Thermal loading on crystals in an x-ray free-electron laser oscillator

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X-ray free electron laser oscillators (XFELOs) are future light sources that produce fully coherent hard x-ray pulses. Based on the low-gain principle, the XFELO traps x-ray pulses in an optical cavity composed of multiple Bragg-reflecting mirrors that have high reflectivity in a bandwidth of about ten meV. The crystal mirrors exposed to intense x-ray beams in the optical cavity are subject to thermal deformations that would shift and distort the Bragg reflectivity curve. Therefore, the stability of the XFELO operation depends on the ability of the mirrors to maintain the Bragg reflection under such thermal load. A new approach was used to analyze the thermal load of the mirrors. The approach utilizes a dedicated Bragg reflection physical process in GEANT4 to obtain precise absorption information of the XFELO pulses in the crystal. Following transient thermal behavior, including single pulse and multiple pulse inputs, was analyzed by finite element analysis software based on the energy absorption information extracted from the GEANT4 simulation. It is shown that, for a typical XFELO pulse depositing about ten microjoules energy the over a spot of tens of micrometers in radius, the thermal relaxation time across the thickness is on tens of nanoseconds scale. In this situation, a simplified heat-load model is then developed to integrate the heat load in the XFELO. With the simplified model, the potential impact of the thermal load on the XFELO operation is estimated. When a large amount of heat remains in the crystal, the pulse energy drops significantly and has large oscillations due to negative feedback of the temperature change on the pulse energy.

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I. INTRODUCTION

In the hard x-ray regime, the operating free electron lasers (FELs) are based on the SASE (self-amplified spontaneous emission) mode [1–[5\],](#page-8-0) which can generate x-ray pulses with unique characteristics, such as ultrahigh peak power and ultrashort pulse duration. However, originating from shot noise, the stochastic nature of SASE leads to a low temporal coherence and poor pulse-to-pulse stability of produced x-ray pulses. To obtain stable, fully coherent x-ray pulses with statistical properties similar to a fully coherent optical laser, an x-ray free electron laser oscillator (XFELO) has been proposed [6–[8\]](#page-8-1). In an XFELO, the x-ray pulses amplified by the interaction with electron bunches in a short undulator, circulate in a low-loss optical cavity formed by multiple Bragg-reflecting crystals, which have a reflectivity close to 100% within a bandwidth of ∼10 meV [\[9,10\]](#page-9-0) in the hard x-ray regime, whereas conventional mirrors are unavailable.

Generally, the operation of XFELOs is based on an accelerator that can deliver electron bunches with a repetition rate of 1 MHz level or above. With high-brightness electron bunches at MHz repetition rate, the intracavity x-ray pulses could have a pulses energy of about 800 μ J and beam radius of about 30 μ m. Such intense x-ray pulses inevitably impose a high heat load on the mirrors, which result in lattice distortions in the crystal. Since the Bragg reflection originates from the x-ray scattering of a periodic atomic lattice, the lattice distortions would decrease the peak of Bragg reflectivity, shift reflectivity curves over the spectrum, and increase sidebands [\[11,12\].](#page-9-1) Therefore, the realization of the optical cavity with Bragg-reflecting mirrors relies highly on the stability of the crystal lattice. The heat load of the crystal mirrors under such intense x-ray pulses is an essential issue of building an XFELO.

In order to investigate the influence of the heat load in crystals, two kinds of experiments have been performed [\[13](#page-9-2)–19]. The first one utilizes a high power conventional laser in the long-wavelength regime to model laser-mirror interactions. The advantage of this method is the ability of precisely controlling the pulse duration and pulse energy, which is crucial for light-material interactions. However, this approach produces a penetration depth at nanometer level [\[14\]](#page-9-3), which is far shorter than that produced by x-rays. Besides, the repetition rate of the high power conventional

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laser is challenging to reach MHz or higher. Another method applies the synchrotron radiation light source. The synchrotron radiation light source can generate hard x-rays with sufficient time-averaged power. However, the synchrotron radiation is polychromatic, and the monochromator is required to obtain high spectral purity. With the monochromator, the pulse energy of the synchrotron radiation light sources decreases significantly [\[15\]](#page-9-4). Besides, if an x-ray beam generated by synchrotron radiation is micro-focused to the XFEL beam level [\[18\]](#page-9-5), its angular spreads may be much larger than Bragg reflection acceptance. Thus, in general, the experimental conditions do not meet the requirements of XFELO operation, and it is still necessary to numerically model the heat load of mirrors and its feedback on XFELO operation.

In previous studies [\[20](#page-9-6)–22], the numerical evaluation of the x-ray absorption in the crystal uses an exponential attenuation model that describes the intensity of the x-rays decreases to 1/e of initial value through one attenuation length. Although the exponential attenuation model is feasible in most of the x-ray interactions [\[23,24\]](#page-9-7), it needs to be improved in the x-ray Bragg reflection. Therefore, this paper proposes a simple new physical process to describe the Bragg reflection in GEANT4, a powerful particle tracker. Combined with the other implemented physical process in the GEANT4, a complete set of interactions can be provided in the simulation of light-material interactions. Based on the heat source obtained by GEANT4 simulations, transient thermal analysis can be conducted with finite-element analysis (FEA) software, ANSYS [\[25\].](#page-9-8) Furthermore, by coupling the temperature evolution of the crystals into the XFELO simulations, one can gain the ability to investigate the impact of the thermal loading on an XFELO operation.

This paper is organized as follows. The Sec. [II](#page-1-0) presents the characteristics of Bragg reflections relating to the thermal load. In the Sec. [III](#page-2-0), the light-matter interaction is simulated in GEANT4, and the following thermal behavior is analyzed. In the Sec. [IV,](#page-6-0) a simplified model to integrate the thermal loading into XFELO simulations is presented. Then, the simulations of thermal loading coupled XFELO are performed. A summary is presented in the Sec. [V.](#page-8-2)

II. CHARACTERISTICS OF BRAGG REFLECTIONS

In an XFELO, the single-pass gain is G , while the total reflectance is R. When $(1 + G)R > 1$, the x-ray pulses could evolve from the initial noise to intense coherent radiation. The oscillator reaches steady state when $(1+G)R = 1$. As the single-pass gain G is governed by the electron beam qualities and undulator configurations, the ability to achieve the stable XFELO operation relies on the reflectivity R , which is sensitive to the lattice structure of the crystals. Adverse effects including the thermal deformations and mounting vibrations of the crystals, may emerge to affect the

FIG. 1. A symmetric x-ray diffraction. The blue lines are the virtual atomic planes. k is the wave number of incident x-ray, and k_H is reflected one. θ_B is the Bragg angle.

lattice and to disrupt the Bragg reflection. Here, we present a brief introduction of the relationship between the lattice distortion and the Bragg reflection.

The Bragg reflection or diffraction originates from the x-ray scattering of atoms in a periodic structure [\[26\],](#page-9-9) which can be imagined as a reflection of imaginary "mirrors" formed by atomic planes in the crystal lattice, seen in Fig. [1](#page-1-1). The Bragg condition can be expressed as

$$
n\lambda_{\rm B} = 2d_{\rm H}(T)\sin\theta_{\rm B},\tag{1}
$$

where *n* is the diffraction order, θ_B is the angle of incidence, λ_B is the radiation wavelength at which the reflection occurs, d_H is the distance between the atomic planes and T is the temperature. The lattice planes is described by the Miller indices, h , k , l , three integers. The spacing between the lattice planes is given by $d_H(T) = \frac{a(T)}{\sqrt{h^2 + k^2 + l^2}}$, where a is the lattice constant. Diamond is a preferred material that can be employed by a high power XFELO to form the x-ray cavity, as it is a unique combination of outstanding thermal and optical properties, including a high thermal diffusivity, low thermal expansion, and high Bragg reflectivity for x-rays [\[27\].](#page-9-10) In the following discussions, the crystal mirror is assumed to be the diamond.

As expected, the high reflectivity in Bragg reflection is only possible if the lattice structure is perfect for a large amount of diffracting planes. In general, the change of crystal lattice acts on both the position and the amplitude of the Bragg reflectivity [\[11,12\]](#page-9-1). Figure [2](#page-2-1) shows three kinds of lattice distortions (up panel) and corresponding effect (bottom panel). In each condition, the undistorted lattice (black dot) is compared with the strained lattice (red dot). The lower plots present the influence of these distortions on the reflectivity curve: (a) Random disorder decreases the peak reflectivity of the Bragg reflection; (b) Homogeneous linear strains shift the reflection; (c) Longitudinal nonlinear strains create sidebands and decrease the peak value of the Bragg reflectivity.

FIG. 2. Three schematic presentations of lattice distortion. In each case, the strained crystal (red dot) is compared with undistorted lattice (black dot). The lower panel shows the impact of the distortion on the reflectivity curve of Bragg reflection. (a) Random disorder decreases the Bragg peaks. (b) Homogeneous deformations lead to spectral shifts depending on the relative change of the lattice parameter. (c) Longitudinal nonlinear strains create sidebands and decrease the Bragg peaks.

When a high-brilliance x-ray pulse is reflected by a Bragg-reflecting crystal, a certain part of the energy is deposited. Different positions and deposited energies would imprint different distortions into the crystal. The type (a) distortion occurs in a case of very high energy density that may impose radiation damage to the crystal. In this case, the crystal may totally lose the long-range order with a set of irreversible changes. In the case of far below the damage threshold, the absorbed energies due to the light-matter interactions translate into the other disorders, type (b) and (c), which are the major issues in the XFELO. When the considered temperature map is complex, the corresponding reflectivity curves can be obtained by solving the multidimensional Takagi–Taupin equations with the strain profile [\[28\].](#page-9-11) If the temperature change is homogeneous, one mainly considers the variation of the lattice constant.

In principle, the relative change of the lattice constant can be described by the thermal expansion coefficient, β . For a cubic crystal like the diamond, the relative change owing to the thermal expansion can be expressed as:

$$
\delta_a = \exp\left(\int_{T_0}^T \beta(T')T'd\right) - 1.
$$
 (2)

While the temperature change is small, β can be assumed constant and relative change of lattice constant can be approximated to $\frac{a(T)-a(T_0)}{a(T_0)} \approx \beta \times (T-T_0)$ with the first order Taylor series. The thermal expansion coefficient $\beta(T)$ can be obtained by an empirical formula from Ref. [\[27\],](#page-9-10) which can be written as $\beta = 4.25 \times 10^{-14} T^3$.

As Eq. [\(2\)](#page-2-2) has shown, a small expansion coefficient is expected to increase the tolerance of the heat load on the crystal. Therefore, it is preferable to use the crystals mirrors at a low temperature, such as 100 K, where the thermal expansion coefficient drops by a few orders of magnitude in comparison with the room temperature value of about 1×10^{-6} . Additionally, the thermal conductivity of diamond would reach the maximum value in the temperature range from 60 K to 100 K [\[29\]](#page-9-12).

III. LIGHT-MATTER INTERACTIONS AND HEAT CONDUCTIONS

In order to investigate the lattice distortions of mirrors, simulating abilities are necessary for obtaining accurate and reliable information of the absorbed energies in the crystals. This raises an essential problem with the interactions between the XFELO radiation pulses and the crystal mirrors. In this section, with the help of a dedicated Bragg reflection model, the light-matter interactions have been numerically simulated within GEANT4 toolkit [\[30\]](#page-9-13), which is developed and actively used for particle experiments and detector designs. GEANT4 is a Monte Carlo based toolkit for simulating the particle interacting with matter. It includes many functionalities like tracking, geometry, physics models, and hits. For convenience and practicality, it already implements many widely used physics models to describe the fundamental light-matter interactions.

Based on the low energy Livermore model in GEANT4 [\[31\]](#page-9-14), an additional Bragg reflection process has been implemented to extend the available range of particle interactions. The implemented Bragg reflection model could redirect photons in the crystal under the Bragg condition. Thus, this model naturally enjoys the feature of the multiple reflecting, which is the most important consideration for Bragg reflection and is absent from the previous exponential attenuation model. As the Bragg reflection occurs in a range of ten meV level, the table that stores the cross section of each interaction in the GEANT4 kernel must be extended to reach the meV level. The cross section of the Bragg reflection, the probabilities considered in the GEANT4 simulation, is calculated via the dynamical theory of x-ray diffraction. The cross section is given by

$$
\sigma_{\rm B}(\text{b}/\text{atom}) = \mu/\rho \, (\text{cm}^2 \text{ g}^{-1}) \{m_{\rm u}(\text{g})A\} \times 10^{24}, \quad (3)
$$

where $1/\mu$ is the photon-energy dependent excitation length calculated by the dynamical theory [\[9,32\]](#page-9-0), m_u = $1.66053886 \times 10^{-24}$ g is the atomic mass unit, and A is the relative atomic mass of the target element. In addition, a parameter defined as the probability that a reflected x-ray will be reflected again can be used to control the multiple Bragg reflections and to correct the reflections in thicker crystals. Overall, in the view of implemented and defined components, all aspects of the Bragg reflection simulation have been involved: (a) the geometry of the diamond crystal; (b) the primary particle (photons); (c) the generation of particles; (d) the tracking of particles through materials; (e) the physics processes modeling particle

FIG. 3. A snapshot of the light-matter interactions in the GEANT4 simulation. The green line is the x-rays. The white box is the frame of the crystal. A large part of incident x-rays is reflected, with crystal thickness of 70 μ m, photon energy of 14.33 keVand Bragg angle of about 83°. The incident direction of the x-rays points to the negative value of the z-axis.

interactions (photoelectric effect, Compton scattering, pair production, bremsstrahlung, and Bragg reflection.).

Figure [3](#page-3-0) shows a snapshot of the particle tracking in the crystals related to (3 3 7) diffracting planes, while the photon energy is about 14.33 keV with an FWHM bandwidth of 3 meV. As it is shown, the Bragg reflections govern the interaction between the x-rays with the crystal. The absorptions of the x-rays mainly result from the photoelectric effect and Compton scattering. According to the statistical data from the GEANT4 simulation, the diamond mirror with a thickness of 70 μ m allows reflecting 87% of the incident x-ray while the transported part remains 11.7%. In other words, only 1.3% of the incident x-ray has been absorbed. Since the dynamical theory of x-ray diffraction predicts the peak reflectance to be 88% and transmittance to be 11.7%, the GEANT4 simulation results are reasonable and acceptable.

Another benchmark about the evolution of the transverse intensity profile is shown in Fig. [4](#page-3-1). The GEANT4 results are compared to those of solving two-dimensional Takagi– Taupin equations. The reflected intensity profiles of the two cases have a slight difference, while the incident pulses have the same profile. These results show that the dedicated model including multiple reflections in GEANT4 can simulate the x-ray crystal interactions with reasonable accuracy.

To obtain the absorption data used for thermal analysis, the GEANT4 simulation utilizes a set of baseline parameters based on the previous studies of XFELO operation for SHINE (Shanghai high repetition rate XFEL and extreme light facility), the first hard x-ray free electron laser in China [\[33](#page-9-15)–35]. The cavity is formed by four diamond mirrors with $(3 \ 3 \ 7)$ diffracting planes while the cavity configuration is optimized to amplify the x-ray pulse with a photon energy of 14.33 keV. The intracavity pulse energy is assumed to be 600 μ J with a 1 MHz repetition rate and 3 meV FWHM spectral width. The temporal distribution of x-ray photons is assumed to a Gaussian distribution with an RMS length of 12 μ m. Meanwhile, the pulse has a

FIG. 4. A benchmark of the GEANT4 (G4) reflection process. The G4 results are compared to those of solving two-dimensional Takagi–Taupin equations (TTE). The reflected intensity profiles of the two cases have a reasonable difference, while the incident pulses have the same profile. The G4 simulation and the TTE calculation use an RMS transverse size of $25 \mu m$.

Gaussian transverse profile with an RMS size of $25 \mu m$, while the divergence angle is assumed to be 0.7 μ rad, which is much smaller than Bragg reflection widths. In addition, the crystal size is 1000 μ m × 1000 μ m × 70 μ m in the GEANT4 simulation.

Considering the pulse energy of $600 \mu J$, the total absorbed energy is about 7.3 μ J in the diamond crystal during the reflection. The transverse distribution of the absorbed energy is mainly determined by the transverse profile of the incident x-ray pulse, seen in Fig. [5.](#page-3-2) With an RMS spot size of about 25 μ m, the maximum energy density reaches $2400 \mu J/mm^2$ corresponding to timeaveraged power of 2.4 kW/mm^2 at 1 MHz repetition rate.

FIG. 5. The density of absorbed energy over the transverse plane. The corresponding marginal distribution is also shown. The maximum energy density reaches 2400 μ J/mm².

Since the crystal rotates about 7° along the x-direction to fulfill the Bragg condition, the irradiation spot size in Fig. [5](#page-3-2) on y-direction is a little larger than that of x-direction. This is due to the transverse shift as x-ray penetrates the crystal at a non-normal angle of incidence. When the Bragg angle is small, this effect would significantly affect the transverse distribution of absorbed energy. It is difficult to evaluate the multiple reflections in theoretical formulas, but simple in this GEANT4 model.

The distributions of the absorbed energy on x-z plane (top) and y-z plane (bottom) are shown in Fig. [6.](#page-4-0) The tilted penetration is presented in the bottom plot of Fig. [6](#page-4-0). Due to the short excitation length associated with a high reflectance, the incident x-ray pulse is significantly attenuated (reflected). The exponential decay model predicts an almost uniform trend when the thickness is only 70 μ m. In contrast, this model shows that the absorption is concentrated in the part of the crystal near the incident plane due to the high reflectivity. For thinner crystals, there are few multiple reflections of x-rays in the crystal, and the absorption is very similar to that predicted by the exponential decay model. Since this model can give more details about the absorbed energy, it could be a powerful tool to study the thermal load of XFELO.

Another advantage is that this scheme allows the study of dynamic absorption, the absorption energy at different times, seen in Fig. [7](#page-4-1). In this case, the time at zero is the

FIG. 6. The distribution of absorbed energy in x-z plane (top) and y-z plane (bottom).

FIG. 7. Absorbed energy density along the thickness at different times.

time when the first photon enters the crystal. At 500 fs, most of the interaction between the x-ray and the crystal has been completed. The duration of the light-matter interaction is determined by the temporal duration of the x-ray pulse and the thickness of the crystal. In addition, with the help of molecular dynamics simulation software, it is possible to study the possible dynamic changes of the crystal during the pulse duration. But, this is not our intent here.

Since the XFELO pulses may be too short on the timescale to affect their own Bragg reflection [\[36,37\]](#page-9-16), one can assume that the crystal temperature variation only acts on the subsequent pulses. Therefore, it can be assumed that the x-ray pulses induce instantaneous heating in the crystal. In other words, the energy deposition resulting from the light-matter interactions is converted into a transient temperature increment.

To study the thermal behavior of the diamond crystal, the commercial finite element analysis software ANSYS was used [\[25\]](#page-9-8). The initial temperatures were obtained from the absorption information from the GEANT4 simulation. In the ANSYS transient thermal analysis, the crystal size is 800 μ m × 800 μ m × 70 μ m, which is large enough to cover the irradiation. The incident and transmitted surface are insulated, while the other surrounding surfaces are fixed at 70 K. The work in Ref. [\[29,38,39\]](#page-9-12) is adopted to obtain the thermal conductivity and specific heat capability. As it is various as a function of the temperature, a fitting linear polynomial formula is utilized for simplicity. The mesh size is optimized, and the time step is fixed to be 0.5 ns.

The thermal conduction calculations for the single pulse input are shown in Figs. [8](#page-5-0) and [9.](#page-5-1) Figure [8](#page-5-0) shows temperature snapshots at 1, 3, 5, 10, 250, and 1000 ns after irradiation. Figure [9](#page-5-1) shows the evolution of temperature changes $(T - 70 \text{ K})$ at the center of the incident $(z = 35 \mu m)$ surfaces and transmitted $(z = -35 \mu m)$ surfaces, while the temperature change averaged over the entire crystal volume is presented by a green dash line.

FIG. 8. Snapshots of temperature field calculated with a single-pulse input. It can be seen that the early temperature is much higher on the incident surface, and the highest temperature drops rapidly to 95 K within 20 ns. At 20 ns, it can be found that the temperature along the thickness becomes homogeneous. The temperature is nearly uniform throughout the entire crystal at 250 ns and drops to about 71.8 K at 1000 ns.

As shown in Fig. [8](#page-5-0), the initial temperature of the incident surface is higher than that of the backside due to the Bragg reflection. The highest temperature near the center of the incident surface is about 1300 K. It decreases rapidly in the first 50 ns and gradually slows down in the following 100 ns. This is due to that the thermal conductivity increases rapidly during the first 50 ns thermal diffusion, and then the thermal diffusion slows down as the temperature gradient gradually decreases. As suggested by the almost identical temperatures of the incident surface and the backside in Fig. [8](#page-5-0) and Fig. [9](#page-5-1), the thermal relaxation time across the whole thickness is considered to be about

FIG. 9. The temperature (minus 70 K) at the center of the incident surface $(z = 35 \mu m)$ and transmitted surface $(z = -35 \mu m)$ is shown. The temperature averaged over the whole crystal is presented by a green dash line.

20 ns. Meanwhile, thermal relaxation time in the radial direction exceeds 150 ns. Indeed, the temperature is nearly homogeneous on the whole crystal after 200 ns. The maximum temperature difference of the whole crystal is about 0.1 K at 250 ns and decreases to 0.05 K at 500 ns. The mean temperature increase is about 4 K, with a total absorbed energy of 7.3 μ J. In the following 1000 ns, the mean temperature change decreases to about 1.8 K, which corresponds to a relative change in lattice parameters of 1×10^{-8} . At 2000 ns, the mean temperature change drops to 0.6 K. It still has a downward trend, but will decrease over a longer period of time.

The multipulse results are shown in Fig. [10](#page-6-1). The power of each laser pulse is added periodically. The time step is fixed at 1 ns, while the other parameters are the same as above. Simulations were performed with 20 μ s, 20 pulses at 1 MHz. The temperature rises instantaneously after each pulse. Its evolution in the interpulse period was similar to the single-pulse case described above. However, the average temperature keeps increasing because some residual heat remains in the crystal. Thus, the temperature increment consists of two components: (1) the heating due to x-ray irradiation and (2) the long-term accumulation of residual heat in the crystal. In our case, the temperature increment due to the irradiation is about 4 K, while the accumulated temperature increment is about 3.5 K. The accumulation of heat places a practical requirement on the cooling system to ensure that the accumulated temperature is trivial. The solution should be established in the specific context of XFELO operation, and is not our intent here. The specific cooling system design can learn experiences from the monochromator design on the synchrotron light source.

FIG. 10. Time histories of the averaged temperature (minus 70 K). The repetition rate is 1 MHz, and the periodic input laser pulse is the same as the single-pulse one.

IV. THERMAL LOADING COUPLED XFELO SIMULATION

The thermal load affects the Bragg reflection mainly through the deformation of the lattice. As mentioned in Sec. [II](#page-1-0), the pulse energy determines that the type (a) distortion is negligible in XFELO. Furthermore, according to the transient thermal simulation above, the type (c) distortion that requires a temperature gradient across the thickness is also negligible. This is because of that the thermal relaxation time across the thickness at low temperatures is very short compared to the interpulse time. Thus, when considering the long-term accumulation, the temperature gradient over the entire thickness is negligible. Therefore, it is reasonable to assume that the most critical issue with the XFELO thermal load is the shift of the reflectivity curve due to the uniform expansion of the crystal.

Therefore, the simplified model could focus on the temperature change averaged over the effective volume, which is $8\sigma_r \times 8\sigma_r \times d$ where σ_r is the RMS photon beam size, *d* is crystal thickness. This assumption is based on the fact that the thermal conductivity of the diamond is very high and that the beam size is much smaller than the crystal size. Then, a simplified model of the averaged temperature can be expressed as [\[40\]:](#page-9-17)

$$
\frac{d\Delta T}{dt} = -\alpha \times \Delta T,\tag{4}
$$

where $\Delta T = T - T_0$ is the temperature change averaged over the effective volume, α is the temperature spreading rate associated with thermal conductivity, specific heat, and crystal size. When the temperature does not change significantly, the value of α can be assumed to be constant. Then, for the single-pulse input, the solution can be written as

$$
\Delta T = \Delta T_p e^{-\alpha t},\qquad(5)
$$

where ΔT_p is the single-pulse heating temperature. With a stable repetition rate, the time for cooling is constant. Thus, we can define $\eta = e^{-\alpha \Delta t}$ as the cooling efficiency, with repetition rate of $1/\Delta t$. *η* describes residual temperature after a period of Δt . Its approximation value can be obtained from the transient thermal analysis of single-pulse input. In the multiple-pulse case, the averaged temperature change can be written as:

$$
\Delta T_{n+1} = (\Delta T_n + \Delta T_p)\eta(\eta < 1),\tag{6}
$$

where n indicates nth pulse. At steady state, the averaged temperature change become stable and we can get a solution, $\Delta T_{n+1} = \Delta T_n = \eta \Delta T_p/(1 - \eta)$. Then, the shift of reflectivity can be written as $\Delta E \approx E_H \beta \Delta T_n$, where E_H is the Bragg energy.

With the help of the simplified thermal loading model, a thermal loading coupled XFELO is performed based on the parameters of SHINE. The main parameters of SHINE are listed in Table [I](#page-6-2). SHINE equips a superconducting linac that can deliver electron bunches with energy of 8 GeV and 100 pC total charge compressed to a peak current of 700 A.

The optical cavity is built from four diamond $(3, 3, 7)$ crystal mirrors, seen in Fig. [11.](#page-7-0) With a thickness of 70 μ m, the downstream diamond mirror C_4 at 14.33 keV has 87% reflectivity (coupling output is 12%). The other three mirrors are expected to reach the total reflection with thick crystals. Since the three crystals do not need to consider the transmission of the output coupling, the temperature is appropriate to be fixed by 70 K. In the simulation, the heat load effect is mainly considered in the downstream mirror C_4 for simplicity.

The simulations are conducted by using the combination of a time-dependent FEL code GENESIS [\[41\],](#page-10-0) a field propagation simulation code OPC [\[42\]](#page-10-1), and a Bragg reflection simulation code BRIGHT [\[43\].](#page-10-2) The simplified thermal loading

TABLE I. The main parameters of XFELO operation for SHINE.

Parameter	Value
Beam energy	8 GeV
Relative energy spread	0.01%
Normalized emittance	$0.4~\text{mm}\cdot\text{mrad}$
Peak current	700 A
Charge	100 pC
Undulator period length	16 mm
Undulator segment length	4 _m
Photon energy	14.33 keV
Mirror material	Diamond
Peak reflectivity	87%
Darwin width	11 meV
Coupling-out mirror thickness	$70 \mu m$

FIG. 11. A schematic XFELO configuration for the SHINE. Four diamond mirrors are used. The photon energy is 14.33 keV. C_1, C_2 , and C_3 are thick crystal to reach a total reflection. The x-ray coupled out from the downstream mirror C_4 which peak reflectivity reaches 87%.

model is integrated into the BRIGHT as the part of the Bragg reflection. In the following analysis, the x-ray pulse is indicated as the intra-cavity one to present how much energy is imposed to the mirrors. The output value needs to be multiplied by an output-coupling factor, the transmittance.

The simulations were conducted while η is 0.1, 0.5, 0.75, and 0.95. The spectrum and power profile are presented in Fig. [12](#page-7-1) and Fig. [13](#page-8-3), respectively. When $\eta = 0.95$, a shift in the reflectivity of C_4 due to the heat load can be found in the top plot of Fig. [12.](#page-7-1) Additionally, the Bragg reflectivity of C_1 is shown by an orange dash line. As expected, in the spectrum, the FEL pulse is trapped in the overlap between the two reflectivity curves. Therefore, the total integral reflectance changes dramatically when the shift in the position of the Bragg reflectivity exceeds half of the Darwin width. This effect significantly influences the stability of the optical cavity and the gain process of XFELOs. Thus, the peak power decreases significantly as η rises, as shown in the bottom plot of Fig. [12.](#page-7-1) Another fact should be noted is that the different reflectivities also yields a slight difference in time delay, which results from the penetration during the Bragg reflection. The mismatch of time delay leads to the imperfect temporal overlap between the x-ray pulses and the electron beam, which can decrease the peak power further. It may be a crucial issue for short electron bunches.

Figure [13](#page-8-3) shows the detailed simulation results, including the spectrum evolution and pulse energy growth when η is 0.1, 0.5, 0.75, and 0.95. The left panel of Fig. [13](#page-8-3) presents the energy evolutions (red line) and the corresponding temperature change (green line). With a small η (0.1), the temperature change is trivial, and the system is stable. The pulse energy reaches about 1 mJ at saturation, while the corresponding temperature change is only 0.8 K. At the steady state, pulse energy has a very small oscillation, which is an XFELO nature. When η is set to 0.5, the temperature change increase to 10 K, while the saturated pulse energy decrease to about 900 μ J. The decrement of the pulse energy is mainly due to the decrement of total integral reflectivity.

FIG. 12. The spectrum of XFELO pulse is shown in the top plot where η is 0.95. The Bragg reflectivity as a function of photon energy for downstream mirror C_4 (green) and upstream mirror C_1 (orange) is presented by dash lines. Shifts in the spectrum and reflectivity can be found. The power profile for each η is shown in the bottom plot. The maximum peak power exceeds 3 GW.

When a large amount of heat remains in the crystal, the pulse energy has large oscillations. The source of the oscillations is the negative feedback of the temperature change on the pulse energy. Since negative feedback generally promotes stability, the XFELO may reach a steady state with the negative feedback of temperature change, such as η is 0.75. However, the steady-state pulse energy is about 350 μ J, only a third of the value of $\eta = 0.1$. In addition, the temperature change always lags behind the pulse energy variation. The combined result is that the temperature peak gradually approaches the trough of the pulse energy in time. Thus, when $\eta = 0.95$ corresponding to a poor cooling capability, the system cannot reach steady state.

The evolution of the spectrums is presented in the right panel of Fig. [13.](#page-8-3) The simulations share the same initial random number. Thus, each η has a similar period for building up the longitudinal coherence of the x-ray pulse. As energy increases, the spectral width decreases further. The final FWHM spectral width is nearly 13 meV when η is 0.1 or 0.5. The value decreases to about 9 mev at $\eta = 0.75$ due to the narrower overlap in reflectivity. The maximum

FIG. 13. The evolution of the pulse energy (red line) and the averaged temperature (minus 70 K, green line) at various η : 0.1, 0.5, 0.75, and 0.95. The evolution of the spectrum at various η : 0.1, 0.5, 0.75, and 0.95. In each round trip, the spectrum is normalized by its maximum value.

available shift is about 6 meV, which is nearly half of the Darwin width. When the reflectivity shift exceeds this value, the spectrum oscillates along with the pulse energy.

V. CONCLUSION

In this paper, we have performed a numerical model to study the heat load of the mirrors in the XFELO operation, including the fundamental light-material interactions, the thermal behavior analysis, and the corresponding heat-load coupled XFELO simulations. The light-material interactions were simulated by means of the particle tracking package GEANT4 with the dedicated Bragg reflection physical process. This dedicated model can give accurate information of x-ray absorptions in the Bragg-reflecting crystals. Following transient thermal behavior, including single pulse and multiple pulse inputs, was analyzed by finite element analysis software based on the energy absorption information extracted from the GEANT4 simulation. It is shown that, for a typical XFELO pulse

depositing about ten microjoules energy the over a spot of tens of micrometers in radius, the thermal relaxation time across the thickness is on tens of nanoseconds scale. In this situation, a simplified heat-load model is then developed to integrate the heat load in the XFELO. With the simplified model, the potential impact of the thermal load on the XFELO operation is estimated. With a large residual fraction of the temperature, the pulse energy and peak power were found to decrease markedly due to the reflectivity shifting and time delay mismatching. The simulations also indicate that the pulse energy and spectrum may have large oscillations due to negative feedback of the temperature change on the pulse energy.

In addition to this work, the analysis of non-Fourier heat conductions as well as the thermal vibrations will be performed to better understand the various thermal load effects. Furthermore, the mirror cooling system for the XFELO operation of SHINE will be designed and optimized in the future.

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