BaHgSn: A Dirac semimetal with surface hourglass fermions

Tan Zhang ^(b), ^{1,2} Zhihai Cui, ^{1,2} Zhijun Wang ^(b), ^{1,2} Hongming Weng, ^{1,2,3,4,*} and Zhong Fang ^{1,2}

¹Beijing National Research Center for Condensed Matter Physics, and Institute of Physics,

Chinese Academy of Sciences, Beijing 100190, China

²School of Physical Sciences, University of Chinese Academy of Sciences, Beijing 100049, China

³Songshan Lake Materials Laboratory, Dongguan, Guangdong 523808, China

⁴CAS Center for Excellence in Topological Quantum Computation, University of Chinese Academy of Sciences, Beijing 100190, China

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We proposed that BaHgSn is a Dirac semimetal (DSM) which can host hourglasslike surface states (HSSs) as protected by nonsymmorphic glide symmetry. Compared to KHgSb, an isostructural topological crystalline insulator with the same HSSs, BaHgSn has an additional band inversion at the Γ point. This band inversion is induced by the stronger interlayer coupling among Hg-Sn honeycomb layers than that among Hg-Sb layers in KHgSb, which leads to bulk Dirac nodes in BaHgSn along the layer stacking direction Γ -A. In addition, the mirror Chern number C_i protected by the mirror plane \overline{M}_z ($k_z = 0$) changes from 2 in KHgSb to 3 in BaHgSn. Therefore, when a compressive uniaxial strain is applied along the *y* axis to break the rotation symmetry protecting the DSM state, BaHgSn becomes a strong topological insulator with Z_2 indices of (1;000) and the topological surface Dirac cone coexists with HSSs on the (010) surface. The Wilson-loop spectra have been calculated to verify these topological features. The calculated surface states, the Fermi surfaces and their quasiparticle interference patterns are ready to be compared with experimental measurements.

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I. INTRODUCTION

Topological quantum states and the topological materials hosting them have been extensively studied since the late 2000s [1–8]. One of the most important and exciting finding is the realization of several relativistic quasiparticles as emergent effects at the boundaries where topological phase transition happens. For example, the two-dimensional (2D) massless Dirac fermions have been realized on the surface of 3D topological insulators (TIs) [1,2] and the Weyl fermions have been realized in 3D Weyl semimetals (WSMs), which can be viewed as the boundary state of a four-dimensional quantum Hall insulators [6,8]. Until now, there have been various quasiparticles theoretically proposed with and without their counterparts in field theory. Several of them have even been experimentally observed in some topological materials, such as the Dirac fermions in Dirac semimetals (DSMs) [9,10], Majorana fermions in topological superconductors [2,11], Weyl fermions in WSMs [12–17], hourglass fermions in topological crystalline insulators [18,19], and spin-1 fermions in triply degenerate nodal-point semimetals [7,20–23].

The hourglass fermions were proposed and observed on the surface of a typical topological crystalline insulator (TCI) KHgSb in nonsymmorphic space group $P6_3/mmc$ (D_{6h}^4) [18,19]. It has symmetry-based indicator $\mathbb{Z}_{12} = 8$ [24–26] and nontrivial topological invariants, namely hourglass invariant $\delta_h = 1$ in glide plane \overline{M}_x and mirror Chern number (mCN) $C_i = 2$ in the \overline{M}_z ($k_z = 0$) plane. These topological invariants determine that it has hourglasslike surface states (HSSs) and Dirac cone like ones protected by these symmetries on a surface preserving them [27]. There have been many compounds of this crystal structure proposed to be topological materials. For example, XYBi (X = Ba, Eu; Y = Cu, Ag, and Au) family [28,29] and XAuTe (X = K, Na, Rb) family [29,30] were proposed to realize DSMs or WSMs after proper tuning. It is noted that NaAuTe was predicted to be a DSM [29] since it has band crossing induced by the stronger bondingantibonding splitting than KHgSb, which also leads to mCN $C_i = 3$ for the mirror plane with $k_z = 0$. The similar way to achieving this band crossing has been proposed by applying compressive pressure along the c axis in KHgSb family compounds [31]. For example, KHgBi becomes a DSM under 11.5% compression of c lattice constant although such large compression is not plausible. Here we propose that BaHgSn is another DSM in the same crystal structure with the same physical mechanism. Therefore, it can host Dirac fermions in bulk and hourglass fermions on the (010) surface. When the rotation symmetry C_{3z} protecting Dirac nodes is broken, it becomes a strong topological insulator [9]. Combining theoretical analysis, first-principles calculations and Wilson loop calculations, we can well understand the resultant nontrivial surface states, including the HSSs and the zigzag like ones protected by mCN $C_i = 3$ proposed in Ref. [18] but not been found in real materials. These novel surface states can be identified by the calculations of surface states, the Fermi surfaces, and their quasiparticle interference (QPI) patterns. These phenomena can be readily measured by angle-resolved photoemission spectroscopy (ARPES) [1,32,33] and scanning tunneling microscope (STM) [34-37].

^{*}hmweng@iphy.ac.cn



FIG. 1. Crystal structure and Brillouin zone (BZ) of BaHgSn. (a) The crystal structure of BaHgSn primitive cell. *a*, *b*, and *c* are lattice vectors. The Ba ion (cyan) is at the spatial origin, and the Hg (red) and Sn (blue) ions form two AA' stacking honeycomb layers. (b) Top-down view of the crystal structure. (c) Bulk BZ of BaHgSn. The blue and red planes are two mirror planes of \overline{M}_z with $k_z = 0$ and $k_z = \pi$, respectively. The (010) and (001) surface BZs are plotted.

II. CRYSTAL STRUCTURE AND METHODOLOGY

The crystal structure of BaHgSn [38,39] is the same as KHgSb [18] as illustrated in Fig. 1. They belong to the same nonsymmorphic space group D_{6h}^4 ($P6_3/mmc$), and the symmetry operators include: an inversion P, a screw rotation \overline{C}_{6z} , mirror operations M_y , mirror operations $\overline{M}_z = t(c \vec{z}/2)M_{z}$, and glide mirror operations $\overline{M}_x = t(c \vec{z}/2)M_x$, where $t(c \vec{z}/2)$ is a c/2 translation along the z axis. The experimental lattice parameters of BaHgSn are a = b = 5.012 Å and c = 9.715 Å. The Ba, Hg, and Sn ions are at the same Wyckoff positions as K, Hg, and Sb ions in KHgSb. The Hg and Sn ions form AA' stacking honeycomb layers intercalated by trigonal layers of Ba ions.

We perform first-principles calculations of BaHgSn based on the projector augmented plane-wave method [40,41]. The Perdew-Burke-Ernzerhof exchange-correlation functional is used under generalized gradient approximation (GGA) [42]. The cutoff energy is 450 eV, and the k-point sampling grid is $12 \times 12 \times 4$ Γ -centered mesh. The hybrid functional calculations in the Heyd-Scuseria-Ernzerhof (HSE) [43,44] scheme are further applied to improve the shortcomings of GGA. To study the (010) and (001) surface states, the surface Green' s function technique is used to simulate the semi-infinite system by constructing the Wannier tight-binding Hamiltonian [45-47], which are generated for s orbitals of Hg and p orbitals of Sn. The Fermi surfaces and spin-dependent scattering probability which can simulate the QPI patterns detected by the STM are calculated according to the formula in Ref. [34] on the basis of Wannier Hamiltonian. Because STM is very sensitive to the surface features, the principle layers containing four unit cells is used to construct the surface Green's function but only the contributions in the outermost one unit cell are weighted. The Wilson-loop spectra are also computed to verify the topological features of surface states [45,46].



FIG. 2. The calculated bands of BaHgSn (a) without and (b) with spin-orbit coupling. The blue circles indicate the projection weight of Sn p_x and p_y orbitals, and the red circles indicate that of the Hg *s* orbitals. The inset of (b) shows the bulk Dirac cone and orbital characters of the wave functions at the Γ point.

III. RESULTS AND DISCUSSIONS

A. Electronic structures: Bulk analysis

The electronic structures are shown in Fig. 2. They are quite similar with KHgSb. The bands from Hg s orbitals are lower than those from Sn (or Sb in KHgSb) p orbitals. In the case without spin-orbit coupling (SOC), the valence and conduction bands at the A point around Fermi energy are degenerate since they are from Sn $p_{x,y}$ orbitals in the two Hg-Sn honeycomb layers in one unit cell. The interlayer coupling leads to the bonding-antiboning splitting along the Γ -A path. When SOC is considered as shown in Fig. 2(b), these bands further split into sub-bands with different angular momenta. It leads to a band inversion and a pair of Dirac nodes along the Γ -A path. The Dirac nodes are around -0.08 eV below the Fermi energy. The band inversion is about 0.48 eV at the Γ point, and further HSE calculation corrects it to be about 0.44 eV. Therefore, this band inversion is most plausible in realistic material.

Due to the existence of inversion symmetry, we can use bonding and antibonding states with definite parity to describe the low-energy Hamiltonian around Fermi energy. The bonding and antibonding states of p orbitals are



FIG. 3. The schematic diagrams show the evolution from the atomic *p* orbitals (blue line) and *s* orbitals (red line) to the conduction and valence states at Γ for (a) KZnP, (b) KHgSb, and (c) BaHgSn. The evolution includes the effects of bonding-antibonding splitting and spin-orbit coupling. Only the states having +*i* eigenvalue of \overline{M}_z ($k_z = 0$) are marked here. The inverted occupied orbitals are marked in pink and the inverted unoccupied ones are in green. The black dashed line represents the Fermi energy.

 $|P_{\alpha}^{\pm}\rangle = \frac{1}{\sqrt{2}}(|\text{Sn1}, p_{\alpha}\rangle \mp |\text{Sn2}, p_{\alpha}\rangle)$, where Sn1 and Sn2 are equivalent Sn atoms related with inversion symmetry, the superscripts \pm indicate the parity (the bonding or antibonding states) and α is x, y, z. When SOC is taken into account, the spin and orbital angular momentum are coupled, and the eigenstates have definite total angular momentum, written as $|P_{\frac{3}{2}}^{\pm}, \pm \frac{3}{2}\rangle$, $|P_{\frac{3}{2}}^{\pm}, \pm \frac{1}{2}\rangle$, and $|P_{\frac{1}{2}}^{\pm}, \pm \frac{1}{2}\rangle$, where the subscript indicates the total angular momentum J, the $\pm \frac{3}{2}, \pm \frac{1}{2}$ indicate the eigenvalues of J_z . The heavy-hole state $|P_{\frac{3}{2}}^{\pm}, \pm \frac{3}{2}\rangle$ and light-hole states $|P_{\frac{3}{2}}^{\pm},\pm\frac{1}{2}\rangle$ are no longer degenerated at the Γ point. The light-hole states $|P_{\frac{3}{2}}^{\pm}, \pm \frac{1}{2}\rangle$ and states $|P_{\frac{1}{2}}^{\pm}, \pm \frac{1}{2}\rangle$ will mix further to form the new eigenstates: $|P_{+}^{-}, \pm \frac{1}{2}\rangle, |P_{+}^{+}, \pm \frac{5}{2}\rangle$, $|P_{-}^{-},\pm\frac{1}{2}\rangle$, and $|P_{-}^{+},\pm\frac{5}{2}\rangle$ [48]. The states included in the band inversion and crossing points are states $|P_{+}^{-},\pm\frac{1}{2}\rangle$ and $|P_{\frac{3}{2}}^+,\pm\frac{3}{2}\rangle$ as shown in Fig. 2(c). Because of inversion and time-reversal symmetries, each band has Kramer degeneracy. The crossing points are fourfold degenerate Dirac nodes, locating at $(0, 0, k_z = \pm 0.35 \times \frac{\pi}{c})$. They are protected by C_{3z} rotation symmetry [9].

We compare the band structures of KZnP, KHgSb, and BaHgSn since they are iso-structural while have different topological quantum states. The key difference is the energy order of the bands around Fermi level at Γ , including *s* bands from Zn or Hg and p bands from of P, Sb, and Sn. The schematic diagrams for the evolution of these bands are shown in Fig. 3. KZnP is a trivial insulator and is used as a reference system having the normal order of bands, namely the cation's s bands are higher than the anion's p bands in energy. KHgSb is a well-known TCI and it has two Hg s orbitals inverting with two Sb p orbitals [18]. BaHgSn has one more band inversion among Sn p bands at the Γ point than KHgSb since the interlay coupling in BaHgSn is stronger than the strength of SOC. Sn and Sb atoms have the similar SOC strength as estimated from the splitting between $|P_{+}^{-}, \pm \frac{1}{2}\rangle$ and $|P_{\frac{3}{2}}^{-}, \pm \frac{3}{2}\rangle$), but BaHgSn has larger energy splitting in the interlayer bonding and antibonding states than KHgSb. This can be seen from the splitting between $|P_+^-, \pm \frac{1}{2}\rangle$ and $|P_+^+, \pm \frac{5}{2}\rangle$ or that between $|P_{\frac{3}{2}}^+,\pm\frac{3}{2}\rangle$ and $|P_{\frac{3}{2}}^-,\pm\frac{3}{2}\rangle$. It is due to the smaller *c* lattice constant of BaHgSn (9.71 Å) than that of KHgSb (10.22 Å). The weak interlayer bonding-antibonding splitting in KHgSb results in a TCI, but the stronger splitting in BaHgSn makes it a DSM. The difference in interlay coupling strength comes from the smaller lattice constant *c* of BaHgSn than that of KHgSb by about 0.51 Å. This may be caused by the stronger attractive interaction of the ±2 valence of Ba ions and Hg-Sn honeycomb layers than that of ±1 valence of K ions and Hg-Sb layers, although the ionic radius of K⁺¹ and Ba⁺² are nearly the same [49].

The states at Γ point are eigenstates of both inversion and \overline{M}_z . For occupied bands, the Z_2 invariant can be easily determined by counting the parity of their eigenstates [1,2]. It is known that the mCN C_i in mirror plane \overline{M}_z with $k_z = 0$ can be obtained by calculating the eigenvalues of rotation symmetry in subspace of eigenstates with either +i or -imirror eigenvalue [18,29,50], as follows:

$$e^{-i\frac{\pi}{3}C} = \prod_{i \in \text{occ.}} (-1)^F \eta_i(\Gamma) \theta_i(K) \zeta_i(M), \qquad (1)$$

where F = 0 (1) for integer-spin (half-integer-spin) system and $\eta_i(\mathbf{k})$, $\theta_i(\mathbf{k})$, and $\zeta_i(\mathbf{k})$ represent the eigenvalues of C_{6z} , C_{3z} , and C_{2z} rotation operators at the corresponding rotational invariant momenta, respectively.

We start from the trivial insulator KZnP in Fig. 3(a). Its Z_2 invariant and mCN for $k_z = 0$ plane are both zero. Compared with KZnP, KHgSb has two band inversions as shown in Fig. 3(b). We only consider the +i mirror eigenvalue subspace. The newly occupied orbitals are $|S^+, -\frac{1}{2}\rangle$ and $|S^-, +\frac{5}{2}\rangle$, and the newly unoccupied orbitals are $|P_{\frac{3}{2}}^-, -\frac{3}{2}\rangle$ and $|P_{\frac{3}{2}}^+, +\frac{3}{2}\rangle$. Obviously, the change in total parity $\Delta \delta = 0$ and that in rotational eigenvalue is $\Delta J_z = 2$, which means that $Z_2 = 0$ and the mCN of KHgSb is calculated from

$$e^{-i\frac{\pi}{3}C_i} = e^{-i\frac{\pi}{3}2}.$$
 (2)

The mCN of KHgSb is determined as $C_i = 2 \mod 6$. Compared with KHgSb, BaHgSn has one additional band inversion for +i mirror eigenvalue subspace in Fig. 3(c). The newly



FIG. 4. (a) The schematic diagram of surface bands for (010) surface. (b) The calculated (010) surface states. The light green dashed line and dot indicate the projection position of the Dirac node. (c) The calculated (001) surface states. Panels (d), (e), and (f) are the same as panels (a), (c), and (e), respectively, except that the 5% compressive strain is applied along the y axis. The horizontal dashed lines in (b) and (e) indicate the Fermi energy at 0.0 (charge neutral level), -0.1 and -0.33 eV, respectively.

occupied orbital is $|P_{\frac{1}{2}}^+, +\frac{3}{2}\rangle$, and the newly unoccupied orbital is $|P_{+}^-, +\frac{1}{2}\rangle$. Thus, the change of parity $\Delta\delta = 1$ and that of $\Delta J_z = 1$, indicating that the $Z_2 = 1$ and mCN $C_i = 3$ from

$$e^{-i\frac{\pi}{3}C_i} = e^{-i\frac{\pi}{3}(2+1)}.$$
(3)

The odd mCN of $k_z = 0$ plane means odd Z_2 , which is consistent with $\Delta \delta = 1$. If 5% compression strain is applied along the *y* axis in BaHgSn to break the C_{3z} rotation symmetry [9], the the Dirac nodes on Γ -*A* will be gapped out by about 9.2 meV, and BaHgSn becomes a strong TI with Z_2 indices of (1;000).

B. Surface analysis

After understanding the band topology of KZnP, KHgSb, and BaHgSn, we turn to their topological surface states. For KZnP, it might have trivial surface states. For KHgSb, on its (010) surface it has HSSs protected by glide symmetry \overline{M}_x and surface states due to $C_i = 2$ as protected by mirror symmetry \overline{M}_z ($k_z = 0$) [18,27]. For BaHgSn, we have calculated the surface states for (010) and (001) surfaces as shown in Fig. 4. Due to additional band inversion and different topological invariant, BaHgSn is expected to have some different features in surface states other than KHgSb.

For the (010) surface, the mirror \overline{M}_z and glide mirror \overline{M}_x are preserved. To explain the nontrivial surface states, we consider each high-symmetrical line in turn. Along the paths of $\tilde{X}\tilde{U}$ and $\tilde{Z}\tilde{\Gamma}$, \overline{M}_x is preserved. Because of time-reversal symmetry T, the eigenstates with +i, -i eigenvalue of \overline{M}_x are degenerated at \tilde{X} and $\tilde{\Gamma}$ ($k_z = 0$), and those of +1, +1 or -1, -1 are degenerated at \tilde{U} and \tilde{Z} $(k_z = \pi/c)$. These constraints imply two topologically distinct connectivities for the surface states, one of which is nontrivial zigzag connectivity across the band gap [51-55], the other one is hourglass-shaped dispersion. We can distinct these two type surface states by Z_2 invariant. For KHgSb, it has trivial Z_2 indices (0, 000), so that the HSSs exist along both $\tilde{X}\tilde{U}$ and $\tilde{Z}\tilde{\Gamma}$ paths. For BaHgSn, the similar HSSs appear only along the $\tilde{X}\tilde{U}$ path. Along the $\tilde{Z}\tilde{\Gamma}$ path, the surface states are not clear since the bulk Dirac node is projected onto this path as shown in Fig. 4(a). However, if the bulk Dirac node is gapped by breaking C_{3z} rotation symmetry, then the surface states along $\tilde{Z}\tilde{\Gamma}$ should be zigzag as shown in Fig. 4(d) since the Z_2 indices are (1;000) as a strong TI. The calculated (010) surface states are shown in Fig. 4(b). They are hourglass-shaped dispersions along the $\tilde{X}\tilde{U}$ path. Because of the overlap of bulk bands, the projection of the bulk Dirac cone and surface states are not clear along the $\tilde{Z}\tilde{\Gamma}$ path.

Along the $\tilde{U}\tilde{Z}$ ($k_z = \pi/c$) path, the symmetry is $T\overline{M}_x$, which results in a Kramers-like degeneracy. The surface states



FIG. 5. (a) The Fermi surface of (010) surface with Fermi energy $E_F = -0.1$ eV in the first surface BZ. The green arrows indicate the prominent scattering vectors. The light green dots indicate the projected positions of the Dirac nodes. (b) The corresponding QPI pattern simulated for (010) surface with $E_F = -0.1$ eV. [(c) and (d)] The Fermi surface and QPI pattern of (010) surface with $E_F = -0.33$ eV. [(e), (f), (g), and (h)] The Fermi surfaces and QPI patterns of (010) surface with $E_F = -0.33$ eV. [(e), applied along the y axis.

of BaHgSn are similar as those of KHgSb as shown in Figs. 4(a) and 4(b).

Along the $\tilde{\Gamma}\tilde{X}$ ($k_z = 0$) path, the surface states are determined by the mCN C_i of \overline{M}_z with $k_z = 0$ [56,57]. For KHgSb, there are two surface bands protected by $C_i = 2$ crossing the Fermi level as shown in Ref. [18]. But for BaHgSn, there are three surface bands protected by $C_i = 3$ crossing the Fermi level as shown in the Figs. 4(a) and 4(d), though such surface states are buried by the projection of bulk bands in Fig. 4(b).

For the (001) surface, the surface states of KHgSb are trivial. But for BaHgSn, the projection of bulk Dirac cone exists at the $\hat{\Gamma}$ point as shown in Fig. 4(c). When 5% compressive strain is applied along the y axis, the gapless Dirac points open gaps and the surface Dirac cone appears in Fig. 4(f).

C. Quasiparticle interference on (010) surface

The Fermi surface of (010) surface with Fermi energy E_F set at -0.1 eV is presented in Fig. 5(a). This makes E_F very close to the energy of Dirac nodes and the projections of two Dirac nodes are indicated with dots on the $\tilde{\Gamma}\tilde{Z}$ path. It can be clearly seen that there are surface states merging into the bulk Dirac cones to form Fermi arcs similar to those in DSMs Na₃Bi, Cd₃As₂, and others [9,10,58]. It is noted that there are two surface states meet each other to be doubly degenerate on the path \tilde{U} - \tilde{Z} , which is consistent with the discussion in former subsection. The weight of scatterings are mainly determined by local density of states, but the scatterings between surface states at **k** and $-\mathbf{k}$ points are prohibited, since the time-reversal symmetry *T* inverses their spin directions [1,32,33]. The green arrows mark the prominent scattering vectors, while the scatterings under time-reversal symmetry and mirror \overline{M}_z symmetry are not marked. The corresponding QPI pattern is shown in Fig. 5(b) with prominent scatterings marked as green arrows. It shows the scattering from surface states along $\tilde{\Gamma}\tilde{Z}$ path labeled as q_4 and crossing-shaped pattern at the Bragg point. The HSSs can be shown by Fermi surface and QPI pattern with E_F at -0.33 eV in Figs. 5(c) and 5(d). Two HSSs are shown along the $\tilde{X}\tilde{U}$ path, and two doubly degenerate surface states are shown along the $\tilde{U}\tilde{Z}$ path. The corresponding QPI pattern becomes ribbon-shaped along the k_z axis due to complicated scatterings. A crossing-shaped pattern still exist at Bragg point, and two small crossing-shaped patterns labeled as q_1 exist beside it.

When 5% compressive strain is applied along y axis, the bulk Dirac nodes are gapped. The Fermi surfaces and QPI patterns at -0.1 eV and -0.33 eV are shown in Figs. 5(e), 5(f), 5(g), and 5(h). The zigzag surface states exist along the $\tilde{\Gamma}\tilde{Z}$ path at -0.1 eV and their scattering q_3 can be found in the QPI pattern. The line-shaped QPI patterns along the k_z axis in Fig. 5(f) are generated from scatterings of line-shaped surface states. The surface states at -0.33 eV become narrow when compressive strain is applied, so that their QPI pattern becomes narrow from Figs. 5(d)-5(h), except the combination of crossing-shaped and rhombus-shaped QPI patterns emerges.

D. Wilson-loop spectra on (010) surface

The (010) surface states can also be described formally by Wilson-loop spectra based on maximally localized Wannier function for the occupied states [45,59]. Each Wannier function is an eigenfunction of the projected position operator [18]. The eigenvalue $y_{n,k_{\parallel}}$ is the center of onedimensional Wannier function along the k_{y} direction, which



FIG. 6. The Wilson-loop spectra for (010) surface of BaHgSn. (a) The case without compressive strain and (b) with 5% compressive strain along the y axis. The light green arrow and dot indicate the projection position of the Dirac node.

is a function of $\mathbf{k}_{\parallel} = (k_x, k_z)$ and is referred as Wilson-loop spectra [45].

We consider the full gapped occupied bands ($n_{occ} = 12$). The eigenvalue $y_{n,k_{\parallel}}$ is calculated without and with 5% compressive strain along the *y* axis in Fig. 6. The spectra have the topologically equivalent features as (010) surface states [60,61]: (i) the hourglass-shaped quadruplets along the $\tilde{X}\tilde{U}$ path; (ii) the degenerate doublets along the $\tilde{U}\tilde{Z}$ path; and (iii) the crossings protected by mCN $C_i = 3$ along the $\tilde{T}\tilde{X}$ path. However, the Wilson-loop spectra are not well defined at the projection point of Dirac node along the $\tilde{Z}\tilde{\Gamma}$ path in Fig. 6(a), since the band gap closes between the occupied and unoccupied bands. When compressive strain is applied, the

Wilson-loop spectra are well defined and the zigzag spectra appear along the $\tilde{Z}\tilde{\Gamma}$ path in Fig. 6(b).

IV. CONCLUSION

We have studied the bulk and surface electron structures of a Dirac semimetal BaHgSn. Though it is isostructural to KHgSb, it has smaller lattice constant c. This leads to larger interlayer bonding-antibonding splitting in the Sn p orbitals and results in additional band inversion at Γ when compared with KHgSb. This additional band inversion leads to bulk Dirac nodes along the Γ -A path and further modifies the topological phases. For KHgSb, its Z₂ indices are (0;000) and the mCN $C_i = 2$ for the mirror plan \overline{M}_z with $k_z = 0$. In contrast, the Z_2 indices are (1;000) and $C_i = 3$ for BaHgSn. On the (010) surface, the HSSs and doubly degenerate surface states exist along with the $\tilde{X}\tilde{U}$ and $\tilde{U}\tilde{Z}$ paths, which are the same as KHgSb. But along the $\tilde{Z}\tilde{\Gamma}$ path, BaHgSn has a projection of bulk Dirac nodes. Under compressive strain along the y axis, the Dirac nodes are gapped out and there are zigzag topological surface states protected by strong Z_2 indices in the gap. Along the $\tilde{\Gamma}\tilde{X}$ path, the topological surface bands are determined by the mCN C_i , which is two for KHgSb and three for BaHgSn. On the (001) surface, KHgSb is trivial. For BaHgSn, there is a projection of bulk Dirac cones in the case without compressive strain. The nontrivial surface Dirac cone due to strong TI appears under compressive strain along the y axis. The Fermi surfaces and QPI patterns by calculations reveal these features, such as the HSSs along $\tilde{X}\tilde{U}$ path, the zigzag surface states along $\tilde{\Gamma}\tilde{Z}$ path, and their corresponding QPI patterns. These results have been further clarified by the Wilson-loop spectra calculations and will help the ARPES and STM measurements to confirm them.

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- [1] M. Z. Hasan and C. L. Kane, Rev. Mod. Phys. 82, 3045 (2010).
- [2] X.-L. Qi and S.-C. Zhang, Rev. Mod. Phys. 83, 1057 (2011).
- [3] H. Weng, R. Yu, X. Hu, X. Dai, and Z. Fang, Adv. Phys. 64, 227 (2015).
- [4] C.-K. Chiu, J. C. Y. Teo, A. P. Schnyder, and S. Ryu, Rev. Mod. Phys. 88, 035005 (2016).
- [5] H. Weng, X. Dai, and Z. Fang, MRS Bull. 39, 849 (2014).
- [6] H. Weng, X. Dai, and Z. Fang, J. Phys.: Condens. Matter 28, 303001 (2016).
- [7] H. Weng, C. Fang, Z. Fang, and X. Dai, Natl. Sci. Rev. 4, 798 (2017).
- [8] N. P. Armitage, E. J. Mele, and A. Vishwanath, Rev. Mod. Phys. 90, 015001 (2018).
- [9] Z. Wang, Y. Sun, X.-Q. Chen, C. Franchini, G. Xu, H. Weng, X. Dai, and Z. Fang, Phys. Rev. B 85, 195320 (2012).
- [10] Z. Wang, H. Weng, Q. Wu, X. Dai, and Z. Fang, Phys. Rev. B 88, 125427 (2013).

- [11] P. Zhang, K. Yaji, T. Hashimoto, Y. Ota, T. Kondo, K. Okazaki, Z. Wang, J. Wen, G. D. Gu, H. Ding, and S. Shin, Science 360, 182 (2018).
- [12] H. Weng, C. Fang, Z. Fang, B. A. Bernevig, and X. Dai, Phys. Rev. X 5, 011029 (2015).
- [13] B. Q. Lv, H. M. Weng, B. B. Fu, X. P. Wang, H. Miao, J. Ma, P. Richard, X. C. Huang, L. X. Zhao, G. F. Chen, Z. Fang, X. Dai, T. Qian, and H. Ding, Phys. Rev. X 5, 031013 (2015).
- [14] S.-Y. Xu, I. Belopolski, N. Alidoust, M. Neupane, G. Bian, C. Zhang, R. Sankar, G. Chang, Z. Yuan, C.-C. Lee, S.-M. Huang, H. Zheng, J. Ma, D. S. Sanchez, B. Wang, A. Bansil, F. Chou, P. P. Shibayev, H. Lin, S. Jia, and M. Z. Hasan, Science 349, 613 (2015).
- [15] L. X. Yang, Z. K. Liu, Y. Sun, H. Peng, H. F. Yang, T. Zhang, B. Zhou, Y. Zhang, Y. F. Guo, M. Rahn, D. Prabhakaran, Z. Hussain, S. K. Mo, C. Felser, B. Yan, and Y. L. Chen, Nat. Phys. 11, 728 (2015).
- [16] Q. Wang, Y. Xu, R. Lou, Z. Liu, M. Li, Y. Huang, D. Shen, H. Weng, S. Wang, and H. Lei, Nat. Commun. 9, 3681 (2018).
- [17] E. Liu, Y. Sun, N. Kumar, L. Muechler, A. Sun, L. Jiao, S.-Y. Yang, D. Liu, A. Liang, Q. Xu, J. Kroder, Vicky, H. Borrmann, C. Shekhar, Z. Wang, C. Xi, W. Wang, W. Schnelle, S. Wirth, Y. Chen, S. T. B. Goennenwein, and C. Felser, Nat. Phys. 14, 1125 (2018).
- [18] Z. Wang, A. Alexandradinata, R. J. Cava, and B. A. Bernevig, Nature 532, 189 (2016).
- [19] J. Ma, C. Yi, B. Lv, Z. Wang, S. Nie, L. Wang, L. Kong, Y. Huang, P. Richard, P. Zhang, K. Yaji, K. Kuroda, S. Shin, H. Weng, B. A. Bernevig, Y. Shi, T. Qian, and H. Ding, Sci. Adv. 3, e1602415 (2017).
- [20] B. Bradlyn, J. Cano, Z. Wang, M. G. Vergniory, C. Felser, R. J. Cava, and B. A. Bernevig, Science 353, aaf5037 (2016).
- [21] H. Weng, C. Fang, Z. Fang, and X. Dai, Phys. Rev. B 93, 241202(R) (2016).
- [22] B. Q. Lv, Z. L. Feng, Q. N. Xu, X. Gao, J. Z. Ma, L. Y. Kong, P. Richard, Y. B. Huang, V. N. Strocov, C. Fang, H. M. Weng, Y. G. Shi, T. Qian, and H. Ding, Nature 546, 627 (2017).
- [23] Z. Zhu, G. W. Winkler, Q. S. Wu, J. Li, and A. A. Soluyanov, Phys. Rev. X 6, 031003 (2016).
- [24] T. Zhang, Y. Jiang, Z. Song, H. Huang, Y. He, Z. Fang, H. Weng, and C. Fang, Nature 566, 475 (2019).
- [25] M. G. Vergniory, L. Elcoro, C. Felser, N. Regnault, B. A. Bernevig, and Z. Wang, Nature 566, 480 (2019).
- [26] F. Tang, H. C. Po, A. Vishwanath, and X. Wan, Nature 566, 486 (2019).
- [27] Z. Song, T. Zhang, Z. Fang, and C. Fang, Nat. Commun. 9, 3530 (2018).
- [28] Y. Du, B. Wan, D. Wang, L. Sheng, C.-G. Duan, and X. Wan, Sci. Rep. 5, 14423 (2015).
- [29] S. Qin, C. Le, X. Wu, and J. Hu, J. Phys. Chem. Solids 128, 218 (2019).
- [30] H. Sun and J. Zhao, Sci. China: Phys., Mech. Astron. 59, 107011 (2016).
- [31] C.-S. Kuo, T.-R. Chang, S.-Y. Xu, and H.-T. Jeng, npj Comput. Mater. 5, 65 (2019).
- [32] D. Hsieh, Y. Xia, D. Qian, L. Wray, J. H. Dil, F. Meier, J. Osterwalder, L. Patthey, J. G. Checkelsky, N. P. Ong, A. V. Fedorov, H. Lin, A. Bansil, D. Grauer, Y. S. Hor, R. J. Cava, and M. Z. Hasan, Nature 460, 1101 (2009).

- [33] Z. Alpichshev, J. G. Analytis, J.-H. Chu, I. R. Fisher, Y. L. Chen, Z. X. Shen, A. Fang, and A. Kapitulnik, Phys. Rev. Lett. 104, 016401 (2010).
- [34] H. Inoue, A. Gyenis, Z. Wang, J. Li, S. W. Oh, S. Jiang, N. Ni, B. A. Bernevig, and A. Yazdani, Science **351**, 1184 (2016).
- [35] S. Kourtis, J. Li, Z. Wang, A. Yazdani, and B. A. Bernevig, Phys. Rev. B 93, 041109(R) (2016).
- [36] H. Li, S. Xu, Z.-C. Rao, L.-Q. Zhou, Z.-J. Wang, S.-M. Zhou, S.-J. Tian, S.-Y. Gao, J.-J. Li, Y.-B. Huang, H.-C. Lei, H.-M. Weng, Y.-J. Sun, T.-L. Xia, T. Qian, and H. Ding, Nat. Commun. 10, 5505 (2019).
- [37] Q.-Q. Yuan, L. Zhou, Z.-C. Rao, S. Tian, W.-M. Zhao, C.-L. Xue, Y. Liu, T. Zhang, C.-Y. Tang, Z.-Q. Shi, Z.-Y. Jia, H. Weng, H. Ding, Y.-J. Sun, H. Lei, and S.-C. Li, Sci. Adv. 5, eaaw9485 (2019).
- [38] F. Merlo, M. Pani, and M. L. Fornasini, J. Alloys Compd. 196, 145 (1993).
- [39] R. Vogel and H. U. Schuster, Zeitschr. Naturforsch. B 35, 114 (1980).
- [40] P. E. Blochl, Phys. Rev. B 50, 17953 (1994).
- [41] G. Kresse and D. Joubert, Phys. Rev. B 59, 1758 (1999).
- [42] J. P. Perdew, J. A. Chevary, S. H. Vosko, K. A. Jackson, M. R. Pederson, D. J. Singh, and C. Fiolhais, Phys. Rev. B 46, 6671 (1992).
- [43] J. Heyd, G. E. Scuseria, and M. Ernzerhof, J. Chem. Phys. 118, 8207 (2003).
- [44] J. Heyd, G. E. Scuseria, and M. Ernzerhof, J. Chem. Phys. 124, 219906 (2006).
- [45] R. D. King-Smith and D. Vanderbilt, Phys. Rev. B 47, 1651 (1993).
- [46] B. J. Wieder, B. Bradlyn, Z. Wang, J. Cano, Y. Kim, H.-S. D. Kim, A. M. Rappe, C. L. Kane, and B. A. Bernevig, Science 361, 246 (2018).
- [47] Q. Wu, S. Zhang, H.-F. Song, M. Troyer, and A. A. Soluyanov, Comput. Phys. Commun. 224, 405 (2018).
- [48] C.-X. Liu, X.-L. Qi, H. Zhang, X. Dai, Z. Fang, and S.-C. Zhang, Phys. Rev. B 82, 045122(2010).
- [49] R. D. Shannon, Acta Crystallogr. A 32, 751 (1976).
- [50] C. Fang, M. J. Gilbert, and B. A. Bernevig, Phys. Rev. B 86, 115112 (2012).
- [51] D. Hsieh, D. Qian, L. Wray, Y. Xia, Y. S. Hor, R. J. Cava, and M. Z. Hasan, Nature 452, 970 (2008).
- [52] M. Koenig, S. Wiedmann, C. Bruene, A. Roth, H. Buhmann, L. W. Molenkamp, X.-L. Qi, and S.-C. Zhang, Science 318, 766 (2007).
- [53] L. Fu, C. L. Kane, and E. J. Mele, Phys. Rev. Lett. 98, 106803(2007).
- [54] J. E. Moore and L. Balents, Phys. Rev. B 75, 121306 (2007).
- [55] R. Roy, Phys. Rev. B 79, 195321 (2009).
- [56] J. C. Y. Teo, L. Fu, and C. L. Kane, Phys. Rev. B 78, 045426 (2008).
- [57] A. Alexandradinata, Z. Wang, and B. A. Bernevig, Phys. Rev. X 6, 021008 (2016).
- [58] B. Peng, C. Yue, H. Zhang, Z. Fang, and H. Weng, npj Comput. Mater. 4, 68 (2018).
- [59] A. Alexandradinata, X. Dai, and B. A. Bernevig, Phys. Rev. B 89, 155114 (2014).
- [60] L. Fidkowski, T. S. Jackson, and I. Klich, Phys. Rev. Lett. 107, 036601 (2011).
- [61] Z. Huang and D. P. Arovas, Phys. Rev. B 86, 245109 (2012).