Giant magneto-optical Schäfer-Hubert effect in the two-dimensional van der Waals antiferromagnets MPS_3 (M = Mn, Fe, Ni)

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The recent discovery of long-range magnetic order in atomically thin films has triggered particular interest in two-dimensional (2D) van der Waals (vdW) magnetic materials. In this paper, we perform a systematic theoretical study of the magneto-optical Schäfer-Hubert effect (MOSHE) in 2D vdW antiferromagnetic MPS_3 (M = Mn, Fe, Ni) with multifold intralayer and interlayer magnetic orders. The formula for evaluating the MOSHE in 2D magnets is derived by considering the influence of a nonmagnetic substrate. The MOSHE of monolayer and bilayer MPS_3 is very large (>2°), originating from the strong anisotropy of in-plane optical conductivity. The Schäfer-Hubert rotation angles are surprisingly insensitive to the orientations of the Néel vector, while the Schäfer-Hubert ellipticities are identified to be a good criterion to distinguish different interlayer magnetic orders. Our work establishes a theoretical framework for exploring novel 2D vdW magnets and facilitates the promising applications of the 2D MPS_3 family in antiferromagnetic nanophotonic devices.

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I. INTRODUCTION

Two-dimensional (2D) van der Waals (vdW) magnetic materials have attracted emerging attention since the discovery of intrinsically long-range ferromagnetic (FM) order in $Cr_2Ge_2Te_6$ and CrI_3 atomic layers [1,2]. The highly tunable magnetism and other exciting physical properties by electric gating [3] and strain engineering [4,5] offer them a promising potential for applications in magnetic sensor, storage, and spintronics. Magneto-optical spectroscopy is a powerful noncontact technique for investigating 2D magnetic materials. For 2D ferromagnets, the magneto-optical Kerr effect (MOKE) signals solid evidence of long-range FM order even down to the monolayer limit [2]. Furthermore, first-principles calculations of MOKE in thin films [6,7] provide a complementary avenue to characterize 2D FM materials [8-11]. For 2D antiferromagnets that have zero net magnetization, the MOKE as a first-order effect is vanishing, and therefore the commonly used magneto-optical techniques are based on second-order effects [12–14]. One option is to probe the difference in absorption or reflectivity for linearly polarized lights parallel and perpendicular to the Néel vector, which is known as magnetic linear dichroism (MLD). Another option is to probe the polarization rotation upon transmission and reflection, which is called the magneto-optical Voigt effect [15] and magnetooptical Schäfer-Hubert effect (MOSHE) [16], respectively. Since the second-order magneto-optical effects in magnetic materials are usually very weak, the characterization of 2D antiferromagnetic (AFM) order has long been considered extremely challenging.

Transition metal thiophosphates MPS_3 (M = Mn, Fe, Ni) are a representative family of 2D vdW materials that host multifold intrinsically intralayer AFM orders [17-19]. In a recent experiment, Zhang et al. [20] observed large MLD in FePS₃ with zigzag-AFM order. The large magneto-optical signals enable the detection of 2D AFM domain orientations [20,21] and the study of ultrafast spin dynamics [22]. Subsequently, the tuning of MLD in FePS₃ was realized by coupling with an optical cavity [23], and the MLD at specific wavelengths can even be enhanced to a near-unity (100%) value. Such an optically anisotropic 2D magnetic material is desirable for achieving densely integrated polarization-selective devices. To date, most of the reports on large linear dichroism and its tuning for 2D materials have been limited to those with in-plane anisotropic crystal structures, such as black phosphorus [24,25] and GeSe [26]. By contrast, anisotropic 2D magnetic materials are more promising for the fast field-effect control since the magnetic orders are sensitive to external stimuli, e.g., magnetic [27] and strain [28] fields. These recent advances call for an exploration of more excellent 2D AFM magneto-optical materials; however, theoretical studies on the second-order magneto-optical effects in thin films remain absent yet.

In this work, we systematically investigate a representative second-order magneto-optical effect, MOSHE, in 2D vdW AFM MPS₃ using first-principles calculation together with magnetic group analysis. A theoretical formula for

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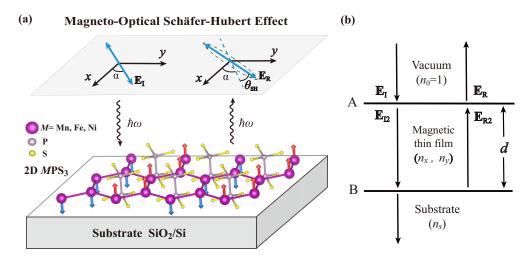


FIG. 1. (a) Schematic illustration of magneto-optical Schäfer-Hubert effect emerged in 2D antiferromagnets MPS_3 (M = Mn, Fe, Ni) prepared on SiO₂/Si substrate. The incident light is linearly polarized with the electric field (\mathbf{E}_I) oriented at an angle of α from the optically anisotropic axis (here, *x* axis). The reflected light becomes elliptically polarized and the polarization plane (\mathbf{E}_R) deflects an angle of θ_{SH} with respect to incident light (\mathbf{E}_I). (b) Optical paths in a magnetic thin film placed on an optically isotropic nonmagnetic substrate. Refractive indices (n_0, n_x, n_y, n_s) in each region and the electric fields ($\mathbf{E}_I, \mathbf{E}_{I2}, \mathbf{E}_R, \mathbf{E}_{R2}$) at the interface A are labeled, and *d* denotes the thickness of magnetic thin film.

evaluating the MOSHE in 2D magnetic materials placed on a nonmagnetic substrate is derived. The MOSHE in FePS₃ and NiPS₃ with the zigzag-AFM order is close to or even exceeds the magnitudes of first-order magneto-optical effects in conventional ferromagnets, especially the Schäfer-Hubert (SH) rotation angle in bilayer NiPS₃ records up to 2.4° . We also find that the MOSHE is insensitive to the magnetization direction, and the SH ellipticity can be used to identify interlayer magnetic structures. Our work deepens the understanding of MOSHE in 2D antiferromagnets and facilitates further exploration of novel AFM magneto-optical devices.

II. RESULTS AND DISCUSSION

When a linearly polarized light normally shines (e.g., along the *z* axis) on a thin film with in-plane magnetic anisotropy, the light propagating in the magnetic thin film can be decomposed into two polarized components along orthogonal anisotropic axes with different refractive indices (n_x, n_y) and reflectivity (r_x, r_y) . The reflected light would become elliptically polarized accompanied by a rotation of the polarization plane with respect to the incident light, namely the MOSHE [Fig. 1(a)]. If the electric field of incident light (\mathbf{E}_I) is placed at an angle of $\alpha = 45^{\circ}$ from the *x* axis, the SH rotation angle (θ_{SH}) and ellipticity (ψ_{SH}) reach up to their maximums, given by [29]

$$\theta_{\rm SH} = \frac{1}{2} \operatorname{atan} \left(\frac{2 \operatorname{Re} \chi}{1 - |\chi|^2} \right) - \frac{\pi}{4},$$

$$\psi_{\rm SH} = \frac{1}{2} \operatorname{asin} \left(\frac{2 \operatorname{Im} \chi}{1 + |\chi|^2} \right), \tag{1}$$

where $\chi = r_y/r_x$. The reflectivity of a magnetic thin film at the interface A [Fig. 1(b)] can be written as

$$r_{x(y)} = \frac{n_0 - \tilde{n}_{x(y)}}{n_0 + \tilde{n}_{x(y)}}.$$
(2)

Here, $n_0 = 1$ is the refractive index of vacuum, and $\tilde{n}_{x(y)}$ is the effective refractive index of a magnetic thin film by considering the influence of its substrate,

$$\tilde{n}_{x(y)} = \frac{1 - r'_{x(y)}\beta_{x(y)}}{1 + r'_{x(y)}\beta_{x(y)}}n_{x(y)},\tag{3}$$

in which $\beta_{x(y)} = \exp(2i\omega dn_{x(y)}/c)$ with the light frequency ω , light speed *c*, and film thickness *d*. The reflectivity of the substrate at the interface B is $r'_{x(y)} = (n_{x(y)} - n_s)/(n_{x(y)} + n_s)$ and n_s is the refractive index of substrate. Plugging Eqs. (2) and (3) into Eq. (1), the complex SH angle can be recast as

$$\theta_{\rm SH} + i\psi_{\rm SH} \approx \frac{\tilde{n}_x - \tilde{n}_y}{1 - \tilde{n}_x \tilde{n}_y}.$$
(4)

The above equation is appropriate for a relatively thick film placed on a substrate. In the case of bulk materials, the effective refractive index $\tilde{n}_{x(y)}$ should be simply replaced by $n_{x(y)}$, disregarding the substrate effect entirely.

For very thin films whose thicknesses are far less than the wavelength of visible light (λ), the effective refractive index can be approximated to $\tilde{n}_{x(y)} \approx n_s - i\frac{2\pi d}{\lambda}(n_{x(y)}^2 - n_s^2)$. In the case of conventional MOSHE induced by in-plane magnetization (e.g., along the *x* axis), the refractive indices by solving the Fresnel equation are given by $n_x = \sqrt{\epsilon_{xx}}$, $n_y = \sqrt{\epsilon_{yy} + \epsilon_{yz}^2/\epsilon_{zz}}$, in which $\epsilon_{\mu\nu}$ with $\mu, \nu \in \{x, y, z\}$ is the permittivity tensor. Then, the complex SH angle can be simplified to

$$\theta_{\rm SH} + i\psi_{\rm SH} \approx \frac{i\omega d}{c(n_s^2 - 1)} \left(\epsilon_{xx} - \epsilon_{yy} - \frac{\epsilon_{yz}^2}{\epsilon_{zz}}\right).$$
(5)

We find that the complex SH angle can be related to the complex Voigt angle [30] via

$$\theta_{\rm SH} + i\psi_{\rm SH} = \frac{2(n_x + n_y)}{1 - n_s^2} (\theta_{\rm V} - i\psi_{\rm V}),$$
 (6)

where θ_V and ψ_V are the Voigt rotation angle and ellipticity, respectively. If the substrate has a relatively small refractive index $(n_s \rightarrow 1)$, the SH angle will be much larger than the Voigt angle, indicating that the optical detection upon reflection is more suitable than upon transmission for studying the second-order magneto-optical effects of magnetic thin films.

The complex SH angle [see Eq. (5)] can also be written in terms of optical conductivity using the relationship between permittivity and optical conductivity, given by $\epsilon_{\mu\nu} = \delta_{\mu\nu} + \frac{4\pi i}{\omega}\sigma_{\mu\nu}$. The off-diagonal elements of the optical conductivity containing the *z* component (e.g., σ_{yz}) have to be zero due to the 2D nature of our considered systems. This can be read from Eq. (A1) since the quenched electron velocity along the *z* direction ($\hat{v}_z = 0$) leads to the vanishing σ_{yz} and σ_{zx} . Therefore, the complex SH angle is simply expressed as

$$\theta_{\rm SH} + i\psi_{\rm SH} \approx \frac{4\pi d}{c(n_{\rm s}^2 - 1)}(\sigma_{yy} - \sigma_{xx}),\tag{7}$$

which is the formula implemented in our first-principles calculations. The 2D vdW magnetic materials are often grown on transparent substrates, such as SiO₂, whose refractive index n_s is a real number. In this case, the SH rotation angle and ellipticity are determined by the real and imaginary parts of conductivity anisotropy (i.e., $\sigma_{yy} - \sigma_{xx}$), respectively. On account of this relationship, the conductivity anisotropy can be accurately measured by MOSHE spectroscopy.

For monolayer MPS_3 , the transition metal atoms M form a flat honeycomb lattice and a bipyramid of P_2S_6 ligand locates at the center of hexagon [Figs. 2(a) and 2(b)]. If removing the magnetic orders, monolayer MPS_3 is in-plane isotropic due to its crystallographic point group of D_{3d} . Nevertheless, the honeycomb lattice can host a variety of magnetic orders, including the FM state as well as Néel-, zigzag-, and stripy-AFM states [Figs. 2(c)–2(e)] [31], depending on the relative strength of intralayer first, second, and third nearest-neighbor exchange interactions. MnPS₃ displays the Néel-AFM order with the out-of-plane (z axis) magnetic easy axis [17]. FePS₃ and NiPS₃ display the zigzag-AFM order with the out-ofplane (z axis) [18] and in-plane (x axis) [32] magnetization,

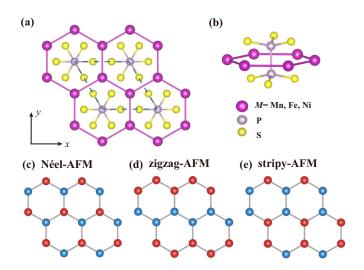


FIG. 2. (a), (b) Top and side views of monolayer MPS_3 . Blue dashed lines draw out the primitive cell of nonmagnetic state. (c)–(e) The Néel-, zigzag-, and stripy-antiferromagnetic orders on a honeycomb lattice. Red and blue spheres represent the M atoms with opposite directions of spin magnetic moments.

respectively. The long-range AFM orders of the exfoliated atomic layers persist down to bilayer or even monolayer limit, and their magnetic critical temperatures are nearly independent of thickness. Moreover, the magnetization for Néel- and zigzag-AFM states can be tuned between the out-of-plane and in-plane directions via atomic substitution [33], and the FM state was predicted to be their ground states under sufficient large carrier density [34].

Before practically calculating MOSHE, we conduct symmetry analysis to evaluate which magnetic order breaks the in-plane optically isotropy of monolayer MPS_3 . The magnetic space groups computed by the ISOTROPY code [35] are listed in Table I, in which the shapes of optical conductivity tensors are identified by the SYMMETR code [36,37]. As expected, all of the magnetic orders with the magnetization along the *x* axis are in-plane anisotropic, which allows the MOSHE.

TABLE I. Magnetic space groups of monolayer MPS_3 with different magnetic orders. The magnetization directions are labeled in brackets. The symbol \checkmark (×) indicates the in-plane optically anisotropy (isotropy). The dipole selection rules at some high-symmetry points (e.g., Γ and K) are listed.

Magnetic orders	Magnetic space group	In-plane anisotropy $(\sigma_{xx} \neq \sigma_{yy})$	Dipole selection rules $(\mathbf{E} \perp z)$
$\overline{\mathrm{FM}}(x)$	C2'/m'	\checkmark	$\Gamma_2^+ \leftrightarrow \Gamma_2^-$
			$\Gamma_4^+ \leftrightarrow \Gamma_5^-, \Gamma_6^- K_4 \leftrightarrow K_5$
FM (<i>z</i>)	$P\overline{3}1m'$	Х	$\Gamma_5^+ \leftrightarrow \Gamma_4^-, \Gamma_6^- \ \ \mathrm{K}_4 \leftrightarrow \ \ \mathrm{K}_6$
			$\Gamma_6^+ \leftrightarrow \Gamma_4^-, \Gamma_5^- \ \ \mathrm{K}_5 \leftrightarrow \ \mathrm{K}_6$
Néel-AFM (x)	C2'/m	\checkmark	$\Gamma_3\Gamma_4 \leftrightarrow \Gamma_3\Gamma_4$
Néel-AFM (z)	$P\overline{3}'1m$	×	$\Gamma_4 \leftrightarrow \Gamma_4 K_4 \leftrightarrow K_4$
			$\Gamma_4 \leftrightarrow \Gamma_5 \Gamma_6 \qquad K_4 \leftrightarrow K_5 K_6$
zigzag-AFM (x, z)	$P_c 2_1/m$	\checkmark	$\Gamma_3^+\Gamma_4^+\leftrightarrow\Gamma_3^-\Gamma_4^-$
stripy-AFM (x, z)	$P_a 2_1/c$	\checkmark	$\Gamma_3^+\Gamma_4^+ \leftrightarrow \Gamma_3^-\Gamma_4^-$

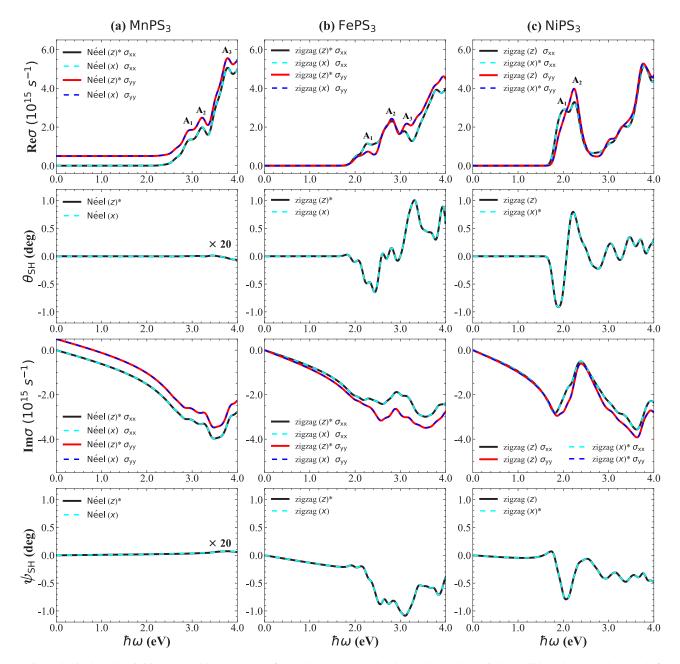


FIG. 3. Optical conductivities and MOSHE spectra of monolayer (a) MnPS₃, (b) FePS₃, and (c) NiPS₃ on SiO₂ substrate. The panels from top to bottom show the real part of optical conductivity (Re σ), SH rotation angle (θ_{SH}), imaginary part of optical conductivity (Im σ), and SH ellipticity (ψ_{SH}), respectively. The magnetization direction of each magnetic order is indicated in brackets, and an asterisk labels the ground state. The Re σ_{yy} and Im σ_{yy} of MnPS₃ are moved upward by $0.5 \times 10^{15} \text{ s}^{-1}$ for a clear observation, and the θ_{SH} and ψ_{SH} of MnPS₃ are multiplied by a factor of 20. A₁, A₂, and A₃ mark several absorption peaks of Re σ in the low-energy range.

For FM and Néel-AFM orders with the spins along the z axis, the in-plane isotropy is preserved by the threefold rotational symmetry in magnetic space groups of $P\overline{3}1m'$ and $P\overline{3}'1m$, respectively. The magnetic space groups of zigzagand stripy-AFM orders with the magnetization along the z axis are the same as that along the x axis, such that the z axis magnetization is also in-plane anisotropic and may also lead to the MOSHE. According to the mirror symmetry M_y in the zigzag-AFM order, the orthogonal anisotropic axes are determined to be the x and y axes as shown in Fig. 2. Figure 3 plots the calculated optical conductivities and MOSHE spectra of monolayer AFM MPS₃. We first discuss the results of each material on its magnetic ground state. For MnPS₃ with the *z* axis Néel-AFM order [Fig. 3(a)], the spectrum of σ_{xx} is identical to that of σ_{yy} , which is governed by the in-plane optical isotropy, and the resulting SH rotation angle (θ_{SH}) and SH ellipticity (ψ_{SH}) are negligibly small. The absorptive parts of the optical conductivity tensor, $\text{Re}\sigma_{xx}$ and $\text{Re}\sigma_{yy}$, are determined by the symmetry-allowed dipole selection rules listed in Table I, from which one can analyze

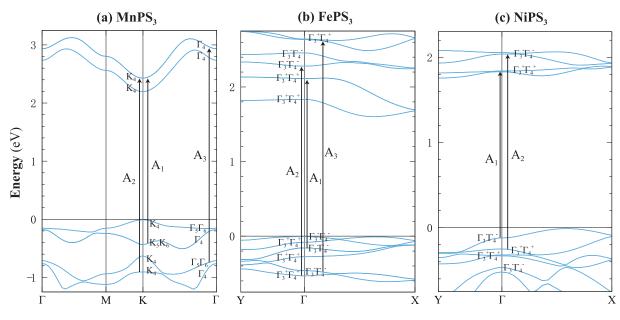


FIG. 4. Relativistic band structures of monolayer (a) MnPS₃ with the *z* axis Néel-AFM order, (b) FePS₃ with the *z* axis zigzag-AFM order, and (c) NiPS₃ with the *x* axis zigzag-AFM order. The irreducible representations of relevant bands at the Γ and K points are labeled. The principal interband transitions A₁, A₂, and A₃ are indicated by arrows, corresponding to the peaks of Re σ_{xx} and Re σ_{yy} in Fig. 3.

the origination of main peaks in conductivity spectra. For example, the A_1 and A_2 peaks at the energies of 2.9 eV and 3.2 eV originate from the interband transitions $K_5K_6 \rightarrow K_4$ and $K_4 \rightarrow K_4$ at the K point, respectively, and the A₃ peak at the energy of 3.7 eV originates from the interband transition $\Gamma_5\Gamma_6 \rightarrow \Gamma_4$ at the Γ point, as depicted in Fig. 4(a). For FePS₃ with the z axis zigzag-AFM order [Fig. 3(b)], one can discern a clear anisotropy in the real and imaginary parts of the optical conductivity above the absorption edge ($\sim 1.8 \text{ eV}$). The spectra of $\text{Re}\sigma_{xx}$ and $\text{Re}\sigma_{yy}$ feature three peaks of A₁, A₂, and A₃ at the energies of 2.3 eV, 2.8 eV, and 3.1 eV, respectively, which come from the interband transitions between the $\Gamma_3^+\Gamma_4^+$ and $\Gamma_3^-\Gamma_4^-$ states at the Γ point [Fig. 4(b)]. The obvious difference in values between $\operatorname{Re}\sigma_{xx}$ and $\operatorname{Re}\sigma_{yy}$ around the A₁ and A₃ peaks generate the maximal SH rotation angles of -0.7° at 2.4 eV and of 1.0° at 3.3 eV, respectively. The SH ellipticity is always negative and reaches up to -1.1° at 3.1 eV. For NiPS₃ with the x axis zigzag-AFM order [Fig. 3(c)], the real part of optical conductivity resembles the experimental detection of its bulk crystal [38]. Both $\text{Re}\sigma_{xx}$ and $\text{Re}\sigma_{yy}$ spectra show the A₂ peak at 2.3 eV due to the interband transition $\Gamma_3^+\Gamma_4^+ \rightarrow$ $\Gamma_3^-\Gamma_4^-$, while an additional peak A₁ appears at 2.0 eV for $\operatorname{Re}\sigma_{xx}$ which is related to the transition from the $\Gamma_3^-\Gamma_4^-$ state (highest valance band) to the $\Gamma_3^+\Gamma_4^+$ state (lowest conduction band) at the Γ point [Fig. 4(c)]. In the energy range of 1.7 \sim 2.5 eV, the significant anisotropy of optical conductivity has to result in large SH rotation angles, e.g., -0.9° at 1.9 eV and 0.8° at 2.2 eV. The corresponding SH ellipticity is also obviously large with a peak of -0.8° at 2.1 eV.

Of particular interest here is that the optical conductivity spectra are almost not changed when the magnetization direction changes from the *z* axis to the *x* axis or vice versa (Fig. 3). This is very similar to the cases of three-dimensional noncollinear AFM Mn_3X (X = Rh, Ir, Pt) [39] and 2D vdW FM Fe_nGeTe₂ (n = 3, 4, 5) [40]. It can be easily understood as the longitudinal optical conductivities (σ_{xx} and σ_{yy}) are closely related to the joint density of states and interband transition probability [41] which are basically not influenced when the angle between adjacent spins keeps fixed. It follows that the SH spectra of MPS₃ are insensitive to magnetization direction, e.g., FePS₃ and NiPS₃ with the z- and x-axis zigzag-AFM orders [Figs. 3(b) and 3(c)]. In the case of MnPS₃ with the x-axis Néel-AFM order, the optical conductivities are identical to the results of z-axis Néel-AFM order, such that the SH rotation angel and ellipticity are also negligibly small [Fig. 3(a)], even though the appearance of MOSHE with the x axis Néel-AFM order is allowed by symmetry (Table I). Similarly, since the z-axis FM order exhibits in-plane isotropy, the MOSHE in all three materials with the x-axis FM order is rather small (see Fig. S1 in Supplemental Material [42]); e.g., the largest SH rotation angle (appearing in FePS₃) is only 0.05° . Therefore, we suggest that it is more likely to observe large secondorder magneto-optical effects in AFM materials that exhibit in-plane anisotropy when the spins are out-of-plane oriented, such as the MPS₃ family with the zigzag-AFM and stripy-AFM [Figs. 3(b) and 3(c)] orders (see supplemental Fig. S2).

Next, we move on to discuss the MOSHE in bilayer FePS₃ and NiPS₃ on their magnetic ground states. For FePS₃, two types of interlayer magnetic structures have long been reported. One is the zigzag-AFM chain along the *x* axis (type A) with AFM interlayer coupling [Fig. 5(a)] [43], while the other one is the zigzag-AFM chain along the x' axis (type B) with FM interlayer coupling [Fig. 5(b)] [44]. Recently, the coexistence of the two types of magnetic structures in multilayer FePS₃ has been confirmed by combining MLD and second-harmonic generation measurements [21]. For NiPS₃ powder and single crystals, as far as we know, only the type-A zigzag chain with FM interlayer coupling has been reported

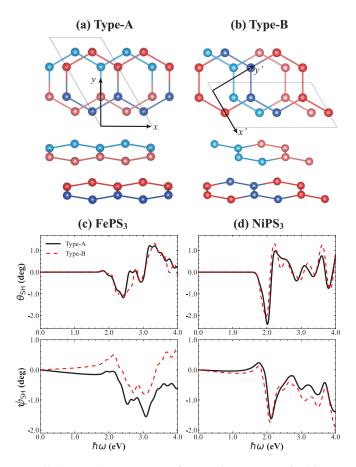


FIG. 5. (a), (b) Two types of magnetic structures for bilayer MPS_3 with the zigzag-AFM chains along the *x* and *x'* axes. Bright (dark) red and blue spheres denote the *M* atoms on bottom (top) layer with opposite spin magnetic moments, whereas P and S atoms are not shown. The solid black lines draw out the 2D primitive cell. (c), (d) Magneto-optical Schäfer-Hubert spectra (θ_{SH} and ψ_{SH}) of bilayer FePS₃ and NiPS₃ with the type-A and type-B magnetic structures.

[45,46]. We speculate that the type-B structure may also exist in bilayer and multilayer NiPS₃.

Here we consider both FM and AFM interlayer coupling for type-A and type-B zigzag chains in bilayer FePS₃ and NiPS₃. The optical conductivities are not shown because they retain the overall trend in monolayers [Figs. 3(b) and 3(c)] with slight change in magnitudes due to the weak interlayer vdW interactions. The calculated SH spectra are plotted in Figs. 5(c) and 5(d), in which the interlayer FM and AFM coupling are not labeled since their spectra are identical to each other. One can observe that for both FePS₃ and NiPS₃, the profiles of SH rotation angles for two types of zigzag chains resemble each other [top panels of Figs. 5(c) and 5(d)]. Moreover, the SH rotation angles for monolayers and bilayers are also similar in the sense that their peaks appear at almost the same photon energy. The calculated SH rotation angles of bilayer FePS₃ (NiPS₃) are surprisingly large recording to -1.2° at 2.4 eV and 1.2° at 3.2 eV (-2.4° at 2.0 eV and 1.0° at 2.2 eV). In contrast to the SH rotation angles, the SH ellipticities are highly correlated to the zigzag chain structures [bottom panels of Figs. 5(c) and 5(d)]. The ellipticity spectra of bilayer FePS₃ with type-A and type-B structures show a

striking contrast in a wide range of photon energy, in particular at 2.4 eV and 3.4 eV, where ψ_{SH} for type-B structure is zero. As well, there is significant difference between the type-A and type-B structures of bilayer NiPS₃ from 3.0 eV to 3.7 eV. We suggest that the dramatic features in SH ellipticity can be used to distinguish the magnetic structures of bilayer *MPS*₃.

III. CONCLUSION AND REMARKS

In summary, our work establishes a simple theoretical framework for studying the magneto-optical Schäfer-Hubert effect in 2D magnetic materials using first-principles calculations, and also proposes second-order magneto-optical spectroscopy to be a powerful technique for accurately detecting the in-plane anisotropy in various magnetic structures. The calculated results demonstrate that monolayer FePS₃ and NiPS₃ with the zigzag antiferromagnetic order exhibit large Schäfer-Hubert angles (up to 1°) in visible light and near ultraviolet range. We further find that the Schäfer-Hubert effect is interestingly insensitive to the orientations of Néel vector. Finally, the magneto-optical response for bilayer FePS₃ and NiPS₃ with different stackings of zigzag antiferromagnetic chains is studied. Surprisingly, the Schäfer-Hubert angle of bilayer NiPS3 records up to 2.4°, and the obvious discrepancy in ellipticity spectra enables a distinction of different interlayer magnetic structures. The excellent properties render the MPS₃ family a novel antiferromagnetic materials platform for nanophotonic devices. More importantly, our theoretical framework allows for high-throughput study of the Schäfer-Hubert effect among 2D antiferromagnetic materials for finding potentially interesting systems.

As is well known, optical second-harmonic generation (SHG) is another powerful technique for exploring electronic and magnetic structures of antiferromagnetic materials [47]. Nevertheless, the Schäfer-Hubert effect studied in this work has at least two distinct aspects in comparison with SHG. First, the SHG response is primarily attributed to electric-dipole and electric-quadrupole contributions. The electric-dipole term, known as the c type, emerges when the underlying magnetic structure breaks inversion symmetry. It is directly proportional to the Néel vector L and changes sign when L is flipped by 180° . The c-type SHG enables the detection of opposite 180° antiferromagnetic domains, whereas it has been observed in only a few materials and is relatively rare. If the inversion symmetry is preserved, the higher-order electric-quadrupole term may appear, but it does not provide any magnetic information as it is timereversal invariant [28]. In contrast, the Schäfer-Hubert effect is more commonly observed in both centrosymmetric and noncentrosymmetric antiferromagnets. This broader applicability makes the Schäfer-Hubert effect suitable for a wider range of materials compared to SHG. Second, unlike SHG, which relies on a high-power laser due to its nonlinear nature and typically requires high light intensities, the Schäfer-Hubert effect is a kind of linear magneto-optical effect (with respect to electric field) that does not necessitate a high-intensity light source. Therefore, the Schäfer-Hubert effect offers a more convenient means of extracting information from antiferromagnetic domains without the need for high-power lasers.

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APPENDIX: COMPUTATIONAL METHODS

1. First-principles calculations

The electronic structure calculations were performed using the projector augmented wave (PAW) method [48], implemented in the Vienna ab initio Simulation Package (VASP) [49]. The exchange-correlation effects were treated using the generalized gradient approximation with the Perdew-Burke-Ernzerhof (GGA-PBE) parametrization [50]. The cutoff energy was set to 300 eV and the energy convergence criterion was chosen to 10^{-6} eV. A k mesh of $12 \times 12 \times 1$ ($12 \times 6 \times 1$) was used for the 1×1 (1×2) unit cell. The spin-orbit coupling effect was included in our calculations. The correlation effects of the d orbitals of Fe, Ni, and Mn atoms were treated by the GGA+U method [51], and the effective Hubbard parameters were set to 3.0 eV, 4.0 eV, and 5.0 eV, respectively. The experimental lattice constants are adopted for MnPS₃ (6.077 Å), FePS₃ (5.947 Å), and NiPS₃ (5.812 Å) [52]. The van der Waals interactions were considered using the DFT-D2 method [53]. A vacuum layer of 15 Å was used to eliminate the interactions between the adjacent atomic layers.

2. Magneto-optical Schäfer-Hubert effect

The complex Schäfer-Hubert angle in two-dimensional (2D) materials was computed according to Eq. (7). We constructed the maximally localized Wannier functions, including

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the *d* orbitals of Mn, Fe, and Ni atoms, the *s* and *p* orbitals of P atoms, and the *p* orbitals of S atoms, using the WANNIER90 package [54]. Then the optical conductivity was calculated using the Kubo-Greenwood formula [55],

$$\sigma_{\mu\nu} = \frac{ie^{2}\hbar}{N_{k}V} \sum_{\mathbf{k}} \sum_{n,m} \frac{f_{m\mathbf{k}} - f_{n\mathbf{k}}}{E_{m\mathbf{k}} - E_{n\mathbf{k}}} \times \frac{\langle \psi_{n\mathbf{k}} | \hat{\upsilon}_{\mu} | \psi_{m\mathbf{k}} \rangle \langle \psi_{m\mathbf{k}} | \hat{\upsilon}_{\nu} | \psi_{n\mathbf{k}} \rangle}{E_{m\mathbf{k}} - E_{n\mathbf{k}} - (\hbar\omega + i\eta)}, \qquad (A1)$$

where $f_{n\mathbf{k}}$, V, N_k , ω , and η are the Fermi-Dirac distribution function, volume of unit cell, total number of k points in the Brillouin zone, photon frequency, and energy smearing parameter, respectively. $\hat{v}_{\mu(\nu)}$ is a velocity operator with subscripts μ , $\nu \in \{x, y, z\}$ denoting Cartesian components. $\psi_{n\mathbf{k}}$ and $E_{n\mathbf{k}}$ are the Wannier functions and interpolated energy at the band index n and momentum \mathbf{k} , respectively. A k mesh of $400 \times 400 \times 1$ was used to converge the optical conductivity and η was set to be 0.1 eV. The effective thicknesses (d) of MnPS₃, FePS₃, and NiPS₃ monolayers were taken from the interlayer distances of their bulk compounds, that is, 6.796 Å, 6.722 Å, and 6.632 Å, respectively [44]. The experimental refractive index of SiO₂ at different photon energies [56] was acquired from an online database [61].

3. Dipole selection rules

The characters of the energy bands at high-symmetry k points were determined using the MagVasp2trace code [57,58]. The corresponding irreducible corepresentations and dipole selection rules were identified by the MSGCorep package [59,60]. Here, we take the magnetic space group $P\vec{3}$ 1m as an example to illustrate how to find out the dipole selection rules. For an in-plane polarized light (i.e., $\mathbf{E} \perp z$), the dipole operators are defined by either $-e\hat{x}$ or $-e\hat{y}$, which together transform as the bases of the irreducible corepresentation Γ_3 of the group $P\vec{3}$ 1m. Using the command "showMSG-CorepDirectProduct" in MSGCorep package, we can obtain the direct products and their decompositions between Γ_3 and other corepresentations (see supplemental Fig. S3). It is easy to find that the dipole selection rules are $\Gamma_4 \leftrightarrow \Gamma_4$ and $\Gamma_4 \leftrightarrow \Gamma_5\Gamma_6$.

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