SrRuO₃-SrTiO₃ heterostructure as a possible platform for studying unconventional superconductivity in Sr₂RuO₄

Bongjae Kim, 1,2,* Sergii Khmelevskyi, 3 Cesare Franchini, 4,5 I. I. Mazin, 6,7 and Kyoo Kim, 1 Department of Physics, Kunsan National University, Gunsan 54150, Korea

2 MPPHC-CPM, Max Planck POSTECH/Korea Research Initiative, Pohang 37673, Korea

3 Center for Computational Materials Science, Institute for Applied Physics, Vienna University of Technology,
Wiedner Hauptstrasse 8-10, A-1040 Vienna, Austria

4 Faculty of Physics and Center for Computational Materials Science, University of Vienna, A-1090 Vienna, Austria

5 Dipartimento di Fisica e Astronomia, Università di Bologna, I-40127 Bologna, Italy

6 Code 6393, Naval Research Laboratory, Washington, DC 20375, USA

7 Quantum Materials Center, George Mason University, Fairfax, Virginia 22030, USA

8 Department of Physics, Pohang University of Science and Technology, Pohang 37673, Korea

9 Korea Atomic Energy Research Institute (KAERI), 111 Daedeok-daero, Daejeon 34057, Korea

(Received 27 January 2020; accepted 22 May 2020; published 8 June 2020)

There is intense controversy around the unconventional superconductivity in strontium ruthenate, where the various theoretical and experimental studies suggest diverse and mutually exclusive pairing symmetries. Currently, the investigation is solely focused on only one material, Sr_2RuO_4 , and the field suffers from the lack of comparison targets. Here, employing a density-functional-theory-based analysis, we show that the heterostructure composed of $SrRuO_3$ and $SrTiO_3$ is endowed with all the key characteristics of Sr_2RuO_4 , and, in principle, can host superconductivity. Furthermore, we show that competing magnetic phases and associated frustration, naturally affecting the superconducting state, can be tuned by epitaxial strain engineering. This system thus offers an excellent platform for gaining more insight into superconductivity in ruthenates.

DOI: 10.1103/PhysRevB.101.220502

Introduction. Transition-metal oxides have long been a fertile ground for novel physics. In particular, unconventional superconductivity, found, for instance, in cuprates and ruthenates, is one of the most intriguing discovery of recent decades, and the physics and mechanisms in these oxides is still highly debated. The latter compound is particularly intriguing; despite tens of years of intensive studies, even a basic understanding of pairing symmetry is still lacking [1], and new surprises are continuously coming up. Arguably, one of the reasons the latter compound has been so difficult to crack is that it is a one-of-a-kind material. As opposed to the Cu- and Fe-based high- T_c superconductors, which have multiple families with many members each, Sr_2RuO_4 is a family of one, and no insightful comparison with any sibling can be effected.

In fact, one of the biggest breakthroughs in the field, which led to dramatic progress, occurred when experimentalists found a way to apply uniaxial strain so as to induce a Lifshitz transition [2,3]. Finding another material, affording even more flexibility in modifying Sr_2RuO_4 properties, including its Fermiology and magnetic response, would open up a cornucopia of new experimental information and may lead to extraordinary progress in understanding the superconductivity in Sr_2RuO_4 .

This idea is not new. Burganov *et al.* demonstrated that it is possible to grow epitaxial Sr_2RuO_4 thin films on a $SrTiO_3$ substrate, with a lattice mismatch of 0.9% [4]. No superconductivity was observed, which may be related to either too many defects at the interface (and the authors were optimistic about reducing their amount in the future), or due to a $\sqrt{2} \times \sqrt{2}$ reconstruction, also present at the free-standing Sr_2RuO_4 surface. In particular, they observed in angle-resolved photoemission spectroscopy a Lifshitz transition not unlike the one seen under uniaxial strain, and speculated that approaching the van Hove singularity brings about a strong enhancement of the uniform magnetic susceptibility via the Stoner mechanism and, therefore, strengthening of ferromagnetic (FM) spin fluctuations.

Very recently, such an enhancement has been observed in the resistivity [5], as well as in nuclear magnetic resonance (NMR) experiments and density functional theory (DFT) calculations [6] for bulk Sr_2RuO_4 under uniaxial stress. Note that, generally speaking, FM fluctuations favor triplet, and antiferromagnetic fluctuations favor singlet pairing. In the unstrained Sr_2RuO_4 , the spin fluctuation spectrum is well documented [7], and it has been demonstrated that, barring other triplet-favoring interactions, the antiferromagnetic fluctuations win hands down. Yet, if pairing is indeed triplet, superconductivity which cannot benefit directly from the increased density of states (DOS) at the van Hove singularity can be boosted by these FM fluctuations, in accordance with the observed T_c behavior [3].

^{*}bongjae.kim@kunsan.ac.kr

[†]kyoo@kaeri.re.kr

Not unexpectedly, the work by Burganov *et al.* [4] has triggered several works based on theoretical modeling employing tight-binding models [8,9]. Unfortunately, such models do not take into account all the complexity of the epitaxial interface, band effects, and structural delicacy, and an extension to other systems has not been made so far. Here, we exploit the idea of heterostructures, employing first-principles calculations, which, differently from model-based approaches, have a capacity to capture much of the aforementioned complexity.

Curiously, while the cuprates, which are already a rather rich family, have been subjected to multiple suggestions and attempts to generate similar materials, including engineered heterostructures [10-18], Sr₂RuO₄, which needs expansion onto other systems much more badly, has hardly been discussed in this context. Yet, from the perspective of the growth technique, epitaxy and heterostructuring are well understood for ruthenates, especially for SrRuO₃ [19]. Particularly promising would be a superlattice composed of a single layer of SrRuO₃ sandwiched between insulating layers, which is now experimentally accessible with fine control [20–22]. It has the same principal structural motif as Sr₂RuO₄, a twodimensional square RuO₂ lattice. As we show below, it shares with Sr₂RuO₄ key features of the electronic structure and magnetic properties, but also has interesting and promising distinctions.

In the following, we report a first-principles computational study of the electronic structure and magnetic properties of the $SrRuO_3$ - $SrTiO_3$ (SRO-STO) heterostructure, grown on a STO substrate (we will also briefly discuss other potential substrate ranges). The role of the substrate is to fix the lateral lattice dimensions, and, thus, to provide biaxial strain. We show that the SRO-STO superlattice has all the key characteristics of Sr_2RuO_4 , and would broaden the exploratory potential dramatically in regard to the superconductivity mechanism and pairing symmetry.

Result and discussion. The most important structural feature of Sr₂RuO₄ is the two-dimensional (2D) RuO₂ layer, spaced between SrO layers, where the corner-shared RuO₆ octahedra form a square net. This active layer can be artificially constructed by sandwiching a SRO monolayer between STO blocks as shown in Fig. 1(a), where the out-of-plane Ru-O-Ru connectivity is broken due to the interspaced STO layers. A small in-plane octahedral rotation (~8.2° for an unstrained case) exists for this superlattice, which produces a $\sqrt{2} \times \sqrt{2}$ supercell reconstruction in the plane [20]. This reconstruction is not found for the pure bulk phase of Sr₂RuO₄, but is observed at the surface, as well as in a Ca-doped system [23–25], reflecting the tendency for RuO₆ octahedra to rotate, if necessary to accommodate the geometrical constraints. Note that even with this octahedral rotation pattern, surface superconductivity has been observed in tunneling [26].

The structural similarity of the two systems is reflected in their electronic structures. In Fig. 1(b), we plotted the orbitally resolved partial density of states (DOS) of Ru-d for both systems. Due to an octahedral crystal field splitting, the three t_{2g} orbitals (xy, yz, and zx) share four electrons, and are responsible for the low-energy physics. The xy orbital forms a quasi-2D band, and a nearly circular Fermi surface (FS), usually labeled as γ , and the other orbitals form two quasi-1D bands, which, upon intersection and rehybridization, create

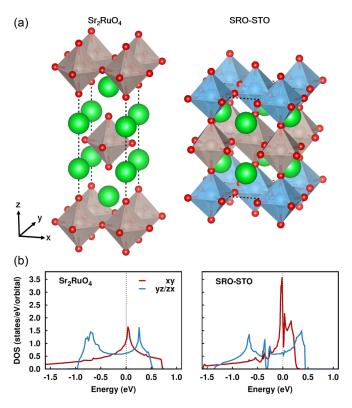


FIG. 1. (a) Crystal structure of Sr_2RuO_4 and SRO-STO. Gray and blue octahedera denote RuO_6 and TiO_6 octahedra, and red and green spheres represent oxygen and strontium atoms, respectively (b) Partial DOS of Ru-d orbitals for Sr_2RuO_4 and SRO-STO for xy and yz/zx.

two rounded-square-shaped Fermi surfaces (α and β). From the partial DOS of SRO-STO, characteristic features observed in Sr₂RuO₄ are present as well. The peak of the xy partial DOS, corresponding to a 2D van Hove singularity (vHs), is located close to the Fermi level, and its contribution at the Fermi energy is much larger than those of the yz/zx bands. Less dispersive yz/zx orbitals have double-peak structures, corresponding to 1D vHs, and have smaller contributions at the Fermi level.

In Fig. 1(b), one can notice that the xy DOS peak in the heterostructure is higher, and the bandwidth is smaller, than in Sr₂RuO₄. This can be attributed to the octahedral rotations, which reduce the effective xy-xy hopping as $\cos^2 \alpha$, where α is the rotational angle, while the xz-xz and yz-yz hopping is unaffected in the same order in α . This fact is immediately relevant for one of the main points of contention regarding superconductivity in Sr₂RuO₄: While one school considers the xy band to have strong pairing, with superconductivity in xz/yz being induced by an interband proximity effect [27,28], another advocates the xz/yz bands as the "active" subsystem, with xy being secondary [29]. Yet others suggested that all bands contribute roughly equally [30,31]. Enhancing selectively the DOS in the xy band only, through octahedral rotations (which, as we show later, can be controlled by the substrate), provides a direct test of these alternatives.

Another issue of relevance is the proximity to magnetism. It is well known that Sr_2RuO_4 is on the verge of an antiferromagnetic instability, driven by the quasi-1D nesting in

the xz/yz bands. The corresponding spin fluctuations favor d-wave pairing [7,31]. However, there is also, albeit a much weaker, tendency to a FM instability, which favors p-wave pairing. A dramatic increase in DOS brings the system much closer to ferromagnetism, thus strongly pushing the system toward a triplet pairing. Note that within the weak-coupling limit, a recent T_c enhancement in the uniaxial strain experiment has been associated with an odd- to even-parity transition, and our study can offer one way to verify this proposition [3].

With this in mind, we have looked at the magnetism of this system in more detail. Note that there have been DFT-based studies on SRO-STO systems both in film and superlattices structures [22,32-34], and they concluded the magnetic and electronic ground states of a single SRO layer to be either a FM metal or Néel-type antiferromagnetic insulator. This issue is not also resolved as judged from the diverse experimental reports [20-22,35]. The problem with these studies is that they take the DFT ground state at its face value, forgetting that it is, by nature, a mean-field approach, and, as such, liable to overestimate the tendency to magnetism. Indeed, straight DFT calculations predict Sr₂RuO₄ to be spin density wave (SDW) type antiferromagnetic (not a Néel type, see below), and, with gradient corrections, even the FM state has a lower energy than the nonmagnetic one [36]. This tendency can be corrected for, phenomenologically, by introducing the concept of a fluctuation-renormalized Stoner factor [37,38], or strongly remedied by switching to dynamical mean-field theory [39]. Applying the same concept to Sr₂RuO₄ and SRO-STO, we observed that the experimentally observed Stoner enhancement for the former is 7 [7], thus the Stoner product $IN(E_F) \approx 0.857$ [at the same time, renormalization at $\mathbf{q} = \mathbf{Q}_{\text{SDW}}$ is about 30, implying that $I(Q_{\text{SDW}})\chi_0(Q_{\text{SDW}}) \approx$ 0.967, where χ_0 is the bare susceptibility]. Our calculations find that the partial density of states of Ru orbitals on the Fermi level is about 50% higher in SRO-STO. This suggests that their $IN(E_F)$ is larger than one and the ground state should be FM. Fixed spin moment calculations confirm that in order to stabilize the paramagnetic state one would need to reduce the Stoner I by 3.5 times, which is completely unphysical.

As discussed below, the DOS at the Fermi level strongly depends on the epitaxial strain, which offers a unique opportunity to tune the system all the way from low DOS and low T_c to higher DOS and high T_c , as in the case of uniaxial strain [2,3], and further to even higher DOS, i.e., the FM state. Furthermore, for a range of strains the FM instability enters a fierce competition with the SDW one, which is even more enhanced for the SRO-STO system [see Figs. 2(a)-2(c)]. Bulk Sr₂RuO₄ is, both experimentally and theoretically, much closer to a SDW instability than to a FM one. Strong magnetic fluctuations have been observed by the neutron diffraction, which persists up to room temperature [40,41]. This had been predicted in DFT studies [31] and a SDW with $\mathbf{q}_3 = (1, 1, 0) \frac{2\pi}{3a}$, a close commensurate approximation to the experimental peak in spin susceptibility, was shown to be stable in calculations [36,42]. On the contrary, for the bulk material the FM state is barely stable in the generalized gradient approximation (GGA), and unstable in the local density approximation (LDA). There is no magnetic frustration in the

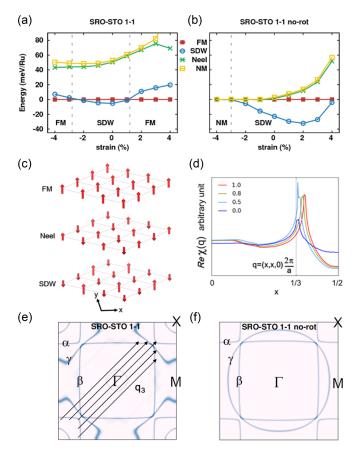


FIG. 2. The energy landscape as a function of strain (a) with and (b) without octahedral rotation of RuO_6 , for three magnetic configurations schematically depicted in (c): FM (ferromagnetic), Néel antiferromagnetic, and SDW [$\mathbf{q}=(1,1,0)\frac{2\pi}{3a}$] phases. (d) The change in Lindhard susceptibility obtained from the unfolded band for Ru yz/zx orbitals. Octahedral rotations are linearly interpolated between the unrotated case (0.0) and the fully rotated one (1.0). (e), (f) The change in Fermi surfaces (e) with and (f) without octahedral rotation. (d)–(f) are for the unstrained (0%) structure. The arrows in the Fermi surfaces indicate \mathbf{q}_3 nesting vector which activates both the xz and yz quasi-1D Fermi surfaces.

system, with the SDW being a clear mean-field ground state, suppressed by quantum fluctuations.

Interestingly, in SRO-STO, while both FM and SDW states gain stability, eventually they become nearly degenerate [Figs. 2(a) and 2(b)], adding a strong parametric frustration to the system. In a number of materials, most notably in Fe-based superconductors, such a strong magnetic frustration leads to the complete destruction of a long-range magnetic order and numerous interesting phenomena, such as nematic order, and, arguably, superconductivity itself [43,44]. It looks plausible that magnetic frustration in SRO-STO would prevent the system from developing a static magnetic order, but would trigger strong spin fluctuations, with all ramifications for the superconductivity. While the jury is still out on whether the spin-fluctuation mechanism is instrumental in cuprates, Fe pnictides, or Sr₂RuO₄, it is considered by many to be a strong contender in all three cases (which, of course, does not imply that the nature of superconductivity is the same in all three instances). Enhanced frustration of magnetism in SRO-STO, in this context, looks especially interesting.

Previous computational studies of the SRO-STO superlattice concentrated upon the competition between FM and Néel orders [32,34], ignoring the SDW $\mathbf{q}_3 = (1, 1, 0) \frac{2\pi}{3a}$ phase. Our calculations show that, in spite of a dramatic stabilization of the FM phase, the SDW still has a lower energy [Fig. 2(a)] over a large range of strains (the Néel order is not competitive at all). The enhanced competition between the two magnetic phases in SRO-STO is promoted by the octahedral rotations, which are absent in Sr₂RuO₄. As shown in Fig. 2(b), the calculations with suppressed octahedral rotations favor the SDW tendency over FM. For strains between $\sim -2\%$ and $\sim 1\%$, the SDW has the lowest energy, but the energy gain is minimal, ~5 meV/Ru for the zero strain case (the corresponding value for Sr_2RuO_4 is ~ 20 meV/Ru). For larger strains ($\lesssim -2\%$ and $\geq 1\%$), the FM state becomes the most stable one. Note that the critical strain for the transition from SDW to FM is within the accessibility of current experimental techniques [45], and even broader epitaxial strain ranges could be reached in recent experiments [46]. This suggests that the SRO-STO system is a suitable platform for the control of different types of magnetic interactions, and, by implication, for tuning the superconducting order parameter. Note that, given that the fluctuation-induced suppression of magnetism may be, and probably is, q dependent, it is hard to say which state has the lowest free energy for which strain.

As can be seen in Fig. 2(d), the real part of the charge susceptibility in the SRO-STO system (see Supplemental Material [47]) shows only a gradual shift of the \mathbf{q}_3 peak upon octahedral rotation. This is consistent with the fact that the α/β FSs remain quasi-one-dimensional, with the same \mathbf{q}_3 nesting vector, as indicated by the arrows in Fig. 2(e). This lets us conclude that the antiferromagnetic SDW fluctuations are highly robust against octahedral rotation, and, consequently, against substrate engineering. However, the two-dimensional ν FS is highly sensitive to octahedral rotations, which can be efficiently tuned by epitaxial strain. In particular, γ FS changes its shape with strain from a circle to a rounded square, which is not related to the van Hove singularity often discussed in strain studies of Sr₂RuO₄ [2,3,6], but is due to mixing between the xy and $x^2 - y^2$ orbitals of neighboring Ru atoms as shown in Figs. 2(e) and 2(f). The hybridization between xy and $x^2 - y^2$ bands due to octahedral rotation is eventually responsible for the additional FSs at around the Γ and X point for SRO-STO [Fig. 2(e)] [47].

Changes in DOS and FSs upon strain are shown in Fig. 3. The width of yz/zx DOS becomes smaller as the tensile strain increases, while the xy DOS is nearly unaffected except for a shift of the peak position toward lower energies. Notably, the peak, mainly from the xy orbital, which is thought to be closely related with the T_c enhancement of the uniaxial strain case, progressively moves down away from the Fermi level for tensile strains while the yz/zx DOS gradually increases at the Fermi level [47,48]. For Sr_2RuO_4 , octahedral distortion shifts the peak down [49], while in SRO-STO, a tensile strain, which generally relieves the distortion, has a similar effect.

This selective evolution of the xy and yz/zx orbital upon strain suggests that future experimental studies of the SRO-STO system can probe the "active" orbitals, important for superconductivity [27–29]. Furthermore, the heterostructural features of the SRO-STO system itself can restrict specific

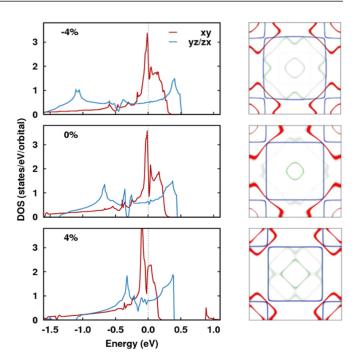


FIG. 3. Partial DOS of Ru xy (red) and Ru yz/zx (blue) bands for strains -4%, 0%, and 4%. Corresponding Fermi surfaces including the Γ point are shown in the right column with the same color code. The contribution of other orbitals at the Fermi surface is marked by the green color.

types of scattering, and, within this structural setup, some superconducting symmetries, such as those with horizontal node lines, cannot survive. The DOS at the Fermi level, which is tunable in our setup, is closely related with T_c behavior for a singlet pairing, but not for a triplet one. Thus, the possibility of a specific pairing such as $d_{x^2-y^2}$, which is promoted by the enhancement of the DOS at the Fermi level, can be directly tested in the SRO-STO system. Being close relatives of the orphan superconductor Sr_2RuO_4 , SRO-STO superlattices offer a unique possibility to controllably tune the Fermi-surface topology, magnetic interactions, and the position of the van Hove peak, thus providing different avenues to probe the paring symmetry of the ruthenates.

Conclusion. SRO-STO heterostructures present an electronic system in many aspects similar to the intriguing putative triplet superconductor Sr₂RuO₄, but also distinctly different. The important common motif is the presence of the three Fermi surfaces, representing three Ru orbitals, xy, yz, and zx, with the first one exhibiting a van Hove singularity at or near the Fermi level. Moreover, the exact topology of the Fermi surfaces can be modified within a reasonable range by varying the substrate and thus exerting different epitaxial strains. An important difference from Sr₂RuO₄ is that SRO-STO appears, in general, more magnetic, and at the same time the tendency to ferromagnetism is enhanced much more strongly than to the SDW. In principle, three scenarios can be realized, all three interesting in their own way. First, the enhanced tendency to SDW could lead to an actual instability. Second, at least for some strains, the material may become FM. The third, and arguably the most interesting, alternative is that the near degeneracy (on a mean-field level) of the FM and the SDW state will lead to a very strong parametric magnetic frustration that will preclude either static order, but will greatly enhance both types of fluctuations. In this case, the system, if manufactured cleanly enough, is likely to become superconducting, possibly in the triplet channel (boosted by FM fluctuations), but the properties of this superconducting state may be even more different from the bulk Sr₂RuO₄ than the critically strained samples [3]. If the symmetry of the superconducting state in Sr₂RuO₄ is indeed triplet (which is, however, strongly questioned by recent NMR experiments [50,51]), the critical temperature of SRO-STO may be dramatically higher, because of stronger fluctuations, than that of Sr₂RuO₄.

Acknowledgments. We thank Beom Hyun Kim and Minjae Kim for fruitful discussions. This work was supported by the NRF Grant (Contracts No. 2018R1D1A1A02086051 and No. 2016R1D1A1B02008461), the research funds of Kunsan National University, and Max-Plank POSTECH/KOREA Research Initiative (Grant No. 2016K1A4A4A01922028). The computing resources from the KISTI supercomputing center (Project No. KSC-2018-CRE-0079) are greatly acknowledged. I.I.M. acknowledges support by ONR through the NRL basic research program. K.K acknowledges support from Internal R&D program at KAERI funded by the Ministry of Science and ICT (MSIT) of the Republic of Korea (Grant No. 524210-20).

- [1] A. P. Mackenzie, T. Scaffidi, C. W. Hicks, and Y. Maeno, npj Quantum Mater. 2, 40 (2017).
- [2] C. W. Hicks, D. O. Brodsky, E. A. Yelland, A. S. Gibbs, J. A. N. Bruin, M. E. Barber, S. D. Edkins, K. Nishimura, S. Yonezawa, Y. Maeno, and A. P. Mackenzie, Science 344, 283 (2014).
- [3] A. Steppke, L. Zhao, M. E. Barber, T. Scaffidi, F. Jerzembeck, H. Rosner, A. S. Gibbs, Y. Maeno, S. H. Simon, A. P. Mackenzie, and C. W. Hicks, Science 355, eaaf9398 (2017).
- [4] B. Burganov, C. Adamo, A. Mulder, M. Uchida, P. D. C. King, J. W. Harter, D. E. Shai, A. S. Gibbs, A. P. Mackenzie, R. Uecker, M. Bruetzam, M. R. Beasley, C. J. Fennie, D. G. Schlom, and K. M. Shen, Phys. Rev. Lett. 116, 197003 (2016).
- [5] M. E. Barber, A. S. Gibbs, Y. Maeno, A. P. Mackenzie, and C. W. Hicks, Phys. Rev. Lett. 120, 076602 (2018).
- [6] Y. Luo, A. Pustogow, P. Guzman, A. P. Dioguardi, S. M. Thomas, F. Ronning, N. Kikugawa, D. A. Sokolov, F. Jerzembeck, A. P. Mackenzie, C. W. Hicks, E. D. Bauer, I. I. Mazin, and S. E. Brown, Phys. Rev. X 9, 021044 (2019).
- [7] P. Steffens, Y. Sidis, J. Kulda, Z. Q. Mao, Y. Maeno, I. I. Mazin, and M. Braden, Phys. Rev. Lett. 122, 047004 (2019).
- [8] Y.-T. Hsu, W. Cho, A. F. Rebola, B. Burganov, C. Adamo, K. M. Shen, D. G. Schlom, C. J. Fennie, and E.-A. Kim, Phys. Rev. B 94, 045118 (2016).
- [9] Y.-C. Liu, W.-S. Wang, F.-C. Zhang, and Q.-H. Wang, Phys. Rev. B 97, 224522 (2018).
- [10] R. Arita, A. Yamasaki, K. Held, J. Matsuno, and K. Kuroki, Phys. Rev. B 75, 174521 (2007).
- [11] V. I. Anisimov, D. Bukhvalov, and T. M. Rice, Phys. Rev. B 59, 7901 (1999).
- [12] J. Chaloupka and G. Khaliullin, Phys. Rev. Lett. 100, 016404 (2008).
- [13] P. Hansmann, X. Yang, A. Toschi, G. Khaliullin, O. K. Andersen, and K. Held, Phys. Rev. Lett. 103, 016401 (2009).
- [14] H. Watanabe, T. Shirakawa, and S. Yunoki, Phys. Rev. Lett. 110, 027002 (2013).
- [15] Y. K. Kim, O. Krupin, J. D. Denlinger, A. Bostwick, E. Rotenberg, Q. Zhao, J. F. Mitchell, J. W. Allen, and B. J. Kim, Science 345, 187 (2014).
- [16] J. Gawraczyński, D. Kurzydłowski, R. A. Ewings, S. Bandaru, W. Gadomski, Z. Mazej, G. Ruani, I. Bergenti, T. Jaroń, A. Ozarowski, S. Hill, P. J. Leszczyński, K. Tokár, M. Derzsi, P. Barone, K. Wohlfeld, J. Lorenzana, and W. Grochala, Proc. Natl. Acad. Sci. USA 116, 1495 (2019).

- [17] E. B. Isaacs and C. Wolverton, Phys. Rev. X 9, 021042 (2019).
- [18] D. Li, K. Lee, B. Y. Wang, M. Osada, S. Crossley, H. R. Lee, Y. Cui, Y. Hikita, and H. Y. Hwang, Nature (London) 572, 624 (2019).
- [19] G. Koster, L. Klein, W. Siemons, G. Rijnders, J. S. Dodge, C.-B. Eom, D. H. A. Blank, and M. R. Beasley, Rev. Mod. Phys. 84, 253 (2012).
- [20] M. Gu, Q. Xie, X. Shen, R. Xie, J. Wang, G. Tang, D. Wu, G. P. Zhang, and X. S. Wu, Phys. Rev. Lett. 109, 157003 (2012).
- [21] H. Boschker, T. Harada, T. Asaba, R. Ashoori, A. V. Boris, H. Hilgenkamp, C. R. Hughes, M. E. Holtz, L. Li, D. A. Muller, H. Nair, P. Reith, X. Renshaw Wang, D. G. Schlom, A. Soukiassian, and J. Mannhart, Phys. Rev. X 9, 011027 (2019).
- [22] S. G. Jeong, T. Min, S. Woo, J. Kim, Y.-Q. Zhang, S. W. Cho, J. Son, Y.-M. Kim, J. H. Han, S. Park, H. Y. Jeong, H. Ohta, S. Lee, T. W. Noh, J. Lee, and W. S. Choi, Phys. Rev. Lett. 124, 026401 (2020).
- [23] R. Matzdorf, Z. Fang, Ismail, J. Zhang, T. Kimura, K. T. Y. Tokura and, and E. W. Plummer, Science 289, 746 (2000).
- [24] A. Damascelli, D. H. Lu, K. M. Shen, N. P. Armitage, F. Ronning, D. L. Feng, C. Kim, Z.-X. Shen, T. Kimura, Y. Tokura, Z. Q. Mao, and Y. Maeno, Phys. Rev. Lett. 85, 5194 (2000).
- [25] S.-C. Wang, H.-B. Yang, A. K. P. Sekharan, S. Souma, H. Matsui, T. Sato, T. Takahashi, C. Lu, J. Zhang, R. Jin, D. Mandrus, E. W. Plummer, Z. Wang, and H. Ding, Phys. Rev. Lett. 93, 177007 (2004).
- [26] I. A. Firmo, S. Lederer, C. Lupien, A. P. Mackenzie, J. C. Davis, and S. A. Kivelson, Phys. Rev. B 88, 134521 (2013).
- [27] D. F. Agterberg, T. M. Rice, and M. Sigrist, Phys. Rev. Lett. 78, 3374 (1997).
- [28] Y. Yanase, T. Jujo, T. Nomura, H. Ikeda, T. Hotta, and K. Yamada, Phys. Rep. **387**, 1 (2003).
- [29] S. Raghu, A. Kapitulnik, and S. A. Kivelson, Phys. Rev. Lett. 105, 136401 (2010).
- [30] I. I. Mazin and D. J. Singh, Phys. Rev. Lett. 79, 733 (1997).
- [31] I. I. Mazin and D. J. Singh, Phys. Rev. Lett. 82, 4324 (1999).
- [32] P. Mahadevan, F. Aryasetiawan, A. Janotti, and T. Sasaki, Phys. Rev. B 80, 035106 (2009).
- [33] M. Verissimo-Alves, P. García-Fernández, D. I. Bilc, P. Ghosez, and J. Junquera, Phys. Rev. Lett. **108**, 107003 (2012).
- [34] L. Si, Z. Zhong, J. M. Tomczak, and K. Held, Phys. Rev. B 92, 041108(R) (2015).

- [35] Y. J. Chang, C. H. Kim, S.-H. Phark, Y. S. Kim, J. Yu, and T. W. Noh, Phys. Rev. Lett. 103, 057201 (2009).
- [36] B. Kim, S. Khmelevskyi, I. I. Mazin, D. F. Agterberg, and C. Franchini, npj Quantum Mater. 2, 37 (2017).
- [37] P. Larson, I. I. Mazin, and D. J. Singh, Phys. Rev. B 69, 064429 (2004).
- [38] L. Ortenzi, I. I. Mazin, P. Blaha, and L. Boeri, Phys. Rev. B 86, 064437 (2012).
- [39] F. B. Kugler, M. Zingl, H. U. R. Strand, S.-S. B. Lee, J. von Delft, and A. Georges, Phys. Rev. Lett. **124**, 016401 (2020).
- [40] Y. Sidis, M. Braden, P. Bourges, B. Hennion, S. NishiZaki, Y. Maeno, and Y. Mori, Phys. Rev. Lett. 83, 3320 (1999).
- [41] K. Iida, M. Kofu, N. Katayama, J. Lee, R. Kajimoto, Y. Inamura, M. Nakamura, M. Arai, Y. Yoshida, M. Fujita, K. Yamada, and S.-H. Lee, Phys. Rev. B 84, 060402(R) (2011).
- [42] S. Cobo, F. Ahn, I. Eremin, and A. Akbari, Phys. Rev. B 94, 224507 (2016).
- [43] I. I. Mazin, M. D. Johannes, L. Boeri, K. Koepernik, and D. J. Singh, Phys. Rev. B 78, 085104 (2008).
- [44] J. K. Glasbrenner, I. I. Mazin, H. O. Jeschke, P. J. Hirschfeld, R. M. Fernandes, and R. Valení, Nat. Phys. 11, 953 (2015).
- [45] A. Vailionis, H. Boschker, W. Siemons, E. P. Houwman, D. H. A. Blank, G. Rijnders, and G. Koster, Phys. Rev. B 83, 064101 (2011).
- [46] D. G. Schlom, L.-Q. Chen, C.-B. Eom, K. M. Rabe, S. K. Streiffer, and J.-M. Triscone, Annu. Rev. Mater. Res. 37, 589 (2007).
- [47] See Supplemental Material at http://link.aps.org/supplemental/ 10.1103/PhysRevB.101.220502 for (i) calculation details, (ii) susceptibility calculation, (iii) strain-dependent changes of structure and DOS, and (iv) the role of spin-orbit coupling on the Fermi surfaces, which includes Refs. [52–60].

- [48] For the detailed structure of the DOS peak at around Fermi level, more accurate treatments are needed. For example, see Luo *et al.* [6]. In the current study, we focus on the general trends upon the substrate strain.
- [49] E. Ko, B. J. Kim, C. Kim, and H. J. Choi, Phys. Rev. Lett. 98, 226401 (2007).
- [50] A. Pustogow, Y. Luo, A. Chronister, Y.-S. Su, D. Sokolov, F. Jerzembeck, A. P. Mackenzie, C. W. Hicks, N. Kikugawa, S. Raghu, E. D. Bauer, and S. E. Brown, Nature 574, 72 (2019).
- [51] K. Ishida, M. Manago, and Y. Maeno, J. Phys. Soc. Jpn. 89, 034712 (2020).
- [52] G. Kresse and J. Hafner, Phys. Rev. B 47, 558 (1993).
- [53] G. Kresse and J. Furthmüller, Phys. Rev. B **54**, 11169 (1996).
- [54] J. P. Perdew, K. Burke, and M. Ernzerhof, Phys. Rev. Lett. 77, 3865 (1996).
- [55] K. Koepernik and H. Eschrig, Phys. Rev. B **59**, 1743 (1999).
- [56] W. Ku, T. Berlijn, and C.-C. Lee, Phys. Rev. Lett. 104, 216401 (2010).
- [57] M. E. Barber, F. Lechermann, S. V. Streltsov, S. L. Skornyakov, S. Ghosh, B. J. Ramshaw, N. Kikugawa, D. A. Sokolov, A. P. Mackenzie, C. W. Hicks, and I. I. Mazin, Phys. Rev. B 100, 245139 (2019).
- [58] H. G. Suh, H. Menke, P. Brydon, C. Timm, A. Ramires, and D. F. Agterberg, arXiv:1912.09525.
- [59] K. M. Shen, A. Damascelli, D. H. Lu, N. P. Armitage, F. Ronning, D. L. Feng, C. Kim, Z.-X. Shen, D. J. Singh, I. I. Mazin, S. Nakatsuji, Z. Q. Mao, Y. Maeno, T. Kimura, and Y. Tokura, Phys. Rev. B 64, 180502(R) (2001).
- [60] C. N. Veenstra, Z.-H. Zhu, B. Ludbrook, M. Capsoni, G. Levy, A. Nicolaou, J. A. Rosen, R. Comin, S. Kittaka, Y. Maeno, I. S. Elfimov, and A. Damascelli, Phys. Rev. Lett. 110, 097004 (2013).