

## Optical and magneto-optical absorption in parabolic quantum wells

L. Brey, N. F. Johnson, and B. I. Halperin

Physics Department, Harvard University, Cambridge, Massachusetts 02138

(Received 17 July 1989)

We show that an  $n$ -doped parabolic quantum well absorbs far infrared radiation at the bare harmonic-oscillator frequency  $\omega_0$  *independently* of the electron-electron interaction and the number of electrons in the well. In the presence of a magnetic field tilted with respect to the plane of the quantum well, we find that the cyclotron resonance becomes coupled to this  $\omega_0$  frequency mode. The absorption then occurs at two frequencies, which are again independent of the electron-electron interaction and the fractional filling of the well.

Wide parabolic quantum wells have been proposed<sup>1-4</sup> as structures in which a high-mobility quasi-three-dimensional electron gas can be realized. A parabolic potential of width  $W$  and height  $\Delta_1$  is equivalent to the potential created by a uniform slab of positive charge with a thickness  $W$  and density  $n_0 = 2\epsilon\Delta_1/W^2 e^2 \pi$ , where  $e$  is the electron charge and  $\epsilon$  is the dielectric constant (taken to be constant in the well). Electrons, which arise from donor impurities located away from the well, enter the well and screen this "fictitious" potential, forming a uniform layer of density  $n_0$  (see Fig. 1). By increasing the number of electrons per unit area  $n_s$  between zero and  $n_0 W$ , the thickness of the electron layer will increase linearly between zero and  $W$ , and a thick slab of electrons ( $> 2000 \text{ \AA}$ ) can be achieved.

These wide parabolic quantum wells have recently been grown<sup>2,3</sup> by tailoring the conduction-band edge of a graded  $\text{Ga}_{1-x}\text{Al}_x\text{As}$  semiconductor, and magnetotransport experiments on these systems<sup>2,3</sup> confirm the existence of a thick slab of high-mobility electron gas. In addition Karrai *et al.*<sup>5</sup> have reported measurements<sup>5</sup> of far-infrared magnetotransmission in these parabolic wells. These experiments reveal a coupling between the cyclotron resonance and a frequency  $\Omega_x$ , as the magnetic field is tilted

with respect to the electron-slab plane. The experiments suggest that  $\Omega_x$  corresponds to the plasma frequency [ $\Omega_x = (4\pi n_0 e^2 / \epsilon m^*)^{1/2}$ ] of a three-dimensional electron gas of density  $n_0$ , which *by construction* is equal to the frequency of the bare harmonic oscillator potential. This frequency  $\Omega_x$  is very different from the energy separation between the subbands of the self-consistent Hartree potential. Karrai *et al.*<sup>5</sup> therefore conclude that the three-dimensional character of the electron gas tends to predominate over the two-dimensional properties.<sup>5</sup>

In this paper we study the optical absorption of an  $n$ -doped parabolic quantum well. In addition we investigate the coupling between the excitations of this system and the cyclotron energy in the presence of a magnetic field which is tilted with respect to the quantum-well plane. We find that in the ideal parabolic well the absorption spectrum is independent of the electron-electron interaction, and also *independent of the number of electrons in the well*.

First, in the absence of any magnetic field, the Hamiltonian of our system (with electrons in the  $x$ - $y$  plane) is given within the effective-mass approximation by

$$H = \frac{1}{2m^*} \sum_{i=1}^N (p_{i,x}^2 + p_{i,y}^2 + p_{i,z}^2) + \sum_{i=1}^N \frac{\omega_0^2 m^*}{2} z_i^2 + U, \quad (1)$$

where  $\mathbf{p}_i$  and  $\mathbf{r}_i$  are respectively the momentum and position operators of the  $i$ th particle,  $m^*$  is the electron effective mass of the host semiconductor,  $\omega_0 = (8\Delta_1/W^2 m^*)^{1/2}$  is the bare harmonic-oscillator frequency of the parabolic well, and  $U$  is the interaction between electrons

$$U \equiv \sum_{i < j} u(\mathbf{r}_i - \mathbf{r}_j).$$

We define raising and lowering operators,

$$\hat{c}^\pm \equiv \sum_{i=1}^N (m^* \omega_0 z_i \mp i p_{i,z}). \quad (2)$$

It then follows from the quadratic form of the man-made potential that

$$[H, \hat{c}^\pm] = \pm \hbar \omega_0 \hat{c}^\pm. \quad (3)$$

If  $\Psi_n$  is an eigenstate of  $H$  with eigenenergy  $E_n$ , Eq. (3) implies

$$H \hat{c}^\pm \Psi_n = (\pm \hbar \omega_0 + E_n) \hat{c}^\pm \Psi_n. \quad (4)$$

Defining  $\Psi_{n\pm 1} \equiv \hat{c}^\pm \Psi_n$ , it is clear that  $\Psi_{n\pm 1}$  is an exact eigenstate of our Hamiltonian with energy  $E_{n\pm 1} = E_n$

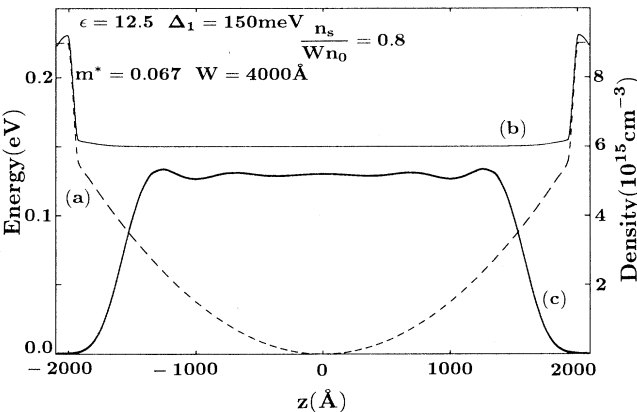


FIG. 1. Calculated electron density profile and self-consistent potential for an 80% full  $\text{Al}_x\text{Ga}_{1-x}\text{As}$  parabolic well of total width 4000  $\text{\AA}$ . (a) Potential for the empty well. (b) Self-consistent total potential in the Hartree approximation. (c) Self-consistent charge density profile as obtained in the Hartree approximation. The parameters of the well are given in the figure.

$\pm \hbar \omega_0$ .

When the system is placed in an electric field  $E$  applied in the  $z$  direction, the following term must be added to the Hamiltonian:

$$H' = \sum_{i=1}^N E e^{-i\omega t} z_i = E \frac{e^{-i\omega t}}{2m^* \omega_0} (\hat{c}^+ + \hat{c}^-). \quad (5)$$

This perturbation will only connect the state  $\Psi_N$  with the states  $\Psi_{N\pm 1}$ . A sharp peak in the absorption spectra is therefore expected at the frequency  $\omega_0$ . This result is not affected by the electron-electron interaction  $U$ . The parabolic well consequently absorbs light at the frequency  $\omega_0$  independently of the number of electrons in the well. We conclude that in the case of a parabolic quantum well the dielectric function diverges at the bare harmonic-oscillator frequency  $\omega_0$ , and not at the frequencies of the intersubband separation obtained in a Hartree or more sophisticated calculation.

The bare harmonic-oscillator frequency  $\omega_0$  is equal by

$$H = \frac{1}{2m^*} \sum_{i=1}^N \{p_{i,x}^2 + [p_{i,y} + (eB/c)(x \cos \theta - z \sin \theta)]^2 + p_{i,z}^2\} + \sum_{i=1}^N \frac{\omega_0^2 m^*}{2} z_i^2 + U. \quad (6)$$

This Hamiltonian can be simplified through a change of coordinates  $x \rightarrow x'$  and  $z \rightarrow z'$  corresponding to a rotation of angle  $\alpha$  with respect to the  $y$  axis,<sup>9,10</sup>

$$H = \frac{1}{2m^*} \sum_{i=1}^N (p_{i,x'}^2 + p_{i,y}^2 + p_{i,z'}^2) + \frac{m^*}{2} \sum_{i=1}^N (\omega_1^2 x_i'^2 + \omega_2^2 z_i'^2) + U + \omega_c \sum_{i=1}^N [\sin(\alpha - \theta) z_i' p_{i,y} + \cos(\alpha - \theta) x_i' p_{i,y}]. \quad (7)$$

The rotation angle  $\alpha$  is obtained from

$$\tan(2\alpha) \equiv \frac{\omega_c^2 \sin 2\theta}{\omega_c^2 \cos 2\theta - \omega_0^2}. \quad (8)$$

The frequencies  $\omega_{1,2}$  are

$$\omega_{1,2} \equiv \left[ \frac{1}{2} (\omega_c^2 + \omega_0^2) \pm \frac{1}{2} (\omega_c^4 + \omega_0^4 - 2\omega_0^2 \omega_c^2 \cos 2\theta)^{1/2} \right]^{1/2}, \quad (9)$$

and  $\omega_c = eB/m^*c$  is the cyclotron frequency. Defining the raising and lowering operators

$$\hat{a}^\pm \equiv \sum_{i=1}^N \left[ m^* \omega_1 x_i' \mp i p_{i,x'} + \frac{\omega_c}{\omega_1} \cos(\alpha - \theta) p_{i,y} \right], \quad (10)$$

$$\hat{b}^\pm \equiv \sum_{i=1}^N \left[ m^* \omega_2 z_i' \mp i p_{i,z'} + \frac{\omega_c}{\omega_2} \sin(\alpha - \theta) p_{i,y} \right],$$

it follows from Eq. (7) that

$$[\hat{a}^\pm, \hat{b}^\pm] = 0, \quad (11)$$

$$[H, \hat{a}^\pm] = \pm \hbar \omega_1 \hat{a}^\pm,$$

$$[H, \hat{b}^\pm] = \pm \hbar \omega_2 \hat{b}^\pm.$$

If  $\Psi_{n_1, n_2}$  is an eigenstate of  $H$  with eigenvalue  $E_{n_1, n_2}$ , then

$$\hat{H} \hat{a}^\pm \Psi_{n_1, n_2} = (\pm \hbar \omega_1 + E_{n_1, n_2}) \hat{a}^\pm \Psi_{n_1, n_2}, \quad (12)$$

$$\hat{H} \hat{b}^\pm \Psi_{n_1, n_2} = (\pm \hbar \omega_2 + E_{n_1, n_2}) \hat{b}^\pm \Psi_{n_1, n_2}.$$

Defining  $\Psi_{n_1 \pm 1, n_2} \equiv \hat{a}^\pm \Psi_{n_1, n_2}$  and  $\Psi_{n_1, n_2 \pm 1} \equiv \hat{b}^\pm \Psi_{n_1, n_2}$ , we see that  $\Psi_{n_1 \pm 1, n_2}$  and  $\Psi_{n_1, n_2 \pm 1}$  are both exact eigenstates with energies  $E_{n_1 \pm 1, n_2} = E_{n_1, n_2} \pm \hbar \omega_1$  and  $E_{n_1, n_2 \pm 1} = E_{n_1, n_2} \pm \hbar \omega_2$ , respectively. When illuminated with long wave-

length light, polarized in the  $z$  direction ( $x$  direction), the optical absorption of the parabolic well will present two peaks at  $\omega_1$  and  $\omega_2$  with intensities proportional to  $\sin^2 \alpha$  ( $\cos^2 \alpha$ ) and  $\cos^2 \alpha$  ( $\sin^2 \alpha$ ), respectively. Once again this result is independent of the electron-electron interaction and is a direct consequence of the parabolic form of the bare quantum-well potential. If the term corresponding to the parabolic well is dropped from the Hamiltonian equation (6), then  $\omega_1 = \omega_c$ ,  $\omega_2 = 0$  and we recover the Kohn result<sup>11</sup> that the cyclotron resonance  $\omega_c$  in a bulk three-dimensional gas is not affected by the electron-electron interaction  $U$ .

In the experimental samples the width of the parabolic well is finite, and the well is confined by an additional barrier (see Fig. 1). This extra confinement will not affect our results as long as the number of electrons in the well is small enough so that the self-consistent charge-density profile remains essentially unchanged. We have checked this numerically in the case of no magnetic field by calculating the dynamical conductivity<sup>12,13</sup> for the potential profile shown in Fig. 1. For a density of electrons per unit area  $n_s$  bigger than  $0.9Wn_0$ , we find that small satellites appear in the optical absorption close to  $\omega_0$ . Of course, impurities or other imperfections in the well can lead to

additional absorption even at smaller values of  $n_s$ . Variation of  $m^*$  across the parabolic well will also give corrections to these results.

In conclusion, we have shown that an ideal parabolic well absorbs light at the bare harmonic-oscillator frequency  $\omega_0 = (8\Delta_1/W^2m^*)^{1/2}$  independently of the electron-electron interaction. In addition, the presence of a tilted magnetic field couples the cyclotron frequency with  $\omega_0$  and *not* with the frequency corresponding to the self-consistent intersubband separation.

We thank D. Vanderbilt, J. Dempsey, R. Westervelt, A. C. Gossard, and R. Meade for useful discussions. L. Brey wishes to acknowledge support from Spain's Ministerio de Educacion y Ciencia. This work was supported by the National Science Foundation through the Harvard Materials Research Laboratory and Grant No. DMR88-17291, U.S. Defense Advanced Research Projects Agency (DARPA) through U.S. Office of Naval Research (ONR) Contract No. N00014-86-K-0033, and by St. John's College, Cambridge, England.

<sup>1</sup>B. I. Halperin, Jpn. J. Appl. Phys. **26**, Suppl. 26-3, 1913 (1987).

<sup>2</sup>E. G. Gwinn, R. M. Westervelt, P. F. Hopkins, A. J. Rimberg, M. Sundaram, and A. G. Gossard, Phys. Rev. B **39**, 6260 (1989); A. J. Rimberg and R. M. Westervelt, *ibid.* **40**, 3970 (1989).

<sup>3</sup>T. Sajoto, J. Jo, L. Engel, M. Santos, and M. Shayegan, Phys. Rev. B **39**, 10464 (1989).

<sup>4</sup>L. Brey and B. I. Halperin, in *Proceedings of the Eighth International Conference on the Electronic Properties of Two Dimensional Systems, Grenoble, France, 1989* (to be published); Phys. Rev. B (to be published).

<sup>5</sup>K. Karrai, H. D. Drew, M. W. Lee, and M. Shayegan, Phys. Rev. B **39**, 1426 (1989).

<sup>6</sup>See, e.g., N. W. Ashcroft and N. D. Mermin, *Solid State Physics* (Holt, New York, 1976), p. 19.

<sup>7</sup>D. A. Dahl and L. J. Sham, Phys. Rev. B **16**, 651 (1977).

<sup>8</sup>D. Pines, *The Many-Body Problem* (Dunod-Wiley, New York, 1959), p. 202.

<sup>9</sup>J. C. Maan, in *Two-Dimensional Systems, Heterostructures and Superlattices*, edited by G. Bauer, F. Kuchar, and H. Heinrich, Solid State Sciences, Vol. 53 (Springer-Verlag, Berlin, 1984).

<sup>10</sup>R. Merlin, Solid State Commun. **64**, 99 (1987).

<sup>11</sup>W. Kohn, Phys. Rev. **123**, 1242 (1961).

<sup>12</sup>T. Ando, Z. Phys. B. **26**, 263 (1977).

<sup>13</sup>K. S. Yi and J. J. Quinn, Phys. Rev. B **27**, 2396 (1983).