TEMPERATURE EFFECTS

What has been discussed so far refers to a situation where all the lattice energy is initially fed in the boundary mode.

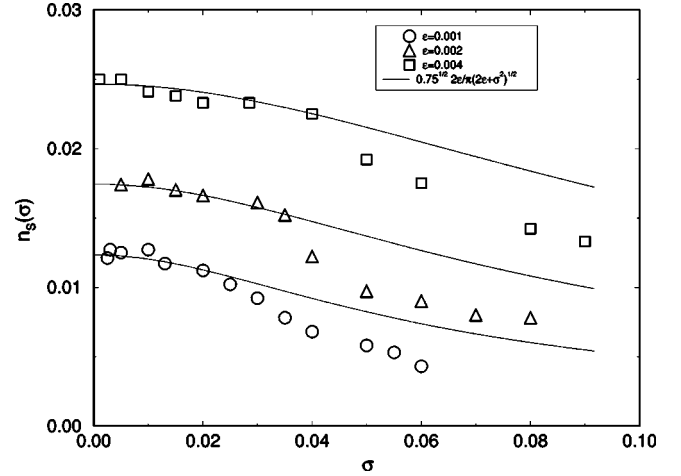


FIG. 6. The maximal initial density of ESs n_s as a function of the parameter σ for several values of N ranging between 920 and 1024. Solid lines correspond to formula (42).

This is clearly not realistic for actual experimental conditions, when the system is in contact with a thermal bath which populates all the other modes even at (classically) low temperatures. One may therefore wonder to what extent the mechanisms described above are effective in localizing the energy.

In the present section we try to approach the problem by first considering a more general class of initial conditions defined by replacing second expression in Eq. (25) by

$$\dot{u}_n(0) = (-1)^n V + \sigma' \xi_n, \quad (41)$$

where ξ_l is a Gaussian random number such that $\langle \xi_l \rangle = 0$ and $\langle \xi_l \xi_{l+m} \rangle = \delta_{m0}$. In order to assign the initial energy per particle to be equal to ϵ , we choose $V = 2\epsilon / \sqrt{2\epsilon + \sigma'^2}$ and $\sigma' = \sigma \sqrt{2\epsilon / (2\epsilon + \sigma'^2)}$. The parameter σ controls thus the ‘‘level of noise’’ in the initial conditions. Indeed, the limit $\sigma'^2 \ll \epsilon$ (exactly the one considered in the previous simulations) amounts to feeding all the available energy in the

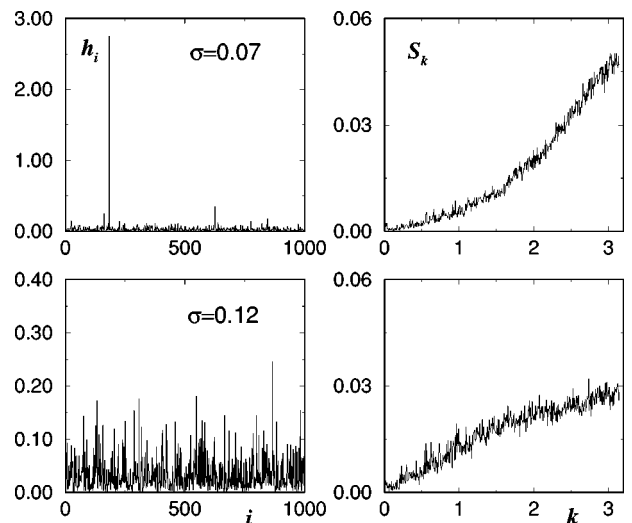


FIG. 7. The distribution of energy densities and power spectra of the velocities for $\epsilon = 0.0330$, $N = 1024$, $t = 8 \cdot 10^4$ and two different amplitudes of initial perturbation $\sigma = 0.07$ (upper panels) and 0.12 (lower panels).

Brillouin-zone-boundary mode whereas upon increasing σ , we can initially excite also the other Fourier modes of the chain. With our definition, equilibrium (i.e., equipartition of energy) is attained for $\sigma \rightarrow \infty$ (actually when $\sigma^2 \rightarrow \sigma_{\max}^2 = 2\epsilon N \gg \epsilon$).

The results presented before can be first extended to the finite σ values by the following reasoning. We showed that the initial density of ESs is basically determined by the vibrational energy available in the Brillouin-zone-boundary mode. Therefore, one can straightforwardly replace $\sqrt{2}\epsilon$ by V in Eq. (31) obtaining

$$n_s(\epsilon, \sigma) = \frac{1}{\pi} \sqrt{\frac{3}{4}} \frac{2\epsilon}{\sqrt{2\epsilon + \sigma^2}}. \quad (42)$$

As seen in Fig. 6, the numerical data are in good agreement with this formula, at least for moderate amplitudes $\sigma^2 \lesssim \epsilon/2$. This indicates that the mechanism of vibrational energy localization, induced by the modulational instability is relatively robust, at least with respect to those type of perturbation in the initial conditions.

Remarkably, as soon as σ overcomes the value $\epsilon/2$, n_s seems to be slightly less (around 10%) than what is expected from the purely energetic argument leading to Eq. (42). Although in this intermediate regime the numerical measurements are definitely less accurate due to the large vibrational background, we can at least give a qualitative justification to this. In fact, in this case we expect some effective dissipative

mechanisms to enter into play and some of the localized excitation will be effectively damped out by the interaction with the ‘‘thermal bath’’ of the other vibrational modes.²⁴

In order to better clarify this, we performed also some measurements of the power spectrum (27) for the relatively large values of σ (see Fig. 7). The results confirm that even for $\sigma^2 \approx \epsilon/2$ (lower panels) some kind of ‘‘localization’’ is still present, although it is hardly detectable by direct measurements.

To summarize, the creation of localized excitations, induced by the modulational instability, is more and more inhibited the closer one starts to the equipartition (thermal equilibrium) state. This is caused both by the less amount of energy available in the corresponding short-wavelength mode [as $V \rightarrow 0$ for $\sigma^2 \gg \epsilon$, see Eq. (41)] and by the effective dissipation caused by the interaction of small-amplitude ESs with the vibrational background. At least in principle, the envelope-function approach could be a useful starting point for a more detailed analysis once one would be able to take into account the fluctuating background by casting it in suitable stochastic terms.

ACKNOWLEDGMENTS

The authors are grateful to I. Barashenkov, S. Flach, A.S. Kovalev, A. Mayer, A. Politi and S. Ruffo for the useful discussions. This work was supported by the Max Planck Society.

*On leave from Moscow State Technological University ‘‘STANKIN,’’ 101472 Moscow, Russia. Electronic address: yukos@mpipks-dresden.mpg.de

†Author to whom correspondence should be addressed. Also at Istituto Nazionale di Fisica della Materia, Unit di Firenze. Electronic address: lepri@avanzi.de.unifi.it

¹A.M. Kosevich and A.S. Kovalev, Zh. ksp. Teor. Fiz. **67**, 1793 (1974) [Sov. Phys. JETP **40**, 891 (1975)].

²A.S. Dolgov, Fiz. Tverd. Tela (Leningrad) **28**, 1641 (1986) [Sov. Phys. Solid State **28**, 907 (1986)]; A.J. Sievers and S. Takeno, Phys. Rev. Lett. **61**, 970 (1988); J.B. Page, Phys. Rev. B **41**, 7835 (1990).

³A.J. Sievers and J.B. Page, in *Dynamical Properties of Solids*, edited by G.K. Horton and A.A. Maradudin (North-Holland, Amsterdam, 1995), Vol. 7, p. 137.

⁴S. Flach and C.R. Willis, Phys. Rep. **295**, 181 (1998).

⁵V.M. Burlakov and S.A. Kiselev, Zh. ksp. Teor. Fiz. **99**, 1526 (1991) [Sov. Phys. JETP **72**, 854 (1991)].

⁶T. Cretegny, T. Dauxois, S. Ruffo, and A. Torcini, Physica D **121**, 106 (1998).

⁷I. Daumont, T. Dauxois, and M. Peyrard, Nonlinearity **10**, 617 (1997).

⁸K.. Rasmussen, A.R. Bishop, and N. Gronbech-Jensen, Phys. Rev. E **58**, R40 (1998).

⁹K.. Rasmussen, S. Aubry, A.R. Bishop, and G.P. Tsironis,

patt-sol/9901002 (unpublished).

¹⁰A.S. Kingsep, L.I. Rudakov, and R.N. Sudan, Phys. Rev. Lett. **31**, 1482 (1973).

¹¹Yu. A. Kosevich, Phys. Lett. A **173**, 257 (1993).

¹²K. Yoshimura and S. Watanabe, J. Phys. Soc. Jpn. **60**, 82 (1991).

¹³Yu. A. Kosevich, Phys. Rev. B **47**, 3138 (1993); **48**, 3580 (1993).

¹⁴Yu. A. Kosevich, Phys. Rev. Lett. **71**, 2058 (1993).

¹⁵L.D. Landau and E.M. Lifshitz, *Mechanics* (Pergamon, Oxford, 1976).

¹⁶K.W. Sandusky and J.B. Page, Phys. Rev. B **50**, 866 (1994).

¹⁷T. Dauxois, S. Ruffo, and A. Torcini, Phys. Rev. E **56**, R6229 (1997).

¹⁸L. Casetti, Phys. Scr. **51**, 29 (1995).

¹⁹A.I. D’yachenko, V.E. Zakharov, A.N. Pushkarev, V.F. Shvets, and V.V. Yan’kov, Zh. ksp. Teor. Fiz. **96**, 2026 (1989) [Sov. Phys. JETP **69**, 1144 (1989)].

²⁰A.M. Kosevich and A.S. Kovalev, *Introduction to Nonlinear Physical Mechanics* (in Russian) (Naukova Dumka, Kiev, 1989).

²¹K.W. Sandusky, J.B. Page, and K.E. Schmidt, Phys. Rev. B **46**, 6161 (1992).

²²R. Bourbonnais and R. Maynard, Int. J. Mod. Phys. B **1**, 233 (1990).

²³T. Cretegny, S. Aubry, and S. Flach, Physica D **119**, 73 (1998).

²⁴S. Lepri, Phys. Rev. E **58**, 7165 (1998).