




Sound attenuation in the hyperhoneycomb Kitaev spin liquidKexin Feng ¹, Aysel Shiraliev ², and Natalia B. Perkins ¹¹*School of Physics and Astronomy, University of Minnesota, Minneapolis, Minnesota 55455, USA*²*Ecole Normale Supérieure, Paris, France*

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In recent years, it has been shown that the phonon dynamics may serve as an indirect probe of fractionalization of spin degrees of freedom. Here we propose that the sound attenuation measurements allows for the characterization and identification of the Kitaev quantum spin liquid on the hyperhoneycomb lattice, which is particularly interesting since the strong Kitaev interaction was observed in the the hyperhoneycomb magnet β -Li₂IrO₃. To this end we consider the low-temperature scattering between acoustic phonons and gapless Majorana fermions with nodal-line band structure. We find that the sound attenuation has a characteristic angular dependence, which is explicitly shown for the high-symmetry planes at temperatures below the flux energy gap.

DOI: [10.1103/PhysRevB.106.144424](https://doi.org/10.1103/PhysRevB.106.144424)**I. INTRODUCTION**

Quantum spin liquids (QSLs) are states of matter in which no symmetry is broken. QSLs are interesting in general because they exhibit a remarkable set of collective phenomena including topological ground-state degeneracy, long-range entanglement, and fractionalized excitations [1–7]. In recent years, much work has been done to understand the nature of QSLs. However, this is not generically an easy task since QSLs in realistic models are usually ensured by frustration, either from a particular geometry of the lattice structure or from competing spin interactions, even identifying the models, which host such states is challenging. In this sense, the exactly solvable Kitaev model on the honeycomb lattice with QSL ground state [8] and its possible realization in strongly spin-orbit couple materials [9,10] helped us both with getting a deeper insight in the nature of QSL state and developing new approaches for detection of this exotic phase of matter in experiment. A promising route for searching for QSL physics in real materials is to look for signatures of spin fractionalizations in various types of dynamical probes, such as inelastic neutron scattering, Raman scattering, resonant inelastic x-ray scattering, ultrafast spectroscopy, and terahertz nonlinear coherent spectroscopy [5–7,11]. A possibility to compute the corresponding response functions analytically in the Kitaev model provides a unique opportunity to explore the characteristic fingerprints of the QSL physics in these dynamical probes on a more quantitative level [12–21]. This is highly significant, because it gives us an opportunity to learn about generic behavior of other QSLs, which are much more difficult to describe.

It was recently shown that a lot of information can be obtained by studying the phonon dynamics in the QSL candidate materials [22–29], since the spin-lattice coupling is inevitable and often rather strong in real materials [30–33]. The characteristic modifications of the phonon dynamics from the Kitaev QSL compared with their nonmagnetic or magnet-

ically ordered analogs can be probed in various observables, including the renormalization of the spectrum of acoustic phonons [33], particular temperature dependence of the sound attenuation pattern and the phonon Hall viscosity [26–28], the Fano lineshapes in the optical phonon Raman spectrum caused by the overlapping of the optical phonon peaks with the continuum of the fractionalized excitations [29,34–39], thermal conductivity and thermal Hall effect [31]. While the presence of the exact solution of the Kitaev model helps to understand the dynamics of the phonons coupled to the underlying QSL qualitatively, first-principles studies highlighted the importance of the magnetoelastic coupling for the description of the various experimental findings in the candidate materials and, in particular, in α -RuCl₃ [40–42].

The Kitaev model can be generalized and defined for various three-coordinated three-dimensional lattices [15,16,19,21,43–47], including the hyperhoneycomb, stripy-honeycomb, hyperhexagon, and hyperoctagon lattices. As a two-dimensional counterpart, these models are exactly solvable and have QSL ground state with fractionalized excitations that are gapless Majorana fermions and gapped Z₂ gauge fluxes for the isotropic coupling parameters. Importantly, the Majorana fermions exhibit a rich variety of nodal structures due to the different (projective) ways symmetries can act on them [44,45]. These nodal structures include nodal lines for the hyperhoneycomb and the stripyhoneycomb models [43], Fermi surfaces for the hyperoctagon model [44], and the Weyl points for the hyperhexagon model [45].

In this paper we performed a study of the phonon dynamics in the Kitaev model on the hyperhoneycomb lattice, which is particularly important among three-dimensional Kitaev models because of the existence of the Kitaev candidate material β -Li₂IrO₃ [48–53], which is realized on the hyperhoneycomb lattice. While we know that other interactions are present in this compound in addition to the dominant Kitaev interaction, here we assume that some good intuition can be obtained by studying the limiting case of the pure Kitaev model. To

this end, we derived the Majorana fermion-phonon coupling vertices using the symmetry considerations and used them for computation of the phonon attenuation.

The rest of the paper is organized as follows. In Sec. II, we present the derivation of the spin-phonon Kitaev Hamiltonian on the hyperhoneycomb lattice. We start by reviewing the Kitaev spin model on the hyperhoneycomb lattice in Sec. II A. We obtain its fermionic band structure and show that the fermions are gapless along the nodal line within the Γ - X - Y plane, for which we obtain an analytical equation. In Sec. II B, we introduce the lattice Hamiltonian for acoustic phonons on the hyperhoneycomb lattice. We obtain the acoustic phonon spectrum for the D_{2h} point group symmetry in the long wavelength approximation. In Sec. II C, we present the explicit microscopic derivation of the Majorana fermion-phonon (MFPh) coupling vertices and show that there are four symmetry channels, which contribute into them. The knowledge of the MFPh couplings allows us to compute the phonon dynamics, so we use the diagrammatic techniques and in Sec. III compute the phonon polarization bubble. In Sec. IV, we present our numerical results for the attenuation coefficient for acoustic phonon modes. To this end, we first discuss the kinematic constraints in Sec. IV A and then in Sec. IV B analyze the angular dependence of the sound attenuation coefficient for acoustic phonons with different polarizations. In Sec. V, we present a short summary and discuss the possibility for the spin fractionalization in the Kitaev hyperhoneycomb model to be seen in the sound attenuation measurements by the ultrasound experiments.

II. THE SPIN-PHONON MODEL

In this section, we introduce the spin-phonon coupled Kitaev model on the hyperhoneycomb lattice and discuss its phonon dynamics. It is described by the following Hamiltonian:

$$\mathcal{H} = \mathcal{H}^s + \mathcal{H}^{\text{ph}} + \mathcal{H}^c. \quad (1)$$

The *first term* in Eq. (1) is the spin Hamiltonian. The *second term* is the bare Hamiltonian for the acoustic phonons. The *third term* is the magnetoelastic coupling.

A. The Kitaev model on the hyperhoneycomb lattice

We start by revisiting the main features of the Kitaev QSL realized on the hyperhoneycomb lattice previously discussed in Refs. [19,45]. The hyperhoneycomb lattice is a face-centered orthorhombic lattice with four sites per primitive unit cell. Apart from translational symmetry, the crystal structure is invariant under the D_{2h} point group symmetry. The conventional orthorhombic unit cell is set by the crystallographic axes $\{\hat{\mathbf{a}}, \hat{\mathbf{b}}, \hat{\mathbf{c}}\}$, as shown in Fig. 1. The Cartesian axes $\{\hat{\mathbf{x}}, \hat{\mathbf{y}}, \hat{\mathbf{z}}\}$ used to write the spin vector field is expressed as $\hat{\mathbf{x}} = (\hat{\mathbf{a}} + \hat{\mathbf{c}})/\sqrt{2}$, $\hat{\mathbf{y}} = (\hat{\mathbf{c}} - \hat{\mathbf{a}})/\sqrt{2}$, and $\hat{\mathbf{z}} = -\hat{\mathbf{b}}$. Different bond types x , y , and z are marked by red, green, and blue, respectively. Note, however, that there are two nonequivalent types of x and y bonds, and the hyperhoneycomb structure can be viewed as a stacking of two types of zigzag chains formed by x, y bonds and x', y' bonds, each pair running along $\mathbf{a} + \mathbf{b}$ and $\mathbf{a} - \mathbf{b}$ directions, respectively. The two types of chains are

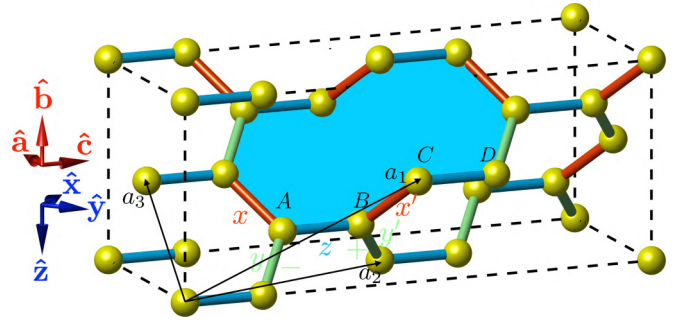


FIG. 1. The sketch of the hyperhoneycomb lattice. The conventional orthorhombic unit cell is set by the crystallographic axes $\hat{\mathbf{a}}$, $\hat{\mathbf{b}}$, and $\hat{\mathbf{c}}$. The three lattice vectors of the primitive face-centered orthorhombic lattice are given by $\mathbf{a}_1 = (0, \sqrt{2}, 3)$, $\mathbf{a}_2 = (1, 0, 3)$, $\mathbf{a}_3 = (1, \sqrt{2}, 0)$, which is written in the crystallographic basis. The four sublattices A , B , C , and D are shown, and we set $\mathbf{r}_A = (0, 0, 0)$. Different bond types x , y , and z are marked by red, green, and blue, respectively. The Cartesian axes $\{\hat{\mathbf{x}}, \hat{\mathbf{y}}, \hat{\mathbf{z}}\}$ used to write the spin Hamiltonian Eq. (2) is related to the crystallographic orthorhombic axes by $\hat{\mathbf{x}} = (\hat{\mathbf{a}} + \hat{\mathbf{c}})/\sqrt{2}$, $\hat{\mathbf{y}} = (\hat{\mathbf{c}} - \hat{\mathbf{a}})/\sqrt{2}$, and $\hat{\mathbf{z}} = -\hat{\mathbf{b}}$. The shaded region denotes a loop on the hyperhoneycomb lattice containing 10 sites.

interconnected with vertical z bonds. Thus, in total, there are five types of nearest neighboring bonds: x , x' , y , y' , and z .

The Kitaev spin model on the hyperhoneycomb lattice reads

$$\mathcal{H}^s = -J \left(\sum_{\langle \mathbf{r}\mathbf{r}' \rangle \in \{x, x'\}} \sigma_{\mathbf{r}}^x \sigma_{\mathbf{r}'}^x + \sum_{\langle \mathbf{r}\mathbf{r}' \rangle \in \{y, y'\}} \sigma_{\mathbf{r}}^y \sigma_{\mathbf{r}'}^y + \sum_{\langle \mathbf{r}\mathbf{r}' \rangle \in \{z\}} \sigma_{\mathbf{r}}^z \sigma_{\mathbf{r}'}^z \right), \quad (2)$$

where \mathbf{r} and \mathbf{r}' are sites on the three-dimensional hyperhoneycomb lattice, which we sketch in Fig. 1 and the summation is done over five types of bonds. We also assumed the isotropic case with $J_x = J_y = J_z = J$. The symmetry of the Hamiltonian (2) involves a combined lattice and spin transformations [54,55] (for detailed mathematical description of the transformations, see Refs. [39,56]). The transformation of the spin operators in Eq. (2) under the three π rotations around the crystallographic axes \mathbf{a} , \mathbf{b} , and \mathbf{c} are the following. Under $C_{2\mathbf{a}}$ rotation spins transform as $[\sigma^x, \sigma^y, \sigma^z] \rightarrow [-\sigma^y, -\sigma^x, -\sigma^z]$, under $C_{2\mathbf{b}}$ rotation $[\sigma^x, \sigma^y, \sigma^z] \rightarrow [-\sigma^x, -\sigma^y, \sigma^z]$, and under $C_{2\mathbf{c}}$ rotation $[\sigma^x, \sigma^y, \sigma^z] \rightarrow [\sigma^y, \sigma^x, -\sigma^z]$. Additionally, D_{2h} group also contains the space inversion I at the middle of x , x' , y , y' bonds, which together with spin transformation leads to $[\sigma^x, \sigma^y, \sigma^z] \rightarrow [\sigma^y, \sigma^x, \sigma^z]$. The transformation $C_{2\mathbf{c}}$, $C_{2\mathbf{b}}$, and I constitute the canonical generators that generate the whole D_{2h} group.

The exact solution of model (2) is based on the macroscopic number of local symmetries in the products of particular components of the spin operators around every plaquette P , which on the hyperhoneycomb lattice consists of ten sites (see shaded region in Fig. 1) and is defined by the following plaquette operator $\hat{W}_p = \prod_{\mathbf{r} \in P} \sigma_{\mathbf{r}}^{\gamma(\mathbf{r})}$, where the spin component $\gamma(\mathbf{r})$ is given by the label of the outgoing bond direction. Since all plaquette operators \hat{W}_p commute with the Hamiltonian, $[\hat{W}_p, \mathcal{H}^s] = 0$, and take eigenvalues of ± 1 , the Hilbert space of the spin Hamiltonian \mathcal{H}^s can be divided into eigenspaces of \hat{W}_p . The ground state of the Kitaev model

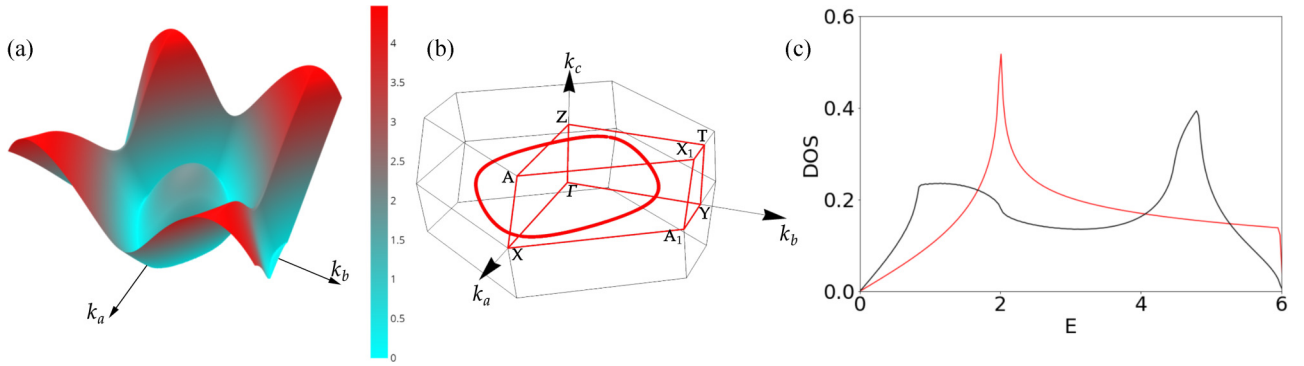


FIG. 2. (a) The dispersion of the lowest branch of the fermionic excitations in the hyperhoneycomb Kitaev model through the plane of the nodal line $\mathbf{K}_0 = (k_a, k_b, 0)$, whose position in the Brillouin zone is explicitly shown in panel (b). (c) One-fermion density of states (DOS) of the isotropic Kitaev models on the honeycomb (red line) and the hyperhoneycomb (black line) lattices. In each case, the density of states is normalized to unity.

on the hyperhoneycomb lattice is the zero-flux state with all $\hat{W}_p = 1$ [43,45]. This, however, can not be derived exactly from the Lieb's theorem [57] but is only based on the numerical calculations [21,45]. Thus, strictly speaking, the Kitaev model on hyperhoneycomb lattice is not exactly solvable. Another striking difference between the hyperhoneycomb Kitaev spin liquid and its two-dimensional counterpart regards the effect of the thermal fluctuations on the stability of the ground-state zero-flux state. While in two-dimensional honeycomb lattice thermal fluctuations immediately destroy the zero-flux order of the Z_2 gauge field [58,59], in three spatial dimensions there is a finite-temperature transition separating it from a high-temperature disordered flux state [21,60,61].

Using the Kitaev's representation of spins in terms of Majorana fermions [8], $\sigma_{\mathbf{r}}^\gamma = ib_{\mathbf{r}}^\gamma c_{\mathbf{r}}$ with $\gamma = x, y, z$ [8], the spin Hamiltonian Eq. (2) can be rewritten as

$$\mathcal{H}^s = \sum_{\gamma} \sum_{\langle \mathbf{r}, \mathbf{r}' \rangle_{\gamma}} iJ_{\gamma} \eta_{\mathbf{r}, \mathbf{r}'}^{\gamma} c_{\mathbf{r}} c_{\mathbf{r}'} = \frac{1}{2} \sum_{\mathbf{r}, \mathbf{r}'} \mathcal{H}_{\mathbf{r}, \mathbf{r}'} c_{\mathbf{r}} c_{\mathbf{r}'}, \quad (3)$$

where $\eta_{\mathbf{r}, \mathbf{r}'}^{\gamma} \equiv ib_{\mathbf{r}}^{\gamma} b_{\mathbf{r}'}^{\gamma} = \pm 1$, $\mathcal{H}_{\mathbf{r}, \mathbf{r}'} = iJ_{\gamma} \eta_{\mathbf{r}, \mathbf{r}'}^{\gamma}$ if \mathbf{r} and \mathbf{r}' are neighboring sites connected by a γ bond and $\mathcal{H}_{\mathbf{r}, \mathbf{r}'} = 0$ otherwise. In the ground-state flux sector, we choose the gauge sector with all $\eta_{\mathbf{r}, \mathbf{r}'}^{\gamma} = 1$, which corresponds to all $\hat{W}_p = 1$. The quadratic fermionic Hamiltonian in Eq. (3) can be diagonalized via a standard procedure [8]. Since the hyperhoneycomb lattice has four sites per unit cell, the resulting band structure has four fermion bands, $\xi = 1 \sim 4$ ($\xi = 1, 2$ are the two positive bands). The diagonal form of the Hamiltonian [19]

$$\mathcal{H}^s = \sum_{\mathbf{k}} \sum_{\xi=1}^4 \varepsilon_{\mathbf{k}, \xi} [\psi_{\mathbf{k}, \xi}^{\dagger} \psi_{\mathbf{k}, \xi} - 1/2] \quad (4)$$

is then obtained by the unitary transformation $\tilde{\mathcal{H}}_{\mathbf{k}} = \mathcal{W}_{\mathbf{k}} \cdot \mathcal{E}_{\mathbf{k}} \cdot \mathcal{W}_{\mathbf{k}}^{\dagger}$ of the Hermitian matrix $\tilde{\mathcal{H}}_{\mathbf{k}}$ with elements $(\tilde{\mathcal{H}}_{\mathbf{k}})_{\nu\nu'} = \frac{1}{N} \sum_{\mathbf{r} \in \nu} \sum_{\mathbf{r}' \in \nu'} \mathcal{H}_{\mathbf{r}, \mathbf{r}'} e^{i\mathbf{k} \cdot (\mathbf{r}' - \mathbf{r})}$, where ν and ν' denote sublattices a, b, c, d shown in Fig. 1, and $\varepsilon_{\mathbf{k}, \xi} = (\hat{\mathcal{E}}_{\mathbf{k}})_{\xi\xi}$ are the fermionic energies. The fermionic eigenmodes are given by

$$\psi_{\mathbf{k}, \xi} = \frac{1}{\sqrt{2N}} \sum_{\nu=1}^4 (\mathcal{W}_{\mathbf{k}}^{\dagger})_{\xi\nu} \sum_{\mathbf{r} \in \nu} c_{\mathbf{r}} e^{-i\mathbf{k} \cdot \mathbf{r}}. \quad (5)$$

Note that only the fermions $\psi_{\mathbf{k}, \xi}$ with energies $\varepsilon_{\mathbf{k}, \xi} > 0$ are physical due to the particle-hole redundancy $\tilde{\mathcal{H}}_{-\mathbf{k}} = -\tilde{\mathcal{H}}_{\mathbf{k}}^*$, which implies $\psi_{-\mathbf{k}, \xi} = \psi_{\mathbf{k}, \xi}^{\dagger}$ and $\varepsilon_{-\mathbf{k}, \xi} = \varepsilon_{\mathbf{k}, \xi}$. Thus, only two branches have positive spectrum. The lowest branch $\varepsilon_{\mathbf{k}, 1}$ [shown in Fig. 2(a)] exhibits the nodal line on the (k_a, k_b) plane [Fig. 2(b)], which is protected by projective time-reversal symmetry [45]. By solving the equation $\varepsilon_{\mathbf{k}, 1} = 0$, we obtained the functional form of the nodal line $\mathbf{K}_0 = (k_a, k_b, 0)$ with

$$k_b = \frac{1}{\sqrt{2}} \arg(1 - 2 \cos k_a \pm i\sqrt{1 + 4 \cos k_a - 2 \cos 2k_a}). \quad (6)$$

The energy dispersion is linear if expanded around the nodal line, i.e., each point of the nodal line represents a Dirac cone. Importantly, the Fermi velocity varies along the nodal line and depends on the direction of the deviation from it, i.e., $v_F = v_F(\mathbf{K}_0, \delta\mathbf{k})$, where $\delta\mathbf{k} = (\delta k_a, \delta k_b, \delta k_c)$. As we will see later, the spacial dependence of the Fermi velocity of the low-energy Majorana fermions will lead to the qualitative difference in the temperature dependence of the sound attenuation coefficient between the hyperhoneycomb model and the honeycomb Kitaev model [27,28].

To further characterized the spectrum of Majorana fermions, in Fig. 2(c) we plot the density of states $\text{DOS}(E) = \sum_{\xi=1,2} \int_{BZ} \delta(E - \varepsilon_{\mathbf{k}, \xi}) d^3\mathbf{k}$ for the hyperhoneycomb Kitaev model (shown by the black line) where the contributions from both branches of Majorana fermions are summed up. The low-energy DOS is linear in energy, which follows directly from the linear low-energy dispersion and the dimension of the Fermi surface [15]. For comparison, in Fig. 2(c) we also plot the DOS for the honeycomb model (shown by the red line). The differences between the DOS for these two lattices can be understood in terms of the number of fermionic bands, one for the honeycomb lattice and two for the hyperhoneycomb lattices, and their nodal structure—two Dirac points for the honeycomb lattice and the closed line of Dirac points for the hyperhoneycomb lattice. The former leads to the absence of the Van Hove singularities and overall more flatten DOS for the hyperhoneycomb lattice. The latter is responsible for a faster growth of the hyperhoneycomb DOS

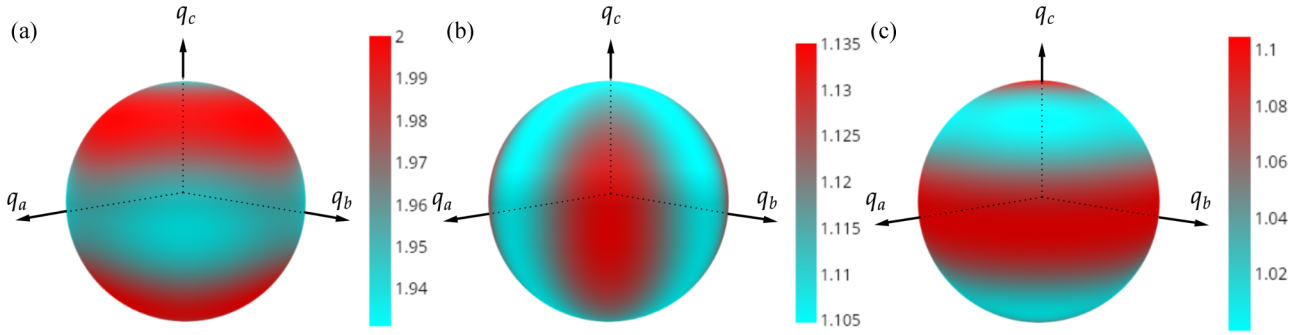


FIG. 3. Angular dependence of the sound velocities [in the unit of 10^4 m/s] for (a) longitudinal mode ($v_s^{\parallel}(\theta, \phi)$) and (b) in-plane transverse mode ($v_s^{\perp 1}(\theta, \phi)$) and (c) out-of-plane transverse mode ($v_s^{\perp 2}(\theta, \phi)$) computed, for the elastic modulus coefficients close to those computed for β - Li_2IrO_3 [64].

at low energies, which is consistent with higher dimensionality of the nodal line and enlarged number of low-energy states.

B. Acoustic phonons on the hyperhoneycomb lattice

Next we find the spectrum of the acoustic phonons on the hyperhoneycomb lattice. The bare Hamiltonian for the acoustic phonons contains the kinetic and elastic energy, $\mathcal{H}^{\text{ph}} = \mathcal{H}_{\text{kinetic}}^{\text{ph}} + \mathcal{H}_{\text{elastic}}^{\text{ph}}$, where $\mathcal{H}_{\text{kinetic}}^{\text{ph}} = \sum_{\mathbf{q}, \mu} \frac{\mathbf{P}_{\mathbf{q}, \mu} \cdot \mathbf{P}_{\mathbf{q}, \mu}}{2\rho\delta_V}$ with $\mathbf{P}_{\mathbf{q}, \mu}$ denoting the momentum of the phonon with polarization μ , δ_V is the area enclosed in one unit cell and ρ is the mass density of the lattice ions. The elastic contribution $\mathcal{H}_{\text{elastic}}^{\text{ph}}$ can be expressed in terms of the strain tensor $\epsilon_{ij} = \frac{1}{2}(\partial_i u_j + \partial_j u_i)$, where $\mathbf{u} = \{u_a, u_b, u_c\}$ describes the displacement of an atom from its original location.

In order to describe the dynamics of the low-energy acoustic phonons, it is convenient to move away from the Hamiltonian formulation and employ instead the long-wavelength effective action \mathcal{S} approach. To lowest order, it

reads [62]

$$\mathcal{S}_{\text{ph}}^{(s)} = \int d^2x d\tau [\rho (\partial_\tau \mathbf{u})^2 + F], \quad F = \frac{1}{2} C_{ijkl} \epsilon_{ij} \epsilon_{lk}, \quad (7)$$

where F is the elastic free energy and C_{ijkl} denote the elements of the elastic modulus tensor. The number of independent nonzero C_{ijkl} is dictated by symmetry. The hyperhoneycomb lattice has $Fddd$ space group [55], which is generated by three glide planes, which are passing through the bond center of either of x, x', y, y' bonds and are orthogonal to the $\mathbf{a}, \mathbf{b}, \mathbf{c}$ axes, respectively. The hyperhoneycomb lattice also has inversion symmetry with respect to the bond center of x, x', y, y' bonds. The inversion thus can be generated by the product of glide mirrors, e.g., the inversion on the x bond can be generated by $d_1^{-1} d_2 d_3^{-1}$, where each d_i glide is accompanied by a half of lattice translation along the primitive lattice vector \mathbf{a}_i [63]. Thus the point group is isomorphic to the D_{2h} , for which there are nine independent nonzero components of the elastic modulus tensor $C_{iiii}, C_{ijij}, C_{ijji}$, where i and j denote a, b, c . Performing the Fourier transform, $\mathbf{u}(\mathbf{r}) = \frac{1}{\sqrt{N}} \sum_{\mathbf{q}} e^{i\mathbf{q}\cdot\mathbf{r}} \mathbf{u}_{\mathbf{q}}$, the elastic free energy can be explicitly written as

$$F = \frac{1}{2} \begin{pmatrix} C_{aaaa}q_a^2 + C_{aacac}q_c^2 + C_{abab}q_b^2 & q_b q_a (C_{aabb} + C_{abab}) & q_a q_c (C_{aacac} + C_{aacac}) \\ q_b q_a (C_{aabb} + C_{abab}) & C_{abab}q_a^2 + C_{bcbcb}q_c^2 + C_{bbbb}q_b^2 & q_b q_c (C_{bbcc} + C_{bcbc}) \\ q_a q_c (C_{aacac} + C_{aacac}) & q_b q_c (C_{bbcc} + C_{bcbc}) & C_{aacac}q_a^2 + C_{cccc}q_c^2 + C_{bcbcb}q_b^2 \end{pmatrix}, \quad (8)$$

where $q_a = q \sin \theta_{\mathbf{q}} \cos \phi_{\mathbf{q}}$, $q_b = q \sin \theta_{\mathbf{q}} \sin \phi_{\mathbf{q}}$, and $q_c = q \cos \theta_{\mathbf{q}}$ are the components of the acoustic vector \mathbf{q} in the orthorhombic reference frame. By diagonalizing the matrix (8), we compute eigenmodes, one longitudinal and two transverse acoustic modes, and the corresponding eigenenergies: the longitudinal and transverse acoustic phonons are then given by

$$\begin{pmatrix} u_{\mathbf{q}, a} \\ u_{\mathbf{q}, b} \\ u_{\mathbf{q}, c} \end{pmatrix} = \begin{pmatrix} R_{11} & R_{12} & R_{13} \\ R_{21} & R_{22} & R_{23} \\ R_{31} & R_{32} & R_{33} \end{pmatrix} \begin{pmatrix} \tilde{u}_{\mathbf{q}}^{\parallel} \\ \tilde{u}_{\mathbf{q}}^{\perp 1} \\ \tilde{u}_{\mathbf{q}}^{\perp 2} \end{pmatrix}, \quad (9)$$

where $\hat{R} \equiv \hat{R}(\theta_{\mathbf{q}}, \phi_{\mathbf{q}})$ is the transformation matrix, $\tilde{u}_{\mathbf{q}}^{\parallel}$, $\tilde{u}_{\mathbf{q}}^{\perp 1}$ and $\tilde{u}_{\mathbf{q}}^{\perp 2}$ are the longitudinal and transverse acoustic eigenmodes, respectively. The energies of the longitudinal and transverse

acoustic phonons are

$$\Omega_{\mathbf{q}}^{\nu} = v_s^{\nu}(\theta_{\mathbf{q}}, \phi_{\mathbf{q}}) q, \quad (10)$$

where the sound velocities v_s^{ν} for the longitudinal acoustic mode, $v_s^{\parallel}(\theta_{\mathbf{q}}, \phi_{\mathbf{q}})$, and two transverse modes, $v_s^{\perp 1}(\theta_{\mathbf{q}}, \phi_{\mathbf{q}})$ and $v_s^{\perp 2}(\theta_{\mathbf{q}}, \phi_{\mathbf{q}})$ are anisotropic in space.

In Fig. 3, we plot the angular dependence of these velocities computed for the elastic modulus tensor coefficients close to those computed for β - Li_2IrO_3 : we set $C_{iiii} = 2800$ kbar, $i = a, b, c$, $C_{aacac} = C_{bbcc} = 1300$ kbar, $C_{aabb} = C_{abab} = C_{aacac} = C_{bcbcb} = 900$ kbar [64]. We see that the angular dependence of the sound velocities is not that strong. In the plot, the maximum sound velocity is estimated to be 2×10^4 m/s, which is in the middle of the sound ve-

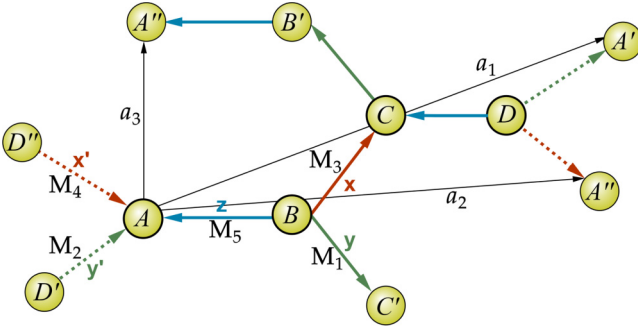


FIG. 4. A, B, C, D denote four sublattices of the hyperhoneycomb lattice. $\mathbf{M}_1 = \frac{1}{2}(1, \sqrt{2}, -1)$, $\mathbf{M}_2 = \frac{1}{2}(1, -\sqrt{2}, -1)$, $\mathbf{M}_3 = \frac{1}{2}(-1, -\sqrt{2}, -1)$, $\mathbf{M}_4 = \frac{1}{2}(-1, \sqrt{2}, -1)$, and $\mathbf{M}_5 = (0, 0, 1)$ are five nearest-neighboring vectors, corresponding to y, y', x, x', z bonds, respectively (all the vectors are given in the crystallographic axes $\hat{\mathbf{a}}, \hat{\mathbf{b}},$ and $\hat{\mathbf{c}}$). We use the following convention: An arrow pointing from site \mathbf{r} to \mathbf{r}' means $u_{\mathbf{r},\mathbf{r}'}$ on the corresponding bond is positive.

localities reported for different directions in α -RuCl₃ [65]. For the elastic modulus tensor given above, and restricting phonon modes to ab, bc and ac crystallographic planes, we numerically checked that the first column of the rotation matrix \hat{R} corresponding to the longitudinal mode indeed gives the vector parallel to \mathbf{q} , i.e., $(R_{11}, R_{21}, R_{31})^T = (\sin\theta_{\mathbf{q}} \cos\phi_{\mathbf{q}}, \sin\theta_{\mathbf{q}} \sin\phi_{\mathbf{q}}, \cos\theta_{\mathbf{q}})^T$, while the second and the third columns are perpendicular to \mathbf{q} (the second column with label \perp_1 corresponds to the in-plane transverse mode, and the third column with label \perp_2 corresponds to the out-of-plane transverse mode).

Knowing the acoustic phonon dispersion relations (10), we can now determine the free phonon propagator in terms of lattice displacement field $\tilde{u}_{\mathbf{q}}^{\nu}$ as

$$D_{\mathbf{q}}^{(0)\nu\nu'}(t) = -i \langle \mathcal{T} \tilde{u}_{-\mathbf{q}}^{\nu}(t) \tilde{u}_{\mathbf{q}}^{\nu'}(0) \rangle^{(0)}, \quad (11)$$

where \mathcal{T} is time ordering operator, the superscript (0) denotes the bare propagator, $\nu = \parallel, \perp_1, \perp_2$ labels the polarization, and $\tilde{u}_{\mathbf{q}}^{\nu}$ are phonon eigenmodes in the corresponding polarization, which in the second quantized form can be written as

$$\tilde{u}_{\mathbf{q}}^{\nu}(t) = i \left(\frac{\hbar}{2\rho \delta_V \Omega_{\mathbf{q}}^{\nu}} \right)^{1/2} (\tilde{a}_{\mathbf{q}} e^{-i\Omega_{\mathbf{q}}^{\nu} t} + \tilde{a}_{-\mathbf{q}}^{\dagger} e^{i\Omega_{\mathbf{q}}^{\nu} t}), \quad (12)$$

$$\begin{aligned} \mathcal{H}^c = \frac{1}{4} \lambda \left[\sum_{\mathbf{r}_A} (4\sigma_{\mathbf{r}_A}^z \sigma_{\mathbf{r}_A+\mathbf{M}_5}^z \epsilon_{cc} + \sigma_{\mathbf{r}_A}^y \sigma_{\mathbf{r}_A+\mathbf{M}_2}^y (\epsilon_{aa} + 2\epsilon_{bb} + \epsilon_{cc} - 2\sqrt{2}(\epsilon_{ab} - \epsilon_{bc}) - 2\epsilon_{ac})) \right. \\ \left. + \sigma_{\mathbf{r}_A}^x \sigma_{\mathbf{r}_A+\mathbf{M}_4}^x (\epsilon_{aa} + 2\epsilon_{bb} + \epsilon_{cc} - 2\sqrt{2}(\epsilon_{ab} + \epsilon_{bc}) + 2\epsilon_{ac}) \right] + \sum_{\mathbf{r}_B} (\sigma_{\mathbf{r}_B}^y \sigma_{\mathbf{r}_B+\mathbf{M}_1}^y (\epsilon_{aa} + 2\epsilon_{bb} + \epsilon_{cc} + 2\sqrt{2}(\epsilon_{ab} - \epsilon_{bc}) - 2\epsilon_{ac}) \\ \left. + \sigma_{\mathbf{r}_B}^x \sigma_{\mathbf{r}_B+\mathbf{M}_3}^x (\epsilon_{aa} + 2\epsilon_{bb} + \epsilon_{cc} + 2\sqrt{2}(\epsilon_{ab} + \epsilon_{bc}) + 2\epsilon_{ac}) \right), \quad (15) \end{aligned}$$

where we use a short notation $\epsilon_{ij} \equiv \epsilon_{ij}(\mathbf{r})$ with $\mathbf{r} = \mathbf{r}_A$ or \mathbf{r}_B depending on the bond and i, j one of the orthorhombic directions a, b, c .

Under the D_{2h} point group symmetry, the spin-phonon Hamiltonian has four independent symmetry channels, $A_g, B_{1g}, B_{2g},$ and B_{3g} , which are inversion-symmetric irreducible representations (IRRs) of this group. The linear combinations of the strain tensors that transform as the D_{2h} are $\epsilon_{aa}, \epsilon_{bb},$ and ϵ_{cc} , in the A_g channel, and $\epsilon_{ab}, \epsilon_{ac},$ and ϵ_{bc} in $B_{1g}, B_{2g},$ and B_{3g} , respectively. By writing the linear combinations of the Kitaev interactions that transform according to these IRRs, we express the spin-phonon

where δ_V is the area enclosed in one unit cell and ρ is the mass density of the lattice ions. In the momentum and frequency space, the bare phonon propagator is then given by

$$D^{(0)\nu\nu}(\mathbf{q}, \Omega) = -\frac{\hbar}{\rho \delta_V} \frac{1}{\Omega^2 - (\Omega_{\mathbf{q}}^{\nu})^2 + i0^+}. \quad (13)$$

The dynamics of phonons will be thus described by the decay and scattering of these eigenmodes on low-energy fractionalized excitations of the Kitaev model, which can be accounted for by the phonon self-energy $\Pi_{\text{ph}}(\mathbf{q}, \Omega)$ [27], which for this case we will discuss later in Sec. III. The renormalized phonon propagator is then given by the Dyson equation $D(\mathbf{q}, \Omega) = [D^{(0)}(\mathbf{q}, \Omega)^{-1} - \Pi_{\text{ph}}(\mathbf{q}, \Omega)]^{-1}$.

C. The Majorana fermion-phonon coupling vertices

In order to study the phonon dynamics in the Kitaev spin liquid, it remains to compute the Majorana fermion-phonon (MFPh) coupling vertices, which we will do in this section. We recall that the magneto-elastic coupling \mathcal{H}^c arises from the change in the Kitaev coupling due to the lattice vibrations. In the long wavelength limit for acoustic phonons, the coupling Hamiltonian on the bond can be written in a differential form as

$$\begin{aligned} \mathcal{H}_{\mathbf{r},\mathbf{r}+\mathbf{M}_{\alpha}}^c &= \lambda \mathbf{M}_{\alpha} \cdot [\mathbf{u}(\mathbf{r} + \mathbf{M}_{\alpha}) - \mathbf{u}(\mathbf{r})] \sigma_{\mathbf{r}}^{\alpha} \sigma_{\mathbf{r}+\mathbf{M}_{\alpha}}^{\alpha} \\ &= \lambda \mathbf{M}_{\alpha} \cdot [(\mathbf{M}_{\alpha} \cdot \nabla) \mathbf{u}(\mathbf{r})] \sigma_{\mathbf{r}}^{\alpha} \sigma_{\mathbf{r}+\mathbf{M}_{\alpha}}^{\alpha}, \quad (14) \end{aligned}$$

where $\lambda \sim (\frac{dJ}{dr})_{\text{eq}} \ell_a$ is the strength of the spin-phonon interaction and ℓ_a is the lattice constant, and $\mathbf{M}_{\alpha} = \mathbf{M}_1, \dots, \mathbf{M}_5$ are five nearest-neighboring vectors corresponding, respectively, to y, y', x, x', z bonds shown in Fig. 4. Note that this is a simplified form of the coupling, since it only takes into account the dependence of the Kitaev interaction between two nearest-neighbor pseudospins on the Ir-Ir distance but not on the Ir-O-Ir bond angle. Using \mathbf{M}_{α} vectors, we can write the spin-phonon coupling Hamiltonian explicitly [66],

coupling Hamiltonian (15) as a sum of four independent contributions, $\mathcal{H}^c = \mathcal{H}_{A_g}^c + \mathcal{H}_{B_{1g}}^c + \mathcal{H}_{B_{2g}}^c + \mathcal{H}_{B_{3g}}^c$ with

$$\begin{aligned}\mathcal{H}_{A_g}^c &= \lambda_{A_g} \sum_{\mathbf{r}_A, \mathbf{r}_B} \left[4\epsilon_{cc} \sigma_{\mathbf{r}_A}^z \sigma_{\mathbf{r}_A + \mathbf{M}_5}^z + (\epsilon_{aa} + 2\epsilon_{bb} + \epsilon_{cc}) (\sigma_{\mathbf{r}_B}^y \sigma_{\mathbf{r}_B + \mathbf{M}_1}^y + \sigma_{\mathbf{r}_A}^y \sigma_{\mathbf{r}_A + \mathbf{M}_2}^y + \sigma_{\mathbf{r}_B}^x \sigma_{\mathbf{r}_B + \mathbf{M}_3}^x + \sigma_{\mathbf{r}_A}^x \sigma_{\mathbf{r}_A + \mathbf{M}_4}^x) \right], \\ \mathcal{H}_{B_{1g}}^c &= \lambda_{B_{1g}} \sum_{\mathbf{r}_A, \mathbf{r}_B} \epsilon_{ab} (\sigma_{\mathbf{r}_B}^y \sigma_{\mathbf{r}_B + \mathbf{M}_1}^y - \sigma_{\mathbf{r}_A}^y \sigma_{\mathbf{r}_A + \mathbf{M}_2}^y + \sigma_{\mathbf{r}_B}^x \sigma_{\mathbf{r}_B + \mathbf{M}_3}^x - \sigma_{\mathbf{r}_A}^x \sigma_{\mathbf{r}_A + \mathbf{M}_4}^x), \\ \mathcal{H}_{B_{2g}}^c &= \lambda_{B_{2g}} \sum_{\mathbf{r}_A, \mathbf{r}_B} \epsilon_{ac} (-\sigma_{\mathbf{r}_B}^y \sigma_{\mathbf{r}_B + \mathbf{M}_1}^y - \sigma_{\mathbf{r}_A}^y \sigma_{\mathbf{r}_A + \mathbf{M}_2}^y + \sigma_{\mathbf{r}_B}^x \sigma_{\mathbf{r}_B + \mathbf{M}_3}^x + \sigma_{\mathbf{r}_A}^x \sigma_{\mathbf{r}_A + \mathbf{M}_4}^x), \\ \mathcal{H}_{B_{3g}}^c &= \lambda_{B_{3g}} \sum_{\mathbf{r}_A, \mathbf{r}_B} \epsilon_{bc} (-\sigma_{\mathbf{r}_B}^y \sigma_{\mathbf{r}_B + \mathbf{M}_1}^y + \sigma_{\mathbf{r}_A}^y \sigma_{\mathbf{r}_A + \mathbf{M}_2}^y + \sigma_{\mathbf{r}_B}^x \sigma_{\mathbf{r}_B + \mathbf{M}_3}^x - \sigma_{\mathbf{r}_A}^x \sigma_{\mathbf{r}_A + \mathbf{M}_4}^x),\end{aligned}\quad (16)$$

where we absorbed numerical prefactors into the definitions of the coupling constants λ_{A_g} , $\lambda_{B_{1g}}$, $\lambda_{B_{2g}}$, and $\lambda_{B_{3g}}$.

Next, we express the spin operators in terms of the Majorana fermions and assume the ground state flux sector. Then we perform the Fourier transformation on both the strain tensor, $\epsilon_{ij}(\mathbf{r}) = \frac{1}{\sqrt{N}} \sum_{\mathbf{q}} \frac{i}{2} (q_i u_{\mathbf{q},j} + q_j u_{\mathbf{q},i}) e^{i\mathbf{q}\cdot\mathbf{r}}$, and the Majorana fermions, $c_{\mathbf{r},\alpha} = \sqrt{\frac{2}{N}} \sum_{\mathbf{k}} c_{\mathbf{k},\alpha} e^{i\mathbf{k}\cdot\mathbf{r}_\alpha}$, where $\alpha = A, B, C, D$ is the sublattice label (see Fig. 4). Now the products of the spin variables on all nonequivalent bonds can be written as (with the long wavelength approximation $\mathbf{q} \rightarrow 0$ applied)

$$\begin{aligned}\sigma_{\mathbf{r}}^y \sigma_{\mathbf{r} + \mathbf{M}_1}^y &\rightarrow \mathbf{A}_{-\mathbf{q}-\mathbf{k}}^T S_{\mathbf{k}}^\dagger \begin{pmatrix} 0 & -ie^{i\mathbf{k}\cdot\mathbf{a}_3} & 0 & 0 \\ ie^{-i\mathbf{k}\cdot\mathbf{a}_3} & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \end{pmatrix} S_{\mathbf{k}} \mathbf{A}_{\mathbf{k}}, \\ \sigma_{\mathbf{r}}^y \sigma_{\mathbf{r} + \mathbf{M}_2}^y &\rightarrow \mathbf{A}_{-\mathbf{q}-\mathbf{k}}^T S_{\mathbf{k}}^\dagger \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & ie^{i\mathbf{k}\cdot\mathbf{a}_1} \\ 0 & 0 & -ie^{-i\mathbf{k}\cdot\mathbf{a}_1} & 0 \end{pmatrix} S_{\mathbf{k}} \mathbf{A}_{\mathbf{k}}, \\ \sigma_{\mathbf{r}}^x \sigma_{\mathbf{r} + \mathbf{M}_3}^x &\rightarrow \mathbf{A}_{-\mathbf{q}-\mathbf{k}}^T S_{\mathbf{k}}^\dagger \begin{pmatrix} 0 & -i & 0 & 0 \\ i & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \end{pmatrix} S_{\mathbf{k}} \mathbf{A}_{\mathbf{k}}, \\ \sigma_{\mathbf{r}}^x \sigma_{\mathbf{r} + \mathbf{M}_4}^x &\rightarrow \mathbf{A}_{-\mathbf{q}-\mathbf{k}}^T S_{\mathbf{k}}^\dagger \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & ie^{i\mathbf{k}\cdot\mathbf{a}_2} \\ 0 & 0 & -ie^{-i\mathbf{k}\cdot\mathbf{a}_2} & 0 \end{pmatrix} S_{\mathbf{k}} \mathbf{A}_{\mathbf{k}}, \\ \sigma_{\mathbf{r}}^z \sigma_{\mathbf{r} + \mathbf{M}_5}^z &\rightarrow \mathbf{A}_{-\mathbf{q}-\mathbf{k}}^T S_{\mathbf{k}}^\dagger \begin{pmatrix} 0 & 0 & -i & 0 \\ 0 & 0 & 0 & i \\ i & 0 & 0 & 0 \\ 0 & -i & 0 & 0 \end{pmatrix} S_{\mathbf{k}} \mathbf{A}_{\mathbf{k}},\end{aligned}$$

where \mathbf{a}_i are the primitive unit vectors, $S_{\mathbf{k}} = \text{diag}\{e^{i\mathbf{k}\cdot\mathbf{r}_\alpha}\}_{\alpha=C,B,D,A}$ is the diagonal matrix in the sublattice basis, and the vector $\mathbf{A}_{\mathbf{k}} = (c_{\mathbf{k},C}, c_{\mathbf{k},B}, c_{\mathbf{k},D}, c_{\mathbf{k},A})^T$. The Majorana-phonon coupling Hamiltonian in the momentum space is now can be written as $\mathcal{H}^c = \sqrt{\frac{2}{N}} \sum_{\mathbf{q}, \mathbf{k}} \mathcal{H}_{\mathbf{q}, \mathbf{k}}$, where each contribution $\mathcal{H}_{\mathbf{q}, \mathbf{k}}$ can be decomposed into the irreducible representations A_g, B_{1g}, B_{2g} and B_{3g} (see Appendix A for explicit expressions). Note also that $\mathbf{A}_{\mathbf{k}}$ is written in this particular permuted basis of the Majorana fermions in order to use the convenience of the auxiliary Pauli matrices in the

representation of the coupling Hamiltonians as shown in Eq. (A2).

Next we express the phonon modes in terms of the transverse and longitudinal eigenmodes defined in Eq. (9). Then $\mathcal{H}_{\mathbf{q}, \mathbf{k}}$ terms in the corresponding polarizations are given by

$$\begin{aligned}\mathcal{H}_{\mathbf{q}, \mathbf{k}}^{\parallel} &= \tilde{u}_{\mathbf{q}}^{\parallel} \mathbf{A}_{-\mathbf{q}-\mathbf{k}}^T S_{\mathbf{k}}^\dagger \hat{\lambda}_{\mathbf{q}, \mathbf{k}}^{\parallel} S_{\mathbf{k}} \mathbf{A}_{\mathbf{k}}, \\ \mathcal{H}_{\mathbf{q}, \mathbf{k}}^{\perp 1} &= \tilde{u}_{\mathbf{q}}^{\perp 1} \mathbf{A}_{-\mathbf{q}-\mathbf{k}}^T S_{\mathbf{k}}^\dagger \hat{\lambda}_{\mathbf{q}, \mathbf{k}}^{\perp 1} S_{\mathbf{k}} \mathbf{A}_{\mathbf{k}}, \\ \mathcal{H}_{\mathbf{q}, \mathbf{k}}^{\perp 2} &= \tilde{u}_{\mathbf{q}}^{\perp 2} \mathbf{A}_{-\mathbf{q}-\mathbf{k}}^T S_{\mathbf{k}}^\dagger \hat{\lambda}_{\mathbf{q}, \mathbf{k}}^{\perp 2} S_{\mathbf{k}} \mathbf{A}_{\mathbf{k}}.\end{aligned}\quad (17)$$

The explicit expressions for the MFPh coupling vertices $\hat{\lambda}_{\mathbf{q}, \mathbf{k}}^{\mu}$ are given by Eq. (A4) in Appendix A. Note also that since we are using the long wavelength limit for the phonons, we only kept the leading in q terms in all $\hat{\lambda}_{\mathbf{q}, \mathbf{k}}^{\mu}$.

III. PHONON POLARIZATION BUBBLE

At the lowest order, the phonon self-energy is given by the polarization bubble [27]

$$\Pi_{\text{ph}}^{\mu\nu}(\mathbf{q}, \Omega) = i\text{Tr}[\tilde{\lambda}_{\mathbf{q}, \mathbf{k}}^{\mu} \mathcal{G}(\mathbf{k}, \omega + \Omega) \tilde{\lambda}_{\mathbf{q}, \mathbf{k}}^{\nu} \mathcal{G}(\mathbf{k} + \mathbf{q}, \omega)], \quad (18)$$

where $\tilde{\lambda}_{\mathbf{q}, \mathbf{k}}^{\mu(v)} = \mathcal{W}_{\mathbf{k}} \hat{\lambda}_{\mathbf{q}, \mathbf{k}}^{\mu(v)} \mathcal{W}_{\mathbf{k}}^\dagger$ are the MFPh coupling vertices, expressed in the basis of the fermion eigenmodes, and $\mathcal{W}_{\mathbf{k}}$ is the unitary transformation matrix. $\mathcal{G}(\mathbf{k}, \omega)_{\xi\xi'} = -i \int_{-\infty}^{+\infty} dt \langle \mathcal{T} \psi_{\mathbf{k}, \xi}(t) \psi_{-\mathbf{k}, \xi'}^\dagger(0) \rangle e^{i\omega t}$ is the fermion Green's function, where ξ indexes the energy branches, making \mathcal{G} a 4×4 matrix. Since we focus on the low-temperature physics, hereafter, we will keep only the lower-positive-energy branch along with its counterpart negative-energy branch. We will also omit the branch index and simply write $\psi_{\mathbf{k}} \equiv \psi_{\mathbf{k}, 1}$. So the fermion Green's function becomes a 2×2 matrix: $\mathcal{G}(\mathbf{k}, \omega) = \begin{pmatrix} g(\mathbf{k}, \omega) & 0 \\ 0 & \bar{g}(\mathbf{k}, \omega) \end{pmatrix}$, where

$$\begin{aligned}g(\mathbf{k}, \omega) &= -i \int dt \langle \mathcal{T} \psi_{\mathbf{k}}(t) \psi_{\mathbf{k}}^\dagger(0) \rangle e^{i\omega t} = \frac{1}{\omega - \varepsilon_{\mathbf{k}} + i0^+}, \\ \bar{g}(\mathbf{k}, \omega) &= -i \int dt \langle \mathcal{T} \psi_{\mathbf{k}}^\dagger(t) \psi_{\mathbf{k}}(0) \rangle e^{i\omega t} = \frac{1}{\omega + \varepsilon_{\mathbf{k}} - i0^+}.\end{aligned}\quad (19)$$

Correspondingly, the MFPh coupling $\tilde{\lambda}_{\mathbf{q}, \mathbf{k}}^{\mu(v)}$ is also truncated to be 2×2 matrix.

Since we are interested in the phonon decay and scattering at finite temperature, it is convenient to express Eq. (18) in

terms of Matsubara frequency representation,

$$\begin{aligned} \Pi_{\text{ph}}^{\mu\nu}(\mathbf{q}, i\Omega_n) &= \int_{\text{BZ}} d\mathbf{k} \text{Tr} [\tilde{\chi}_{\mathbf{q},\mathbf{k}}^{\mu} \mathcal{G}(\mathbf{k}, i\omega_m + i\Omega_n) \tilde{\chi}_{\mathbf{q},\mathbf{k}}^{\nu}] \\ &\quad \times \mathcal{G}(\mathbf{k} + \mathbf{q}, i\omega_m)] \\ &= \int_{\text{BZ}} d\mathbf{k} \sum_{ij} \sum_l [\tilde{\chi}_{\mathbf{q},\mathbf{k}}^{\mu} \hat{E}_i \tilde{\chi}_{\mathbf{q},\mathbf{k}}^{\nu} \hat{E}_j]_{ll} P_{\mathbf{k},ij} \end{aligned} \quad (20)$$

where $\text{Tr}[\dots]$ in the first two lines sums over the Matsubara frequencies $i\omega_m$, and $\hat{E}_i = \begin{pmatrix} \delta_{i1} & 0 \\ 0 & \delta_{i2} \end{pmatrix}$ serves to pick up a specific entry of a matrix. The summation over the Matsubara frequencies gives the dynamic part of the matrix entries

$$P_{\mathbf{k},ij} = T \sum_{i\omega_m} [\mathcal{G}(\mathbf{k}, i\omega_m + i\Omega_n)_{ii} \mathcal{G}(\mathbf{k} + \mathbf{q}, i\omega_m)_{jj}]. \quad (21)$$

Their explicit expressions are given in Appendix C.

IV. ANGULAR DEPENDENCE OF THE ATTENUATION COEFFICIENT

In this section, we will compute the attenuation coefficient for the lossy acoustic wave function, which decays with distance away from the driving source as

$$\mathbf{u}(\mathbf{x}, t) = \mathbf{u}_0 e^{-\alpha_s(\mathbf{q})x} e^{i(\Omega_{\mathbf{q}}t - \mathbf{q} \cdot \mathbf{x})}, \quad (22)$$

where $\mathbf{u}(\mathbf{x}, t)$ is the lattice displacement vector, $\mathbf{u}_0 = \mathbf{u}(\mathbf{x} = 0, t = 0)$, $\Omega_{\mathbf{q}}$ is the acoustic wave frequency and $\mathbf{q} = q(\sin \theta_{\mathbf{q}} \cos \phi_{\mathbf{q}}, \sin \theta_{\mathbf{q}} \sin \phi_{\mathbf{q}}, \cos \theta_{\mathbf{q}})$ is the propagation vector. The attenuation coefficient $\alpha_s^{\mu}(\mathbf{q})$ for a given phonon polarization $\mu = \parallel, \perp_1, \perp_2$, defined as the inverse of the phonon mean free path, can be calculated from the diagonal component of the imaginary part of the phonon self-energy as [27]

$$\alpha_s^{\mu}(\mathbf{q}) = -\frac{1}{2\rho\delta_V [v_s^{\mu}(\theta_{\mathbf{q}}, \phi_{\mathbf{q}})]^2 q} \text{Im} \Pi_{\text{ph}}^{\mu\mu}(\mathbf{q}, \Omega) \Big|_{\Omega=v_s^{\mu}(\theta_{\mathbf{q}}, \phi_{\mathbf{q}})q}. \quad (23)$$

A. Kinematic constraints and the estimates for the sound and Fermi velocities in $\beta\text{-Li}_2\text{IrO}_3$

Before analyzing the angular dependence of the sound attenuation coefficient, we need first discuss the kinematic constraints determining type of the processes involved in sound attenuation. In the zero-flux low-temperature phase, both momentum and energy are conserved and kinematic constraints are primarily determined by the relative strength of acoustic phonon velocity $v_s(\theta_{\mathbf{q}}, \phi_{\mathbf{q}})$ and Fermi velocity $v_F(\mathbf{K}_0, \theta_{\delta\mathbf{k}}, \phi_{\delta\mathbf{k}})$ (the slope of the Dirac cone at each point of the nodal line), which in the most general case are both angular dependent. These constraints determine whether the decay of the acoustic phonon happens in the particle-hole (ph) or in the particle-particle (pp) channel. Here, by particle and hole, we mean if the state of the Majorana fermion at $\varepsilon_{\mathbf{k}}$ ($\varepsilon_{\mathbf{k}} = \varepsilon_{\mathbf{k},1}$) is occupied or empty. In other words, the particle number refers to that of the complex fermion $\psi_{\mathbf{k}}$ ($\psi_{\mathbf{k}} = \psi_{\mathbf{k},1}$) in Eq. (4).

Here we assume that the angular dependence of $v_s(\theta_{\mathbf{q}}, \phi_{\mathbf{q}})$ in $\beta\text{-Li}_2\text{IrO}_3$ is weak (see the magnitude scale bars in Fig. 3) and consider it to be equal to v_s . However, the Fermi velocity

$v_F(\mathbf{K}_0, \theta_{\delta\mathbf{k}}, \phi_{\delta\mathbf{k}})$ varies strongly between $v_F = 0$ along the nodal line and $\max(v_F)$, which can be estimated from the magnitude of the Kitaev coupling, which in $\beta\text{-Li}_2\text{IrO}_3$ is $J \simeq 20$ meV [51–53,67]. Taking the lattice constant to be equal to $\ell = 0.23$ nm [68], we estimate $\max(v_F) = 3J\ell = 2.1 \times 10^4$ m/s. According to the estimation of the sound velocity in Sec. II B, $v_s^{\parallel} \simeq 2.0 \times 10^4$ m/s $\lesssim \max(v_F)$, and $v_s^{\perp_{1,2}} \simeq 1.1 \times 10^4$ m/s $< \max(v_F)$. When $v_s < \max(v_F)$, the ph processes are allowed but since they require finite occupation number, they scale with T at low temperatures. However, due to the existence of the nodal line along which the Fermi velocity $v_F = 0$, the pp processes are always allowed [28]. Since they do not require finite occupation number, they are nonzero even at zero temperature. Therefore, both the pp processes and ph processes should be included into consideration.

B. Numerical results

Considering the estimations above, we set $v_s^{\parallel} = 3J\ell$ and $v_s^{\perp_{1,2}} \approx 1.6J\ell$. We also take $T = 0.02J$, which is below the flux energy gap. In the long wavelength limit, the angular dependence of the sound attenuation coefficient is scale invariant and is more experimentally relevant than the dependence on the magnitude of the momentum q . Thus, we fix $q = 0.005 \ell^{-1}$ and show the polar plots of the angular dependence of the sound attenuation (where the radius represents the magnitude of the sound attenuation coefficient). This angular dependence is a direct reflection of the Majorana-phonon couplings (16) constructed based on symmetry.

We compute the sound attenuation coefficient in the four symmetry channels, $A_g, B_{1g}, B_{2g}, B_{3g}$, considering separately the contributions from the pp- and ph-scattering processes. In Figs. 5–7, we present our results for the sound attenuation's angular dependence patterns for the phonon modes in the three crystallographic planes, correspondingly, ab , ac , and bc , for three phonon's polarizations, $\parallel, \perp_1, \perp_2$. The explicit expressions for the Majorana-phonon coupling vertices in these special geometries are presented in Appendix B. These expressions show that for each of the phonon polarizations, the coupling vertex has contributions from only two symmetry channels and another two symmetry channels give exactly zero contribution. Furthermore, we find that some symmetry channels have higher order (dominant) contributions in the long wavelength limit $q \rightarrow 0$. So, below we will only show the results from the leading order contributions into sound attenuation for each crystallographic plane.

Phonon within the ab plane. The contributions from the pp- and ph-scattering processes for the attenuation coefficient for the phonon propagating in the ab plane are shown in Figs. 5(a)–5(c) and 5(d), and 5(e), respectively. This plane is special compared with bc and ac planes, because of the presence of the nodal line in the fermionic spectrum [see Fig. 2(a)] as well as the crystallographic structure shown in Fig. 1. As such, there always exist zero Fermi velocities along the nodal line and small Fermi velocities in the vicinity of the nodal line. Therefore, the sound velocities along these directions are larger than the Fermi velocities, which gives rise to the nonzero pp processes. In this scattering geometry, the pp processes contribute only in the attenuation of the out-of-plane transverse phonon mode in the \perp_2 polarization

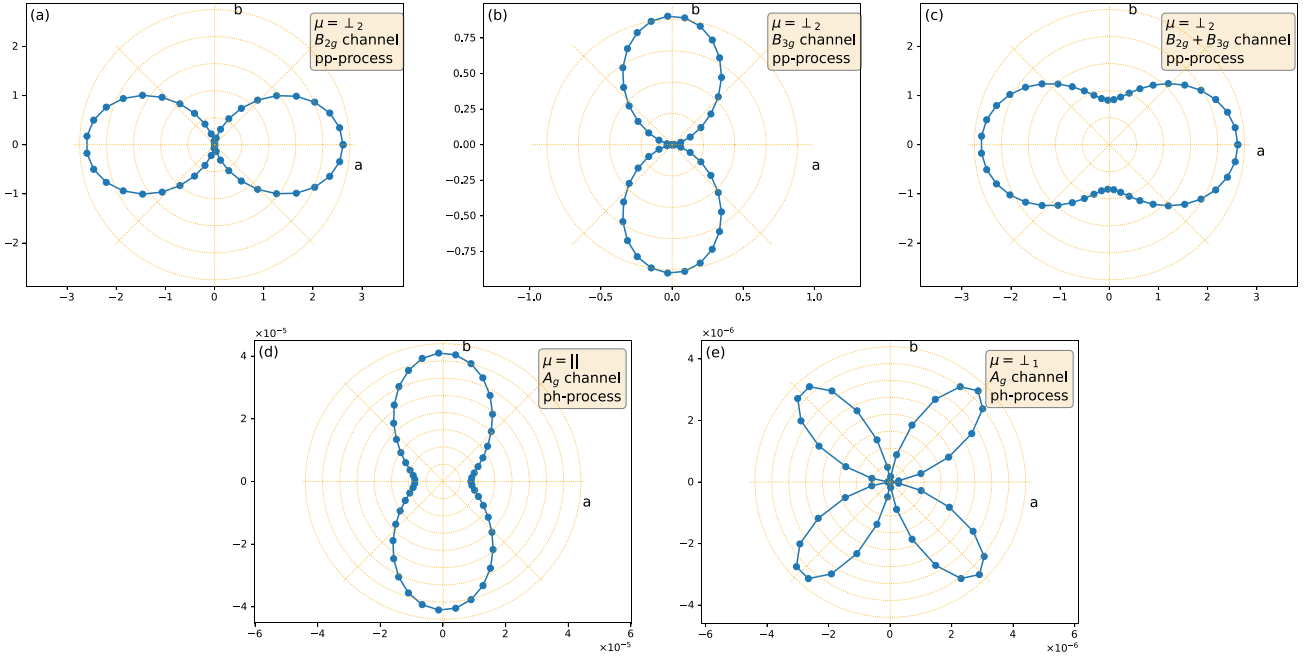


FIG. 5. The angular dependence of sound attenuation coefficient $\alpha_s^\mu(\theta_q = 90^\circ, \phi_q)$ in the ab plane. The pp processes contribute to the attenuation of the out-of-plane transverse phonon mode ($\alpha_s^{\perp 2}$) in (a) B_{2g} channel, (b) B_{3g} channel, and (c) combined B_{2g} and B_{3g} channels. The ph processes contribute to (d) the attenuation of the longitudinal phonon (α_s^\parallel) and to (e) the attenuation of the in-plane transverse phonon mode ($\alpha_s^{\perp 1}$) in the A_g channel. The radius represents the magnitude of α_s^μ in the units of $10^{-5}\rho\delta v$. The calculation is performed at $T = 0.02J$.

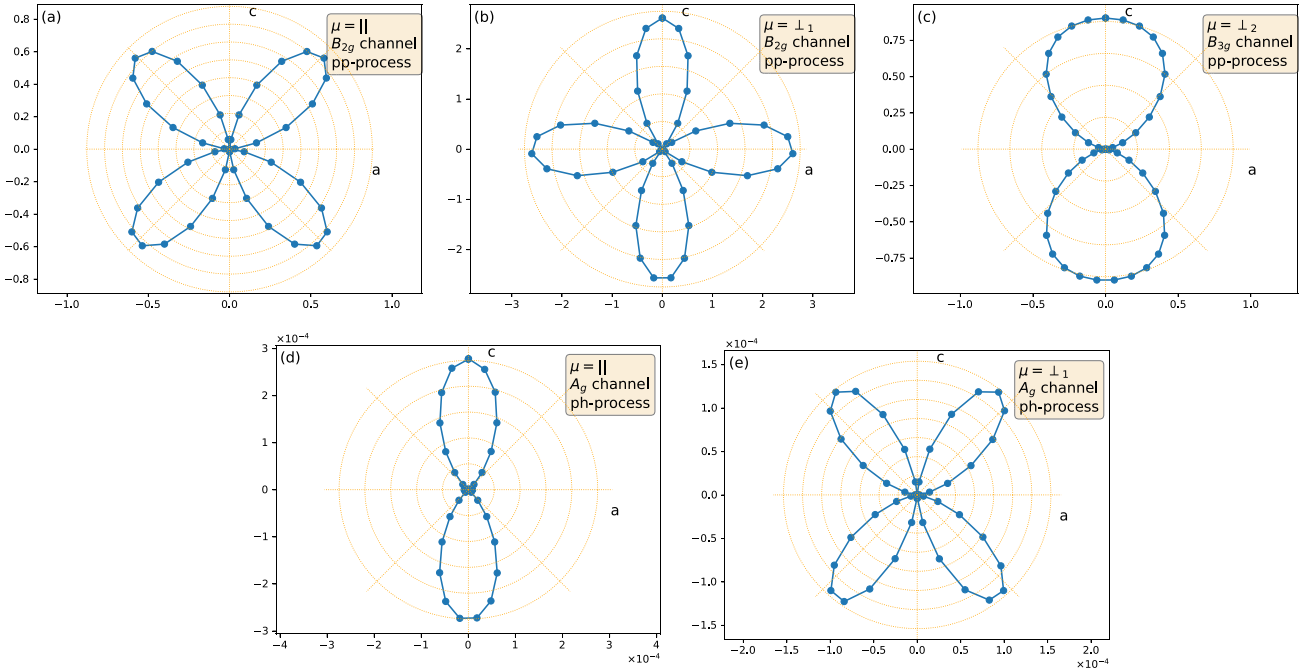


FIG. 6. The angular dependence of sound attenuation coefficient $\alpha_s^\mu(\theta_q, \phi_q = 0^\circ)$ in the ac plane. The pp processes contribute to (a) the attenuation of the longitudinal phonon (α_s^\parallel) in B_{2g} channel, (b) the attenuation of the in-plane transverse phonon ($\alpha_s^{\perp 1}$) in the B_{2g} channel, and (c) the attenuation of the out-of-plane transverse phonon ($\alpha_s^{\perp 2}$) in the B_{3g} channel. The ph processes contribute to (d) the attenuation of the longitudinal phonon (α_s^\parallel) and (e) of the in-plane transverse phonon ($\alpha_s^{\perp 1}$) in the A_g channel. The radius represents the magnitude of α_s^μ in the units of $10^{-5}\rho\delta v$. The calculation is performed at $T = 0.02J$.

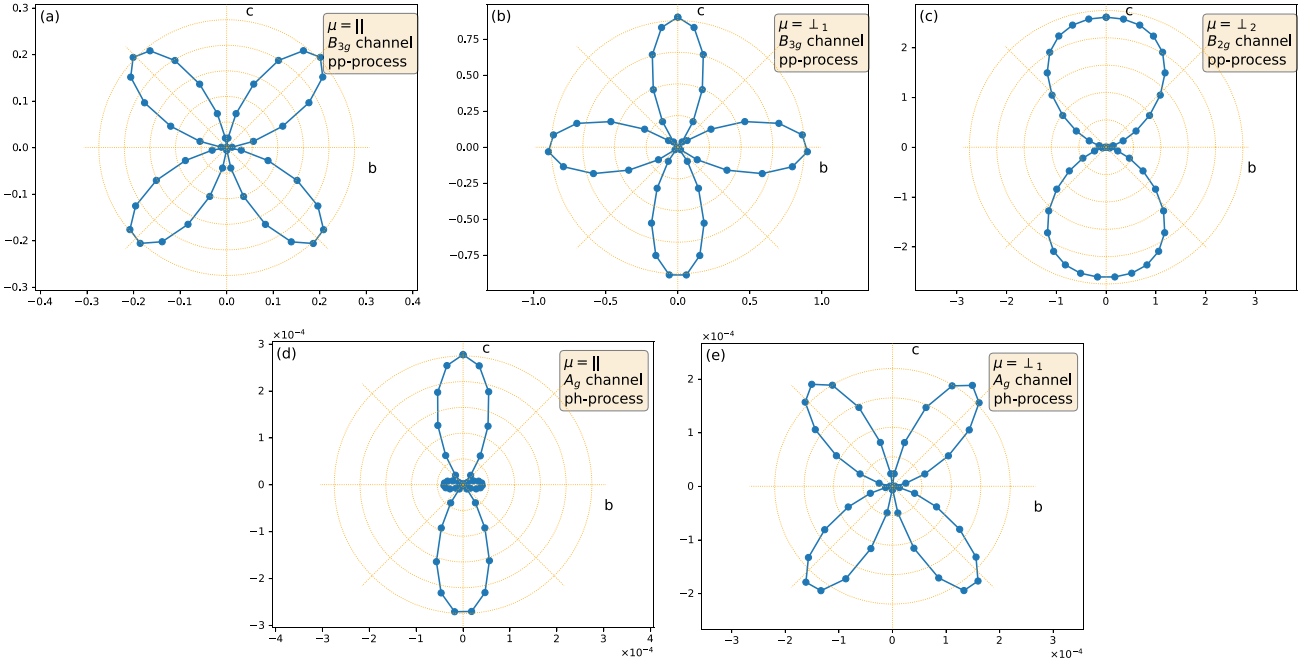


FIG. 7. The angular dependence of sound attenuation coefficient $\alpha_s^\mu(\theta_q, \phi_q = 90^\circ)$ in the bc plane. The pp processes contribute to (a) the attenuation of the longitudinal phonon (α_s^\parallel) in B_{3g} channel, (b) the attenuation of the in-plane transverse phonon ($\alpha_s^{\perp 1}$) in the B_{3g} channel, and (c) the attenuation of the out-of-plane transverse phonon ($\alpha_s^{\perp 2}$) in the B_{2g} channel. The ph processes contribute to (d) the attenuation of the longitudinal phonon (α_s^\parallel) and (e) of the in-plane transverse phonon ($\alpha_s^{\perp 1}$) in the A_g channel. The radius represents the magnitude of α_s^μ in the units of $10^{-5} \rho \delta_V$. The calculation is performed at $T = 0.02 J$.

$\alpha_s^{\perp 2}(\mathbf{q})$. As follows from Eq. (B3), $\alpha_s^{\perp 2}(\mathbf{q})$ has two contributions, one from the B_{2g} channel [Fig. 5(a)], describing the attenuation of lattice vibrations in the ac plane, and from the B_{3g} channel [Fig. 5(b)], describing the attenuation of lattice vibrations in the bc plane, with the former being a bit stronger. The total sound attenuation of the out-of-plane transverse phonon mode $\alpha_s^{\perp 2}(\mathbf{q})$ shown in Fig. 5(c) is the sum of these two contributions, and its angular dependence looks like twofold symmetric four-petal pattern.

Since $v_s^\parallel \lesssim \max(v_F)$ and $v_s^{\perp 1,2} < \max(v_F)$, the ph processes are also allowed. They contribute to the attenuation of the longitudinal phonon mode $\alpha_s^\parallel(\mathbf{q})$ shown in Fig. 5(d) and of the in-plane transverse mode $\alpha_s^{\perp 1}(\mathbf{q})$ shown in Fig. 5(e). According to the form of Majorana-phonon coupling vertices in this geometry given by Eq. (B1) and Eq. (B2), the attenuation of both the longitudinal and the in-plane transverse phonons comes from the dominant A_g and subdominant B_{1g} channels. However, at $T = 0.02 J$ both contributions are very small compared with the one from the pp processes. Thus in Figs. 5(d) and 5(e) we only show the angular dependence of the attenuation computed from the A_g contribution, which displays a vertical dumbbell pattern for $\alpha_s^\parallel(\mathbf{q})$, and the diagonal four-petal pattern for $\alpha_s^{\perp 1}(\mathbf{q})$.

Note, however, that the comparison of the magnitude of the sound attenuation coefficients between ph and pp processes needs to take into consideration the temperature effect [28]. Since the pp processes do not require finite particle number, the low-temperature scaling behavior of its contribution to the sound attenuation coefficient does not depend on temperature, i.e., $\alpha_s^{\text{pp}} \sim T^0$. The ph processes require finite particle occupation, and its low-temperature scaling behavior

is $\alpha_s^{\text{ph}} \sim T^1$ same as in the 2D Kitaev model [27]. This is a direct result of the fact that the low-energy DOS(E) $\sim E^1$ as shown in Fig. 5(c). This low-energy scaling behavior of DOS can also be analytically obtained by evaluating $\text{DOS}(E) = \int_{\text{BZ}} d^3 \mathbf{k} \delta(E - \epsilon_{\mathbf{k}})$. Then at low energy, if we expand the fermionic spectrum around the nodal line, then

$$\text{DOS}(E) = \int_{\text{BZ}} d^3 \mathbf{k} \delta(E - v_F(k_\phi, \delta k_\theta) \delta k_r), \quad (24)$$

where k_ϕ uniquely specifies a point on the nodal line by its orientation ϕ . Around this nodal point, on a neighboring disk locally perpendicular to the nodal line, $(\delta k_\theta, \delta k_r)$ uniquely specifies the \mathbf{k} point that contributes to the DOS. Then the integration Eq. (24) is equivalent to stringing the local disks together along the nodal line. So it is easy to see that the low-energy scaling behavior of DOS(E) is decided by the co-dimension, i.e., the dimension of the BZ space minus the nodal dimension. Thus, the low-energy behavior of $\text{DOS}(E) \sim E^1$ is the same for both 2D plane model and 3D hyperhoneycomb model, so is the low-temperature behavior of the sound attenuation coefficient.

The low-temperature behaviours of both pp- and ph processes distinguish themselves from the attenuation of other interaction channels, such as the channel due to phonon-phonon interactions, which scales as $\sim T^3$ in 2D and $\sim T^5$ in 3D, so they are promising for experimental detection at low enough temperature.

Our numerical calculation shows that, even though the temperature dependence of attenuation from ph process has larger power than that from pp process, pp process still dominates at

high temperatures. The main reason is that Fermi velocities range from 0 to $\max(v_F) = 3J\ell$, so the sound velocities $v_s^{\parallel} = 3J\ell$ and $v_s^{\perp 1,2} \approx 1.6J\ell$, which we use to describe the phonons in β -Li₂IrO₃ compound, are still larger than a significant portion of Fermi velocities, which is consistent with an existence of the nearly-zero Fermi velocities along the nodal line. If we use fictitious smaller sound velocities, the contribution from the ph processes will become larger [28].

Phonon within the ac or bc planes. As shown in Fig. 6 and Fig. 7, the attenuation of the phonons propagating in the ac and bc planes are similar, which is consistent with the crystallographic structure displayed in Fig. 1. Because the Fermi velocities for small deviations \mathbf{k} from the nodal line $\mathbf{K}_0 = (k_a, k_b, 0)$ either into the ac or into the bc planes are small, at low temperatures the pp processes dominate over the ph processes and thus define the angular dependence of the sound attenuation. In both geometries, the pp processes contribute to the attenuation of the phonons with all three different polarizations angular, with similar angular patterns. However, while for the phonon in the ac plane, the strongest attenuation is for the in-plane transverse phonon ($\alpha_s^{\perp 1}$), for the phonon in the bc plane, the strongest attenuation is for out-of-plane transverse polarization ($\alpha_s^{\perp 2}$). In both geometries, attenuation of phonons with out-of-plane transverse polarization only happen through pp processes and displays the vertical dumbbell pattern. The four-petal angular patterns of attenuation of the longitudinal and the in-plane transverse phonons are rotated by 45° with respect to each other. As mentioned before, these distinct patterns directly reflect the spin-phonon couplings from different symmetry channels, probed by different phonon polarization modes. The temperature dependence of the sound attenuation of the phonons propagating in ac or bc planes is similar to that in ab plane.

V. SUMMARY

In this paper, we studied the three-dimensional Kitaev spin-phonon model on the hyperhoneycomb lattice. In this model, the sound attenuation is determined by the decay of a phonon into a pair of Majorana fermions and can be calculated from the imaginary part of the phonon self-energy, which at the lowest order is given by the polarization bubble. Thus, we argued that the phonon attenuation, measurable by the ultrasound experiments, can serve as an effective indirect probe of the spin fractionalization.

In our paper we considered only low temperatures below the flux disordering transition [60], in which only Majorana fermions contribute to the phonon self-energy. We showed

that Majorana semimetal with nodal line band structure leaves distinct characteristic fingerprints in the temperature dependence of the phonon attenuation coefficient as a function of incident phonon momentum. First, it allows the presence of the pp processes of the phonon decay in all three considered scattering geometries with the phonon propagating in one of the three crystallographic planes. Second, since the pp processes of the phonon decay is allowed at all temperatures, the sound attenuation is nonzero even at zero temperature and is almost temperature independent ($\sim T^0$) at lowest temperatures. Combining both pp processes and ph processes that are allowed by symmetry constraints for each scattering geometry and phonon polarization, the temperature dependence of attenuation coefficient can be schematically described by $a_T T^0 + b_T T^1$ with $a_T > b_T$. Thus, the sound attenuation contributed from the decay into fractionalized excitations will be the dominant one at low enough temperatures, distinguishing itself from the contribution due to the phonon-phonon interactions, which scales as $\propto T^5$ in the three-dimensional system. We anticipate that the Z₂ fluxes will play an important role on the phonon dynamics at temperatures above the flux ordering transition temperature. We also obtained that the sound attenuation shows a strong angular dependence at the leading order in phonon momentum q . It is determined by the anisotropic form of the MFPh coupling and the nodal structure of the low-energy fermionic excitations.

Finally, we note that our study was performed for the pure Kitaev model. Of course, real Kitaev materials feature additional weak time-reversal-invariant non-Kitaev interactions, which give rise to other magnetic phases competing with the Kitaev spin liquid. In particular, the minimal spin Hamiltonian for the β -Li₂IrO₃ compound in addition to the Kitaev coupling has contains antiferromagnetic Heisenberg interaction and off-diagonal Γ exchange term [69]. Nevertheless, we believe that the temperature evolution of the sound attenuation will remain similar to the one in the pure Kitaev model as long as these perturbations do not break time reversal symmetry protecting the nodal line [46] and are small enough that the material is in the proximity to the spin liquid phase.

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APPENDIX A: DETAILS OF THE MFPH COUPLING'S DERIVATION

In this Appendix we present the technical details of the derivation of the Majorana fermion-phonon (MFPh) coupling. In the momentum space, the Majorana-phonon coupling Hamiltonian is can be written as

$$\mathcal{H}^c = \sqrt{\frac{2}{N}} \sum_{\mathbf{q}, \mathbf{k}} (\mathcal{H}_{\mathbf{q}, \mathbf{k}}^{A_g} + \mathcal{H}_{\mathbf{q}, \mathbf{k}}^{B_{1g}} + \mathcal{H}_{\mathbf{q}, \mathbf{k}}^{B_{2g}} + \mathcal{H}_{\mathbf{q}, \mathbf{k}}^{B_{3g}}), \quad (\text{A1})$$

where the explicit expressions for the contributions from different symmetry channels are given by

$$\begin{aligned}
\mathcal{H}_{\mathbf{q},\mathbf{k}}^{A_g} &= i\lambda_{A_g} \left(4q_c u_{\mathbf{q},c} \mathbf{A}_{-\mathbf{q}-\mathbf{k}}^\top S_{\mathbf{k}}^\dagger \begin{pmatrix} \hat{O} & -i\hat{\sigma}_3 \\ i\hat{\sigma}_3 & \hat{O} \end{pmatrix} S_{\mathbf{k}} \mathbf{A}_{\mathbf{k}} + (q_c u_{\mathbf{q},c} + 2q_b u_{\mathbf{q},b} + q_a u_{\mathbf{q},a}) \mathbf{A}_{-\mathbf{q}-\mathbf{k}}^\top S_{\mathbf{k}}^\dagger \hat{Q}_{\mathbf{k},1} S_{\mathbf{k}} \mathbf{A}_{\mathbf{k}} \right), \\
\mathcal{H}_{\mathbf{q},\mathbf{k}}^{B_{1g}} &= \frac{i\lambda_{B_{1g}}}{2} (q_a u_{\mathbf{q},b} + q_b u_{\mathbf{q},a}) \mathbf{A}_{-\mathbf{q}-\mathbf{k}}^\top S_{\mathbf{k}}^\dagger \hat{Q}_{\mathbf{k},2} S_{\mathbf{k}} \mathbf{A}_{\mathbf{k}}, \\
\mathcal{H}_{\mathbf{q},\mathbf{k}}^{B_{2g}} &= \frac{i\lambda_{B_{2g}}}{2} (q_a u_{\mathbf{q},c} + q_c u_{\mathbf{q},a}) \mathbf{A}_{-\mathbf{q}-\mathbf{k}}^\top S_{\mathbf{k}}^\dagger \hat{Q}_{\mathbf{k},3} S_{\mathbf{k}} \mathbf{A}_{\mathbf{k}}, \\
\mathcal{H}_{\mathbf{q},\mathbf{k}}^{B_{3g}} &= \frac{i\lambda_{B_{3g}}}{2} (q_b u_{\mathbf{q},c} + q_c u_{\mathbf{q},b}) \mathbf{A}_{-\mathbf{q}-\mathbf{k}}^\top S_{\mathbf{k}}^\dagger \hat{Q}_{\mathbf{k},4} S_{\mathbf{k}} \mathbf{A}_{\mathbf{k}}.
\end{aligned} \tag{A2}$$

Here $S_{\mathbf{k}} = \text{diag}\{e^{i\mathbf{k}\cdot\mathbf{r}_\alpha}\}_{\alpha=C,B,D,A}$ is the diagonal matrix in the sublattice basis, $\hat{O} = \begin{pmatrix} 0 & 0 \\ 0 & 0 \end{pmatrix}$ is the zero 2 by 2 matrix, $\hat{\sigma}_i$ are the auxiliary Pauli matrices, and the explicit expressions for $\hat{Q}_{\mathbf{k}}$ -matrices are given by

$$\begin{aligned}
\hat{Q}_{\mathbf{k},1} &= \begin{pmatrix} (1 + \cos(\mathbf{k} \cdot \mathbf{a}_3))\hat{\sigma}_2 + \sin(\mathbf{k} \cdot \mathbf{a}_3)\hat{\sigma}_1 & \hat{O} \\ \hat{O} & -(\cos(\mathbf{k} \cdot \mathbf{a}_1) + \cos(\mathbf{k} \cdot \mathbf{a}_2))\hat{\sigma}_2 - (\sin(\mathbf{k} \cdot \mathbf{a}_1) + \sin(\mathbf{k} \cdot \mathbf{a}_2))\hat{\sigma}_1 \end{pmatrix}, \\
\hat{Q}_{\mathbf{k},2} &= \begin{pmatrix} (1 + \cos(\mathbf{k} \cdot \mathbf{a}_3))\hat{\sigma}_2 + \sin(\mathbf{k} \cdot \mathbf{a}_3)\hat{\sigma}_1 & \hat{O} \\ \hat{O} & (\cos(\mathbf{k} \cdot \mathbf{a}_1) + \cos(\mathbf{k} \cdot \mathbf{a}_2))\hat{\sigma}_2 + (\sin(\mathbf{k} \cdot \mathbf{a}_1) + \sin(\mathbf{k} \cdot \mathbf{a}_2))\hat{\sigma}_1 \end{pmatrix}, \\
\hat{Q}_{\mathbf{k},3} &= \begin{pmatrix} (1 - \cos(\mathbf{k} \cdot \mathbf{a}_3))\hat{\sigma}_2 - \sin(\mathbf{k} \cdot \mathbf{a}_3)\hat{\sigma}_1 & \hat{O} \\ \hat{O} & (\cos(\mathbf{k} \cdot \mathbf{a}_1) - \cos(\mathbf{k} \cdot \mathbf{a}_2))\hat{\sigma}_2 + (\sin(\mathbf{k} \cdot \mathbf{a}_1) - \sin(\mathbf{k} \cdot \mathbf{a}_2))\hat{\sigma}_1 \end{pmatrix}, \\
\hat{Q}_{\mathbf{k},4} &= \begin{pmatrix} (1 - \cos(\mathbf{k} \cdot \mathbf{a}_3))\hat{\sigma}_2 - \sin(\mathbf{k} \cdot \mathbf{a}_3)\hat{\sigma}_1 & \hat{O} \\ \hat{O} & (\cos(\mathbf{k} \cdot \mathbf{a}_2) - \cos(\mathbf{k} \cdot \mathbf{a}_1))\hat{\sigma}_2 + (\sin(\mathbf{k} \cdot \mathbf{a}_2) - \sin(\mathbf{k} \cdot \mathbf{a}_1))\hat{\sigma}_1 \end{pmatrix}.
\end{aligned} \tag{A3}$$

Next we rewrite \mathcal{H}^c in terms of the transverse and longitudinal eigenmodes as in Eq. (17), where the corresponding MFPh coupling vertices are given by

$$\begin{aligned}
\hat{\lambda}_{\mathbf{q},\mathbf{k}}^{\parallel} &= i\lambda_{A_g} \left(4q_c R_{31} \begin{pmatrix} \hat{O} & -i\hat{\sigma}_3 \\ i\hat{\sigma}_3 & \hat{O} \end{pmatrix} + (q_c R_{31} + 2q_b R_{21} + q_a R_{11}) \hat{Q}_{\mathbf{k},1} \right) \\
&\quad + \frac{i\lambda_{B_{1g}}}{2} (q_a R_{21} + q_b R_{11}) \hat{Q}_{\mathbf{k},2} + \frac{i\lambda_{B_{2g}}}{2} (q_a R_{31} + q_c R_{11}) \hat{Q}_{\mathbf{k},3} + \frac{i\lambda_{B_{3g}}}{2} (q_b R_{31} + q_c R_{21}) \hat{Q}_{\mathbf{k},4}, \\
\hat{\lambda}_{\mathbf{q},\mathbf{k}}^{\perp_1} &= i\lambda_{A_g} \left(4q_c R_{32} \begin{pmatrix} \hat{O} & -i\hat{\sigma}_3 \\ i\hat{\sigma}_3 & \hat{O} \end{pmatrix} + (q_c R_{32} + 2q_b R_{22} + q_a R_{12}) \hat{Q}_{\mathbf{k},1} \right) \\
&\quad + \frac{i\lambda_{B_{1g}}}{2} (q_a R_{22} + q_b R_{12}) \hat{Q}_{\mathbf{k},2} + \frac{i\lambda_{B_{2g}}}{2} (q_a R_{32} + q_c R_{12}) \hat{Q}_{\mathbf{k},3} + \frac{i\lambda_{B_{3g}}}{2} (q_b R_{32} + q_c R_{22}) \hat{Q}_{\mathbf{k},4}, \\
\hat{\lambda}_{\mathbf{q},\mathbf{k}}^{\perp_2} &= i\lambda_{A_g} \left(4q_c R_{33} \begin{pmatrix} \hat{O} & -i\hat{\sigma}_3 \\ i\hat{\sigma}_3 & \hat{O} \end{pmatrix} + (q_c R_{33} + 2q_b R_{23} + q_a R_{13}) \hat{Q}_{\mathbf{k},1} \right) \\
&\quad + \frac{i\lambda_{B_{1g}}}{2} (q_a R_{23} + q_b R_{13}) \hat{Q}_{\mathbf{k},2} + \frac{i\lambda_{B_{2g}}}{2} (q_a R_{33} + q_c R_{13}) \hat{Q}_{\mathbf{k},3} + \frac{i\lambda_{B_{3g}}}{2} (q_b R_{33} + q_c R_{23}) \hat{Q}_{\mathbf{k},4}.
\end{aligned} \tag{A4}$$

Note also that since we are using the long wavelength limit for the phonons, we only kept the leading in q terms in all the expressions.

APPENDIX B: MFPH COUPLINGS IN VARIOUS POLARIZATIONS

For phonon in the ab plane, the rotation matrix is given by $\hat{R} = \begin{pmatrix} \cos \phi_{\mathbf{q}} & -\sin \phi_{\mathbf{q}} & 0 \\ \sin \phi_{\mathbf{q}} & \cos \phi_{\mathbf{q}} & 0 \\ 0 & 0 & 1 \end{pmatrix}$, which simplifies the general expressions for the MFPh coupling vertices to

$$\hat{\lambda}_{\mathbf{q},\mathbf{k}}^{\parallel} = i\lambda_{A_g} q_a R_{11} \hat{Q}_{\mathbf{k},1} + \frac{i\lambda_{B_{1g}}}{2} (q_a R_{21} + q_b R_{11}) \hat{Q}_{\mathbf{k},2}, \tag{B1}$$

$$\hat{\lambda}_{\mathbf{q},\mathbf{k}}^{\perp_1} = i\lambda_{A_g} (2q_b R_{22} + q_a R_{12}) \hat{Q}_{\mathbf{k},1} + \frac{i\lambda_{B_{1g}}}{2} (q_a R_{22} + q_b R_{12}) \hat{Q}_{\mathbf{k},2}, \tag{B2}$$

$$\hat{\lambda}_{\mathbf{q},\mathbf{k}}^{\perp_2} = \frac{i\lambda_{B_{2g}}}{2} q_a R_{33} \hat{Q}_{\mathbf{k},3} + \frac{i\lambda_{B_{3g}}}{2} q_b R_{33} \hat{Q}_{\mathbf{k},4}. \tag{B3}$$

Similarly, for the phonon in the ac plane, the rotation matrix is given by $\hat{R} = \begin{pmatrix} \sin \theta_{\mathbf{q}} & \cos \theta_{\mathbf{q}} & 0 \\ 0 & 0 & 1 \\ \cos \theta_{\mathbf{q}} & -\sin \theta_{\mathbf{q}} & 0 \end{pmatrix}$, so the MFPh coupling vertices are given by

$$\hat{\lambda}_{\mathbf{q},\mathbf{k}}^{\parallel} = i\lambda_{A_g} \left(4q_c R_{31} \begin{pmatrix} \hat{O} & -i\hat{\sigma}_3 \\ i\hat{\sigma}_3 & \hat{O} \end{pmatrix} + (q_c R_{31} + q_a R_{11}) \hat{Q}_{\mathbf{k},1} \right) + \frac{i\lambda_{B_{2g}}}{2} (q_a R_{31} + q_c R_{11}) \hat{Q}_{\mathbf{k},3}, \quad (\text{B4})$$

$$\hat{\lambda}_{\mathbf{q},\mathbf{k}}^{\perp_1} = i\lambda_{A_g} \left(4q_c R_{32} \begin{pmatrix} \hat{O} & -i\hat{\sigma}_3 \\ i\hat{\sigma}_3 & \hat{O} \end{pmatrix} + (q_c R_{32} + q_a R_{12}) \hat{Q}_{\mathbf{k},1} \right) + \frac{i\lambda_{B_{2g}}}{2} (q_a R_{32} + q_c R_{12}) \hat{Q}_{\mathbf{k},3}, \quad (\text{B5})$$

$$\hat{\lambda}_{\mathbf{q},\mathbf{k}}^{\perp_2} = \frac{i\lambda_{B_{1g}}}{2} q_a R_{23} \hat{Q}_{\mathbf{k},2} + \frac{i\lambda_{B_{3g}}}{2} q_c R_{23} \hat{Q}_{\mathbf{k},4}. \quad (\text{B6})$$

For the phonon in the bc plane, the rotation matrix is $\hat{R} = \begin{pmatrix} 0 & 0 & 1 \\ \sin \theta_{\mathbf{q}} & \cos \theta_{\mathbf{q}} & 0 \\ \cos \theta_{\mathbf{q}} & -\sin \theta_{\mathbf{q}} & 0 \end{pmatrix}$, and the MFPh coupling vertices are given by

$$\hat{\lambda}_{\mathbf{q},\mathbf{k}}^{\parallel} = i\lambda_{A_g} \left(4q_c R_{31} \begin{pmatrix} \hat{O} & -i\hat{\sigma}_3 \\ i\hat{\sigma}_3 & \hat{O} \end{pmatrix} + (q_c R_{31} + 2q_b R_{21}) \hat{Q}_{\mathbf{k},1} \right) + \frac{i\lambda_{B_{3g}}}{2} (q_b R_{31} + q_c R_{21}) \hat{Q}_{\mathbf{k},4}, \quad (\text{B7})$$

$$\hat{\lambda}_{\mathbf{q},\mathbf{k}}^{\perp_1} = i\lambda_{A_g} \left(4q_c R_{32} \begin{pmatrix} \hat{O} & -i\hat{\sigma}_3 \\ i\hat{\sigma}_3 & \hat{O} \end{pmatrix} + (q_c R_{32} + 2q_b R_{22}) \hat{Q}_{\mathbf{k},1} \right) + \frac{i\lambda_{B_{3g}}}{2} (q_b R_{32} + q_c R_{22}) \hat{Q}_{\mathbf{k},4}, \quad (\text{B8})$$

$$\hat{\lambda}_{\mathbf{q},\mathbf{k}}^{\perp_2} = \frac{i\lambda_{B_{1g}}}{2} q_b R_{13} \hat{Q}_{\mathbf{k},2} + \frac{i\lambda_{B_{2g}}}{2} q_c R_{13} \hat{Q}_{\mathbf{k},3}. \quad (\text{B9})$$

So in each plane, only two of the four symmetry channels are active. And as shown in the numerical calculations presented in the main text, in the long wavelength limit, one of the two channels dominates over the other. Similar situation was observed in the analysis of the 2D spin-phonon Kitaev model [27].

APPENDIX C: EXPLICIT EXPRESSIONS FOR THE DYNAMICAL FACTORS IN (20)

The dynamic factors in Eq. (21) are evaluated as follows:

$$\begin{aligned} P_{\mathbf{k},11} &= T \sum_{i\omega_m} \frac{1}{(i\Omega_n + i\omega_m) - \varepsilon_{\mathbf{k}}} \frac{1}{i\omega_m - \varepsilon_{\mathbf{k}+\mathbf{q}}} = \frac{n_F(\varepsilon_{\mathbf{k}}) - n_F(\varepsilon_{\mathbf{k}+\mathbf{q}})}{i\Omega_n - \varepsilon_{\mathbf{k}} + \varepsilon_{\mathbf{k}+\mathbf{q}}}, \\ P_{\mathbf{k},22} &= T \sum_{i\omega_m} \frac{1}{(i\Omega_n + i\omega_m) + \varepsilon_{\mathbf{k}}} \frac{1}{i\omega_m + \varepsilon_{\mathbf{k}+\mathbf{q}}} = \frac{n_F(-\varepsilon_{\mathbf{k}}) - n_F(-\varepsilon_{\mathbf{k}+\mathbf{q}})}{i\Omega_n + \varepsilon_{\mathbf{k}} - \varepsilon_{\mathbf{k}+\mathbf{q}}}, \\ P_{\mathbf{k},21} &= T \sum_{i\omega_m} \frac{1}{(i\Omega_n + i\omega_m) + \varepsilon_{\mathbf{k}}} \frac{1}{i\omega_m - \varepsilon_{\mathbf{k}+\mathbf{q}}} = \frac{n_F(-\varepsilon_{\mathbf{k}}) - n_F(\varepsilon_{\mathbf{k}+\mathbf{q}})}{i\Omega_n + \varepsilon_{\mathbf{k}} + \varepsilon_{\mathbf{k}+\mathbf{q}}}, \\ P_{\mathbf{k},12} &= T \sum_{i\omega_m} \frac{1}{(i\Omega_n + i\omega_m) - \varepsilon_{\mathbf{k}}} \frac{1}{i\omega_m + \varepsilon_{\mathbf{k}+\mathbf{q}}} = \frac{n_F(\varepsilon_{\mathbf{k}}) - n_F(-\varepsilon_{\mathbf{k}+\mathbf{q}})}{i\Omega_n - \varepsilon_{\mathbf{k}} - \varepsilon_{\mathbf{k}+\mathbf{q}}}. \end{aligned} \quad (\text{C1})$$

APPENDIX D: VEGAS+ MONTE CARLO INTEGRATION

In this paper, we applied an efficient Monte Carlo algorithm for multidimensional integration Vegas+ [70,71] to evaluate the phase space integration in the polarization bubble Eq. (20). In this section, we will briefly discuss the technical aspect of this algorithm.

Vegas+ is an adaptive stratified sampling algorithm, which is very effective for the integrands with multiple peaks or diagonal nodal (significant) structures. In general, an importance sampling (as in the original Vegas algorithm) is a basic variance reduction technique in Monte Carlo integration, where the probability space is transformed, such that the sampling is concentrated on the important region of the integrand. For example, suppose we need to compute a 1D integral

$$I = \int_a^b f(x) dx. \quad (\text{D1})$$

Different from directly sampling $x \in [a, b]$, as is done in a standard Monte Carlo technique, importance sampling introduces a measurable map from y to x , $x = G^{-1}(y)$, where $y \in [0, 1]$. Then, the integration is equivalently written as

$$I = \int_0^1 f(x(y)) \frac{dx}{dy} dy, \quad (\text{D2})$$

and, instead of uniformly sampling $x \in [a, b]$, one uniformly samples $y \in [0, 1]$. The result of this probability space transformation is such that the distribution of x is described by function $g(x) = G'(x)$ (known from inverse transform sampling). If $g(x)$ is well designed to be of similar shape to $f(x)$, i.e., $g(x)$ is large where $f(x)$ is large, then the x samples will be concentrated in the important region of $f(x)$.

What the Vegas algorithm [72] does is to numerically obtain the map $G^{-1} : y \rightarrow x$, which gives the probability

distribution function $g(x)$, in the following adaptive way. First the x integration space is partitioned into N_p intervals, and Δx_i is the length of each interval (not necessarily uniform). Then the functional form is chosen such that x monotonically increases with y , and within each partition of x , the increase is linear with a rate (Jacobian) J_i , i.e., $\Delta x_i = J_i \Delta y_i$, (again not necessarily uniform). So the measurable map $G^{-1} : y \rightarrow x$ is specified by the set of variables $\{\Delta x_i, J_i\}$, which are under the constraints $\sum_{i=1}^{N_p} \Delta x_i = b - a$, $\sum_{i=1}^{N_p} \Delta y_i = 1$. The objective of designing the distribution function $g(x)$ is to minimize the variation of the integrand (seen as a function of random variable y),

$$\begin{aligned} \sigma_I^2 &= \mathbb{V}_{\text{ar}_{y \in [0,1]}} \left[f(x(y)) \frac{dx}{dy} \right] \\ &= \int_0^1 \left[f(x(y)) \frac{dx}{dy} \right]^2 dy - I^2 \\ &= \sum_i J_i \int_{x_i}^{x_i + \Delta x_i} f(x)^2 dx - I^2, \end{aligned} \quad (\text{D3})$$

where x_i is the left end of each interval partitioned from $x \in [a, b]$. So now designing the map $G^{-1} : y \rightarrow x$ becomes a constrained optimization problem

$$\min_{\{\Delta x_i, J_i\}} \sigma_I^2(\{\Delta x_i, J_i\}). \quad (\text{D4})$$

From here, it is easy to get the necessary optimal condition [70]

$$\frac{1}{\Delta x_i} \int_{x_i}^{x_i + \Delta x_i} J_i^2 f(x)^2 dx = \text{constant}, \quad (\text{D5})$$

i.e., the optimal partition grid $\{x_i\}$ is such that the average of $J_i^2 f(x)^2$ over each interval Δx_i is uniform across the partitions. Without loss of generality, we can introduce uniform grid (partition) in y space, i.e., $\Delta y_i = 1/N_p$. Then, $J_i = \frac{\Delta x_i}{\Delta y_i} = \Delta x_i \cdot N_p$. Then, the objective becomes finding the grid in x space, such that the average of $\Delta x_i^2 f(x)^2$ over Δx_i is uniform, the result of which leads to importance sampling.

The uniform $\Delta x_i^2 f(x)^2$ is achieved by an adaptive numerical algorithm, which can be intuitively understood as follows. First, the average $w_i = \langle \Delta x_i^2 f(x)^2 \rangle_{\Delta x_i}$ on Δx_i is defined to be the weight of the i -th partition. We also define the center weight of all partitions to be $c = \frac{1}{N_p} \sum_i w_i$. Then the uniform weight $\{w_i\}$ condition is equivalent to requiring $\sum_i |w_i - c|^2$ to be minimized. In other words, we have the following optimization problem:

$$L(\{x_i\}) = \min_{\{x_i\}} \min_c \sum_i |w_i - c|^2. \quad (\text{D6})$$

We can easily verify that uniform $\{w_i\}$ is indeed the saddle point solution, i.e., if $\{w_i\}$ is uniform, then $L = 0$. This problem is solved by an alternating optimization algorithm, which alternatively updates $\{x_i\}$ and c in an adaptive procedure [70]. The optimal solution yields a grid of x , which is the most dense in the importance region of the integrand. Thus Vegas is considered an adaptive importance sampling method.

Next, we introduce Vegas+, the enhanced version of vegas with stratified sampling. In the above algorithm, we have obtained the uniform grid of $\{y_i = 1/N_p\}$, so it is natural to stratify the sampling according to this partition. To obtain the optimal number of samples allocated to each stratum $\{n_i\}$, we can minimize the Monte Carlo standard deviation $\sigma_{\text{MC}}^2 = \sum_{i=1}^{N_p} \frac{\sigma_i^2(fJ_i)}{n_i}$ with the constraint $\sum_i n_i = N_{\text{total}}$, where $\sigma_i(fJ_i)$ is the variance of $f(x)J_i$ in the i th partition. This gives that the optimal stratification is $n_i \propto \sigma_i(fJ_i)$. The same optimization method was also applied in the stratified Monte Carlo simulations in the 2D Kitaev QSL [28,56], except that there $\sigma_{\text{MC}}^2 = \sum_i \frac{p_i^2 \sigma_i^2}{n_i}$, where p_i is the normalized probability of the i th partition, and the optimal stratification is $n_i \propto p_i \sigma_i \approx p_i$.

Finally, in this paper the integration was done in 3D \mathbf{k} space with the important region centered around the nodal line. The Vegas algorithm was used to make sure that the samples are concentrated near the 2D plane. But within that 2D plane, partition grid is basically uniform. At this point, the adaptive stratified sampling of Vegas+ was used to assure that the dominant contribution comes from the samples only in the important hypercubes near the nodal line.

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