Nonlocal free-energy density functional for a broad range of warm dense matter simulations

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Finite-temperature orbital-free density functional theory (FT-OFDFT) holds significant promise for simulating warm dense matter due to its favorable scaling with both system size and temperature. However, the lack of the numerically accurate and transferable noninteracting free-energy functionals results in a limit on the application of FT-OFDFT for warm dense matter simulations. Here, a nonlocal free-energy functional, named XWMF, was derived by line integrals for FT-OFDFT simulations. Particularly, a designed integral path, wherein the electronic density varies from uniform to inhomogeneous, was employed to accurately describe deviations in response behavior from the uniform electron gas. The XWMF has been benchmarked by a range of warm dense matter systems, including the Si, Al, H, He, and H-He mixture. The simulated results demonstrate that FT-OFDFT within XWMF achieves remarkable performance for accuracy and numerical stability. It is worth noting that XWMF exhibits a low computational cost for large-scale *ab initio* simulations, offering exciting opportunities for the realistic simulations of warm dense matter systems covering a broad range of temperatures and pressures.

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I. INTRODUCTION

As a bridge between cold condensed matter and hot plasma, warm dense matter (WDM) represents an extreme state comprising a matter regime characterized by high temperatures (\sim 10 eV) and high pressures (\sim 1 Mbar or higher). The WDM has attracted tremendous attention in various fields, including planet science, high-energy-density physics, and materials science, due to its ubiquity throughout nature, such as exoplanet interiors, the path to inertial confinement fusion, and neutron stellar atmospheres [1,2]. In the WDM regime, electrons are usually in a highly excited state and partially degenerate, exhibiting non-negligible quantum effects. Therefore, *ab initio*–based simulations have become an indispensable approach for describing the phenomena and mechanisms of WDM.

Currently, several approaches, including finite-temperature Kohn-Sham density functional theory (FT-KSDFT) [3–10], path-integral Monte Carlo (PIMC) [11–14], extended first-principles molecular dynamics (ext-FPMD) [15,16], and stochastic density functional theory (sDFT) [17–20], have been established and yielded successes to simulate WDM. Particularly, a combination of FT-KSDFT and molecular dynamics (MD) has been proposed as an ideal framework for simulations of WDM, where electrons and ions are treated by quantum mechanical KSDFT and classical MD, respectively.

However, the heavy computational demands for FT-KSDFT simulations at high temperature make its use for WDM problematic [19]. Even with modern supercomputers, FT-KSDFT can only be used to treat the WDM at relatively low temperature without further tricks.

The finite-temperature orbital-free density functional theory (FT-OFDFT) [21–26] combined with MD offers a promising alternative to simulate the WDM, because its computational cost grows linearly with system size almost irrespective of temperatures. It is necessary to note that the performance of FT-OFDFT critically depends on the noninteracting free-energy density functional (FEDF). In this context, the development of accurate and reliable FEDF is highly desirable.

The early FEDFs based on finite-temperature Thomas-Fermi (TF) theory [27–31] were derived and successfully applied to warm dense plasma at ultrahigh temperatures [32–34]. Later, several semilocal FEDFs, such as VT84F [35] and Luo-Karasiev-Trickey FEDF (LKTF) [36], were derived from the generalized-gradient approximation (GGA) [23]. These GGA-FEDFs show significant accuracy improvement in WDM simulations compared to TF-FEDFs [37-39], while the GGA functionals show a lack of nonlocality and cannot reproduce the exact Lindhard response [25,40-42]. A recent nonlocal [43-46] FEDF, denoted as WTF [41], has been constructed by a direct extension of the Wang-Teter kinetic energy density functional [44] and shown high accuracy across a broad range of temperatures. However, the WTF was found to be numerically problematic under specific conditions [47]. To address this issue, a modified functional, WTF β vW [47], was later proposed and exhibited remarkable performance in WDM simulations. Despite the advances in the FEDF within

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FT-OFDFTs over past decades, there is huge room for future development, and any progress could impact the FT-OFDFT application in WDM simulations.

In this work we derived an advanced nonlocal FEDF using line integrals, following the theoretical formalism of the Xu-Wang-Ma (XWM) kinetic energy density functional [48]. Numerical assessments reveal that XWM-FEDF (XWMF) exhibits numerical stability and high accuracy, as demonstrated by fact that results calculated by XWMF were in good agreement with FT-KSDFT results.

The reminder of this manuscript is organized as follows: Sec. II offers the definition and construction of the XWMF functional. The computational details are provided in Sec. III. The calculated results and discussion are presented in Sec. IV, followed by conclusions in Sec. V.

II. METHODS

In contrast to the simulation at zero temperature, the grand-canonical potential instead of energy [5,6] is the critical physical quantity of interest at finite temperature. The grand potential (Ω) of elections can be expressed as a density functional:

$$\Omega[\rho;T] = F[\rho;T] + \int d^3 \mathbf{r} v_{\text{ext}}(\mathbf{r})\rho(\mathbf{r}) - \mu N_e, \quad (1)$$

where ρ , *T*, v_{ext} , μ , and N_e are the electron density, absolute temperature, external potential, chemical potential, and number of electrons, respectively. The universal functional $F[\rho]$ is an unknown term but could be decomposed into three terms [6]:

$$F[\rho; T] = F_s[\rho; T] + F_H[\rho] + F_{\rm XC}[\rho; T], \qquad (2)$$

where F_s , F_H , and $F_{\rm XC}$ denote the noninteracting free energy, the classical Coulomb energy (Hartree energy), and exchange-correlation energy, respectively. Generally, F_H explicitly depends on the electron density and can be calculated by $F_H[\rho] = \int \int \frac{\rho(\mathbf{r})\rho(\mathbf{r}')}{|\mathbf{r}-\mathbf{r}'|} d^3\mathbf{r} d^3\mathbf{r}'$, while the approximate functional forms should be required to express the $F_s[\rho; T]$ and $F_{\rm XC}[\rho; T]$. Several approximate forms of $F_{\rm XC}$, including the local-density approximation [49], GGA [50], hybrids [51], and meta-GGA [52] functionals, have been proposed for finite-temperature simulations. Moreover, the adiabatic approximation ($F_{\rm XC}[\rho; T] \approx E_{\rm XC}[\rho]$) has also been widely used in such simulations [39,53,54]. These exchange-correlation functionals usually yield reasonable results since the $F_{\rm XC}$ is an order of magnitude smaller than F [23].

Since the F_s has the same order of magnitude as the functional F, it plays a crucial role in determining accuracy of FT-OFDFT simulations [23]. Generally, the nonlocal FEDF can be written as [41,44,47]

$$F_{s}[\rho;T] = F_{s}^{\text{TF}}[\rho;T] + F_{s}^{\text{vW}}[\rho;T] + F_{s}^{\text{NL}}[\rho;T], \quad (3)$$

where F_s^{TF} is the TF-FEDF, which can be derived from the finite-temperature TF theory:

$$F_s^{\rm TF}[\rho;T] = \int f_s^{\rm TF}(\mathbf{r};T) d^3\mathbf{r}, \qquad (4)$$

where $f_s^{\text{TF}}(\mathbf{r}; T) = t_0^{\text{TF}}(\mathbf{r})k(\tau, \mathbf{r})$ is the TF free-energy density. $t_0^{\text{TF}} = \frac{3}{10}(3\pi^2)^{2/3}\rho^{5/3}(\mathbf{r})$ is the TF kinetic energy density at 0 K, and the factor $k(\tau, \mathbf{r})$ has an analytical form [31,55] at finite temperature. $\tau = \frac{2k_BT}{[3\pi^2\rho]^{2/3}}$ and k_B denote the reduced temperature and Boltzmann constant, respectively.

The second term in Eq. (3) is the von Weizsäcker (vW) functional [56]. At finite temperature, three vW-functional forms have been proposed: (i) the adiabatic approximation form adopted as vW kinetic energy [56] $F_s^{vW}[\rho;T] = T_s^{vW}[\rho] \equiv \int \frac{|\nabla \rho(\mathbf{r})|^2}{8\rho(\mathbf{r})} d^3\mathbf{r}$; (ii) the vW functional scaled by a reduced temperature function derived from finite-temperature gradient correction [23,30,31]; and (iii) βvW form employed in WTF βvW [47]. Note that the adiabatic vW functional $(F_s^{vW} = T_s^{vW})$ is adopted in this work because of its correct asymptotic behavior in the linear response of the uniform electron gas [41].

The third term in Eq. (3) is the nonlocal part free energy (NLFE). Following the formalism of the XWM kinetic energy density functional [48], NLFE can be derived from the line integral:

$$F_{s}^{\mathrm{NL}}[\rho;T] = F_{s}^{\mathrm{NL}}[\rho_{t=0};T] + \int d^{3}\mathbf{r} \int_{0}^{1} dt \frac{\delta F_{s}^{\mathrm{NL}}[\rho;T]}{\delta \rho_{t}(\mathbf{r})} \frac{d\rho_{t}(\mathbf{r})}{dt}.$$
 (5)

Note that the first-order functional derivative in the integrand can also be evaluated from the line integral:

$$\frac{\delta F_s^{\text{NL}}[\rho;T]}{\delta \rho(\mathbf{r})} = \frac{\delta F_s^{\text{NL}}[\rho;T]}{\delta \rho(\mathbf{r})}|_{\rho_{t'=0}} + \int d^3 \mathbf{r}' \int_0^1 dt' \frac{\delta^2 F_s^{\text{NL}}[\rho;T]}{\delta \rho(\mathbf{r}) \delta \rho_{t'}(\mathbf{r}')} \frac{d\rho_{t'}(\mathbf{r})}{dt'}.$$
 (6)

In principle, one can derive an exact NLFE functional expression according to Eqs. (5) and (6) if the second derivative of NLFE, denoted as $(F_s^{\text{NL}})''(\mathbf{r}, \mathbf{r}') \equiv \frac{\delta^2 F_s^{\text{NL}}}{\delta \rho(\mathbf{r}) \delta \rho(\mathbf{r}')}$, is known along a given density path ρ_t . However, the exact condition is only available by the Lindhard theory [40,57] at uniform electron gas, where $\rho = \rho_0$. Taking this into consideration, the density integral path is selected by the following form [48,58]:

$$\rho_t(\mathbf{r}) = \rho_0 + t[\rho(\mathbf{r}) - \rho_0], \tag{7}$$

where $t \in [0, 1]$. The peculiar path in Eq. (7) has proven to be successfully applied in describing near-free electron gas systems [48]. Furthermore, the initial values in Eqs. (5) and (6), $F_s^{\text{NL}}[\rho_{t=0}]$ and $\frac{\delta F_s^{\text{NL}}}{\delta \rho(\mathbf{r})}|_{\rho_{t'=0}}$, naturally become zero for this density path. However, the general form of the second-order derivative of NLFE in Eq. (6) is unknown along this path, except for the uniform electron gas [40]:

$$\left(F_s^{\mathrm{NL}} \right)''[\rho_0; T](\mathbf{r}, \mathbf{r}') = -\chi_{\mathrm{NL}}^{-1}[\rho_0; T](\mathbf{r}, \mathbf{r}')$$

$$\equiv \rho_0^{-1/3} G\left(k_F^0 |\mathbf{r} - \mathbf{r}'|; T \right), \qquad (8)$$

where $\chi_{\rm NL}^{-1} = \chi_L^{-1} - \chi_{\rm TF}^{-1} - \chi_{\rm vW}^{-1}$. χ_L^{-1} , $\chi_{\rm TF}^{-1}$, and $\chi_{\rm vW}^{-1}$ represent the Lindhard-response, TF-response, and vW-response functions, respectively [21,40,57]. The *G*-function term of Eq. (8) is a dimensionless function and has an analytical form in momentum space, see Supplemental Material [59]. For a general inhomogeneous electron density, we extended Eq. (7) to a general form by introducing $\rho_0^{-1/3} \rightarrow \rho^{-1/6}(\mathbf{r})\rho^{-1/6}(\mathbf{r}')$



FIG. 1. Electronic pressure prediction comparison for various free-energy functionals compared with FT-KSDFT results for the static lattice of (a), (b) fcc H, (c), (d) fcc Al, and (e), (f) cd Si. The LKTF data are adopted from Ref. [36].

and
$$k_F^0 \to \varepsilon_{\gamma}(\mathbf{r}, \mathbf{r}') = \left[\frac{k_F^{\gamma}(\mathbf{r}) + k_F^{\gamma}(\mathbf{r}')}{2}\right]^{1/\gamma}$$
:
 $\left(F_s^{\mathrm{NL}}\right)''[\rho; T](\mathbf{r}, \mathbf{r}')$
 $= I^{\kappa}(\mathbf{r})\rho^{-\frac{1}{6}}(\mathbf{r})G(\varepsilon_{\gamma}|\mathbf{r} - \mathbf{r}'|; T)\rho^{-\frac{1}{6}}(\mathbf{r}')I^{\kappa}(\mathbf{r}'),$ (9)

where $I^{\kappa}(\mathbf{r}) = [\rho(\mathbf{r})/\rho_0]^{\kappa}$ and κ is an adjustable parameter. Note that Eq. (9) will naturally degenerate into Eq. (8) when the electron density is a uniform electron gas $[\rho_{t=0} = \rho_0]$. Therefore, the additional term $I^{\kappa}(\mathbf{r})I^{\kappa}(\mathbf{r}')$ in Eq. (9) is introduced to tune the second derivative of NLFE for inhomogeneous electron gas along the integral path ($t \neq 0$). To reduce the complexity, we simplify the density-dependent *G* function in Eq. (9) by a Taylor series expansion:

$$G(\varepsilon_{\gamma}|\mathbf{r} - \mathbf{r}'|;T) \simeq G(k_{F}^{*}|\mathbf{r} - \mathbf{r}'|;T) + \frac{\partial G(\varepsilon_{\gamma}|\mathbf{r} - \mathbf{r}'|;T)}{\partial \rho(\mathbf{r})} |_{\rho^{*}} \Delta \rho(\mathbf{r}) + \frac{\partial G(\varepsilon_{\gamma}|\mathbf{r} - \mathbf{r}'|;T)}{\partial \rho(\mathbf{r}')} |_{\rho^{*}} \Delta \rho(\mathbf{r}'), \quad (10)$$

where ρ^* is a reference charge density, and $\Delta \rho(\mathbf{r}) = \rho(\mathbf{r}) - \rho^*$. So far, the nonlocal part of the kinetic functional can be derived by combining Eqs. (5)–(10),

$$F_{s}^{\rm NL}[\rho;T] = F_{0}^{\rm NL}[\rho;T] + F_{1}^{\rm NL}[\rho;T], \qquad (11)$$

where the zeroth-order NLFE (F_0^{NL}) is given by

$$F_0^{\mathrm{NL}}[\rho;T] = \iint d^3 \mathbf{r} d^3 \mathbf{r}' \rho^{\kappa + \frac{5}{6}}(\mathbf{r}) \omega_0(\mathbf{r},\mathbf{r}') \rho^{\kappa + \frac{5}{6}}(\mathbf{r}'), \quad (12)$$

where $\omega_0 = \frac{18G(k_F^*|\mathbf{r}-\mathbf{r}'|;T)}{(6\kappa+5)^2\rho_0^{2\kappa}}$. Note that F_0^{NL} shows the same form as the nonlocal part of WTF when $\kappa = 0$ and $\rho^* = \rho_0$. The first-order term is given by

$$F_{1}^{\mathrm{NL}}[\rho;T] = \iint d^{3}\mathbf{r} d^{3}\mathbf{r}' \rho^{\kappa + \frac{11}{6}}(\mathbf{r})\omega_{11}(\mathbf{r},\mathbf{r}')\rho^{\kappa + \frac{5}{6}}(\mathbf{r}') + \iint d^{3}\mathbf{r} d^{3}\mathbf{r}' \rho^{\kappa + \frac{5}{6}}(\mathbf{r})\omega_{12}(\mathbf{r},\mathbf{r}')\rho^{\kappa + \frac{5}{6}}(\mathbf{r}'),$$
(13)

where $\omega_{11} = \frac{G'(\mathbf{r}, \mathbf{r}'; T)}{(\kappa + 5/6)(\kappa + 11/6)\rho_0^{2\kappa}}$, $\omega_{12} = -\frac{\rho^* G'(\mathbf{r}, \mathbf{r}'; T)}{(\kappa + 5/6)^2 \rho_0^{2\kappa}}$, and $G' \equiv \frac{\partial G(\varepsilon_{\gamma} | \mathbf{r} - \mathbf{r}' |; T)}{\partial \rho(\mathbf{r})} |_{\rho^*}$. In practice, a compact form of XWMF is used for implementation, with high computational efficiency, by combining Eq. (12) and the second line of Eq. (13) [60]:

$$F_{s}^{\mathrm{NL}}[\rho;T] = \iint d^{3}\mathbf{r} d^{3}\mathbf{r}' \rho^{\kappa+\frac{5}{6}}(\mathbf{r})\omega_{1}(\mathbf{r},\mathbf{r}')\rho^{\kappa+\frac{5}{6}}(\mathbf{r}') + \iint d^{3}\mathbf{r} d^{3}\mathbf{r}' \rho^{\kappa+\frac{11}{6}}(\mathbf{r})\omega_{2}(\mathbf{r},\mathbf{r}')\rho^{\kappa+\frac{5}{6}}(\mathbf{r}'),$$
(14)

where $\omega_1 = \omega_0 + \omega_{12}$, and $\omega_2 = \omega_{11}$. We should point out here that XWMF will naturally degenerate into an XWM kinetic energy density functional [48] at T = 0 K; a numerical comparison between XWMF and XWM is provided in the Supplemental Material [59]. The detailed derivation and implementation of XWMF are provided in the Supplemental Material [59].

III. COMPUTATIONAL DETAILS

The various FEDFs, including TF, WTF, WTF β vW, and XWMF, have been implemented in ATLAS [61,62]. To validate the accuracy of the XWMF functional, we conducted static lattice calculations and *ab initio* molecular dynamics (AIMD) using both FT-OFDFT and FT-KSDFT, employing identical pseudopotentials. For uniformity, we employed the adiabatic approximation using the adiabatic Perdew-Zunger local density approximation [63] as the exchange-correlation (XC) functional across all calculations. For XWMF, the parameters of $\rho^* = \rho_0$ and $\kappa = 0$ are adopted for all cases, and a brief discussion about κ is provided in the Fig. S1 of the Supplemental Material [59]. In all FT-KSDFT calculations, appropriate band numbers were chosen to encompass all bands with occupations exceeding 5×10^{-5} , as detailed in the Supplemental Material [59].

For the static lattice equation-of-state (EOS) calculations, the bulk-derived local pseudopotentials (BLPSs) [64,65] were adopted for fcc Al and cd Si; Heine-Abarenkov pseudopotentials (HAPPs) [23,66,67] were employed for fcc H. The grid spacing of 0.08, 0.08, and 0.05 Å for Al, Si, and H was employed in ATLAS to ensure convergence of the free energy to within 1 meV/atom. FT-KSDFT calculations were performed using CASTEP 8.0 [68], employing a kinetic energy cutoff of 1000 eV for Al and Si, and 4000 eV for H. The k-point meshes were generated via the Monkhorst-Pack method [69] with a spacing of 0.016 Å^{-1} . The AIMD calculations were performed to obtain the EOS of hydrogen (H), helium (He), and the H-He mixture. The HAPPs [23,66,67] with $r_{\text{cut}} = 0.25$ Bohr were adopted for H with a bulk density of less than 8.0 g/cm³. To simulate denser H and He, the Troullier-Martins norm-conserving nonlocal pseudopotentials (NLPPs) [70] were generated by FHI98PP [71] containing only s-channel with 0.10 and 0.25 Bohr cutoffs. In FT-OFDFT calculations, 0.05 and 0.03 Å grid spacings were employed for HAPPs and NLPPs. The FT-KSDFT-AIMD was performed with ABACUS 3.5.3 [72,73]. The kinetic energy cutoff and k-point meshes were varied from 200 to 300 Ry and Γ -only to $3 \times 3 \times 3$. The number of atoms was 108 for H less than 8.0 g/cm³ and 125 atoms for the denser case. The time step varied

from 0.005 to 0.1 fs. All AIMD were performed with the *NVT* ensemble controlled by Nose-Hoover thermostat [74,75]. Each system was run of 10 000 steps for initial equilibration; after that, pressures were averaged over the next 10 000 steps.

IV. RESULTS AND DISCUSSION

A. Static lattice equations of state

To evaluate the accuracy of XWMF, we initially applied it to obtain the static lattice equations of state of fcc H, fcc Al, and cd Si. Comparisons of the curves of pressure vs temperature and density obtained by FT-OFDFT using XWMF, TF, LKTF, and WTF β vW, along with the FT-KSDFT results, are shown in Fig. 1. Overall, the nonlocal functionals were superior to or comparable to the local (TF) and semilocal (LKTF) functionals for most of the considered systems. It can be clearly seen that the curves obtained by FT-KSDFT were well reproduced by the nonlocal functionals of WTF β vW and XWMF across the considered temperature and density ranges for Al and H with fcc structures, as presented in Figs. 1(a)-1(d). For cd Si, the pressure deviations evaluated by FT-OFDFT within XWMF with respect to those of FT-KSDFT were generally small for the considered temperatures and densities. This was in contrast to $WTF\beta vW$, which produces a large deviation at low temperatures ranging from 0.01 to 1 eV, as shown in Figs. 1(e) and 1(f). For example, the pressure deviation between WTF β vW and FT-KSDFT is 6.47 GPa at $\rho = 2.3$ g/cm³ and T = 1 eV, while XWMF has contributed to a significant reduction of this deviation to 2.05 GPa. These results illustrated that FT-OFDFT within XWMF can give credibility for simulations of WDM over a wide range of pressures and temperatures.

B. Ab initio molecular dynamics

The FT-OFDFT has been proved to yield a powerful predictive capability in determining the thermophysical properties, offering the same accuracy as the FT-KSDFT. Thus the XWMF makes FT-OFDFT to be sufficient in both accuracy and cost savings for the simulations of WDM, allowing us to investigate the properties of warm dense H and H-He

Temperature	Density	FT-KSDFT	$FT - KSDFT^{a}$	LKTF ^a	WTF	WTFβvW	XWMF
	0.6	2.1	2.1	1.9	-	-	1.9
	1.0	4.9	5.0	4.6	-	5.1	4.8
25 kK	2.0	16.7	16.9	16.3	-	16.9	16.7
	4.0	58.8	59.1	58.5	59.3	59.0	58.7
	8.0	205.7	207.2	205.8	206.2	205.8	205.7
50 kK	0.6	3.9	3.9	3.6	-	4.1	3.6
	1.0	7.9	8.0	7.5	-	8.1	7.6
	2.0	22.5	22.7	22.2	-	22.9	22.4
	4.0	70.1	70.6	69.9	71.1	70.5	70.1
	8.0	228.0	229.5	228.3	229.0	228.5	228.0
Mean absolute deviation		-	-	0.28	0.75	0.27	0.11

TABLE I. The evaluated pressure (in Mbar) of warm dense H with various densities (in g/cm³) via FT-KSDFT-MD and FT-OFDFT-MD simulations using WTF/WTF β vW/XWMF at T = 25 and 50 kK.^{*a*} The FT-KSDFT and LKTF results are adopted from Ref. [36].

mixtures under relevant planetary conditions. To illustrate the validity of XWMF, the equilibrium pressures of H at 25 and 50 kK were calculated by FT-KSDFT-MD and FT-OFDFT-MD using WTF, WTF β vW, and XWMF. The calculated results, along with those obtained by LKTF and FT-KSDFT in Ref. [36], were presented in Table I. Generally, the accuracy of XWMF implemented in FT-OFDFT-MD is comparable to or better than that of previous FEDFs with minimal mean absolute deviation. In particular, there was remarkable agreement between the XWMF and WTF β vW for the warm dense H at 2.0 g/cm³ and 25 kK. The estimated equilibrium pressures were 16.7 and 16.9 Mbar for XWMF and WTF β vW, respectively. The results were well consistent with that obtained by FT-KSDFT, but LKTF given a relatively large discrepancy with respect to FT-KSDFT (0.4 Mbar).

Generally, the pair distribution functions are of importance to represent the structure properties of matter, such as local bonding and composition. Therefore, we calculated the pair distribution functions for warm dense H with various densities at 25 and 50 kK. Just as shown in Fig. 2, our estimates of the pair distribution functions using FT-OFDFT-MD with all FEDFs were in perfect agreement with the estimates of FT-KSDFT-MD for the WDM of H at high densities. However, the stark differences between XWMF and other previous FEDFs were observed at lower densities. XWMF provides better predictions than WTF β vW for pair distribution functions at densities of 0.6–2.0 g/cm³. Our findings indicated that the XWMF has a dramatic improvement in numerical accuracy for predicting equilibrium pressures and pair distribution functions of WDM.

Besides accuracy, the numerical stability is an alternative important issue for the FEDFs. According to our results, FT-OFDFT-MD simulations using the WTF and WTF β vW FEDFs exhibit numerical instabilities. The situation was more dramatic for the cases of WDM at lower densities, while XWMF enabled us to effectively overcome this problem to a large extent, as evidenced by fact that the simulations for all the considered systems with various densities were successfully converged.

It is well known that H and He are ubiquitous in giant planets and numerous exoplanets [77]. As a consequence, deep investigations of the H-He phase behavior under relevant planetary conditions are highly desirable [10,76–86]. However, it remains challenging experimentally because the extremes of temperature and pressure are generally difficult to be currently accessible. Previously, FT-OFDFT has been applied to WDM simulations (such as H [36,38], D [36,47], Fe [32], SiO₂ [87], T-D mixture [88], and CHON resin material [37]) and has become routine in recent years. High-quality EOS (less than $\sim 1\%$) is required to provide an accurate model to answer some fundamental questions regarding the composition and formation of planets [77].

Here the H, He, and H-He mixtures with two typical densities in the solar interior, 10 and 160 g/cm³, were selected for investigating the matter states of the solar radiative zone and core. As shown in Figs. 3(a) and 3(b), the EOS curves calculated by FT-OFDFT using XWMF almost overlap with those calculated by FT-KSDFT. It should be stressed that the mean percentage error of pressure $(\overline{\Delta P^{OF}/P^{KS}} = \frac{1}{N} \sum \frac{|P^{OF} - P^{KS}|}{P^{KS}})$



FIG. 2. The pair distribution functions of H calculated by FT-KSDFT, and FT-OFDFT within WTF, WTF β vW, and XWMF at 25 kK (left) and 50 kK (right).

was $\simeq 1.4\%$. Due to the superior performance of XWMF in accuracy and generality, FT-OFDFT within XWMF holds great promise for offering a reliable prediction of properties of WDM formed by H and He.

C. Computational cost

To evaluate the computational efficiency of XWMF, we considered fcc Al supercells with sizes ranging from 108 to 13500 atoms and temperatures ranging from 1 to 10 eV as benchmarks. We compared the wall time for the static lattice



FIG. 3. Pressures vs temperatures for H, He, and H-He mixture with 0.4 mass abundance of H calculated by FT-OFDFT using XWMF and FT-KSDFT at (a) $\rho = 10 \text{ g/cm}^3$ and (b) $\rho = 160 \text{ g/cm}^3$. FT-KSDFT results are adopted from Ref. [76].

energy calculations using the FT-OFDFT code against that using the FT-KSDFT of the VASP 6.1.0 package [89,90]. Note that the projector augmented-wave (PAW) potentials [91] and gamma-point sampling were employed in FT-KSDFT calculations. All calculations were performed on a node with 2 Intel(R) Xeon(R) Gold 9242 CPUs (48 cores, 2.30 GHz) and 384 GB of RAM. At elevated temperatures, enough states should be required to accurately describe a thermal ensemble of electrons. Hence, the large computational cost of FT-KSDFT was a major obstacle to practical simulations of WDM. Just as shown in Figs. 4(a) and 4(b), the wall times of FT-KSDFT grew rapidly as both the number of atoms and temperatures increased, while the wall time of FT-OFDFT using XWMF was quasilinear scaling with system size [Fig. 4(a)] and it did not scale with temperatures [Fig. 4(b)]. Thus, the XWMF method can be used in FT-OFDFT simulations of the WDM at higher temperatures than that accessible in the FT-KSDFT. We also observed that the computational costs of XWMF are comparable to those of WTF β vW, which is slightly higher than the TF model, as shown in Fig. S2 of the Supplemental Material [59]. The low



FIG. 4. Wall times of FT-OFDFT using XWMF and FT-KSDFT for fcc Al supercell static lattice calculations. (a) Wall times as a function of the number of atoms varying from 108 to 13 500 at a temperature of 1 eV. (b) Wall times as a function of temperature for 108 atom supercells.

computational cost of XWMF, as well as its high accuracy, leads us to expect that XWMF is well suited for simulations of WDM across the wide temperature regime.

V. CONCLUSIONS

In summary, a nonlocal finite-temperature free-energy energy density functional of XWMF is derived via line integrals and has been systemically applied to several warm dense matters. The accuracy of XWMF is validated by correctly reproducing equations of state and pair distribution functions predicted by FT-KSDFT for various warm dense matters. Remarkably, our findings reveal that the performance of XWMF in accuracy and efficiency is comparable or superior to the existing free-energy density functionals (e.g., WTF, WTF β vW, and LKTF). Despite its enormous success for expanding the applicability of FT-OFDFT over a very wide temperature range, XWMF derived by uniform electron gas yields a considerable margin of error at lower temperature regimes accessible with a conventional FT-KSDFT (temperature less than 1 eV). It is expected that further improvement of the free-energy energy density functional by incorporation of more appropriate electron density response behaviors will be built into standard tools for investigation of properties of warm dense matter.

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