Feshbach resonance of heavy exciton-polaritons

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We study interactions between polaritons formed by hybridization of excitons in a two-dimensional semiconductor with surface optical phonons or plasmons. These quasiparticles have a high effective mass and can bind into bipolaritons near a Feshbach-like scattering resonance. We analyze the phase diagram of a many-body condensate of heavy polaritons and bipolaritons and calculate their absorption and luminescence spectra, which can be measured experimentally.

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I. INTRODUCTION

Exciton-polaritons, formed by coherent coupling of excitons with photons in semiconductor microcavities, have been a subject of active research [1,2]. These hybrid lightmatter quasiparticles are of interest due to their potential device applications [3] and solid-state realizations of condensation and superfluidity [4,5]. Many phenomena unique to exciton-polaritons stem from strong interactions induced by their excitonic component. Recent pump-probe optical experiments revealed that exciton-polaritons can exhibit a Feshbach scattering resonance mediated by the biexciton state [6–11], which in some sense allows one to control the strength and sign of polariton interactions, akin to the manipulation performed in experiments with atomic gases.

The origin of the exciton-polariton Feshbach resonance is associated with formation of bipolaritons, i.e., bound states of polaritons. One might expect that bipolaritons are realized when the Rabi splitting (Sec. II A) exceeds the biexciton binding energy. However, due to the strong dispersion of the photon, exciton-polaritons have a very small effective mass and the bipolariton binding energy is exponentially suppressed in two dimensions. Thus, bipolariton states are not observed [12], although radiative corrections can significantly modify the biexciton dispersion [13–16]. Note that while such radiatively renormalized biexciton states are often called bipolaritons in the literature, in this paper we use the term bipolariton to mean a true bound state of two polaritons with energy below the two-particle continuum.

Polaritons in two-dimensional (2D) materials, such as phonon-polaritons in hexagonal boron nitride and exciton-polaritons in transitional-metal dichalcogenide (TMD) monolayers, are a growing research field, which provides a controllable platform to study polaritons and their interactions [17,18]. Motivated by this, we consider a novel type of exciton-polariton formed by strong coupling of excitons in a 2D semiconductor with surface optical phonons or surface plasmons, see Figs. 1(a) and 1(b). In contrast to previous investigations of resonant exciton-phonon coupling [19–21], we consider the regime where the phonon frequency is resonant with transitions across the band gap, not between internal states of the exciton. This is similar to the case of plexcitons, which are hybrid quasiparticles resulting from resonant coupling between excitons and plasmons in metallic nanostructures [22–24]. By analogy to heavy fermions, we use a common term heavy exciton-polaritons for these quasiparticles to emphasize that they have a large effective mass inherited from a nearly flat dispersion of phonons or plasmons. This is why heavy exciton-polaritons can be confined in nanocavities of size much smaller than the diffraction limit of light. Note that such highly confined heavy polaritons have a very small photonic component. In this regard, we wish to clarify that we use the term polariton in a broad sense, meaning a hybrid excitation of a polarizable medium.

In this paper, we analyze the Feshbach scattering resonance between heavy polaritons and show that they can form bipolariton states near the resonance. In contrast to previously studied exciton-polaritons in photonic cavities, where paired states are essentially biexcitons, heavy polaritons retain their hybrid nature in the bound state [Fig. 1(c)] assuming that Rabi frequency exceeds the biexciton binding energy. We subsequently consider a system with finite densities of polaritons and bipolaritons that can exhibit Bose condensation and superfluidity. We study possible phases of such a condensate, the polariton superfluid (PSF) and bipolariton superfluid (BSF) [25–28], and the associated excitation spectra within a meanfield approximation.

The rest of this paper is organized as follows. In Sec. II we first discuss single-particle states and compare heavy exciton-polaritons to photon exciton-polaritons. We then include exciton-exciton interactions and show how bipolariton states emerge near the Feshbach resonance. In the last part of Sec. II, we study the consequences of heavy bipolariton formation in many-body condensates. In Sec. III, we compute the absorption and luminescence spectra of the polariton systems and show that they are dominated by collective excitations of the condensate. We discuss additional experimentally testable predictions of our theory in Sec. IV. Details of our calculations are provided in the Appendix.



FIG. 1. Schematic diagrams of the (a) photon cavity and (b) phonon/plasmon cavity. The lines indicate the electric field direction. (c) Schematic of the heavy exciton-polariton (left) and bipolariton (right). The ring surrounding the exciton indicates the characteristic polariton radius $R_c = (1/m_x \Omega)^{1/2}$, with m_x the exciton mass and Ω the polariton Rabi frequency. Bipolaritons are well defined when the biexciton binding energy E_{xx} is much smaller than Ω , or equivalently R_c is much smaller than the length scale $r_{xx} = (1/m_x E_{xx})^{1/2}$, which is of order of the bipolariton radius.

II. MODEL

A. Single-particle states

We consider a simple model of the heavy polariton that captures the essential physics. The Hamiltonian is

$$H_0 = \sum_{k} \left[\omega_{x,k} b_k^{\dagger} b_k + \omega_{c,k} a_k^{\dagger} a_k + \frac{\Omega_k}{2} (b_k^{\dagger} a_k + \text{H.c.}) \right], \quad (1)$$

which is analogous to what is commonly used to model exciton-polaritons in photonic cavities [2]. In that context the operators a_k^{\dagger} and b_k^{\dagger} , respectively, create cavity mode and exciton states with momentum k. The exciton kinetic energy is $\omega_{x,k} = k^2/2m_x$, where m_x is the exciton mass. In our case a_k^{\dagger} creates a phonon or a plasmon (Fig. 1) with energy $\omega_{c,k} = \delta$ independent of k since the dispersion of a surface optical phonon (plasmon) is nearly flat compared to that of the exciton. We refer to parameter δ as the cavity detuning. The strength of the coupling is characterized by the Rabi frequency Ω_k , i.e., the rate of energy transfer between the two modes, which we take to be momentum independent for simplicity:



FIG. 2. Plots of the polariton energy (left) and the exciton fraction (right) versus momentum for (a) photon exciton-polaritons and (b) heavy exciton-polaritons.

 $\Omega_k = \Omega$. We also ignore the polarization degree of freedom of the exciton. We use units such that $\hbar = 1$ throughout. The spectrum consists of upper and lower polariton branches with energies

$$\omega_{\pm,\boldsymbol{k}} = \frac{1}{2} \left[\omega_{c,\boldsymbol{k}} + \omega_{x,\boldsymbol{k}} \pm \sqrt{(\omega_{x,\boldsymbol{k}} - \omega_{c,\boldsymbol{k}})^2 + \Omega_{\boldsymbol{k}}^2} \right], \quad (2)$$

and the exciton fraction in each branch is given by the squares of the Hopfield coefficients

$$U_{\pm,k}^{2} = \frac{1}{2} \Big[1 \pm (\omega_{x,k} - \omega_{c,k}) \Big/ \sqrt{(\omega_{x,k} - \omega_{c,k})^{2} + \Omega_{k}^{2}} \Big].$$
(3)

In Fig. 2, we compare the momentum dependence of the polariton energy and Hopfield coefficients for conventional photon exciton-polaritons and heavy exciton-polaritons. Cavity photons have a very steep energy-momentum dispersion, which takes the form $\omega_{p,k} \simeq \delta + k^2/2m_p$, with an effective mass $m_p \sim 10^{-4}m_x$. Therefore, for a photonic cavity, the polaritons have a strong dispersion at small k, but with increasing k outside the light cone the lower polariton rapidly becomes mostly exciton and the upper polariton becomes mostly photon. This is contrasted with the heavy polaritons, which have an effective mass of order of the exciton mass m_x and retain their hybrid character at much larger k.

B. Two-particle states

The interaction between two excitons, viewed as rigid structureless bosons, takes the form

$$H_{\rm int} = \frac{1}{2A} \sum_{kk'q} W_q \, b^{\dagger}_{k+q} b^{\dagger}_{k'-q} b_{k'} b_k, \qquad (4)$$

where A is the area of the system and W_q is an attractive interaction that supports the biexciton bound state. The structureless boson approximation for excitons is justified if their binding energy is much greater than all other relevant energy scales in the problems, including the Rabi frequency, the biexciton binding energy, etc. If this is not true, a more sophisticated formalism is required [29]. The total Hamiltonian is $H = H_0 + H_{int}$, with H_0 defined in Eq. (1). We may estimate the bipolariton binding energy *E* from the well-known formula for weak attractive interactions in two dimensions [30]:

$$E \simeq E_{\Lambda} \exp\left(-\frac{2\pi}{|V_p|m_r}\right),$$
 (5)

with $V_p < 0$ the polariton interaction strength at zero momentum, m_r the reduced mass of the two polaritons forming the bound state, and E_{Λ} a high-energy cutoff determined by the range of the interaction. For photon exciton-polaritons, $m_r \sim m_p \sim 10^{-4} m_x$, so the binding energy is exponentially small, as mentioned earlier in Sec. I (see also Appendix A). In contrast, for heavy polaritons m_r and V_p are of the same order as those of the exciton. Therefore, strongly bound bipolaritons are expected near the biexciton resonance. We find the bipolariton energies from the poles of the exciton two-particle scattering matrix. To simplify the calculation we choose W_a to be momentum independent, i.e., a contact interaction, and subsequently eliminate it in favor of the biexciton binding energy E_{xx} to obtain the renormalized scattering matrix $\mathcal{T}(\mathbf{K}, \omega)$, which depends on the total incoming momentum K and energy ω of the particles. We give an explicit formula for the scattering matrix, along with details of the calculation, in Appendix A.

Here and in the following calculations we choose $\Omega =$ $3E_{xx}$ for the Rabi frequency. In Fig. 3(a) we plot the bipolariton energies versus cavity detuning δ . There are two Feshbach resonances. The first one occurs when the energy of a biexciton (the horizontal dotted line) coincides with the energy of two lower polaritons (the lower boundary of the continuum), $2\omega_{-} = -E_{xx}$, all at momenta k = 0. The second resonance is found where the biexciton energy is equal to the sum of the energies of a lower and upper polariton, $\omega_+ + \omega_- = -E_{xx}$. In both cases there are bipolariton states lying below the continuum of unbound two-particle states, e.g., $\omega = \omega_{-}(\mathbf{k}) + \omega_{-}(\mathbf{k})$ $\omega_{-}(-k)$. In principle, two upper polaritons can also form a quasibound state but it lies in the continuum and is damped. Figure 3(b) shows the binding energies and radii of the states outside the continuum versus δ . The bipolariton states become more loosely bound with decreasing δ as the exciton fraction u_{-}^2 decreases, e.g., for two lower polaritons $m_r \propto u_{-}^{-2}$ and $V_p \propto u_{\perp}^4$, hence their binding energy decreases towards zero in accordance with Eq. (5) while their radius diverges.

C. Phase diagram and excitations

Many-body physics of a system with a finite concentration of polaritons may be described by a two-channel effective Hamiltonian [28] that explicitly includes polariton (ξ_1) and bipolariton (ξ_2) fields:

$$H_{\text{eff}} = \sum_{i=1,2} \sum_{k} (\omega_{i,k} - \mu_i) \xi_{i,k}^{\dagger} \xi_{i,k} + \frac{1}{2A} \sum_{ij} \sum_{kk'q} g_{kk'q}^{ij} \xi_{i,k+q}^{\dagger} \xi_{j,k'-q}^{\dagger} \xi_{j,k'} \xi_{i,k} + \frac{1}{\sqrt{2A}} \sum_{kq} (\alpha_{kq} \xi_{1,q/2+k}^{\dagger} \xi_{1,q/2-k}^{\dagger} \xi_{2,q} + \text{H.c.}). \quad (6)$$



FIG. 3. (a) Energies of the bipolariton states versus detuning δ for $\Omega = 3E_{xx}$. The solid and dashed black lines are the energies of lower-lower, upper-lower, and upper-upper bipolaritons, denoted respectively by ω_{--} , ω_{+-} , and ω_{++} . The horizontal dotted line is the biexciton energy, the negative of the biexciton binding energy E_{xx} . The dashed-dotted lines are the sum of the energies of two polaritons at zero momentum, with the two-polariton continuum shown by the shaded regions above. (b) The binding energies $E_{--} = 2\omega_{-} - \omega_{--}$ and $E_{+-} = \omega_{+} + \omega_{-} - \omega_{+-}$ and corresponding radii of the bipolariton states versus detuning δ . The radii $r_{xx'}$ are defined by $r_{xy'}^2 = \int d^2r r^2 |\psi_{xy'}(\mathbf{r})|^2$, with $\psi_{xy'}(\mathbf{r})$ the bipolariton wave function, and $r'_{xx} = (2/3m_x E_{xx})^{1/2}$ is the biexciton radius for the contact potential model.

We assume that only the "–" polaritons, which are the lower-energy states, are present. Parameter μ is the chemical potential; $\mu_1 = \mu$, $\mu_2 = 2\mu$; $\omega_{1,k} = \omega_{-,k}$ and $\omega_{2,k}$ [defined by Eq. (8) below] are the energies of the polariton and bipolariton, respectively, and α is the polariton-bipolariton coupling. Parameters g^{ij} are repulsive background interactions, which depend on the Hopfield coefficients according to

$$g_{kk'q}^{ij} = U_{-,k}U_{-,k'}U_{-,k+q}U_{-,k'-q}\tilde{g}^{ij}/m_x,$$
(7)

and \tilde{g}^{ij} are dimensionless interaction strengths, which we assume to be constant. The parameters α and $\omega_{2,k}$ are determined by an expansion of the polariton-polariton scattering matrix near the bipolariton pole (see Appendix A):

$$\mathcal{T}_{--}(\boldsymbol{k}, \boldsymbol{k}', \boldsymbol{K}, \omega) = U_{-,\boldsymbol{K}/2+\boldsymbol{k}} U_{-,\boldsymbol{K}/2-\boldsymbol{k}} U_{-,\boldsymbol{K}/2+\boldsymbol{k}'} U_{-,\boldsymbol{K}/2-\boldsymbol{k}'} \\ \times \mathcal{T}(\boldsymbol{K}, \omega) \simeq \frac{\alpha_{\boldsymbol{k}\boldsymbol{K}}\alpha_{\boldsymbol{k}'\boldsymbol{K}}}{\omega - \omega_{2} \kappa}.$$
(8)



FIG. 4. (a) Schematic of polariton distribution in the BSF and PSF phases. (b), (c) Ratio of single polariton condensate density to the total density at zero temperature as a function of cavity detuning δ and chemical potential μ (b) or density *n* (c). The background scattering parameters are $\tilde{g}^{11} = 6$, $\tilde{g}^{12} = 9$, and $\tilde{g}^{22} = 20$. The normal phase is labeled by N, the polariton superfluid by PSF, the bipolariton superfluid by BSF, and phase separated regions by PS.

The system exhibits a transition from a polariton to bipolariton superfluid with changing density and detuning. In the polariton superfluid phase (PSF) both polaritons and bipolaritons condense; in the bipolariton superfluid phase (BSF) only bipolaritons condense. Following Refs. [26,28], we calculate the mean-field phase diagram at zero temperature using the two-channel model of Eq. (6). We caution the reader that these phase diagrams are not suitable for immediate comparison with experiment because quantum fluctuations beyond meanfield theory can significantly shift the phase boundaries [28]. Additionally, there is a significant uncertainty in the values of phenomenological parameters \tilde{g}_{ii} (see more below).

We plot the ratio of the single polariton condensate density to the total density in Figs. 4(b) and 4(c) [note that $r_{xx} = (1/m_x E_{xx})^{1/2}$], showing the BSF to PSF phase transition with decreasing δ , which can be either first order or continuous. In the case of a first-order transition, there is a region of phase separation between the PSF phase and the BSF or normal phase. The mean-field phase boundary of the continuous transition is described by the equation

$$E_{--} = (g_{22}/2 - g_{12})n + \alpha \sqrt{2n}, \qquad (9)$$

with E_{--} the bipolariton binding energy. This equation relates the total density *n* to the cavity detuning δ . Note that the parameters \tilde{g}^{ij} are difficult to determine even for simplified microscopic models; furthermore, they depend crucially on details such as the exciton spin structure not included in our simple model. We crudely estimate these parameters by requiring that the BSF-PSF transition occur at a critical density n_M in accordance with the Mott criterion $n_M r_{--}^2 \sim 0.03$ [31,32]. Here r_{--} is the bipolariton radius, which is plotted versus detuning in Fig. 3(b). Associated with this criterion is the picture of particle distribution in real space, Fig. 4(a). In the BSF phase, polaritons form a dilute gas of bound pairs. As the system approaches the phase transition, the binding energy of these pairs decreases while their radius becomes larger so they begin to overlap in space. As a result, a significant fraction of these pairs dissociate leading to the formation of a single-polariton condensate in addition to the bipolariton condensate in the PSF phase.

As discussed below in Sec. III, the optical response of the PSF and BSF phases is dominated by their collective excitations. The excitation spectrum has one gapless acoustic mode and one gapped mode in each phase [26]. In the BSF phase, the gapless mode is due to phase oscillations of the bipolariton condensate and the gapped mode is due to pair breaking of bipolaritons into two polaritons. The energy of the gapped mode is

$$E_1^{\text{BSF}} = \left\{ \left[\frac{1}{2} (E_{--} - g_{22} n_2 + 2g_{12} n_2) \right]^2 - \alpha^2 n_2 \right\}^{1/2}, \quad (10)$$

where n_1 is the single polariton density and n_2 is the bipolariton density. In the PSF phase, the gapless and gapped modes are due to in-phase and out-of-phase fluctuations of polariton and bipolariton fields, respectively. The energy of the gapped mode is

$$E_1^{\text{PSF}} = \{ [E_{--} - g_{12}n_1 - g_{22}n_2 + 2g_{11}n_1 + 2g_{12}n_2 - \alpha\sqrt{n_2}]^2 + \alpha n_1 [\alpha - 2g_{12}\sqrt{n_1} + (4g_{11} + g_{22})\sqrt{n_2}] \}^{1/2}.$$
(11)

Both E_1^{BSF} and E_1^{PSF} vanish along the phase boundary given by Eq. (9). Far from the phase transition, and neglecting the Feshbach coupling α , the energies of the gapped modes have a simple interpretation. We may write $E_1^{\text{PSF}} = \tilde{\omega}_2 - 2\tilde{\omega}_1$, where $\tilde{\omega}_1 = g_{11}n_1 + g_{12}n_2$ and $\tilde{\omega}_2 = -E_{--} + g_{12}n_1 + g_{22}n_2$ are, respectively, the polariton and bipolariton energies renormalized by interaction. This corresponds to the energy to create a bipolariton. Similarly, $E_1^{\text{BSF}} = \frac{1}{2}(2\tilde{\omega}_1 - \tilde{\omega}_2)$ corresponds to the energy to break a bipolariton into its constituent polaritons.

III. EXPERIMENTAL SIGNATURES

One way to distinguish PSF and BSF phases is by measuring the optical absorption or luminescence spectrum. The spectrum in both normal and condensed states is determined from the collective excitations of the system, which we calculate from the effective Hamiltonian of Eq. (6). Details of the calculation are given in Appendix B. There are two contributions: direct coupling of polaritons to photons, which gives rise to sharp emission lines, and coupling of a bipolariton to a photon and polariton, which yields a broad continuum. In Fig. 5, we show plots of the absorption and luminescence spectra for a polariton condensate at zero temperature.

The energy of a photon measured in an optical experiment is the sum (difference) of the chemical potential μ and excitation energy corresponding to absorption (emission) of a photon along with emission of a collective excitation. Therefore, the gapless modes follow the lines $\omega = \mu$ (the solid lines in Fig. 5), and the gapped modes are positioned above or below these lines. Within the mean-field theory, the chemical potential is given by

$$\mu = \begin{cases} g_{11}n_1 + g_{12}n_2 - \alpha \sqrt{n_1 n_2} & (\text{PSF}) \\ \frac{1}{2}(-E_{--} + g_{22}n_2) & (\text{BSF}). \end{cases}$$
(12)



FIG. 5. (a) Absorption and (b) luminescence spectra at zero temperature for density $nr_{xx}^2 = 0.015$ (left) and $nr_{xx}^2 = 0.03$ (right). The chemical potential μ is denoted by the solid lines, and dashed gray lines indicate thresholds associated with the emission of collective excitations. The absorption spectrum is negative for frequencies below μ , corresponding to optical gain, and positive for frequencies above μ . Since only emission (and not absorption) of collective excitations is possible at zero temperature, luminescence only occurs at frequencies below μ .

The energy $\omega = \mu + E_1^{\text{PSF}}$ of the collective mode above μ in the PSF phase and the energy $\omega = \mu - E_1^{\text{BSF}}$ below μ in the BSF phase, can be viewed as the renormalized energy of the bipolariton spectral line. Far from the transition, this energy is given by the formula

$$\omega = \tilde{\omega}_2 - \tilde{\omega}_1 = -E_{--} + (g_{12} - g_{11})n_1 + (g_{22} - g_{12})n_2.$$
(13)

At large positive δ , the system is in the BSF phase, where bipolaritons are energetically favored. The gap between μ and the absorption or emission threshold is the energy of the gapped excitation mode, given by Eq. (10). The gapless mode, which is due to phase oscillations of the bipolariton condensate, is not observed in this phase because bipolaritons do not couple directly to light. With decreasing detuning, the gap decreases and closes at the phase transition, where single polariton condensation becomes favorable. In the PSF phase, both gapless and gapped modes are observed. The energy of the gapped mode is given by Eq. (11).

We now discuss the relation of our results to previous experiments on the polaritonic Feshbach resonance [8–10]. Photon exciton-polaritons studied in those experiments have a polarization or pseudospin degree of freedom since there are two countercircular polarization states of the photon that couple to excitons with the same polarization. Since the interaction between two excitons is attractive (repulsive) when they have antiparallel (parallel) spin, only polaritons with antiparallel spin can form biexcitons. In the experiments, the system is pumped with circularly polarized light to create a condensate of spin-up polaritons. A probe beam with the opposite circular polarization then excites a few spin-down polaritons, which interact with the spin-up condensate to form biexcitons. This results in a shift in the spin-down polariton energy observed in the probe transmission spectrum. The



FIG. 6. Density profiles for polariton condensate in a harmonic trap for first-order transition (left) and continuous transition (right). The plots correspond, respectively, to $\delta/E_{xx} = 0$ and $\delta/E_{xx} = 2$ in the phase diagram of Fig. 4(b). The trap potential is $V(x) = \frac{1}{2}E_{xx}(x/r_{trap})^2$ with $r_{trap} = 300r_{xx}$. Here n_1 is the single polariton density, n_2 is the bipolariton density, and $n = n_1 + 2n_2$ is the total density.

shift changes from positive to negative with decreasing cavity detuning as the system is tuned across the Feshbach resonance, which is often colloquially described as the interaction changing from repulsive to attractive. The corresponding effect in our model is the change in the collective mode energy from $-E_1^{\text{BSF}}$ to E_1^{PSF} in Fig. 5(a) and crossing between the photon energy ω in Eq. (13) and chemical potential μ . In other words, the renormalized bipolariton binding energy effectively changes from positive in the BSF phase to negative in the PSF phase. However, the aforementioned experiments have studied a transient state rather than the true ground state of the polariton system, since in principle biexcitons should also condense.

In addition to photoluminescence spectroscopy of a uniform density system, the PSF to BSF transition can be detected by imaging the polariton condensate confined in a trap. In Fig. 6, we show the density profile in a harmonic trap calculated in the Thomas-Fermi approximation by solving $\mu[n(r)] + V(r) = \mu_0$ for n(r), where $\mu[n]$ is the local chemical potential as a function of density, V(r) = $\frac{1}{2}E_{xx}(r/r_{\text{trap}})^2$ is the trap potential with $r_{\text{trap}} = 300r_{xx}$, and μ_0 is the chemical potential at the center of the trap. Since the local chemical potential $\mu(r)$ decreases with increasing radial coordinate r, there is a transition from the PSF to BSF phase, which results in a discontinuity in the density n(r) or its derivative $\frac{dn}{dr}$ for a first-order or continuous transition, respectively. Since μ is continuous across the phase boundary of the continuous transition, $\Delta \mu = 0$, we may relate the discontinuity in $\frac{\partial \mu}{\partial n}$ to that of $\frac{\partial \mu}{\partial \delta}$ by $\Delta \frac{\partial \mu}{\partial \delta} = -\frac{dn}{d\delta} \Delta \frac{\partial \mu}{\partial n}$, with $\frac{dn}{d\delta}$ determined from the phase boundary in Fig. 4(c). From Fig. 5, $\frac{\partial \mu}{\partial \delta}|_{c-} = -0.022$ and $\frac{\partial \mu}{\partial \delta}|_{c+} = -0.026$, where c - (c+) indicates that the derivative is evaluated just below (above) the transition point. This gives $\Delta \frac{\partial \mu}{\partial \delta} \approx -0.004$, and since $\frac{dn}{d\delta} \approx 0.002/(E_{xx}r_{xx}^2)$, we have $\Delta \frac{\partial \mu}{\partial n} \approx 2E_{xx}r_{xx}^2$. Therefore, $\Delta(\frac{dn}{dr})^{-1} = \frac{r_{trap}^2}{E_{xx}r} \Delta \frac{\partial \mu}{\partial n} \approx 10 r_{trap} r_{xx}^2$, which is consistent with the slopes in Fig. 6: $\frac{dn}{dr}|_{c-} \approx -0.5/(r_{\text{trap}}r_{xx}^2)$ and $\frac{dn}{dr}|_{c+} \approx$ $-0.05/(r_{\text{trap}}r_{xx}^2)$. This explains the large kink seen in the density profile of Fig. 6 (right) despite the small discontinuity of $\frac{\partial \mu}{\partial \delta}$ in Fig. 5.

IV. DISCUSSION AND OUTLOOK

We have investigated a novel type of quasiparticle, heavy polaritons, formed by strong coupling of excitons in a 2D semiconductor with surface optical phonons or plasmons. Similar to systems of cold atoms, where this phenomenon was first studied, heavy polaritons exhibit a Feshbach resonance when the biexciton energy becomes resonant with that of two polaritons. Such resonances have been recently probed in experiments with exciton-polaritons in photonic cavities. However, in those experiments the polaritons did not form true bound states. In contrast, we predict that heavy polaritons bind into bipolaritons near the Feshbach resonance. For a system with a finite density of polaritons and bipolaritons we have analyzed the possible phases, polariton (PSF) and bipolariton (BSF) superfluids, and their collective excitations within a mean-field approximation. We have computed the absorption and luminescence spectra in these phases.

Another observable signature of polariton condensation is formation of quantized vortices. Such vortices can be induced in the system by external perturbations or nonequilibrium flow and subsequently detected by optical imaging [33-35]. According to Ref. [26], in the PSF phase the Feshbach coupling of polaritons and bipolaritons induces a splitting of a 2π polariton vortex into two π vortices connected by a domain wall. Observation of this vortex splitting would confirm the existence of the PSF phase.

In conclusion, we briefly mention possible materials realizations of our results. Since phonons have frequencies in the terahertz region, resonantly coupling them to excitons requires a narrow band gap semiconductor. One candidate is gapped bilayer graphene, where tunable excitons have been observed with energies around 100 meV [36,37]. These excitons can be tuned into resonance with the low-loss hyperbolic phonon modes in hBN [38,39]. Regarding plasmons, several experiments have demonstrated strong coupling of excitons in TMD monolayers with plasmons in metallic nanostructures [24,40]. TMDs host strongly bound biexcitons with $E_{xx} \approx$ 50 meV [41]. Also, a highly controllable realization of plasmonic strong local coupling with excitons can be achieved using a nano-optical antenna [42,43]. Note that previously the term heavy polariton has been used to describe hybridization of excitons with heavy photons, i.e., flat photonic bands, in an optical lattice [44]. Even though the range of momenta where this band dispersion remains flat is relatively narrow, this could potentially be another path to realization of the physics described in this paper.

Possible extensions of our work include studying a Josephson-like effect where the optical spectrum in the PSF phase depends on the relative phase of polariton and bipolariton condensates and analyzing the collective modes of the polariton condensate in a trap. A potential application of the heavy polariton Feshbach resonance is the generation of entangled pairs of polaritons [45], which is another interesting subject for future study. In addition, experiments with atomic Bose gases near a Feshbach resonance have observed novel phenomena including Efimov states [46], universal dynamical behavior [47], and formation of solitons [48] or quantum droplets [49]. The system of heavy polaritons studied

in this paper provides another potential avenue to realize these phenomena.

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APPENDIX A: SCATTERING MATRIX CALCULATION

The scattering matrix satisfies the Bethe-Salpeter equation, which with the notation $k = (\mathbf{k}, i\omega_n)$ takes the form [11]

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$$\mathcal{T}(k, k'; k+q, k'-q) = W(k, k', q) + \sum_{q'} W(k, k'q') G_x(k+q') \times G_x(k'-q') \mathcal{T}(k+q', k'-q'; k+q; k'-q).$$
(A1)

Here W(k, k', q) is the interaction vertex and $G_x(k)$ is the exciton Green's function. If $W(k, k', q) = W_0$ is frequency and momentum independent, then \mathcal{T} depends only on the total incoming momentum and energy (\mathbf{K}, ω) and Eq. (A1) is solved immediately:

$$\mathcal{T}(\boldsymbol{K},\omega) = \left[W_0^{-1} - \Pi(\boldsymbol{K},\omega)\right]^{-1}, \qquad (A2)$$

with the two-exciton propagator (at zero temperature)

$$\Pi(\mathbf{K},\omega) = \sum_{\mathbf{q}} \int \frac{d\omega'}{2\pi i} G_x(\mathbf{K}-\mathbf{q},\omega-\omega')G_x(\mathbf{q},\omega').$$
(A3)

Consider first the case of free excitons, where the Green's function is $G_x^0(\mathbf{k}, \omega) = (\omega - \omega_{x,\mathbf{k}})^{-1}$, and the propagator is

$$\Pi_{0}(\boldsymbol{K},\omega) = \sum_{\boldsymbol{q}} \int \frac{d\omega'}{2\pi i} G_{x}^{0}(\boldsymbol{K}-\boldsymbol{q},\omega-\omega') G_{x}^{0}(\boldsymbol{q},\omega')$$
$$= \sum_{\boldsymbol{q}} \frac{1}{\omega - \omega_{x,\boldsymbol{K}-\boldsymbol{q}} - \omega_{x,\boldsymbol{q}}}.$$
(A4)

Since the sum diverges logarithmically at large q, we must impose a momentum cutoff Λ , and we have $\Pi_0(\mathbf{K} = 0, \omega) =$ $-\frac{\dot{m}_x}{4\pi}\ln(-E_\Lambda/\omega)$, with $E_\Lambda = \Lambda^2/2m_x$. We eliminate W_0 by requiring that the biexciton binding energy E_{xx} is the pole of the of the zero-momentum scattering matrix, so $W_0^{-1} =$ $-\frac{m_x}{4\pi}\ln(E_{\Lambda}/E_{xx})$, and the renormalized free exciton scattering matrix takes the form

$$\mathcal{T}_0(\mathbf{K}=0,\omega) = \frac{4\pi/m_x}{\ln\left(-E_{xx}/\omega\right)},\tag{A5}$$

which is the universal result for low-energy scattering in two dimensions [50]. We now consider the case of excitonpolaritons, where the exciton Green's function is

$$G_x(\boldsymbol{k},\omega) = \sum_{s=\pm} \frac{U_{s,\boldsymbol{k}}^2}{\omega - \omega_{s,\boldsymbol{k}}}.$$
 (A6)

The polariton energies $\omega_{\pm,k}$ and Hopfield coefficients $U_{\pm,k}$ are given, respectively, in Eqs. (2) and (3) of the main text. The two-exciton propagator is

$$\Pi(\mathbf{K},\omega) = \sum_{\mathbf{q}} \sum_{s,s'=\pm} \frac{U_{s,\mathbf{K}-\mathbf{q}}^2 U_{s',\mathbf{q}}^2}{\omega - \omega_{s,\mathbf{K}-\mathbf{q}} - \omega_{s',\mathbf{q}}}.$$
 (A7)



FIG. 7. Diagrams for calculating photon self-energy including (a) normal and (b) anomalous contributions. In the normal state only the first two diagrams in (a) are nonvanishing. Single and double lines denote exciton and biexciton Green's functions, respectively. The filled circle denotes the photon-polariton coupling vertex, and the open circle is the photon-polariton-bipolariton coupling vertex. Dotted lines indicate particles going into or out of the condensate.

For photon exciton-polaritons, the cavity mode dispersion takes the form $\omega_{p,k} = \delta + k^2/2m_p$, with $m_p \ll m_x$. We expand $\Pi(\mathbf{K} = 0, \omega)$ in $m_p/m_x \ll 1$ and find

$$\mathcal{T}(\mathbf{K}=0,\omega) = \frac{4\pi}{m_x} \left\{ \ln(-E_{xx}/\omega) + \frac{m_p}{m_x} \frac{\Omega^2}{\omega^2} \left[-1 + \ln[-(m_x/2m_p)\omega/E_{xx}] - \frac{1}{\omega^2 + \Omega^2} (\omega^2 \ln[(\omega_+ + \omega_- - \omega)/E_{xx}] + \Omega^2 \{ -\ln(-2\omega/E_{xx}) + \ln[(2\omega_+ - \omega)/E_{xx}] + \ln[(2\omega_- - \omega)/E_{xx}] \} \right) \right] + \dots \right\}^{-1}.$$
(A8)

Here $\omega_{\pm} = \frac{1}{2}(\delta \pm \sqrt{\delta^2 + \Omega^2})$ are the polariton energies at zero momentum. This shows that $\mathcal{T}(\mathbf{K} = 0, \omega)$ has a pole at the biexciton binding energy, which is slightly shifted from E_{xx} and acquires a small imaginary part in the two-polariton continuum. In addition, there are bipolariton poles at energies below $\omega_s + \omega_{s'}$, but it is apparent from Eq. (A8) that the bipolariton binding energy and spectral weight are exponentially small in the large mass ratio m_x/m_p , in agreement with Eq. (5) of the main text.

For heavy polaritons we take $\omega_{c,k} = \delta$ constant. Then the integral in Eq. (A7) may be evaluated analytically at K = 0 and we find

$$\mathcal{T}(\mathbf{K}=0,\omega) = \frac{-(4\pi/m_x)[(2\delta-\omega)^2 + \Omega^2]}{\Omega^2 \ln[2(\omega_+ + \omega_- - \omega)/E_{xx}] + (2\delta-\omega)^2 \{-\ln[(2\delta-\omega)/E_{xx}] + \ln[(2\omega_+ - \omega)/E_{xx}] + \ln[(2\omega_- - \omega)/E_{xx}]\}},$$
(A9)

The bipolariton energies $\omega_{ss',K}$ are poles of $\mathcal{T}(K, \omega)$. The scattering matrix between two polaritons s and s' with incoming momenta K/2 + k and K/2 - k and outgoing momenta K/2 + k' and K/2 - k' is given by

$$\mathcal{T}_{ss'}(\boldsymbol{k}, \boldsymbol{k}', \boldsymbol{K}, \omega) = U_{s,\boldsymbol{K}/2+\boldsymbol{k}} U_{s',\boldsymbol{K}/2-\boldsymbol{k}} U_{s,\boldsymbol{K}/2+\boldsymbol{k}'} U_{s',\boldsymbol{K}/2-\boldsymbol{k}'} \mathcal{T}(\boldsymbol{K}, \omega). \quad (A10)$$

The polariton Green's functions are $G_s(\mathbf{k}, \omega) = (\omega - \omega_{s,\mathbf{k}})^{-1}$ and we define $\tilde{G}_{ss'}(\mathbf{k}, \mathbf{K}, \omega) = \int \frac{d\omega'}{2\pi i} G_s(\omega - \omega', \mathbf{K/2} - \mathbf{k}) G_{s'}(\omega', \mathbf{K/2} + \mathbf{k})$, or

$$\tilde{G}_{ss'}(\boldsymbol{k},\boldsymbol{K},\omega) = \frac{1}{\omega - \omega_{s,\boldsymbol{K}/2-\boldsymbol{k}} - \omega_{s',\boldsymbol{K}/2+\boldsymbol{k}}}.$$
 (A11)

In the spectral vicinity of the bipolariton resonance, we have [15]

$$\tilde{G}_{ss'}(\boldsymbol{k},\boldsymbol{K},\omega)\mathcal{T}_{ss'}(\boldsymbol{k},\boldsymbol{k}',\boldsymbol{K},\omega)\tilde{G}_{ss'}(\boldsymbol{k}',\boldsymbol{K},\omega)$$

$$\simeq \frac{\psi_{ss'}(\boldsymbol{k},\boldsymbol{K})\psi_{ss'}(\boldsymbol{k}',\boldsymbol{K})}{\omega - \omega_{ss',\boldsymbol{K}}},$$
(A12)

which defines the bipolariton wave function $\psi_{ss'}(\boldsymbol{k}, \boldsymbol{K})$. In our approximation

$$\psi_{ss'}(\boldsymbol{k}, \boldsymbol{K}) = \mathcal{N}_{\boldsymbol{K}} \frac{U_{s, \boldsymbol{K}/2-\boldsymbol{k}} U_{s', \boldsymbol{K}/2+\boldsymbol{k}}}{\omega_{ss', \boldsymbol{K}} - \omega_{s, \boldsymbol{K}/2-\boldsymbol{k}} - \omega_{s', \boldsymbol{K}/2+\boldsymbol{k}}}, \quad (A13)$$

with \mathcal{N}_{K} a normalization factor such that $\sum_{k} |\psi_{ss'}(k, K)|^{2} = 1$. The bipolariton wave function is used to compute the optical spectra, see Appendix B.

APPENDIX B: CALCULATION OF ABSORPTION AND LUMINESCENCE SPECTRA

Two processes contribute to the absorption and luminescence spectra: direct coupling of polaritons to photons and coupling of a bipolariton to a photon and polariton [51,52]. The relevant diagrams for calculating photon self-energy are shown in Fig. 7. The filled circle denotes the photon-polariton coupling vertex. In terms of the photon-exciton coupling μ_x , it is given by

$$\mu_{1,\boldsymbol{q}} = U_{-,\boldsymbol{q}}\mu_{\boldsymbol{x}},\tag{B1}$$

where $U_{-,q}$ is the Hopfield coefficient. The open circle denotes the photon-polariton-bipolariton coupling vertex, given by

$$\mu_{12,k,q} = U_{-,k} U_{-,q+k}^2 \psi((q-k)/2, q+k) \mu_x, \qquad (B2)$$

with ψ the bipolariton wave function defined by Eq. (A12).

1. Normal state

In the normal state, only the first two diagrams of Fig. 7 contribute. The Matsubara Green's functions are

$$G_i(\boldsymbol{k}, i\omega_n) = \frac{1}{i\omega_n - \tilde{\omega}_{i,\boldsymbol{k}}},$$
(B3)

where i = 1, 2 for the polariton and bipolariton, respectively. Here $\tilde{\omega}_{1,k}$ and $\tilde{\omega}_{2,k}$ are the polariton and bipolariton energies shifted by interaction (see Sec. II). The photon self-energy is

$$\Pi(\boldsymbol{q}, i\omega_n) = |\mu_{1,\boldsymbol{q}}|^2 G_1(\boldsymbol{q}, i\omega_n) - \sum_{\boldsymbol{k}} \sum_{i\omega'_n} |\mu_{12,\boldsymbol{k},\boldsymbol{q}}|^2 G_1(\boldsymbol{k}, i\omega'_n)$$
$$\times G_2(\boldsymbol{q} + \boldsymbol{k}, i\omega_n + i\omega'_n). \tag{B4}$$

Evaluating the sum over $i\omega'_n$, we obtain the absorption spectrum:

$$A(\boldsymbol{q},\omega) = -2\mathrm{Im}\,\Pi(\boldsymbol{q},\omega-\mu) = |\mu_{1,\boldsymbol{q}}|^2 2\pi \delta(\omega-\mu-\tilde{\omega}_{1,\boldsymbol{q}}) + \sum_{\boldsymbol{k}} |\mu_{12,\boldsymbol{k},\boldsymbol{q}}|^2 [n_B(\tilde{\omega}_{1,\boldsymbol{k}}) - n_B(\tilde{\omega}_{2,\boldsymbol{q}+\boldsymbol{k}})] \times 2\pi \delta(\omega-\mu+\tilde{\omega}_{1,\boldsymbol{k}}-\tilde{\omega}_{2,\boldsymbol{q}+\boldsymbol{k}}).$$
(B5)

with

$$n_B(\omega) = (e^{\beta\omega} - 1)^{-1}$$
(B6)

the Bose distribution function, β being the inverse temperature. The luminescence spectrum is [51]

$$I(\boldsymbol{q},\omega) = n_B(\omega - \mu)A(\boldsymbol{q},\omega)$$

= $|\mu_{1,\boldsymbol{q}}|^2 n_B(\tilde{\omega}_{1,\boldsymbol{q}})2\pi\delta(\omega - \mu - \tilde{\omega}_{1,\boldsymbol{q}})$
+ $\sum_k |\mu_{12,k,\boldsymbol{q}}|^2 [1 + n_B(\tilde{\omega}_{1,k})]$
 $\times n_B(\tilde{\omega}_{2,\boldsymbol{q}+k})2\pi\delta(\omega - \mu + \tilde{\omega}_{1,k} - \tilde{\omega}_{2,\boldsymbol{q}+k}).$ (B7)

For q = 0 and far from the Feshbach resonance, this spectrum consists of a sharp peak at $\omega = \mu + g_{11}n_1 + g_{12}n_2$ and a continuum below the energy $\omega = \mu - E_{--} + (g_{11} - g_{12})n_1 + (g_{12} - g_{22})n_2$, similar to Eq. (13). In principle, this spectrum is qualitatively different from what is shown in Fig. 5(b), and so it can be used to identify the condensation transition.

2. Condensed state

In the condensed state, there are two modes: a gapless mode $E_{0,k}$ and gapped mode $E_{1,k}$. The Matsubara Green's functions may be written

$$G_i(\boldsymbol{k}, i\omega_n) = \sum_{\sigma=0,1} \sum_{s=\pm} \frac{s(u_{i,k}^{\sigma,s})^2}{i\omega_n - sE_k^{\sigma}}.$$
 (B8)

Here the $u_{i,k}^{\sigma,s}$ are Bogoliubov coefficients that satisfy $\sum_{\sigma,s} s(u_{i,k}^{\sigma,s})^2 = 1$. We also need the anomalous Green's functions

$$F_{ij}(\boldsymbol{k}, i\omega_n) = -\sum_{\sigma=0,1} \sum_{s=\pm} \frac{s u_{i,\boldsymbol{k}}^{\sigma,s} u_{j,\boldsymbol{k}}^{\sigma,-s}}{i\omega_n - s E_{\boldsymbol{k}}^{\sigma}}.$$
 (B9)

Expressions for the energies and Bogoliubov coefficients in terms of the parameters in the effective Hamiltonian of Eq. (6) are given in Ref. [26]. The photon self-energy is

$$\Pi(\boldsymbol{q}, i\omega_n) = |\mu_{1,\boldsymbol{q}}|^2 G_1(\boldsymbol{q}, i\omega_n) - \sum_{\boldsymbol{k}} \sum_{i\omega'_n} |\mu_{12,\boldsymbol{k},\boldsymbol{q}}|^2 G_1(\boldsymbol{k}, i\omega'_n) G_2(\boldsymbol{q} + \boldsymbol{k}, i\omega_n + i\omega'_n) - \sum_{\boldsymbol{k}} \sum_{i\omega'_n} |\mu_{12,\boldsymbol{k},\boldsymbol{q}}|^2 F_{12}(\boldsymbol{k}, i\omega'_n) F_{21}^*(\boldsymbol{q} + \boldsymbol{k}, i\omega_n + i\omega'_n) + |\mu_{12,\boldsymbol{0},\boldsymbol{q}}|^2 [n_1 G_2(\boldsymbol{q}, i\omega_n) + n_2 G_1(\boldsymbol{q}, -i\omega_n) - 2\sqrt{n_1 n_2} F_{12}(\boldsymbol{q}, i\omega_n)].$$
(B10)

We find for the absorption spectrum

$$A(\boldsymbol{q},\omega) = -2\mathrm{Im}\,\Pi(\boldsymbol{q},\omega-\mu) = \sum_{n=1}^{3} A_{n}(\boldsymbol{q},\omega),$$

$$A_{1}(\boldsymbol{q},\omega) = -|\mu_{1,\boldsymbol{q}}|^{2} \sum_{\sigma,s} s(u_{1,\boldsymbol{q}}^{\sigma,-s})^{2} 2\pi \delta(\omega-\mu+sE_{\boldsymbol{q}}^{\sigma}),$$

$$A_{2}(\boldsymbol{q},\omega) = \sum_{\sigma,\sigma'} \sum_{s,s'} \sum_{\boldsymbol{k}} |\mu_{12,\boldsymbol{k},\boldsymbol{q}}|^{2} ss' \Big[(u_{1,\boldsymbol{k}}^{\sigma,s})^{2} (u_{2,\boldsymbol{q}+\boldsymbol{k}}^{\sigma',s'})^{2} + u_{1,\boldsymbol{k}}^{\sigma,s} u_{2,\boldsymbol{k}}^{\sigma',-s'} u_{1,\boldsymbol{q}+\boldsymbol{k}}^{\sigma',s'} \Big] \Big[n_{B}(sE_{\boldsymbol{k}}^{\sigma}) - n_{B}(s'E_{\boldsymbol{q}+\boldsymbol{k}}^{\sigma'}) \Big]$$

$$\times 2\pi \delta(\omega-\mu+sE_{\boldsymbol{k}}^{\sigma}-s'E_{\boldsymbol{q}+\boldsymbol{k}}^{\sigma'}),$$

$$A_{3}(\boldsymbol{q},\omega) = -|\mu_{12,\boldsymbol{0},\boldsymbol{q}}|^{2} \sum_{\sigma,s} s \Big[n_{1}(u_{2,\boldsymbol{q}}^{\sigma,-s})^{2} + n_{2}(u_{1,\boldsymbol{q}}^{\sigma,s})^{2} - 2\sqrt{n_{1}n_{2}}u_{1,\boldsymbol{q}}^{\sigma,s} u_{2,\boldsymbol{q}}^{\sigma,-s} \Big] 2\pi \delta(\omega-\mu+sE_{\boldsymbol{q}}^{\sigma}). \tag{B11}$$

The luminescence spectrum is

$$I(\boldsymbol{q},\omega) = I_{0}(\boldsymbol{q},\omega) + n_{B}(\omega-\mu)A(\boldsymbol{q},\omega) = \sum_{n=0}^{3} I_{n}(\boldsymbol{q},\omega),$$

$$I_{0}(\boldsymbol{q},\omega) = \left(|\mu_{1,0}|^{2}n_{1} + |\mu_{12,0,0}|^{2}n_{1}n_{2}\right)(2\pi)^{3}\delta(\omega-\mu)\delta(\boldsymbol{q}),$$

$$I_{1}(\boldsymbol{q},\omega) = |\mu_{1,q}|^{2}\sum_{\sigma,s} s\left(u_{1,q}^{\sigma,-s}\right)^{2} \left[1 + n_{B}\left(sE_{q}^{\sigma}\right)\right]2\pi\delta(\omega-\mu+sE_{q}^{\sigma}),$$

$$I_{2}(\boldsymbol{q},\omega) = \sum_{\sigma,\sigma'} \sum_{s,s'} \sum_{k} |\mu_{12,k,q}|^{2}ss' \left[\left(u_{1,k}^{\sigma,s}\right)^{2}\left(u_{2,q+k}^{\sigma',-s'}\right)^{2} + u_{1,k}^{\sigma,s}u_{2,k}^{\sigma,-s}u_{1,q+k}^{\sigma',-s'}\right] \left[1 + n_{B}\left(sE_{k}^{\sigma}\right)\right] \left[1 + n_{B}\left(s'E_{q+k}^{i\sigma}\right)\right]$$

$$\times 2\pi\delta\left(\omega-\mu+sE_{k}^{\sigma}+s'E_{q+k}^{\sigma'}\right),$$

$$I_{3}(\boldsymbol{q},\omega) = |\mu_{12,0,q}|^{2}\sum_{\sigma,s} s\left[n_{1}\left(u_{2,q}^{\sigma,-s}\right)^{2} + n_{2}\left(u_{1,q}^{\sigma,s}\right)^{2} - 2\sqrt{n_{1}n_{2}}u_{1,q}^{\sigma,s}u_{2,q}^{\sigma,-s}\right] \left[1 + n_{B}\left(sE_{q}^{\sigma}\right)\right] 2\pi\delta\left(\omega-\mu+sE_{q}^{\sigma}\right).$$
(B12)

The first term is due to coherent spontaneous emission of photons from the condensate, and is not present in the absorption spectrum. Actually, when $q \rightarrow 0$ the weight of the peak at $\omega = \mu$ diverges due to the divergence of the Bogoliubov coefficients. We deal with this by integrating over a finite range of q, $q < q_c$. Physically, the cutoff q_c represents a typical momentum due to scattering by inhomogeneities, phonons, etc., or the inverse trap size in the case of confined polaritons (Sec. III).

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