

# Quantum chaos associated with an emergent ergosurface in the transition layer between type-I and type-II Weyl semimetals

Yu-Ge Chen,<sup>1,2,3,\*</sup> Xi Luo,<sup>4,\*</sup> Fei-Ye Li,<sup>1,2,3</sup> Bin Chen,<sup>4</sup> and Yue Yu<sup>1,2,3</sup>

<sup>1</sup>State Key Laboratory of Surface Physics, Fudan University, Shanghai 200433, People's Republic of China

<sup>2</sup>Center for Field Theory and Particle Physics, Department of Physics, Fudan University, Shanghai 200433, People's Republic of China

<sup>3</sup>Collaborative Innovation Center of Advanced Microstructures, Nanjing 210093, People's Republic of China

<sup>4</sup>College of Science, University of Shanghai for Science and Technology, Shanghai 200093, People's Republic of China



(Received 4 June 2019; published 17 January 2020)

We present emergent ergosurfaces (ES) in a transition layer between type-I and type-II Weyl semimetals (WSMs). The Hawking temperature defined by the surface gravity at the acoustic event horizon which coincides with the ES when the tangent velocity  $v_{\parallel}$  is small is in a measurable interval. On the type-II WSM side, i.e., inside the ES when  $v_{\parallel}$  is large, the motion of the quasiparticles may be chaotic after a critical surface as they are governed by an effective inverted oscillator potential induced by the mismatch between the type-I and type-II Weyl nodes. In a relevant lattice model, we calculate out-of-time-ordered correlators (OTOCs). We find that the OTOCs are fast scrambling with a quantum Lyapunov exponent in high temperature and the scrambling is saturated after the Ehrenfest time. This confirms the quantum chaotic behavior.

DOI: [10.1103/PhysRevB.101.035130](https://doi.org/10.1103/PhysRevB.101.035130)

## I. INTRODUCTIONS

Recent developments on the frontier of theoretical physics research have entangled black hole horizon physics with quantum chaos [1–7].

In condensed matter systems, the black hole analog had been pioneered by Unruh in transsonic fluid flow [8]. A quantum analog had been provided in Bose-Einstein condensation [9,10], and a magnonic black hole was predicted [11]. The exotic emergent geometries (or gravities) have also been studied in Fermi surface [12], fractional quantum Hall effects [13–16], graphene [17], deformed crystal [18], type-II Weyl semimetal (WSM) [19–22], and type-III Dirac semimetal [23]. Several systems may simulate the Hawking evaporation at a black hole event horizon [24,25]. However, the chaotic behaviors at the horizon do not appear easily. For instance, for the trajectory of a moving particle outside of the horizon to be chaotic classically, there must be an external potential near the horizon [26]. The Lyapunov exponent is subject to a maximal chaotic bound, the surface gravity [1]. This bound is saturated for theories with anti-de Sitter/conformal field theory duality, such as the case in the Sachdev-Ye-Kitaev model [2–4].

Volovik *et al.* recently studied the emergent metric in the inhomogeneous WSM [19–21]. They found a general correspondence between the emergent vielbein  $e^i_{\mu}$  and the effective Hamiltonian near a Weyl node  $\mathbf{K} = (K_x, K_y, K_z)$ , i.e.,

$$H(\vec{q}) = e^i_{\mu} q_i \sigma^{\mu}, \quad (1)$$

where  $\mu = 0, i$  with  $i = x, y, z$  and  $q_i = k_i - K_i$  and  $\sigma^{\mu} = (I, \sigma^i)$  are the identity and Pauli matrices. Here the Einstein's summation convention for repeated indices is used.

The “speed of light”  $v_F$  (the Fermi velocity), the electron effective mass  $m_b$ , and  $\hbar$  are set to be 1 unless they are explicitly restored. The spectrum is given by

$$E(Q_i)_{\pm} = e^j_0 [e^{-1}]^j_i Q_i \pm |\mathbf{Q}|, \quad (2)$$

with  $Q_i \equiv e^j_i q_j$ . The vielbein components  $e^i_{\mu}$  together with  $e^0_0 = -1$  and  $e^0_i = 0$  define an emergent acoustic metric  $g^{\mu\nu} = \eta^{\alpha\beta} e^{\mu}_{\alpha} e^{\nu}_{\beta}$  with the signature  $\eta^{\mu\nu} = \text{diag}(-1, 1, 1, 1)$  [19]. With this vielbein choice, the conical spectrum (2) can be written as

$$g^{\mu\nu} q_{\mu} q_{\nu} = 0, \quad (3)$$

where  $q_0 = E$  and the dual metric  $g_{\mu\nu}$  defines the light cone with  $g_{\mu\nu} x^{\mu} x^{\nu} = 0$ . Therefore, the ergosurface (ES) of this emergent geometry is determined by  $g_{00} = -(1 - \mathbf{v}^2) = 0$  [27], i.e.,  $|\mathbf{v}| = 1$ , where  $v^i = e^j_0 [e^{-1}]^j_i$ . A Weyl node is called type I (type II) if  $|\mathbf{v}| < 1$  ( $|\mathbf{v}| > 1$ ). Notice that the type of each one in a pairwise nodes can be arbitrary [28,29]. From this effective geometry point of view, quasiparticles associated with the type-I (type-II) Weyl node effectively live outside (inside) the ES.

Volovik uses the spherically symmetric metric to study inhomogeneous WSM [20], which is difficult to be realized in reality. Both type-I and type-II WSMs have been found [30–36]. In order to simulate an acoustic event horizon (which coincides with the ES when the normal velocity  $v_{\perp}$  dominates [27]) in a realizable geometry and study the chaos phenomena, we consider a type-I/type-II WSM transition layer (TL) where the type-I and type-II Weyl nodes may mismatch [see Fig. 1(a)]. In the following, we will assume that  $v_{\perp}$  dominates whenever the horizon is mentioned. Within the TL, a black/white hole planar horizon emerges effectively at a plane with  $|\mathbf{v}_z| = 1$  in the TL. A fermionic shock wave and the Hawking evaporation are possible to be detected because

\*These two authors contribute equally.

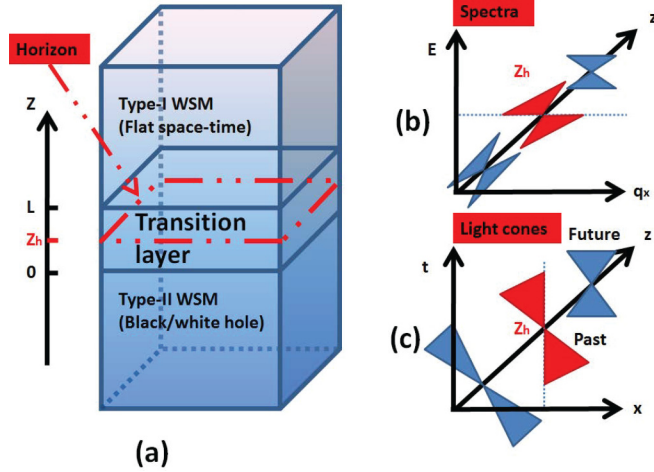


FIG. 1. (a) The TL between Type-I and Type-II WSMs. The acoustic event horizon is located at  $z = z_h$  with  $v_z(z_h) = 1$ . (b) The spectra outside, at, and inside the horizon, corresponding to type-I, critical, and type-II WSM's. (c) The light cones outside, at, and inside the horizon. In (b) and (c), we take  $\alpha_z(z) = 0$  and project to the  $x$  direction for simplicity.

the Hawking temperature may be as high as several tens of Kelvin.

In the effective Hamiltonian approach, we further find that when the Weyl nodes of the type I and type II do not completely match, the TL states inside and outside the ES may be totally different: On the type-I WSM side, the TL state is described by the Landau levels of the WSM in an effective magnetic field. On the type-II side, after a characteristic plane, the TL states are governed by an effective inverted oscillator potential [37–39] and thus are chaotic due to the nonintegrability caused by coupling to the environment potential [26].

To further study the quantum chaos of the quasiparticles in the TL, we study a lattice model which reduces to the effective chaotic model in the TL in the long-wavelength limit. We calculate out-of-time-ordered correlators (OTOCs) of the quantum states which characterize the quantum chaotic behavior [1–7,40]. For a proper hopping strength in the TL with a sufficient layer thickness, we find that in a lower temperature, the OTOC does not show a fast scrambling over a short time while it does when the temperature is higher and a quantum Lyapunov exponent is obtained. The OTOCs saturate after the Ehrenfest time. These two properties of OTOC indicate the existence of quantum chaos inside the ES [41].

## II. EFFECTIVE MODEL FOR TL

We consider that type-I and type-II WSMs couple as shown in Fig. 1(a). The TL is located between  $z = 0$  and  $z = L$ . Assuming  $\mathbf{K}^I$  ( $\mathbf{K}^{II}$ ) is a Weyl node of the type-I (type-II) WSM in the regime  $z > L$  ( $z < 0$ ), and  $|\mathbf{K}^I - \mathbf{K}^{II}| \ll K_0$ , where  $K_0$  is the distance between a pair of Weyl nodes in the bulk. The effective Hamiltonian in the bulk of the type-I/II WSM is given by

$$H_{I/II} = (k_i - K_i^{I/II})\sigma^i - \alpha_{x,I/II}(k_x - K_x^{I/II}) - \alpha_{z,I/II}(k_z - K_z^{I/II}), \quad (4)$$

where  $\alpha_{I/II} = (\alpha_{x,I/II}, 0, \alpha_{z,I/II})$  are parameters that tilt the Dirac cones. For the type-I WSM,  $|\alpha_I| = \sqrt{\alpha_{xI}^2 + \alpha_{zI}^2} < 1$  and  $|\alpha_{II}| = \sqrt{\alpha_{xII}^2 + \alpha_{zII}^2} > 1$  for the type-II WSM. We further assume  $|\Delta\alpha| = |\alpha_{II} - \alpha_I| \ll 1$ . The small difference in the Weyl nodes can be viewed as a soft edge between the WSMs. The soft edge is achieved by a small linear deformation between the lattice constants, which corresponds to a linear interpolation between the Weyl nodes in the lowest order. Therefore, the TL effective Hamiltonian can be approximated by a linear interpolation between  $H_I$  and  $H_{II}$  [42],

$$H(z) = H_{II} + (H_I - H_{II})\frac{z}{L}, \quad (5)$$

where  $0 \leq z \leq L$ . Defining  $\mathbf{q} = \mathbf{k} - \mathbf{K}^{II}$  and  $B_i = \frac{K_i^I - K_i^{II}}{L} = \frac{\Delta K_i}{L}$ ,

$$H(z) = (q_x + B_x z)\sigma^x + (q_y + B_y z)\sigma^y + (-i\partial_z + B_z z)\sigma^z - \alpha_x(z)(q_x + B_x z) - \alpha_z(z)(-i\partial_z + B_z z) + \dots, \quad (6)$$

where “...” are high-order terms including  $O(\Delta\mathbf{K} \cdot \Delta\alpha)$ , etc.;  $q_z$  is replaced by  $-i\partial_z$ ; and  $B_x$  ( $B_y$ ) can be thought as an effective magnetic field in the  $y$  ( $x$ ) direction. One can perform a gauge transformation to change  $-i\partial_z + B_z z$  to  $-i\partial_z$  if the wave function  $\psi$  changes to  $e^{-i\frac{B_z z^2}{2}}\psi$  accordingly.  $\alpha(z)$  is a linear interpolation like  $H(z)$ , i.e.,  $\alpha(z) = \alpha_{II} - \Delta\alpha\frac{z}{L}$ . Since  $|\Delta\alpha| \ll 1$ ,  $\alpha(z)$  is a slow-varying vector function of  $z$ .  $|\alpha(z_c)| = 1$  determines the critical surface with  $z = z_c$  which separates the type-I and type-II WSM, where

$$\frac{z_c}{L} = \frac{\Delta\alpha \cdot \alpha_{II} - \sqrt{(\Delta\alpha \cdot \alpha_{II})^2 - (\Delta\alpha)^2(\alpha_{II}^2 - 1)}}{(\Delta\alpha)^2}. \quad (7)$$

## III. EMERGENT GEOMETRY AND HORIZON

To see the emergent geometry, we first consider  $B_i = \frac{\Delta K_i}{L} = 0$  for simplicity. Comparing the TL Hamiltonian (6) with the effective Hamiltonian (1), we can identify the vielbein:  $e_0^x = -\alpha_x(z)$ ,  $e_0^y = -\alpha_y(z) = 0$ ,  $e_0^z = -\alpha_z(z)$ , and  $e_j^i = \delta_j^i$ . Therefore the effective geometry in the TL is described by the line element of the induced metric [19]

$$ds^2 = -[1 - |\mathbf{v}(z)|^2]dt^2 - 2v^i(z)dr_i dt + (dr)^2, \\ = -d^2t + (dr_i - v_i dt)^2, \quad (8)$$

with  $v^i = e_0^i = -\alpha_i(z)$  and  $|\mathbf{v}| = |\alpha(z)| = \sqrt{\alpha_x^2 + \alpha_z^2}$ . The ES, which is denoted as  $z = z_c$ , is defined by  $g_{00} = -[1 - \mathbf{v}(z)^2] = 0$  and the acoustic event horizon is defined by  $v_z(z_h) = 1$  [27,43]. In the limit  $\frac{v_z^2}{v_z^2} \rightarrow 0$ , the ES coincides with the event horizon which we denote as  $z = z_h$ . For instance, we sketch the spectra and light cones with  $\alpha_z(z) = 1$  in Figs. 1(b) and 1(c). In that emergent metric, the Ricci scalar  $R$  at the

acoustic event horizon is given by

$$R_{z_h} = \frac{\Delta\alpha_x^2[-2(1-\alpha_{IIz})^2\Delta\alpha_x^2 - 4\alpha_{IIx}(1-\alpha_{IIz})\Delta\alpha_1\Delta\alpha_3 + (1-2\alpha_{IIx}^2)\Delta\alpha_z^2]}{2\Delta\alpha_z^2}, \quad (9)$$

which reduces to zero when  $\Delta\alpha_x \rightarrow 0$ . Furthermore, the surface gravity  $\eta_H$  can be identified by introducing  $L^\mu = (1, v_x, 0, 0)$  which is the null vector field on the acoustic event horizon with the integral curves generating the horizon [43],

$$L^\nu \partial_\nu L^\mu|_{z=z_h} = \frac{\partial(1-v_z^2)}{2\partial z} L^\mu = \eta_H L^\mu, \quad (10)$$

namely

$$\eta_H = \frac{\partial(1-v_z^2)}{2\partial z}|_{z=z_h} = \Delta\alpha \cdot \alpha(z_h) \frac{1}{L}. \quad (11)$$

The Hawking temperature of this emergent black hole horizon is defined by [43]

$$T_H = \frac{\hbar\eta_H}{2\pi k_B} = \frac{\Delta\alpha \cdot \alpha(z_h)\hbar v_F}{2L\pi k_B}, \quad (12)$$

where we have explicitly restored the Fermi velocity whose typical value for WSM is  $v_F \sim 10^6$  m/s. If the thickness of the TL  $L \lesssim 10$  nm,  $T_H \gtrsim 10^2 \Delta\alpha \cdot \alpha(z_h)$  K. If  $\Delta\alpha \cdot \alpha(z_h) \sim 0.1$ , then the Hawking temperature may arrive at several tens of Kelvin which are experimentally reachable.

In these discussions, we have assumed the Fermi velocity  $v_F = 1$  is a constant in the low-energy effective theory. If the interactions and quantum fluctuations are considered, then  $v_F$  may be renormalized or even spatial dependent. To exactly calculate the Hawking temperature, one needs to solve the Weyl equation with the emergent metric (8), and then a Hawking evaporation may be observed by a thermal spectrum of the TL fermion when the layer is forming [20,44]. Similarly to the transsonic wave in the bosonic fluid [8,45], we shall have a fermion shock wave with a velocity  $\mathbf{v}(z)$ , which will be presented elsewhere.

#### IV. CHAOTIC TRANSITION LAYER

To show the chaotic behavior in the effective theory, we consider the Dirac equation with mismatching Weyl nodes  $0 < \Delta K_i \ll K_0$ ,

$$H(z)\psi(z, t) = i\dot{\psi}(t), \quad (13)$$

where  $H(z)$  is given by Eq. (6) and  $\psi^T = (\chi, \zeta)$  is a two-component spinor. Without loss of generality, we take  $q_y = 0$ . Because  $\alpha(z)$  slowly varies, we neglect the  $\partial_z \alpha(z)$  term. Then the dominant part of  $\chi$ 's equation in the large  $\tilde{z}$  limit reads

$$\ddot{\chi} - \partial_{\tilde{z}}^2 \chi + \left[ \kappa \tilde{z}^2 + B_y^2 \left( q_x^2 - \frac{2\sqrt{1-\alpha_z^2}}{B_x} q_x \tilde{z} \right) \right] \chi + \dots = 0, \quad (14)$$

where  $\tilde{z} = \frac{q_x + B_{xz}}{B_x \sqrt{1-\alpha_z^2}}$  and  $\kappa = (1-\alpha^2)B_x^2 + (1-\alpha_z^2)B_y^2$  (see Appendix A for “...” terms in more detail). For  $\kappa = \omega^2 > 0$  with a fixed  $q_x$ , Eq. (14) coincides with the Landau levels of a Weyl fermion outside the ES.

Equation (14) will become an inverted oscillator when  $\kappa < 0$  [37–39]. This can be satisfied for  $|\alpha| > 1$  and  $|\alpha_z| < 1$  after a characterized plane  $z_c$  inside the ES [ $\mathbf{v}(z_c)^2 = 1$  if  $B_y = 0$ ] [46]. An inverted oscillator is known to lead to the chaotic dynamics when it couples to the environment  $V(x)$  [26,47,48]. The detailed form of  $V(x)$  is not important as long as the potential  $V(x)$  is a confining one. We therefore introduce an environment potential  $V(x)$  and then Eq. (14) becomes

$$\ddot{\chi} - \partial_{\tilde{z}}^2 \chi - \lambda_L^2 \tilde{z}^2 \chi - B_y^2 \partial_x^2 + \epsilon_3 \tilde{z}^3 \chi + \epsilon_4 \tilde{z}^4 \chi + 2iB_y^2 B_x^{-1} \sqrt{1-\alpha_z^2} \tilde{z} \partial_x \chi + V(x)\chi = 0, \quad (15)$$

where  $\lambda_L^2 = (\alpha_{II}^2 - 1)B_x^2 + (\alpha_{z,II}^2 - 1)B_y^2 > 0$ ,  $\epsilon_3 \sim O(\Delta\alpha \cdot \alpha_{II})$ , and  $\epsilon_4 \sim O[(\Delta\alpha)^2] > 0$  (see Appendix A). If  $V(x) = 0$ , then Eq. (15) is integrable with a positive Lyapunov exponent  $\lambda_L$  and basically describes an inverted oscillator. To explicitly show the classical chaos, we study the steady-state solution of (15) and treat  $E^2 = \mathcal{E}$  as a classical energy. The Poincaré sections for the classical limit of (15) with  $V(x) \sim x^2$  show that the Kolmogorov-Arnold-Moser tori keep nice periodic orbits in lower  $\mathcal{E}$  but break down as  $\mathcal{E}$  increases, which indicates the existence of classical chaos. For details, see Appendix A.

At the end of this section, we would like to point out that, in the Hawking radiation (12), the temperature is proportional to  $|\alpha_z(z_h)| = v_F$ , while the Lyapunov exponent remains positive as long as  $|\alpha_x| > 1$  even if  $\alpha_z = 0$ . The Hawking radiation is not directly related to the chaotic behavior in the TL. In other words, the chaos survives even if the ES does not coincide with the event horizon.

#### V. LATTICE MODEL

To further show the quantum chaos in the TL, we study a lattice model on a cubic lattice with a lattice spacing  $a = 1$ . For simplicity, we do not consider the environment disorder while the Lyapunov exponent remains positive. The periodic boundary conditions are implicated while the thickness in the  $z$  direction is finite. We label the sizes of type-I/II WSMs and the TL as  $L_{I/II}$  and  $L$ , respectively. We consider minimal models describing the type-I/II WSMs on a periodic cubic lattice with Hamiltonian  $H_{I/II}^I$  [49]

$$H_{I/II}^I = d_{I/II,\mu} \sigma^\mu, \quad (16)$$

where

$$\begin{aligned} d_{I/II,x} &= 2\tilde{t}(\cos k_x - \cos K_x^{I/II}) + 2\tilde{t}[1 - \cos(k_y - K_y^{I/II})] \\ &\quad + \frac{\tilde{t}}{2}(\cos 3k_x - \cos 3K_x^{I/II}) + 2\tilde{t}\gamma(1 - \cos k_z), \\ d_{I/II,y} &= 2\tilde{t} \sin(k_y - K_y^{I/II}), \quad d_{I/II,z} = -2\tilde{t}\gamma \sin k_z, \\ d_{I,0} &= 0, \quad d_{II,0} = -2\eta_{II}(\cos k_x - \cos K_x^{II}). \end{aligned} \quad (17)$$

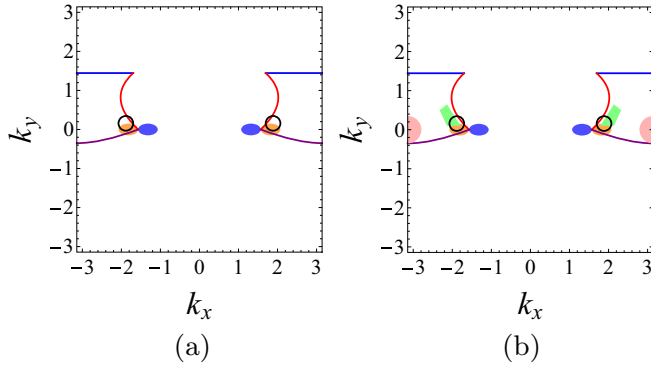


FIG. 2. The zero-energy Fermi arcs and pockets in the  $k_x$ - $k_y$  plane. (a)  $\gamma_I = \gamma_{II} = \gamma_{TL} = 1$ . (b)  $\gamma_I = \gamma_{II} = 1$  and  $\gamma_{TL} = 0.7$ . The red Fermi arcs lie in the TL and connect to the Weyl nodes' images of the WSMs. The orange and blue regions correspond to the quasiparticle and quasihole Fermi pockets in the type-II WSM. Approximately, the states inside the black circle are described by the effective inverted oscillator and lie in the ES. The additional areas in (b) are the pink quasiparticle pockets belonging to the interval inside the ES in the TL, and the green quasiparticle pockets are very close to the ES from the type-II WSM side.

Here  $\tilde{t}$  is the hopping strength. We introduce  $\gamma$ , which is a  $z$ -dependent parameter [see Eq. (C2)], to fine-tune the hopping strength in the  $z$  direction. The high-energy states which may be quantum chaotic and far away from the Weyl points can be moved near the Fermi level by tuning  $\gamma$ . The Weyl nodes are located at  $\mathbf{K}^{I/II} = (K_x^{I/II}, K_y^{I/II}, 0)$ . The lattice Hamiltonian we use here is given by

$$H^I = \begin{cases} H_{II}^I, & \gamma = \gamma_{II}, n_z \leq L_{II}, \\ H_{II}^I + \xi(H_I^I - H_{II}^I), & \gamma = \gamma_{TL}, n_z \in (L_{II}, L + L_{II}], \\ H_I^I, & \gamma = \gamma_I, n_z > L + L_{II}, \end{cases} \quad (18)$$

where  $n_z$  is the lattice site index in the  $z$  direction and  $\xi = \frac{n_z - L_{II}}{L}$ . In the long-wavelength limit, the effective Hamiltonian in the TL reduces to the form of Eq. (5) to a  $z$ -dependent function which leads to a line of Weyl nodes (see Appendix C).

Taking  $\mathbf{K}^I = (0.54\pi, 0.46\pi, 0)$ ,  $\mathbf{K}^{II} = (0.5\pi, 0, 0)$ , and  $\eta_{II} = 0.8\tilde{t}$  and considering a half filling lattice with  $L = L_I = L_{II} = 100$ , we sketch the Fermi arcs and pockets projected to the  $k_x$ - $k_y$  plane with the colored regions in Fig. 2. There are the Fermi arcs on the surfaces of WSMs and the Fermi pockets in the type-II WSM as expected. In the TL, there are Fermi arcs that connect the type-I Weyl nodes' images with the type-II ones with the same chirality. The Weyl nodes' lines slightly deviate from these Fermi arcs and the ES is at  $\xi_h \approx 0.18$  (see Appendix C). According to the band structures of the lattice model (see Appendix B), there is no Fermi pocket in the TL for  $\gamma_I = \gamma_{II} = \gamma_{TL} = 1$  [Fig. 2(a)], while Fermi pockets emerge in the TL for  $\gamma_I = \gamma_{II} = 1$  and  $\gamma_{TL} = 0.7$  by tuning  $\gamma$  [Fig. 2(b)].

## VI. OTOCS

The sensitivity of the initial value in quantum chaos could be measured by OTOCs [1]. We consider the simplest OTOC between the position and momentum operators in the  $z$  direction which is measurable in the transport experiment [40]. In the semiclassical limit, this OTOC describes the dependence on the initial condition of the particle motion. On the lattice, these operators correspond to the site operator  $n_z$  and the difference operator  $\hat{P}_z$ , which is defined by

$$\hat{P}_z |\psi_{\mathbf{k}}(n_z)\rangle \equiv \frac{1}{2} [|\psi_{\mathbf{k}}(n_z - 1)\rangle - |\psi_{\mathbf{k}}(n_z + 1)\rangle], \quad (19)$$

where  $\mathbf{k} = (k_x, k_y)$  and  $|\psi_{\mathbf{k}}(n_z)\rangle$  is the state on site  $n_z$ . The OTOCs are defined by

$$\begin{aligned} C_T(t) &= \frac{1}{Z} \text{Tr}[e^{-\beta H^I} [n_z(t), \hat{P}_z(0)]^2] \\ &= \sum_{\psi} f_{\psi} \langle \psi | [n_z(t), \hat{P}_z(0)]^2 | \psi \rangle, \end{aligned} \quad (20)$$

where  $n_z(t)$  and  $\hat{P}_z(t)$  are the time-dependent operators in Heisenberg's picture and  $f_{\psi} = \frac{1}{e^{\beta E_{\psi}} + 1}$  is the Fermi distribution function. More explicitly, the OTOCs are given by

$$C_T(t) = \sum_{\mathbf{k}} C_T(t, \psi_{\mathbf{k}}), \quad (21)$$

where the  $\mathbf{k}$ -dependent OTOC is given by

$$\begin{aligned} C_T(t, \psi_{\mathbf{k}}) &= \sum_{\psi_1, \psi_2, \psi_3} f_{\psi} (1 - f_{\psi_1}) \\ &\quad \times (1 - f_{\psi_2})(1 - f_{\psi_3}) \rho_{\psi\psi_1} \rho_{\psi_1\psi_2} \rho_{\psi_2\psi_3} \rho_{\psi_3\psi} \\ &\quad \times [n_z(t)_{\psi\psi_1} P_z(0)_{\psi_1\psi_2} n_z(t)_{\psi_2\psi_3} P_z(0)_{\psi_3\psi} + \dots], \end{aligned} \quad (22)$$

$O_{\psi\psi'} = \langle \psi | \hat{O} | \psi' \rangle$  for  $\hat{O} = n_z(t)$  and  $\hat{P}_z(0)$ .  $\rho_{\psi\psi'}$  is the probability that the state  $|\psi'\rangle$  transits to the state  $|\psi\rangle$  due to a disturbance. For convenience, we take  $\rho_{\psi\psi'} = e^{-\beta(E_{\psi} - E_{\psi'})}$  for  $E_{\psi} > E_{\psi'}$  or 1 otherwise;  $O_{fi} = \langle f | \hat{O} | i \rangle$  for  $\hat{O} = n_z(t)$  and  $\hat{P}_z(0)$ . “ $\dots$ ” is the other three terms when explicitly writing out  $[n_z(t), \hat{P}_z(0)]^2$ .

In numerical calculations, we take a lower temperature  $T = 1/\beta = \tilde{t}/4$  and a higher  $T = 2\tilde{t}/3$ . For the model with  $\gamma_I = \gamma_{II} = \gamma_{TL} = 1$ , the OTOC with the lower  $T$  has a small magnitude and does not have an exponential fast scrambling, as shown in Fig. 3(a). In the inset in Fig. 3(a), the negative slope of the  $\ln C_T(t)$  in the very short time implies an exponential decay of the OTOC. The reason lies in the binding potential of the quasiparticles on the lattice site at the beginning. After that time period,  $\ln C_T(t)$  behaves logarithmic. As the temperature raises, the fast scrambling of the OTOC appears and its magnitude becomes three orders larger than that in the lower temperature as shown in Fig. 3(b). In the inset of Fig. 3(b), the OTOC still decays exponentially in a very short time. After this very short time period, a time-dependent positive quantum Lyapunov exponent  $\lambda_{\text{QL}}(t)$  can be read out according to an exponential fitting  $C_T(t) \sim e^{2\lambda_{\text{QL}}(t)t}$ . There is a time period  $t_Q \approx 100$  in which  $\lambda_{\text{QL}} = 0.006$  is constant [50]. This implies the fast scrambling of the OTOC. The emergence of the quantum chaotic behavior in the higher temperature is consistent with the classical chaos in the effective model

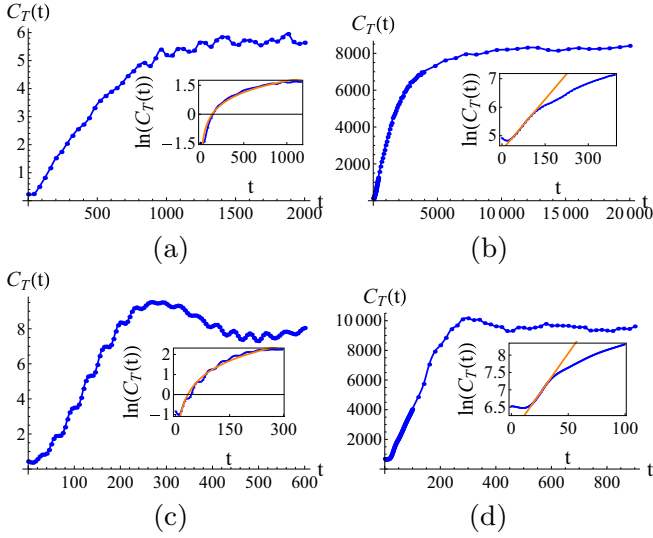


FIG. 3. The OTOCs  $C_T(t)$ . (a)  $\gamma_{TL} = 1$  and  $T = \tilde{t}/4$ . (b)  $\gamma_{TL} = 1$  and  $T = 2\tilde{t}/3$ . (c)  $\gamma_{TL} = 0.7$  and  $T = \tilde{t}/4$ . (d)  $\gamma_{TL} = 0.7$  and  $T = 2\tilde{t}/3$ . The insets are a logarithmic fitting (red curve) of the  $\ln C_T(t)$  in the short time for (a) and (c), while the ones in (b) and (d) include the linear region (red line) of the  $\ln C_T(t)$  with  $\lambda_{QL} = 0.006$  and  $\lambda_{QL} = 0.023$ , respectively.

because the chaos appears at higher energy. After  $t_Q$ , due to the quantum fluctuation, the fast scrambling begins to be suppressed and  $\ln C_T(t)$  deviates from the linear fitting. After a long time  $t_E \gtrsim 5000$ , the OTOC is saturated instead of growing everlastingly in the classical chaos. This time  $t_E$  corresponds to the Ehrenfest time. This shows a possible distinction between the quantum chaos and the classical one [41].

For  $\gamma_I = \gamma_{II} = 1$  and  $\gamma_{TL} = 0.7$ , as shown in Figs. 3(c) and 3(d), the fast scrambling in the OTOC is absent for the lower temperature while it exists in the higher temperature, similar to that for  $\gamma_{TL} = 1$ . However, the time  $t_Q \approx 40$  and the saturated time  $t_E \approx 300$  are shorter than those for  $\gamma_{TL} = 1$ . This is because the green pockets emerge in Fig. 2(b), which increases the density of states and then enhances the quantum fluctuation. Meanwhile, the higher density of states also raises the quantum Lyapunov exponent in linear region about 4 times,  $\lambda_{QL} = 0.023$ . This means that the scrambling is much faster than that for  $\gamma_{TL} = 1$ . In both cases of  $\gamma_{TL} = 1$  and  $\gamma_{TL} = 0.7$ , the OTOCs are dominated by the states inside the

black circle which corresponds to the inverted oscillator in the effective theory. Besides the black circle area, there are other states contribute to the positive Lyapunov exponent, such as the pink area in Fig. 2(b) which emerge by tuning  $\gamma_{TL}$ . These states are not predicted by the effective Hamiltonian (5).

We have yet to consider the  $x$ - $y$  plane environment potential in the lattice model. The integrability of the model is lost if it is turned on and then the fast scrambling of the OTOC and the Ehrenfest time identify the quantum chaos.

## VII. CONCLUSIONS

We pointed out that there are emergent ESs which coincides with acoustic event horizons under certain conditions in the TL between type-I and type-II WSMs. And the corresponding Hawking temperature may be measurable. When the Weyl nodes of type I and type II mismatch, the quasiparticles moving inside the ES may be quantum chaotic, even if the event horizon and ES do not coincide. We confirm this quantum chaotic behavior in a lattice model by calculating the OTOCs of the system. Two diagnostic quantities, the quantum Lyapunov exponent and the Ehrenfest time, which characterize the quantum chaos, were determined.

## ACKNOWLEDGMENTS

This work is supported by NNSF of China with Grant No. 11474061 (Y.G.C., X.L., and Y.Y.) and Grant No. 11804223 (X.L.) and the Ministry of Science and Technology of China with Grant No. 2016YFA0301001 (F.Y.L.).

## APPENDIX A: DETAILS ON SOLVING EQ. (13) IN THE MAIN TEXT AND ITS CLASSICAL POINCARÉ SECTION

The equations of motion (13) in the main text for the wave function  $\psi = (\chi, \lambda)^T$  read

$$(q_x + B_x z)\lambda - i(q_y + B_y z)\lambda + (-i\partial_z + B_z z)\chi - \alpha_x(q_x + B_x z)\chi - \alpha_z(z)(-i\partial_z + B_z z)\chi = i\partial_t \chi, \quad (\text{A1})$$

$$(q_x + B_x z)\chi + i(q_y + B_y z)\chi - (-i\partial_z + B_z z)\lambda - \alpha_x(q_x + B_x z)\lambda - \alpha_z(z)(-i\partial_z + B_z z)\lambda = i\partial_t \lambda. \quad (\text{A2})$$

Without loss of generality, we take  $q_y = 0$ . Notice that the  $B_z z$  term is a pure gauge and can be gauged away by a gauge transformation  $\psi \rightarrow \exp(-i\frac{B_z z^2}{2})\psi$ . Solving Eq. (A1), we have  $\lambda = \lambda(\chi)$  and then substitute it into Eq. (A2),

$$\left\{ [(1 - \alpha_x^2)(q_x + B_x z)^2 + B_y^2 z^2] - \partial_z^2 + 2i\alpha_x(q_x + B_x z)(\alpha_z \partial_z - \partial_t) - 2\alpha_x \partial_z \partial_t + \alpha_z^2 \partial_z^2 + \partial_t^2 \right\} \chi - \frac{1}{q_x + (B_x - iB_y)z} \left\{ \alpha_x(1 + \alpha_z)B_y q_x + (\alpha_z^2 B_x - B_x + iB_y - i\alpha_z^2 B_y)\partial_z + [(1 - \alpha_z)B_x + i(1 + \alpha_z)B_y \partial_t] \right\} \chi = 0. \quad (\text{A3})$$

When  $q_x + B_x z \gg 1$ , the second line of Eq. (A3) can be neglected. Making a transformation  $\chi \rightarrow \exp(i\frac{\alpha_x \alpha_z (q_x + B_x z)^2}{2B_x(1 - \alpha_z^2)})\chi$ , Eq. (A3) becomes

$$\left\{ i\alpha_x \alpha_z (1 - \alpha_z^2) B_x + \alpha_x^2 (q_x + B_x z)^2 - (1 - \alpha_z^2) [(q_x + B_x z)^2 + B_y^2 z^2] \right\} \chi + (1 - \alpha_z^2)^2 \partial_z^2 \chi + 2i\alpha_x (q_x + B_x z) \partial_t \chi + 2\alpha_z (1 - \alpha_z^2) \partial_t \partial_z \chi - (1 - \alpha_z^2) \partial_t^2 \chi = 0. \quad (\text{A4})$$

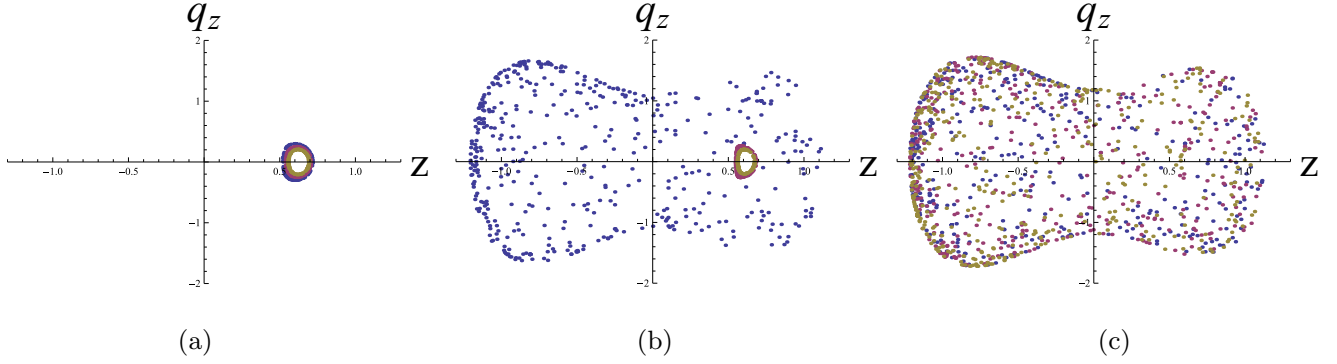


FIG. 4. The Poincaré section for Hamiltonian (A10). For simplicity, we choose  $B_x = -1$ ,  $B_y = 1$ ,  $\alpha_x = 2$ ,  $\alpha_z = 0$ ,  $\epsilon_3 = 0$ ,  $\epsilon_4 = 4$ , and  $\epsilon_x = 4$ . The colors in the diagrams represent the KAM (Kolmogorov–Arnold–Moser) tori for different initial conditions:  $q_z(0) = 0.3, 0.25, 0.2, z(0) = 0.6$ , and  $x(0) = 0$ . The energies are chosen as follows: (a)  $E = 1$ , (b)  $E = 1.2$ , and (c)  $E = 1.4$ . The breaking down of the KAM tori indicates the existence of classical chaos.

Defining  $\tilde{z} = \frac{q_x + B_x z}{B_x \sqrt{1 - \alpha_z^2}}$ , Eq. (A4) reads

$$\partial_t^2 \chi - \partial_{\tilde{z}}^2 \chi + \left[ (1 - \alpha_x^2 - \alpha_z^2) B_x^2 \tilde{z}^2 + B_y^2 \left( \sqrt{1 - \alpha_z^2} \tilde{z} - \frac{q_x}{B_x} \right)^2 \right] \chi - i \alpha_x \alpha_z B_x \chi - 2 \alpha_z \sqrt{1 - \alpha_z^2} \partial_0 \partial_{\tilde{z}} \chi - \frac{2i \alpha_x B_x \tilde{z}}{\sqrt{1 - \alpha_z^2}} \partial_t \chi = 0. \quad (\text{A5})$$

Equation (A5) describes a harmonic oscillator when  $\omega^2 = (1 - \alpha_x^2 - \alpha_z^2) B_x^2 + (1 - \alpha_z^2) B_y^2 > 0$  or an inverted one with positive Lyapunov exponent  $\lambda_L$  when  $-\lambda_L^2 = (1 - \alpha_x^2 - \alpha_z^2) B_x^2 + (1 - \alpha_z^2) B_y^2 < 0$ . The leading-order terms are the first line of Eq. (A5), and the last three terms only shift the oscillating center and the zero point energy. Since we have treated the slow-varying functions  $\alpha = \alpha_{II} - \Delta \alpha_L^z$  as constants in the derivation of Eq. (A5), the Lyapunov exponent  $\Lambda_L$  can be expressed as

$$\begin{aligned} \lambda_L^2(\tilde{z}) &= \left[ \alpha_{II}^2 - 2 \alpha_{II} \cdot \Delta \alpha_L^z + \left( \Delta \alpha_L^z \right)^2 - 1 \right] B_x^2 + \left[ \alpha_{zII}^2 - 2 \alpha_{zII} \Delta \alpha_z \frac{z}{L} + \left( \Delta \alpha_z \frac{z}{L} \right)^2 - 1 \right] B_y^2 \\ &= \left[ (\alpha_{II}^2 - 1) + \left( \frac{2 \alpha_{II} \cdot \Delta \alpha q_x}{B_x L} + \frac{\Delta \alpha^2 q_x^2}{B_x^2 L^2} \right) - \frac{2 \sqrt{1 - \alpha_{zII}^2} (B_x \Delta \alpha \cdot \alpha_{II} L + \Delta \alpha^2 q_x) \tilde{z}}{B_x L^2} + \frac{(\alpha_{zII}^2 - 1) \Delta \alpha^2 z^2}{L^2} \right] B_x^2 \\ &\quad + \left[ (\alpha_{zII}^2 - 1) + \left( \frac{2 \alpha_{zII} \Delta \alpha_z q_x}{B_x L} + \frac{\Delta \alpha_z^2 q_x^2}{B_x^2 L^2} \right) - \frac{2 \sqrt{1 - \alpha_{zII}^2} (B_x \Delta \alpha_z \alpha_{zII} L + \Delta \alpha_z^2 q_x) \tilde{z}}{B_x L^2} + \frac{(\alpha_{zII}^2 - 1) \Delta \alpha_z^2 z^2}{L^2} \right] B_y^2. \quad (\text{A6}) \end{aligned}$$

Therefore the chaos-related terms in Eq. (A5) read

$$\partial_t^2 \chi - \partial_{\tilde{z}}^2 \chi - \lambda_{LII}^2 \tilde{z}^2 \chi - 2 B_y^2 B_x^{-1} \sqrt{1 - \alpha_z^2} q_x \tilde{z} \chi + B_y^2 B_x^{-2} q_x^2 \chi + \epsilon_3 \tilde{z}^3 \chi + \epsilon_4 \tilde{z}^4 \chi + \dots = 0, \quad (\text{A7})$$

where

$$\epsilon_3 = \frac{2 \sqrt{1 - \alpha_{zII}^2} [(B_x \Delta \alpha \cdot \alpha_{II} L + \Delta \alpha^2 q_x) B_x^2 + (B_x \Delta \alpha_z \alpha_{zII} L + \Delta \alpha_z^2 q_x) B_y^2]}{B_x L^2}, \quad (\text{A8})$$

$$\epsilon_4 = \frac{(1 - \alpha_{zII}^2) (\Delta \alpha^2 B_x^2 + \Delta \alpha_z^2 B_y^2)}{L^2}. \quad (\text{A9})$$

The first three terms in Eq. (A7) represent an inverted oscillator if  $\Lambda_{LII}^2 > 0$ . The  $\tilde{z}^3$  and  $\tilde{z}^4$  terms are the quantum perturbations, and the  $q_x \tilde{z}$  term is the coupling between  $\tilde{z}$  and  $q_x$ , the “environment.” Therefor Eq. (A7) can be viewed as an intrinsic quantum open system and chaos raises.

To be more specific about the emergence of chaos, we consider the Poincaré section of the classical limit of Eq. (A7) up to the  $\tilde{z}^4$  terms, i.e., the classical Hamiltonian reads,

$$\begin{aligned} H &= q_{\tilde{z}}^2 - \lambda_{LII}^2 \tilde{z}^2 + \epsilon_3 \tilde{z}^3 + \epsilon_4 \tilde{z}^4 - 2 B_y^2 B_x^{-1} \sqrt{1 - \alpha_z^2} q_x \tilde{z} \\ &\quad + B_y^2 B_x^{-2} q_x^2 + \epsilon_x x^2, \quad (\text{A10}) \end{aligned}$$

where  $q_{\tilde{z}}$  is the momentum for  $\tilde{z}$ , and we also add a harmonic potential along the  $x$  direction which simulates the effects of disorders. We draw its Poincaré section numerically in Fig. 4 which clearly shows the existence of classical chaos.

## APPENDIX B: BAND STRUCTURE AND FERMI ENERGY

We give the band structures of the lattice model. Figure 5(a) is for  $\gamma_I = \gamma_{II} = \gamma_{TL} = 1$  and Fig. 5(b) is for  $\gamma_I = \gamma_{II} = \gamma_{TL} = 1$ . Corresponding zero-energy Fermi arcs and

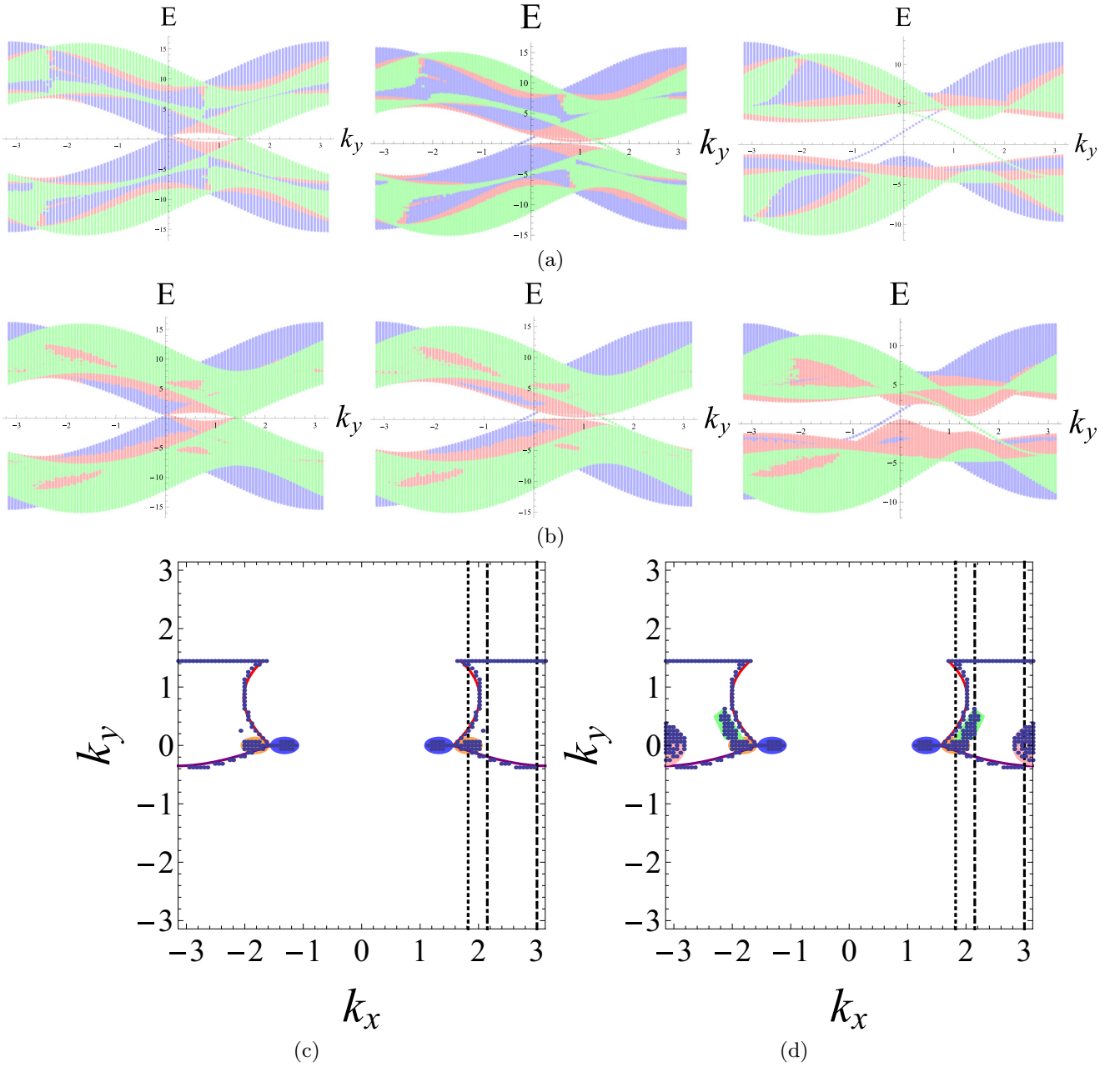


FIG. 5. The band structures of the Hamiltonian (17) in the main text with  $L_I = L_{II} = L = 100$ . The first panel is for  $\gamma_I = \gamma_{II} = \gamma_{TL} = 1$  with  $k_x = 1.8, 2.15, 3$  (see the vertical dashed lines in the third panel). The second panel is for  $\gamma_I = \gamma_{II} = 1$  and  $\gamma_{TL} = 0.7$  with  $k_x = 1.8, 2.15, 3$ . The third panel is the original data images of Fig. 2 in the main text. The zero-energy state projections crossing with the vertical dashed lines can be read out from the first and second panels. The states in the red (green and blue) bands belong to the transition layer (the bulk of type-I WSM and the bulk of type-II WSM).

pockets projected to the  $k_x$ - $k_y$  are shown in Fig. 5(c) and Fig. 5(d), respectively, and they are the original data for sketching Figs. 2(a) and 2(b).

### APPENDIX C: WEYL POINTS AND ES

Writing the lattice Hamiltonian (16) in the main text explicitly

$$H^l = d_\mu^l \sigma^\mu, \quad (\text{C1})$$

where

$$d_\mu^l = \begin{cases} d_{\mu,II}^l, & \gamma = \gamma_{II}, \text{ for } n_z < n_{II}, \\ d_{\mu,II}^l + \xi(d_{\mu,I}^l - d_{\mu,II}^l), & \gamma = \gamma_{TL}, \text{ for } n_{II} < n_z < n_I, \\ d_{\mu,I}^l, & \gamma = \gamma_I, \text{ for } n_z > n_I, \end{cases} \quad (\text{C2})$$

and

$$d_{I/II,x} = 2\tilde{t}(\cos k_x - \cos K_x^{I/II}) + 2\tilde{t}[1 - \cos(k_y - K_y^{I/II})] + \frac{\tilde{t}}{2}(\cos 3k_x - \cos 3K_x^{I/II}) + 2\tilde{t}\gamma(1 - \cos k_z),$$

$$d_{I/II,y} = 2\tilde{t} \sin(k_y - K_y^{I/II}), d_{I/II,z} = -2\tilde{t}\gamma \sin k_z,$$

$$d_{I,0} = 0, d_{II,0} = -2\eta_{II}(\cos k_x - \cos K_x^{II}), \quad (C3)$$

where  $n_z$  is the lattice site index in the  $z$  direction and  $\xi = \frac{n_z - L_{II}}{L}$ . Then the Weyl points are determined by

$$d_i^l = 0, i = x, y, z. \quad (C4)$$

At Weyl points,

$$\partial_i d_j^l = 0, \quad \text{if } i \neq j. \quad (C5)$$

The Weyl points in the transition layer are running as  $\xi$  and their projection to the  $k_x$ - $k_y$  plane is depicted in Fig. 6(a). Notice that these Weyl points do not lie on the Fermi surface because of the interlayer coupling. If the transition layer is replaced by an insulating layer with a large band gap, then there will be ordinary Fermi arcs that connect Weyl nodes with opposite chiralities in the type-I (-II) WSM on the top (bottom) surface of the insulating layer. After turning on the hopping term between the type-I and type-II WSMs in the transition layer, the Fermi arcs will reconstruct from the surface Fermi arcs into the ones connecting type-I and type-II Weyl nodes with the same chirality inside the transition layer [see the red Fermi arcs in Fig. 6(a)].

Expanding the Hamiltonian (C2) around the Weyl points and comparing the result with the effective Hamiltonians (3) and (4) in the main text, one obtains,

$$|\alpha| = \sqrt{\sum_{i=x,y,z} (\partial_i d_0^l / \partial_i d_i^l)^2} = |\partial_x d_0^l / \partial_x d_x^l|. \quad (C6)$$

The ES can be obtained by solving  $|\alpha| = 1$ , which is at  $\xi_h \approx 0.18$ , reading out from 6(b).

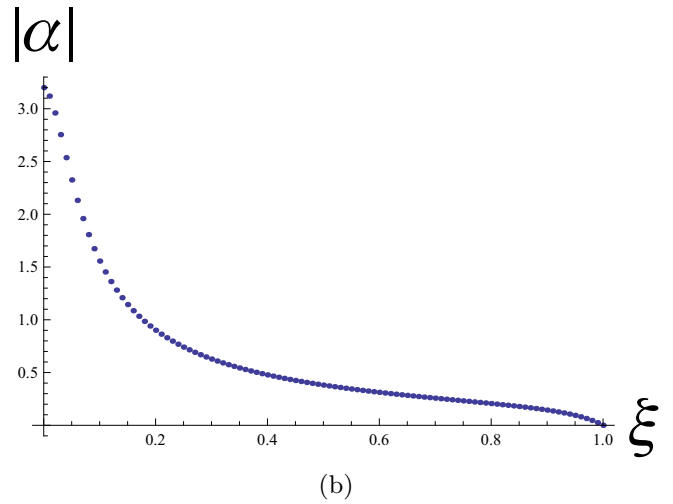
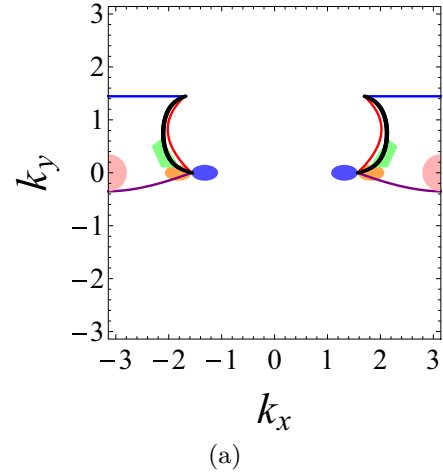


FIG. 6. (a) The Weyl points in the transition layer (black curves). We copy the zero-energy regions in Fig. 2(b) of the main text for the reference. (b) The function  $|\alpha|$  of  $\xi$ .  $|\alpha(\xi_h)| = 1$  determines the ES.

- [1] J. Maldacena, S. H. Shenker, and D. Stanford, *J. High Energy Phys.* **1608** (2016) 106.
- [2] A. Kitaev, <http://online.kitp.ucsb.edu/online/joint98/kitaev/>.
- [3] A. Kitaev, <http://online.kitp.ucsb.edu/online/entangled15/kitaev/>, <http://online.kitp.ucsb.edu/online/entangled15/kitaev2/>.
- [4] J. Maldacena and D. Stanford, *Phys. Rev. D* **94**, 106002 (2016).
- [5] S. Sachdev, *Phys. Rev. Lett.* **105**, 151602 (2010).
- [6] S. H. Shenker and D. Stanford, *J. High Energy Phys.* **03** (2014) 067.
- [7] S. H. Shenker and D. Stanford, *J. High Energy Phys.* **05** (2015) 132.
- [8] W. G. Unruh, *Phys. Rev. Lett.* **46**, 1351 (1981).
- [9] O. Lahav, A. Itah, A. Blumkin, C. Gordon, S. Rinott, A. Zayats, and J. Steinhauer, *Phys. Rev. Lett.* **105**, 240401 (2010).
- [10] J. Raman, M. de Nova, K. Golubkov, V. Kolobov, and J. Steinhauer, *Nature* **569**, 688 (2019).
- [11] A. Roldan-Molina, A. S. Nunez, and R. A. Duine, *Phys. Rev. Lett.* **118**, 061301 (2017).
- [12] P. Horava, *Phys. Rev. Lett.* **95**, 016405 (2005).
- [13] F. D. M. Haldane, *Phys. Rev. Lett.* **107**, 116801 (2011).
- [14] R. Z. Qiu, S. P. Kou, Z. X. Hu, X. Wan, and S. Yi, *Phys. Rev. A* **83**, 063633 (2011).
- [15] K. Yang, *Phys. Rev. B* **93**, 161302(R) (2016).
- [16] X. Luo, Y.-S. Wu, and Y. Yu, *Phys. Rev. D* **93**, 125005 (2016).
- [17] G. E. Volovik and M. A. Zubkov, *Ann. Phys.* **356**, 255 (2015).
- [18] L. Dong and Q. Niu, *Phys. Rev. B* **98**, 115162 (2018).
- [19] G. E. Volovik, *JETP Lett.* **104**, 645 (2016).
- [20] G. E. Volovik, *Phys. Usp.* **61**, 89 (2018).
- [21] J. Nissinen and G. E. Volovik, *JETP Lett.* **105**, 447 (2017).
- [22] K. Zhang and G.E. Volovik, *J. Low Temp. Phys.* **189**, 276 (2017).
- [23] H. Q. Huang, K.-H. Jin, and F. Liu, *Phys. Rev. B* **98**, 121110(R) (2018).
- [24] S. W. Hawking, *Nature* **248**, 30 (1974); *Commun. Math. Phys.* **43**, 199 (1975).
- [25] J. D. Bekenstein, *Phys. Rev. D* **7**, 2333 (1973).



- [26] K. Hashimoto and N. Tanahashi, *Phys. Rev. D* **95**, 024007 (2017).
- [27] U. R. Fischer and G. E. Volovik, *Int. J. Mod. Phys. D* **10**, 57 (2001).
- [28] H. B. Nielsen and M. Ninomiya, *Nucl. Phys. B* **185**, 20 (1981); **193**, 173 (1981); *Phys. Lett. B* **130**, 389 (1983).
- [29] F.-Y. Li, X. Luo, X. Dai, Y. Yu, F. Zhang, and G. Chen, *Phys. Rev. B* **94**, 121105(R) (2016).
- [30] X. Wan, A. M. Turner, A. Vishwanath, and S. Y. Savrasov, *Phys. Rev. B* **83**, 205101 (2011).
- [31] H. M. Weng, C. Fang, Z. Fang, B. A. Bernevig, and X. Dai, *Phys. Rev. X* **5**, 011029 (2015).
- [32] S.-M. Huang, S.-Y. Xu, I. Belopolski, C.-C. Lee, G. Q. Chang, B. K. Wang, N. Alidoust, G. Bian, M. Neupane, C. L. Zhang, S. Jia, A. Bansil, H. Lin, and M. Z. Hasan, *Nat. Commun.* **6**, 8373 (2015).
- [33] S.-Y. Xu, I. Belopolski, N. Alidoust, M. Neupane, G. Bian, C. L. Zhang, R. Sankar, G. Q. Chang, Z. J. Yuan, C.-C. Lee *et al.*, *Science* **349**, 613 (2015).
- [34] B. Q. Lv, H. M. Weng, B. B. Fu, X. P. Wang, H. Miao, J. Ma, P. Richard, X. C. Huang, L. X. Zhao, G. F. Chen *et al.*, *Phys. Rev. X* **5**, 031013 (2015).
- [35] L. X. Yang, Z. K. Liu, Y. Sun, H. Peng, H. F. Yang, T. Zhang, B. Zhou, Y. Zhang, Y. F. Guo, M. Rahn *et al.*, *Nat. Phys.* **11**, 728 (2015).
- [36] A. A. Soluyanov, D. Gresch, Z. J. Wang, Q. S. Wu, M. Troyer, X. Dai, and B. A. Bernevig, *Nature* **527**, 495 (2015).
- [37] G. Barton, *Ann. Phys.* **166**, 322 (1986).
- [38] C. Yuce, A. Kilic, and A. Coruh, *Phys. Scr.* **74**, 114 (2006).
- [39] T. Morita, *Phys. Rev. Lett.* **122**, 101603 (2019).
- [40] A. I. Larkin and Y. N. Ovchinnikov, *Sov. Phys. JETP* **28**, 1200 (1969).
- [41] K. Hashimoto, K. Murata, and R. Yoshii, *J. High Energy Phys.* **10** (2017) 138.
- [42] S. Tchoumakov, M. Civelli, and M. O. Goerbig, *Phys. Rev. B* **95**, 125306 (2017).
- [43] M. Visser, *Class. Quant. Grav.* **15**, 1767 (1998).
- [44] H. Liu, J.-T. Sun, H. Q. Huang, F. Liu, and S. Meng, [arXiv:1809.00479v2](https://arxiv.org/abs/1809.00479v2).
- [45] G. E. Volovik, *The Universe in a Helium Droplet* (Clarendon Press, Oxford, 2003).
- [46] We do not discuss the case with  $|\alpha_z| > 1$ . The ES is not stable in this case.
- [47] W. H. Zurek and J. P. Paz, *Phys. Rev. Lett.* **72**, 2508 (1994).
- [48] P. A. Miller and S. Sarkar, *Phys. Rev. E* **58**, 4217 (1998).
- [49] T. M. McCormick, I. Kimchi, and N. Trivedi, *Phys. Rev. B* **95**, 075133 (2017).
- [50] The quantum Lyapunov exponent  $\lambda_{\text{QL}}$  does not need to be exactly the same as  $\lambda_L$  due to the quantum correction.