Third-order polarizability of interlayer excitons in heterobilayers

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(Received 22 June 2021; revised 6 October 2021; accepted 11 November 2021; published 29 November 2021)

In this paper, we employ a fully microscopic approach to the study of interlayer excitons in heterobilayers. We use Fowler's and Karplus' method to access the dynamical polarizability of noninteracting interlayer excitons in a WSe_2/WS_2 -based van der Waals heterostructure. Following from the calculation of the linear polarizability, we consider Svendsen's variational method to the calculation of the dynamic third-order polarizability. With this variational method, we study both two-photon absorption and third-harmonic generation processes for interlayer excitons in a WSe_2/WS_2 heterobilayer, discussing the various selection rules of intraexcitonic energy level transitions.

DOI: 10.1103/PhysRevB.104.205433

I. INTRODUCTION

The advent of the study of two-dimensional layers with atomic thickness has renewed the interest in transition-metal dichalcogenides (TMDs) and their optical/optoelectronic properties [1]. These materials, which have been studied in their bulk form since the 1960's [2,3], have been shown to be good candidates for various optical and optoelectronic applications [4–7].

The optical response of these materials is dominated by excitons [8], and optically bright exciton absorption peaks have been shown to correspond to the excitation of states in the *ns* series [9,10]. As for *np*-series excitons, these can be controlled magnetically despite being optically dark in TMDs [11]. The selection rules of excitons in TMDs for absorption experiments have been recently thoroughly studied [12].

The linear dielectric response of TMDs consists of two distinct regimes: the interband regime, where electrons from the valence band are excited to the conduction band, leaving being a hole [6,13]. The attractive electrostatic interaction between this newly formed electron-hole pair leads to the formation of a bound state (exciton), and its *ns* states are optically active and can be observed in absorption measurements [14]. The second regime, characterized by intraexciton transitions, consists of the transition between the excitonic ground state (1s) and the empty *np* states [15,16]. Each of the $1s \rightarrow np$ transitions [17–19] is characterized by a peak in the dynamical polarizability which, in turn, determines the dielectric response of the system in a pump-probe experiment [20,21].

Interlayer excitons are formed when two monolayers are brought together with type-II band alignment, where the conduction band minimum and the valence band maximum are located in different layers [22,23]. This allows the intralayer electron-hole pairs to tunnel into interlayer excitons (Fig. 1, left), which have a significantly longer lifetime due to the small overlap of the individual electron and hole wave functions [24,25]. Additionally, interlayer excitons exhibit luminescence at lower energies than their intralayer counterpart [26] which, together with their lower binding energies, allows for an easier identification of the specific species in question in polarizability measurements [21]. Recently, Merkl *et al.* [21] were able to observe the transition between interlayer and intralayer excitonic phases in van der Waals heterostructures. This observation was performed via the measurement of the linear dielectric function of the excitons in a pump-probe experiment (Fig. 1, right), observing a significant shift in maximum of the optical conductivity.

The excitonic energy levels are known to depend on whether the system is confined or not (for example, the Rydberg series for the two-dimensional Hydrogen atom is strongly affected by radial confinement [27]). Following recent advancements in the fabrication of TMD-based metamaterials with atomic precision [28,29], we focus our attention on radially confined interlayer excitons in a WSe_2/WS_2 bilayer.

From studying the dynamic polarizability of the excitonic states, the full dielectric response of the excited sample can be extracted. This transition from polarizability to dielectric response is one from the microscopic (individual exciton) to the macroscopic (when many excitons form a low-density exciton gas). The susceptibility of this exciton gas would be given, in a simplified form, as $\chi_{ex}^{inter} = \frac{N_X \alpha}{\epsilon_q}$, with N_X the exciton density in the two-dimensional material, α the linear polarizability, and ϵ_0 the vacuum dielectric constant. Separating the real and imaginary part of the susceptibility, these can be experimentally measured [20,21,30,31] and, therefore, would provide a good comparison point for the results we obtain for the linear polarizability.

The nonlinear polarizabilities of the hydrogen atom have been thoroughly studied throughout the years [32,33], providing a good starting point for the study of these same nonlinear processes in excitons. Recently, the nonlinear optical effects from intraexcitonic transitions have been the focus



FIG. 1. (Left) Interlayer exciton in a van der Waals heterostructure. (Right) Diagram of pump-probe experiments for observing np (dark) excitonic states.

of many theoretical and experimental works. This has been performed for several TMDs, including but not limited to WSe_2 [34,35] and WS_2 [36,37], as well as hBN homostructures [38]. Through doping processes, the optical nonlinear properties of TMDs can be tuned [39] which, along with their large nonlinear optical coefficients [12], increases their feasibility for practical applications [40] (e.g., as optical modulators).

We begin this paper by discussing the necessary modifications to the Rytova-Keldysh potential [41,42] when dealing with interlayer excitons, followed by analyzing a harmonic oscillator approximation to this potential which will motivate the chosen variational wave function. After defining this *ansatz*, we outline the complete set of basis functions we will use to approximate the excited states of the exciton. The necessary material-dependent set of parameters characteristic of the specific van der Waals heterostructure in question will also be discussed, obtaining the desired energy levels.

In Sec. III we outline Fowler's and Karplus' method [43,44] to compute the dynamical polarizability of noninteracting excitons. We then apply this same method to interlayer excitons in a WSe₂/WS₂-based van der Waals heterostructure, comparing the obtained resonances against the transitions calculated both numerically and variationally via a finite-basis approach.

In Sec. IV, we turn to the nonlinear response of the excitons [45,46]. We focus our discussion on the third-order polarizability, first outlining a variational procedure based on the ideas of both Karplus and Svendsen [43,47,48] for the calculation of the intraexcitonic third-order polarizability. This procedure was recently studied by Henriques et al., applied to the study of the two-photon absorption for excitons in WSe₂ [49]. After outlining this variational method, we look at the two-photon absorption for excitons in WSe₂/WS₂, discussing the various intraexcitonic transitions observed. Finally, we consider the third-harmonic generation, first with an extremely small broadening as to clearly differentiate each individual resonance, and afterwards with two much larger broadenings as to ascertain the feasibility of experimental detection of each peak. A diagrammatic representation of the various transitions is also presented for both the two-photon absorption and third-harmonic generation processes. We finish the paper with our closing remarks in Sec. V.

II. INTERLAYER RYTOVA-KELDYSH POTENTIAL

We begin this section by discussing the Hamiltonian in which we focus our attention. This will lead to the discussion of the necessary modifications to the Rytova-Keldysh potential when dealing with interlayer excitons. We define the basis for our variational approach to the excitonic states (starting by the *np*-series states for Sec. III and then the *ns* and *nd* series for Sec. IV). We finish this section by comparing the various variational energies for the excitonic ground state against the result from numerical integration of the Schrödinger equation in a log-grid [50–52] via the Numerov shooting method.

A. Model Hamiltonian in the dipole approximation

Let us now begin by considering the following Hamiltonian (in atomic units, as will be used throughout this paper) in the dipole approximation

$$H = H_0 - \mathbf{r} \cdot \mathbf{F}(t) = -\frac{1}{2\mu} \nabla^2 + V(r) - \mathbf{r} \cdot \mathbf{F}(t), \quad (1)$$

where μ is the reduced mass of the electron-hole system, ∇^2 is the Laplacian operator (taken in polar coordinates throughout this paper), V(r) is a potential energy term, and $\mathbf{F}(t)$ is an external time-dependent field. This external field will be initially ignored as to first outline the variational methods and wave functions that will be used, and its action will be studied in Secs. III and IV, when the methods to obtain both the linear and the third-order dynamical polarizability are described, respectively.

Considering this same electron-hole system in a van der Waals (vdW) heterostructure, their potential energy is accurately modeled by the Rytova-Keldysh potential [41,42]

$$V_{RK}(r) = -\frac{\pi}{2r_0} \left[\mathbf{H}_0 \left(\kappa \frac{r}{r_0} \right) - \mathbf{Y}_0 \left(\kappa \frac{r}{r_0} \right) \right], \qquad (2)$$

with κ the mean dielectric constant of the media, r_0 an intrinsic parameter of the 2D material (interpretable as an in-plane screening length), and \mathbf{H}_0 , \mathbf{Y}_0 the Struve-H and Bessel-Y (second kind) special functions of zeroth order, respectively. The polarizability of intralayer excitons is discussed in detail in Refs. [49,53].

When considering interlayer excitons, a minimum separation originating from the physical distance between the two layers appears in the Rytova-Keldysh potential. The interlayer modified Rytova-Keldysh potential then reads [54]

$$V_{i-RK}(r) = -\frac{\pi}{2r_0} \left[\mathbf{H}_0 \left(\kappa \frac{\sqrt{r^2 + d^2}}{r_0} \right) - \mathbf{Y}_0 \left(\kappa \frac{\sqrt{r^2 + d^2}}{r_0} \right) \right],$$
(3)

with d this interlayer separation distance.

Presenting the same Coulomb tail behavior as the Rytova-Keldysh potential [Eq. (2)] at large distances, the finite interlayer separation d eliminates the logarithmic divergence at the origin. The absence of this divergence makes it so that an inverse-exponential *ansatz* for the excitonic wave function [55] is not the most adequate. This, together with the parabolic nature of the interlayer modified Rytova-Keldysh potential near r = 0, motivates the *ansatz* that will be obtained in the following section.

B. Gaussian-based ansatz

Ahead we introduce a Gaussian wave function to describe the ground state of the interlayer 1s state analytically. In order to motivate this choice, we show below that the potential Eq. (3) can, indeed, be approximated by a parabolic potential near r = 0.

In this regime, Eq. (3) can be expanded up to second order in *r* as

$$V(r) \approx -V_0 + \gamma r^2, \tag{4}$$

where

$$V_{0} = \frac{\pi}{2r_{0}} \left[\mathbf{H}_{0} \left(\kappa \frac{d}{r_{0}} \right) - \mathbf{Y}_{0} \left(\kappa \frac{d}{r_{0}} \right) \right],$$

$$\gamma = -\frac{\pi\kappa}{4dr_{0}^{2}} \left[\mathbf{H}_{1} \left(\kappa \frac{d}{r_{0}} \right) - \mathbf{Y}_{-1} \left(\kappa \frac{d}{r_{0}} \right) \right].$$
(5)

Considering this potential, the Hamiltonian will be

$$H = -\frac{\hbar^2}{2\mu}\nabla^2 - V_0 + \gamma r^2.$$

For the harmonic oscillator it is well known [56–58] that the 1s function is proportional to a Gaussian, that is, we have

$$\Psi(\eta) \propto e^{-\eta/2},$$

where $\eta = \mu \omega r^2 / \hbar$ and $\omega = \sqrt{2\gamma / \mu}$. As such, an *ansatz* based on this solution is given by

$$\psi(r) = \mathcal{C}e^{-\frac{r^2}{2\beta}},\tag{6}$$

where C is a normalization constant and β is a variational parameter. As we are interested in studying interlayer excitons confined to a finite radius, a multiplicative factor is added to impose boundary conditions,

$$\psi_0(r) = Ce^{-\frac{r^2}{2\beta}}(R-r),$$
 (7)

where *R* is the radius of the enclosure, considered R = 1200 throughout this paper (in atomic units). This radius is considered sufficiently large such that the variational wave functions become zero significantly before the boundary is reached (clearly seen in Fig. 2, where the wave functions are already zero at around $\sim R/3$).

Similarly to the *ansatz* defined by Pedersen in [55], we modify Eq. (7) into a sum of two Gaussians, given by

$$\psi_0(r) = \mathcal{C}(e^{-\frac{r^2}{2a}} + be^{-\frac{r^2}{2c}})(R-r), \tag{8}$$

which will the be *ansatz* of the excitonic ground state we will consider throughout this paper. Additionally, a visual comparison of Eq. (7) against Eq. (8), as well as the approximate wave function obtained via the method described in Sec. II C, is present in Fig. 2. Clearly the *ansatz* (8) performs much better than the *ansatz* (7).

C. Variational approach from boundary conditions

In certain conditions, namely for a small interlayer separation d, the obtained eigenvalues from minimization of the Hamiltonian

$$H_0 = -\frac{1}{2\mu} \nabla^2 + V_{i-RK}(r)$$
(9)



FIG. 2. Radial profile of the double-Gaussian ansatz [Eq. (8)], the Gaussian ansatz [Eq. (7)], and the finite-basis approximation (Sec. II C) of the ground-state excitonic wave function.

with the *ansatz* from Eq. (8) are not significantly different from those obtained from numerical integration (a small difference is present for the present case in Table II). Additionally, obtaining higher energy states involves orthogonalization against Eqs. (7) or (8), followed by normalization, a process whose complexity increases substantially when increasing the number of excited states [59].

A different approach to obtaining variational solutions to the Schrödinger equation with the interlayer modified Rytova-Keldysh is by integrating the Hamiltonian matrix elements for a complete basis whose elements obey the necessary boundary conditions. Truncating this basis and diagonalizing the Hamiltonian matrix, we obtain a set of variational wave functions that converge towards the real wave functions of the system as the basis size increases.

Following Ref. [60], a basis of functions can be constructed from the solutions of the circular infinite well, given by

$$\psi_{n,l}(r,\varphi) = \frac{C_{n,l}}{\sqrt{2\pi}} e^{il\varphi} J_l\left(z_{n,l}\frac{r}{R}\right),\tag{10}$$

where $J_l(x)$ is the Bessel function of the first kind of order l, $z_{n,l}$ is the *n*th zero of $J_l(z)$, and $C_{n,l}$ is a normalization constant given by

$$C_{n,l} = \sqrt{\frac{2}{R^2 J_{|l|+1}^2(z_{n,l})}}$$

As both $V_{RK}(r)$ and $V_{i-RK}(r)$ are invariant under rotations, the quantum number l is well defined. As such, we can integrate the Hamiltonian and diagonalize it for a finite number of functions with fixed l. Knowing this, we will start the discussion by outlining the necessary material dependent parameters for a WSe₂/WS₂ van der Waals heterostructure.

D. Material dependent parameters in WSe₂/WS₂

To perform the necessary calculations and obtain the wave functions and energy eigenvalues, we must substitute the material-specific parameters characteristic of a WSe₂/WS₂ van der Waals (vdW) heterostructure [21]. The reduced mass of interlayer excitons in this vdW heterostructure is $\mu = 0.15$

TABLE I. In-plane, out-of-plane, and total relative permittivities of WSe_2 and WS_2 .

	WSe ₂	WS_2
$\overline{\epsilon_{(n)}^{\perp}}$	7.5	6.3
$\epsilon_{(n)}^{\parallel}$	13.36	11.75
$\epsilon_{(n)}$	10.01	8.604

and, as such, the effective Bohr radius will be $a_0 \approx 0.581$ Å. The thickness of each layer is considered $d_{\text{mat}} = 5.7$ Å and the gap distance is $d_{\text{gap}} = 1$ Å. As such, the effective interlayer distance will be

$$d_{\rm eff} = \frac{1}{2}d_{\rm mat} + d_{\rm gap} + \frac{1}{2}d_{\rm mat} = 6.7$$
 Å. (11)

The permittivities for each material are given by

$$\epsilon_{(n)} = \sqrt{\epsilon_{(n)}^{\perp} \epsilon_{(n)}^{\parallel}}, \qquad (12)$$

where $\epsilon_{(n)}^{\parallel}$ and $\epsilon_{(n)}^{\perp}$ are the in- and out-of-plane relative permittivities of the material *n*. The numerical values of each component of the relative permittivities in each material is present in Table I.

Knowing the permittivities in each material, the average permittivity of the vdW heterostructure is

$$\kappa_{\rm mat} = \frac{\epsilon_{\rm WSe_2} + \epsilon_{\rm WS_2}}{2} = 9.31$$

Regarding the screening length r_0 , this parameter can be separated as a sum for each layer (as described in Ref. [54])

$$r_0 \to r_0^{(1)} + r_0^{(2)},$$

where $r_0^{(n)}$ denotes the screening length for the layer *n*. The screening length for each individual layer can be obtained as [20]

$$r_0^{(n)} = \frac{\epsilon_n^2 - 1}{2\epsilon_n} d_n \sqrt{\frac{\epsilon_n^{\parallel}}{\epsilon_n^{\perp}}},\tag{13}$$

where d is the thickness of the layer. As the in- and out-ofplane relative permittivities of the two materials are known, the total screening length is

$$r_0 = 70.73$$
 Å.

Choosing a basis size of 120 and the same enclosure radius as in Sec. II B (R = 1200), the ground-state energy eigenvalue is (in both atomic units and meV)

$$E_{1,0} = -0.00143378 = -39.0152$$
 meV.

A comparison against both variational *ansatze* [Eqs. (7) and (8)] and the numerical results from considering the shooting method in a log-grid is given in Table II.

III. DYNAMICAL VARIATIONAL METHOD FOR LINEAR POLARIZABILITY

In this section we will briefly outline Fowler's and Karplus' [43,44] method to compute the dynamical polarizability of various systems. This discussion will be along the same lines as in Ref. [53], serving as a quick outline of the procedure

TABLE II. Comparison of the different estimates for the excitonic ground-state energy eigenvalue for a WSe_2/WS_2 heterostructure. The interlayer distance is given in Eq. (11).

	1s state
Finite basis	-0.00143378
Shooting method	-0.00144375
Double-Gaussian [Eq. (8)]	-0.00142879
Gaussian [Eq. (7)]	-0.00136314

as some results will be necessary further ahead. We apply this method to calculate the linear polarizability of interlayer excitons in a WSe₂/WS₂ circular dot, comparing the obtained peaks with the calculated $1s \rightarrow np$ transitions.

Looking back to Eq. (1), we consider (without loss of generality) the external field pointing along the -x direction. Closely following the approach presented in Ref. [53], first outlined by Fowler and Karplus [43,44], the time-dependent Schrödinger equation reads

$$[H_0 + xF(t)]|\psi(t)\rangle = i\frac{\partial}{\partial t}|\psi(t)\rangle, \qquad (14)$$

where $|\psi(t)\rangle$ describes the wave function of the system in the presence of the electric field **F**. The dynamical polarizability is defined as

$$\alpha(\omega) = -\langle \psi_0 | x | \psi_1^+ \rangle - \langle \psi_0 | x | \psi_1^- \rangle \tag{15}$$

with ψ_0 the Gaussian *ansatz* of Eq. (8) and ψ_1^{\pm} the excited *p*-series states, such that

$$E(\omega) = E_0 - \frac{1}{2}\alpha(\omega)F^2.$$
 (16)

As we are working on a finite disk of radius *R*, Bessel functions of the first kind are an appropriate complete set of functions to describe a problem in such a geometry, as defined in Eq. (10). Using these functions as a basis, we write $\psi_1^{\pm}(\mathbf{r})$ as

$$\psi_1^{\pm}(\mathbf{r}) = \cos\theta \sum_{n=1}^N c_n^{\pm} J_1\left(\frac{z_{1,n}r}{R}\right),\tag{17}$$

where $J_1(z)$ is the Bessel function of the first kind of order 1, $z_{1,n}$ is the *n*th zero of $J_1(z)$, *N* is the number of Bessel functions we choose to use, and $\{c_n^{\pm}\}$ are a set of coefficients yet to be determined. As proposed in Refs. [43,61–63], the values of $\{c_n^{\pm}\}$ are determined from the minimization of the functional

$$\mathcal{I}_{\pm} = \int d\mathbf{r} \,\psi_{1,n}^{\pm}(\mathbf{r})[H_0 - E_0 \pm \hbar\omega]\psi_{1,n}^{\pm}(\mathbf{r}) + 2 \int d\mathbf{r} \,\psi_{1,n}^{\pm}(\mathbf{r})r\cos\theta\psi_0(\mathbf{r}).$$
(18)

The discussion of the minimization of the \mathcal{J} functional and the computation of the c_n^{\pm} coefficients is performed in Appendix A.

The dynamical polarizability is then computed by substituting Eq. (17) into Eq. (15), together with the c_n^{\pm} coefficients

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FIG. 3. Real (blue) and imaginary (red) parts of the normalized polarizability for interlayer excitons in a WSe₂/WS₂ heterostructure. The vertical black lines represent the excitonic $1s \rightarrow np$ transitions.

discussed in Appendix A, being written as

$$\alpha(\omega) = -g_v \pi \sum_{n=1}^{N} (c_n^+ + c_n^-) \\ \times \int_0^R J_1\left(\frac{z_{1n}r}{R}\right) r \,\psi_0(r) r \,dr,$$
(19)

where $g_v = 2$ is the valley degeneracy. The ω dependence on the right hand side is present in the coefficients c_n^{\pm} [more explicitly, this dependence is present in the M matrix Eq. (A10) from the $g_i^{\pm}(\omega)$ term, as defined in Eq. (A9)]. We remark that choosing for the ground state wave function a variational function allows us to write Eq. (19) exclusively in terms of analytical functions. This considerably simplifies the numerical minimization.

The polarizability for interlayer excitons in a WSe₂/WS₂ heterostructure is plotted in Fig. 3, normalized by its value at the maximum for the excitonic $1s \rightarrow 2p$ transition (with $E_{1s\rightarrow 2p} \sim 29.0576$ meV) for easier visualization. Addition-

ally, a brief comparison of the linear polarizability for interlayer and intralayer excitons is made in Appendix B. Having outlined and obtained the linear polarizability in Fig. 3, we now turn our attention to the third-order polarizability of interlayer excitons in this same vdW heterostructure.

IV. SVENDSEN'S METHOD FOR THE THIRD-ORDER POLARIZABILITY

Having outlined Fowler's and Karplus' method in Sec. III, we will now consider a similar method for the third-order polarizability, following the procedure delineated by Svendsen [47,48] and recently applied by Henriques *et al.* [49] to intralayer excitons in TMDs.

Maintaining the definition for the c^{\pm} coefficients given in Eq. (A10), and following closely [43,48,49], we write the functional

$$\mathcal{K} = \langle \xi_{\alpha\beta}(\omega_a, \omega_b) | H_0 - E_0 + \omega_a + \omega_b | \xi_{\alpha\beta}(\omega_a, \omega_b) \rangle + \langle \xi_{\alpha\beta}(\omega_a, \omega_b) | \mathbf{d}_\beta | \psi_\alpha(\omega_a) \rangle + \langle \psi_\alpha(\omega_a) | \mathbf{d}_\beta | \xi_{\alpha\beta}(\omega_a, \omega_b) \rangle$$
(20)

where **d** is the dipole operator defined previously, E_0 the ground-state energy, H_0 the unperturbed Hamiltonian [defined in Eq. (9)] and $|\psi_{\alpha}\rangle$ the basis obtained via the \mathbf{c}^{\pm} coefficients and the Bessel functions $J_1(z_{1,n})$ in Eq. (17). This functional is then minimized with respect to $|\xi_{\alpha\beta}\rangle$, defined analogously to Eq. (17) as

$$\xi_{\alpha\beta}(\omega_a,\omega_b;r) = \sum_{n=1}^{+\infty} \sum_{|l|=0}^{+\infty} \zeta_{l,n}^{\alpha,\beta}(\omega_a,\omega_b) J_l\left(z_{l,n}\frac{r}{R}\right) \frac{e^{il\theta}}{\sqrt{2\pi}},$$
(21)

where we will focus our attention on the *xx* component of $\xi_{\alpha\beta}$. The discussion of minimization of this \mathcal{K} functional, done in a similar fashion as for the \mathcal{J} functional, is performed in depth in Appendix C.

Having defined both $\zeta_0^{x,x}$ and $\zeta_2^{x,x}$ in Eqs. (C9)–(C12), respectively, the third-order susceptibility follows from Ref. [49] (with $\omega_{\sigma} = \omega_1 + \omega_2 + \omega_3$) as

$$\chi_{x,x,x,x}^{(3)}(-\omega_{\sigma};\omega_{1},\omega_{2},\omega_{3}) = \frac{1}{3!} \mathcal{P} \left\{ -\mathbf{c}^{\dagger}((-\omega_{\sigma})^{\dagger}) \cdot \left[\mathbb{T}_{0} \cdot \zeta_{0}^{x,x}(-\omega_{2},-\omega_{3}) + \mathbb{T}_{2} \cdot \zeta_{2}^{x,x}(-\omega_{2},-\omega_{3}) \right] + \left[\mathbf{S} \cdot \mathbf{c}(-\omega_{\sigma}) \right] \times \left[\sum_{n=1}^{N} c_{n}^{\dagger}((-\omega_{2})^{\dagger})c_{n}(\omega_{1})2\pi \frac{R^{2}}{2} [J_{2}(\mathbf{z}_{1,n})]^{2} \right] \right\},$$
(22)

with \mathcal{P} denoting the different permutations of the frequencies $(-\omega_{\sigma}; \omega_1, \omega_2, \omega_3)$, $\mathbb{T}_{0/2}$ is defined in Eq. (C3), and S is the vector in Eq. (A4). Having defined Eq. (22), we will now utilize it to compute two different regimes for the third-order polarizability of interlayer excitons in WSe₂/WS₂.

A. Two-photon absorption in WSe₂/WS₂

We will now compute the *xxxx* component of the two-photon absorption third-order susceptibility, defined as $\chi_{xxxx}^{TPA}(\omega) = \chi_{xxxx}^{(3)}(-\omega; \omega, -\omega, \omega)$. As the two-photon absorption was considered in [49], we will not discuss it as in depth

as we will with the third-harmonic generation. Similarly to what was done in Fig. 3, we plot the real and imaginary parts of the two-photon absorption third-order susceptibility for a broadening small enough (0.05 meV) such that each individual peak can be clearly observed. This plot is visible in Fig. 4,



FIG. 4. (Left) Real (blue) and imaginary (red) parts of the two-photon absorption third-order susceptibility (normalized by the peak value for the $1 \rightarrow 3d$ transition at ~17.5 meV) as a function of the incident photons' energies for a broadening $\Delta = 0.05$ meV and basis size N = 120. The resonances correspond to $1s \rightarrow ns$ transitions (black dashed lines) and to $1s \rightarrow nd$ transitions (dark orange dashed lines). The dominant $1s \rightarrow 2p$ transition is at ~29.0576 meV, outside of the frequency range plotted in this figure. (Right) Diagrammatic representation of the $1s \rightarrow ns$ (blue arrows) and $1s \rightarrow nd$ (red arrows) transitions. Topmost black line represents the resonance when $\hbar\omega = E_{1s\rightarrow 2p}$. We remark that the red and blue arrows represent the selection rules of the system, that is, the only allowed transitions.

normalized by the peak value for the $1 \rightarrow 3d$ transition at ~16.6 meV, along with a diagram representing the $1s \rightarrow ns$ and $1s \rightarrow nd$ transitions.

B. Third-harmonic generation

To finalize this paper, we will now compute the *xxxx* component of the third-order susceptibility third-harmonic generation of interlayer excitons in WSe₂/WS₂, which is defined as $\chi_{xxxx}^{THG}(3\omega) = \chi_{xxxx}^{(3)}(-3\omega; \omega, \omega, \omega)$. Following the same approach as in Sec. IV A, we will begin by computing $\chi_{xxxx}^{THG}(3\omega)$ for a small enough broadening such that each

resonance is clear. The left-hand side of Fig. 5 has been normalized by the peak value for the resonance at $3\hbar\omega = E_{1s\to 2p}$, while the right-hand side features a diagram representing the various transitions.

Spectral broadening as low as $\Delta \sim 2 \text{ meV}$ (i.e., the broadening that was considered for Fig. 3) can be achieved for low temperature encapsulated systems [64]. Considering much higher values of the broadening than in Fig. 5, namely $\Delta = 1 \text{ meV}$ and $\Delta = 2 \text{ meV}$, we recompute $\chi_{xxxx}^{\text{THG}}(3\omega)$, obtaining Fig. 6.

Although some $1s \rightarrow np$ resonances are already apparent at $\Delta = 2 \text{ meV}$, namely at $\sim 40 \text{ meV}$ $(1s \rightarrow 5p)$ and



FIG. 5. (Left) Real (blue) and imaginary (red) parts of the third-harmonic generation third-order susceptibility (normalized by the peak value for the $1 \rightarrow 2p$ resonance at ~29.0576, meV) as a function of the incident photons' energies for a broadening $\Delta = 0.05$ meV and basis size N = 120. The resonances correspond to $1s \rightarrow np$ transitions (black dashed lines), $1s \rightarrow ns$ transitions (green dashed lines), and $1s \rightarrow nd$ transitions (dark-orange dashed lines). The ~87.1727 meV resonance (black dot-dashed line) corresponds to the response at frequency $3\hbar\omega = 3E_{1s\rightarrow 2p}$. (Right) Diagrammatic representation of the various transitions visible in the plot on the left-hand side. The colors of the arrows represent the different transitions: $1s \rightarrow np$ (black), $1s \rightarrow ns$ (green), and $1s \rightarrow nd$ (dark-orange), while the number of arrows represents the frequency of the transition: 3 arrows for $3\hbar\omega = E_{i\rightarrow f}$, 2 arrows for $2\hbar\omega = E_{i\rightarrow f}$, and 1 arrow for $\hbar\omega = E_{i\rightarrow f}$. Again, as in Fig. 4, the arrows represent the selection rules associated with the third-harmonic generation (THG) nonlinear process.



FIG. 6. Real (left) and imaginary (right) parts of the normalized third-harmonic generation third-order susceptibility for different broadening values: $\Delta = 2 \text{ meV}$ (solid lines) and 1 meV (dashed lines). The basis size is fixed at N = 120. The resonances correspond to $1s \rightarrow np$ transitions (black dashed lines), $1s \rightarrow ns$ transitions (green dashed lines) and $1s \rightarrow nd$ transitions (dark orange dashed lines). The ~87.1727 meV resonance (black dot-dashed line) corresponds to the response at $3\hbar\omega = 3E_{1s\rightarrow 2p}$.

~60 meV $(1s \rightarrow 8p)$, most $1s \rightarrow ns$ and $1s \rightarrow nd$ transitions only become clearly visible in the sub-meV regime. At those broadenings, the most clear peaks are at $3\hbar\omega \approx 50$ meV $(1s \rightarrow 3d)$ and at $3\hbar\omega \approx 56$ meV $(1s \rightarrow 4s)$. The two resonances at ~72 meV and ~75 meV are joined into one at $\Delta =$ 1 meV, although their presence is noticeable from comparison with the adjacent $1s \rightarrow np$ transition peaks.

The $1s \rightarrow np$ resonance peaks occur for $3\hbar\omega = E_{1s\rightarrow np}$, while the $1s \rightarrow ns$ and $1s \rightarrow nd$ peaks are situated at $3\hbar\omega = \frac{3}{2}E_{1s\rightarrow ns/nd}$. Additionally, the peak at $3\hbar\omega = 87.1727$ meV corresponds to $3\hbar\omega = 3E_{1s\rightarrow 2p}$. This is due to the response at both 2ω [65] and ω also being present in this process, described in detail in Sec. 3.2 of Ref. [66], i.e., $2\hbar\omega = E_{1s\rightarrow ns/nd}$ and $\hbar\omega = E_{1s\rightarrow 2p}$. No other peaks for $\hbar\omega = E_{1s\rightarrow np}$ are visible, as they would already be outside the considered frequency domain ($3E_{1s\rightarrow 3p} \sim 105.872$ meV, barely outside of the plot).

V. CONCLUSIONS

In this paper, we discussed interlayer excitons in a WSe_2/WS_2 -based van der Waals heterostructure, for which recent experimental measurements of the polarizability through pump-probe experiments [21] have been performed. In these experiments, the transition from bright to dark states has been accessed for both interlayer and intralayer excitons. In this paper, we focused ourselves solely on the transitions between intraexcitonic states for interlayer excitons in a circular WSe₂/WS₂ dot.

Approximating the interlayer modified Rytova-Keldysh potential in a power series, we motivate a Gaussian variational *ansatz* for the ground state of the system. After discussing the necessary material-dependent parameters, we consider both this *ansatz* and a complete set of basis functions the Bessel functions of the first kind to estimate the excitonic wave functions through variational methods. We apply Fowler's and Karplus' method variational method to access the linear polarizability of two-dimensional interlayer excitons, focusing on the transition from bright to dark states (more specifically, $1s \rightarrow np$ transitions) in a circular dot of a WSe₂/WS₂-based van der Waals heterostructure with Dirichlet boundary conditions.

For the linear regime we consider a broadening of $\Delta = 2$ meV, allowing for the clear observation of the multiple

peaks in the excitonic polarizability. We observe an almost perfect agreement between the frequency of each resonance and the energy differences between the ground state and the excited states of the exciton. These energy differences were calculated both numerically (via the Numerov shooting method) and variationally (via both a Gaussian *ansatz* for the ground-state and a finite basis of Bessel function of the first kind), allowing for a greater confidence in their values. As expected [21], comparing the linear polarizability of interlayer excitons in the heterobilayer against intralayer excitons in each of the individual layers shows that the binding energies are much lower in the interlayer regime [Figs. 3 and 7, Table II, and Eq. (B2)].

Onto the third-order polarizability, we began by providing an in-depth discussion of Svendsen's variational method [47,48]. After outlining this method and analyzing the required modifications due to the necessity of orthogonality between the various states, we work on the algebraization of the method, arriving at a purely vectorial problem after the necessary integrals are computed. The obtained expression is then tested first with the two-photon absorption and then with third-harmonic generation.

Starting by two-photon absorption, a process which was discussed in-depth by Henriques *et al.* [49], we perform the necessary calculations with a broadening small enough such that the resonance associated with each individual intraexcitonic energy level transition can be clearly identifiable ($\Delta = 0.05 \text{ meV}$). The two-photon absorption third-order susceptibility is then normalized by the value at the $E_{1s \rightarrow 3d}$ resonance as to facilitate the comparison of each peak. Of note is the amplitude of the resonance associated with the $E_{1s \rightarrow 7d}$ transition, peaking at around 40% of the maximum and being dramatically larger than the adjacent transitions. Additionally, the individual selection rules for this system were identified.

Finally turning to the third-harmonic generation in the third-order susceptibility, we first perform the calculations for a very small broadening ($\Delta = 0.05 \text{ meV}$). Each individual transition was clearly identifiable, with resonances at $\hbar\omega = E_{i\rightarrow f}$ and $3\hbar\omega = E_{i\rightarrow f}$ for $1s \rightarrow np$ transitions, and at $2\hbar\omega = E_{i\rightarrow f}$ for both $1s \rightarrow ns$ and $1s \rightarrow nd$ transitions. The individual selection rules for the third-harmonic generation process were clearly identifiable.

Considering experimentally obtainable values of the broadening, we compare the third-harmonic generation third-order susceptibility for $\Delta = 1$ meV and 2 meV. We argue for the feasibility of experimentally observing the three well-resolved resonance peaks, associated with $1s \rightarrow 2p$, $1s \rightarrow 3d$, and $1s \rightarrow 4s$ transitions, where the latter two could be an experimental indication of third-harmonic generation due to intraexcitonic transitions in this specific van der Waals heterostructure.

ACKNOWLEDGMENTS

M.F.C.M.Q. acknowledges the International Nanotechnology Laboratory for the Quantum Portugal Initiative Grant No. SFRH/BD/151114/2021. N.M.R.P acknowledges support by the Portuguese Foundation for Science and Technology (FCT) in the framework of the Strategic Funding UIDB/04650/2020. J.C.G.H. acknowledges the Center of Physics for a grant funded by the UIDB/04650/2020 strategic project. N.M.R.P. also acknowledges support from the European Commission through the project Graphene-Driven Revolutions in ICT and Beyond (Ref. No. 881603, CORE 3), COMPETE 2020, PORTUGAL 2020, FEDER and the FCT through projects POCI-01-0145-FEDER- 028114, POCI-01-0145-FEDER-02888 and PTDC/NAN-OPT/29265/2017.

APPENDIX A: VARIATIONAL COEFFICIENTS FOR THE LINEAR POLARIZABILITY

Recalling the orthogonality relation of Bessel functions on a finite disk of radius R [67],

$$\int_0^R J_\nu \left(\frac{z_{\nu m} r}{R}\right) J_\nu \left(\frac{z_{\nu n} r}{R}\right) r dr = \frac{R^2}{2} \delta_{nm} [J_{|\nu|+1}(z_{\nu m})]^2,$$
(A1)

with z_{vm} the *m*th zero of $J_{\nu}(z)$, one easily shows that the functional can be rewritten as

$$\begin{aligned} \mathcal{J}_{\pm} &= \pi \sum_{n=1}^{\infty} \sum_{k=1}^{\infty} c_n^{\pm} c_k^{\pm} \mathcal{I}_{kn} + 2\pi \sum_{n=1}^{\infty} c_n^{\pm} \mathcal{S}_n \\ &+ \frac{\pi R^2}{2} \sum_{n=1}^{\infty} (c_n^{\pm})^2 \bigg[\frac{\hbar^2}{2\mu} \frac{z_{1n}^2}{R^2} - E_0 \pm \hbar \omega \bigg] [J_2(z_{1n})]^2, \end{aligned}$$
(A2)

where \mathcal{I}_{kn} and \mathcal{S}_n refer to the following integrals involving two and one Bessel functions, respectively,

$$\mathcal{I}_{kn} = \int_0^R J_1\left(\frac{z_{1k}r}{R}\right) V_{\mathrm{RK}}(r) J_1\left(\frac{z_{1n}r}{R}\right) r \, dr, \qquad (A3)$$

$$S_n = \int_0^R J_1\left(\frac{z_{1n}r}{R}\right) r \,\psi_0^{\text{fin.}}(r) r dr. \tag{A4}$$

Differentiating \mathcal{J}_{\pm} with respect to the different c_n^{\pm} coefficients, one finds

$$c_{j}^{\pm} \left\{ \frac{R^{2} [J_{2}(z_{1j})]^{2}}{2} \left[\frac{z_{1j}^{2}}{2\mu R^{2}} - E_{0} \pm \omega \right] + \mathcal{I}_{jj} \right\} + \sum_{n \neq j}^{N} c_{n}^{\pm} \mathcal{I}_{jn} = -\mathcal{S}_{j},$$
(A5)

with $j \in \{1, 2, ..., N\}$. This equation defines a linear system of equations whose solution determines the values of the coefficients c_n^{\pm} .

We can write Eq. (A5) in a more concise manner, using matrix notation, as

$$\mathbb{M} \cdot \mathbf{c}^{\pm} = -\mathbf{S},\tag{A6}$$

where \mathbf{c}^{\pm} and \mathbf{S} are column vectors defined as

$$[\mathbf{c}^{\pm}]^{\mathrm{T}} = (c_{1}^{\pm}, c_{2}^{\pm}, \dots, c_{N}^{\pm}), \quad \mathbf{S}^{\mathrm{T}} = (\mathcal{S}_{1}, \mathcal{S}_{2}, \dots, \mathcal{S}_{N}), \quad (A7)$$

and \mathbb{M} is a $N \times N$ matrix with:

$$(\mathbb{M})_{ij} = g_i^{\pm}(\omega)\delta_{ij} + \mathcal{I}_{ij}, \qquad (A8)$$

where δ_{ij} is the Kronecker delta and

$$g_i^{\pm}(\omega) = \frac{R^2 [J_2(z_{1i})]^2}{2} \left[\frac{z_{1i}^2}{2\mu R^2} - E_0 \pm \omega \right].$$
(A9)

After \mathbb{M} and **S** are computed using Eqs. (A3) and (A9) and Eq. (A4), respectively, the coefficients that determine $\psi_1^{\pm}(\mathbf{r})$ are readily obtained as

$$\mathbf{c}^{\pm} = -\mathbf{M}^{-1} \cdot \mathbf{S},\tag{A10}$$

and the solution of the differential equation is found.

APPENDIX B: LINEAR POLARIZABILITY FOR INTERLAYER AND INTRALAYER EXCITONS

In this Appendix we will make a quick comparison between the linear polarizability of both interlayer and intralayer excitons in a WSe₂/WS₂ heterobilayer. Intralayer excitons are considered for both the WSe₂ and the WS₂ layers, with the relevant material-dependent parameters. Additionally, due to the Coulomb-like divergence of the potential at r = 0, we consider the more appropriate variational *ansatz*, defined by Pedersen [55], given by

$$\psi_0(r) = \mathcal{C}(e^{-ar} - be^{-\gamma ar})(R - r). \tag{B1}$$

Minimizing this functional, the ground-state energy for excitons each layer will be

$$E_{WS_2} = -0.00438405 = -119.296 \text{ meV},$$

 $E_{WS_2} = -0.00541680 = -147.399 \text{ meV},$ (B2)

both closely matching the energies obtained via the finite basis approach from Sec. II C. These binding energies are also substantially larger than the ground-state energy for interlayer excitons in the heterobilayer (Table II), as expected from experimental studies [21]. Computing the linear polarizability with the formalism described in Appendix A, normalized by the value at the $1s \rightarrow 2p$ peak, we display the results in Fig. 7.

APPENDIX C: VARIATIONAL COEFFICIENTS FOR THE THIRD-ORDER POLARIZABILITY

Defining $g_n^{(l)}(\omega)$ as a more general form of Eq. (A9)

$$g_n^{(l)}(\omega) = \frac{R^2 [J_{|l|+1}(z_{l,n})]^2}{2} \left[\frac{z_{l,n}^2}{2\mu R^2} - E_0 + \omega \right]$$



FIG. 7. Imaginary part of the normalized linear polarizability for interlayer excitons in a WSe_2/WS_2 heterostructure and intralayer excitons in both a WS_2 and a WSe_2 monolayers. The peaks correspond to transitions between the 1*s* state (ground-state) and the excited *np* states.

and $\mathcal{V}_{nm}^{(l)}$ analogously to Eq. (A3)

$$\begin{aligned} \mathcal{V}_{nm}^{(l)} &= \left\langle J_l \left(z_{l,n} \frac{r}{R} \right) \middle| V_{i-RK}(r) \middle| J_l \left(z_{l,m} \frac{r}{R} \right) \right\rangle \\ &= \int_0^R J_l \left(z_{l,n} \frac{r}{R} \right) V_{i-RK} J_l \left(z_{l,m} \frac{r}{R} \right) r dr, \end{aligned}$$

the functional of Eq. (20) can then be rewritten using the both the c_n^{\pm} and the $\zeta_{l,n}$ coefficients as

$$\begin{aligned} \mathcal{K} &= \sum_{n=1}^{+\infty} \sum_{|l|=0}^{+\infty} \left[\zeta_{l,n} \right]^{\dagger} g_{n}^{(l)} (\omega_{a} + \omega_{b}) \zeta_{l,n} \\ &+ \sum_{n,m=1}^{+\infty} \sum_{|l|=0}^{+\infty} \left[\zeta_{l,n} \right]^{\dagger} \mathcal{V}_{nm}^{(l)} \zeta_{l,m} \\ &+ \frac{1}{2} \sum_{n,m=1}^{+\infty} \left\{ \left[\zeta_{0,n} \right]^{\dagger} \mathcal{T}_{nm}^{l=0} \cdot (c_{m}^{+} + c_{m}^{-}) \right. \\ &+ \left[\zeta_{|l|=2,n} \right]^{\dagger} \left(\mathcal{T}_{nm}^{l=2} \cdot c_{m}^{+} + \mathcal{T}_{nm}^{l=-2} \cdot c_{m}^{-} \right) + \text{c.c.} \right\}. \end{aligned}$$

In Eq. (C1), \mathcal{T}_{nm}^{l} is the dipole transition amplitude between the functions of the basis, denoted by $\langle \psi_{\alpha}(\omega_{a}) | \mathbf{d}_{\beta} | \xi_{\alpha\beta}(\omega_{a}, \omega_{b}) \rangle$ in Eq. (20) and given by

$$\mathcal{T}_{nm}^{l} = \left\langle J_{l}\left(z_{l,n}\frac{r}{R}\right) \middle| r \left| J_{1}\left(z_{1,m}\frac{r}{R}\right) \right\rangle$$
$$= \int_{0}^{R} J_{l}\left(z_{l,n}\frac{r}{R}\right) r J_{1}\left(z_{1,m}\frac{r}{R}\right) r dr.$$

The solution of this integral can be written in a closed form by applying the more general expression

$$\int_{0}^{1} J_{\nu}(\alpha r) r J_{\nu+1}(\beta r) r dr$$

= $\frac{\alpha J_{\nu+1}(\alpha)}{(\alpha^{2} - \beta^{2})^{2}} [-2\beta J_{\nu}(\beta) + (\alpha^{2} - \beta^{2}) J_{\nu+1}(\beta)],$ (C2)

valid as long as $J_{\nu}(\alpha) = 0$. Comparing with the definition of \mathcal{T}_{nm}^{l} , Eq. (C2) can be simplified further as $J_{\nu+1}(\beta) = 0$. Performing the necessary changes of variable, and substituting $J_{\nu+1}(\beta)$ into Eq. (C2), \mathcal{T}_{nm}^{l} can be written as

$$\mathcal{T}_{nm}^{(l=0)} = -2R^3 \frac{z_{0,n} z_{1,m} J_1(z_{0,n}) J_0(z_{1,m})}{\left(z_{0,n}^2 - z_{1,m}^2\right)^2},$$

$$\mathcal{T}_{nm}^{(l=2)} = -2R^3 \frac{z_{1,n} z_{2,m} J_2(z_{1,n}) J_1(z_{2,m})}{\left(z_{1,n}^2 - z_{2,m}^2\right)^2}.$$
 (C3)

As done in the Appendix of Ref. [49], we separate the l = 0and the l = 2 cases. Looking first at the more complicated case of l = 0, we must ensure that $|\xi_{\alpha\beta}\rangle$ is orthogonal to the considered ground-state [(in our case, the *ansatz* of Eq. (8)]. As such, to ensure this, we must have

$$\langle \psi_0 | \xi_{\alpha\beta} \rangle = \sum_{n=1}^{+\infty} \zeta_{0,n}^{\alpha,\beta}(\omega_a, \omega_b) \Big\langle \psi_0 \Big| J_0 \Big(z_{0,n} \frac{r}{R} \Big) \Big\rangle = 0, \quad (C4)$$

where nonzero angular momentum terms vanish upon angular integration due to the isotropic nature of the ground-state *ansatz*. Separating the first term of the sum, this condition is given by

$$\begin{split} & \zeta_{0,1}^{\alpha,\beta}(\omega_a,\omega_b) \Big\langle \psi_0 \Big| J_0\Big(z_{0,1}\frac{r}{R}\Big) \Big\rangle \\ & + \sum_{n=2}^{+\infty} \zeta_{0,n}^{\alpha,\beta}(\omega_a,\omega_b) \Big\langle \psi_0 \Big| J_0\Big(z_{0,n}\frac{r}{R}\Big) \Big\rangle = 0, \end{split}$$

which can then be rewritten as

$$\zeta_{0,1}^{\alpha,\beta}(\omega_a,\omega_b) = -\sum_{n=2}^{+\infty} \zeta_{0,n}^{\alpha,\beta}(\omega_a,\omega_b) \frac{\left\langle \psi_0 | J_0\left(z_{0,n}\frac{r}{R}\right)\right\rangle}{\left\langle \psi_0 | J_0\left(z_{0,1}\frac{r}{R}\right)\right\rangle}.$$
 (C5)

This means that $\zeta_{0,1}$ is no longer considered an independent variable and, as such, we will focus our attention on the $n \ge 2$ terms. For compactness, we define

$$\mathbf{f}_n = \frac{\left\langle \psi_0 | J_0\left(z_{0,n} \frac{r}{R}\right) \right\rangle}{\left\langle \psi_0 | J_0\left(z_{0,1} \frac{r}{R}\right) \right\rangle},\tag{C6}$$

which lets us rewrite Eq. (C5) as

$$\zeta_{0,1}^{\alpha,\beta}(\omega_a,\omega_b) = -\sum_{n=2}^{+\infty} \zeta_{0,n}^{\alpha,\beta}(\omega_a,\omega_b) \mathbf{f}_n.$$
 (C7)

Substituting Eq. (C7) into Eq. (C1) and then differentiating the resulting expression with respect to $[\zeta_{l,n}]^{\dagger}$, we obtain for

 $n \ge 2$,

$$\begin{split} & \left[g_{1}^{(0)}(\omega_{a}+\omega_{b})\mathbf{f}_{n}\mathbf{f}_{m}+\mathcal{V}_{1,1}^{(0)}\mathbf{f}_{n}\mathbf{f}_{m}\right]\zeta_{0,n}-\mathcal{V}_{1,m}^{(0)}\mathbf{f}_{n}\zeta_{0,n}-\mathcal{V}_{n,1}^{(0)}\mathbf{f}_{m}\zeta_{0,n} \\ & -\frac{1}{2}\mathbf{f}_{n}\sum_{m=1}^{+\infty}\mathcal{T}_{nm}^{l=0}\cdot(c_{m}^{+}+c_{m}^{-}) \\ & +\sum_{n=2}^{+\infty}g_{n}^{(0)}(\omega_{a}+\omega_{b})\zeta_{0,n}+\sum_{n,m=2}^{+\infty}\mathcal{V}_{nm}^{(0)}\zeta_{0,m} \\ & +\frac{1}{2}\sum_{n,m=2}^{+\infty}\left\{\mathcal{T}_{nm}^{l=0}\cdot(c_{m}^{+}+c_{m}^{-})+\text{c.c.}\right\}=0. \end{split}$$
(C8)

This means that the system can be easily solved algebraically as

$$= [\mathbb{F} + \mathbb{M}^{(0)}(\omega_a + \omega_b)]^{-1} \cdot \left[\mathbf{W}_0^{x,x}(\omega_a) + \mathbf{f}_0^{x,x}(\omega_a) \right],$$
(C9)

with $\mathbb{M}^{(0)}(\omega_a + \omega_b)$ defined as in Eq. (A7) and

$$\mathbf{W}_{0}^{x,x} = -\mathbb{T}^{0} \cdot \mathbf{c}, \quad \left(\mathbf{f}_{0}^{x,x}\right)_{n} = \mathbf{f}_{n} \sum_{m=1}^{N} c_{m} \mathcal{T}_{1,m}^{(0)} c_{m},$$

$$(\mathbb{F})_{i,j} = \left[g_{1}^{(0)}(\omega_{a} + \omega_{b}) + \mathcal{V}_{1,1}^{(0)}\right] \mathbf{f}_{i} \mathbf{f}_{j}$$

$$-\mathcal{V}_{1,j}^{(0)} \mathbf{f}_{i} - \mathcal{V}_{i,1}^{(0)} \mathbf{f}_{j}.$$
 (C10)

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The missing coefficient $\zeta_{0,1}^{\alpha,\beta}(\omega_a, \omega_b)$ is obtained by calculating Eq. (C7) with the solution of Eq. (C9). Several expressions have been already simplified by taking advantage of the parity of Bessel functions of the first kind,

$$J_{-l}(x) = (-1)^l J_l(x).$$

Looking now at the simpler l = 2 case, and as no orthogonality-based restrictions need to be applied to its coefficients, we can directly differentiate Eq. (C1) with respect to $[\zeta_{l,n}]^{\dagger}$ and then minimize the resulting expression, obtaining

$$\sum_{n=1}^{+\infty} g_n^{(2)}(\omega_a + \omega_b)\zeta_{2,n} + \sum_{n,m=1}^{+\infty} \mathcal{V}_{nm}^{(2)}\zeta_{2,m} + \frac{1}{2} \sum_{n,m=1}^{+\infty} \left\{ \mathcal{T}_{nm}^{l=2} \cdot c_m^+ + \mathcal{T}_{nm}^{l=-2} \cdot c_m^- \right\} = 0.$$
(C11)

The $\zeta_{2,n}$ variational coefficients can then be obtained, in vector form, as

$$\zeta_2^{x,x}(\omega_a,\omega_b) = \left[\mathbb{M}^{(2)}(\omega_a + \omega_b)\right]^{-1} \cdot \mathbf{W}_2^{x,x}(\omega_a), \qquad (C12)$$

a similar system to the one in Eq. (A10) with $\mathbf{W}_2^{x,x}$ defined analogously to $\mathbf{W}_0^{x,x}$, after simplifications from parity of Bessel functions.

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