Spin dynamics for the one-dimensional XY model at infinite temperature^{*†}

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The infinite-temperature space- and time-dependent spin-correlation functions $g_r^x(t)$ are studied for the onedimensional XY model. Numerical calculations are performed to obtain the exact autocorrelation function $g_0^x(t)$ for chains containing 5, 7, and 9 spins (S = 1/2). This yields exact results for the first 16 moments of the frequency autocorrelation function of the infinite chain, and estimates for a few of the higher moments as well. The analysis suggests that $g_0^x(t)$ for the infinite chain is identical to $\exp(-J^2t^2)/4$. We show that $g_r^x(t)$ for $r \neq 0$ vanishes identically for all values of time, implying a wave-vector-independent relaxation shape function. Our result for $g_0^x(t)$ is compared with that obtained by Huber for the classical ($S = \infty$) chain.

I. INTRODUCTION

The one-dimensional XY model has attracted considerable attention as a nontrivial many-body problem having interesting physical properties.¹ The Hamiltonian is described by

$$\mathcal{G}_{\gamma} = 2J \sum_{i} \left[(1+\gamma) S_{i}^{x} S_{i+1}^{x} + (1-\gamma) S_{i}^{y} S_{i+1}^{y} \right], \qquad (1)$$

where J is the strength of the interaction, \tilde{S}_i is the spin vector of the *i*th particle $(S = \frac{1}{2})$, and γ is the anisotropy parameter. Lieb, Schultz, and Mattis² showed that \mathcal{K}_{r} can be transformed into one describing a set of noninteracting fermions and made a detailed study of its ground-state (T=0) properties. Katsura³ introduced a magnetic field in the z direction and obtained exact results for the temperature and magnetic field dependence of the various thermodynamic properties. McCoy⁴ and later Barouch and McCoy⁵ made a detailed investigation of the various space-dependent spin-correlation functions at different temperatures. Suzuki⁶ studied the effect of introducing a staggered magnetic field along the z direction. He later introduced a generalized XY model and examined its relationship to other exactly soluble magnetic models.⁷

Recently, the dynamic behavior of this system has also received a great deal of attention. Barouch, McCoy, and Dresden⁸ introduced a timedependent external magnetic field in the z direction and solved the Liouville equation exactly. They demonstrated that for $\gamma \neq 0$ the magnetization shows nonergodic behavior. This rather peculiar behavior was also observed by Mazur, ⁹ Suzuki, ¹⁰ and Girardeau.¹¹ The space- and time-dependent spin pair-correlation functions $\tilde{g}'_{r}(t)$, defined by

$$\tilde{g}_{r}^{\nu}(t) = \langle S_{0}^{\nu}(t) S_{r}^{\nu} \rangle, \qquad (2)$$

are of considerable interest also. Niemeijer¹² and also Katsura, Horiguchi, and Suzuki¹³ obtained exactly the longitudinal correlation functions $\tilde{g}_r^{e}(t)$. Because of mathematical complexity, however, the

knowledge of transverse correlation functions $\tilde{g}_{r}^{x}(t)$ is rather limited. Barouch and McCoy¹⁴ examined the asymptotic behavior of $\tilde{g}_{r}^{x}(t)$ at T = 0 as $r \to \infty$ and $t \to \infty$. They found it to be an oscillatory factor times t^{-p} , where p is a fraction depending on the values of the parameters of the Hamiltonian. Very recently, Capel, Van Dongen, and Siskens¹⁵ obtained analytically a few of the high-temperature and short-time series expansion coefficients of $\tilde{g}_{r}^{*}(t)$ for $r \leq 5$.

Here we focus our attention on the isotropic XYmodel where the Hamiltonian \mathcal{K}_0 is given by Eq. (1) with $\gamma = 0$. We also confine ourselves to the infinite-temperature correlation functions

$$g_{\mathbf{r}}^{\nu}(t) = \operatorname{Tr}(e^{i\mathcal{R}_{0}t}S_{0}^{\nu}e^{-i\mathcal{R}_{0}t}S_{\mathbf{r}}^{\nu})/\operatorname{Tr}(1) = \overline{g}_{\mathbf{r}}^{\nu}(t)|_{\mathbf{T}=\infty} . (3)$$

and also its frequency Fourier transform $f_r^{\nu}(\omega)$, defined by

$$f_{r}^{\nu}(\omega) = \frac{1}{2\pi} \int_{-\infty}^{+\infty} e^{-i\omega t} g_{r}^{\nu}(t) dt, \quad \nu = x, y, z.$$
 (4)

The trace in Eq. (3) can be taken over any complete set of states.

We present here a study of the autocorrelation functions $g_0^{\epsilon}(t)$ and $g_0^{\star}(t)$ for the short-time region $(Jt \sim 3)$ for the infinite linear chain on the basis of *ab initio* calculation for finite linear chains containing N spins for N=5,7, and 9. Our result for $g_0^{\epsilon}(t)$ is in excellent agreement with the exact result and confirms the accuracy of our method. The first 16 moments of $f_0^{\star}(\omega)$ for the infinite chain are found to be identical to those of a Gaussian. The transverse crosscorrelation functions $g_{\star}^{\star}(t)$ for $r \neq 0$ are shown to vanish identically for all values of t. This, however, is not true for the general anisotropic Heisenberg model.

The results are presented in Sec. II and some concluding remarks are added in Sec. III.

II. RESULTS

The method of obtaining $g_0^{\nu}(t)$ ($\nu = x$ or z) for a finite chain containing N spins is identical to that

12

3845

of a recent study of linear Heisenberg chains.¹⁶ We first introduce the periodic boundary condition.

$$\vec{\mathbf{S}}_{i+N} = \vec{\mathbf{S}}_i \tag{5}$$

for each *i*. This enables one to diagonalize \mathcal{K}_0 numerically in a representation in which the translation operator *T* and the *z* component of the total spin S^x are diagonal. The trace in Eq. (3) is then evaluated over the eigenfunctions of \mathcal{K}_0 . The moments $(\mu_r^x)_{2k}$ of $f_r^x(\omega)$, which are given by the Taylor series expansion of $g_r^x(t)$ around t=0, namely.

$$g_{r}^{x}(t) = (\mu_{r}^{x})_{0} + \sum_{k=1}^{\infty} (-1)^{k} (\mu_{r}^{x})_{2k} \frac{t^{2k}}{(2k)!} ,$$

$$(\mu_{r}^{x})_{0} = \frac{1}{4} \delta_{0,r} ,$$
(6)

can also be evaluated in this representation. We shall on occasion use the normalized correlation functions $G_0^{\nu}(t)$, defined by

$$G_0^{\nu}(t) = g_0^{\nu}(t) / g_0^{\nu}(0) = 4g_0^{\nu}(t) . \tag{7}$$

To illustrate the accuracy of our method, we have plotted in Fig. 1 our numerical results for $G_0^{\epsilon}(t)$ for N=5,7, and 9, and also the analytic result for the infinite chain obtained by Niemeijer, ¹² namely,

$$G_0^{z}(t) = [J_0(2Jt)]^2, \tag{8}$$

where J_0 is the Bessel function of order **0**. As has been observed previously,¹⁶ addition of spins to a chain modifies $G_0^{e}(t)$ at successively longer times, keeping the short-time behavior relatively unchanged (see Fig. 1). We define τ_N to be the time domain up to which the N-spin result for $G_0^{e}(t)$ reproduces that of the infinite chain accurately. Figure 1 indicates that $J\tau_5 \sim 1$, $J\tau_7 \sim 2$, and $J\tau_9 \sim 3$. The result for $G_0^{e}(t)$ for each value of N investigated is



FIG. 1. Exact normalized infinite-temperature longitudinal autocorrelation function $G_0^{\varepsilon}(t)$ for the nearestneighbor one-dimensional XY model for different values of N. The solid curve is the analytical result of Niemeijer (Ref. 12) for the infinite chain.

in excellent agreement (within the accuracy of our computations¹⁷) with the exact result for $N = \infty$, if $t < \tau_N$.

The moments of $f_0^x(\omega)$ are shown in Table I. The *N*-spin result is expected to reproduce the moments of the infinite chain up through $(\mu_0^x)_{2N-2}$ exactly.¹⁶ Empirically (see Table I), it also seems to provide lower bounds for a few of the higher moments as well. A comparison of our results for N=5 and N=7 shows that the moments $(\mu_0^x)_2$ through $(\mu_0^x)_8$ are identical in these two cases. One also notes that $(\mu_0^x)_{10}$ and $(\mu_0^x)_{12}$ for N=5 are within 1.2% and 5%, respectively, of those obtained for N=7. Similarly, $(\mu_0^x)_{14}$ and $(\mu_0^x)_{16}$ for N=7 are within 0.03% and 0.2%, respectively, of those obtained for N=9. We believe that this trend continues for higher values of N as well and probably arises because of the

TABLE I. Moments of the infinite-temperature transverse autocorrelation function $f_{\delta}^{*}(\omega)$ for the nearest-neighbor one-dimensional XY model. Entries up through those marked (a) are identical to those of the infinite chain for each finite value of N investigated. In the fifth column $(N=\infty)$ entries up through $(\mu_{\delta}^{*})_{1\delta}$ are identical to those for N=9. The $(\mu_{\delta}^{*})_{1\delta}$, marked (b) in the table, and the higher entries in the same column should be considered as lower bounds of the exact results only.

			$\mu_0^{\mathbf{x}} 2k/J^{2k}$		
2k	N=5	N=7	N=9	<i>N</i> = ∞	$\frac{1}{4}(e-J^2t^2)$
0	0.25	0.25	0.25	0.25	0.25
2	0.500000	0.500000	0.500000	0.500000	0.50
4	3.00000	3.00000	3.00000	3.00000	3.00
6	30.0000	30.0000	30,0000	30.0000	30.00
8	420.000 ^a	420.000	420.000	420.000	420.00
10	7469.00	7560.00	7560.00	7560.00	7560.00
12	$0.157878 imes 10^6$	0.166320×10^{6} a	$0.166320 imes 10^6$	166320.0	166320.00
14	$0.378744 imes 10^7$	$0.432307 imes 10^7$	$0.432432 imes 10^7$	$0.432432 imes 10^7$	4324320.00
16	$0.998628 imes 10^8$	$0.129498 imes 10^9$	$0.129730 imes 10^{9}$ a	$0.129730 imes 10^{9}$	0.12972960×10^{9}
18	0.282815×10^{10}	$0.438355 imes 10^{10}$	0.441079×10^{10}	$0.44108 imes 10^{10}$ b	$0.44108064 imes 10^{10}$
20	$0.845415 imes 10^{11}$	$0.164966 imes 10^{12}$	0.167605×10^{12}	$0.1676 imes 10^{12}$	$0.16761064 imes 10^{12}$
22			$0.703861 imes 10^{13}$	$0.704 imes 10^{13}$	$0.70396470 imes 10^{13}$
24			0.323669×10^{15}	0.32×10^{15}	$0.32382376 \times 10^{15}$



FIG. 2. Test for Gaussian hypothesis for the normalized infinite-temperature transverse autocorrelation function $G_0^x(t)$ for the nearest-neighbor one-dimensional XY model.

lower dimensionality of our system. In column 5 of Table I, we provide exact results for the first 16 moments and also estimate lower bounds for $(\mu_0^x)_{18}$ through $(\mu_0^x)_{24}$ for the infinite chain. Our results for $(\mu_0^x)_2$ and $(\mu_0^x)_4$ are in excellent agreement with those obtained by McFadden and Tahir-Kheli.¹⁸ The $(\mu_0^x)_6$ and higher moments have not been calculated previously. The last column in Table I contains the moments of a Gaussian f(t), defined by

$$f(t) = \frac{1}{4} e^{-J^2 t^2} . (9)$$

A comparison of the last two columns in Table I shows that each of the known moments for $N = \infty$ is identical to that of f(t). This suggests that $g_0^x(t)$ for the infinite chain may be identical to f(t). We know of no rigorous proof of this result.

To investigate how $g_{0}^{*}(t)$ for finite chains approaches the Gaussian behavior, we have plotted h(t), defined by

$$h(t) = -\left[\log G_0^*(t)\right] / J^2 t^2, \tag{10}$$

against Jt for N=5,7, and 9, as shown in Fig. 2, which indicates $J\tau_5 \sim 1$, and $J\tau_7 \sim 2$. Assuming the same trend to continue for N=9, one obtains $J\tau_9 \sim 3$. We note that for t within $\tau_N h(t)$ is identical to 1 (within the accuracy of our calculations). Our numerical result on $G_0^*(t)$ for N=9 shows that it has already decayed to 10^{-4} when $Jt \sim 3$. We therefore hypothesize that $G_0^*(t)$ for the infinite chain is given by

$$G_0^{\mathbf{x}}(t) = e^{-J^2 t^2}.$$
 (11)

Recently, Huber¹⁹ has obtained numerical results on $g_0^x(t)$ for a classical $(S = \infty)$ chain by Monte Carlo methods. To compare our result with his, we scale his result according to the transformation

$$J = J_H / \sqrt{3} \quad (12)$$

where J_H is the exchange constant, as defined in Ref. 19. This ensures that the two models have the same second moment. We also properly nor-

malize his result. A comparison of the two cases is illustrated in Fig. 3. The asymptotic behavior for large values of t is quite different for the two cases, in contrast to the Heisenberg chain, where it was found to be rather insensitive to the S value.¹⁶

At infinite temperatures, the transverse off-site (or crosscorrelation) functions $g_r^{x}(t)$ for $r \neq 0$ vanish identically. This can be seen term by term in, say, the diagram expansion discussed by Wortis, ²⁰ where classes of diagrams vanish either because $J_{\mu} = J_{g} = 0$ or because $\beta = 1/k_B T = 0$. The correlations for r an odd integer can be seen to vanish by the usual symmetry arguments.²¹ These results are of course reproduced explicitly in the calculations on the finite chains.

III. CONCLUDING REMARKS

Since $g_r^*(t)$ vanishes for $r \neq 0$, the wave-vectordependent transverse relaxation shape function²² $F_a^*(t)$ for our model is given by

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$$F_{q}^{x}(t) = e^{-J^{2}t^{2}},$$
(13)

independent of q. In the theory of magnetic resonance, the q = 0 mode of $F_q^x(t)$ is referred to as the free-induction decay function.²³ Its frequency Fourier transform is identical to the NMR line shape.²⁴ We therefore conclude that the NMR line shape for the one-dimensional XY model is a Gaussian, provided that the magnetic field is applied along the z direction. The usual phenomenological arguments²⁵ which predict a Gaussian line shape, however, cannot be valid for this system. A simple physical interpretation and a rigorous proof of our result is still lacking.



FIG. 3. Comparison of normalized infinite-temperature transverse autocorrelation function for the nearestneighbor one-dimensional XY model for different values of S. The solid curve is our present result for $S = \frac{1}{2}$. The dashed curve is the numerical result obtained by Huber (Ref. 19) for $S = \infty$.

Our arguments that the cross-correlation functions vanish can be generalized to other loosepacked lattices. In particular, this shows that $F_q^{x}(t)$ for the square and simple-cubic lattices²⁶ are also independent of q. Results for the second (m_2) and fourth (m_4) moments of the frequency Fourier transform of $G_0^{x}(t)$ for hypercubic lattices¹⁸ show that

$$m_4/(m_2)^2 = (\mu_0^x)_4/4[(\mu_0^x)_2]^2 = 5 - 2/d$$
, (14)

where *d* is the dimensionality of the lattice. Since the ratio in Eq. (14) must be exactly three for a Gaussian function, $F_q^x(t)$ cannot be Gaussian for all times when d > 1.

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Note added in proof. The question as to what happens when $\gamma \neq 0$ in Eq. (1) has been raised (B. M. McCoy, private communication). One can show (see Ref. 15) that for this case $m_4/(m_2)^2 3[1+\gamma(\gamma -\frac{2}{3})]$. This shows that for a general nonzero value of γ , other than $\gamma = \frac{2}{3}$, G_0^x cannot be a Gaussian. It remains to be seen if the higher moments for $\gamma = \frac{2}{3}$ are also equal to those of a Gaussian.

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