## Projection of infinite-U Hubbard model and algebraic sign structure

Yunqing Ouyang<sup>1,2</sup> and Xiao Yan Xu<sup>3,4,\*</sup>

<sup>1</sup>State Key Laboratory of Surface Physics, Fudan University, Shanghai 200433, China

<sup>2</sup>Center for Field Theory and Particle Physics, Department of Physics, Fudan University, Shanghai 200433, China

<sup>3</sup>Key Laboratory of Artificial Structures and Quantum Control (Ministry of Education), School of Physics and Astronomy,

Shanghai Jiao Tong University, Shanghai 200240, China

<sup>4</sup>Department of Physics, University of California at San Diego, La Jolla, California 92093, USA

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We propose a projection approach to perform quantum Monte Carlo (QMC) simulation on the infinite-UHubbard model at some integer fillings where either it is sign problem free or surprisingly has an algebraic sign structure—a power law dependence of average sign on system size. We demonstrate our scheme on the infinite-U SU(2N) fermionic Hubbard model on both a square and honeycomb lattice at half filling, where it is sign problem free, and suggest possible correlated ground states. The method can be generalized to study certain extended Hubbard models applying to cluster Mott insulators or two-dimensional Moiré systems; among one of them at certain non-half-integer filling, the sign has an algebraic behavior such that it can be numerically solved within a polynomial time. Further, our projection scheme can also be generalized to implement the Gutzwiller projection to spin basis such that SU(2N) quantum spin models and Kondo lattice models may be studied in the framework of fermionic QMC simulations.

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*Introduction.* The Hubbard model [1–4] provides a starting point to understand physics in strongly correlated electronic systems, such as cuprates [5–9], iron-based superconductors [10,11], and heavy-fermion materials [12,13], as well as ultracold-atom-simulated correlated systems [14] and recently found multiple Morié superlattice systems [15–17]. In those strongly correlated electronic systems, the possible phases of the strong correlation limit are extremely important, outlining possible topology of phase diagrams or serving as mother states to generate more exotic phases.

The infinite-U Hubbard model provides an important perspective on strong correlation physics [18–20]. Solid conclusions can only be made on limited and specific cases, e.g., Nagaoka's theorem [18] applies to low hole density limit of the infinite-U Hubbard model on bipartite lattice and Lieb's theorem [21] imposes constraints on the ground state of the Hubbard model with bipartite hopping at half filling. While numerical methods may provide important hints [19,20], disputes still exist on questions such as which phase is the true ground state of the large-U SU(N) Hubbard model on several lattices [22–26]. Recently, large-U SU(N) Hubbard models are getting more and more attention in ultracold atom simulations [27–34].

Alternatively, an infinite-U Hubbard model can be used as a constraint on local Hilbert space; a typical example is a study of Kondo lattice models, where the local spin is written in terms of fermion operators and a constraint is imposed to restore the local spin Hilbert space by introducing a Hubbard-U term [35,36]. A finite Hubbard-U term plays as a soft constraint, while if U goes to infinity, it becomes an exact constraint, and an elegant form for the SU(2) case is pointed out in Ref. [37].

Inspired by the mentioned former works [35,37], in this Letter we propose a general projection approach, such that controllable large-scale quantum Monte Carlo (QMC) simulations on various infinite-U Hubbard models may be done at some integer fillings. Our scheme can be well demonstrated on the infinite-U SU(2N) fermionic Hubbard models on both square and honeycomb lattices at integer fillings, and the Monte Carlo results are presented at half filling where it is sign problem free. We found the infinite-U SU(4) Hubbard model with Dirac dispersion on square and honeycomb lattices may stabilize a spin liquid (SL) state. We further show how to generalize our scheme to study extended Hubbard models, such as a cluster charge Hubbard model on both square and honeycomb lattices, and obtain possible ground states at half filling. Finally, we apply the projection approach to more exotic SU(4) extended Hubbard models with only an interaction term. It is sign problem free at half filling and it has an SU(4)ferromagnetic ground state. For a certain non-half-integer filling, there is a sign problem, but the average sign happens to be only power law dependence on system size, such that it is also simulatable. This finding inspires a different perspective on finding Monte Carlo simulatable models.

*Projection approach.* We implement our projection approach in the framework of determinant QMC (DQMC) [38]. We illustrate our projection approach through an  $SU(N_f)$  Hubbard model with Hamiltonian  $H = H_t + H_U$  on a general lattice, with the kinetic part  $H_t = -\sum_{ij\alpha} [t_{ij}c_{i,\alpha}^{\dagger}c_{j,\alpha} + \text{H.c.}]$ , and Hubbard interaction part  $H_U = \frac{U}{2} \sum_i (n_i - \nu)^2$ . Here the fermion density operator  $n_i = \sum_{\alpha} n_{i,\alpha}$  at each site is a sum

<sup>\*</sup>xiaoyanxu@sjtu.edu.cn



FIG. 1. Different fermiology considered. (a) Square lattice with uniform hopping. (b) Square lattice with  $\pi$ -flux hopping. We choose a gauge where the solid line denotes  $t_{ij} = t$  and the dashed line denotes  $t_{ij} = -t$ . (c) Honeycomb lattice with uniform hopping. (d) Energy band for (a) with nesting FS, denoted as  $\Box$ -nesting-FS. (e) Energy band for (b) with Dirac dispersion, denoted as  $\Box$ -Dirac. (f) Energy band for (c) with Dirac dispersion, denoted as  $\Box$ -Dirac

over fermion flavor density operator  $n_{i,\alpha} = c_{i,\alpha}^{\dagger} c_{i,\alpha}$  with  $\alpha = 1, \ldots, N_f$ . We focus on the repulsive Hubbard interaction case (U > 0). In DQMC, one starts with partition function  $Z = \text{tr}[e^{-\beta H}]$  and observables  $\langle O \rangle = \text{tr}[O e^{-\beta H}]/Z$ , where  $\beta$ , the inverse temperature, is Trotter decomposed into  $L_{\tau}$  slices, i.e.,  $\beta = L_{\tau} \Delta_{\tau}$ . One needs to further make a Trotter decomposition, i.e.,  $e^{-\Delta_{\tau} H} \approx e^{-\frac{1}{2}\Delta_{\tau} H_i} e^{-\Delta_{\tau} H_i} e^{-\frac{1}{2}\Delta_{\tau} H_i}$ . For the integer-filling infinite-*U* case,  $H_U$  plays the role of constraint on local Hilbert space, i.e., it defines a projection operator, and we observe the following *exact* relation in the infinite-*U* limit [39]:

$$e^{-\frac{\Delta_{\tau}U}{2}(n_{i}-\nu)^{2}}\big|_{U\to+\infty} = \frac{1}{M} \sum_{s_{i}=1}^{M} e^{\frac{i2\pi s_{i}}{M}(n_{i}-\nu)},$$
 (1)

with  $M = \frac{N_t}{2} + |\tilde{v}| + 1$ , and an effective filling  $\tilde{v} \equiv v - \frac{N_t}{2}$ in reference to half filling. As we only focus on integer fillings, the effective filling  $\tilde{v}$  takes values  $\tilde{v} = -\frac{N_t}{2}, -\frac{N_t}{2} + 1, \dots, \frac{N_t}{2}$ , where  $\tilde{v} = 0$  ( $v = \frac{N_t}{2}$ ) corresponds to half filling. The projection is done by introducing a sum over auxiliary fields  $s_i$ . With the above projection operator, the trace over fermions can be easily performed [38,40], and the partition function  $Z = \sum_c w_c$  and observables  $\langle O \rangle = \frac{\sum_c O_c w_c}{\sum_c w_c}$  depend on auxiliary fields  $c = \{s_{i,l}\}$  (*l* is the time slice index), and the sampling over auxiliary fields  $\{s_{i,l}\}$  can be done with Monte Carlo simulation. We note the above finite temperature DQMC scheme can be easily adapted to the zero-temperature projection DQMC [41–43]. As the auxiliary fields here only take finite values, a local update with Metropolis algorithm is efficient. One caution here on the sign problem: for the Hubbard model with bipartite hopping at half filling, one can easily prove the sign problem free [44], and numerical rigorous results can be obtained; for some special cases we find very likely ground state candidates at certain non-half-integer fillings even with the sign problem.

Infinite-USU(2N) Hubbard model on bipartite lattice. We apply the above projection approach to the SU(2N) Hubbard model both on two dimensional (2D) square and honeycomb lattices. We use a zero-temperature projection DQMC, with  $L_{\tau} = 600$  and  $\Delta_{\tau} t = 0.1$ , in our calculation. We have performed about  $0.3 \times 10^4$  warmup sweeps and  $1.2 \times 10^4$ measurement sweeps (grouped into 20 bins) for each parameter running. For a square lattice, we consider two conventional fermionologies-one is the typical uniform nearest neighbor hopping which gives a nesting Fermi surface at half filling (denoted as  $\Box$ -nesting-FS) as shown in Figs. 1(a) and 1(d) and the other is  $\pi$ -flux hopping, which gives Dirac dispersion (denoted as  $\Box$ -Dirac) as shown in Figs. 1(b) and 1(e). For the honeycomb lattice, we consider uniform nearest neighbor hopping which also gives Dirac dispersion (denoted as  $\bigcirc$ -Dirac) as shown in Figs. 1(c) and 1(f). We performed simulations on the infinite-U SU(2N) Hubbard model with the above mentioned fermionology and identify a possible ground state as shown in Table I.

Before we discuss more details of all cases listed in Table I, we consider some analytical arguments. For the SU(2N)Hubbard model on bipartite lattice, one can perform a t/Uexpansion, giving Heisenberg interaction  $J \sum_{\langle ij \rangle} \mathbf{S}_i \cdot \mathbf{S}_j$  at second order, with effective exchange coupling  $J \sim \frac{t^2}{U} > 0$ . Therefore, the SU(2N) Heisenberg model may capture some physics of the SU(2N) Hubbard model at half filling, but be cautioned that they are different and may have different ground states in the infinite-U limit, as the t/U expansion

TABLE I. Possible ground state of infinite-U SU(2N) Hubbard model at half filling on bipartite lattice with different fermiology. In the table,  $U^{\infty}$  denotes infinite U.

$SU(2N)$ - $U^{\infty}$	□-nesting-FS	□-Dirac	⊘-Dirac
<i>SU</i> (2)	Néel	Néel	Néel
SU(4)	Néel	SL?	SL?
$SU(\geq 6)$	VBS	VBS	VBS

will give zero J in that limit. For the SU(2) case, it is well known that the SU(2) Heisenberg model on bipartite lattice has a Néel type ordered ground state. For N larger than a certain value, it has a valence bond solid (VBS) order (also called spin-Peierls state) [24,45–49], and the critical  $N_c$  is estimated about  $2N_c = 4.57(5)$  through QMC calculations [47]. Comparing with our numerics, the SU(2) and  $SU(\ge 6)$  are quite consistent with the SU(2N) Heisenberg model, while the SU(4) case is very special. For  $\Box$ -nesting-FS, Néel type order is favored also for SU(4), but for  $\Box$ -Dirac and  $\bigcirc$ -Dirac, it is very likely that a SL state is stabilized.

In the following, we investigate possible ordered states. One possible order is Néel type spin order. As the generators of  $SU(N_{\rm f})$  can be written as  $S^{\alpha}_{\beta}(\mathbf{r}_i) \equiv c^{\dagger}_{i,\alpha}c_{j,\beta} - \frac{\delta_{\alpha,\beta}}{N_{\rm f}} \sum_{\gamma} c^{\dagger}_{i,\gamma}c_{i,\gamma}$ , the matrix form of the Néel operator can be defined as  $N^{\alpha}_{\beta}(\mathbf{r}_i) \equiv \frac{1}{N_{\rm f}}e^{-i\mathbf{Q}\cdot\mathbf{r}_i}S^{\alpha}_{\beta}(\mathbf{r}_i)$  with  $\mathbf{Q} = (\pi, \pi)$  on square lattice and  $N^{\alpha}_{\beta}(\mathbf{r}_i) \equiv \frac{1}{N_{\rm f}}[S^{\alpha}_{\beta}(\mathbf{r}_i + \tau_1) - S^{\alpha}_{\beta}(\mathbf{r}_i + \tau_2)]$  on honeycomb lattice, where  $\tau_1$  and  $\tau_2$  are inner-cell coordinates of two independent sites of each unit cell of honeycomb lattice. Another possible order is the VBS order, with the gauge invariant bond operator defined as  $B(\mathbf{r}_i) \equiv$  $\frac{1}{N_{\rm f}}e^{-i\mathbf{Q}\cdot\mathbf{r}_i}\sum_{\alpha}\sum_{i,i+\delta}c^{\dagger}_{i,\alpha}c_{i+\delta,\alpha} + \text{H.c., where } i, i+\delta \text{ are a pair of sites of two ends of the nearest neighbor (NN) bond in a fixed}$ direction, with  $\mathbf{Q} = (\pi, 0)$  corresponding to columnar VBS for square lattice and  $\mathbf{Q} = (\frac{2}{3}\pi, \frac{2}{3}\pi)$  for Kekulé type VBS for honeycomb lattice. In the simulation, we measure Néel operator correlations  $C_N(\mathbf{r}_i - \mathbf{r}_j) = \sum_{\alpha\beta} \langle N^{\alpha}_{\beta}(\mathbf{r}_i) N^{\beta}_{\alpha}(\mathbf{r}_j) \rangle \langle N^{\alpha}_{\beta}(\mathbf{r}_i)\rangle\langle N^{\beta}_{\alpha}(\mathbf{r}_i)\rangle$  as well as bond operator correlations  $C_B(\mathbf{r}_i - \mathbf{r}_j) = \langle B(\mathbf{r}_i)B(\mathbf{r}_j) \rangle - \langle B(\mathbf{r}_i) \rangle \langle B(\mathbf{r}_j) \rangle$ . With those correlations, we can extract Néel order parameter  $m_N$  and the VBS order parameter  $m_B$ . The square of the Néel order parameter can be calculated as  $m_N^2 = \frac{1}{L^4} \sum_{i,j} C_N(\mathbf{r}_i - \mathbf{r}_j)$  and the square of the VBS bond order parameter can be calculated as  $m_B^2 = \frac{1}{L^4} \sum_{i,j} C_B(\mathbf{r}_i - \mathbf{r}_j)$ . As shown in Fig. 2, we plot the 1/L extrapolation of the Néel and VBS order parameters for SU(2), SU(4), and SU(6) infinite-U Hubbard models with different fermiology. For SU(2), we have a finite Néel order parameter and, for SU(6), we have a finite VBS order parameter, while, for SU(4), we have a finite Néel order parameter for  $\Box$ -nesting-FS, but for  $\Box$ -Dirac and  $\bigcirc$ -Dirac both the Néel and VBS order are very likely zero in the thermodynamic limit. The softening of static form factors for both spin and bond further rules out any possible trend to ordinary magnetic orders for the infinite-U SU(4) Hubbard model with Dirac dispersion. We conjecture a SL may be stablized here. To further identify the possible ground state of the infinite-U SU(4)Hubbard model with Dirac dispersion, we plot the real space decay of spin-spin and bond-bond correlations as shown in Fig. 3, and found they have algebraic behavior approximately,



FIG. 2. 1/*L* extrapolation of Néel order parameter  $m_N$  and VBS order parameter  $m_B$ . Panels (a) and (b) are for  $\Box$ -nesting-FS, (c) and (d) for  $\Box$ -Dirac, and (e) and (f) for  $\bigcirc$ -Dirac.



FIG. 3. Real space decay of Néel and VBS correlations. Panels (a) and (b) are for  $\Box$ -Dirac and (c) and (d) for  $\bigcirc$ -Dirac. The green solid line is a power law fitting of L = 20 data points (x = 2, ..., 7) for  $\Box$ -Dirac and L = 15 data points (x = 2, ..., 5) for  $\bigcirc$ -Dirac. Due to size limit, a systematic finite size scaling of the power law decay still cannot be realized.

which may indicate an algebraic SL state [9,50-54]. Based on limited system sizes, it is hard to determine precisely the scaling dimensions of spin and bond operators and thus it is hard to tell whether the SL stabilized here is a Dirac SU(4), U(1),  $Z_4$ , or  $Z_2$  SL [9,50-54]. The absence of a Néel or VBS order parameter for the SU(4) infinite-U Hubbard model is also well anticipated from former studies [24,25], where people find VBS order at intermediate U, but the VBS order parameter decreases when people further increase U. Therefore, a transition from VBS to SL is expected, but whether it will happen at a finite larger U or only happen in the infinite-U limit is an open question.

Infinite-USU(2N) extended Hubbard model on bipartite lattice. Further, we apply our projection approach to extended Hubbard models  $H = H_t + H_U$ , where the interaction part is defined as  $H_U = U_p \sum_p (Q_p - v_p)^2$ , where  $Q_p$  is the plaquette charge operator defined on the elemental plaquette of the lattice,  $Q_p = \frac{1}{z} \sum_{i \in p} n_i$ , where the factor *z* is used to normalize the filling of each plaquette as each site is shared by *z* plaquettes: z = 4 for square lattice and z = 3 for honeycomb lattice. Similar to the Hubbard model, we have the following relation in the infinite-*U* limit for the extended Hubbard model [39]:

$$e^{-\Delta_{\tau} U_{p}(Q_{p}-\nu)^{2}}\Big|_{U_{p}\to+\infty} = \frac{1}{M} \sum_{s_{p}=1}^{M} e^{\frac{i2\pi z s_{p}}{M}(Q_{p}-\nu_{p})}, \qquad (2)$$

with  $M = \frac{z\eta N_f}{2} + z|\tilde{v}_p| + 1$ , where  $\eta$  is the effective number of sites per plaquette,  $\eta = 1$  for square lattice, and  $\eta = 2$  for honeycomb lattice. Here  $v_p$  is defined as the filling per plaquette, which is different from the  $\nu$  defined in the on site Hubbard

TABLE II. Ground state of infinite-U SU(2N) extended Hubbard model [denoted as  $SU(2N)-U_p^{\infty}$ ] at half filling on bipartite lattice with different fermiology. For  $N \ge 3$  we have VBS as the ground state for all three cases.

$SU(2N)$ - $U_p^{\infty}$	□-nesting-FS	□-Dirac	⊘-Dirac
<i>SU</i> (2)	Néel	Néel	Néel
SU(4)	Néel	SL?	VBS
$SU(\geq 6)$	VBS	VBS	VBS

model, where  $\nu$  is the filling per site. The cluster charge model is originally motivated to describe magic angle twisted bilayer graphene [55,56], and we will explore the infinite-*U* correlated ground state. Again, we consider several different kinds of fermiology, including  $\Box$ -nesting-FS,  $\Box$ -Dirac, and  $\bigcirc$ -Dirac. The possible ground states are listed in Table II. For *SU*(2) and *SU*( $\geq$  6), the ground states are the same with the Hubbard model as shown in Table I, while for the *SU*(4) case with  $\bigcirc$ -Dirac dispersion, the extended Hubbard favors a Kekulé type VBS order. This is consistent with the large-*U* result in Ref. [57].

Strong coupling SU(2N) extended Hubbard models with assisted hopping term on bipartite lattice. Another interesting application of our projection approach is for strong coupling SU(2N) extended Hubbard models with assisted hopping term on bipartite lattice, where the kinetic part  $H_t$  is turned off,  $H \equiv H_U = U_p \sum_p (\tilde{Q}_p - \nu_p)^2$  with  $\tilde{Q}_p \equiv Q_p + \alpha T_p$ , and where the assisted hopping term  $T_p$  results from topological obstruction when people try to construct an effective real space lattice model for magic angle twisted bilayer graphene (TBG) [58]. We will focus on the SU(4) case on a honeycomb lattice, which is directly related to magic angle TBG. It would be quite interesting to explore the possible ground states at each integer filling, where correlated insulator phases are found almost at all integer fillings ( $\tilde{\nu}_p = 0, \pm 1, \pm 2, \pm 3$ ) [59–61]. We have the following relation to implement the projection [39]:

$$e^{-\Delta_{\tau} U_{p}(\tilde{Q}_{p}-\nu_{p})^{2}}\big|_{U_{p}\to+\infty} = \frac{1}{M} \sum_{s_{p}=1}^{M} e^{\frac{i2\pi z_{p}}{M}(\tilde{Q}_{p}-\nu_{p})}.$$
 (3)

As the kinetic part  $H_t$  is turned off, we only have  $\beta U_p$  as an independent parameter and, in the DQMC simulation, we divide  $\beta U_{\rm p}$  into  $L_{\tau}$  slices  $\beta U_{\rm p} \equiv L_{\tau} \Delta_{\tau} U_{\rm p}$ , and we let  $L_{\tau}$  scale with L,  $L_{\tau} = 10L$ . When we take the infinite- $U_{\rm p}$  limit, it corresponds to the zero temperature properties of the model. For  $\tilde{\nu}_{p} = 0$ , it is sign problem free, as pointed out by one of us [62]. It favors an intervalley coherent state, when the kinetic part  $H_t$  is added back, which breaks SU(4) into two SU(2) for each valley and a valley U(1) [62]. For the strong coupling extended Hubbard model with assisted hopping term, we have full SU(4), and we found an SU(4) ferromagnetic state is stabilized for any finite  $\alpha$ . For other integer fillings, there is a sign problem. We define a reference *bosonic* system with partition function  $Z_b =$  $\sum_{c} |w_{c}|$  [63], and the observables become  $\langle O \rangle = \frac{\langle O \rangle_{b}}{\langle \operatorname{sgn} \rangle_{b}}$ , where  $\langle \cdots \rangle_b$  denotes sampling according to the reference *bosonic* system, i.e.,  $\langle O \rangle_b = \sum_c O_c |w_c|/Z_b$ . In general,  $\langle \text{sgn} \rangle_b$  decays exponentially with system size, as  $\langle \text{sgn} \rangle_b \sim e^{-\beta N_s \Delta f}$  [64], where  $N_s$  is the total number of sites and  $\Delta f$  is the free energy



FIG. 4. Linear system size dependence of average sign (sign). The red dashed line is a power law fitting, getting a  $L^{-2.0(1)}$  dependence. The power law system size dependence of average sign indicating an  $\sim \frac{\ln(\beta N_s)}{\beta N_s}$  dependence of the free energy density difference of the original system and the bosonic reference system.

density difference between the original system (with weight  $w_c$ ) and the reference system (with weight  $|w_c|$ ). However, we found a very interesting phenomenon at  $|\tilde{\nu}_p| = 2$ , i.e., the average sign decays algebraically instead of exponentially with system size as shown in Fig. 4, such that a power law computation complexity is expected and reliable QMC results can be obtained. Our simulation suggests an SU(4) ferromagnetic state, confirming the analytical exact argument in Ref. [58]. We conjecture that the less severe sign problem may come from little fluctuations of the SU(4) ferromagnetic ground state [58]. In detail,  $\langle \text{sgn} \rangle_b = \langle \text{sgn}[\exp(4i\pi z \sum_{p,l} s_{p,l}/M)] \rangle_b$ and the sign problem is mild if  $\sum_{p,l} s_{p,l}$  has little fluctuations due to an SU(4) ferromagnetic ground state which usually has less fluctuations. The algebraic sign behavior also implicitly indicates  $\Delta f$  is  $\sim \frac{\ln(\beta N_s)}{\beta N_s}$  small, such that a faithful mapping of a *fermionic* system to a *bosonic* system in the thermodynamic limit is obtained.

Discussion and conclusion. Our projection approach paves a way to study the infinite-U Hubbard model at integer fillings. We also show how to apply our projection approach to the extended Hubbard model. As the infinite-U Hubbard term is usually used to make a constraint on the local Hilbert space, such as for quantum spin models and Kondo lattice models, our projection approach can be further used to implement those constraints, such that they can be simulated in the framework of fermionic QMC simulations [65].

One important issue of QMC simulations is the sign problem. It is generally believed that, if there is a sign problem, the average sign will decay exponentially with system size. Our finding provides a counterexample, inspiring a different direction to find Monte Carlo simulatable models, where the sign does not have to be always semipositive, as long as its average has an equal or better scaling than algebraic scaling with system size.

We further remark that our projection approach can also be extended to any rational filling, while the price is that we may need a high component of auxiliary fields depending on the filling factors [39]. In addition, we can also use a Hubbard-type term to impose the conservation of total number of particles, such that a hard constrained canonical ensemble Monte Carlo method is obtained, going beyond a soft constrained one [66].

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- [1] M. C. Gutzwiller, Phys. Rev. Lett. 10, 159 (1963).
- [2] J. Hubbard, Proc. R. Soc. London A 276, 238 (1963).
- [3] D. P. Arovas, E. Berg, S. Kivelson, and S. Raghu, arXiv:2103.12097 [cond-mat.str-el].
- [4] M. Qin, T. Schäfer, S. Andergassen, P. Corboz, and E. Gull, arXiv:2104.00064 [cond-mat.str-el].
- [5] P. W. Anderson, Science 235, 1196 (1987).
- [6] V. J. Emery, Phys. Rev. Lett. 58, 2794 (1987).
- [7] E. Dagotto, Rev. Mod. Phys. 66, 763 (1994).
- [8] M. Imada, A. Fujimori, and Y. Tokura, Rev. Mod. Phys. 70, 1039 (1998).
- [9] P. A. Lee, N. Nagaosa, and X.-G. Wen, Rev. Mod. Phys. 78, 17 (2006).
- [10] P. J. Hirschfeld, M. M. Korshunov, and I. I. Mazin, Rep. Prog. Phys. 74, 124508 (2011).
- [11] R. M. Fernandes and A. V. Chubukov, Rep. Prog. Phys. 80, 014503 (2016).
- [12] J. R. Schrieffer and P. A. Wolff, Phys. Rev. 149, 491 (1966).
- [13] D. J. Scalapino, Rev. Mod. Phys. 84, 1383 (2012).

- [14] A. Mazurenko, C. S. Chiu, G. Ji, M. F. Parsons, M. Kanász-Nagy, R. Schmidt, F. Grusdt, E. Demler, D. Greif, and M. Greiner, Nature (London) 545, 462 (2017).
- [15] L. Balents, C. R. Dean, D. K. Efetov, and A. F. Young, Nat. Phys. 16, 725 (2020).
- [16] E. Y. Andrei, D. K. Efetov, P. Jarillo-Herrero, A. H. MacDonald, K. F. Mak, T. Senthil, E. Tutuc, A. Yazdani, and A. F. Young, Nat. Rev. Mater. 6, 201 (2021).
- [17] T. Li, S. Jiang, L. Li, Y. Zhang, K. Kang, J. Zhu, K. Watanabe, T. Taniguchi, D. Chowdhury, L. Fu, J. Shan, and K. F. Mak, Nature (London) 597, 350 (2021).
- [18] Y. Nagaoka, Phys. Rev. 147, 392 (1966).
- [19] L. Liu, H. Yao, E. Berg, S. R. White, and S. A. Kivelson, Phys. Rev. Lett. 108, 126406 (2012).
- [20] A. Tandon, Z. Wang, and G. Kotliar, Phys. Rev. Lett. 83, 2046 (1999).
- [21] E. H. Lieb, Phys. Rev. Lett. 62, 1201 (1989).
- [22] F. F. Assaad, Phys. Rev. B **71**, 075103 (2005).
- [23] A. Paramekanti and J. B. Marston, J. Phys.: Condens. Matter 19, 125215 (2007).

- [24] Z. Zhou, D. Wang, Z. Y. Meng, Y. Wang, and C. Wu, Phys. Rev. B 93, 245157 (2016).
- [25] Z. Zhou, C. Wu, and Y. Wang, Phys. Rev. B 97, 195122 (2018).
- [26] F. H. Kim, F. F. Assaad, K. Penc, and F. Mila, Phys. Rev. B 100, 085103 (2019).
- [27] S. Taie, R. Yamazaki, S. Sugawa, and Y. Takahashi, Nat. Phys. 8, 825 (2012).
- [28] Z. Cai, H.-h. Hung, L. Wang, D. Zheng, and C. Wu, Phys. Rev. Lett. 110, 220401 (2013).
- [29] C. Hofrichter, L. Riegger, F. Scazza, M. Höfer, D. R. Fernandes, I. Bloch, and S. Fölling, Phys. Rev. X 6, 021030 (2016).
- [30] H. Ozawa, S. Taie, Y. Takasu, and Y. Takahashi, Phys. Rev. Lett. 121, 225303 (2018).
- [31] S. Taie, E. Ibarra-García-Padilla, N. Nishizawa, Y. Takasu, Y. Kuno, H.-T. Wei, R. T. Scalettar, K. R. Hazzard, and Y. Takahashi, arXiv:2010.07730 [cond-mat.quant-gas].
- [32] E. Ibarra-García-Padilla, S. Dasgupta, H.-T. Wei, S. Taie, Y. Takahashi, R. T. Scalettar, and K. R. A. Hazzard, Phys. Rev. A 104, 043316 (2021).
- [33] D. Tusi, L. Franchi, L. F. Livi, K. Baumann, D. B. Orenes, L. Del Re, R. E. Barfknecht, T. Zhou, M. Inguscio, G. Cappellini et al., arXiv:2104.13338 [cond-mat.quant-gas].
- [34] E. Altman, K. R. Brown, G. Carleo, L. D. Carr, E. Demler, C. Chin, B. DeMarco, S. E. Economou, M. A. Eriksson, K.-M. C. Fu, M. Greiner, K. R. A. Hazzard, R. G. Hulet, A. J. Kollár, B. L. Lev, M. D. Lukin, R. Ma, X. Mi, S. Misra, C. Monroe, K. Murch, Z. Nazario, K.-K. Ni, A. C. Potter, P. Roushan, M. Saffman, M. Schleier-Smith, I. Siddiqi, R. Simmonds, M. Singh, I. B. Spielman, K. Temme, D. S. Weiss, J. Vučković, V. Vuletić, J. Ye, and M. Zwierlein, PRX Quantum 2, 017003 (2021).
- [35] F. F. Assaad, Phys. Rev. Lett. 83, 796 (1999).
- [36] M. Raczkowski and F. F. Assaad, Phys. Rev. Res. 2, 013276 (2020).
- [37] S. Capponi and F. F. Assaad, Phys. Rev. B 63, 155114 (2001).
- [38] R. Blankenbecler, D. J. Scalapino, and R. L. Sugar, Phys. Rev. D 24, 2278 (1981).
- [39] See Supplemental Material (SM) at http://link.aps.org/ supplemental/10.1103/PhysRevB.104.L241104 for more details of the projection approach and the calculations.
- [40] F. Assaad and H. Evertz, in *Computational Many-Particle Physics*, Lecture Notes in Physics Vol. 739, edited by H. Fehske, R. Schneider, and A. Weiße (Springer, Berlin, 2008), pp. 277–356.
- [41] G. Sugiyama and S. Koonin, Ann. Phys. (NY) 168, 1 (1986).
- [42] S. Sorella, S. Baroni, R. Car, and M. Parrinello, Europhys. Lett. (EPL) 8, 663 (1989).
- [43] S. Sorella, E. Tosatti, S. Baroni, R. Car, and M. Parrinello, Int. J. Mod. Phys. B 02, 993 (1988).

- [44] C. Wu and S.-C. Zhang, Phys. Rev. B 71, 155115 (2005).
- [45] N. Read and S. Sachdev, Phys. Rev. B 42, 4568 (1990).
- [46] K. Harada, N. Kawashima, and M. Troyer, Phys. Rev. Lett. 90, 117203 (2003).
- [47] K. S. D. Beach, F. Alet, M. Mambrini, and S. Capponi, Phys. Rev. B 80, 184401 (2009).
- [48] T. C. Lang, Z. Y. Meng, A. Muramatsu, S. Wessel, and F. F. Assaad, Phys. Rev. Lett. 111, 066401 (2013).
- [49] Z.-X. Li, Y.-F. Jiang, S.-K. Jian, and H. Yao, Nat. Commun. 8, 314 (2017).
- [50] M. Hermele, T. Senthil, M. P. A. Fisher, P. A. Lee, N. Nagaosa, and X.-G. Wen, Phys. Rev. B 70, 214437 (2004).
- [51] M. Hermele, T. Senthil, and M. P. A. Fisher, Phys. Rev. B 72, 104404 (2005).
- [52] X. Y. Xu, Y. Qi, L. Zhang, F. F. Assaad, C. Xu, and Z. Y. Meng, Phys. Rev. X 9, 021022 (2019).
- [53] X.-Y. Song, Y.-C. He, A. Vishwanath, and C. Wang, Phys. Rev. X 10, 011033 (2020).
- [54] V. Calvera and C. Wang, arXiv:2103.13405 [cond-mat.str-el].
- [55] H. C. Po, L. Zou, A. Vishwanath, and T. Senthil, Phys. Rev. X 8, 031089 (2018).
- [56] X. Y. Xu, K. T. Law, and P. A. Lee, Phys. Rev. B 98, 121406(R) (2018).
- [57] Y. Da Liao, Z. Y. Meng, and X. Y. Xu, Phys. Rev. Lett. 123, 157601 (2019).
- [58] J. Kang and O. Vafek, Phys. Rev. Lett. 122, 246401 (2019).
- [59] Y. Cao, V. Fatemi, A. Demir, S. Fang, S. L. Tomarken, J. Y. Luo, J. D. Sanchez-Yamagishi, K. Watanabe, T. Taniguchi, E. Kaxiras *et al.*, Nature (London) 556, 80 (2018).
- [60] M. Yankowitz, S. Chen, H. Polshyn, Y. Zhang, K. Watanabe, T. Taniguchi, D. Graf, A. F. Young, and C. R. Dean, Science 363, 1059 (2019).
- [61] X. Lu, P. Stepanov, W. Yang, M. Xie, M. A. Aamir, I. Das, C. Urgell, K. Watanabe, T. Taniguchi, G. Zhang *et al.*, Nature (London) 574, 653 (2019).
- [62] Y. D. Liao, J. Kang, C. N. Breiø, X. Y. Xu, H.-Q. Wu, B. M. Andersen, R. M. Fernandes, and Z. Y. Meng, Phys. Rev. X 11, 011014 (2021).
- [63] If  $w_c$  is a complex number, we use  $|\text{Re}(w_c)|$ .
- [64] M. Troyer and U.-J. Wiese, Phys. Rev. Lett. 94, 170201 (2005).
- [65] X. Y. Xu et al. (unpublished).
- [66] Z. Wang, F. F. Assaad, and F. Parisen Toldin, Phys. Rev. E 96, 042131 (2017).
- [67] J. Towns, T. Cockerill, M. Dahan, I. Foster, K. Gaither, A. Grimshaw, V. Hazlewood, S. Lathrop, D. Lifka, G. D. Peterson, R. Roskies, J. Scott, and N. Wilkins-Diehr, Comput. Sci. Eng. 16, 62 (2014).