Singularities of Floquet scattering and tunneling

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We study quasibound states and scattering with short-range potentials in three dimensions, subject to an axial periodic driving. We find that poles of the scattering S matrix can cross the real energy axis as a function of the drive amplitude, making the S matrix nonanalytic at a singular point. For the corresponding quasibound states that can tunnel out of (or get captured within) a potential well, this results in a discontinuous jump in both the angular momentum and energy of emitted (absorbed) waves. We also analyze elastic and inelastic scattering of slow particles in the time-dependent potential. For a drive amplitude at the singular point, there is a total absorption of incoming low-energy (s wave) particles and their conversion to high-energy outgoing (mostly p) waves. We examine the relation of such Floquet singularities, lacking in an effective time-independent approximation, with well-known "spectral singularities" (or "exceptional points"). These results are based on an analytic approach for obtaining eigensolutions of time-dependent periodic Hamiltonians with mixed cylindrical and spherical symmetry, and apply broadly to particles interacting via power-law forces and subject to periodic fields, e.g., co-trapped ions and atoms.

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I. INTRODUCTION AND MAIN RESULTS

The main object of this paper is a time-dependent Schrödinger equation in three dimensions, that can be brought to the form

$$i\dot{\phi}(\vec{r},t) = \left[-\frac{1}{2}\nabla^2 + V_{\rm in}(\vec{r},t) + V_{\rm out}(\vec{r},t)\right]\phi(\vec{r},t),$$
 (1)

where each potential term is dominant in a different spatial region, and both are π -periodic in time (in rescaled units in which the fundamental angular frequency is 2, and $\hbar = m = 1$). We present an approach for obtaining approximate quasiperiodic, Floquet eigenfunctions of Eq. (1), starting with explicitly known families of solutions for each separate Schrödinger equation with V_{in} or V_{out} , one possibly being time-independent as a particular case. This method allows us to explore a regime of parameters inaccessible to perturbation methods.

In particular we study solutions to a problem that can be formulated in two equivalent ways; one is given by the equation

$$i\dot{\psi}(\vec{R},t) = \left\{-\frac{1}{2}\nabla^2 + V_{\text{int}}[|\vec{R} - \vec{R}^{\pi}(t)|]\right\}\psi(\vec{R},t), \quad (2)$$

where V_{int} is a spherically symmetric interaction potential and $\vec{R}^{\pi}(t)$ is a prescribed π -periodic trajectory of the center of force. To obtain a form compatible with Eq. (1), we apply the (unitary) change of coordinates

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$$\vec{r} = \vec{R} - \vec{R}^{\pi}(t), \quad \partial_t \to -\vec{R}^{\pi}(t) \cdot \vec{\nabla} + \partial_t,$$
 (3)

and then a second unitary transformation

$$\psi = \exp\{i R^{\pi}(t) \cdot \vec{r}\}\phi, \qquad (4)$$

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reducing Eq. (2) to the sum of a time-independent central force and an additional π -periodic linear force,

$$i\dot{\phi}(\vec{r},t) = \left[-\frac{1}{2}\nabla^2 + V_{\text{int}}(r) - \vec{F}(t)\cdot\vec{r} + V_F(t)\right]\phi(\vec{r},t), \quad (5)$$

where $r \equiv |\vec{r}|$ and

$$\vec{F}(t) = -\vec{R}^{\pi}(t), \quad V_F(t) = -\frac{1}{2}\vec{R}^{\pi}(t)^2.$$
 (6)

Here, the choice of what constitutes V_{in} and V_{out} depends on the approximation that is required in order to obtain the solution. The term $V_F(t)$ comes from the change of frame starting from Eq. (2), and if the starting point is Eq. (5), it will be absent. Both of these aspects will be further discussed in Sec. III, and here we keep the discussion general.

If we consider $-\vec{F}(t)$ to be a monochromatic electric field amplitude, and $V_{int}(r)$ the Coulomb potential for an electron with coordinate \vec{r} , Eq. (5) with $V_F = 0$ describes the wellstudied problem of an atom in an ac field (the ac Stark effect), written in the length gauge within the dipole approximation. Then the bound states of V_{int} are known to turn into resonances. These are quasibound states with a finite lifetime determined by the imaginary part of their complex energy. This happens generally under the effect of a periodic perturbation, for any Hamiltonian with a continuous spectrum of scattering states [1-3]. The reason is that the periodic perturbation makes every bound state with energy $(-|\epsilon|)$ resonant with unbound states from the continuum, under absorption of at least N quanta from the external drive (whose frequency is 2), where

$$(-|\epsilon|) + 2N > 0, \tag{7}$$

and 2N gives the exponent of the power-law dependence of the resonance width on the perturbation amplitude.

Studies of nonperturbative violations of this picture go back to Keldysh theory [4] and the extensive intense-laser

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literature [5]. The limit where the frequency and intensity of the oscillating field are much higher than the atomic potential can be studied by using an effective averaged potential, known as the Kramers-Hanneberger (KH) approximation. This approach has led to the prediction of the remarkable phenomenon of stabilization of the atom against ionization [6-8], with renewed interest in recent years following experimental results [9–11] and theoretical investigations [12,13]. Other recent works have also revisited the systematic expansion of an effective timeindependent Hamiltonian in the high-frequency limit in a general setup [14,15], and effects related to the potential's initial phase [16]. For intermediate laser intensities and frequencies, the problem is inherently difficult and most approaches are based on numerical integration in some form, e.g., using close-coupled equations [17–20], or Floquet *R*-matrix theory, dividing space into two regions and connecting the numerically integrated solutions at the boundary [21,22]. There is renewed interest in the modeling and measuring of ac Stark shifts of trapped atoms [23], in calculating and directly probing the angular distribution of photoelectron spectra [24,25], and in the momentum distribution of emitted electrons [26-31], where cusps in the transverse momentum distribution curves were attributed to the long-range nature of the Coulomb interaction.

In contrast, in this work we focus on the case of a short-range potential V_{int} , for which Floquet resonances are much less studied. At the same time, the singularities that result from the periodic driving can be clearly identified, avoiding the additional complexity related to the Coulomb force and the accumulation of bound states near the threshold. Indeed some detailed studies of resonances in periodically driven problems were restricted to simpler one-dimensional (1D) models [32– 42], and include the appearance and annihilation of bound states in the dressed potential, resonant coupling between internal levels, and coherent destruction of tunneling.

In order to study a truly 3D geometry [43,44], in Sec. II we present the main tool employed in this work, an expansion for problems with mixed cylindrical and spherical symmetry, as in Eq. (5). In Sec. III we apply this expansion to V_{int} that is a spherically symmetric square-well potential [Eq. (54)], with an additional axial force, directed along the *z* axis, harmonic in time,

$$\dot{F} = F_2 \cos(2t)\hat{z}.$$
(8)

The study of the time-independent square-well potential [45] constitutes one of the few examples of a detailed enumeration of the poles of the *S* matrix, and their evolution in complex momentum space as a function of the universal parameter of the problem. The *S* matrix is the operator that relates incoming spherical waves to outgoing, regarded as a function of complex energy or momentum [46], and whose poles give the bound or quasibound states of the system, discussed more in Sec. II D. A systematic study of the evolution of resonances subject to a periodic drive would allow a deeper understanding of the nonperturbative regime up to the high-frequency stabilization limit, and we take here the first step in this direction.

Figure 1 depicts the scenario standing at the center of this work [with the specific parameters given in Eq. (60) in Sec. III C]. By the Bloch-Floquet theorem, solutions of Eq. (1) can always be written as a superposition of quasiperiodic wave



FIG. 1. The complex quasienergy ω (in nondimensional units) of two quasibound solutions to Eq. (2) with a spherically symmetric square-well potential and an axial periodic drive (see text for details), continued with small increments of the drive amplitude F_2 from a near-threshold s-wave bound state. The two complex-conjugate (time-reversed) resonances are poles of the S matrix that correspond for Im $\omega < 0$ (Im $\omega > 0$) to an escape (capture) process. Perturbatively in F_2 , the escape pole is a superposition of bound components (channels of energy $\omega + 2j$ with $j \leq 0$), and outgoing waves $[j \geq 1,$ predominantly a p wave with energy $(2 + \omega) \gg \omega$]. At the crossing of $Im\omega = 0$ the two poles lie on different edges of the branch cut of the energy plane, the scattering amplitude becomes nonanalytic at a real energy, and the solution characteristics change abruptly. This is partly similar to a "spectral singularity" (or an "exceptional point"); however unitarity is obeyed and such a solution cannot be obtained from a time-independent effective potential. After crossing, the pole coming from above is now the emitting solution with $Im\omega < 0$, but radiates only in j = 0 channels (mostly *s* wave of energy ω); i.e., the radiation is emitted by tunneling without any quanta being absorbed from the drive. In a scattering experiment, the singular point corresponds to total absorption of s waves at the critical energy and their conversion into (mostly) higher energy p waves.

functions of the Floquet form

$$\phi(\vec{r},t) = e^{-i\omega t} \phi^{\pi}(\vec{r},t), \tag{9}$$

where $\phi^{\pi}(\vec{r},t)$ is π -periodic. Wave functions of the form of Eq. (9) constitute the equivalent of the eigenfunctions of a time-independent Hamiltonian, being characterized by a single frequency ω , the (quasi)energy. Hence, in Fig. 1, two exact quasiperiodic solutions are followed in the complex ω plane by continuation as F_2 is increased. These solutions give poles of the *S* matrix, as defined by the boundary conditions. For $F_2 = 0$ both coincide at an *s*-wave bound state of the time-independent square well, and thus lie initially within the physical sheet of complex energy. Since we consider a time-reversal-invariant Hamiltonian, the two poles are related by complex conjugation in the ω plane.

A quasibound state with a general complex ω is a coherent superposition of components bound to the potential well, and components which are asymptotically (for $r \to \infty$) traveling waves (going inwards or outwards). As we discuss in Sec. II C, since Eq. (2) becomes in this limit the equation of a free particle, the solutions (in that frame) tend to a superposition of free spherical waves [in the frame of Eq. (5) the waves remain periodically driven]. According to Eq. (9) each such component must have an energy equal to $\omega + 2j$ with $j \in \mathbb{Z}$. The form of a wave function of a complex (quasi)energy can be understood in the limit of $F_2 \ll 1$ using perturbation theory as mentioned above. For the pole with $\text{Im}\omega < 0$ the probability of measuring the state in one of the localized components decreases with time and hence there must be a corresponding escaping probability flux. The open channels (all partial waves with energy $\text{Re}\omega + 2j > 0$) are therefore outgoing waves, diverging at $r \to \infty$. The corresponding momentum k_{2j} , defined by

$$\frac{1}{2}k_{2j}^2 = \omega + 2j,$$
 (10)

must have $\operatorname{Re} k_{2j} > 0$ and $\operatorname{Im} k_{2j} < 0$ for $j \ge 1$ at $F_2 \to 0$. For the time-reversed solution pole, ω goes into the upper half plane as F_2 is increased from 0, the probability increases with time, and the open channels must correspond to (diverging) incoming waves [with the root of Eq. (10) chosen to have $\operatorname{Re} k_{2j} < 0$ and $\operatorname{Im} k_{2j} < 0$]; this is a capture process. The closed (bound) channels are exponentially decaying in both cases with $\operatorname{Im} k_{2j} > 0$. This choice of boundary conditions is then held fixed and defines the solutions that are continuously followed as F_2 varies. Each channel is followed separately and continuously; each k_{2j} moves on its own Riemann surface [47,48] as the parameter of the continuation is varied.

In contrast to free waves with real momentum, which can be obtained as the limit of square-integrable eigenfunctions of the free-particle Hamiltonian, the diverging (Gamow-Siegert [49]) waves with complex momentum are eigenfunctions of an (asymptotically) non-Hermitian free-particle Hamiltonian. Non-Hermitian Hamiltonians [50] very generally are the result of tracing out some degrees of freedom, and the resulting nonunitarity is a consequence of probability flux going into those degrees of freedom, which now lie outside the Hilbert space. The diverging waves carry a well-defined probability current density, and thus can be used to calculate a relative probability flux. If a quasibound state is decaying, the relative probability to detect an outgoing wave in one of the open channels will be proportional to the square of its amplitude (see Sec. IID and Fig. 6). In the time-reversed process, the bound state can form if particles are sent in towards the center, with the probability for this process again proportional to the overlap of available free-particle states with the Floquet solution, and (half) the rate of formation is given by $\text{Im}\omega$.

Alternatively, a standard scattering formulation (discussed in Sec. IIE) is obtained by setting $\phi(\vec{r},t)$ to be asymptotically the sum of a plane wave and scattered outgoing spherical waves, and restricting ω to be the real, positive energy of the plane wave. Resonances in the scattering cross section can be typically related to quasibound (resonance) states [47], and hence the nomenclature coincidence. The relation results from the scattering amplitude (and the *S* matrix) being a meromorphic function of the energy (or momentum) in some region in the complex plane (that depends on the potential), and hence its values on the real axis are significantly influenced by nearby poles. We follow both aspects and their relation in our analysis of singularities in the Floquet setup.

Indeed, as F_2 is increased, the absolute value of Im ω in Fig. 1 initially increases and then decreases; this is an example of a nonperturbative stabilization at higher field amplitudes. As the pole in the lower half plane passes to Re $\omega > 0$, the j = 0 channels open, but consist of asymptotically decaying

incoming waves (in momentum space, k_0 just crosses the bisector of its quarter plane), capturing back-reflected (mostly *s*) waves. At a critical value of the drive strength the two poles reach the real ω axis and at this singular point with a real energy, the nature of the quasibound states changes abruptly. We denote the parameters of this point by

$$F_2 = \bar{F}^c, \quad \omega = \bar{\omega}^c. \tag{11}$$

Just before crossing the real line, the solution with $\text{Im}\omega < 0$ radiates partial waves of energy $\omega + 2j$ with $j \ge 1$ after absorbing at least one quantum from the drive, with the dominant partial wave being the p wave with j = 1. After crossing, this solution becomes the capture solution (with $\text{Im}\omega > 0$) and all previously open radiating channels are now decaying $(\text{Im}k_{2i} > 0 \text{ for } i \ge 1)$ and carry no current asymptotically. This solution is capturing incoming (mostly s) waves in the j =0 channels (which have crossed to $\text{Im}k_0 < 0$, while remaining with $\operatorname{Re}k_0 < 0$). The radiating solution now, for $F_2 > \overline{F}^c$, is the solution that came from the upper (energy) half plane. All channels with $j \ge 1$ (which were incoming, diverging waves) are now asymptotically decaying; in the frame of Eq. (5) they are back-reflected by the oscillating drive [in the frame of Eq. (2) it is destructive interference of waves emitted during the oscillations]. The only open, radiating channels are with j = 0(mostly the s wave), which was previously asymptotically decaying and are now diverging. Thus the bound state radiates solely by tunneling, without absorbing quanta from the drive.

Exactly at Im $\omega = 0$ (for \overline{F}^c) there is one pole on the upper edge of the cut of the complex energy plane (this is the pole that moved through the upper half of the plane), and one pole directly below it, on the lower edge of the cut. Both are about to leave the physical sheet. In momentum space, the two corresponding k_0 poles are crossing from the upper to the lower half plane, on both sides of the imaginary axis. The two solutions have a real (degenerate) energy and the $j \ge 0$ channels are (driven) free-particle waves. The solutions describe a balanced flux of incoming and outgoing waves in the respective channels. This is a "self-sustaining" standing wave that exists with open boundary conditions. The S matrix becomes nonanalytic on the real energy axis for this critical parameter, and an effective, time-averaged Hermitian approximation of the potential (as in the KH approximation discussed above) cannot result in such a solution, as this would violate unitarity of the resulting elastic scattering.

In the full time-dependent setup however, the scattering is inelastic and unitarity is obviously not violated. In terms of the *S* matrix, paired with each pole there is a zero of the *S* matrix that in momentum space is located at $-k_{2j}$, which follows the same trajectory in the energy plane. In a scattering experiment at $F_2 = \overline{F}^c$ where the poles and zeros all coincide at $\omega = \overline{\omega}^c$, an incoming (j = 0) s wave with this value of energy is completely absorbed and removed from the scattered wave function, which becomes predominantly *p* wave with j = 1. This occurs as the scattering amplitude of the incoming *s* wave in the j = 0 channel goes through 0 at the singular point, and in a 2D parameter space composed of F_2 and the real scattering energy ω , a 2π phase is accumulated around this point (see Fig. 7).

Hence the time-dependent solution at the critical point of crossing the real ω axis shares some properties with singularities discussed mostly in the context of timeindependent complex potentials. These include in particular "spectral singularities," which occur with two scattering states with a real energy in a complex potential [51]. However with spectral singularities the manifestly complex potential violates unitarity, which is not the case in the current setup. Similarly, an "exceptional point" typically refers to the coalescence of two discrete states of a complex Hamiltonian [52], a problem that continues to be studied theoretically [53-55], with interesting recent realizations and implications [56-58]. In contrast, in the current problem, despite the coincidence of two poles at the same (real) value of ω , there is no coalescence of the eigenvectors. The coincidence of a pole and a zero in the Floquet problem is also similar to singular points of laserabsorber \mathcal{PT} -symmetric systems [59–61], which are again nonunitary. In such optical systems with gain and loss also Floquet setups attract increasing interest [62,63]. Exceptional points in a Floquet unitary scattering setup have been discussed before [64,65], and we further discuss Floquet exceptional points in Sec. IV. We note also that the solution at the critical point is not a "bound state in the continuum" [47,66], since it is not square-integrable.

Further relevant examples where the results presented here are applicable include quantum wires and dots [67] that have been modeled by similar finite-barrier potentials, and the expansion presented here can be used to solve a mixedtype system, periodically driven. Interacting cold atoms or molecules [68] are often subject to oscillating fields [69,70]. The generalization to settings with a potential of spherical symmetry in the exterior region is straightforward, and the case of zero-range interaction has been recently treated in [71]. Overlapping a trap for neutral atoms with a periodically driven Paul trap for ions [72] was suggested and realized [73,74], followed by the demonstration of a trapped ion immersed in a dilute atomic Bose-Einstein condensate [75,76], and many other experiments. The effect of the periodic drive of the ion has been analyzed for classical collisions with the atom [77] or Rydberg atoms [78], for quantum scattering employing a master equation description [79], and for an ion and atom in separate traps [80]. Quantum defect theory (QDT) [81,82] is a very important theoretical tool for modeling atomic scattering, and continues to evolve [83-92], together with new models and methods [93,94], applied to many-body states as well [95-97]. As we argue in Sec. IV, the results presented in this work hold for short-range power-law potentials, and calculations using QDT, which can naturally be used in the interior region, show that they can potentially be observed with a co-trapped ion and atom system [98].

This paper is organized as follows. Section II develops the theory. In Sec. II A we introduce the expansion that is used in Sec. II B to relate the solution matching conditions. Some general properties of the solutions are discussed in Sec. II C, and then Sec. II D discusses in more detail quasibound (pole) solutions and their characterization, while Sec. II E introduces scattering solutions and cross sections. We conclude the formalism with a review of some analytic properties of the wave functions used in the expansion, and of the partial waves expansion, in Sec. II F. The results of applying the theory to the problem of a driven square well are presented in Sec. III, with a driven loosely bound s wave studied in Sec. III A and a deeper bound p wave in Sec. III B. In Sec. III C we discuss some aspects of the method and compare approximations that can be achieved with it, concluding in Sec. IV with a discussion of the applicability of this work to more physical atomic potentials, and the relation to other singularities of non-Hermitian Hamiltonians. The appendices contain, in addition to some details of the derivation, a few general expressions useful for the calculation of expectation values using the solution wave functions.

II. FLOQUET WAVE FUNCTIONS

A. Floquet waves with cylindrical symmetry

Starting with the general time-dependent Schrödinger equation

$$i\dot{\phi}(\vec{r},t) = \left[-\frac{1}{2}\nabla^2 - \vec{F}(t)\cdot\vec{r} + V_1(t)\right]\phi(\vec{r},t), \quad (12)$$

a family of solutions can be written in the form

$$\phi(\vec{r},t) \propto e^{i\vec{q}(t)\cdot\vec{r}-ig(t)}, \quad \vec{q}(t) = \vec{F} + \vec{k},$$
 (13)

where \vec{k} is the (possibly complex) constant of integration, and

$$g(t) = \frac{1}{2}\vec{k}^2 t + \int^t \left[\vec{k}\cdot\vec{F}(t') + \frac{1}{2}\vec{F}(t')^2 + V_1(t')\right]dt'.$$
 (14)

In the rest of this paper, we assume the drive to be π -periodic and coaxial at any time, and choose a cylindrical coordinate system, $\vec{r} = (\rho, z, \varphi)$, in which

$$\vec{F}(t) = F^{\pi}(t)\hat{z}.$$
(15)

Furthermore, if Eq. (2) is the physical starting point we have using Eq. (6)

$$V_1(t) = V_F(t) = -\frac{1}{2}\vec{F}(t)^2,$$
 (16)

which also simplifies the current expressions by canceling the \vec{k} -independent term in Eq. (14). We return to this point in Sec. III C.

The cylindrical waves are eigenfunctions of the free-particle Hamiltonian, $H = -\frac{1}{2}\nabla^2$, given by

$$\chi_m^{(1,2,J)}(\vec{r};k,\alpha) = e^{im\varphi} H_m^{(1,2,J)}(k\rho\sin\alpha) e^{ikz\cos\alpha}, \qquad (17)$$

where *m* is the magnetic quantum number, *k* is the (possibly complex) wave number, and α is a complex parameter. $H_m^{(1,2,J)}$ is a Hankel function of the first or second kind (corresponding to outgoing and incoming traveling waves, respectively), or a Bessel function (which we denote with a superscript *J*). Thus outgoing and incoming traveling-wave solutions to Eq. (12), subject to Eqs. (15) and (16), can be written using Eq. (17) in the form

$$\phi_m^{(1,2)}(\vec{r},t;k,\alpha) = e^{-i\frac{1}{2}k^2t}e^{i\vec{F}^{\pi}(t)z}\chi_m^{(1,2)}(\vec{r};k,\alpha)e^{-iF^{\pi}(t)k\cos\alpha}.$$
(18)

Equation (18) is a particular quasiperiodic solution of Eq. (12), taking the Floquet form of Eq. (9). With k_{2j} defined in Eq. (10), the most general quasiperiodic solution to Eq. (12) is

$$\phi_m(\vec{r},t) = \sum_{\substack{j \in \mathbb{Z} \\ a=1,2}} \int_{C^{(a)}} d\alpha \sin \alpha b_{2j}^{(a)}(\alpha) \phi_m^{(a)}(\vec{r},t;k_{2j},\alpha), \quad (19)$$

which consists at each value of k_{2j} of a superposition of outgoing and incoming cylindrical waves, parametrized by integrals in the complex α plane along the contours $C^{(a)}$ with weight functions $b_{2j}^{(a)}(\alpha)$.

To transform the solution of Eq. (19) to spherical coordinates $\vec{r} = (r, \theta, \varphi)$, we take the arbitrary weight functions for cylindrical Floquet waves in Eq. (19) to be

$$b_{2j}^{(a)}(\alpha) = \sum_{l_1} b_{2j,l_1} N_{2j,l_1}^m S_{2j,l_1}^{(a)} P_{l_1}(\cos \alpha), \qquad (20)$$

with N_{2j,l_1}^m a normalization constant to be defined in Eq. (37), $S_{2j,l_1}^{(a)}$ will depend on the boundary conditions at $r \to \infty$, and b_{2j,l_1} will be matching coefficients for quasibound states, as elaborated in the following subsections. In Appendix A we show that by plugging Eq. (20) into Eq. (19), each term of the series within the latter equation can be rewritten as

$$\int_{C^{(a)}} d\alpha \sin \alpha b_{2j}^{(a)}(\alpha) \phi_m^{(a)}(\vec{r},t;k_{2j},\alpha)$$

= $e^{-i\frac{1}{2}k_{2j}^2 t} \sum_{l_1,l} b_{2j,l_1} N_{2j,l_1}^m S_{2j,l_1}^{(a)} R_{2j,l,l_1}^{(a)}(r,t) Y_l^m(\theta,\varphi),$ (21)

where Y_l^m are normalized spherical harmonics [Eq. (A7)] and the radial functions are

$$R_{2j,l_{1,l}}^{(a)}(r,t) = \sum_{l_{2},l_{3},l_{4}} c_{l_{1},l_{2},l_{3},l_{4},l} j_{l_{2}} [F^{\pi}(t)k_{2j}] \\ \times j_{l_{4}} [\dot{F}^{\pi}(t)r] h_{l_{3}}^{(a)}(k_{2j}r), \quad (22)$$

with the coefficients $c_{l_1,l_2,l_3,l_4,l}$ being defined in Eq. (A6), and the spherical Hankel functions of the first (second) kind, $h_l^{(1)}$ ($h_l^{(2)}$), correspond to outgoing (incoming) spherical waves

(for Rek > 0). This expansion forms the essential tool that allows one, in conjunction with the matching described in the following subsection, to obtain the results in this work.

B. Floquet wave functions with mixed cylindrical and spherical symmetry

Adding a spherically symmetric interaction potential to Eq. (12) [with V_1 of Eq. (16)], we regain Eq. (5),

$$i\dot{\phi}(\vec{r},t) = \left[-\frac{1}{2}\nabla^2 + V_{\text{int}}(r) - \vec{F}(t) \cdot \vec{r} + V_F(t)\right] \phi(\vec{r},t).$$
(23)

Depending on V_{int} , this equation may be solvable either exactly or only approximately. The solution proceeds by assuming that Eq. (23) can be replaced by an equation having the form of Eq. (1),

$$i\dot{\phi}(\vec{r},t) = \left[-\frac{1}{2}\nabla^2 + V_{\rm in}^{(1)}(\vec{r},t) + V_{\rm out}^{(1)}(\vec{r},t)\right]\phi(\vec{r},t),\qquad(24)$$

and dividing space into two regions, interior and exterior to sphere $|\vec{r}| = d$, where the Schrödinger equation can be solved exactly with either one of the potentials above. In this section we will focus on the case when an approximation is required, taking the form

$$V_{\rm in}^{(1)}(\vec{r},t) = V_{\rm int}(r)\Theta(d-r),$$

$$V_{\rm out}^{(1)}(\vec{r},t) = [-\vec{F}(t)\cdot\vec{r} + V_F(t)]\Theta(r-d), \qquad (25)$$

where $\Theta(\cdot)$ is the Heaviside function. Thus in Eq. (23) we have truncated V_{int} at a finite radius and removed the external drive from the interior region, leaving it to modulate the free-particle exterior region. A further discussion of this approximation (and a comparison to the exact solution for a square well) will follow in Sec. III C. For finding quasiperiodic solutions in the Floquet form of Eq. (9), we can employ the ansatz

$$\phi_{m}(\vec{r},t) = \begin{cases} \sum_{n,l} a_{2n,l} e^{-i(\omega+2n)t} \phi_{\mathrm{in},\omega+2n,l}(r) Y_{l}^{m}(\theta,\varphi), & r < d, \\ \sum_{j,l_{1}} b_{2j,l_{1}} e^{-i(\omega+2j)t} \sum_{l} \phi_{\mathrm{out},2j,l_{1},l}^{\pi}(r,t) Y_{l}^{m}(\theta,\varphi), & r > d, \end{cases}$$
(26)

with

$$\phi_{\text{out},2j,l_1,l}^{\pi}(r,t) = N_{2j,l_1}^{m} \Big[S_{2j,l_1}^{(2)} R_{2j,l_1,l}^{(2)}(r,t) + S_{2j,l_1}^{(1)} R_{2j,l_1,l}^{(1)}(r,t) \Big].$$
(27)

This wave function is the most general Floquet superposition of solutions in both the interior and exterior region and is an exact solution of the original problem [Eq. (23)] in both limits $r \to 0$ and $r \to \infty$. In the interior region each wave function $\phi_{in,\omega+2n,l}(r)$ is (locally) a solution with energy $\omega + 2n$ of the Schrödinger equation with the potential $V_{in}^{(1)}(\vec{r})$. The Fourier expansion (that is tractable if it can be truncated of course) takes all integers $n \in \mathbb{Z}$ (and similarly for *j*), without any *a priori* restriction to positive (real part) energies. The required boundary conditions at $r \to 0$ are assumed to have been imposed, which determine a unique linear combination of the two linearly independent solutions at each value of $\omega + 2n$. This can be the condition of regularity at the origin, or a quantum defect theory parametrization. The boundary conditions at $r \to \infty$ are discussed in the following subsections. Finally, Eq. (27) does not indicate explicitly a summation over any degeneracy of the wave functions, which might require independent matching coefficients.

Requiring the continuity of the wave function across the surface of the sphere $|\vec{r}| = d$ gives the equation (trivially independent of θ, φ due to the identical expansion in spherical harmonics on both sides)

$$\sum_{n,l} e^{-i(\omega+2n)t} a_{2n,l} \phi_{\mathrm{in},\omega+2n,l}(d)$$

$$= \sum_{j,l_1} e^{-i(\omega+2j)t} b_{2j,l_1} \sum_{l} \phi_{\mathrm{out},2j,l_1,l}^{\pi}(d,t)$$

$$= \sum_{j,l_1} e^{-i(\omega+2j)t} b_{2j,l_1} \sum_{l,p} d_{2p,l,2j,l_1} e^{-i2pt}$$

$$= \sum_{n,l} e^{-i(\omega+2n)t} \sum_{j,l_1} b_{2j,l_1} d_{2(n-j),l,2j,l_1}, \qquad (28)$$

where $d_{2p,l,2j,l_1}$ are the expansion coefficients of the Fourier series of $\phi_{\text{out},2j,l_1,l}^{\pi}(d,t)$, which in general must be obtained numerically. Then the first matching condition is

$$c_{2n,l}a_{2n,l} \equiv \phi_{\mathrm{in},\omega+2n,l}(d)a_{2n,l} = \sum_{j,l_1} d_{2(n-j),l,2j,l_1}b_{2j,l_1}.$$
 (29)

$$\partial_r R_{2j,l_1,l}^{(a)}(r,t) = \sum_{l_2,l_3,l_4} c_{l_1,l_2,l_3,l_4,l} j_{l_2} [F^{\pi}(t)k_{2j}] \\ \times \left\{ -\dot{F}^{\pi}(t)j_{l_4+1}[\dot{F}^{\pi}(t)r]h_{l_3}^{(a)}(k_{2j}r) + j_{l_4}[\dot{F}^{\pi}(t)r] \left[\frac{1}{r}(l_3+l_4)h_{l_3}^{(a)}(k_{2j}r) - k_{2j}h_{l_3+1}^{(a)}(k_{2j}r) \right] \right\}$$

In the following subsections we elaborate on the properties of the wave function, and we will arrange these recursion formulas in a matrix form, to facilitate their solution.

C. General properties of the solutions

Some general properties of the Schrödinger equation and its solutions will be used in the following. First, we assume that F(t) [or $R^{\pi}(t)$] is time-reversal invariant (including the possibility that this requires a trivial shift of t). Then the Schrödinger equation is invariant under a simultaneous change $t \rightarrow -t$ and complex conjugation, and hence if $\phi(t)$ is a solution, so is $\phi^*(-t)$, which may be the same wave function or an independent one. In addition, the equation conserves probability locally in time and space, such that the continuity equation holds,

$$\partial_t n(\vec{r},t) + \vec{\nabla} \cdot \vec{j}(\vec{r},t) = 0, \qquad (32)$$

with the density and probability current density defined by

$$n(\vec{r},t) = |\phi|^2, \quad \vec{j}(\vec{r},t) = \frac{1}{2i} [\phi^* \nabla \phi - \phi \nabla \phi^*].$$
 (33)

The continuity equation holds irrespective of the Hermiticity of the boundary conditions (or whether ω is real or complex), as long as the Hamiltonian is real. Then, if the wave function is square-integrable in configuration space, its norm remains constant (and finite) in time. However, even if the wave function is not normalizable, a meaning can be attributed to the relative probability amplitude of each asymptotic channel, as detailed in the following subsection.

In order to gain more insight into the physical meaning of the wave function of Eq. (26), we can use its connection to the Schrödinger equation in the "lab" frame, Eq. (2), which is reproduced here again,

$$i\dot{\psi}(\vec{R},t) = \left\{-\frac{1}{2}\nabla^2 + V_{\text{int}}[|\vec{R} - \vec{R}^{\pi}(t)|]\right\}\psi(\vec{R},t).$$
 (34)

In the asymptotic $\vec{R} \to \infty$ region, V_{int} decays and the solutions of Eq. (34) reduce to free-particle solutions. By starting with a free spherical wave of momentum k_{2j} (dropping $e^{-i(\omega+2j)t}$ for simplicity),

$$\psi_{k_{2j}l_{1m}}^{(a)}(\vec{R}) = N_{2j,l_{1}}^{m} h_{l_{1}}^{(a)}(k_{2j}R) P_{l_{1}}^{m}(\cos\theta) e^{im\varphi}, \quad (35)$$

effecting the unitary transformation of Eqs. (3) and (4), and then using the representation of Eq. (A1), it is seen that in fact Similarly, the second matching condition comes from the continuity of the radial derivative (∂_r) , which gives

$$f_{2n,l}a_{2n,l} \equiv \partial_r \phi_{\mathrm{in},\omega+2n,l}(d)a_{2n,l} = \sum_{j,l_1} g_{2(n-j),l,2j,l_1}b_{2j,l_1},$$
(30)

with $g_{2p,l,2j,l_1}$ the expansion coefficients of the Fourier series of $\partial_r \phi_{\text{out},2j,l_1,l}^{\pi}(d,t)$. The latter can be written as

$$^{\tau}(t)j_{l_{4}+1}[\dot{F}^{\pi}(t)r]h_{l_{3}}^{(a)}(k_{2j}r) + j_{l_{4}}[\dot{F}^{\pi}(t)r]\left[\frac{1}{r}(l_{3}+l_{4})h_{l_{3}}^{(a)}(k_{2j}r) - k_{2j}h_{l_{3}+1}^{(a)}(k_{2j}r)\right] \right\}.$$
(31)

the transformation carries

$$\psi_{k_{2j}l_{1}m}^{(a)}(\vec{R}) \to N_{2j,l_{1}}^{m} \sum_{l} R_{2j,l_{1},l}^{(a)}(r,t) Y_{l}^{m}(\theta,\varphi), \quad (36)$$

so that $b_{2j,l_1} S_{2j,l_1}^{(a)}$ in the solution Eq. (26) is the coefficient of the asymptotically free spherical wave in the "lab" frame (with R), with energy $2j + \omega$ and angular momentum quantum number l_1 . A similar conclusion can also be obtained in the frame of Eq. (23). In this frame, in any physical realization, the periodic drive $F(t) \cdot \vec{r}$ cannot continue to infinity. Then if beyond some large enough distance r outside the range of V_{int} the amplitude of periodic drive F(t) is adiabatically diminishing (in space), then the solutions will become asymptotically free spherical waves again. However, in this frame, Eq. (16) has to be modified $(V_1 = 0)$, and the asymptotic momenta will be different. We do not include this calculation explicitly although we return to this point in Sec. III C.

The wave of Eq. (35), in the nondecaying channels, carries a momentum current density proportional to $\operatorname{Re}k_{2i}$. Thus for the solutions of Eq. (2) [Eq. (34)] we can set for the normalization constant of Eq. (20) and Eq. (27),

$$N_{2j,l_1}^m = (\operatorname{Re} v_{2j})^{-1/2} (\operatorname{Re} k_{2j}) N_{l_1}^m = (\operatorname{Re} k_{2j})^{1/2} N_{l_1}^m, \quad (37)$$

with N_i^m defined in Eq. (A7) and the velocity of a particle is related to its wave number in nondimensional units simply by $\operatorname{Re} v_{2i} = \operatorname{Re} k_{2i}$. This makes the probability current density \vec{j} of Eq. (33) in one outgoing, nondecaying channel wave function [Eq. (35)], asymptotically normalized to unit flux on the sphere,

$$\vec{j} \sim \left| Y_{l_1}^m(\theta, \varphi) \right|^2 \hat{r}.$$
 (38)

This normalization will be used in the following two subsections.

D. Quasibound wave functions

Quasibound states are defined by fixing the boundary conditions in all channels, by setting each of the constants $S_{2j,l_1}^{(a)}$ in Eq. (27), to either 0 or a modulus 1 value, $|S_{2j,l_1}^{(a)}|^2 = \{0,1\}.$ The two matching relations can be written in matrix form (once a finite truncation has been applied),

$$C\vec{a} = D\vec{b}, \quad F\vec{a} = G\vec{b},$$
 (39)

where \vec{a} and \vec{b} denote the expansion coefficients $a_{2n,l}$ and b_{2i,l_1} whose indexes are "flattened" in vector notation; C, D are matrices whose elements [using Eq. (29)] are

$$(C)_{(2n,l),(2j,l_1)} = c_{2n,l}\delta_{n,j}\delta_{l,l_1},$$

$$(D)_{(2n,l),(2j,l_1)} = d_{(2(n-j),l),(2j,l_1)},$$
(40)

and similarly for F and G using Eq. (30). By writing the two equations in block form

$$K(\omega) \begin{pmatrix} \vec{a} \\ \vec{b} \end{pmatrix} \equiv \begin{pmatrix} C & -D \\ F & -G \end{pmatrix} \begin{pmatrix} \vec{a} \\ \vec{b} \end{pmatrix} = 0, \tag{41}$$

the compatibility of the two matching conditions implies the vanishing of (at least) one eigenvalue (or, more generally, singular value in the singular value decomposition) of $K(\omega)$. A (complex in general) value of ω compatible with the imposed boundary condition has to be searched, and the corresponding kernel vector then gives the expansion coefficients. In practice it is possible to work with the smaller matrix (*C* and *F* would in general be invertible)

$$FC^{-1}D\vec{b} = G\vec{b} \quad \Rightarrow (G - FC^{-1}D)\vec{b} = 0, \qquad (42)$$

whose kernel vectors give the exterior region coefficients b, from which \vec{a} immediately follows. The normalization of the wave function is discussed in Appendix C, and in Appendix B we lay down for completeness the expansion of integrals which are required in order to calculate expectation values of some general operators (we restrict the expressions to axially symmetric wave functions with m = 0).

The constants $S_{2j,l_1}^{(a)}$ determine the boundary conditions in the asymptotic region. For the partial waves with $\text{Re}\omega + 2j < 0$, setting $S_{2j,l_1}^{(2)} = 0$ and $S_{2j,l_1}^{(1)} = 1$ gives waves exponentially decaying in space (Imk > 0) that carry no flux asymptotically. These are the bound, square-integrable components of the wave function, which represents the probability density localized to the well. The boundary conditions for $\text{Re}\omega + 2j > 0$ depend on whether the problem is to be Hermitian or non-Hermitian.

In order to impose Hermitian boundary conditions with a real quasienergy ω , the terms with $\omega + 2j > 0$ must include both outgoing and incoming waves. Then we can let $S_{2j,l_1}^{(1)} = 1$ and set $S_{2j,l_1}^{(2)}$ to the relative phase of waves reflected inwards from the boundary at infinity (assuming that it depends only on the energy and the angular momentum quantum number l_1 in the asymptotic, drive-free region). This phase can be used for expansion of the wave functions of any Hamiltonian in the region that is far from the scattering center (such as an external particle trap). The solution describes a steady state with a superposition of bound components and traveling waves, incoming and outgoing.

When solving for a complex ω , the boundary condition make the problem non-Hermitian. Resonances with Im $\omega > 0$ describe a capture process by the oscillating well, with the probability exponentially increasing in time. For the components with Re $\omega + 2j > 0$ both the imaginary and real parts of k_{2j} can be chosen to be positive. Setting $S_{2j,l_1}^{(2)} = 1$ gives incoming waves whose amplitude exponentially diverges at infinity; the incoming flux of these waves is being captured by the state within the well. Setting $S_{2j,l_1}^{(1)} = 1$ gives outgoing waves whose amplitude exponentially decays at infinity. Alternatively, resonances with Im $\omega < 0$, describing an escape out of the well (exponential decay with time), would have the real part of k_{2j} (for the components with $\text{Re}\omega + 2j > 0$) necessarily negative (for the root with positive imaginary part chosen to have the asymptotic decay or divergence as above), which inverts the roles of incoming and outgoing waves; $S_{2j,l_1}^{(2)}$ are outgoing waves diverging at infinity and $S_{2j,l_1}^{(1)}$ are incoming waves decaying at infinity.

The probability density of measuring in the asymptotic region an emitted particle with (real) momentum \vec{k} is given by the squared absolute value of the probability amplitude of the corresponding free-particle wave function. Taking into account the interference of the angular harmonics with different values of l_1 at each value of k_{2j} , we can define the joint probability density in spherical momentum coordinates

$$f(k,\theta,\varphi) = \frac{1}{\mathcal{N}} \sum_{j}' \left| \sum_{l_1} b_{2j,l_1} Y_{l_1}^m(\theta,\varphi) \right|^2 \frac{\delta(k - \operatorname{Re}k_{2j})}{k^2}$$
(43)

with the summation index *j* extending over all outgoing channels that do not decay asymptotically. The normalization of Eq. (38) guarantees that each channel is weighted correctly, and \mathcal{N} sets the overall normalization. The probability to measure a particle with momentum in the volume element of momentum space between (k, θ, φ) and $(k + dk, \theta + d\theta, \varphi + d\varphi)$ is then

$$p(k,\theta,\varphi) = f(k,\theta,\varphi)k^2 \sin\theta dk d\theta d\varphi, \qquad (44)$$

with the normalization $\int p = 1$ determining the value of \mathcal{N} . We can also define the axially symmetric marginal probability density

$$f(k,\theta) = \int d\varphi f(k,\theta,\varphi), \qquad (45)$$

which will be used in the following. From $f(k,\theta)$ it is a simple change of coordinates to $f(k_{\rho},k_z)$ which is the joint distribution in terms of axial and transverse momentum, from which the marginal distributions can also be obtained.

E. Scattering wave functions

In a scattering problem the imposed boundary conditions are composed of a given form of free-particle waves in the "input" channel, with a well-defined energy value $\omega > 0$, in the asymptotic region. The given wave is interacting with the potential within its range of effect, resulting in superimposed scattered outgoing spherical waves in the asymptotic region. We will pose the boundary conditions in terms of ingoing spherical waves, which facilitates the calculation of the unitary *S* matrix.

Specifying the values of $S_{2j,l_1}^{(2)}$ in Eq. (27) and treating $S_{2j,l_1}^{(1)}$ as free parameters, the matching condition becomes the inhomogeneous linear equation

$$[G^{(1)} - FC^{-1}D^{(1)}]\vec{S}^{(1)} = -[G^{(2)} - FC^{-1}D^{(2)}]\vec{S}^{(2)}.$$
 (46)

Setting $S_{2j,l_1}^{(2)} = \delta_{j,j'}\delta_{l_1,l'_1}$ describes an incoming spherical wave with asymptotic unit flux in the channel (j',l'_1) , as the normalization of Eq. (37) makes $|\vec{j}| = |S_{2j,l_1}^{(a)}|^2$ in each channel. The resulting $S_{2j,l_1}^{(1)}$ gives the $[(j,l_1),(j',l'_1)]$ matrix-element of the unitary scattering *S* matrix that transforms an incoming wave to an outgoing. To simplify the expressions in the following, we further assume that the incoming wave is purely *s* wave in a single energy channel, i.e., $(j',l'_1) = (0,0)$. This corresponds

to the limit of scattering of slow particles, if the energy is also low enough [46]. Then the elastic cross section is

$$\sigma_{e,0}/(2\pi) = |1 - S_{0,0}|^2/(4\omega), \tag{47}$$

the inelastic cross section is given by

$$\sigma_{r,0}/(2\pi) = (1 - |S_{0,0}|^2)/(4\omega), \tag{48}$$

and the total cross section is

$$\sigma_{t,0}/(2\pi) = 2(1 - \operatorname{Re} S_{0,0})/(4\omega),$$
 (49)

where we have removed the superscript from $S_{0,0}^{(1)}$ to simplify the notation. We note that for scattering solutions, time reversal (with complex conjugation) interchanges the initial and final states and reverses the direction of wave propagation, and for a Hamiltonian which is invariant, the scattering amplitude must remain the same; this is the reciprocity theorem.

F. Analytic properties of the solutions

The asymptotic form of V_{int} determines important properties of the scattering solutions, bound states, and poles of the *S* matrix in complex energy and momentum planes. In this subsection we review a few of these properties [46,47,99], which will be used in the following sections. We will restrict the discussion to two forms for the interaction potential: either V_{int} that vanishes identically beyond a certain distance (a finiterange potential) or is asymptotically an (attractive) power-law potential $V_{int} \sim -C/r^{\alpha}$ with C > 0 and $\alpha > 3$ which is a restricted form of what is typically referred to as a short-range (or "shorter-ranged") potential [100].

If we consider the time-independent Schrödinger equation (with potential V_{int}), we can write its solutions in the form

$$\phi_{\varepsilon lm}(\vec{r}) = \frac{1}{r} u_{\varepsilon l}(r) Y_l^m(\theta, \varphi), \qquad (50)$$

where $u_{\varepsilon l}(r)$ are solutions of the reduced equation

$$\left[-\frac{1}{2}\frac{d^2}{dr^2} + \frac{l(l+1)}{2r^2} + V_{\text{int}}(r) - \varepsilon\right]u_{\varepsilon l}(r) = 0.$$
 (51)

At each (complex) value of ε , Eq. (51) has two linearly independent solutions, $u_{\varepsilon l}^{(1,2)}$. If the potential at $r \to \infty$ behaves like $-C/r^{\alpha}$ with $\alpha > 2$ then the two linearly independent solutions $u_{\varepsilon l}^{(1,2)}$ can be chosen to have an ε -independent limit at $r \to 0$ for all energies. In addition, the potential cannot have an infinity of bound states accumulating near the threshold $\varepsilon = 0$. For $\alpha > 3$, $u_{\varepsilon l}^{(1,2)}$ are entire functions [83] of the complex momentum k or of the energy (on its two-sheeted Riemann surface, with the cut extending on the positive real axis), at any fixed r. This property, which holds also for the solutions of a finite-range potential, will be alluded to in Sec. IV. When this property fails, $u_{\varepsilon l}^{(1,2)}$ will not be analytic in the entire k plane, only in parts of it, and there may exist cuts on the imaginary k axis (at negative energy).

At the same time, a potential with a tail with $\alpha > 2$ cannot keep the same power-law behavior down to the origin, because then the energy spectrum will not be bounded from below (due to an infinity of bound states at decreasing energies). Hence physically relevant short-range atomic potentials will rise near the origin, supporting a finite number of bound states, which is again true for finite-range potentials as well. The physical solution to Eq. (51) is defined to be the unique linear combination that is regular at the origin, normalized by a definite condition, e.g., $r^{-l-1}u_{\varepsilon l} \rightarrow 1$. Since this boundary condition is ε - (and k-) independent, the physical solution is an entire function of ε (k) for any fixed r. When both of the functions $u_{\varepsilon l}^{(1,2)}$ are entire, this immediately implies that the scattering matrix is meromorphic in complex energy (or momentum space) with isolated poles. Poles on the imaginary k axis correspond to bound states, and there are no other poles in the upper momentum half plane (i.e., poles in the physical energy sheet can only be bound states with negative energy). If the *S* matrix has a pole at k, then it also has a pole in $-k^*$, and zeros at -k and k^* . Hence a bound state actually hides two coinciding poles. In Fig. 1 and in Fig. 15 of Sec. IV it can be clearly seen how these two poles separate at the presence of the periodic perturbation.

Finally, the partial wave series of the scattering solution is convergent and the scattering amplitude (and cross sections) finite at all angles for a potential V_{int} that decreases asymptotically faster than $1/r^3$. In that case, the scattering in the limit of low velocity ($\omega \rightarrow 0$) is isotropic and independent of the energy. This can be expressed using the *s*-wave scattering length *a* (that is well defined), and the fact that $\sigma_e \rightarrow 4\pi |a|^2$. At the presence of inelastic interactions, the scattering length is not necessarily real but can have a (negative) imaginary part that gives the inelastic cross section, $\sigma_r \rightarrow 4\pi |\text{Im}a|/k$, that is inversely proportional to the velocity in the low-velocity limit (the 1/v law). These limiting behaviors of the cross sections will be shown to hold in Fig. 8 and Fig. 11.

III. POLES AND SINGULARITIES WITH PERIODIC DRIVING

In this section we employ the methods presented in the previous sections to study a model system consisting of a spherically symmetric square-well potential and a time-dependent periodic linear drive which acts outside of the well. Using the fundamental frequency of the periodic drive, Ω , we can define the length and energy scales

$$d_o = \sqrt{2\hbar/m\Omega}, \quad E_o = \hbar\Omega/2,$$
 (52)

and originally dimensional variables become nondimensional by rescaling according to

$$\vec{r} \to \vec{r}/d_o, \ \vec{k} \to \vec{k}d_o, \ t \to t\Omega/2,$$
 (53)

after which we have explicitly $\hbar = m = 1$ and the drive's frequency in these units is $\Omega = 2$. With a spherical square-well potential,

$$V_{\text{int}}(|\vec{r}|) = \begin{cases} -V_0, & |\vec{r}| < d, \\ 0, & |\vec{r}| > d \end{cases}$$
(54)

[where d and V_0 are nondimensional, measured in the units of Eq. (52)], and the regular solution inside the well is a spherical Bessel function,

$$\Phi_{\{k,l,m\}}(\vec{r},t) \propto e^{-i(\frac{1}{2}k^2 - V_0)t} j_l(kr)Y_l^m.$$
 (55)

In those units, we take the periodic force of Eq. (15) to be a simple harmonic drive with amplitude F_2 ,

$$F^{\pi}(t) = F_2 \cos 2t,$$
 (56)

regaining Eq. (8).



FIG. 2. The spectrum of the square-well potential [Eq. (54)] as a function of the universal parameter A [Eq. (57)]. As $-A/\pi$ crosses an integer (half-integer) value, a new bound p wave (s wave) appears in the well. Higher order l states appear progressively within these intervals. The nondimensional energy scale is fixed by setting the well size to d = 2.

We define the well parameter

$$A = -\sqrt{2V_0 d^2},\tag{57}$$

which is the single parameter that characterizes the square well [45]. In Fig. 2 we show the spectrum of bound states of the time-independent square well over a small range as a function of the universal parameter A/π . The energy scale is obtained using Eq. (52), fixing d = 2. It can be immediately seen from Eq. (57) that varying Ω is equivalent to leaving A invariant while scaling both V_0 and d, as the energy and distance units are rescaled.

A. Driving a loosely bound s wave

In this subsection we set $-A/\pi \approx 0.504$; i.e., the well is shallow with a single bound *s*-wave state close to threshold. We solve the problem by plugging the potential [Eq. (54)] into Eq. (25); i.e., the drive is restricted to act outside the well. This is in contrast to the results presented in Fig. 1, where the drive was solved for in all space. Choosing this form is motivated by the fact that for a general interaction potential, the problem cannot be solved exactly together with the periodic drive, and some sort of approximation is required. A possible choice [introduced in Eq. (1)] is to divide space into two regions where either the interaction or the periodic drive act. A further discussion and comparison of the two models will be presented in Sec. III C.

Figure 3 shows the values of ω for the (initially) escape (radiating) pole followed by continuation from the *s*-wave bound state at $F_2 = 0$, at 4 different values of V_0 (with A fixed). It can be seen that the crossing of the real ω axis is generic and can be realized at different values of the parameters. For low drive amplitude, the quasibound state's decay rate grows quadratically (as can be inferred from a log-log plot, not shown), which is the expected perturbative result [Eq. (7)]. In the nonperturbative regime the decay rate is clearly nonmonotonic; for a strong enough drive it decreases and reaches 0, as the poles reach the real ω axis.

Figure 4 shows the momentum space values of k_0 for the same solutions, lying on nearly straight lines. The slope of the



FIG. 3. Complex ω space showing 4 resonances of the squarewell potential [Eq. (54)] followed as a function of F_2 [Eq. (56)] by solving Eqs. (24) and (25); i.e., the drive is restricted to act outside the well (in contrast to Fig. 1). The well parameter of Eq. (57) is fixed at $-A/\pi \approx 0.504$, with V_0 taking 4 different values shown in the legend, *d* being adjusted accordingly. Starting from the (single) loosely bound *s*-wave state (see Fig. 2), the poles of the *S* matrix correspond initially (while Im $\omega < 0$) to the escape process (emission out of the well). The pole for $V_0 = 0.557$ crosses the real ω axis outside of the figure boundaries, at the critical parameters [Eq. (11)] $\bar{F}^c \approx 0.260$, $\bar{\omega}^c \approx 3.35 \times 10^{-3}$ (see Fig. 6).

lines varies continuously with the parameters. The properties of the solutions were described at length in Sec. I, and we further discuss the parametric dependence in Sec. IV. Since in momentum space, the poles that start on the imaginary k axis must cross the bisector of their quarter plane before reaching the Imk = 0 line, in the ω plane the poles follow a curved trajectory around the origin, crossing to $\text{Re}\omega > 0$. This guarantees that after the critical point, the solutions that have switched roles (between capture and emission) remain valid—with $\text{Im}\omega \neq 0$ and one channel (j = 0) that does not decay asymptotically, as required by the continuity equation.

The superposition of different bound and diverging components can be seen in the solution coefficients of the expansion in Eq. (26), which are depicted in Fig. 5 for a particular state. The



FIG. 4. Complex k space showing the value of k_0 at the poles of Fig. 3. The pole trajectories can be seen to lie almost along straight lines in this case (see however Fig. 14). The poles of the corresponding capture process (related by time-reversal invariance of the Hamiltonian) have mirror-imaged trajectories at the right half plane (not shown). For each (capture or emission) pole of the *S* matrix, there is also a zero at -k.



FIG. 5. Absolute value of the matching coefficients of the solution [Eq. (26)], in the interior region $(a_{2n,l})$ and in the exterior region (b_{2j,l_1}) , for the lowest pole Fig. 3 with $V_0 = 0.557$, at $F_2 = 0.03$. The superposition of (a small number of) components can be seen, both inside and outside the well, of decaying $(j \leq 0)$ as well as traveling waves $(j \geq 1)$.

quasibound s-wave state which for $F_2 = 0$ would have its entire amplitude at (n = 0, l = 0) and $(j = 0, l_1 = 0)$ has developed a superposition of partial waves (here mostly outside of the well). The "checkerboard" pattern is the result of the dipolar nature of the coupling, which conserves $(-1)^{n+l}$ [or $(-1)^{j+l_1}$], and can be used in practice to speed up the numerical calculations. Figure 6 shows the radiation pattern in the leading nondecaying channel for the lowest radiating pole of Fig. 3 with $V_0 = 0.557$. The joint probability density of the radiated waves given by $f(k,\theta)$ of Eq. (44) multiplied by Im ω is plotted at discrete steps of F_2 which determines the value of Rek. On the left, for F_2 lower than the critical value $\bar{F}^c \approx 0.260$, the radiation is an odd Legendre polynomial of $\cos \theta$ (mostly p wave) in j = 1 channels, with the flux initially increasing and then decreasing. On the right-hand side, immediately after the pole crosses the real ω axis (at $\bar{\omega}^c \approx 3.35 \times 10^{-3}$), the radiation



FIG. 6. The leading order contribution to the the probability density $f(k,\theta)$ in spherical coordinates of momentum flux [see Eq. (44)], multiplied by (half) the rate of emission out of the well, plotted for increasing values of F_2 , for the lowest pole in Fig. 3 with $V_0 = 0.557$. The curve colors (from dark blue to bright yellow) encode the value of F_2 . (a) For F_2 up to the critical value $\overline{F}^c \approx 0.260$ the radiation is mostly p wave in the j = 1 channel. (b) At the critical value the distribution switches abruptly to the j = 0 channel (mostly s wave). The multiplication by the emission rate that goes through 0 makes the physical process smoothly decaying and then rising, despite the fact that the distribution $f(k,\theta)$ changes abruptly.



FIG. 7. (a) The squared absolute value of the *S*-matrix element of the incoming *s* wave $(|S_{0,0}|^2)$, color coded on a logarithmic scale as function of F_2 and the real incoming wave energy ω . $S_{0,0}$ vanishes at the value of F_2 and energy for which the pole lies on the real energy axis ($\bar{F}^c \approx 0.260$, $\bar{\omega}^c \approx 3.35 \times 10^{-3}$, for the same pole as Fig. 6), with the finite grid size of the figure giving a small nonzero value. This corresponds to total absorption in this channel; all incoming waves are scattered into the other output channels. (b) The argument of the *S*-matrix element, arg $S_{0,0}$, showing that a 2π phase is accumulated around the critical point in this parameter space. (c) The amplitude of outgoing *p* waves in the j = 1 channel, given by $|S_{2,1}|^2$ that approaches unity near the critical point (in a logarithmic scale). (d) The argument arg $S_{2,1}$, which shows no feature like that of panel (b).

abruptly collapses into the j = 0 channels, composed mostly of *s* waves (and other even harmonics), with the flux increasing as F_2 further increases.

Figure 7 shows some quantities of the S matrix of a scattering formulation as a function of F_2 and the energy $\omega \ll 1$ of an incoming s wave (which determines the limit of scattering of slow particles; see Sec. IIE). The parameters are the same as for the pole followed in Fig. 6. The critical point [defined in Eq. (11)] where the two complex conjugate poles (and zeros) of the S matrix coincide on the real energy axis can be identified as at this point the S-matrix element $S_{0,0}$ vanishes (making the argument undefined), and a 2π phase is accumulated if going around this point in (F_2, ω) parameter space. arg $S_{0,0}$ is defined by continuity from $\omega \rightarrow 0$ for each fixed value of F_2 . Along such a line for which $F_2 < \overline{F}^c$ there is a sharp decrease of the argument as function of ω , while for $F_2 > \overline{F}^c$ there is a sharp rise. The S-matrix element $S_{2,1}$ of scattered p waves in the j = 1 channel approaches unit modulus (in a large region of the parameter plane). At the critical point there is total absorption of the incoming s wave, and it is being radiated out as a (mostly) p wave, with energy higher by at least a drive quantum.

In Fig. 8 the total scattering cross section is plotted for *s* waves with $\omega \ll 1$. The features of scattering of slow particles discussed in Sec. II F (the energy-independent elastic scattering length and the 1/v law for inelastic scattering) hold, as can be deduced from an examination of the elastic and inelastic cross section curves (not shown). The total cross section changes its



FIG. 8. The base-10 logarithm of the total scattering cross sections divided by 2π [Eq. (49)], as a function of the real incoming wave energy ω , for the input *s*-wave channel (which gives the scattering in the limit of slow particles). The parameters are identical to Fig. 7. The curves at several F_2 values, plotted on a log-log scale, show how for $F_2 \rightarrow 0$ the scattering is predominantly elastic (with the $\omega \rightarrow 0$ dependence being that of $\sigma_{e,0}$, i.e., is energy-independent), and becoming increasingly dominated by the inelastic scattering ($\propto \omega^{-1/2}$). The cross section does not show any nonmonotonic features however; compare with Fig. 10 and Fig. 11.

behavior as a function of the drive amplitude; for $F_2 \rightarrow 0$ the scattering is elastic, and it becomes increasingly inelastic as F_2 is increased. The cross section however is monotonic as a function of the energy (as is typical for an *s*-wave resonance) throughout the large variation of F_2 . This is very different from the results of driving a bound *p*-wave state to be discussed in the next subsection.

B. Driving a deeper bound *p* wave

In this subsection we set $-A/\pi \approx 2.565$ and focus on a *p* wave (with magnetic quantum number m = 0) whose binding



FIG. 9. Complex ω space (as in Fig. 3) showing 4 resonances of the square-well potential for the well parameter fixed at $-A/\pi \approx$ 2.565. Each being initially a *p*-wave bound state with energy $\omega(F_2 =$ $0) \leq -2$, at least $N \geq 2$ quanta of the external drive (of frequency 2) must be absorbed in each radiating channel (Re $\omega + 2j > 0, j \geq 2$). Even as the three *S*-matrix poles which are pushed towards a larger quasienergy cross to Re $\omega \geq -2$, the same channels ($j \geq 2$) are the only radiating ones, since the j = 1 channel remains incoming and decaying. At the crossing of the real ω axis the solutions are singular in a similar manner to the near-threshold *s* waves.

energy (for $F_2 = 0$) is a little larger than 2; absorption of two quanta is necessary (for $F_2 \ll 1$) to emit outgoing waves. The same equations are solved as in the previous subsection. Figure 9 shows the values of ω for four poles followed by continuation. Three of those poles cross towards Re $\omega > -2$ while there is one pole that is pushed towards negative energies

Figure 9 shows the values of ω for four poles followed by continuation. Three of those poles cross towards $\text{Re}\omega > -2$ while there is one pole that is pushed towards negative energies. An exceptional point, discussed further in Sec. IV, separates the poles going left and right. As in the previous subsection, as the poles in the lower half plane cross the line $\text{Re}\omega = -2$, the channels whose energy becomes positive (and for that could be termed "open") remain in fact asymptotically decaying (and in fact incoming) so the radiation pattern does not show a qualitative change.

As the poles cross to the upper half ω -plane, the solutions change their nature abruptly and they are singular on the real energy line. The pole trajectories in momentum space will be shown in Sec. IV and present a nontrivial behavior. The radiation pattern (not shown) presents similar features to that of Fig. 6 (with the required modifications of the momentum values and the distribution shape).

Figures 10 and 11 show the characteristics of a scattering setup with slow particles. The limiting low-energy behaviors of the cross sections, discussed in Sec. IIF, are clearly visible: an energy-independent elastic cross section and the 1/v law for the inelastic scattering cross section. Sharp features and nonmonotonicity of the scattering as function of F_2 and in particular of ω are present, resembling shape and Fano resonances (we return to this point briefly in Sec. IV). The presence of the pole at $\bar{\omega}^c \approx -1.9995$ suggests that at the critical point, an incoming *s* wave with low energy $\omega = \bar{\omega}^c + 2$ emits one quantum of energy into the drive and is completely captured into the (long-lived) bound state, only to be radiated as a (mostly) *p* wave after absorbing two quanta from the drive.

C. Approximations

We now discuss the truncation of the potential introduced in Eq. (25) that we repeat here,

$$V_{\rm in}^{(1)}(\vec{r},t) = [V_{\rm int}(r)]\Theta(d-r),$$

$$V_{\rm out}^{(1)}(\vec{r},t) = [-\ddot{\vec{F}}(t)\cdot\vec{r} + V_F(t)]\Theta(r-d), \qquad (58)$$

which corresponds to truncating the axial drive inside the well. The solutions presented in Secs. III A and III B were all obtained using this form of the potential. The solutions presented in Fig. 1 do not employ this truncation, but rather solve the full problem with

$$V_{\rm in}^{(2)}(\vec{r},t) = [V_{\rm int}(r) - \vec{F}(t) \cdot \vec{r} + V_F(t)]\Theta(d-r),$$

$$V_{\rm out}^{(2)}(\vec{r},t) = [-\ddot{\vec{F}}(t) \cdot \vec{r} + V_F(t)]\Theta(r-d),$$
 (59)

with the parameters for Fig. 1 being

$$-A/\pi \approx 0.504, \quad V_0 = 1.977.$$
 (60)

For $V_{in}^{(2)}$ of Eq. (59), the wave functions in the interior region become time-dependent and have to be Fourier-expanded as in the exterior region [making the matrices *C* and *F* of Eq. (39) nondiagonal]. As noted above, $V_F(t)$ is the result of taking Eq. (2) as the physical starting point (an oscillating center of



FIG. 10. (a) As in Fig. 7, $|S_{0,0}|^2$ color coded on a logarithmic scale as function of F_2 and ω , corresponding to the pole with $V_0 = 6.75$ in Fig. 9 (with $\bar{F}^c \approx 0.0898$, $\bar{\omega}^c \approx -1.9995$). (b) The argument of the *S*-matrix element, arg $S_{0,0}$, showing a 2π phase jump, around \bar{F}^c and $\omega \approx 5 \times 10^{-4}$, which plausibly corresponds to an incoming *s* wave completely transferred into the quasibound *p*-wave state by emitting a quantum of energy into the drive, and then being radiated as a *p* wave, 2 quanta higher in energy. We note the different range of the argument as compared with Fig. 7, determined by the $\omega \rightarrow 0$ limit. (c) The base-10 logarithm of the elastic [Eq. (47)] and (d) the inelastic [Eq. (48)] cross sections divided by 2π , for the input *s*-wave channel (in the limit of scattering of slow particles). Sharp features exist in the cross sections (more details are discernible in Fig. 11) in a relatively small range of F_2 and a large range of energy, to be contrasted with Fig. 8.

the potential). V_F then cancels the prefactor in Eq. (14),

$$e^{-i\int' \frac{1}{2}F^2(t')dt'},$$
 (61)

that would otherwise multiply the wave function. If Eq. (5) is the physical potential (a static potential with an external, periodically modulated linear force), we can define $V_{in}^{(3,4)}$ and $V_{out}^{(3,4)}$ to be equal to $V_{in}^{(1,2)}$ and $V_{out}^{(1,2)}$ with $V_F(t)$ set to 0. Then the solutions have to be matched including the term of Eq. (61) that shifts the quasienergy and modifies also the Fourier expansion.

Figures 12 and 13 show the pole trajectories for a nearthreshold *s*-wave bound state of a well with $-A/\pi \approx 2.5037$, solved for three increasing different well depths V_0 , and for each value, comparing the four potentials $V_{\text{in}}^{(p)} + V_{\text{out}}^{(p)}$ with p = 1,2,3,4.

The most notable result is that the singularities studied in this work exist with all of the studied forms of the potential. They are the result of tunneling and interference, and since they exist with p = 3, they are not hindered by the possibility to directly couple to the continuum by absorbing energy within the well. The solution with p = 1 approximates the potential with p = 2, which can be considered as the "true" potential in a realization that starts from Eq. (2), although there are noticeable differences in the slopes in momentum space, which lead to some deviations in ω space as well. This approximation



FIG. 11. (a) The elastic, (b) inelastic, and (c) total scattering cross sections divided by 2π [Eqs. (47)–(49)], vs the energy of the incoming wave (on a log-log plot), taken at a few values of F_2 around the critical value $\bar{F}^c \approx 0.0898$, for the same parameters as in Fig. 10. The nonmonotonic features of the cross sections as function of energy and the strong alteration of the curves as function of F_2 are evident, as well as the $\omega \rightarrow 0$ dependence of $\sigma_{e,0}$ (energy-independent) and of $\sigma_{r,0} (\propto \omega^{-1/2})$, as required by the slow-particle scattering limits.

depends on the norm of the wave function in the exterior region (which is subject to drive). The poles with p = 3 and p = 4are closer to each other (for deeper well depths) in energy space because of the importance of V_F , although in momentum space they do not coincide; in fact the poles for p = 2 and p = 4 coincide exactly in momentum space, because when V_F appears both within and outside the well, its value does not enter the momentum matching.

We note the difference between the expansion presented in the previous sections and the well-known Floquet formalism for treating periodic Hamiltonians [101-103], or timedependent perturbation theory [104]. The periodicity of the Hamiltonian allows defining an extended Hilbert space in position and time, which can be spanned by a set of spatially orthogonal wave functions and a Fourier basis for timeperiodic functions. The current expansion however, employs exact wave functions (at least asymptotically) that vary with ω and are not separable in time and space, in contrast to an expansion using a fixed separable basis that would typically require significantly more basis functions. The connection to the physical solutions is transparent and the explicit use of analytic wave functions in each region gives access to details of the spectrum which may be hard to locate otherwise, and in particular quantum defect theory (QDT), discussed in Sec. I, can be used for the expansion of wave functions in the interior region. Our approach is nonperturbative in both potentials, but neglects the effect of either potential in some region of space and the obtained solutions can be considered, if necessary, as a starting point for an expansion that will correct for the neglected contributions.

Finally we note that the presented expansion and numerical results have been verified by using two different numerical routines in two different programming languages (MATLAB and MATHEMATICA), and then directly by plugging the explicit



FIG. 12. Complex *k* space comparing poles solved using different potentials defined in Sec. III C, studying the approximation involved in truncating the drive inside the well, and also the effect of V_F of Eq. (6). See the text for the analysis.

wave function into the time-dependent Schrödinger equation and verifying the vanishing of both sides of the equation, to the numerical precision possible in the calculation and according to the truncated components.

IV. OUTLOOK

The momentum space values of k_2 for the poles followed in Sec. III B (Fig. 9) are shown in Fig. 14. In contrast to Fig. 4 of Sec. III A, here the pole trajectories in momentum space deviate significantly from straight lines. This may correspond to the proximity of other poles ("resonance interaction") and to the existence of the exceptional point nearby [52,105], discussed below.

Figure 15 compares two scenarios for the pole trajectories in complex momentum and energy planes when varying V_0 continuously. In panels (a) and (b), a smooth rotation of the pole trajectories can be seen as their slope changes continuously with V_0 , going through a point which plausibly shows an interchange of the poles which become the capture and emission poles (which are indistinguishable here at $F_2 = 0$, but possibly could be distinguished with a Floquet invariant like the Krein signature [106]). In panels (c) and (d), A is



FIG. 13. Complex ω space for the poles of Fig. 12 solved using different potentials (see text for details).



FIG. 14. Complex *k* space showing the value of k_2 at the poles of Fig. 9. The pole trajectories can be seen to deviate significantly from the almost straight lines seen (for different parameters) in Fig. 4. A further study is required in order to explain this and possibly relate it to "resonance interaction" and to the nearby exceptional point (see the text and Fig. 15).

fixed as in Fig. 14, and V_0 is varied in small range around the point at which the pole trajectories of Fig. 14 seem to "branch." The existence of an exceptional point is clearly seen, where two bound states for $F_2 = 0$ coincide in energy (mod 2), around which the parametric dependence appears to be nonanalytic. An initial study indicates an entire line of exceptional points that emanates from this point in (V_0, F_2) parameter space. Further study is required in order to check whether an exceptional point can be followed up to the



FIG. 15. (a) Complex ω space and (b) complex k plane for a series of poles with a fixed value of A, as in Fig. 3, and V_0 varied between 1.15 and 1.3 (the 9 nearly straight lines starting at adjacent values of ω or Imk, respectively). The axis markings and the scale of $F_2 \ll 1$ (given by the color code) have been removed for clarity. A smooth rotation of the two pole curves (in both energy and momentum planes) can be seen as function of V_0 . This is to be contrasted with the seemingly nonanalytic point visible in (c) complex ω space and (d) complex k plane for a series of p-wave poles with A fixed as in Fig. 9 and Fig. 14, and $V_0 \in [7.078, 7.091]$. Here it is a Floquet-induced (timeindependent) exceptional point, where the p-wave bound state with energy $\omega \approx -2.125$ coincides in energy (mod 2) with the s wave of $\omega \approx -0.125$, for $V_0 \approx 7.09$. A further study is required to confirm a possible line of exceptional points passing through this point.

real ω axis as the Floquet poles studied above, whence it may share further similarities with a spectral singularity of scattering with a real energy. As discussed in Sec. I, the role of exceptional points in non-Hermitian (open) systems is attracting increasing attention, and new effects are being actively explored.

The cross sections in Fig. 11 show features resembling shape and Fano resonances, which are important in atomic and also nanoscale structures [107-109] (with the distinction that here the scattering is inelastic, and the potential is timedependent and noncentral, both of those aspects playing an important role). As discussed in Sec. IIF, the two linearly independent solutions of a power-law potential $-C/r^{\alpha}$ with $\alpha > 3$ are analytic in complex energy and momentum planes. Hence, when applying the expansion of the current work to match such wave functions in the interior region with the driven particle wave functions in the exterior region, no nonanalyticity is introduced in the entire k plane (at any fixed $r \neq d$). Therefore, following the poles of the driven problem in the k plane for such a potential should be possible (apriori) as shown here for a finite-range potential. A further study of the possibility to tailor scattering and resonances using Floquet driving, in particular of atomic systems [71], is a promising direction to apply the techniques developed in this work. Indeed, the singular points analyzed in the current work can be found in an explicit calculation employing QDT with the polarization interaction ($\alpha = 4$) of a cotrapped ion-atom system [98]. More broadly, Floquet driven systems are often analyzed in terms of a time-independent, effective approximation. Our results show a scenario where accounting for the time-dependent nature of the drive is essential, and the implications to many-body Floquet systems [110–113] present an intriguing future direction to explore.

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APPENDIX A: EXPANSION OF LINEARLY DRIVEN CYLINDRICAL WAVES IN SPHERICAL WAVES

In the following we will use a representation of the spherical Hankel function of the first kind as an integral over cylindrical waves [114,115] in the form

$$h_l^{(1)}(kr)P_l^m(\cos\theta)e^{im\varphi} = \int_{C^{(1)}} d\alpha \sin\alpha \frac{1}{2}(-i)^{l-m}P_l^m \times (\cos\alpha)\chi_m^{(1)}(\vec{r};k,\alpha), \quad (A1)$$

where P_l^m are the associated Legendre polynomials and the directed contour of integration $C^{(1)}$ lies in complex α plane. For *k* with a positive imaginary part we must take $C^{(1)} = \pi/2 + i(\infty, -\infty)$, on which $\cos \alpha \in i(\infty, -\infty)$, and $\sin \alpha \in (0, \infty)$. Then $h_l^{(1)}(kr)$ decays asymptotically as $e^{-(\operatorname{Im}k)r}/r$ [the integral which gives $h_l^{(1)}(kr)$ is well defined for any $\rho > 0$, and decays for $r \to \infty$, which is just what is required for the validity of the solution]. For *k* real and positive the contour of integration is given by $C^{(1)} = i(\infty, 0) + [0, \pi] + \{\pi + i(0, -\infty)\}$, on which $\cos \alpha \in (\infty, -\infty)$, and Im $\sin \alpha \ge 0$. For any value of *k*, we have similarly to Eq. (A1)

$$j_l(kr)P_l^m(\cos\theta)e^{im\varphi} = \int_{[0,\pi]} d\alpha \sin\alpha \frac{1}{2}(-i)^{l-m}P_l^m \times (\cos\alpha)\chi_m^{(J)}(\vec{r};k,\alpha),$$
(A2)

where j_l is a spherical Bessel function. Writing $h_l^{(2)} = 2j_l - h_l^{(1)}$, and using Eq. (A2) and the fact that the l,m-dependent coefficients in Eqs. (A1) and (A2) are identical, we get an expression identical in form to Eq. (A1), with the outgoing waves replaced by incoming waves $h_l^{(2)}$, and a contour $C^{(2)}$. A similar derivation can be repeated in the lower half of the complex k plane, making the expansion valid for every complex k (which also follows by analytic continuation).

The proof of Eq. (21) proceeds by using Eq. (18) to write

$$e^{i\dot{F}^{\pi}(t)z} \int_{C^{(a)}} d\alpha \sin \alpha \chi_{m}^{(a)}(\vec{r};k_{2j},\alpha) b_{2j}^{(a)}(\alpha) e^{-iF^{\pi}(t)k_{2j}\cos\alpha}$$

$$= e^{i\dot{F}^{\pi}(t)z} \int_{C^{(a)}} d\alpha \sin \alpha \chi_{m}^{(a)}(\vec{r};k_{2j},\alpha) \sum_{l_{1}} b_{2j,l_{1}} P_{l_{1}}(\cos\alpha) \sum_{l_{2}} (-i)^{l_{2}} (2l_{2}+1) j_{l_{2}}[F^{\pi}(t)k_{2j}] P_{l_{2}}(\cos\alpha)$$

$$= e^{i\dot{F}^{\pi}(t)z} \sum_{l_{1},l_{2}} b_{2j,l_{1}}(-i)^{l_{2}} (2l_{2}+1) j_{l_{2}}[F^{\pi}(t)k_{2j}] \sum_{l_{3}} W(P_{l_{1}},P_{l_{2}},P_{l_{3}}^{m}) \int_{C^{(a)}} d\alpha \sin \alpha \chi_{m}^{(a)}(\vec{r};k_{2j},\alpha) P_{l_{3}}^{m}(\cos\alpha)$$

$$= e^{i\dot{F}^{\pi}(t)z} \sum_{l_{1},l_{2}} b_{2j,l_{1}} j_{l_{2}}[F^{\pi}(t)k_{2j}] \sum_{l_{3}} c_{l_{1},l_{2},l_{3}} h_{l_{3}}^{(a)}(k_{2j}r) P_{l_{3}}^{m}(\cos\theta) e^{im\varphi}$$

$$= \sum_{l_{4}} i^{l_{4}} (2l_{4}+1) j_{l_{4}}[\dot{F}^{\pi}(t)r] P_{l_{4}}(\cos\theta) \sum_{l_{1},l_{2}} b_{2j,l_{1}} j_{l_{2}}[F^{\pi}(t)k_{2j}] \sum_{l_{3}} c_{l_{1},l_{2},l_{3},l_{4},l} Y_{l}^{m}(\theta,\varphi), \qquad (A3)$$

where the multiplicative factors $e^{-i\frac{1}{2}k_{2j}^2t}$, N_{2j,l_1}^m , and $S_{2j,l_1}^{(a)}$ have been omitted for simplicity, and by using the definition of $R_{2j,l_1,l}^{(a)}(r,t)$ given in Eq. (22), Eq. (A3) results in Eq. (21). In the derivation of Eq. (A3), the plane-wave expansion in terms of spherical Bessel functions has been used (twice), the coefficients of expansion of a product of two (associated) Legendre polynomials (which can be written using Wigner 3-j symbols) are defined by

$$W(P_{l_1}^{m_1}, P_{l_2}^{m_2}, P_{l_3}^{m_3}) = \{2(l_3 + m_3)! / [(2l_3 + 1)(l_3 - m_3)!]\}^{-1} \\ \times \int_{-1}^{1} P_{l_1}^{m_1}(w) P_{l_2}^{m_2}(w) P_{l_3}^{m_3}(w) dw,$$
(A4)

the coefficients c_{l_1,l_2,l_3} are obtained using Eq. (A1) and Eq. (A4) and given by

$$c_{l_1,l_2,l_3} = 2(2l_2+1)(-i)^{l_2}i^{l_3-m}W(P_{l_1},P_{l_2},P_{l_3}^m), \quad (A5)$$

and the coefficients $c_{l_1,l_2,l_3,l_4,l}$ are similarly

$$c_{l_1,l_2,l_3,l_4,l} = c_{l_1,l_2,l_3}(2l_4+1)i^{l_4}W(P_{l_3}^m, P_{l_4}, P_l^m)/N_l^m, \quad (A6)$$

with the definitions

$$Y_l^m(\theta,\varphi) = N_l^m P_l^m(\cos\theta) e^{im\varphi},$$

$$N_l^m = (-1)^m \sqrt{(2l+1)/4\pi} \sqrt{(l-m)!/(l+m)!}.$$
 (A7)

APPENDIX B: THE EXPECTATION VALUE OF TENSOR OPERATORS

In this appendix we give explicitly the expansion of integrals which are required in order to calculate expectation values of general tensor operators, in the Floquet eigensolutions of Sec. II B. For simplicity we treat here only the most useful case of axially symmetric wave functions, with m = 0 (no φ dependence). Using the notation of Eq. (50), we start by writing the π -periodic part of the wave function in the form

$$\phi^{\pi}(\vec{r}) = \sum_{n,l} a_{2n,l} e^{-i2nt} \frac{1}{r} u_{2n,l}(r) Y_l^0, \tag{B1}$$

which corresponds to the expansion in Eq. (26) of wave functions in the interior region. For such wave functions, we define the (unnormalized) expectation value in the interior region of a purely radial operator O(r),

$$\begin{aligned} \mathfrak{I}_{0}[\mathcal{O}(r)] &= \int d^{3}\vec{r} |\phi^{\pi}(\vec{r},t)|^{2} \mathcal{O}(r) \\ &= \sum_{(n,l),(n',l')} \delta_{l,l'} e^{2i(n-n')t} a_{2n,l}^{*} a_{2n',l'} \\ &\times \int dr [u_{2n,l}]^{*} \mathcal{O}(r) u_{2n',l'}. \end{aligned} \tag{B2}$$

The above expression can be rewritten as

$$\mathfrak{I}_0[\mathcal{O}(r)] = \sum_l I_{l,l}[\mathcal{O}(r)], \qquad (B3)$$

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where we have defined for convenience the functional [symmetric under the exchange $(n,l) \leftrightarrow (n',l')$]

$$I_{l,l'}[\mathcal{O}(r)] = \sum_{n \leqslant n'} (2 - \delta_{n,n'}) \operatorname{Re} \left\{ e^{2i(n-n')t} a_{2n,l}^* a_{2n',l'} \\ \times \int dr [u_{2n,l}]^* \mathcal{O}(r) u_{2n',l'} \right\},$$
(B4)

with the summation taken over pairs of states enumerated by $\{(n,l), (n',l')\}$ with fixed *l* and *l'* obeying $n \leq n'$.

For example, the normalization integral calculated for any time (see Appendix C) can be written as

$$\mathfrak{I}_0[\hat{1}] = \sum_l I_{l,l}[\hat{1}],$$
 (B5)

with $\hat{1}$ the identity operator. Any other expectation value must then be divided by the value of this normalization integral. Similarly, the expectation value of the squared angular momentum operator \vec{L}^2 is given by

$$\Im_0[\vec{L}^2] = \sum_l l(l+1)I_{l,l}[\hat{1}].$$
 (B6)

For an operator of a general radial part multiplied by the position vector, $O(r)\vec{r}$, only the Cartesian z component survives the integral (for axially symmetric wave functions), and we can write using $z/r = \cos \theta$

$$\mathcal{D}_{1}[\mathcal{O}(r)\vec{r}] = \int d^{3}\vec{r} |\phi^{\pi}(\vec{r},t)|^{2} \mathcal{O}(r)\vec{r}$$

$$= \hat{z} \sum_{(n,l),(n',l')} p_{l,l'} e^{2i(n-n')t} a_{2n,l}^{*} a_{2n',l'}$$

$$\times \int dr [u_{2n,l}]^{*} \mathcal{O}(r)r \, u_{2n',l'}$$
(B7)

with the coefficients being

$$p_{l,l'} = 2\pi N_l^0 N_{l'}^0 \int d\theta \sin \theta \cos \theta P_l(\cos \theta) P_{l'}(\cos \theta).$$
(B8)

Using the fact that $p_{l,l'} = p_{l',l}$ and since nonzero terms will have |l - l'| = 1, we find

$$\Im_{1}[\mathcal{O}(r)\vec{r}] = \hat{z} \sum_{l} p_{l,l+1}\{I_{l,l+1}[\mathcal{O}(r)r] + I_{l+1,l}[\mathcal{O}(r)r]\}.$$
(B9)

For an operator with a general radial part multiplied by a bilinear combination of \vec{r} components, $\mathcal{O}(r)\vec{r}_{\alpha}\vec{r}_{\beta}$, where $\alpha,\beta \in \{x,y,z\}$, only the diagonal terms with $\alpha = \beta$ survive the integration (for m = 0), with the result

$$\Im_{2}[\mathcal{O}(r)\vec{r}_{\alpha}\vec{r}_{\beta}] = \delta_{\alpha,\beta} \sum_{l,l'} q_{\alpha,l,l'} I_{l,l'}[\mathcal{O}(r)r^{2}], \qquad (B10)$$

where

$$q_{\alpha,l,l'} = 2\pi N_l^0 N_{l'}^0 \int d\theta \sin \theta \bigg[\cos^2 \theta \delta_{\alpha,z} + \frac{1}{2} \sin^2 \theta (\delta_{\alpha,x} + \delta_{\alpha,y}) \bigg] P_l(\cos \theta) P_{l'}(\cos \theta).$$
(B11)

In all of the above expressions, $I_{l,l'}[\mathcal{O}(r)]$ as defined in Eq. (B4) is valid in the interior region. To get the complete

result for expectation values in whole space, the integration over the exterior region must be added, where the wave functions are expanded differently in Eq. (26). In this case, Eq. (B1) is to be replaced by

$$\phi^{\pi}(\vec{r},t) = \sum_{j,l_1} b_{2j,l_1} e^{-i2jt} \sum_l \frac{1}{r} u^{\pi}_{2j,l_1,l}(r,t) Y^0_l, \qquad (B12)$$

and accordingly, Eq. (B4) becomes in the exterior region

$$I_{l,l'}[\mathcal{O}(r)] = \sum_{(j,l_1),(j',l'_1)} e^{2i(j-j')t} b_{2j,l_1}^* b_{2j',l'_1}$$
$$\times \int dr [u_{2j,l_1,l}^\pi]^* \mathcal{O}(r) u_{2j',l'_1,l'}^\pi, \quad (B13)$$

with the summation taken over pairs of states enumerated by $\{(j,l_1,l),(j',l'_1,l')\}$ with fixed l and l'. Finally, we note that in the above expressions, the imaginary part of the energy has been omitted; it gives an exponential envelope of the decay or formation rate of the quasibound state. Moreover, all integrals can be performed only on the square-integrable part of the wave function, with the nonnormalizable traveling waves omitted from the sums above, in accordance with the interpretation that these belong to the inaccessible part of the Hilbert space.

APPENDIX C: NORMALIZATION OF THE WAVE FUNCTION

The expectation value of any time-independent (or π periodic) operator is π -periodic for the Floquet eigenstates, possibly with an exponential envelope for complex ω . The normalization integral is not constant in time but rather π periodic because the relative weight of the nonnormalizable components oscillates in time (as they are emitted and reflected back during a period of the drive). In order to calculate an expectation value of an operator (determined by the bound components), its integral must be divided by the squared norm, both of which being π -periodic functions that can be calculated using Appendix B (after which averaging is possible). The normalization in the interior region can be obtained without explicitly performing the integration, directly from the wave functions and their gradients at the matching point. This can be useful especially when the interior wave functions are not explicitly known close to the origin, but rather are determined within a QDT formulation [82,83,88,89,92]. The projection of two eigenfunctions ϕ_1 and ϕ_2 of the interior Hamiltonian with

energies ε_1 and ε_2 correspondingly is shown in Appendix D to be

$$2\operatorname{Re}\int_{0}^{d}\phi_{1}^{*}\phi_{2}r^{2}dr = (\varepsilon_{1} - \varepsilon_{2})^{-1}\operatorname{Re}\{u_{1}^{*}u_{2}' - u_{2}^{*}u_{1}'\}|_{d}, \quad (C1)$$

where $u_1' \equiv \partial_r u_1$, and ε_1 , ε_2 are assumed to have equal imaginary parts. The left-hand side of Eq. (C1) gives the integrals required for the normalization, with the factor of 2 relevant for the off-diagonal projections (when $\phi_1 \neq \phi_2$). In the limit of $\phi_1 \rightarrow \phi_2$ we have for the diagonal normalization terms

$$\int_0^d |\phi_1|^2 r^2 dr = \frac{1}{2} \lim_{\epsilon_1 \to \epsilon_2} (\epsilon_1 - \epsilon_2)^{-1} [u_1^* u_2' - u_2^* u_1']|_d.$$
(C2)

APPENDIX D: THE PROJECTION OF TWO EIGENFUNCTIONS OF THE INTERNAL HAMILTONIAN

In order to derive Eq. (C1), let $\varepsilon_1, \varepsilon_2$ be the (possibly complex) energies of two complex eigenfunctions ϕ_1, ϕ_2 of the interior Hamiltonian $H_{\rm in} = -\frac{1}{2}\nabla^2 + V_{\rm in}$. For the projection of the two within the interior region, we can write

$$0 = \langle \phi_2 | (H_{\rm in} - \varepsilon_1) | \phi_1 \rangle - \langle \phi_1 | (H_{\rm in} - \varepsilon_2) | \phi_2 \rangle.$$
 (D1)

By canceling the potential energy terms, we get after rearranging the kinetic terms and terminating the integration at an arbitrary point d (which is allowed since the equality above holds identically in space)

$$\int_{0}^{d} (\varepsilon_{1}\phi_{2}^{*}\phi_{1} - \varepsilon_{2}\phi_{1}^{*}\phi_{2})r^{2}dr = -\frac{1}{2}\int_{0}^{d} (u_{2}^{*}\partial_{r}^{2}u_{1} - u_{1}^{*}\partial_{r}^{2}u_{2})dr$$
$$= -\frac{1}{2}[(u_{2}^{*}\partial_{r}u_{1} - u_{1}^{*}\partial_{r}u_{2})|_{d} - \int_{0}^{d} (\partial_{r}u_{2}^{*}\partial_{r}u_{1} - \partial_{r}u_{1}^{*}\partial_{r}u_{2})dr],$$
(D2)

where the factor of 1/2 is the prefactor in the kinetic energy term $-\frac{1}{2}\nabla^2$, as in Eq. (51). In the second line of the above equation, the integrated term is purely imaginary being the difference of two complex conjugates. Taking the complex conjugate of the entire equation and adding, this term drops and we get

$$\int_{0}^{d} 2\operatorname{Re}\{(\varepsilon_{1} - \varepsilon_{2})\phi_{2}\phi_{1}^{*}\}r^{2}dr = -\frac{1}{2}2\operatorname{Re}\{u_{2}\partial_{r}u_{1}^{*} - u_{1}\partial_{r}u_{2}^{*}\}|_{d},$$
(D3)

which gives immediately Eq. (C1).

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