Effect of finite impurity mass on the Anderson orthogonality catastrophe in one dimension

H. Castella

Institut Romand de Recherche Nume´rique en Physique des Mate´riaux (IRRMA), PHB-Ecublens, CH-1015 Lausanne, Switzerland

*and Department of Physics, Ohio State University, 174 West 18th Avenue, Columbus, Ohio 43210-1106**

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A one-dimensional tight-binding Hamiltonian describes the evolution of a single impurity interacting locally with *N* electrons. The impurity spectral function has a power-law singularity $A(\omega) \propto |\omega - \omega_0|^{-1+\beta}$ with the same exponent β that characterizes the logarithmic decay of the quasiparticle weight *Z* with the number of electrons *N*, $Z \propto N^{-\beta}$. The exponent β is computed by (1) perturbation theory in the interaction strength and (2) numerical evaluations with exact results for small systems and variational results for larger systems. A nonanalytical behavior of β is observed in the limit of infinite impurity mass. For large interaction strength, the exponent depends strongly on the mass of the impurity in contrast to the perturbative result. $[$ S0163-1829(96)01348-3]

I. INTRODUCTION

Anderson studied the effect of a static impurity potential on conduction electrons in metals, $¹$ and showed that the</sup> ground state of the electrons is strongly renormalized by the local potential of the impurity and has an overlap with the unperturbed state, or quasiparticle weight *Z*, vanishing as *Z* \propto *N*^{- β} with an increasing number of electrons *N*. This effect, known as the Anderson orthogonality catastrophe, has its origin in an infrared singularity due to shake-up processes of the electron sea in the presence of the impurity potential, λ and signals the failure of the quasiparticle picture to describe the low-energy excitations.

The infrared singularity also affects the optical properties of metals. The core-level hole created by an x ray disturbs the conduction electrons similarly to an impurity potential. On one hand, the x-ray photoemission spectrum is asymmetrically broadened above the threshold.^{3,4} On the other hand, the x-ray-absorption spectrum has a strongly enhanced threshold, the so-called Fermi-edge singularity. $\frac{5}{5}$

These singularities in the optical spectra apply for a static core hole, i.e., an infinite-mass hole (or impurity). For a finite-mass hole and an isotropic band dispersion, the infrared singularity does not occur in three dimensions, because the hole recoil strongly restricts the number of low-energy excitations; $6,7$ as a consequence, the edge singularities disappear. In one dimension, however, the infrared singularity persists even for a finite-mass hole.⁸ The observation of an enhanced threshold in UV-absorption spectra of doped semiconductor quantum wires was interpreted as a Fermi-edge singularity, \int and stimulated renewed interest in the onedimensional problem.^{10,11}

The present work studies the Anderson orthogonality for a finite-mass impurity in a simple one-dimensional model, and focuses on the impurity recoil. Although the infrared singularity occurs for both infinite- and finite-mass impurities in one dimension, the recoil plays an important role on the critical exponent β which does not extrapolate to the staticimpurity value in the infinite-mass limit. The dependence of the exponent on the impurity mass is investigated analytically by perturbation theory, and numerically by a variational approach in the nonperturbative regime. The results are eventually compared to analytical calculations from Ref. 12, which studied the heavy-mass and strong-coupling regime using a path-integral formalism.¹³

Section II presents the model and the results. In Sec. III, a perturbation analysis of the impurity spectral function is performed along the line of Ref. 8. Section IV numerically calculates the critical exponent using a variational approach.

II. MODEL AND RESULTS

This section presents the model and summarizes the known results on the impurity spectral function and the Anderson orthogonality catastrophe. At the end of the section, the main results of this work are briefly described.

The model describes a single impurity and *N* spinless electrons moving on a chain of *L* sites with periodic boundary conditions. Within a tight-binding approximation with nearest-neighbor hopping, the band energies are $-2t_h \text{cos}k$ and $-2t \cos k$ for the impurity and the electrons, respectively. Further, the impurity and electrons feel an interaction *U* when they sit on the same site. Although this study is restricted to a single impurity, it is convenient to write the Hamiltonian \hat{H} in second-quantized form with creation operators c_i^{\dagger} for an electron on site *i*, and d_i^{\dagger} for an impurity:

$$
\hat{H} = -t \sum_{i=1}^{L} (c_i^{\dagger} c_{i+1} + \text{H.c.}) - t_h \sum_{i=1}^{L} (d_i^{\dagger} d_{i+1} + \text{H.c.})
$$

+ $U \sum_{i=1}^{L} d_i^{\dagger} d_i c_i^{\dagger} c_i$. (1)

The interaction is attractive in order to describe a hole in a valence band. For this particular model, however, the repulsive and attractive cases are related by a particle-hole transformation for the electrons $\tilde{c}_j = (-1)^j c_j^{\dagger}$. In the rest of the paper all the results are presented for the repulsive case; the corresponding results for the attractive interaction are obtained by the transformation $\rho \rightarrow 1-\rho$, where $\rho = N/L$ is the density of conduction electrons.

The spectral function of the hole (or impurity) $A(q,\omega)$ describes the photoemission response within the sudden approximation, i.e., neglecting interaction with the outgoing electron.¹⁴ It has a spectral decomposition in terms of eigenstates, $|\psi_n\rangle$, and eigenenergies, E_n , of the Hamiltonian (1) in the presence of the impurity, and of the ground-state wavethe presence of the impurity, and of the ground-state wave-
function $|\phi_0\rangle$ and energy \widetilde{E}_0 in the absence of the impurity:

$$
A(q,\omega) = \frac{1}{\pi} \operatorname{Im} G_q(\omega) = \sum_n |\langle \psi_n | d_q^\dagger | \phi_0 \rangle|^2 \delta(\omega - E_n + \widetilde{E}_0).
$$
\n(2)

In the spectral decomposition of $A(q=0,\omega)$, the spectral weight of the ground state $Z = |\langle \psi_0 | d_{q=0}^{\dagger} | \phi_0 \rangle|^2$ may remain finite in the thermodynamic limit, giving rise to a quasiparticle peak in the spectral function. This is the usual situation when the quasiparticle picture applies.

Static impurity $(t_h=0)$. The interaction causes the quasiparticle picture to break down: $¹$ the spectral weight scales to</sup> zero with increasing number of fermions *N* as $Z \propto N^{-\beta}$, the ground state of the interacting system being orthogonal to the quasiparticle state $d^{\dagger}|\phi_0\rangle$ for $N \rightarrow \infty$. This is known as the Anderson orthogonality catastrophe. The exponent is related to the phase shift δ_F of an electron at the Fermi energy scattered off the static impurity:

$$
\beta(t_h=0) = (\delta_F/\pi)^2. \tag{3}
$$

For the present model, the phase shift depends on *U* and on the density of states at the Fermi energy $N_F = 1/(2\pi \sin k_F)$: $\delta_F = -\arctan(\pi U N_F)$. Furthermore, the spectral function has, instead of a quasiparticle peak, a power-law singularity at instead of a quasiparticle peak, a power-law singularity at
the threshold $\omega_0 = E_0 - \tilde{E_0}$, with the critical exponent $1-\beta(t_h=0)$:⁴

$$
A(\omega) \propto \frac{1}{|\omega - \omega_0|^{1-\beta}}.\tag{4}
$$

This singularity in $A(\omega)$ is observed in x-ray photoemission of metals.³

Finite-mass impurity $(t_h > 0)$. While an infinite-mass impurity acts as an external potential on the electrons, the impurity recoil further complicates the many-body problem. Despite this complexity, the eigenenergies and eigenstates of \hat{H} are known exactly for the special case $t_h = t$ ¹⁵. Using the exact solution, the spectral function for $q \approx 0$ is computed in Ref. 16. It has no quasiparticle peak because of the Anderson orthogonality catastrophe, and has a power-law singularity with an exponent

$$
\beta(t_h = t) = 2(\delta'_F/\pi)^2.
$$
 (5)

The exponent is given by the phase shift of a single electron at k_F scattering off a finite-mass impurity δ'_F $=$ - arctan($\pi U N_F/2$). Notice the similarity in the exponents for $t_h = t$ and $t_h = 0$, which are both expressed in terms of phase shifts. The phase shifts, however, differ since the number of states contributing to the Anderson orthogonality is reduced from UN_F to $UN_F/2$ between a static and finite mass impurity, respectively. The origin of this difference is discussed in Sec. III for the perturbative results.

FIG. 1. Exponent $\beta(t_h)/\beta(t_h=t)$ normalized to the analytical value $\beta(t_h=t)$ in Eq. (5) as a function of the mass ratio t_h/t for different interaction strengths U and densities ρ . The filled symbols at $t_h=0$ are the exact results in Eq. (3) for the static impurity. The open symbols for t_h >0 are the exponents extracted as $N \rightarrow \infty$ from the variational results for a sequence of finite systems with $N_1/N_2=0.7$, the largest sizes being $N_1=159$ and 121 at halfand one-third-filling, respectively.

The present work presents calculations of the exponent $\beta(t_h)$ for different hopping parameters t_h , interaction strengths *U*, and electron densities ρ . The exponent $\beta(t_h)$ is computed (1) analytically using perturbation theory in U , and (2) numerically in the nonperturbative regime using a variational wave function proposed by Edwards.¹⁷ The exponent is extracted numerically from finite-size results using a precise scaling law for *Z* as a function of *N* and a numerical fit of the data as $N \rightarrow \infty$. The main results of this study are summarized now. Sections III and IV will give a detailed description of the perturbation calculations and the numerical simulations, respectively.

Mass dependence of the exponent. The perturbative results indicate that the exponent β , to order $(U/t)^2$, is independent of t_h for $t_h > 0$, and equals $(UN_F)^2/2$. For a finite *U*, however, the exponent $\beta(t_h)$ calculated numerically does depend on t_h , and its dependence increases with increasing *U*, as illustrated in Fig. 1, which shows the exponent normalized to its value for $t_h = t$, $\beta(t_h)/\beta(t_h = t)$, as a function of t_h . In the strong-coupling limit, $U = \infty$, β varies quasilinearly with t_h .

Discontinuous exponent in the heavy-mass limit. The perturbative calculations predict a discontinuous exponent in the limit of a flat impurity dispersion: $\lim_{t_h\to 0} \beta(t_h)$ $= \beta(t_h=0)/2$. This discontinuity is due to the irrelevance of backscattering processes for t_h > 0 because of the finite recoil energy involved. The numerical results in Fig. 1 illustrate the nonanalyticity for a finite *U*: when $t_h \rightarrow 0$ the exponent does not extrapolate to the static value which is indicated by the filled symbols at $t_h=0$. The numerical results can be compared to calculations by Rosch and $Kopp¹²$ for the exponent in the heavy-mass and strong-coupling regime. Their analysis, based on an effective action for the long-time behavior of the impurity propagator, also predicts a discontinuity $\lim_{t_h \to 0} \beta(t_h)/\beta(t) = \alpha \neq \beta(0)/\beta(t)$, and gives $\alpha = \frac{1}{4}$, at half-

FIG. 2. Logarithmic decay of the quasiparticle weight *Z* with number of electrons N from the perturbative result in Eq. (6) at half-filling, and for different impurity hopping parameters t_h . The dashed lines are the asymptotic behaviors with the analytical exponent $\beta(t_h)$ in Eq. (8). The arrows indicate the estimate of the crossover size N_c from Eq. (A10).

filling and $\alpha = \frac{19}{56} \approx 0.339$ at one-third-filling. The numerical results in Fig. 1 suggest $\alpha \approx 0.25$ for both half- and one-thirdfillings. The result at half-filling is therefore in good agreement with their prediction, while at third filling the value of α is significantly smaller. Note, however, that the present work relies on a variational approach.

Crossover behavior in the heavy-mass limit. A detailed analysis of the perturbative results for $t_h \leq t$ reveals a crossover in the scaling behavior of ln*Z* as a function of ln*N*, as shown in Fig. 2. While ln*Z* closely follows the staticimpurity behavior with a slope $\beta(t_h=0)$ for a number of fermions smaller than a crossover value N_c , it adjusts to the finite-mass behavior only for $N>N_c$. Further, N_c diverges as t/t_h for $t_h \rightarrow 0$. Therefore, the discontinuity is only an asymptotic result, valid in the limit $N \rightarrow \infty$.

III. PERTURBATION THEORY

This section evaluates the impurity spectral function and spectral weight *Z* perturbatively in *U* and for an arbitrary hopping parameter t_h in one dimension, following Ref. 8, that computes $A(q=0,\omega)$ for the equal-masses case $(t_h=t)$ in connection with the stability of the ferromagnetic state in the Hubbard model.

Spectral weight. In the perturbative expansion, the first terms that renormalize the ground-state wave function correspond to the creation of a single particle-hole pair within the Fermi sea by an impurity of momentum $q=0$. The excitation energy is $\Delta \epsilon(k_1, k_2) = 2t \cos k_1 - 2t \cos k_2 + 2t_h - 2t_h \cos(k_1$ $-k_2$). The spectral weight has a cumulant expansion⁶ that, up to second order in *U*, involves only these excitations:

$$
\ln Z = -\left(\frac{U}{L}\right)^2 \sum_{k_1, k_2} \frac{\Theta(k_F - |k_1|) \Theta(|k_2| - k_F)}{\Delta \epsilon(k_1, k_2)^2}.
$$
 (6)

The sum over k_1, k_2 diverges logarithmically with an increasing number of electrons, *N*. As shown in the Appendix, a large-*N* expansion gives

$$
ln Z = -\beta(t_h) ln N + \alpha_0(t_h) + \alpha_1(t_h)/N + O(1/N^2).
$$
 (7)

The logarithmic term dominates for large *N*, and gives rise to the Anderson orthogonality. The finite-size corrections are used in Sec. IV for the numerical study of β .

The main result of this section is the evaluation of $\beta(t_h)$:

$$
\beta(0) = (UN_F)^2 \quad \text{for } t_h = 0,
$$

\n
$$
\beta(t_h) = \frac{1}{2} (UN_F)^2 \quad \text{for } t_h > 0.
$$
\n(8)

The exponent $\beta(t_h)$ is independent of t_h as far as $t_h > 0$. The infrared singularity is caused by forward-scattering processes with small momentum transfer $|k_2 - k_1| \ll k_F$, which are gapless excitations for any t_h . The hopping t_h is irrelevant since the impurity recoil energy $2t_h(cos(k₁-k₂)-1)$ is negligible as compared to the particle-hole energy $2t(\cos k_2 - \cos k_1)$.

Further, the exponent has a discontinuity in the heavymass limit $\lim_{t_h\to 0} \beta(t_h) = \beta(t_h=0)/2$. The difference between the infinite- and finite-mass exponents is simply related to scattering of one electron from one side of the Fermi surface to the other. These so-called backscattering processes, which involve a large momentum transfer $|k_2 - k_1| \approx 2k_F$, do not contribute to the infrared divergence for $t_h > 0$, since the impurity recoil opens a gap $2t_h(\cos 2k_F-1)$. For $t_h=0$ however, both backscattering and forward-scattering processes are gapless. The number of low-energy excitations contributing to the infrared singularity is thus reduced by a factor of 2 for a finite-mass impurity, as compared to its value for the static impurity.

The discontinuity in the exponent is an asymptotic result valid only for $N \rightarrow \infty$. For a finite number of fermions and a large but finite mass $0 \lt t_h \ll t$, however, ln*Z* has a crossover as a function of ln*N*, illustrated in Fig. 2, where the spectral weight is computed numerically from Eq. (6) . The logarithm of the spectral weight has the slope $\beta(t_h)$ only for a number of electrons larger than a crossover value N_c , while for a small *N* it follows the static-impurity behavior with a slope $\beta(t_h=0)$. The dashed lines indicate the asymptotic behaviors $-\ln Z = \beta(t_h)\ln N - \alpha_0(t_h)$, with the analytical β from Eq. (8) , and α_0 fitted to the value of lnZ for the largest size. The intercept of the asymptotes gives the crossover size N_c , which agrees very well with the estimate $N_c \approx 0.3244t/t_h$, presented in the Appendix. Notice that N_c diverges as t/t_h for $t_h \rightarrow 0$, and the asymptotic regime is reached for a larger number of electrons the lower t_h .

Spectral function. The spectral function is computed only at $q=0$, where it has a power-law singularity. The propagator is

$$
G_{q=0}(\tau) = -i \langle \phi_0 | d_{q=0} \exp(-i\hat{H}\tau) d_{q=0}^{\dagger} | \phi_0 \rangle e^{i\tilde{E}_0 \tau} \Theta(\tau). \tag{9}
$$

The propagator also has a cumulant expansion, and is written in terms of the density of particle-hole excitations $S(\omega)$ and a renormalized impurity energy $\tilde{\epsilon}_0$:

$$
G_{q=0}(\tau) = -i \exp\left[-i\tilde{\epsilon}_0 \tau - U^2 \int_0^\infty S(\omega) \times \frac{1 - \exp(-i\omega \tau)}{\omega^2} d\omega\right],
$$
 (10)

$$
S(\omega) = \frac{1}{(2\pi)^2} \int_{-\pi}^{\pi} \int_{-\pi}^{\pi} \Theta(|k_1| - k_F) \Theta(k_F - |k_2|)
$$

$$
\times \delta(\omega - \Delta \epsilon(k_1, k_2)) dk_1 dk_2.
$$

The density *S* has a linear frequency dependence for small $\omega: S(\omega) = (\beta(t_h)/U^2)\omega$, where $\beta(t_h)$ is the exponent of the Anderson orthogonality in Eq. (8) . This linear behavior determines the low-frequency spectral function which has a power-law singularity at the threshold with the exponent $1-\beta$:⁴

$$
A(q=0,\omega) = \frac{\sin(\pi\beta)\Gamma(1-\beta)}{\pi(\omega-\tilde{\epsilon}_0)^{1-\beta}}\Theta(\omega-\tilde{\epsilon}_0).
$$
 (11)

This is, however, only an asymptotic result for frequencies smaller than a cutoff *W*. For $t_h=0$, the cutoff is of the order of the Fermi energy: $W \approx 2t(1-\cos k_F)$. For $t_h > 0$ however, the linear behavior of *S* holds only for frequencies smaller than the impurity recoil energy, and the cutoff is given by $W \approx min(2t(1-\cos k_F), 2t_h(1-\cos 2k_F))$. For a heavy impurity the cutoff is of the order of the impurity recoil energy $2t_h(1-\cos 2k_F)$ rather than the Fermi energy, and the asymptotic result is valid only in a very narrow frequency range. Furthermore, the density of excitations exhibits a crossover similarly to the spectral weight. This might give rise to a crossover in the spectral function as well.

In summary the exponent β in Eq. (8) characterizes the power-law singularity of the spectral function and the logarithmic decay of the quasiparticle weight. The exponent does not depend on the mass of the impurity except in the static limit $t_h=0$, and it has a discontinuity at $t_h=0$. Notice that the perturbative results agree with the small-*U* expansions of the exponent for $t_h=0$ and $t_h=t$.

IV. NUMERICAL STUDY

This section presents a numerical study of the exponent $\beta(t_h)$ based on a variational approach. The variational predictions for the energy and correlation functions are compared to results from Lanczos exact diagonalizations and projection quantum Monte Carlo simulations. Then the variational calculations are used to extract the exponent.

The variational wave function was originally proposed for the single spin-flip problem in the two-dimensional Hubbard model in reference to the stability of the ferromagnetic state.¹⁷ It was also used to study numerically the quasiparticle weight in two dimensions.¹⁸ Furthermore, this variational approach is equivalent to the approximation used in Ref. 11. In one dimension, the variational class of wave functions contains all the eigenstates of the model (1) for $t_h = t$, as shown by Edwards.¹⁷ The variational approach is thus expected to include much of the relevant correlations even for $t_h \neq t$.

In the reference frame comoving with the impurity, a

wave function $|\Psi_{q}\rangle$ of total momentum *q* is represented by a function $f(j_1, \ldots, j_N)$ depending only on the positions of the electrons:

$$
|\Psi_{q}\rangle = \frac{1}{\sqrt{L}} \sum_{j_{0}=1}^{L} e^{iqj_{0}} d_{j_{0}}^{\dagger}
$$

$$
\times \sum_{j_{1}, \dots, j_{N}=1}^{L} f(j_{1} - j_{0}, \dots, j_{N} - j_{0}) c_{j_{1}}^{\dagger} \dots c_{j_{N}}^{\dagger} |0\rangle.
$$

(12)

The variational ansatz for f is a determinant of singleparticle wave functions ϕ_m :

$$
f(j_1, ..., j_N) = \frac{1}{\sqrt{N!}} \det[\phi_m(j_l)]_{m,l=1,...,N}.
$$
 (13)

The expectation value of the energy is

$$
\langle \hat{H} \rangle = \sum_{l=1}^{N} \left[-t \sum_{j=0}^{L-1} (\phi_l^*(j)\phi_l(j+1) + \text{c.c.}) + U|\phi_l(0)|^2 \right]
$$

- $t_h[\exp(-iq)\det(\mathbf{S}) + \text{c.c.}],$

$$
\mathbf{S}_{mn} = \sum_{j=0}^{L-1} \phi_m^*(j+1)\phi_n(j).
$$
 (14)

The variational parameters $\phi_m(j)$ are found by minimization of the energy using a steepest-descent algorithm. If one chooses to start with $\phi_l(i)$ as the exact solution for $t_h = t$, convergence is reached after a relatively small number of iterations even for hopping parameters very different from *t*.

Comparison of variational results to Lanczos exact diagonalizations and projection quantum Monte Carlo simulations. Only small systems are accessible by exact diagonalization because the dimension of the Hilbert space increases very rapidly with the number of lattice sites. For the halffilled band, the biggest closed-shell system studied has $L=22$, whereas for $\rho = \frac{1}{4}$ it has $L=20$. For larger systems, Monte Carlo simulations are required.

The projection quantum Monte Carlo simulation gives a statistical estimate of ground-state expectation values of observables $\ddot{\theta}$ by a projection of a trial wave function $|\psi_T\rangle$ onto the ground state with the operator $\exp(-\eta \hat{H})$:¹⁹

$$
\langle \psi_0 | \hat{O} | \psi_0 \rangle = \lim_{\eta \to \infty} \frac{\langle \psi_T | \exp(-\eta \hat{H}) \hat{O} \exp(-\eta \hat{H}) | \psi_T \rangle}{\| \exp(-\eta H) | \psi_T \rangle \|^2}.
$$
\n(15)

The algorithm used here closely follows Ref. 20. The imaginary-time evolution $exp(-nH)$ is performed sequentially for small time intervals $\Delta \tau$, and a Trotter decomposition is used for the kinetic-energy and interaction terms in the Hamiltonian. Eventually the two-body term is represented by discrete Hubbard-Stratonovitch fields that mediate the interaction. For the present calculations a rather large η =15 is necessary in order to converge the relevant correlation functions, and the results are extrapolated to $\Delta \tau \rightarrow 0$ using several values of the time interval.

TABLE I. Comparison of correlation energy, $e_c = E_0 - \widetilde{E}_0 - U\rho$, where E_0 and \widetilde{E}_0 are the ground-state energies of the interacting and noninteracting systems, respectively, computed by Lanczos exact diagonalizations, projection quantum Monte Carlo (PQMC), and Edwards's variational wave function, as a function of the number of fermions *N* at half- and quarter-filling ($\rho = N/L = \frac{1}{2}$ and $\frac{1}{4}$, respectively), for $t_h = 0.5t$ and $U=4t$. The last line gives the energy extrapolated for the infinite system both from either Lanczos or PQMC results, and the variational results by a linear fit in 1/*N*.

e_c				e_c			
N (ρ =1/2)	Lanczos	POMC	Variational	N (ρ =1/4)	Lanczos	POMC	Variational
5	-0.9028		-0.9025	5	-0.6102		-0.6101
7	-0.9230		-0.9227	7		$-0.619(4)$	-0.6190
9	-0.9340		-0.9336	9		$-0.624(3)$	-0.6239
11	-0.9408		-0.9404	11		$-0.627(2)$	-0.6271
13		$-0.945(2)$	-0.9451	15		$-0.632(2)$	-0.6309
∞	-0.9727		-0.9720	∞		$-0.642(4)$	-0.6413

The energy is computed for the ground state, which has zero total momentum. The relevant quantity is the correlation energy $e_c = E_0 - \widetilde{E}_0 - U\rho$, where E_0 and \widetilde{E}_0 are the interacting and noninteracting energies, respectively. Notice that e_c is of order 1, while E_0 is of order *N*. Table I gives e_c for several system sizes at half- and quarter-filling, and for an interaction equal to the bandwidth $U=4t$ and a hopping parameter $t_h=0.5t$. The variational ansatz is not exact for t_h $\neq t$, since the variational energy departs significantly from the exact energy. Nevertheless it always remains very close to the ground-state energy. The relative difference between the exact and the variational energy remains smaller than 0.1% even when extrapolating to the infinite-system limit with a $1/N$ scaling law. For the sake of comparison the correlation energy for the unrestricted Hartree-Fock solution is $-0.4380t$ at quarter-filling, a value much larger than the variational energy $-0.6413t$.

The $k=0$ component of the momentum distribution function of the impurity $n(k=0)$ has the same scaling law as Z when the Anderson orthogonality catastrophe occurs. Both *Z* and $n(k=0)$ are computed in order to test the relevance of the variational calculations to extract the exponent β . Figure 3 shows the ratio of the exact momentum distribution $n(k=0)$, calculated by either exact diagonalization or quantum Monte Carlo methods depending on the system size, to the variational estimate $n(k=0)_{var}$ as a function of $1/N$ at quarter-filling and half-filling, for $U=4t$ and $t_h=0.5t$. If the variational result reproduced correctly the scaling behavior of $n(k=0)$ with *N*, the ratio $n(k=0)/n(k=0)_{var}$ should be constant as a function of 1/*N*. At half-filling, the ratio remains indeed always close to 1, and the variational approach seems to correctly describe the scaling behavior. At quarterfilling, however, the comparison relies mostly on Monte Carlo simulations, where the estimate of $n(k=0)$ has strong statistical fluctuations for large system sizes, large autocorrelation times having an essential contribution to the error bars. The comparison to the variational results is therefore delicate, but still the ratio remains close to 1 within the error bars. Finally, exact-diagonalization results for *Z* are presented in Fig. 4 for $U = \infty$, $\rho = \frac{1}{2}$ and different t_h . The slope of ln*Z* as a function of ln*N* is an estimate of the exponent β , and both exact-diagonalization results and variational calculations are in good agreement. These last results illustrate the robustness of the variational wave function even in the strong-coupling regime.

Finite-size scaling analysis of the exponent. The finitesize corrections to the Anderson orthogonality catastrophe in Eq. (7) , which were derived from the perturbative analysis, are used now to extract numerically the exponent. It should be stressed that these 1/*N* corrections contrast with the slowly decaying ln ln*N* corrections expected for the paramagnetic phase of the one-dimensional Hubbard model. 21

The exponent is extracted from numerical data on finite systems by scaling the slope of ln*Z* as a function of ln*N*: the spectral weights Z_1 and Z_2 are computed for systems with numbers of fermions N_1 and N_2 , respectively, at a fixed density and a fixed ratio $r = N_1 / N_2$; from the perturbative analysis in Eq. (7), the slopes $(lnZ_1 - lnZ_2)/(lnN_2 - lnN_1)$ have a polynomial expansion in $1/N$, *N* being the mean number of electrons $N = (N_1 + N_2)/2$:

FIG. 3. Ratio of the momentum distribution function $n(k)$ at $k=0$, calculated with the Lanczos diagonalization or projection quantum Monte Carlo, to its variational estimate $n(k=0)_{var}$ as a function of $1/N$ for half- and quarter-filling, and for $U=4t$ and $t_h=0.5t$. The circles are the Lanczos results and the squares the Monte Carlo results with the corresponding error bars. The left axis refers to half-filling, and the right axis to quarter-filling.

FIG. 4. Comparison of the exact spectral weight *Z* from Lanczos diagonalization to its variational estimate as a function of number of electrons *N* in the log-log plot for $U = \infty$, $\rho = \frac{1}{2}$ and different t_h . The open symbols are the variational results, and the filled symbols the Lanczos results.

$$
\frac{\ln Z_1 - \ln Z_2}{\ln N_1 - \ln N_2} = -\beta + \alpha_1 \frac{1 - r^2}{2r \ln r} \frac{1}{\overline{N}} + O\left(\frac{1}{\overline{N}^2}\right).
$$
 (16)

The exponent β is estimated by a numerical fit of the slopes as $N \rightarrow \infty$. In practice, it is not possible to keep the ratio N_1/N_2 exactly fixed while increasing *N*, but the number of electrons can be adjusted such that N_1/N_2 approaches a fixed value for large *N*. All the results presented here have been obtained for $N_1/N_2 \approx 0.7$.

The scaling procedure is tested both in the perturbative regime for different t_h and at finite *U* for $t_h = t$, where β is known analytically. Figure 5 illustrates the scaling procedure for the perturbative regime where ln*Z* is computed from Eq.

FIG. 5. Slopes $d \ln(Z)/d \ln(N) = (\ln Z_1 - \ln Z_2)/(\ln N_1 - \ln N_2)$ normalized to the analytical value of $\beta(t_h)$ in Eq. (8) as a function of the inverse mean number of electrons $2/(N_1 + N_2)$ in the perturbative regime and for different hopping parameters t_h . The spectral weight Z is computed numerically from Eq. (6) . The dashed lines are numerical fits of the data as polynomials in $2/(N_1 + N_2)$. The inset is an enlargement around the origin.

FIG. 6. Slopes *d* ln(*Z*)/*d* ln(*N*) normalized to the analytical value of $\beta(t_h=t)$ in Eq. (5) as a function of the inverse mean number of electrons $2/(N_1+N_2)$ for $t_h=t$ and different interaction strengths $U=2$, 4, and 8*t* at half-filling, and for $N_1/N_2=0.7$. The spectral weight *Z* is computed with the Bethe's ansatz wave function in Eq. (13). The dashed lines are numerical fits of the data as a third-order polynomial in $2/(N_1+N_2)$. The inset is an enlargement around the origin.

 (6) . The inset shows that the scaling procedure is well be-(6). The inset shows that the scaling procedure is well behaved even for very large \overline{N} . Furthermore, the scaling analysis is essential to determine the dependence of β on t_h , since the finite-size corrections differ in both sign and magnitude for different t_h . Note also that the nonlinear terms in $1/N$ become increasingly important with decreasing t_h . The spectral weight *Z* is computed numerically for $t_h = t$ and different *U* using the Bethe's ansatz wave function (13) , and the results are fitted with a third-order polynomial in $1/N$. The relative accuracy of the fitting procedure in extracting β remains of the order of 10^{-4} even for a strong interaction $U=8t$, as illustrated in Fig. 6. Although the scaling behavior was derived in the perturbative regime, it seems to hold for any interaction strength *U*.

For a finite *U* and $t_h \neq t$, the exponent is not known, and its evaluation relies on the variational approach and the finite-size scaling analysis. Figure 7 illustrates the scaling procedure in the strong-coupling regime $U = \infty$, for $\rho = \frac{1}{2}$ and different t_h . For $t_h > 0.1t$, the exponent, which is obtained by an extrapolation of the data as $N \rightarrow \infty$, depends only slightly on the order of the polynomial used in the fit. For $t_h=0.1t$, however, the fitting procedure is not well behaved, and a relative error of a few percent is expected in the extraction of β . A precise investigation of the heavy-mass regime $t_h \leq t$ would require the study of even larger systems due to the important nonlinear corrections in 1/*N*.

Mass dependence of the exponent. The study of $\beta(t_h)$ is based on the variational approach and the scaling analysis presented above. Figure 1 presents the exponent normalized to its value at $t_h = t$, $\beta(t_h)/\beta(t_h = t)$, as a function of t_h/t . For a finite *U*, the exponent does depend on the hopping parameter t_h , in contrast to the perturbative result $\beta(t_h)/\beta(t_h=t)$ indicated by the dotted line. Furthermore, this dependence increases with increasing *U*, and in the strong-coupling regime $U = \infty$ the exponent varies quasilin-

FIG. 7. Slopes $d \ln(Z)/d \ln(N)$ normalized to the fitted $\beta(t_h)$ as a function of the inverse mean number of electrons $2/(N_1 + N_2)$ for $U = \infty$, different hopping parameters $t_h = 0.5$, 0.25, and 0.1*t*, at halffilling. The spectral weight is computed with the variational wave function. The solid and dashed lines are numerical fits of the data as third-order and fourth-order polynomials in $2/(N_1 + N_2)$, respectively.

early with t_h/t . For small t_h and at half-filling, the exponent only slightly departs from the linear behavior. Furthermore, the exponent depends on the density, since the data for $U = \infty$ and $\rho = \frac{1}{2}$ significantly differ from the exponents for $\rho = \frac{1}{3}$. This contrasts with the exact exponent for $t_h = 0$ and t_h = *t*, which are independent of the density. However, the limiting value $\lim_{t_h\to 0} \beta(t_h)$ seems to be independent of the density.

The numerical results demonstrate the discontinuity of the exponent at $t_h=0$ for a finite *U*. The numerical data, indeed, do not extrapolate to the exact results for $t_h=0$ indicated by the filled symbols in Fig. 1. A precise extraction of the limiting value $\lim_{t_h\to 0} \beta(t_h)$, however, would require an investigation of larger systems.

The occurrence of the nonanalyticity at $t_h=0$ is not surprising, since the translational symmetry is broken at this point. Still the role played by the recoil of the impurity in the discontinuity of the exponent is not clear. The perturbation calculations indicate that the discontinuity at $t_h=0$ is due to the irrelevance of the backscattering processes whenever t_h > 0 ; yet a simple argument can persuade us that this is true only in the small-*U* regime. Indeed let us first assume backscattering to be responsible for the nonanalyticity for all *U*. One can devise an effective model for the heavy-impurity limit $(t_h \ll t)$, where the impurity is considered static but the interaction with the electrons is restricted to forward scattering:

$$
H = -2t\sum_{k} \cos k c_{k}^{\dagger} c_{k} + \frac{U}{L} \sum_{k k' > 0} c_{k}^{\dagger} c_{k'}.
$$
 (17)

In this picture, the only effect of the impurity recoil is the restriction to forward-scattering processes. Since the potential is static, the exponent is expressed in terms of phase shifts. A calculation of the phase shift gives the same exponent as the result for $t_h = t$ in Eq. (5). This is not compatible with the numerical results for $t_h \leq t$. Therefore forwardscattering processes alone cannot account for the discontinuity, and backscattering has to be invoked.

V. CONCLUSIONS

This study of the one-dimensional Anderson orthogonality catastrophe, combining analytical and numerical calculations, has focused on the effect of the impurity recoil. The numerical study requires a finite-size scaling analysis, since the Anderson orthogonality catastrophe results from a logarithmic decay of the quasiparticle weight with the number of fermions. The present work shows, however, that a reliable numerical analysis of the quasiparticle renormalization can be achieved if a precise scaling hypothesis is established, and large enough systems are accessed.

Within perturbation theory, the infrared singularity that signals the orthogonality catastrophe occurs for an impurity band of any dispersion. Still, there is a discontinuity between the zero-band-width exponent and the finite-band-width exponent. This discontinuity is related to the impurity recoil, which opens a gap in the spectrum of particle-hole excitations for backscattering of one electron from the Fermi momentum k_F to $-k_F$.

Outside the perturbative regime, the numerical analysis demonstrates the discontinuous behavior of the exponent, and agrees at half-filling with a study of the heavy-mass and strong-coupling regime in Ref. 12. At one-third-filling, however, the numerical results differ from the analytical prediction.

The discontinuity of the exponent is an asymptotic result valid only in the limit of an infinite system. For a finite number of electrons and a heavy but finite-mass impurity, the quasiparticle weight has the same logarithmic behavior as a static impurity up to a critical number of electrons N_c where the former weight shows a crossover to the true asymptotic decay for a finite-mass impurity. Furthermore, N_c diverges with increasing mass of the impurity. This crossover is also expected in the low-frequency behavior of the spectral function.

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APPENDIX

This appendix derives the scaling law (7) from the perturbative expression for the spectral weight *Z*, and estimates the crossover size N_c . The calculations are presented for the half-filled band in detail, and the density dependence of N_c is briefly discussed at the end.

Forward-*scattering processes.* An electron with momentum k_1 is scattered into an empty state with momentum k_2 ,

$$
\Delta \epsilon(k_1, k_2) \approx 4 \pi t \frac{n_1 + n_2}{2N} \left(1 + \frac{t_h \pi(n_1 + n_2)}{2Nt} \right). \tag{A1}
$$

While the linearization of the energies allows an exact calculation of the exponent $\beta(t_h)$, it provides only an estimation of the remaining terms in Eq. (7) . The contribution of forward-scattering processes to the spectral weight is

$$
-\left(\frac{U}{L}\right)_{k_1k_2>0} \frac{1}{\Delta \epsilon(k_1, k_2)^2}
$$

= $-\frac{1}{2} (UN_F)^2 \sum_{n_1=0}^{N/2} \sum_{n_2=1}^{N/2}$
 $\times \frac{1}{(n_1+n_2)^2 (1+\pi t_h(n_1+n_2)/(2Nt))^2}.$ (A2)

The fraction in the sum is expanded in four different terms:

$$
\frac{1}{(n_1 + n_2)^2 (1 + \pi t_h (n_1 + n_2)/(2Nt))^2}
$$
\n
$$
= \frac{1}{(n_1 + n_2)^2} + \frac{1}{(n_1 + n_2 + 2Nt/(\pi t_h))^2} - \frac{\pi t_h}{Nt(n_1 + n_2)}
$$
\n
$$
+ \frac{\pi t_h}{Nt(n_1 + n_2 + 2Nt/(\pi t_h))}.
$$
\n(A3)

The first term gives the logarithmic divergence²²

$$
\sum_{n_1=0}^{N/2} \sum_{n_2=1}^{N/2} \frac{1}{(n_1+n_2)^2} = \ln N + 1 + C - 2 \ln 2 + \frac{1}{N} + O\left(\frac{1}{N^2}\right),\tag{A4}
$$

where $C \approx 0.5772$ is Euler's constant. The contributions of all other terms remain finite for $N \rightarrow \infty$. For the second and fourth terms in Eq. $(A3)$, the discrete sum can be replaced by an integral. Further, the finite-size corrections are of order

*Present address.

- ¹P. W. Anderson, Phys. Rev. Lett. **18**, 1049 (1967); Phys. Rev. 164, 352 (1967).
- 2 J. J. Hopfield, Comments Solid State Phys. 2, 40 (1969) .
- 3G. K. Wertheim and P. H. Citrin, in *Photoemission in Solids II*, edited by L. Ley and M. Cardona (Springer, Berlin, 1979); for a review, see K. Ohtaka and Y. Tanabe, Rev. Mod. Phys. **62**, 929 $(1990).$
- ⁴S. Doniach and M. Sunjić, J. Phys. C 3, 285 (1970).
- ⁵G. D. Mahan, Phys. Rev. **163**, 612 (1967); P. Nozières, J. Gavoret, and B. Roulet, *ibid*. **178**, 1084 (1969); P. Nozières and C. T. DeDominicis, *ibid.* **178**, 1097 (1969).
- ${}^{6}E$. Müller-Hartmann, T. V. Ramakrishnan, and G. Toulouse, Phys. Rev. B 3, 1102 (1971).

1/*N*. The extraction of the finite-size corrections for the third term, however, requires the evaluation of the discrete sum²²

$$
\frac{1}{N} \sum_{n_1=0}^{N/2} \sum_{n_2=1}^{N/2} \frac{1}{n_1 + n_2} = \ln 2 + \left(\ln 2 - \frac{1}{2} \right) \frac{1}{N} + O\left(\frac{1}{N^2}\right). \tag{A5}
$$

Backscattering processes. The initial and final momenta have a different sign, $k_1k_2<0$, and, for $k_1>0$ and $k_2<0$, the momenta are written as $k_1 = k_F - 2\pi n_1 / L$ and $k_2 = -k_F - 2\pi n_2 / L$. The linearized excitation energy is

$$
\Delta \epsilon(k_1, k_2,) = 4 \left(t_h + \frac{t \pi(n_1 + n_2)}{2N} \right). \tag{A6}
$$

For $t_h=0$, backscattering processes are gapless excitations, and make the same contribution to *Z* as forward-scattering processes, as given in Eq. $(A4)$. For $t_h > 0$, however, they have a gap, and their contribution, which is finite, is computed by an integral representation of the sum.

Finally, all the terms, for both forward-scattering and backscattering, give finite-size corrections of order 1/*N*. Furthermore the coefficient α_0 in Eq. (7) is evaluated explicitly in order to extract the crossover size N_c :

$$
\alpha_0(t_h=0) = -\beta(0)(1 + C - 2\ln 2), \quad (A7)
$$

$$
\alpha_0(t_h > 0) = -\beta(t_h) \left[1 + C + \ln \frac{2t(\pi t_h + 2t)}{(\pi t_h + 4t)^2} + \frac{\pi t_h}{t} \ln \frac{\pi t_h + 2t}{\pi t_h + 4t} + \ln \frac{(\pi t + 4t_h)^2}{8t_h(\pi t + 2t_h)} \right].
$$
\n(A8)

The crossover size N_c is the solution of the equation,

$$
- \beta(0) \ln N_c + \alpha_0(0) = - \beta(t_h) \ln N_c + \alpha_0(t_h).
$$
 (A9)

The crossover occurs for $t_h \ll t$, where $\alpha_0(t_h)$ diverges logarithmically. As a consequence, N_c is proportional to t_h/t :

$$
N_c = \frac{\pi t}{2t_h} \exp(-1 - C) \approx 0.3244 \frac{t}{t_h}.
$$
 (A10)

Arbitrary density. The scaling law (7) applies to any density. Furthemore, the crossover size is proportional to t/t_h , and diverges as $1-\rho$ in the limit $\rho \rightarrow 1$.

- $⁷$ J. Gavoret, P. Nozières, B. Roulet, and M. Combescot, J. Phys.</sup> (Paris) **30**, 987 (1969); S. Doniach, Phys. Rev. B 2, 3898 (1970); A. E. Ruckenstein and S. Schmitt-Rink, *ibid.* **35**, 7551 (1987).
- 8T. Kopp, A. E. Ruckenstein, and S. Schmitt-Rink, Phys. Rev. B 42, 6850 (1990).
- ⁹ J. M. Calleja, A. R. Goñi, B. S. Dennis, J. S. Weiner, A. Pinczuk, S. Schmitt-Rink, L. N. Pfeiffer, K. W. West, J. F. Müller, and A. E. Ruckenstein, Solid State Commun. 79, 911 (1991); Surf. Sci. 263, 346 (1992); A. R. Goni *et al.*, Phys. Rev. Lett. 67, 3298 $(1991).$
- 10 J. F. Müller, A. E. Ruckenstein, and S. Schmitt-Rink, Phys. Rev. B 45, 8902 (1991); T. Ogawa, A. Furusaki, and N. Nagaosa, Phys. Rev. Lett. **68**, 3638 (1992); F. J. Rodriguez and C. Tejedor, Phys. Rev. B 47, 1506 (1993); 47, 13 015 (1993).
- 11 I. E. Perakis and Y. C. Chang, Phys. Rev. B 47, 6573 (1993).
- 12 A. Rosch and T. Kopp, Phys. Rev. Lett. **75**, 1988 (1995).
- 13 F. Sols and F. Guinea, Phys. Rev. B 36, 7775 (1987); K. Vladár, G. T. Zima´nyi, and A. Zawadowski, Phys. Rev. Lett. **56**, 286 (1986); N. V. Prokof'ev, J. Mosc. Phys. Soc. 2, 157 (1992); P. Hedegård and A. Caldeira, Phys. Scr. 35, 609 (1987).
- 14C.-O.Almbladh and L. Hedin, in *Handbook on Synchrotron Ra*diation, edited by E. E. Koch (North-Holland, Amsterdam, 1983), Vol. B, p. 611.
- ¹⁵ J. B. McGuire, J. Math. Phys. **6**, 432 (1965); **7**, 123 (1966).
- 16 H. Castella and X. Zotos, Phys. Rev. B 47, 16 186 (1993).
- 17 D. M. Edwards, Prog. Theor. Phys. Suppl. 101 , 453 (1990); W.

von der Linden and D. M. Edwards, J. Phys. C 3, 4917 (1991). ¹⁸ S. Sorella, Phys. Rev. B **49**, 12 373 (1994).

- 19S. Sorella, S. Baroni, R. Car, and M. Parinello, Europhys. Lett. **8**, 663 (1989).
- 20E. Y. Loh and J. E. Gubernatis, in *Electrons Phase Transitions*, edited by B. S. Shastry, S. S. Jha, and V. Singh (Springer, Berlin, 1985).
- 21S. Sorella, A. Parola, M. Parinello, and E. Tosatti, Europhys. Lett. **12**, 721 (1990).
- ²²E. R. Hansen, *A Table of Series and Products* (Prentice-Hall, Englewood Cliffs, NJ, 1975).