

Entanglement Hamiltonian of a nonrelativistic Fermi gas

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(Received 11 December 2023; revised 29 February 2024; accepted 6 May 2024; published 13 May 2024)

We study the entanglement Hamiltonian for a spherical domain in the ground state of a nonrelativistic free-fermion gas in arbitrary dimensions. Decomposed into a set of radial entanglement Hamiltonians, we show that the entanglement spectrum in each sector is identical to that of a hopping chain in a linear potential, with the angular momentum playing the role of the subsystem boundary. Furthermore, the eigenfunctions follow from a commuting differential operator that has exactly the form predicted by conformal field theory. Rescaled by the radial Fermi velocity, this operator gives a perfect approximation of the entanglement Hamiltonian, except for large angular momenta that belong to the edge regime in the analogous gradient chain. One thus finds that the conformal field theory result becomes asymptotically exact only in one dimension.

DOI: [10.1103/PhysRevB.109.L201113](https://doi.org/10.1103/PhysRevB.109.L201113)

Entanglement plays a key role in characterizing the distinct phases of quantum matter in ground states of many-body systems [1–4]. The intricate nature of quantum correlations is encoded in the reduced density matrix of a subsystem, or equivalently, written in an exponential form, in the entanglement Hamiltonian (EH) [5]. One of the most remarkable properties that has been uncovered in a broad range of many-body systems is the locality of the EH [5]. Its precise structure is, however, not only of theoretical interest, but also fundamental to novel techniques aiming at a more efficient spectroscopy and tomography of the reduced density matrix in quantum simulators [6–9]. These protocols perform a variational learning of the EH from the available measurement data, and have recently led to breakthrough results in ion-trap [10] and cold-atom [11] experiments.

In the above-mentioned applications, it is crucial to have an educated ansatz for the EH, which is mainly guided by the Bisognano-Wichmann theorem of relativistic quantum field theory [12,13]. This provides the EH of a half-infinite system via the physical energy density, weighted by an inverse temperature that increases linearly from the entanglement cut, and is valid in arbitrary dimensions. Generalizations to different geometries exist within conformal field theory (CFT), and yield again a local result with a modified weight function [14–17].

In practice, however, one typically faces a problem, where Lorentz invariance is explicitly broken by the presence of a lattice. Although quantum field theory may still provide an effective low-energy description, it is crucial to address the robustness of the results for the EH. In particular, the analytical solution for a free-fermion chain shows [18] that the lattice EH indeed deviates from the CFT prediction, which can only be recovered after taking a proper continuum limit [19–22]. Nevertheless, it has been demonstrated on a number of examples that the simple lattice discretization of the CFT ansatz provides an excellent approximation of the actual EH at low energies and for large subsystems [6,23–25].

Here we explore a different scenario, where the model is defined in continuous space, but described by the

nonrelativistic Schrödinger equation. We focus on the free-fermion gas, where the entanglement entropy has been studied before [26–28], and shows a logarithmic area-law violation in arbitrary dimensions due to the presence of a Fermi surface [29–33]. Although this result was interpreted via the contributions of independent gapless modes building up the Fermi surface [34], the precise applicability of a CFT description in higher dimensions remained elusive.

Our main goal here is to directly address the EH of the Fermi gas for a d -dimensional spherical domain A with radius R , and compare it to the CFT prediction [14,15]

$$\hat{\mathcal{H}}_{\text{CFT}} = \frac{\pi R}{v} \int_A d^d \mathbf{x} \left(1 - \frac{|\mathbf{x}|^2}{R^2} \right) T_{00}(\mathbf{x}), \quad (1)$$

where $T_{00}(\mathbf{x})$ is the energy density and v is the speed of excitations, which makes $\hat{\mathcal{H}}_{\text{CFT}}$ dimensionless. Its form thus corresponds to an inverse temperature that varies parabolically in the radius and vanishes at the surface of the sphere. The numerical check of Eq. (1) for a free massless scalar field was carried out by first decomposing the EH into angular momentum sectors, and then discretizing the remaining radial problem [35]. While a good agreement with CFT was found at low angular momenta, for higher ones the results are inconclusive.

Our main result is that, in any dimension $d > 1$, the CFT description of the nonrelativistic Fermi gas breaks down at large angular momenta. In particular, we show the equivalence of the entanglement spectra in continuous free space to those of a lattice problem with a linear potential [36]. The mapping identifies the angular momentum with the subsystem boundary on the chain, whereas the radius R sets the length of the region with nontrivial fermion density. While the bulk of this region admits an effective CFT description [37], characterized by a spatially varying Fermi velocity, the fine structure close to the dilute edge is not properly captured. The discrepancy is demonstrated by comparing the actual entanglement spectra and entropies to those that follow from parabolic deformations

Eq. (1) of the physical Hamiltonian, which commute exactly with the EH [38–40].

The free Fermi gas in d dimensions is described by the single-particle Hamiltonian

$$\hat{H} = \frac{\hat{\mathbf{p}}^2}{2m} - \mu, \quad (2)$$

where $\hat{\mathbf{p}} = -i\nabla$ is the momentum operator and the chemical potential $\mu = q_F^2/2m$ sets the filling via the Fermi wave number q_F . The ground state is given by a Fermi sea F , with the plane-wave modes occupied in a spherical domain $|\mathbf{q}| < q_F$. We are interested in a spherical subsystem A of radius R centered around the origin, $|\mathbf{x}| < R$. The entanglement Hamiltonian $\hat{\mathcal{H}}$ is then defined via the reduced density matrix and can be written as [41]

$$\hat{\rho}_A = \frac{1}{Z} e^{-\hat{\mathcal{H}}}, \quad \hat{\mathcal{H}} = \ln(\hat{\mathcal{K}}_A^{-1} - 1), \quad (3)$$

in terms of an integral operator

$$(\hat{\mathcal{K}}_A \psi)(\mathbf{x}) = \int_A d^d \mathbf{x}' K(\mathbf{x}, \mathbf{x}') \psi(\mathbf{x}'), \quad (4)$$

that acts on wave functions in the domain A , with the kernel given by the two-point correlation function

$$K(\mathbf{x}, \mathbf{x}') = \int_F \frac{d^d \mathbf{q}}{(2\pi)^d} e^{i\mathbf{q}(\mathbf{x}-\mathbf{x}')}. \quad (5)$$

We first discuss the simplest case of a one-dimensional (1D) system, where $A = [-R, R]$. After a rescaling $y = x/R$, the integral operator Eq. (4) is given by the famous sine kernel

$$K(y, y') = \frac{\sin c(y - y')}{\pi(y - y')}, \quad (6)$$

which depends on the dimensionless parameter $c = q_F R$. To construct the EH via Eq. (3), one needs to solve $\hat{\mathcal{K}}_A \psi_k = \zeta_k \psi_k$ to find the eigenvalues and eigenfunctions of $\hat{\mathcal{K}}_A$. This can be done by considering instead the differential operator [38,42,43]

$$\hat{D} = -\frac{d}{dy}(1 - y^2)\frac{d}{dy} - c^2(1 - y^2), \quad (7)$$

which commutes with the integral operator, $[\hat{\mathcal{K}}_A, \hat{D}] = 0$. The bounded solutions of the equation $\hat{D} \psi_k = \chi_k \psi_k$ within the domain $|y| < 1$ are known as the angular prolate spheroidal wave functions [44,45], $\psi_k(y) = S_{0k}(c, y)$, and exist for a discrete set of eigenvalues χ_k with $k = 0, 1, \dots$. The eigenvalues of $\hat{\mathcal{K}}_A$ then follow from the radial spheroidal wave functions as $\zeta_k = \frac{2c}{\pi} [R_{0k}(c, 1)]^2$ [38].

It is easy to see that the operator Eq. (7) is a simple parabolic deformation of the original Hamiltonian Eq. (2). Comparing with Eq. (1), one can identify it with the CFT expression after proper rescaling

$$\hat{\mathcal{H}}_{\text{CFT}} = \frac{\pi R}{v_F} \frac{\hat{D}}{2mR^2} = \frac{\pi}{2c} \hat{D}, \quad (8)$$

where the speed must be identified with the Fermi velocity $v_F = q_F/m$. The spheroidal eigenvalues χ_k can be computed using *Mathematica*, and thus the spectrum of $\hat{\mathcal{H}}_{\text{CFT}}$ can be compared against that $\varepsilon_k = \ln(\zeta_k^{-1} - 1)$ of the actual EH in

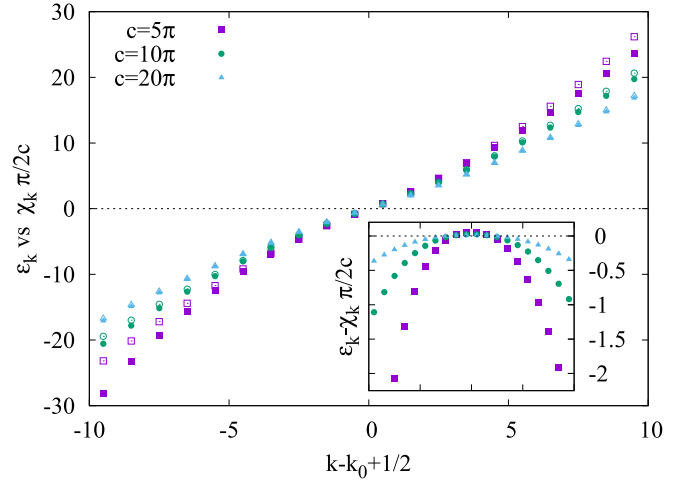


FIG. 1. Single-particle entanglement spectra ε_k of $\hat{\mathcal{H}}$ (full symbols) in one dimension, compared against the spectra of $\hat{\mathcal{H}}_{\text{CFT}}$ (empty symbols) for various values of c . The inset shows the deviations.

Eq. (3). These are shown in Fig. 1, with the full/empty symbols corresponding to ε_k and $\frac{\pi}{2c} \chi_k$, respectively, while the inset shows their difference. Note that the index k was shifted by $k_0 - 1/2$, with $k_0 = 2c/\pi$, to align the low-energy part of the spectra. One clearly observes that the deviation diminishes for increasing c , suggesting the asymptotic equivalence $\hat{\mathcal{H}} \rightarrow \hat{\mathcal{H}}_{\text{CFT}}$ of the operators. This is supported by analytical results [46–48], as well as further numerical evidence [49]. In particular, for finite c one has a series expansion

$$\hat{\mathcal{H}} = \hat{\mathcal{H}}_{\text{CFT}} + \sum_{n=1}^{\infty} \frac{1}{c^n} P_{n+1}(\hat{\mathcal{H}}_{\text{CFT}}), \quad (9)$$

where P_n is an n th-order polynomial. The n th correction term is thus an increasingly nonlocal differential operator of order $2(n+1)$, which is, however, suppressed by c^n . Furthermore, using the lowest-order terms in Eq. (9), we find that the entanglement entropy $S = -\text{Tr}[\rho_A \ln \rho_A]$ is reproduced by $\hat{\mathcal{H}}_{\text{CFT}}$ up to a correction scaling as $\delta S \propto \ln(c)/c^2$, which agrees well with our numerics [49].

We now proceed to the case $d \geq 2$, which considerably simplifies using the rotational symmetry of both A and F . Indeed, setting $\mathbf{x} = r \mathbf{n}$, the Hamiltonian can be decomposed by considering the ansatz for the wave function

$$\psi(r \mathbf{n}) = \frac{\Phi(r)}{r^{(d-1)/2}} Y_{\ell,i}(\mathbf{n}), \quad (10)$$

where $Y_{\ell,i}(\mathbf{n})$ are d -dimensional spherical harmonics [50], with \mathbf{n} being a vector on the surface of the unit sphere, parametrized by $d-1$ angular coordinates. The quantum number $\ell = 0, 1, \dots$ corresponds to the angular momentum, and $i = 1, \dots, M_\ell$ indexes the linearly independent spherical harmonics with fixed ℓ . In this basis, the Hamiltonian $\hat{H} = \bigoplus_{\ell,i} \hat{H}_{\ell,i}$ becomes block diagonal and in the respective sector reads

$$\hat{H}_{\ell,i} = \frac{1}{2m} \left(-\frac{d^2}{dr^2} - q_F^2 + \frac{(\ell + \frac{d-2}{2})^2 - 1/4}{r^2} \right). \quad (11)$$

Note that $\hat{H}_{\ell,i}$ does not depend on the quantum number i , such that one simply has a degeneracy in each sector $\ell \geq 1$ with corresponding multiplicity

$$M_\ell = \frac{2\ell + d - 2}{\ell} \binom{\ell + d - 3}{\ell - 1}, \quad (12)$$

while $M_0 = 1$. Thus the problem boils down to treating the one-dimensional Hamiltonian Eq. (11), where one has an extra contribution from the centrifugal potential. The dimensionality enters via the multiplicities Eq. (12) and a shift of the angular momentum index ℓ . For simplicity, we will discuss the 2D case below, as the generalization to $d > 2$ is trivial.

The eigenvalue problem of the kernel Eq. (5) was considered in Ref. [39], see Ref. [49] for details. One first rewrites it as the absolute square of an exponential kernel $K'(\mathbf{y}, \mathbf{z}) = e^{i\mathbf{c}\mathbf{y}\mathbf{z}}$ in the scaled coordinates $\mathbf{y} = \mathbf{x}/R$ and $\mathbf{z} = \mathbf{q}/q_F$. Separating variables using the ansatz Eq. (10), one is led to consider the radial eigenvalue problem $\hat{K}'_\ell \Phi_{\ell,k} = \gamma_{\ell,k} \Phi_{\ell,k}$, with the kernel given by $K'_\ell(y, z) = J_\ell(cyz) \sqrt{c^2 yz}$. Note that $y = |\mathbf{y}| \leq 1$, $z = |\mathbf{z}| \leq 1$, and the eigenvalues of the original operator \hat{K}_ℓ follow as $\zeta_{\ell,k} = |\gamma_{\ell,k}|^2$. The squared kernel can then be written as

$$K_\ell(y, y') = 2c^2 \sqrt{yy'} K_{\text{Be},\ell}(c^2 y^2, c^2 y'^2) \quad (13)$$

via the Bessel kernel defined as [51]

$$K_{\text{Be},\ell}(u, v) = \frac{\sqrt{v} J_\ell(\sqrt{u}) J'_\ell(\sqrt{v}) - \sqrt{u} J_\ell(\sqrt{v}) J'_\ell(\sqrt{u})}{2(u - v)}. \quad (14)$$

Note that the factor in Eq. (13) in front of the Bessel kernel can be absorbed by a change of variables $u = c^2 y^2$ and $v = c^2 y'^2$, such that the spectrum of \hat{K}_ℓ on the domain $[0, 1]$ is identical to that of $\hat{K}_{\text{Be},\ell}$ on $[0, c^2]$.

Analogously to the 1D case, one can find again a commuting differential operator in each angular momentum sector, $[\hat{K}_\ell, \hat{D}_\ell] = 0$, which reads [39,52]

$$\hat{D}_\ell = -\frac{d}{dy} \beta(y) \frac{d}{dy} - \left(c^2 - \frac{\ell^2 - 1/4}{y^2} \right) \beta(y), \quad (15)$$

with $\beta(y) = 1 - y^2$. Clearly, Eq. (15) can be interpreted as the parabolic deformation of the radial Hamiltonian Eq. (11). Its eigenvalue equation reads $\hat{D}_\ell \Phi_{\ell,k} = \chi_{\ell,k} \Phi_{\ell,k}$, and the eigenfunctions were dubbed generalized prolate spheroidal wave functions. Their asymptotic expressions for $c, k \gg 1$ were studied in Ref. [39]. Moreover, high-precision numerical computation of the eigenvalues $\zeta_{\ell,k}$ and $\chi_{\ell,k}$ is available via an open-source MATLAB code [53,54].

Before turning to the numerics, however, one needs an argument to fix the velocity in the CFT expression Eq. (1). Indeed, the inhomogeneous part of the radial Hamiltonian Eq. (11) can be interpreted as a spatially varying chemical potential $\mu_\ell(r)$. In other words, the effective Fermi energy of the radial motion is reduced by the centrifugal energy of the orbital one. Furthermore, we argue that the only relevant radius in our problem is that of our subsystem, and thus the effective chemical potential should be evaluated at $r = R$. Assuming $R \gg 1$, one obtains for the radial Fermi velocity

$$v_{F,\ell} = \sqrt{\frac{2\mu_\ell(R)}{m}} = v_F \sqrt{1 - \frac{\ell^2}{c^2}}. \quad (16)$$

In particular, $v_{F,\ell}$ vanishes at $\ell = c$, which corresponds to the angular momentum where the classical turning point is given by R . For all $\ell > c$, the eigenfunctions of Eq. (11) have exponentially small amplitudes within A , and thus their contribution to the EH should be negligible.

Alternatively, the emergence of the Fermi velocity Eq. (16) can be understood by mapping the problem to that of an inhomogeneous quantum chain. This can be achieved using a remarkable identity found in Ref. [55], which establishes a connection between the Bessel kernel Eq. (14) and the analogous discrete Bessel kernel

$$\mathcal{K}_{\text{dBe},c}(i, j) = \frac{c J_{i-1}(c) J_j(c) - c J_i(c) J_{j-1}(c)}{2(i - j)}, \quad (17)$$

where $i, j \in \mathbb{Z}$. The identity relates the trace of an integer power of the corresponding operators [55]

$$\text{Tr}_{[0,c^2]}(\hat{K}_{\text{Be},\ell}^n) = \text{Tr}_{[\ell+1,\infty)}(\hat{K}_{\text{dBe},c}^n), \quad (18)$$

where the subscripts denote the domains of the respective kernels, over which the trace is carried out, with the right-hand side being the trace of an ordinary matrix. Since the relation holds for arbitrary n , this implies that the spectra of the two operators are identical.

The matrix defined in Eq. (17) is precisely the correlation matrix of a hopping chain with a linear potential [36], and unitary equivalent to the one describing domain-wall melting [56]. The parameter c now plays the role of the half-width of the front region, where the fermion density differs from one and zero. Moreover, the angular momentum ℓ is identified with the position of the entanglement cut. In turn, the expression Eq. (16) simply corresponds to the spatial dependence of the Fermi velocity due to the variation of the filling within the front region [57,58]. The CFT prediction for the respective EH thus reads

$$\hat{\mathcal{H}}_{\ell,\text{CFT}} = \frac{\pi}{2\sqrt{c^2 - \ell^2}} \hat{D}_\ell. \quad (19)$$

To test the validity of the ansatz Eq. (19), we evaluate and compare the entropies obtained from $\hat{\mathcal{H}}_\ell$ and $\hat{\mathcal{H}}_{\ell,\text{CFT}}$, as shown in Fig. 2. The agreement is excellent in the bulk of the profile, where the asymptotics of the spectra $\varepsilon_{\ell,k}$ with ℓ/c fixed were studied numerically for the gradient chain [59]. The resulting entropy profile

$$S_\ell = \frac{1}{6} \ln(c) + \frac{1}{4} \ln[1 - (\ell/c)^2] + \mathcal{C}, \quad (20)$$

where $\mathcal{C} \approx 0.4785$ is a nonuniversal constant [60,61], is shown by the red line. In fact, Eq. (20) can also be derived using a curved-space CFT approach [37], where the inhomogeneous metric is chosen to absorb the spatial variation of the Fermi velocity. While Eq. (20) gives an accurate description of the bulk entropy profile, it does not capture the fine structure around the edge $\ell \approx c$, where also the ansatz Eq. (19) seems to break down. Indeed, using the scaling variable $(\ell - c)/c^{1/3}$, the correlation matrix Eq. (17) can be approximated by the Airy kernel [62], and S_ℓ displays a corresponding edge scaling [63]. As shown by the inset of Fig. 2, the same holds true for the difference $\delta S_\ell = S_\ell - S_{\ell,\text{CFT}}$, which shows a data collapse for various values of c .

The situation is very similar in $d > 2$ dimensions, where the index of the Bessel kernel is $\ell + (d - 2)/2$. This is a

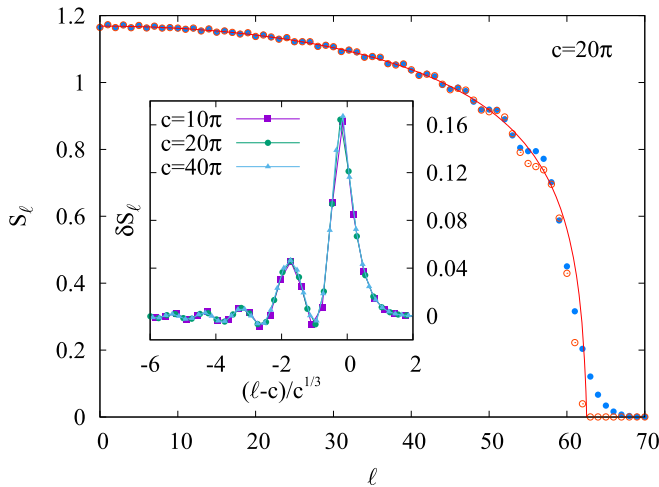


FIG. 2. Entanglement entropies S_ℓ (full symbols) and $S_{\ell,\text{CFT}}$ (empty symbols), calculated using the ansatz Eq. (19) for $c = 20\pi$. The red solid line shows the result Eq. (20). Inset: Difference of the entropies in the appropriately rescaled edge regime, for various values of c .

half-integer in odd dimensions, such that the one-to-one correspondence with the gradient chain is lost. Nevertheless, when plotted against the shifted index $\ell + (d-2)/2$, the entropy profile S_ℓ smoothly interpolates between the data points of the $d=2$ case. Applying the same shift in the scaling factor in Eq. (19), the plot of the 3D case is almost identical to Fig. 2. One thus concludes that the CFT ansatz breaks down for high angular momenta $\ell \approx c - (d-2)/2$. Due to the increasing multiplicities M_ℓ with the dimensionality, however, the leading-order mismatch of the total entropy in $d \geq 2$ scales as

$$\delta S = \sum_{\ell} M_{\ell} \delta S_{\ell} \propto \frac{c^{d-2}}{(d-2)!} c^{1/3}. \quad (21)$$

Thus, in sharp contrast to the 1D case, the entropy deviation becomes divergent in the $c \rightarrow \infty$ limit. This is a consequence of the edge-scaling regime in angular-momentum space, which is not properly described by CFT. The scaling Eq. (21) is consistent with our numerics in Fig. 3, albeit with strong subleading corrections.

The mapping to the gradient chain, with resulting entropy profile Eq. (20), also allows us to obtain the analytical result for the total entropy

$$S = \sum_{\ell} M_{\ell} S_{\ell} \simeq \sigma_d c^{d-1} \ln c + A_d c^{d-1}, \quad (22)$$

where the prefactors can be calculated as [49]

$$\sigma_d = \frac{1}{3(d-1)!}, \quad A_d = \frac{4\mathcal{C} - \psi\left(\frac{d+1}{2}\right) - \gamma}{2(d-1)!}, \quad (23)$$

with $\psi(x)$ being the digamma function and γ the Euler-Mascheroni constant. It is easy to check that the prefactor of the area-law violating term agrees with the general expression found in Refs. [29,32]. The area-law contribution is nonuniversal, and follows from the summation of the second and third terms in Eq. (20) [49]. We tested the prediction Eq. (22) by adding a subleading term $B_d c^{d-2}$ and fitting to

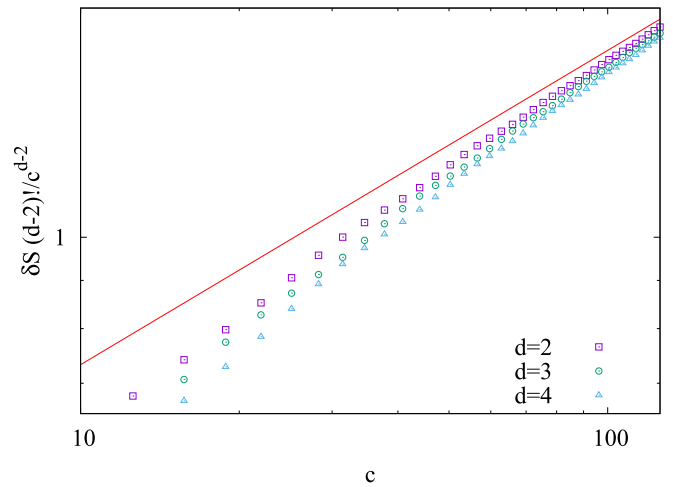


FIG. 3. Deviation of the total entropy from the one calculated via the CFT EH for various dimensions, scaled according to Eq. (21). The red solid line with slope $1/3$ is a guide to the eye.

our numerical data. The results $\sigma_2 = 0.3332$, $A_2 = 0.651$ and $\sigma_3 = 0.1667$, $A_3 = 0.2285$ for the 2D and 3D cases, respectively, are in excellent agreement with Eq. (23). Note that our result on A_d also agrees with the conjecture formulated in Ref. [64]. One should also remark that, in free massless relativistic theories, no violation of the area law occurs [65–67].

In conclusion, we have found that the EH of a nonrelativistic Fermi gas is well reproduced by the appropriately rescaled parabolic deformation of the physical Hamiltonian. While in 1D the relation becomes asymptotically exact in the limit of large subsystems, the situation in higher dimensions is much more subtle. First, the CFT prediction can only be applied in the angular momentum sectors, after rescaling with the Fermi velocity of the radial motion. Since this velocity carries a nontrivial dependence on ℓ , the relation cannot be lifted back to the total EH and rewritten as a deformation of the total energy density as in Eq. (1). In sharp contrast, for relativistic Dirac fermions the form of the total EH is identical to those in the sectors [67]. Second, for the nonrelativistic case some deviations persist even in the sectors for large angular momenta. Using the mapping to the gradient chain, these discrepancies can be traced back to the dilute edge regime of the fermionic density, where the fine structure of the correlations does not admit a CFT description. Hence, translating Eq. (1) to nonrelativistic systems requires proper insight and care.

Our work opens up various directions for future research. One could address how the shape of the Fermi surface, which is known to be crucial for the entropy scaling [29,32], affects the results for the EH. A further natural extension would be the study of a trapped Fermi gas [68], where the CFT predictions for the EH are also available [69]. Finally, one should investigate how the results generalize to particles with bosonic statistics.

I thank I. Peschel and E. Tonni for fruitful discussions and correspondence. In my numerical calculations I used the open-source code available online [54]. The author acknowledges funding from the Austrian Science Fund (FWF) through Project No. P35434-N.

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