Effect of nuclear quadrupole moments on parity nonconservation in atoms

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Nuclei with spin $I \ge 1$ have a weak quadrupole moment which leads to tensor contribution to the parity nonconserving interaction between nuclei and electrons. We calculate this contribution for atoms of current experiment interest Yb⁺, Fr, and Ra⁺. We have also performed order of magnitude estimates and found strong enhancement of the weak quadrupole effects due to the close levels of opposite parity in many lanthanoids (e.g., Nd, Gd, Dy, Ho, Er, Pr, Sm) and Ra. Another possibility is to measure the parity-nonconservation (PNC) transitions between the hyperfine components of the ground state of Bi. Since nuclear weak charge is dominated by neutrons this opens a way of measuring quadrupole moments of neutron distribution in nuclei.

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I. INTRODUCTION

Studying parity-nonconservation (PNC) in atoms is a way of testing the standard model at low energy as well as searching for new physics beyond the standard model (see, e.g., [1,2]). The most precise measurements of the PNC in Cs [3] supported by accurate atomic calculations [4–9] show no significant deviation from the standard model. Atomic PNC experiments may also measure the nuclear anapole moments [3,10–13] and the ratio of PNC amplitudes for different isotopes which is not sensitive to the accuracy of atomic calculations [14,15].

Atomic PNC measurements can also be used to study the neutron distribution in nuclei. Several studies looked at the effect of *neutron skin* (the difference in radius of the proton and neutron distributions) on the PNC in atoms and demonstrated that it can give a small but measurable contribution to the PNC amplitude (see, e.g., [16]). The study of the neutron distribution should help to establish the equation of state for the nuclear matter and properties of neutron stars including the mass boundary for the stability of neutron stars (the neutron repulsion at short distances prevents collapse of a neutron star to a black hole).

In the present paper we provide a theory for a different method to study the neutron distribution in atomic PNC experiments. It was noted in Ref. [17] (see also [1]) that the nuclear quadrupole moment induces a tensor PNC weak interaction between the nucleus and electrons in atoms and molecules. In Ref. [18] it was shown that the combined action of the weak charge and the quadrupole hyperfine interaction produces a similar effect but of a significantly smaller amplitude. Note however, that such effect may be enhanced if there are close levels mixed by the quadrupole hyperfine interaction.

In Ref. [19] it was argued that the tensor effects of the weak quadrupole moments are strongly enhanced for deformed nuclei and may get a significant additional enhancement due to the close atomic and molecular levels of opposite parity with a difference of the electron angular momenta $|J_1 - J_2| \le 2$. These selection rules are similar to that for the effects of the time reversal (T) and parity (P) violating nuclear magnetic quadrupole moment (MQM). Therefore, nuclei, molecules, and molecular levels suggested for the MQM search in Ref. [20], for example, $|\Omega| = 1$ doublets in the molecules

¹⁷⁷HfF+, ²²⁹ThO, ¹⁸¹TaN, will also have enhanced effects of the weak quadrupole.

Differences in the selection rules for the scalar weak charge $(J_1-J_2=0)$, vector anapole moment $(|J_1-J_2|\leqslant 1)$, and the tensor weak quadrupole moment $(|J_1-J_2|\leqslant 2)$, or the difference in the dependence of the PNC effect on the hyperfine components of an atomic transition if more than one operator contribute, allows one to separate the contribution of the weak quadrupole.

The weak charge of the neutron (-1) exceeds the weak charge of the proton (0.08) by more than an order of magnitude. Therefore, the measurements of the PNC effects produced by the weak quadrupole moment allows one to measure the quadrupole moments of the neutron distribution in nuclei.

In present paper we perform the relativistic many-body calculations of the weak quadrupole effects in atoms of current experiment interest Yb⁺, Fr, and Ra⁺. We have also performed order of magnitude estimates and found strong enhancement of the weak quadrupole effects due to the close levels of opposite parity in many lanthanoids (e.g., Nd, Gd, Dy, Ho, Er, Pr, Sm) and Ra. Finally, we have calculated the weak quadrupole effects in the PNC transitions between the hyperfine components of the ground state of Bi.

II. THEORY

An effective single-electron interaction operator that is responsible for parity nonconservation (PNC) in atom is given by

$$h_{\text{PNC}} = -\frac{G_F}{\sqrt{2}} \gamma_5 [ZC_{1p} \rho_{0p}(r) + NC_{1n} \rho_{0n}(r)] - \frac{G_F}{\sqrt{2}} \gamma_5 Y_{20} [ZC_{1p} \rho_{2p}(r) + NC_{1n} \rho_{2n}(r)], \quad (1)$$

where $G_F \approx 2.2225 \times 10^{-14}$ in atomic units (a.u.) is the Fermi constant, the Dirac matrix γ_5 is defined as in Ref. [1], Z and N are the number of protons and neutrons, the coefficients $2C_{1p} = (1 - 4\sin^2\theta_W) \approx 0.08$, $2C_{1n} = -1$ are the proton and neutron weak charges, $\rho_p(\mathbf{r}) \approx \rho_{0p}(r) + \rho_{2p}(r)Y_{20}(\theta,\phi)$ and $\rho_n(\mathbf{r}) \approx \rho_{0n}(r) + \rho_{2n}(r)Y_{20}(\theta,\phi)$ are proton and neutron densities in a nucleus normalized to unity, $\int \rho(r)d^3r = 1$. We have taken into account that if the nuclear spin has the maximal

(or any fixed) projection on the z axis the quadrupole part of the density is proportional to $Y_{20}(\theta, \phi)$.

Below we will concentrate on the neutron contribution since the proton contribution to the weak charge is small and may be treated as a correction. Therefore, to simplify the formulas we assume that the spherical part of the proton density distribution is equal to that for neutrons: $\rho_{0p} = \rho_{0n} = \rho_0$. Anyway, the neutron skin is small.

If we assume that $\rho_{2n}(r) = K_n \rho_0(r)$, the proportionality constant can be expressed in terms of the quadrupole moment $Q_n = Q_{zz} = N \int (3z^2 - r^2)\rho(r)d^3r$: $K_n = \sqrt{5}Q_n/(4N\sqrt{\pi}\int\rho_0r^4dr)$ and the tensor part of the weak interaction is

$$h_Q = -\frac{G_F}{2\sqrt{2}} \gamma_5 Y_{20} \rho_0 \frac{\sqrt{5\pi} Q^{\text{TW}}}{\langle r^2 \rangle},$$
 (2)

where $Q^{\mathrm{TW}}=2C_{1n}Q_n+2C_{1p}Q_p=-Q_n+0.08Q_p$ is the weak quadrupole moment, $\langle r^2\rangle=4\pi\int\rho_0r^4dr\approx 3R_N^2/5$ is the mean squared nuclear radius, R_N is the nuclear radius. The quadrupoles Q_p of the proton distribution in nuclei are measured and tabulated in the literature. The neutron quadrupoles Q_n have never been measured. In deformed nuclei $Q_n\approx (N/Z)Q_p$.

In the electromagnetic transitions between the hyperfine components the nuclear spin projection changes, and we should present the PNC interaction Hamiltonian in terms of the irreducible tensor components:

$$h_Q = -\frac{5G_F}{2\sqrt{2}\langle r^2 \rangle} \sum_q (-1)^q T_q^{(2)} Q_{-q}^{\text{TW}}, \tag{3}$$

where $T_q^{(2)}=C_q^{(2)}\gamma_5\rho_0(r)$ is the electronic part of the operator, $C_q^{(2)}=\sqrt{4\pi/5}Y_{2q}$ and for the second rank tensor $Q^{\rm TW}=2Q_0^{\rm TW}$.

The PNC electric dipole amplitude between states $(|i\rangle \rightarrow |f\rangle)$ with the same parity due to the tensor weak interaction is

$$E_{\text{PNC}}^{i \to f} = \sum_{n} \left[\frac{\langle f | \boldsymbol{d} | n \rangle \langle n | h_{Q} | i \rangle}{E_{i} - E_{n}} + \frac{\langle f | h_{Q} | n \rangle \langle n | \boldsymbol{d} | i \rangle}{E_{f} - E_{n}} \right], \quad (4)$$

where $|a\rangle \equiv |J_a F_a M_a\rangle$ is a hyperfine sate and $d = -e \sum_i r_i$ is the electric dipole operator, F = J + I is the total angular momentum of an atom, J is the electron angular momentum, and I is the nuclear spin. More detailed formulas are presented in the Appendix.

In performing numerical calculations we follow our earlier work [21] on spin-dependent PNC in single-valence-electron atoms. We include the nuclear anapole moment contribution as well, so that in most of cases the total PNC amplitudes consist of three terms: the spin-independent contribution due to weak nuclear charge, the anapole moment contribution, and the weak quadrupole moment contribution. This allows us to fix the relative sign of all three terms. Random-phase approximation (RPA) is used for all operators of external fields, including the PNC operators and the electric dipole operator. Brueckner orbitals are used to include the core-valence correlations (see [21] for details).

We also use analytical estimations to check numerical results and their uncertainty. To do this we use the radial wave functions near the nucleus from Ref. [1]:

$$f_{n\kappa} = \frac{\kappa}{|\kappa|} (\kappa - \gamma) \left(\frac{Z}{a_0^3 v^3} \right)^{1/2} \frac{2}{\Gamma(2\gamma + 1)} \left(\frac{a_0}{2Z} \right)^{1-\gamma} r^{\gamma}, \quad (5)$$

$$g_{n\kappa} = \frac{\kappa}{|\kappa|} Z \left(\frac{Z}{a_0^3 v^3}\right)^{1/2} \frac{2}{\Gamma(2\gamma + 1)} \left(\frac{a_0}{2Z}\right)^{1-\gamma} r^{\gamma},$$
 (6)

where $\gamma = \sqrt{\kappa^2 - Z^2 \alpha^2}$, a_0 is Bohr radius, $v_n^2 = -1/(2\epsilon_n)$ is the effective principle quantum number, ϵ_n is the orbital energy in a.u., and $\Gamma(x)$ is the Γ function.

Analytical and numerical results agree on the level of 30% or better. The accuracy of the numerical results is few percent for Fr and Ra⁺ and \sim 30% for Yb⁺. A detailed analysis of accuracy of the calculations can be found in Ref. [21]. We believe that the accuracy of our present calculations is the same as in Ref. [21].

III. RESULTS AND DISCUSSION

A. s-s and s-d transitions

Calculated PNC amplitudes between different hfs components of s and d states of $^{173}{\rm Yb}^+$, $^{223}{\rm Fr}$, and $^{223}{\rm Ra}^+$ are presented in Table I. The amplitudes consist of three contributions, the spin-independent contribution due to nuclear weak charge Q_W , the contribution of the nuclear anapole moment x, and the contribution of the neutron quadrupole moment q. We have chosen these atoms because they are considered for the PNC measurements (see, e.g., [21–27]) and because some isotopes of these atoms have deformed nucleus and therefore large quadrupole moments for both proton and neutron distributions. Electric quadropole moments (Q_p) are known and tabulated [28]. The values for considered isotopes are $Q_p(^{173}\text{Yb}) = 2.80(4)b$, $Q_p(^{223}\text{Fr}) = 1.17(2)b$, and $Q_p(^{223}\text{Ra}) = 1.25(7)b$. Using the estimate $Q_n \approx (N/Z)Q_p$ we see that the largest contributions of the neutron quadrupole term to the PNC amplitude is $\sim 10^{-4}$ of the spin-independent contribution. This is a relatively small value which probably means that one should look for enhancement factors, such as, e.g., close states of opposite parity. The atoms considered above do not have such enhancement. They were originally chosen for the measurements of the spin-independent PNC. They have large Z (PNC scales as $\sim Z^3$) and relatively simple electron structure (one electron above closed shells) which allow for accurate interpretation of the measurements. The study of neutron quadrupole moments needs different criteria for choosing the objects for measurements. One could search, e.g., for close states of opposite parity with $\Delta J = 2$. Such states can be only mixed by the neutron quadrupole moment and PNC amplitudes involving such states can be enhanced to the measurable level by small energy intervals. Note also that high accuracy of the calculations is not needed at this stage. Therefore, promising candidates can probably be found in atoms with dense spectra such as atoms with open d or f shells. Molecules can be good candidates too.

TABLE I. PNC amplitudes $\langle a, F_1 | E_z^{\text{PNC}} | b, F_2 \rangle$ (z components) for the s-s and s-d transitions in $^{173}\text{Yb}^+$ (I = 7/2, $Q_W = -96.84$), ^{223}Fr (I = 3/2, $Q_W = -128.25$), and $^{223}\text{Ra}^+$ (I = 3/2, $Q_W = -127.2$). Weak nuclear charge (Q_W), nuclear anapole moment (\varkappa), and neutron quadrupole moment (Q_n) contributions are presented. The unit for Q_n is barn (1 b = 10^{-24} cm²).

Isotope/			PNC amplitude
transition	F_1	F_2	$10^{-10}iea_0$
¹⁷³ Yb ⁺	1	2	$-0.41 \times [1 - 0.022\varkappa - 7.5 \times 10^{-6}Q_n]$
$5d_{3/2}$ -6s	2	2	$-0.53 \times [1 - 0.016\varkappa - 2.7 \times 10^{-6}Q_n]$
	2	3	$-0.17 \times [1 - 0.005 \times + 1.4 \times 10^{-5} Q_n]$
	3	2	$0.28 \times [1 + 0.007 \times -1.1 \times 10^{-5} Q_n]$
	3	3	$-0.48 \times [1 + 0.004 \times + 2.1 \times 10^{-6} Q_n]$
	4	3	$0.37 \times [1 - 0.016 \times + 2.7 \times 10^{-6} Q_n]$
$^{173}Yb^{+}$	1	2	$-6.8 \times 10^{-4} \times +4.8 \times 10^{-6} Q_n$
$5d_{5/2}$ -6s	2	2	$-1.4 \times 10^{-3} \kappa + 1.1 \times 10^{-5} Q_n$
,	2	3	$-4.6 \times 10^{-4} \times -3.1 \times 10^{-6} Q_n$
	3	2	$9.4 \times 10^{-4} \times -8.8 \times 10^{-6} Q_n$
	3	3	$-1.6 \times 10^{-3} \kappa - 8.5 \times 10^{-6} Q_n$
	4	3	$1.2 \times 10^{-3} \kappa + 4.1 \times 10^{-6} Q_n$
²²³ Fr	1	1	$-0.31 \times [1 - 0.023 \times + 3.4 \times 10^{-5} Q_n]$
7s-8s	1	2	$0.54 \times [1 + 0.17\varkappa + 6.7 \times 10^{-6}Q_n]$
	2	1	$0.54 \times [1 - 0.16 \times + 6.7 \times 10^{-6} Q_n]$
	2	2	$0.63 \times [1 - 0.014 \times -6.7 \times 10^{-6} Q_n]$
²²³ Fr	1	0	$-4.5 \times [1 - 0.026 \times -1.1 \times 10^{-5} Q_n]$
$7s-6d_{3/2}$	1	1	$-5.0 \times [1 - 0.026 \times -4.3 \times 10^{-6} Q_n]$
- ,	1	2	$3.9 \times [1 + 0.026 \times -8.9 \times 10^{-6} Q_n]$
	2	1	$-1.7 \times [1 + 0.016 \times + 7.8 \times 10^{-6} Q_n]$
	2	2	$-4.5 \times [1 + 0.016 \times + 4.8 \times 10^{-6} Q_n]$
	2	3	$4.5 \times [1 - 0.015 \times + 2.2 \times 10^{-6} Q_n]$
²²³ Fr	1	1	$4.9 \times 10^{-3} \kappa + 1.0 \times 10^{-4} Q_n$
$7s-6d_{5/2}$	1	2	$-5.8 \times 10^{-3} \kappa - 1.2 \times 10^{-4} Q_n$
•	2	1	$1.7 \times 10^{-3} \kappa - 9.1 \times 10^{-6} Q_n$
	2	2	$6.7 \times 10^{-3} \kappa - 4.0 \times 10^{-5} Q_n$
	2	3	$-7.2 \times 10^{-3} \varkappa + 5.0 \times 10^{-5} Q_n$
$^{223}Ra^{+}$	1	0	$-3.0 \times [1 - 0.025\varkappa - 9.5 \times 10^{-6}Q_n]$
$7s-6d_{3/2}$	1	1	$-3.4 \times [1 - 0.022 \times -5.2 \times 10^{-6} Q_n]$
	1	2	$2.6 \times [1 + 0.016\varkappa - 8.9 \times 10^{-6}Q_n]$
	2	1	$-1.2 \times [1 + 0.0003 \times + 1.4 \times 10^{-5} Q_n]$
	2	2	$-3.0 \times [1 + 0.0062 \times + 2.4 \times 10^{-6} Q_n]$
	2	3	$3.0 \times [1 - 0.015\varkappa + 1.9 \times 10^{-6}Q_n]$
$^{223}Ra^{+}$	1	1	$1.4 \times 10^{-3} \varkappa + 6.2 \times 10^{-5} Q_n$
$7s-6d_{5/2}$	1	2	$-1.6 \times 10^{-3} \varkappa - 7.8 \times 10^{-5} Q_n$
•	2	1	$4.7 \times 10^{-4} \varkappa - 1.0 \times 10^{-5} Q_n$
	2	2	$1.8 \times 10^{-3} \varkappa - 3.5 \times 10^{-5} Q_n$
	2	3	$-1.9 \times 10^{-3} \varkappa + 2.9 \times 10^{-5} Q_n$

B. Hyperfine transitions

Similar to the anapole moment contribution, the neutron quadrupole moment can lead to PNC transition between different hyperfine components of the same state. We found that to have a nonzero contribution of the weak quadrupole moment the minimal value of the electron angular momentum in a single-valence-electron atom is J=3/2.

This means that the effect is zero in the ground state of all atoms considered above. Therefore, we consider the Bi atom instead for which first measurements of atomic PNC were performed [29]. The results of the calculations are presented

TABLE II. Nuclear anapole and neutron quadrupole contributions to the PNC transition between hfs components of the ground state of 209 Bi (I = 9/2, $Q_W = -118.65$).

F_1	F_2	PNC amplitude $10^{-10}iea_0$
3	4	$-2.0 \times 10^{-4} \varkappa + 2.6 \times 10^{-6} Q_n$
4	5	$-2.4 \times 10^{-4} \varkappa + 7.6 \times 10^{-7} Q_n$
5	6	$-2.1 \times 10^{-4} \varkappa - 1.8 \times 10^{-6} Q_n$

in Table II. Using estimations $Q_n \approx (N/Z)Q_p \approx -0.9b$, and $\varkappa(\text{Bi}) \sim 0.1$ [11] we see that the neutron quadrupole contribution is only about one order of magnitude smaller than the anapole contribution.

C. Close levels of opposite parity in lanthanoids

As we discussed above the quadrupole PNC contributions is at least four orders of magnitude smaller than the scalar one. This makes it hard to measure and one should look for enhancement factors. Strong enhancement can take place when a pair of states of opposite parity is separated by a small energy interval. Such pairs can be found in lanthanoid atoms. For the atoms considered above the typical energy denominator [see formula (4)] is $\sim 10\,000$ cm⁻¹. Therefore, for the quadrupole contribution being similar in value with the scalar one we need to look for energy intervals between states of opposite parity, ~ 1 cm⁻¹. We shall consider close states with the difference in the value of the total angular momentum $\Delta J = 1,2$. The opposite parity states with $\Delta J = 2$ can only be mixed by weak quadrupole making it the only contribution to the PNC amplitude. This is a clear case for the weak quadrupole study. In contrast, the states with $\Delta J = 1$ can be mixed by both weak quadrupole and nuclear anapole. We consider these two cases separately.

1. Close states with $\Delta J = 2$

Table III shows some examples of the pairs of states for lanthanoid atoms separated by energy interval $\Delta E \leq 10~\mathrm{cm^{-1}}$ and having the values of the total angular momentum J which differ by 2. The data have been obtained by analyzing the NIST databases [30]. We include only states which seem to be promising for the study of the PNC caused by neutron quadrupole moment. We excluded atoms where all stable isotopes have small nuclear spin (I < 1) and thus no quadrupole moment. We excluded highly excited states and pairs of close states if an electron configuration for at least one state is not known.

Neither scalar nor anapole PNC interactions can mix the states with $\Delta J=2$. The weak quadrupole is the only contribution to the PNC involving the states. This makes them good candidates for the study of the neutron quadrupole moments. If one of the states is connected to the ground state by an electric dipole transition (E1) one can study the interference between Stark-induced and PNC-induced amplitudes of the transition to the ground state similar to what was measured in Cs [3]. Otherwise, one can study the interference between the hyperfine or Stark-induced and the PNC-induced amplitudes

TABLE III. Ground states and pairs of close states of opposite parity with $\Delta J = 2$ in some lanthanoid atoms and corresponding PNC amplitudes.

Atom/ $Q_p(b)$	States		J	E (cm ⁻¹)	τ	ΔE (cm ⁻¹)	h_Q m.e.	<i>E</i> _{PNC} (a.u.)
¹⁴³ Nd	4 f 4.6s2	5 <i>I</i>	4.0	0.0				
-0.61(2)	$4f3.(4I^*).5d2.(3F) (6L^*).6s$	$7L^*$	5.0	11108.813	$1 \mu s$		$\langle 4f6s h_Q 5d^2\rangle$	1.1×10^{-13}
0.01(2)	4 f 4.6s2	5 <i>F</i>	3.0	11118.466	338 μ s	9.652	(1) 00/110/04	111 // 10
	4f4.(5I).5d.6s.(3D)	7G	2.0	11990.020	273 μs	,,,,,	$\langle 4f h_O 6s\rangle$	5.4×10^{-13}
	4f3.(4I*).5d.6s2	*	4.0	11992.388	$1 \mu s$	2.368	() () ()	
¹⁵⁵ Gd	4 f 7.(8S*).5d.6s2	$9D^{*}$	2.0	0.0	•			
1.27(3)	$4f7.(8S^*).5d2.(3P)(10P^*).6s$	$9P^*$	3.0	15173.639	407 ns		$\langle 5d h_O 6p\rangle$	3.9×10^{-11}
(-)	$4f7.(8S^*).5d (9D^*).6s.6p.(3P^*)$	11 <i>F</i>	5.0	15174.000	$2 \mu s$	0.361	() 2 1 7	
	$4f7.(8S^*).5d2.(1D)(8D^*).6s$	$9D^{*}$	6.0	17906.736	655 ns		$\langle 5d h_Q 6p\rangle$	4.3×10^{-12}
	$4f7.(8S^*).5d~(9D^*).6s.6p.(3P^*)$	11 <i>F</i>	8.0	17909.943	$7 \mu s$	3.207		
¹⁶¹ Dy	4 f 10.6s2	5 <i>I</i>	8.0	0.0	·			
2.51(2)	4f10.(5I(8)).5d.6s.(3D)	3[9]	10.0	18462.650	$11 \mu s$		$\langle 4f h_Q 5d\rangle$	2.0×10^{-12}
	$4f9.(6H^*).5d2.(3F) (8G^*).6s$	$9G^*$	8.0	18472.711	819 ns	10.061		
¹⁶⁵ Ho	4 f 11.6s2	$4I^*$	7.5	0.0				
3.58(2)	$4f11.(4I^*).5d.6s.(3D)$	*	6.5	20493.770	771 ns		$\langle 4f h_O 5d\rangle$	5.8×10^{-12}
	4f 10.(5I).5d2.(3F) (7H).6s	8H	8.5	20498.730	$1 \mu s$	4.961		
	$4f11.(4I^*\langle 13/2\rangle).6s.6p.(3P^*\langle 1\rangle)$	(13/2,1)	6.5	22157.859	512 ns		$\langle 6p h_O 5d\rangle$	1.9×10^{-9}
	$4f11.(4I^*).5d.6s.(3D)$	*	4.5	22157.881	$3 \mu s$	0.021	~	
¹⁶⁷ Er	4f12.6s2	3H	6.0	0.0				
3.57(3)	4f11.(4I*).5d.6s.6p		7.0	25861.232	567 ns		$\langle 6p h_Q 5d\rangle$	1.8×10^{-11}
	$4f11.(4I^*).5d2.(3P) (6I^*).6s$	$7I^*$	9.0	25863.453	507 ns	2.221	~	

of the transition between these two states similar to what was done for Dy [31]. In the latter case one needs metastable states. Therefore, we performed estimations of the lifetimes of each state in the table. The estimations are approximate. We consider only E1 transitions, using experimental energies and assuming that all E1 amplitudes are equal to 1 a.u. The results are presented in Table III.

The E^{PNC} amplitude is estimated using the formula

$$E_{ag}^{\mathrm{PNC}} \sim c_0 \frac{\langle a|h_Q|b\rangle\langle b|D|g\rangle}{\Delta E} Q_n.$$
 (7)

Here a and b is a pair of the close-energy states, state g is the ground state, D is an operator of the electric dipole transition (E1), c_0 is angular coefficient [see formula (A3)], Q_n is the neutron quadrupole moment. For the estimations we assume $\langle b|D|c\rangle=1$ a.u., $c_0=0.1$, $Q_n=(N/Z)Q_p$. The values of the electric quadrupole moment Q_p are taken from Ref. [28].

Estimations of the $\langle a|h_Q|b\rangle$ matrix elements are more complicated. Calculations show that all of them apart from only the s- $p_{3/2}$ matrix elements are very sensitive to manybody effects. This is a well-known feature of any short-range interaction of atomic electrons with the nucleus. The wave functions of states with angular momentum l>1 are negligibly small on the nucleus and s states of other electrons must come into play via many-body effects to make a dominant contribution. Table III indicates that we need to deal with the s-f, p-d, and d-f weak matrix elements which are sensitive to many-body effects. Table IV shows the values of the weak matrix elements calculated in the relativistic Hartree-Fock (RHF) and RPA approximations (we use Gd atom as an example). Taking into account the core polarization via the RPA calculations increases the value of most matrix elements

by many orders of magnitude. Further increase can be found if the configuration mixing is taken into account. Configuration mixing brings into play configurations which make possible the 6s- $np_{3/2}$ contribution to the weak matrix element. Sample diagrams are presented in Fig. 1. Note that the configuration mixing is due to the Coulomb interaction. Therefore, we call corresponding corrections to the weak matrix elements the Coulomb corrections. For example, the Coulomb corrections to the $\langle 5d|h_Q|6p\rangle$ and $\langle 5d|h_Q|4f\rangle$ matrix elements in Gd, Ho, and Er are given by the diagram in Fig. 1(a) and the correction to the $\langle 6s|h_Q|4f\rangle$ matrix element is given by the diagram in Fig. 1(b). Note that the weak matrix element between first pair of states in Nb is zero in the single-electron approximation

TABLE IV. Matrix elements of the neutron quadrupole operator h_Q and Coulomb corrections to them (a.u.). RHF stands for relativistic Hartree-Fock, RPA is the random-phase approximation. Numbers in square brackets stand for powers of 10.

	$\langle a H$	$ Q b\rangle$	
Transition	RHF	RPA	$\langle \tilde{6s}, a \frac{e^2}{ r_1 - r_2 } 6s, b \rangle$
$4f_{5/2}$ - $6s_{1/2}$	2.24[-19]	2.50[-19]	7.78[-17]
$4f_{5/2}-5d_{3/2}$	5.83[-25]	2.95[-19]	2.14[-16]
$4f_{5/2}-5d_{5/2}$	-1.89[-25]	2.54[-18]	-5.50[-17]
$4f_{7/2}-5d_{5/2}$	5.57[-30]	4.60[-18]	2.54[-16]
$6p_{3/2}-6s_{1/2}$	2.11[-16]	5.62[-16]	4.40[-16]
$6p_{1/2}-5d_{3/2}$	5.00[-18]	1.46[-17]	-3.46[-16]
$6p_{3/2}$ – $5d_{3/2}$	4.30[-19]	-1.72[-17]	1.39[-16]
$6p_{3/2}$ - $5d_{5/2}$	4.27[-23]	-4.15[-17]	4.27[-16]

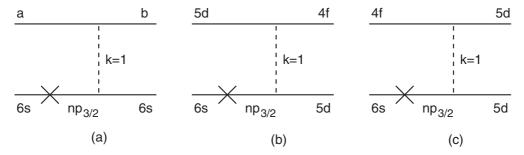


FIG. 1. Coulomb corrections to the weak matrix elements. Cross stands for the weak quadrupole interaction; summation over complete set of $np_{3/2}$ states is assumed. (a) The 6p-5d or 5d-4f matrix elements. (b) The 6s-4f matrix element. (c) The two-electron matrix element between the 4f6s and $5d^2$ states (e.g., in Nb).

since the states differ by two electron orbitals. In this case the diagram in Fig. 1(c) is the lowest-order contribution.

We estimate the diagrams (Fig. 1) by calculating Coulomb integrals in which one 6s wave function is replaced by a correction induced by the h_q operator. The correction is calculated in the RPA approximation

$$(H^{\rm RHF} - \epsilon_{6s})\delta\psi_{6s} = -(h_Q + \delta V^{\rm RHF})\psi_{6s}. \tag{8}$$

Corresponding Coulomb integrals are

$$\langle \tilde{6s}, a | r_{<}/r_{>}^{2} | 6s, b \rangle$$
 Fig. 1(a),
 $\langle \tilde{6s}, 5d | r_{<}/r_{>}^{2} | 5d, 4f \rangle$ Fig. 1(b),
 $\langle \tilde{6s}, 4f | r_{<}/r_{>}^{2} | 5d, 5d \rangle$ Fig. 1(c).

Here $r_{<} = \min(r_1, r_2)$, $r_{>} = \max(r_1, r_2)$, and $|\delta s\rangle \equiv \delta \psi_{6s}$. Calculated values of the Coulomb integrals are presented in the last column of Table IV. Substituting these numbers into (7) we get estimations for the PNC amplitudes. The results are presented in Table III. Note that in contrast to the amplitudes considered in Secs. III A and III B the amplitudes here are relatively large. In most cases they are larger than the PNC amplitude in Cs [3]. In the case of the second pair of close states in Ho, the amplitude is as large as in Yb, the largest PNC atomic amplitude which has been measured so far [15].

In all cases considered above one can measure the transition rate between the two close states of opposite parity and study the interference between the PNC amplitude (7) and the electric dipole transition induced by the hyperfine interaction. In addition, when one of the states is connected to the ground state by the magnetic dipole (M1) transition (first pairs of states in Nd, Gd, Ho, and Er) or an electric quadrupole (E2) transition (second pair of states in Nd) one can study the interference between these M1 or E2 amplitudes and the PNC amplitude (7) to the ground state.

2. Close states with $\Delta J = 1$

Close states of opposite parity with $\Delta J=1$ are also important. Here both the anapole moment and the weak quadrupole moments contribute to the PNC effect. Measuring both these moments are equally important. The anapole moment has been measured for Cs only [3]. The limit on

the anapole moment of Tl obtained in the PNC measurements [12] has also been obtained. Measuring more anapole moments may help to extract constants of the weak interaction between nucleons and to get better understanding of nuclear structure. Measuring the PNC effect which has both anapole and quadrupole contributions may have some advantages. The effect is expected to be larger while different dependence of two contributions on the quantum numbers (e.g., on total angular momentum F, $\mathbf{F} = \mathbf{J} + \mathbf{I}$) allows one to separate the contributions.

Table V shows pairs of opposite parity states of lanthanoids separated by the energy interval $\Delta E < 10 \, \mathrm{cm}^{-1}$ with values of the total angular momentum J which differ by 1. The pairs have been found by analyzing the NIST database [30]. We also included Ra which was studied in Ref. [32].

It is clear that many of the systems listed in Table V are as good as those considered in the previous section. Estimations can be also done in a similar way. The most important parameters defining the value of the PNC amplitude are the energy interval between states of opposite parity and the type of the weak matrix element. The values for different types of weak matrix elements are presented in Table IV. The energy intervals are presented in Table V. More detailed study of the PNC amplitudes for all systems listed in Table V goes beyond the scope of present work. The analysis can be done for a particular system which is of the greatest interest to experimentalists. In our view there are many systems which look very promising but require careful consideration from the experimental point of view.

IV. CONCLUSION

We argue that the measuring PNC in atoms can be used to study the neutron distribution in nuclei via measuring the parity-nonconserving weak quadrupole moment. The effect is small in atoms which have been already used to study PNC. However, a strong enhancement due to close states of opposite parity can be found in lanthanoids and in Ra. Here the neutron quadrupole moments can be studied together with the nuclear anapole moments. There are many systems where the weak quadrupole moment is the only enhanced contribution to the PNC effect. The enhancement is sufficiently strong to make the prospects of the measurements very realistic.

TABLE V. Ground states and pairs of close states of opposite parity with $\Delta J = 1$ in some lanthanoid atoms and Ra.

Atom	States		J	$E \text{ (cm}^{-1})$	τ	$\Delta E (\mathrm{cm}^{-1})$
Pr	4 <i>f</i> 3.6 <i>s</i> 2	41*	4.5	0.0		
Pr	4f2.(3H).5d.6s2	4H	5.5	9675.010	$6 \mu \mathrm{s}$	
	$4f3.(4I^*).5d.6s.(3D)$	$6K^*$	6.5	9684.240	$5 \mu s$	9.230
Pr			6.5	10423.680	$306 \mu s$	
	$4f3.(4I^*).5d.6s.(3D)$	$4K^*$	5.5	10431.750	$3 \mu s$	8.070
Pr	4f2.(3H).6s2.6p	$4I^*$	4.5	19339.859	113 ns	
			5.5	19343.250	356 ns	3.391
Nd	4f4.6s2	5 <i>I</i>	4.0	0.0		
Nd	$4f3.(4I^*).5d2.(3F)(6L^*).6s$	$7L^*$	5.0	11108.813	$1 \mu s$	
	4f4.(5I).5d.6s.(3D)	7 <i>K</i>	6.0	11109.167	29 μs	0.354
Nd	4f4.(5I).5d.6s.(3D)	7I	6.0	12917.422	$7 \mu \mathrm{s}$	
	4 f 3.(4I*).5d.6s2	$5I^*$	7.0	12927.232	$2 \mu s$	9.811
Sm	4f6.6s2	7F	0.0	0.0	,	
Sm	$4f6.(7F).6s.6p.(3P^*)$	$9G^{*}$	5.0	16344.770	753 ns	
	4f6.(7F).5d(8D).6s	7D	4.0	16354.600	1 ms	9.830
Gd	$4f7.(8S^*).5d.6s2$	$9D^*$	2.0	0.0		
Gd	$4f7.(8S^*).5d(9D^*).6s.6p.(3P^*)$	7D	3.0	19399.840	252 ns	
	$4f7.(8S^*).5d2.(1G)(8G^*).6s$	$9G^*$	2.0	19403.104	164 ns	3.264
Gd	$4f7.(8S^*).5d2.(3F)(6F^*).6s$	$5F^*$	3.0	20299.869	121 ns	
			2.0	20303.801	169 ns	3.932
Tb	4 f 9.6s2	$6H^*$	7.5	0.0		
Tb	$4f8.(7F\langle 6\rangle).6s2.6p\langle 1/2\rangle$	$(6,1/2)^*$	5.5	13616.270	430 ns	
		, ,	6.5	13622.690	$1 \mu s$	6.421
Dy	4 f 10.6s2	5 <i>I</i>	8.0	0.0	,	
Dy	$4f10.(5I\langle 8\rangle).6s.6p.(3P^*\langle 2\rangle)$	$(8,2)^*$	10.0	17513.330	metastable	
3	4f10.(5I(8)).5d.6s.(3D)	3[8]	9.0	17514.500	$5 \mu s$	1.170
Dy	$4f9.(6H^*).5d2.(3P)(8I^*).6s$	*	9.0	23271.740	697 ns	
3	$4f10.(5I\langle 7\rangle).5d.6s.(3D)$		8.0	23280.461	556 ns	8.721
Dy	4f10.(5I(7)).5d.6s.(3D)		6.0	23333.920	526 ns	
,	$4f9.(6H^*).5d2.(3F)(8G^*).6s$	*	7.0	23340.119	263 ns	6.199
Dy	$4f9.(6H^*).5d2.(3F)(8F^*).6s$	*	6.0	23359.820	359 ns	
3	$4f10.(5I\langle 7\rangle).5d.6s.(3D)$		7.0	23360.660	435 ns	0.840
Но	4 f 11.6s2	$4I^*$	7.5	0.0		
Но	$4f10.(5I\langle 6 \rangle).5d\langle 3/2 \rangle.6s2$	(6,3/2)	6.5	18564.900	935 ns	
	$4f10.(5I\langle 8\rangle).6s2.6p\langle 1/2\rangle$	(8,1/2)	7.5	18572.279	$1 \mu s$	7.3
Ra	7s2	18	0.0	0.0	r	
Ra	7s6d	3D	2.0	13993.94	metastable	
· -	7s7 p	$3P^*$	1.0	13999.3569	500 ns	5.42

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APPENDIX: MATRIX ELEMENTS

The projection M dependence of the amplitude can be factorized by using the Wigner-Eckart theorem:

$$E_{\text{PNC}}^{i \to f} = (-1)^{F_f - M_f} \begin{pmatrix} F_f & 1 & F_i \\ -M_f & q & M_i \end{pmatrix} \langle J_f F_f \| d_{\text{PNC}} \| J_i F_i \rangle.$$
(A1)

By means of the standard angular momentum technique, the matrix element of h_Q between the hyperfine states $|(JI)FM\rangle$

and $|(J'I)F'M'\rangle$ can be written as a product of the reduced matrix elements of the electronic part and the nuclear part of the interaction:

$$\langle (J'I)F'M'|h_O|(JI)FM\rangle \propto \delta_{F'F}\delta_{M'M}(-1)^{F+J+I}$$

$$\times \left\{ \begin{matrix} J' & J & 2 \\ I & I & F \end{matrix} \right\} \langle J' \| \boldsymbol{T} \| J \rangle \langle I \| \boldsymbol{Q}^{\mathrm{TW}} \| I \rangle. \tag{A2}$$

The formula for the reduced matrix element of the PNC amplitude induced by the weak quadrupole $Q^{\rm TW}$ can be derived similar to the derivation of the nuclear-spin-dependent (SD) PNC amplitude in Refs. [33] and [34].

The result is

$$\langle J_{f}F_{f}\|d_{Q}\|J_{i}F_{i}\rangle = \sqrt{\frac{(2I+3)(2I+1)(I+1)}{I(2I-1)}}\sqrt{[F_{i}][F_{f}]}\sum_{n}\left[(-1)^{J_{f}-J_{i}}\begin{cases}J_{n} & J_{f} & 1\\F_{f} & F_{i} & I\end{cases}\right]\begin{cases}J_{n} & J_{i} & 2\\I & I & F_{i}\end{cases}$$

$$\times \frac{\langle J_{f}\|\boldsymbol{d}\|nJ_{n}\rangle\langle nJ_{n}\|\boldsymbol{h}_{Q}^{e}\|Ji\rangle}{E_{n}-E_{i}} + (-1)^{F_{f}-F_{i}}\begin{cases}J_{n} & J_{i} & 1\\F_{i} & F_{f} & I\end{cases}\begin{cases}J_{n} & J_{f} & 2\\I & I & F_{f}\end{cases}\frac{\langle J_{f}\|\boldsymbol{h}_{Q}^{e}\|nJ_{n}\rangle\langle nJ_{n}\|\boldsymbol{d}\|Ji\rangle}{E_{n}-E_{f}}\right].$$
(A3)

Here

$$\boldsymbol{h}_{Q}^{e} = -\frac{5G_{F}Q^{\mathrm{TW}}}{4\sqrt{2}\langle r^{2}\rangle}\boldsymbol{C}^{(2)}\gamma_{5}\rho_{0}(r)$$

is the electronic tensor part of the weak interaction and the notation $[F_a] \equiv 2F_a + 1$ is used.

For comparison we present two other contributions to the PNC amplitudes in atoms, namely the nuclear spin independent (SI) weak charge Q_W contribution and the nuclear spin dependent (SD) contribution dominated by the magnetic interaction of atomic electrons with the nuclear anapole moment (AM) [10,11]. They have been measured and calculated in many atomic systems—see, e.g., [3,5–9,12,15,21,35–47]. The reduced matrix elements of SI and SD PNC amplitudes are presented, e.g., in Ref. [48]:

$$\langle J_{f}F_{f}\|d_{SD}\|J_{i}F_{i}\rangle = \sqrt{\frac{(2I+1)(I+1)}{I}}\sqrt{[F_{i}][F_{f}]}\sum_{n}\left[(-1)^{J_{f}-J_{i}}\begin{cases}J_{n} & J_{f} & 1\\F_{f} & F_{i} & I\end{cases}\begin{cases}J_{n} & J_{i} & 1\\I & I & F_{i}\end{cases}\right] \times \frac{\langle J_{f}\|d\|nJ_{n}\rangle\langle nJ_{n}\|\mathbf{h}_{SD}\|Ji\rangle}{E_{n}-E_{i}} + (-1)^{F_{f}-F_{i}}\begin{cases}J_{n} & J_{i} & 1\\F_{i} & F_{f} & I\end{cases}\begin{cases}J_{n} & J_{f} & 1\\I & I & F_{f}\end{cases}\frac{\langle J_{f}\|\mathbf{h}_{SD}\|nJ_{n}\rangle\langle nJ_{n}\|\mathbf{d}\|Ji\rangle}{E_{n}-E_{f}},$$
(A4)

where the vector operator is the electronic part of the SD interaction $h_{\rm SD} = (G_F/\sqrt{2})\kappa\alpha\rho_0(r)$ and the Dirac matrix is defined by

$$\boldsymbol{\alpha} = \begin{pmatrix} 0 & \boldsymbol{\sigma} \\ \boldsymbol{\sigma} & 0 \end{pmatrix}.$$

The dimensionless parameter κ determines the strength of the SD PNC interaction. The three major contributions to κ come from the electromagnetic interaction of the atomic electrons with the nuclear anapole moment [10,11], the electron-nucleus SD weak interaction [46], and the combined effect of the SI weak interaction and the magnetic hyperfine interaction [49]. For the SI PNC reduced amplitude we have

$$\langle J_{f}F_{f}\|d_{\mathrm{SI}}\|J_{i}F_{i}\rangle = (-1)^{I+F_{i}+J_{f}}\sqrt{[F_{i}][F_{f}]} \begin{cases} J_{i} & J_{f} & 1\\ F_{f} & F_{i} & I \end{cases} \sum_{n} \left[\frac{\langle J_{f}\|\boldsymbol{d}\|nJ_{n}\rangle\langle nJ_{n}|H_{\mathrm{SI}}|Ji\rangle}{E_{n}-E_{i}} + \frac{\langle J_{f}|H_{\mathrm{SI}}|nJ_{n}\rangle\langle nJ_{n}\|\boldsymbol{d}\|Ji\rangle}{E_{n}-E_{f}} \right], \tag{A5}$$

where the weak interaction is $H_{\rm SI} = -G_F Q_W/(2\sqrt{2})\gamma_5\rho_0(r)$ and Q_W is the nuclear weak charge. Note that the weak matrix elements $\langle nJ_n|H_{\rm SI}|Ji\rangle$ are not reduced ones in Eq. (A5).

The single-electron orbitals used to calculate the matrix elements are

$$\varphi_{n\kappa m}(\mathbf{r}) = \frac{1}{r} \begin{pmatrix} f_{n\kappa}(r)\Omega_{\kappa m}(\theta, \phi) \\ ig_{n\kappa}(r)\Omega_{-\kappa m}(\theta, \phi) \end{pmatrix}, \tag{A6}$$

where n is the principle quantum number and $\kappa = \mp (j+1/2)$ (for $j=l\pm 1/2$) is the angular quantum number for the Dirac spinor. The relativistic single-particle matrix elements of the PNC operators are

$$\langle \kappa_1 \| \boldsymbol{h}_{Q}^{e} \| \kappa_2 \rangle = i \frac{5G_F Q^{\text{TW}}}{4\sqrt{2}\langle r^2 \rangle} \langle \kappa_1 \| C^{(2)} \| - \kappa_2 \rangle$$

$$\times \int (f_1 g_2 - g_1 f_2) \rho_0 dr \tag{A7}$$

$$\langle \kappa_1 || \mathbf{h}_{SD} || \kappa_2 \rangle = -i \frac{G_F \kappa}{\sqrt{2}} \langle \kappa_1 || C^{(1)} || \kappa_2 \rangle \int [(\kappa_1 - \kappa_2 + 1) \times g_1 f_2 - (\kappa_2 - \kappa_1 + 1) f_1 g_2) \rho_0 dr$$
(A8)

$$\langle \kappa_1 | H_{\rm SI} | \kappa_2 \rangle = i \frac{G_F Q_W}{2\sqrt{2}} \delta_{-\kappa_1, \kappa_2} \int (f_1 g_2 - g_1 f_2) \rho_0 dr. \quad (A9)$$

Note that all weak matrix elements have imaginary values. The reduced matrix element of $C^{(k)}$ is given by

$$\langle \kappa_1 || \mathbf{C}^{(k)} || \kappa_2 \rangle = (-1)^{j_2 + 1/2} \sqrt{[j_1][j_2]} \times \xi(l_1 + l_2 + k) \begin{pmatrix} j_2 & j_1 & k \\ -1/2 & 1/2 & 0 \end{pmatrix}, \tag{A10}$$

where $\xi(L) = 1$ if L is an even number, otherwise it is 0.

- [1] I. B. Khriplovich, *Parity Nonconservation in Atomic Phenomena* (Gordon and Breach, New York, 1991).
- [2] J. S. M. Ginges and V. V. Flambaum, Phys. Rep. 397, 63 (2004).
- [3] C. S. Wood, S. C. Bennett, D. Cho, B. P. Masterson, J. L. Roberts,C. E. Tanner, and C. E. Wieman, Science 275, 1759 (1997).
- [4] V. A. Dzuba, V. V. Flambaum, and O. P. Sushkov, Phys. Lett. A 141, 147 (1989).
- [5] S. A. Blundell, W. R. Johnson, and J. Sapirstein, Phys. Rev. Lett. 65, 1411 (1990).
- [6] M. G. Kozlov, S. G. Porsev, and I. I. Tupitsyn, Phys. Rev. Lett. 86, 3260 (2001).
- [7] V. A. Dzuba, V. V. Flambaum, and J. S. M. Ginges, Phys. Rev. D 66, 076013 (2002).
- [8] S. G. Porsev, K. Beloy, and A. Derevianko, Phys. Rev. Lett. 102, 181601 (2009).
- [9] V. A. Dzuba, J. C. Berengut, V. V. Flambaum, and B. M. Roberts, Phys. Rev. Lett. 109, 203003 (2012).
- [10] V. V. Flambaum and I. B. Khriplovich, Sov. Phys. JETP 52, 835 (1980) [Zh. Eksp. Teor. Fiz. 79, 1656 1980].
- [11] V. V. Flambaum, I. B. Khriplovich, and O. P. Sushkov, Phys. Lett. B 146, 367 (1984).
- [12] P. A. Vetter, D. M. Meekhof, P. K. Majumder, S. K. Lamoreaux, and E. N. Fortson, Phys. Rev. Lett. 74, 2658 (1995).
- [13] V. V. Flambaum and D. W. Murray, Phys. Rev. C 56, 1641 (1997).
- [14] V. A. Dzuba, V. V. Flambaum, and I. B. Khriplovich, Z. Phys. D 1, 243 (1986).
- [15] K. Tsigutkin, D. R. Dounas-Frazer, A. Family, J. E. Stalnaker, V. V. Yashchuk, and D. Budker, Phys. Rev. Lett. 103, 071601 (2009).
- [16] B. A. Brown, A. Derevianko, and V. V. Flambaum, Phys. Rev. C 79, 035501 (2009).
- [17] O. P. Sushkov and V. V. Flambaum, Sov. Phys. JETP 48, 453 (1978) [Zh. Eksp. Teor. Fiz. 75, 1208 (1978)].
- [18] I. B. Khriplovich and M. E. Pospelov, Z. Phys. D 22, 367 (1991).
- [19] V. V. Flambaum, Phys. Rev. Lett. 117, 072501 (2016).
- [20] V. V. Flambaum, D. DeMille, and M. G. Kozlov, Phys. Rev. Lett. **113**, 103003 (2014).
- [21] V. A. Dzuba and V. V. Flambaum, Phys. Rev. A 83, 052513 (2011).
- [22] M. Tandecki, J. Zhang, R. Collister et al., J. Instrum. 8, P12006 (2013).
- [23] E. Mariotti, A. Khanbekyan, C. Marinelli *et al.*, Int. J. Mod. Phys. E 23, 1430009 (2014).
- [24] J. Choi and D. S. Elliott, Phys. Rev. A 93, 023432 (2016).
- [25] B. K. Sahoo and B. P. Das, Phys. Rev. A 92, 052511 (2015).
- [26] L. W. Wansbeek, S. Schlesser, B. K. Sahoo, A. E. L. Dieperink, C. J. G. Onderwater, and R. G. E. Timmermans, Phys. Rev. C 86, 015503 (2012).

- [27] M. Nunez Portela, E. A. Dijck, A. Mohanty *et al.*, Appl. Phys. B **114**, 173 (2014).
- [28] N. J. Stone, At. Data Nucl. Data Tables 111-112, 1 (2016).
- [29] L. M. Barkov and M. S. Zolotorev, Phys. Lett. B 85, 308 (1979).
- [30] A. Kramida, Yu. Ralchenko, J. Reader, and NIST ASD Team (2013). NIST Atomic Spectra Database (ver. 5.1) (online). Available: http://physics.nist.gov/asd [2013, December 12]. National Institute of Standards and Technology, Gaithersburg, MD.
- [31] A. T. Nguyen, D. Budker, D. DeMille, and M. Zolotorev, Phys. Rev. A 56, 3453 (1997).
- [32] V. A. Dzuba, V. V. Flambaum, and J. S. M. Ginges, Phys. Rev. A 61, 062509 (2000).
- [33] S. G. Porsev and M. G. Kozlov, Phys. Rev. A 64, 064101 (2001).
- [34] W. R. Johnson, M. S. Safronova, and U. I. Safronova, Phys. Rev. A 67, 062106 (2003).
- [35] M. J. D. Macpherson, K. P. Zetie, R. B. Warrington, D. N. Stacey, and J. P. Hoare, Phys. Rev. Lett. 67, 2784 (1991).
- [36] R. B. Warrington, C. D. Thompson, and D. N. Stacey, Europhys. Lett. 24, 641 (1993).
- [37] D. M. Meekhof, P. A. Vetter, P. K. Majumder, S. K. Lamoreaux, and E. N. Fortson, Phys. Rev. Lett. 71, 3442 (1993).
- [38] S. J. Phipp, N. H. Edwards, P. E. G. Baird, and S. Nakayama, J. Phys. B 29, 1861 (1996).
- [39] N. H. Edwards, S. J. Phipp, P. E. G. Baird, and S. Nakayama, Phys. Rev. Lett. 74, 2654 (1995).
- [40] J. Guéna, D. Chauvat, P. Jacquier, E. Jahier, M. Lintz, S. Sanguinetti, A. Wasan, M.-A. Bouchiat, A. V. Papoyan, and D. Sarkisyan, Phys. Rev. Lett. 90, 143001 (2003).
- [41] J. Guéna, M. Lintz, and M.-A. Bouchiat, Phys. Rev. A 71, 042108 (2005).
- [42] V. A. Dzuba, V. V. Flambaum, P. G. Silvestrov, and O. P. Sushkov, Europhys. Lett. 7, 413 (1988).
- [43] V. A. Dzuba, V. V. Flambaum, P. G. Silvestrov, and O. P. Sushkov, J. Phys. B 20, 3297 (1987).
- [44] M. G. Kozlov, S. G. Porsev, and W. R. Johnson, Phys. Rev. A 64, 052107 (2001).
- [45] V. A. Dzuba and V. V. Flambaum, Phys. Rev. A **83**, 042514
- [46] V. N. Novikov, O. P. Sushkov, V. V. Flambaum, and I. B. Khriplovich, Sov. Phys. JETP 46, 420 (1977) [Zh. Eksp. Teor. Fiz. 73, 802 (1977)].
- [47] V. N. Novikov and I. B. Khriplovich, JETP Lett. 22, 74 (1975)
 [Pis'ma Zh. Eksp. Teor. Fiz. 22, 162 (1975)].
- [48] V. A. Dzuba, V. V. Flambaum, and C. Harabati, Phys. Rev. A 84, 052108 (2011).
- [49] V. V. Flambaum and I. B. Khriplovich, Sov. Phys. JETP 62, 872 (1985) [Zh. Eksp. Teor. Fiz. 89, 1505 (1985)].