Field dependences of magnetization and specific-heat coefficient in a nearly magnetic system: Liquid ³He and strong Pauli paramagnets

M. T. Béal-Monod and E. Daniel*

Laboratoire de Physique des Solides, associé au Centre National de la Recherche Scientifique, Université de Paris-Sud, 91400 Orsay, France

(Received 6 July 1982)

We analytically derive the first-order magnetic field dependences of the magnetization M and of the coefficient γ of the specific heat, in a nearly magnetic paramagnet, at low temperature.

The contribution of spin fluctuations, "paramagnons," to the temperature dependence of the spin susceptibility of a nearly magnetic Fermi liquid was calculated long ago,¹ for a parabolic band, with no adjustable parameter, in terms of the Stoner enhancement S given by experiments. The agreement with the data on normal liquid ³He was very good for all available values of S. An extension was recently made² to account for a more general band structure. Polarized liquid ³He as well as strongly exchangeenhanced metallic paramagnets, involving strong spin-spin interaction I among fermions, have received renewed interest due to recent measurements in high magnetic fields³: deviations of the magnetization M(T,H) from linearity with the field H,Hdependence of the coefficient of the specific heat $(\gamma = [C(T,H)/T]_{T=0})$. As recalled in Ref. 3, due to quantum effects,⁴ the critical exponents at the ferromagnetic transition at 0 K, should assume their mean-field values. The critical spin fluctuations will not affect the T = 0 powers of the Stoner enhancement S. It was thus expected³ that the H dependence of M would qualitatively be the same at 0 K, as that of the Stoner-type result.⁵ But at finite T, paramagnons modify by one power of S the zero-field Tdependence of the susceptibility,¹ compared to the Stoner result.

It is the purpose of the present paper to provide

analytical formulas, at finite H and T, for M(T,H)and $\gamma(H)$, taking paramagnons into account. We give the result for a parabolic band but it could be generalized for a given band structure. Other authors⁶ recently computed such field dependences, but their results were given under the form of complicated integrals difficult to compare with our present results; we note though that in Ref. 6 H always appears through the combination SH while here it rather involves $S^{3/2}H$. This last combination, we believe, ought to be the correct one due to the abovementioned quantum effects imposing the critical exponent in $H \sim M^8$ to assume its mean-field value $\delta = 3$.

We start with the same paramagnon formalism as that in Ref. 1 (hereafter referred to as BMMF) with the same notations, although M is used for the total magnetization, and to simplify we suppose that the number of particles per unit volume N/V = 1, which does not affect the final result:

$$F(T,M) = G_0(T,B) + I(1-M^2)/4 + \Delta F(T,B) - MB$$
(1)

B is defined by $(\partial F/\partial B)_{T,M} = 0$; besides, *H* = $\partial F/\partial M$. These two equations yield

$$M = \left[-\chi_{\text{Pauli}}B + \left(\frac{\partial\Delta F}{\partial B}\right)\right]_{-B=H+M/2} \quad (2)$$

 ΔF is given by the same diagrams as in BMMF:

$$\Delta F = (T/2) \sum_{\vec{0}, \vec{v}} \{ \ln(1 - I^2 \chi^{0++} \chi^{0--}) + I^2 \chi^{0++} \chi^{0--} + \ln(1 - I \chi^{0+-}) + I \chi^{0+-} + \ln(1 - I \chi^{0-+}) + I \chi^{0-+} \} , \qquad (3)$$

in terms of the dynamic susceptibilities in the absence of interactions,

$$\chi^{0\alpha,\beta} = -\sum_{\vec{p}} \left\{ \left(f\frac{\alpha}{\vec{p}} - f\frac{\beta}{\vec{p}} + \vec{q} \right) / \left[\omega + \xi\frac{\alpha}{\vec{p}} - \xi\frac{\beta}{\vec{p}} + \vec{q} + i\eta \operatorname{sgn}\left(\xi\frac{\beta}{\vec{p}} + \vec{q} - \xi\frac{\alpha}{\vec{p}}\right) \right] \right\} ;$$

 α , β are spin indices, the f 's are the Fermi functions expressed in terms of the fermions energies in atomic units $\xi_{\vec{p}}^{\alpha,\beta} = p^2/2 - \mu + (\alpha,\beta)B$, μ is the chemical potential, B is in units of energy. All that is reproduced from BMMF. Now straightforward algebra gives

$$\operatorname{Rex}^{0++} = N(E_F) \left[\frac{h_-}{2} + \frac{1}{8\bar{q}} (h_-^2 - x^2) \ln \frac{h_- + x}{h_- - x} - \frac{1}{8\bar{q}} (h_-^2 - y^2) \ln \frac{h_- + y}{h_- - y} \right] ,$$

$$\operatorname{Rex}^{0--}(h) = \operatorname{Rex}^{0++}(-h) , \qquad (4)$$

$$h_{\mp} = (1 \mp h)^{1/2}, \ h = B/E_F, \ x = \frac{\bar{\omega}}{4\bar{q}} + \bar{q}, \ y = \frac{\bar{\omega}}{4\bar{q}} - \bar{q}, \ \bar{\omega} = \frac{\omega}{E_F}, \ \bar{q} = \frac{q}{2k_F}$$

4467

© 1983 The American Physical Society

BRIEF REPORTS

$$\operatorname{Re} \chi^{0+-} = N(E_F) \left\{ \frac{h_+ x_+ - h_- y_+}{4\bar{q}} + \frac{1}{8\bar{q}} (h_+^2 - x_+^2) \ln \frac{h_+ + x_+}{h_+ - x_+} - \frac{1}{8\bar{q}} (h_-^2 - y_+^2) \ln \frac{h_- + y_+}{h_- - y_+} \right\} ,$$

$$\operatorname{Re} \chi^{0+-}(\bar{\omega}, h) = \operatorname{Re} \chi^{0+-}(-\bar{\omega}, h) = \operatorname{Re} \chi^{0+-}(\bar{\omega}, -h) , \qquad (5)$$

$$x_+ = x \pm h/2\bar{q}, \ y_+ = y \pm h/2\bar{q} .$$

 $N(E_F)$ is the density of states per spin at the Fermi level, k_F the Fermi momentum. In the region of interest, where $Im \chi^{0\alpha\beta}$ is linear with $\overline{\omega}$, one has

$$\mathrm{Im} \chi^{0\alpha\beta} = N(E_F)(\pi/8)(\overline{\omega}/\overline{q})\lambda^{\alpha\beta} , \ \lambda^{\pm\pm} = \theta(h_{\pm}^2 - y^2)\theta(h_{\pm}^2 - x^2) , \ \lambda^{\pm\pm} = \theta(h_{\pm}^2 - y_{\pm}^2)\theta(h_{\pm}^2 - x_{\pm}^2) .$$
(6)

 λ^{\mp} is defined in a domain symmetric from the one of λ^{\pm} with respect to the \bar{q} axis. The above-defined domains played a key role in the further integrations over \bar{q} and $\bar{\omega}$. Note that in the $(\bar{\omega}, \bar{q})$ plan there are gaps for small \bar{q} and $\bar{\omega}$ around the origin, for the transverse susceptibilities due to the presence of *B*. The $\bar{\omega} = 0$, $\bar{q} < 1$, and h < 1 expansions are useful:

$$\operatorname{Re} \chi^{0+-}(\overline{\omega}=0) = \operatorname{Re} \chi^{0-+}(\overline{\omega}=0) \approx N(E_F) \left[1 - \frac{\overline{q}^2}{3} - \frac{h^2}{8} \left(\frac{1}{3} + \frac{\overline{q}^2}{5} \right) + \cdots \right] ,$$

$$\left(\operatorname{Re} \chi^{0++} \operatorname{Re} \chi^{0--}\right)_{\overline{\omega}=0} \approx N^2(E_F) \left[1 - \frac{2\overline{q}^2}{3} - \frac{h^2}{2} \left(1 + \frac{2\overline{q}^2}{3} \right) + \cdots \right] ,$$

$$\operatorname{Re} \chi^{0++}(\overline{\omega}=0) + \operatorname{Re} \chi^{0--}(\overline{\omega}=0) \approx 2N(E_F) \left[1 - \frac{\overline{q}^2}{3} - \frac{h^2}{8} (1 + \overline{q}^2) + \cdots \right] .$$

$$(7)$$

We found it more convenient to calculate $[\Delta F(B) - \Delta F(0)]$ which makes immediately evident the finite terms in B. We find

$$\Delta F(T,B) - \Delta F(T,0) = \chi_{\text{Pauli}}[\alpha(\overline{T},\overline{I})B^2/2 + \beta(\overline{T},\overline{I})B^4/(4E_F^2) + \cdots]$$
(8)

in appropriate units, where α and β are functions of $\overline{T} = T/T_F$ and of the dimensionless interaction $\overline{I} = IN(E_F)$, or the Stoner factor $S = (1 - \overline{I})^{-1}$; in units of the square of the magnetic moment, χ_{Pauli} is $2N(E_F)$. Then (2) and (8) yield

$$M(T,H) = \chi_{\text{Pauli}} H \left[\frac{1-\alpha}{1-\bar{I}+\alpha\bar{I}} - \beta \frac{H^2}{E_F^2} \frac{1}{(1-\bar{I}+\alpha\bar{I})^4} \cdots \right] .$$
(9)

Expanding further, for low $T \ll T_{sf}$, and low $H \ll T_{sf}/\sqrt{S}$, with the spin fluctuation temperature $T_{sf} = (1 - \overline{I})T_F$ (T_F the bare Fermi temperature), we obtain

$$M(T,H) = S\chi_{\text{Pauli}}H\left(1 - \alpha_1 S^2 \frac{T^2}{T_F^2} - \beta_0 S^3 \frac{H^2}{T_F^2} + (\beta_1 + 4\alpha_1 \beta_0) S^2 \frac{T^2}{T_F^2} S^3 \frac{H^2}{T_F^2} \cdots\right)$$
(10)

On the other hand we find (for a parabolic band)

$$\beta_0 \approx \frac{1}{6}, \quad \alpha_1 \approx \pi^2/6, \quad \beta_1 \approx 23\pi^2/24^2, \quad H \ll T ,$$
(11)

$$\beta_0 = \frac{1}{6}, \quad \alpha_1 \approx \pi^2/4, \quad \beta_1 \approx 27\pi^2/24^2, \quad T \ll H .$$

Several remarks arise at that point:

(i) As recalled above and detailed in Ref. 4, the fluctuations do not affect the T = 0 behaviors: $\beta(T=0) \equiv \beta_0$ is the value computed in the Stoner-Wohlfarth theory⁵; $\alpha(T=0) \equiv \alpha_0$ is assumed to have been incorporated in the definition of \overline{I} in (9) and is extracted from experiments. In contrast, the T dependences of α and β do diverge with S: The fluctuations greatly enhance the finite temperature dependence of M. For comparsion, the analog of (10) in the absence of fluctuations, as given in Ref. 5, would contain ST^2/T_F^2 instead of S^2T^2/T_F^2 .

(ii) $(M/H)_{H=0}$ must be compared to the result $\chi(T,0)$ in BMMF. The $H \ll T$ limit of α in (11) is $\pi^2/6 = 4\pi^2/24$ to be compared with the result in BMMF, $3.2 \pi^2/24 = [4 - (8/\pi^2)]\pi^2/24$. The extra term $8/\pi^2$ arose from less-important contributions in the integrals that we have neglected here; nevertheless the coefficient in BMMF, $3.2\pi^2/24 \approx 1.3$, was more accurate than (although very close to) the one we have here, $4\pi^2/24 \approx 1.6$.

(iii) The first term in (9) appeared with such a form in the theory of Moriya and Kawabata.⁷ As already pointed out elsewhere,⁸ their formalism is identical to the one derived earlier in BMMF that we use here.

4468

27

(iv) As emphasized in Refs. 3, 8, and 9, one notes the scaling in ST/T_F and the one in $S^{3/2}H/T_F$

 $=\sqrt{S}H/T_{sf}$ which is in agreement with the relation, at the ferromagnetic transition $(T = 0, \overline{I} = 1)$, $H \sim M^{3-3}$, where one replaces M by $\chi H \sim SH$. Thus the results of Ref. 6, which do not contain the same scaling, appear doubtful to us.

(v) From (10), it is clear that the deviations of M from linearity with respect to H are less pronounced when T increases. To observe the nonlinear

behavior, one would need higher fields at high temperatures than at low T. For a different band structure, when α_1 , β_0 , and β_1 would be all negative instead of positive here, then M(T,H)/H would increase with H (instead of decreasing here), but less and less so when the temperature increases; this is in qualitative agreement with what is found in TiBE₂.¹⁰ (vi) From the Maxwell relation^{1,3} $\partial M/\partial T$

 $= \partial S/\partial H$, we deduce $\partial^2 M/\partial T^2 = \partial \gamma/\partial H$, which thus yields with (10) (for a parabolic band)

$$\gamma(H) - \gamma(0) = -\chi_{\text{Pauli}}\alpha_1 \left(\frac{S^{3/2}H}{T_F}\right)^2 \left[1 - \left(\frac{\beta_1}{2\alpha_1} + 2\beta_0\right) \left(\frac{S^{3/2}H}{T_F}\right)^2 \cdots\right]$$
(12)

As pointed out in Ref. 3, the lowest field dependence of $\gamma(H) - \gamma(0)$, i.e., the first term in (12), follows from the curvature of $\chi(T, H = 0)$ at T = 0. However, at higher fields, or strong enough values of $[\beta_1/(2\alpha_1) + 2\beta_0]$, the H⁴ term in (12) being of opposite sign, may counterbalance the first one. This seems to be the case in TiBe₂ (Ref. 11) where, at $H = 7 \text{ T}, \gamma(H) \leq \gamma(0), \text{ while } \chi(T, H = 0) \text{ first in-}$ creases with T; however, it was pointed out to us^{12} that at 7 T, $\chi(T,H)$ decreases rather than increases when T increases. Therefore the predicted values given in Ref. 3 for the relative variation of $\gamma(H)$ at 7 T was erroneous since it corresponded to the extrapolation of the first term in (12) to a field value (7) T) where $\chi(T,H) - \chi(0,H)$ does not have the same sign that $\chi(T,H\simeq 0) - \chi(0,H\simeq 0)$; in that case, the 2nd term in (12), but also higher-order ones (not computed here), ought to be checked. In contrast, for Pd, where M remains linear in H up to about 35 T (see Ref. 39 in Ref. 3), the H^2 term in (10) and the H^4 term in $\gamma(H)$ must be very tiny and, from $\chi(T, 0)$ increasing with T, one expected³ a small increase of $\gamma(H)$ with H; some experiments (Ref. 34 of Ref. 3) found a strong decrease while others¹³ did observe recently a small increase.

(vii) A discussion was given elsewhere⁸ concerning the T variation of the nuclear relaxation rate T_1^{-1} in strongly enhanced paramagnets, recovering the proportionality to $T\chi(T)$ already proposed by Moriya and Ueda¹⁴ but also pointing out a scaling in T/T_{sf} below and above T_{sf} . If $\chi(T,H)$ replaces $\chi(T,0)$ in the finite field dependence of T_1^{-1} , then one can write at low $T \ll T_{sf}$ and $H \ll T_{sf}/\sqrt{S}$, T_1^{-1} $\propto T[M(T,H)/H]$, with M(T,H) given by (10). This is qualitatively analogous to the result derived in Ref. 15 but with different numerical coefficients. We note however that Ref. 15 used a supposedly general expansion for small \bar{q} and $\bar{\omega}/\bar{q}$ for the transverse susceptibility χ^{0-+} , which appears incomplete to us; such terms as h/\bar{q}^2 [that we have in (5)] are missing, for instance. The respective ratios of the four small quantities \bar{q} , $\bar{\omega}/\bar{q}$, \bar{T} , and h must be handled with special care since the domain of integration in the $(\bar{\omega}, \bar{q})$ plan is very tricky as can be seen from (6).

Our main result (10), with positive coefficients (11) can apply to two physical cases: liquid ³He, where there is no band structure and (11) applies as such, and UA1₂ which behaves the same way (Ref. 7 of Ref. 3), but whose band structure has not been calculated so far. Recent experiments on this last material¹⁶ exhibit a variation of M vs H in qualitative agreement with our result: M deviates from linearity with H around 15 T at low T but remains linear at much higher fields at high T, for which we expect that one would need higher fields to observe nonlinearity. A quantitative comparison is difficult though, since our result only applies for $T \ll T_{sf}$, and for $H \ll T_{\rm sf}/(S\beta_0)^{1/2}$, conditions which are not all fulfilled in the experiments. One remark: for this compound where $S \sim 4$, $(S\beta_0)^{1/2} \approx 1$ so that the characteristic field equal to $T_{\rm sf}/(S\beta_0)^{1/2}$ is practically equal to T_{sf} ; but for higher values of S, for instance in TiBe₂ ($S \sim 65$), the two values will be different. For liquid ³He, we expect deviations from linearity for M vs H to be of order 2% at 100 kG at 15 mK, at melting pressure where $S \sim 20$; but in confined geometries, if S can be as large as 60,¹⁷ then the corresponding effects are expected to be larger, $\sim 25\%$. The characteristic field where nonlinearity is expected to occur is estimated to be about 90 T at melting pressure for the bulk, but may be reached below ~ 18 T in confined geometry if an S value of ~ 60 can be achieved.

ACKNOWLEDGMENTS

We thank P. Nozières for having suggested that calculation to us, in connection with further polarized ³He studies, and we acknowledge useful discussions with K. Levin and O. Valls.

- *Permanent address: Université Louis Pasteur, 4 rue Blaise Pascal, F-6700 Strasbourg, France.
- ¹M. T. Béal-Monod, S. K. Ma, and D. R. Fredkin, Phys. Rev. Lett. <u>20</u>, 929 (1968); H. Ramm *et al.*, J. Low Temp. Phys. <u>2</u>, 539 (1970).
- ²M. T. Béal-Monod and J. Lawrence, Phys. Rev. B <u>21</u>, 5400 (1980).
- ³See a review in M. T. Béal-Mond, Physica (Utrecht), <u>109</u> and <u>110B</u>, 1837 (1982).
- ⁴M. T. Béal-Monod and K. Maki, Phys. Rev. Lett. <u>34</u>, 1461 (1975); J. Hertz, Phys. Rev. B <u>14</u>, 1165 (1976).
- ⁵See E. P. Wohlfarth and P. Rhodes, Philos. Mag. <u>7</u>, 1817 (1962); M. Shimizu, Rep. Prog. Phys. <u>44</u>, 329 (1981).
- ⁶P. Hertel, J. Appel, and D. Fay, Phys. Rev. B <u>22</u>, 534 (1980).
- ⁷T. Moriya and A. Kawabata, J. Phys. Soc. Jpn. <u>34</u>, 639 (1973).

- ⁸M. T. Béal-Monod (unpublished).
- ⁹J. Lawrence and M. T. Béal-Monod, in Valence Fluctuations in Solids, edited by L. Falicov, W. Hanke, and M. Maple (North-Holland, Amsterdam, 1981), p. 53.
- ¹⁰See Ref. 10 in Ref. 3 above.
- ¹¹G. R. Stewart, J. L. Smith, and B. L. Brandt, Phys. Rev. B <u>25</u>, 5907 (1982).
- ¹²F. Acker (private communication).
- ¹³J. J. M. Franse (private communication).
- ¹⁴T. Moriya and K. Ueda, Solid State Commun. <u>15</u>, 169 (1974).
- ¹⁵K. Ueda, Solid State Commun. <u>19</u>, 965 (1976).
- ¹⁶J. J. M. Franse et al., Phys. Rev. Lett. <u>48</u>, 1749 (1982).
- ¹⁷For instance, the case described, at high pressures, in B. Herbal *et al.*, Phys. Rev. Lett. <u>46</u>, 42 (1981).