Field dependences of magnetization and specific-heat coefficient in a nearly magnetic system: Liquid 3He and strong Pauli paramagnets

M. T. Béal-Monod and E. Daniel*

Laboratoire de Physique des Solides, associé au Centre National de la Recherche Scientifique, Universite de Paris-Sud, 91400 Orsay, France

(Received 6 July 1982)

We analytically derive the first-order magnetic field dependences of the magnetization M and of the coefficient γ of the specific heat, in a nearly magnetic paramagnet, at low temperature.

The contribution of spin fluctuations, "paramag-The contribution of spin fluctuations, "paramagnons," to the temperature dependence of the spin susceptibility of a nearly magnetic Fermi liquid was calculated long ago,¹ for a parabolic band, with no adjustable parameter, in terms of the Stoner enhancement S given by experiments. The agreement with the data on normal liquid 3 He was very good for all available values of S. An extension was recently made² to account for a more general band structure. Polarized liquid ³He as well as strongly exchangeenhanced metallic paramagnets, involving strong spin-spin interaction I among fermions, have received renewed interest due to recent measurements in high magnetic fields³: deviations of the magnetization $M(T,H)$ from linearity with the field H, H dependence of the coefficient of the specific heat $(\gamma = [C(T,H)/T]_{T=0})$. As recalled in Ref. 3, due to quantum effects,⁴ the critical exponents at the ferromagnetic transition at 0 K, should assume their mean-field values. The critical spin fluctuations will not affect the $T = 0$ powers of the Stoner enhancement S. It was thus expected³ that the H dependence of M would qualitatively be the same at 0 K, as that of the Stoner-type result.⁵ But at finite T , paramagnons modify by one power of S the zero-field T dependence of the susceptibility,¹ compared to the Stoner result.

It is the purpose of the present paper to provide

analytical formulas, at finite H and T, for $M(T,H)$ and $\gamma(H)$, taking paramagnons into account. We give the result for a parabolic band but it could be generalized for a given band structure. Other authors⁶ recently computed such field dependences, but their results were given under the form of complicated integrals difficult to compare with our present results; we note though that in Ref. 6 H always appears through the combination SH while here it rather involves $S^{3/2}H$. This last combination, we believe, ought to be the correct one due to the abovementioned quantum effects imposing the critical exponent in $H \sim M^8$ to assume its mean-field value $\delta = 3$.

We start with the same paramagnon formalism as that in Ref. 1 (hereafter referred to as BMMF} with the same notations, although M is used for the total magnetization, and to simplify we suppose that the number of particles per unit volume $N/V = 1$, which does not affect the final result:

$$
F(T,M) = G_0(T,B) + I(1-M^2)/4 + \Delta F(T,B) - MB
$$
\n(1)

B is defined by $(\partial F/\partial B)_{T,M} = 0$; besides, H $=\partial F/\partial M$. These two equations yield

$$
M = [-\chi_{\text{Pauli}}B + (\partial \Delta F/\partial B)]_{-B=H+IM/2} .
$$
 (2)

 ΔF is given by the same diagrams as in BMMF:

$$
\Delta F = (T/2) \sum_{\vec{q}, \omega} \{ \ln(1 - I^2 \chi^{0++} \chi^{0-}) + I^2 \chi^{0++} \chi^{0-} + \ln(1 - I \chi^{0+}) + I \chi^{0+} + \ln(1 - I \chi^{0-}) + I \chi^{0-} \} ,
$$
 (3)

in terms of the dynamic susceptibilities in the absence of interactions,

$$
\chi^{0\alpha,\beta}=-\sum_{\vec{p}}\left\{\left(f\frac{\alpha}{\vec{p}}-f\frac{\beta}{\vec{p}+\vec{q}}\right)/\left[\omega+\xi\frac{\alpha}{\vec{p}}-\xi\frac{\beta}{\vec{p}+\vec{q}}+i\,\eta\,\text{sgn}\left(\xi\frac{\beta}{\vec{p}+\vec{q}}-\xi\frac{\alpha}{\vec{p}}\right)\right]\right\}\ ;
$$

 α , β are spin indices, the f's are the Fermi functions expressed in terms of the fermions energies in atomic unit $\xi_{\overline{n}}^{\alpha,\beta} = p^2/2 - \mu + (\alpha, \beta)B$, μ is the chemical potential, B is in units of energy. All that is reproduced from BMMF. Now straightforward algebra gives

$$
Rex^{0++} = N(E_F) \left(\frac{h_-}{2} + \frac{1}{8\bar{q}} (h_-^2 - x^2) \ln \frac{h_- + x}{h_- - x} - \frac{1}{8\bar{q}} (h_-^2 - y^2) \ln \frac{h_- + y}{h_- - y} \right),
$$

\n
$$
Rex^{0--}(h) = Rex^{0++}(-h) ,
$$

\n
$$
h_{\mp} = (1 \mp h)^{1/2}, h = B/E_F, x = \frac{\bar{\omega}}{4\bar{q}} + \bar{q}, y = \frac{\bar{\omega}}{4\bar{q}} - \bar{q}, \bar{\omega} = \frac{\omega}{E_F}, \bar{q} = \frac{q}{2k_F}
$$
 (4)

$$
27 \quad \overline{}
$$

4467 1983 The American Physical Society

BRIEF REPORTS 27

$$
Rex^{0+-} = N(E_F) \left[\frac{h_+x_+ - h_-y_+}{4\overline{q}} + \frac{1}{8\overline{q}} (h_+^2 - x_+^2) \ln \frac{h_+ + x_+}{h_+ - x_+} - \frac{1}{8\overline{q}} (h_-^2 - y_+^2) \ln \frac{h_- + y_+}{h_- - y_+} \right],
$$

\n
$$
Rex^{0-}(\overline{\omega}, h) = Rex^{0+}(-\overline{\omega}, h) = Rex^{0+}(\overline{\omega}, -h) ,
$$

\n
$$
x_{\pm} = x_{\pm} h/2\overline{q}, y_{\pm} = y_{\pm} h/2\overline{q} .
$$
\n(5)

 $N(E_F)$ is the density of states per spin at the Fermi level, k_F the Fermi momentum. In the region of interest, where Im $\chi^{0\alpha\beta}$ is linear with $\bar{\omega}$, one has

$$
\text{Im}\chi^{0\alpha\beta} = N(E_F)(\pi/8)(\bar{\omega}/\bar{q})\lambda^{\alpha\beta} , \quad \lambda^{\pm \pm} = \theta(h_{\mp}^2 - y^2)\theta(h_{\mp}^2 - x^2) , \quad \lambda^{\pm \mp} = \theta(h_{\mp}^2 - y_{\pm}^2)\theta(h_{\pm}^2 - x_{\pm}^2) . \tag{6}
$$

 λ^{\pm} is defined in a domain symmetric from the one of λ^{\pm} with respect to the \bar{q} axis. The above-defined domains played a key role in the further integrations over \bar{q} and $\bar{\omega}$. Note that in the $(\bar{\omega}, \bar{q})$ plan there are gaps for small \bar{q} and $\bar{\omega}$ around the origin, for the transverse susceptibilities due to the presence of B. The $\bar{\omega} = 0$, $\bar{q} < 1$, and $h < 1$ expansions are useful:

$$
Re x^{0+-(\overline{\omega}=0)} = Re x^{0-+(\overline{\omega}=0)} \approx N(E_F) \left[1 - \frac{\overline{q}^2}{3} - \frac{h^2}{8} \left(\frac{1}{3} + \frac{\overline{q}^2}{5} \right) + \cdots \right],
$$

\n
$$
(Re x^{0++} Re x^{0-})_{\overline{\omega}=0} \approx N^2(E_F) \left[1 - \frac{2\overline{q}^2}{3} - \frac{h^2}{2} \left[1 + \frac{2\overline{q}^2}{3} \right] + \cdots \right],
$$

\n
$$
Re x^{0++(\overline{\omega}=0)} + Re x^{0--(\overline{\omega}=0)} \approx 2N(E_F) \left[1 - \frac{\overline{q}^2}{3} - \frac{h^2}{8} (1 + \overline{q}^2) + \cdots \right].
$$

\n(7)

We found it more convenient to calculate $[\Delta F(B) - \Delta F(0)]$ which makes immediately evident the finite terms in 8. We find

$$
\Delta F(T, B) - \Delta F(T, 0) = \chi_{\text{Pauli}}[\alpha(\overline{T}, \overline{I})B^2/2 + \beta(\overline{T}, \overline{I})B^4/(4E_f^2) + \cdots] \tag{8}
$$

in appropriate units, where α and β are functions of $\bar{T} = T/T_F$ and of the dimensionless interaction $\bar{I} = IN(E_F)$, or the Stoner factor $S = (1 - \overline{I})^{-1}$; in units of the square of the magnetic moment, χ_{Pauli} is $2N(E_F)$. Then (2) and (8) yield

$$
M(T,H) = \chi_{\text{Pauli}} H \left[\frac{1-\alpha}{1-\overline{I}+\alpha\overline{I}} - \beta \frac{H^2}{E_f^2} \frac{1}{(1-\overline{I}+\alpha\overline{I})^4} \cdots \right] \tag{9}
$$

Expanding further, for low $T \ll T_{sf}$, and low $H \ll T_{sf}/\sqrt{S}$, with the spin fluctuation temperature $T_{\text{sf}} = (1 - \overline{I}) T_F$ (T_F the bare Fermi temperature), we obtain

$$
M(T,H) = S\chi_{\text{Pauli}}H\left[1-\alpha_1S^2\frac{T^2}{T_f^2}-\beta_0S^3\frac{H^2}{T_f^2}+(\beta_1+4\alpha_1\beta_0)S^2\frac{T^2}{T_f^2}S^3\frac{H^2}{T_f^2}+\cdots\right].
$$
\n(10)

On the other hand we find (for a parabolic bang)

$$
\beta_0 \approx \frac{1}{6}, \quad \alpha_1 \approx \pi^2/6, \quad \beta_1 \approx 23\pi^2/24^2, \quad H << T
$$
\n
$$
\beta_0 = \frac{1}{6}, \quad \alpha_1 \approx \pi^2/4, \quad \beta_1 \approx 27\pi^2/24^2, \quad T << H \quad . \tag{11}
$$

Several remarks arise at that point:

(i) As recalled above and detailed in Ref. 4, the fluctuations do not affect the $T = 0$ behaviors: $\beta(T=0) = \beta_0$ is the value computed in the Stoner-Wohlfarth theory⁵; $\alpha(T=0) = \alpha_0$ is assumed to have been incorporated in the definition of I in (9) and is extracted from experiments. In contrast, the T dependences of α and β do diverge with S: The fluctuations greatly enhance the finite temperature dependence of M. For comparsion, the analog of

(10) in the absence of fluctuations, as given in Ref. 5, would contain ST^2/T_F^2 instead of S^2T^2/T_F^2 .

(ii) $(M/H)_{H=0}$ must be compared to the result $\chi(T, 0)$ in BMMF. The $H \ll T$ limit of α in (11) is $\pi^2/6 = 4\pi^2/24$ to be compared with the result in BMMF, 3.2 $\pi^2/24 = [4 - (8/\pi^2)]\pi^2/24$. The extra term $8/\pi^2$ arose from less-important contributions in the integrals that we have neglected here; nevertheless the coefficient in BMMF, $3.2\pi^2/24 \approx 1.3$, was more accurate than (although very close to) the one we have here, $4\pi^2/24 \approx 1.6$.

(iii) The first term in (9) appeared with such a form in the theory of Moriya and Kawabata.^{\prime} As already pointed out elsewhere, $⁸$ their formalism is</sup> identical to the one derived earlier in BMMF that we use here.

4468

(iv) As emphasized in Refs. 3, 8, and 9, one notes the scaling in ST/T_F and the one in $S^{3/2}H/T_F$

 $=\sqrt{S}H/T_{\rm sf}$ which is in agreement with the relation, at the ferromagnetic transition $(T = 0, \overline{I} = 1)$, $H \sim M^{3-3}$, where one replaces M by $\chi H \sim S H$. Thus the results of Ref. 6, which do not contain the same scaling, appear doubtful to us.

(v) From (10) , it is clear that the deviations of M from linearity with respect to H are less pronounced when T increases. To observe the nonlinear

behavior, one would need higher fields at high temperatures than at low T. For a different band structure, when α_1 , β_0 , and β_1 would be all negative instead of positive here, then $M(T,H)/H$ would increase with H (instead of decreasing here), but less and less so when the temperature increases; this is in qualitative agreement with what is found in TiBE₂.¹⁰ (vi) From the Maxwell relation^{1,3} $\partial M/\partial T$

 $= \frac{\partial S}{\partial H}$, we deduce $\frac{\partial^2 M}{\partial T^2} = \frac{\partial \gamma}{\partial H}$, which thus yields with (10) (for a parabolic band)

$$
\gamma(H) - \gamma(0) = -\chi_{\text{Pauli}}\alpha_1 \left[\frac{S^{3/2}H}{T_F}\right]^2 \left[1 - \left(\frac{\beta_1}{2\alpha_1} + 2\beta_0\right) \left(\frac{S^{3/2}H}{T_F}\right)^2 \cdots \right] \ . \tag{12}
$$

As pointed out in Ref. 3, the lowest field dependence of $\gamma(H) - \gamma(0)$, i.e., the first term in (12), follows from the curvature of $\chi(T,H=0)$ at $T=0$. However, at higher fields, or strong enough values of $[\beta_1/(2\alpha_1) + 2\beta_0]$, the H^4 term in (12) being of opposite sign, may counterbalance the first one. This seems to be the case in $TiBe₂$ (Ref. 11) where, at $H = 7$ T, $\gamma(H) \leq \gamma(0)$, while $\chi(T, H = 0)$ first increases with T ; however, it was pointed out to us¹² that at 7 T, $\chi(T,H)$ decreases rather than increases when T increases. Therefore the predicted values given in Ref. 3 for the relative variation of $\gamma(H)$ at 7 T was erroneous since it corresponded to the extrapolation of the first term in (12) to a field value (7 T) where $\chi(T,H) - \chi(0,H)$ does not have the same sign that $\chi(T,H \simeq 0) - \chi(0,H \simeq 0)$; in that case, the 2nd term in (12), but also higher-order ones (not computed here), ought to be checked. In contrast, for Pd, where M remains linear in H up to about 35 T (see Ref. 39 in Ref. 3), the $H²$ term in (10) and the H^4 term in $\gamma(H)$ must be very tiny and, from $\chi(T, 0)$ increasing with T, one expected³ a small increase of $\gamma(H)$ with H; some experiments (Ref. 34) of Ref. 3) found a strong decrease while others¹³ did observe recently a small increase.

(vii) A discussion was given elsewhere' concerning the T variation of the nuclear relaxation rate T_1^{-1} in strongly enhanced paramagnets, recovering the proportionality to $T\chi(T)$ already proposed by Moriya and Ueda¹⁴ but also pointing out a scaling in T/T_{sf} below and above T_{sf} . If $\chi(T,H)$ replaces $\chi(T, 0)$ in the finite field dependence of T_1^{-1} , then one can
write at low $T \ll T_{\text{sf}}$ and $H \ll T_{\text{sf}}/\sqrt{S}$, T_1^{-1} $\propto T[M(T,H)/H]$, with $M(T,H)$ given by (10). This is qualitatively analogous to the result derived in Ref. 15 but with different numerical coefficients. We note however that Ref. 15 used a supposedly general expansion for small \bar{q} and $\bar{\omega}/\bar{q}$ for the transverse susceptibility x^{0-+} , which appears incomplete to us; such terms as h/\bar{q}^2 [that we have in (5)] are missing, for

instance. The respective ratios of the four small quantities \bar{q} , $\bar{\omega}/\bar{q}$, \bar{T} , and h must be handled with special care since the domain of integration in the $(\bar{\omega}, \bar{q})$ plan is very tricky as can be seen from (6).

Our main result (10), with positive coefficients (11) can apply to two physical cases: liquid 3 He, where there is no band structure and (11) applies as such, and $UA1₂$ which behaves the same way (Ref. 7) of Ref. 3), but whose band structure has not been calculated so far. Recent experiments on this last material¹⁶ exhibit a variation of M vs H in qualitative agreement with our result: M deviates from linearity with H around 15 T at low T but remains linear at much higher fields at high T , for which we expect that one would need higher fields to observe nonlinearity. A quantitative comparison is difficult though, since our result only applies for $T \ll T_{\rm sf}$, and for $H \ll T_{\rm sf}/(S\beta_0)^{1/2}$, conditions which are not all fulfilled in the experiments. One remark: for this compound where $S \sim 4$, $(S\beta_0)^{1/2} \approx 1$ so that the characteristic field equal to $T_{sf}/(S\beta_0)^{1/2}$ is practically equal to T_{sf} , but for higher values of S, for instance in TiBe₂ ($S \sim 65$), the two values will be different. For liquid 3 He, we expect deviations from linearity for M vs H to be of order 2% at 100 kG at 15 mK, at melting pressure where $S \sim 20$; but in confined melting pressure where $S \sim 20$; but in confined geometries, if S can be as large as 60 ,¹⁷ then the corresponding effects are expected to be larger, \sim 25%. The characteristic field where nonlinearity is expected to occur is estimated to be about 90 T at melting pressure for the bulk, but may be reached below \sim 18 T in confined geometry if an S value of ~ 60 can be achieved.

ACKNOWLEDGMENTS

We thank P. Nozieres for having suggested that calculation to us, in connection with further polarized ³He studies, and we acknowledge useful discussions with K. Levin and O. Valls.

- *Permanent address: Université Louis Pasteur, 4 rue Blaise Pascal, F-6700 Strasbourg, France.
- ¹M. T. Béal-Monod, S. K. Ma, and D. R. Fredkin, Phys. Rev. Lett. 20, 929 (1968); H. Ramm et al., J. Low Temp. Phys. 2, 539 (1970).
- ²M. T. Béal-Monod and J. Lawrence, Phys. Rev. B 21, 5400 (1980).
- ³See a review in M. T. Béal-Mond, Physica (Utrecht), 109 and 110B, 1837 (1982).
- ⁴M. T. Beal-Monod and K. Maki, Phys. Rev. Lett. 34, 1461 (1975); J. Hertz, Phys. Rev. B 14, 1165 (1976).
- ⁵See E. P. Wohlfarth and P. Rhodes, Philos. Mag. 7, 1817 (1962); M. Shimizu, Rep. Prog. Phys. 44, 329 (1981).
- 6P. Hertel, J. Appel, and D. Fay, Phys. Rev. B 22, 534 (1980).
- ⁷T. Moriya and A. Kawabata, J. Phys. Soc. Jpn. 34, 639 (1973).
- 8M. T. Beal-Monod (unpublished).
- ⁹J. Lawrence and M. T. Béal-Monod, in Valence Fluctuations in Solids, edited by L. Falicov, W. Hanke, and M. Maple (North-Holland, Amsterdam, 1981), p. 53.
- ¹⁰See Ref. 10 in Ref. 3 above
- G. R. Stewart, J. L. Smith, and B. L. Brandt, Phys. Rev. B 25, 5907 (1982).
- 12 F. Acker (private communication
- ¹³J. J. M. Franse (private communication).
- ¹⁴T. Moriya and K. Ueda, Solid State Commun. 15, 169 $(1974).$
- $15K$. Ueda, Solid State Commun. 19, 965 (1976).
- 16 J. J. M. Franse et al., Phys. Rev. Lett. $48, 1749$ (1982).
- 17 For instance, the case described, at high pressures, in B. Herbal et al., Phys. Rev. Lett. 46, 42 (1981).