

# Autoionization resonances in the neon isoelectronic sequence using relativistic multichannel quantum-defect theory

M. Nrisimhamurty,<sup>1</sup> G. Aravind,<sup>1</sup> P. C. Deshmukh,<sup>1,\*</sup> and S. T. Manson<sup>2</sup>

<sup>1</sup>*Department of Physics, Indian Institute of Technology Madras, Chennai 600036, India*

<sup>2</sup>*Department of Physics and Astronomy, Georgia State University, Atlanta, Georgia 30303, USA*

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Studies of  $2s \rightarrow np$  autoionization resonances in the neon isoelectronic sequence using relativistic multichannel quantum-defect theory are reported. The relativistic random-phase approximation is used to calculate the quantum-defect parameters. The autoionization resonances are characterized using Fano resonance parameters. The results are compared with available experimental and theoretical data, and the behavior of the resonances as a function of  $Z$  is discussed.

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## I. INTRODUCTION

Qualitative and quantitative knowledge of various observables resulting from the photoabsorption process is an important source of data for various applications, such as modeling of the interstellar and intergalactic media and development of x-ray free-electron lasers. The recent upsurge of interest in atomic photoabsorption [1–4] requires accurate atomic spectroscopic data. Quantum-mechanical interference between bound-continuum channels and bound-bound channels leads to autoionization resonances [5–7]. Generally, autoionization is described as a two-step process that occurs when an excited atom or ion consisting of many electrons decays nonradiatively, with the excited electron filling in a hole in a deeper bound state, enabling one of the other electrons bound with energy *less* than the excitation energy to escape into the unbound continuum. Autoionization is thus essentially a consequence of electron correlation and requires a many-body theory for its calculation. Studies of the autoionization resonances in noble-gas atoms [5] have long attracted experimental and theoretical scrutiny. The significant cosmic abundance of neonlike highly charged ions (HCI) and other similar atomic or ionic many-electron systems has drawn special attention to their photoionization, with particular emphasis upon the asymmetric line profiles exhibited by the autoionization resonances [8–31]. Motivated by the astrophysical importance of HCIs and by the anomalous behavior in Ne-like silicon and argon [21–23], the present work focuses primarily on the  $2s \rightarrow np$  autoionizing resonance transitions in the neon isoelectronic sequence.

The relativistic random-phase approximation (RRPA) [32], which is a *whole-space* correlation theory, is employed to calculate the quantum-defect parameters for the neutral neon atom and various isoelectronic ions. We then use the relativistic multichannel quantum-defect theory (RMQDT) [33], a relativistic *partitioned-space* correlation theory based on Seaton's formalism [34,35], to study the autoionization resonances. The resonances are then characterized by their energies, widths, and shape profiles in the form of the Fano parameters [4,6].

## II. METHODOLOGY

The RRPA builds electron correlations by carrying out a linearization of the time-dependent Dirac-Hartree-Fock (DHF) formalism. Initial-state correlations are included via the time-backward ring (and corresponding exchange) diagrams, and the correlations in the final state are included via interchannel coupling corresponding to the time-forward ring (and corresponding exchange) diagrams. In the *full RRPA*, all allowed relativistic dipole channels are coupled. However, in the present work, we have coupled only the important channels in a *truncated RRPA*. In particular, for the present study, we have included interchannel coupling among only the following seven channels:

$$\begin{aligned} 2p_{3/2} &\rightarrow kd_{5/2}, kd_{3/2}, ks_{1/2}, \\ 2p_{1/2} &\rightarrow kd_{3/2}, ks_{1/2}, \\ 2s_{1/2} &\rightarrow kp_{3/2}, kp_{1/2}, \end{aligned}$$

where  $k = n$  for discrete bound-bound channels and  $k = \epsilon$  for the bound-continuum channels. In principle, the truncation causes loss of the gauge invariance of the dipole matrix element, resulting in a disagreement between length and velocity forms. Nevertheless, the disagreement between the two forms is rarely any worse than 5% and is usually much less since the  $1s$  photoionization channels are so far away energetically from the energy region of interest. Hence, in the present work, we have presented the results in the length form alone.

For linearly polarized incident light, the differential cross section for an atomic subshell with quantum numbers  $n, \kappa$  is given by [32]

$$\frac{d\sigma_{n\kappa}}{d\Omega} = \frac{\sigma_{n\kappa}(\omega)}{4\pi} [1 + \beta_{n\kappa}(\omega)P_2(\cos\theta)], \quad (1)$$

where  $\beta_{n\kappa}$  is the angular distribution asymmetry parameter,  $P_2$  is a Legendre polynomial,  $\theta$  is the angle between the photon polarization and the photoelectron momentum, and  $\sigma_{n\kappa}$  is the total (angle-integrated) cross section for photoionization from the subshell ( $n\kappa$ ) and is written in terms of dipole matrix elements for the  $nj \rightarrow j'$  transitions as

$$\sigma_{n\kappa} = \frac{4\pi^2\alpha\omega}{3} (|D_{nj \rightarrow j-1}|^2 + |D_{nj \rightarrow j}|^2 + |D_{nj \rightarrow j+1}|^2), \quad (2)$$

\*pcd@physics.iitm.ac.in

with the matrix elements, calculated within the RRPA framework, given by

$$D_{nj \rightarrow \bar{j}} = i^{1-l} e^{i\delta_{\bar{k}}} \langle \bar{k} || Q_1^{(1)} || \kappa \rangle_{RRPA} \quad (3)$$

in terms of the reduced matrix elements  $\langle || || \rangle$  and phase shifts  $\delta_{\bar{k}}$ . The RMQDT parameters [33], such as the eigen-dipole amplitudes  $D_\alpha$ , eigen-quantum defects  $\mu_\alpha$ , and the frame-transformation matrix elements  $U_{i\alpha}$  depend only very weakly on the energy, thereby allowing interpolation of their values in the region of the autoionization resonances from values just outside this region calculated using RRPA. The autoionization resonances are then computed with RMQDT using these parameters. The values obtained using the RRPA and RMQDT methodologies, which are reported here, have 5% uncertainty.

### III. RESULTS AND DISCUSSION

The DHF threshold energies used in these calculations for various ions and the neutral neon atom, along with other theoretical and experimental energies, are given in the Table I. Since electron correlations are not included in the DHF methodology, the DHF threshold energies are slightly higher than the experimental energies. Note, however, that it is not absolute energies we seek, but rather resonance energies with respect to the relevant thresholds, essentially, quantum defects, and these will be seen to be much more accurate. However, to illustrate, the absolute energies are listed and compared with experiment and other calculations in Table I.

To set the stage for a discussion of the resonances, the RRPA nonresonant photoionization cross sections, which include coupling among the seven relativistic dipole channels described above, for atomic Ne, Na<sup>+</sup>, Mg<sup>2+</sup>, Al<sup>3+</sup>, Si<sup>4+</sup>, and Ar<sup>8+</sup> above their respective  $2p_{1/2}$  thresholds are presented in Fig. 1(a). Note that the calculations have been carried out for a much larger set of ions, and only an illustrative subset is shown in detail. Note also that some of these cross sections have been presented earlier [21] but are shown here for completeness. In any case, the cross sections shown are the sum of the partial cross sections from the  $2p_{3/2}$  and the  $2p_{1/2}$  subshells; the autoionizing region between  $2p_{3/2}$  and the  $2p_{1/2}$  of each cross section is omitted for clarity. The cross section for neutral Ne shows a *delayed maximum* at about  $\sim 38$  eV, in agreement with earlier work [21,41–44]; it occurs owing to the centrifugal potential barrier faced by the final-state continuum  $d$  waves, which overwhelm the attractive potential of the Ne<sup>+</sup> ion. This *delayed maximum* appears even for the Na<sup>+</sup> ion but is not as pronounced as in the case of the neutral Ne atom because of the stronger attractive nature of the Ne<sup>2+</sup> ion in the final state. With increasing ionicity in the isoelectronic sequence, this *delay* disappears, as expected, since the ionic potential becomes strong enough to pull the continuum  $d$  wave in, even at threshold. In Si<sup>4+</sup>, the cross section, however, increases rapidly from threshold from near zero and then decreases; this is due to a  $2s \rightarrow 3p$  resonance just at threshold which was discussed (and corroborated experimentally) earlier [21]. By Ar<sup>8+</sup>, the hint of an increase in the cross section at threshold is not the remnants of the shape resonance, but rather the effect of the nearby  $2s \rightarrow 4p$  resonance that lies slightly below threshold.

TABLE I. Thresholds of outer subshells (eV). E = experiment; T = theory.

Atom or ion	Thresholds of different subshells (eV)		
	$2p_{3/2}$	$2p_{1/2}$	$2s$
Ne	23.083 <sup>a</sup>	23.207 <sup>a</sup>	52.677 <sup>a</sup>
	21.566 <sup>b</sup>	21.663 <sup>b</sup>	48.475(E) <sup>c</sup>
	21.602 <sup>c</sup>	21.699 <sup>c</sup>	48.029(T) <sup>e</sup>
Na <sup>+</sup>	48.818 <sup>a</sup>	49.020 <sup>a</sup>	83.877 <sup>a</sup>
	47.287 <sup>b</sup>	47.465 <sup>b</sup>	80.073 <sup>f</sup>
	47.285 <sup>f</sup>	47.455 <sup>f</sup>	
Mg <sup>2+</sup>	81.682 <sup>a</sup>	81.998 <sup>a</sup>	122.356 <sup>a</sup>
	80.144 <sup>b</sup>	80.410 <sup>b</sup>	
Al <sup>3+</sup>	121.538 <sup>a</sup>	122.013 <sup>a</sup>	167.931 <sup>a</sup>
	119.9924(19) <sup>g</sup>	120.5249(19) <sup>g</sup>	167.0470(19) <sup>g</sup>
	119.9924(19) <sup>h</sup>	120.4192(19) <sup>h</sup>	164.4799(19) <sup>h</sup>
Si <sup>4+</sup>	168.318 <sup>a</sup>	169.01 <sup>a</sup>	220.526 <sup>a</sup>
	166.77 <sup>h</sup>	167.4 <sup>h</sup>	
	166.8(2) <sup>i</sup>	167.2(3) <sup>i</sup>	
P <sup>5+</sup>	166.7 <sup>j</sup>	167.2 <sup>j</sup>	
	221.987 <sup>a</sup>	222.962 <sup>a</sup>	280.106 <sup>a</sup>
Ar <sup>8+</sup>	424.142 <sup>a</sup>	426.51 <sup>a</sup>	500.675 <sup>a</sup>
	420.197 <sup>k</sup>	422.466 <sup>k</sup>	497.39 <sup>k</sup>
	422.54 <sup>h</sup> ; 422.20(12) <sup>l</sup> ( $2p$ thresholds)		497.44 <sup>h</sup>
			497.80(40) <sup>l</sup>
	425.0 <sup>m</sup>	427.31 <sup>m</sup>	501.92 <sup>m</sup>
K <sup>9+</sup>	505.212 <sup>a</sup>	508.276 <sup>a</sup>	588.158 <sup>a</sup>
Ca <sup>10+</sup>	593.117 <sup>a</sup>	597.020 <sup>a</sup>	682.642 <sup>a</sup>
V <sup>13+</sup>	897.825 <sup>a</sup>	905.308 <sup>a</sup>	1008.297 <sup>a</sup>
Cr <sup>14+</sup>	1013.060 <sup>a</sup>	1022.162 <sup>a</sup>	1130.993 <sup>a</sup>
Fe <sup>16+</sup>	1264.041 <sup>a</sup>	1277.161 <sup>a</sup>	1397.766 <sup>a</sup>
	1262.7(7) <sup>h</sup>	1278.80 <sup>h</sup>	1394.71(7) <sup>h</sup>
	1265.8 <sup>m</sup>	1275.42(7) <sup>m</sup>	1399.31 <sup>m</sup>
Ni <sup>18+</sup>	1542.387 <sup>a</sup>	1560.743 <sup>a</sup>	1693.251 <sup>a</sup>
Ge <sup>22+</sup>	2181.300 <sup>a</sup>	2214.738 <sup>a</sup>	2371.491 <sup>a</sup>
Kr <sup>26+</sup>	2930.110 <sup>a</sup>	2986.607 <sup>a</sup>	3168.284 <sup>a</sup>
	2928.90(17) <sup>h</sup>	2984.25(17) <sup>h</sup>	
	2932.1 <sup>m</sup>	2988.36 <sup>m</sup>	3170.08 <sup>m</sup>
Zr <sup>30+</sup>	3789.213 <sup>a</sup>	3879.321 <sup>a</sup>	4086.719 <sup>a</sup>
Cd <sup>38+</sup>	5840.278 <sup>a</sup>	6041.968 <sup>a</sup>	6303.720 <sup>a</sup>
Xe <sup>44+</sup>	7672.183 <sup>a</sup>	8011.974 <sup>a</sup>	8317.617 <sup>a</sup>
	7660(4) <sup>h</sup>		
	7672.8 <sup>m</sup>	8011.98 <sup>m</sup>	8318.53 <sup>m</sup>
Nd <sup>50+</sup>	9758.418 <sup>a</sup>	10301.724 <sup>a</sup>	10654.559 <sup>a</sup>
Dy <sup>56+</sup>	12101.798 <sup>a</sup>	12935.758 <sup>a</sup>	13339.722 <sup>a</sup>
Ta <sup>63+</sup>	15165.047 <sup>a</sup>	16485.428 <sup>a</sup>	16954.945 <sup>a</sup>
Hg <sup>70+</sup>	18588.463 <sup>a</sup>	20609.196 <sup>a</sup>	21151.655 <sup>a</sup>
Bi <sup>73+</sup>	20167.490 <sup>a</sup>	22572.300 <sup>a</sup>	23148.485 <sup>a</sup>

<sup>a</sup>Present DHF thresholds.

<sup>b</sup>Reference [36] (E).

<sup>c</sup>Reference [16] (E).

<sup>d</sup>Reference [15] (E).

<sup>e</sup>Reference [37] (E).

<sup>f</sup>Reference [38] (T)

<sup>g</sup>Reference [31] (T).

<sup>h</sup>NIST [39,40].

<sup>i</sup>Reference [22] (E).

<sup>j</sup>Reference [21] (E).

<sup>k</sup>Reference [25] (T).

<sup>l</sup>Reference [23] (E).

<sup>m</sup>Reference [29] (T; Dirac atomic  $R$ -matrix code (DARC) results).

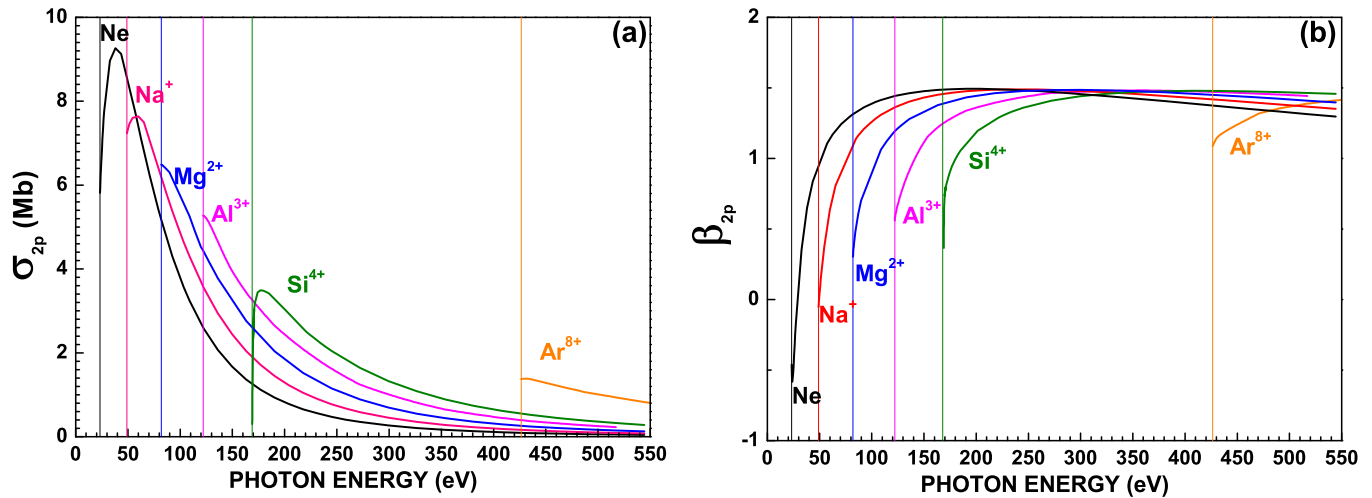


FIG. 1. (Color online) (a) Total  $2p$  ( $2p_{3/2} + 2p_{1/2}$ ) nonresonant (background) photoionization cross section for Ne,  $\text{Na}^+$ ,  $\text{Mg}^{2+}$ ,  $\text{Al}^{3+}$ ,  $\text{Si}^{4+}$ , and  $\text{Ar}^{8+}$ . (b) Weighted average (as described in the text) of the associated  $2p_{3/2}$  and  $2p_{1/2}$  photoelectron angular distribution asymmetry parameter. The vertical lines represent the respective ionization thresholds.

Also shown in Fig. 1(b) are the angular distribution asymmetry parameters  $\beta$  as a function of photon energy as a weighted (by respective partial cross sections) average of the angular distribution asymmetry parameter for the  $2p_{3/2}$  and the  $2p_{1/2}$  subshells. The evolution of  $\beta$  as a function of  $Z$  is clearly seen; they become flatter and flatter as a function of energy as the shape resonance moves further and further below threshold, except for the anomaly at  $\text{Si}^{4+}$  owing to the Feshbach resonance just at threshold. For neutral Ne, these results have previously been shown to be in excellent agreement with experiment [45]. Thus, it is rather likely that the calculations for the rest of the sequence are quantitatively accurate as well. This is of interest since the recent advent of UV and x-ray laser sources has made it possible to measure photoelectron angular distributions resulting from ionic photoionization; there are essentially no extant measurements of photoelectron angular distributions resulting from ionic photoionization.

Turning our attention to the  $2s \rightarrow np$  resonances, it is crucial to note that, in general, MQDT parameters are almost

insensitive to energy in the threshold region. As an example, for  $\text{Na}^+$ , we present, in Fig. 2(a), the eigenamplitudes  $D_\alpha$  and, in Fig. 2(b), the corresponding quantum defects  $\mu_\alpha$  for the seven  $\alpha$  channels arising from the  $2s$ ,  $2p_{1/2}$ , and  $2p_{3/2}$  subshells of  $\text{Na}^+$  to illustrate the near insensitivity to energy of the MQDT parameters obtained from the application of RRPA. Results for other members of the Ne isoelectronic sequence studied are qualitatively similar. The  $D_\alpha$  data for  $\text{Na}^+$  show that channels  $n = 1, 3$ , and  $5$ , the optically allowed (mostly) singlet channels, are relatively stronger than the remaining four triplet channels and are possible at all results from the relativistic interactions.

Using RRPA to calculate the RMQDT parameters and the latter to obtain the physical observables, the  $2s \rightarrow np$  autoionization resonances in the  $2p$  cross section and the corresponding angular distribution asymmetry parameter  $\beta$  have been obtained. To give some idea of how the spectrum changes along the sequence, the results for Ne,  $\text{Na}^+$ ,  $\text{Si}^{4+}$ , and  $\text{Ar}^{8+}$  are presented in Figs. 3 to 6, respectively. In each of these figures, panels (a), (b), (c), and (d) provide, respectively,

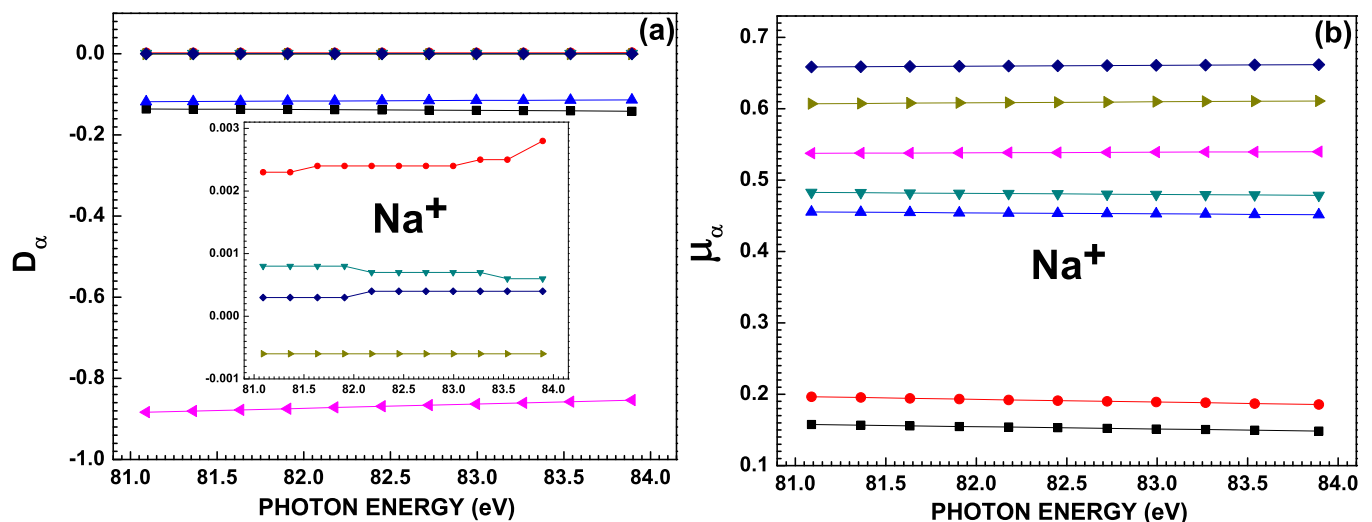


FIG. 2. (Color online) Variation of (a) eigenamplitudes  $D_\alpha$  and (b) eigen-quantum defects  $\mu_\alpha$  of  $\text{Na}^+$  as a function of photon energy (eV).

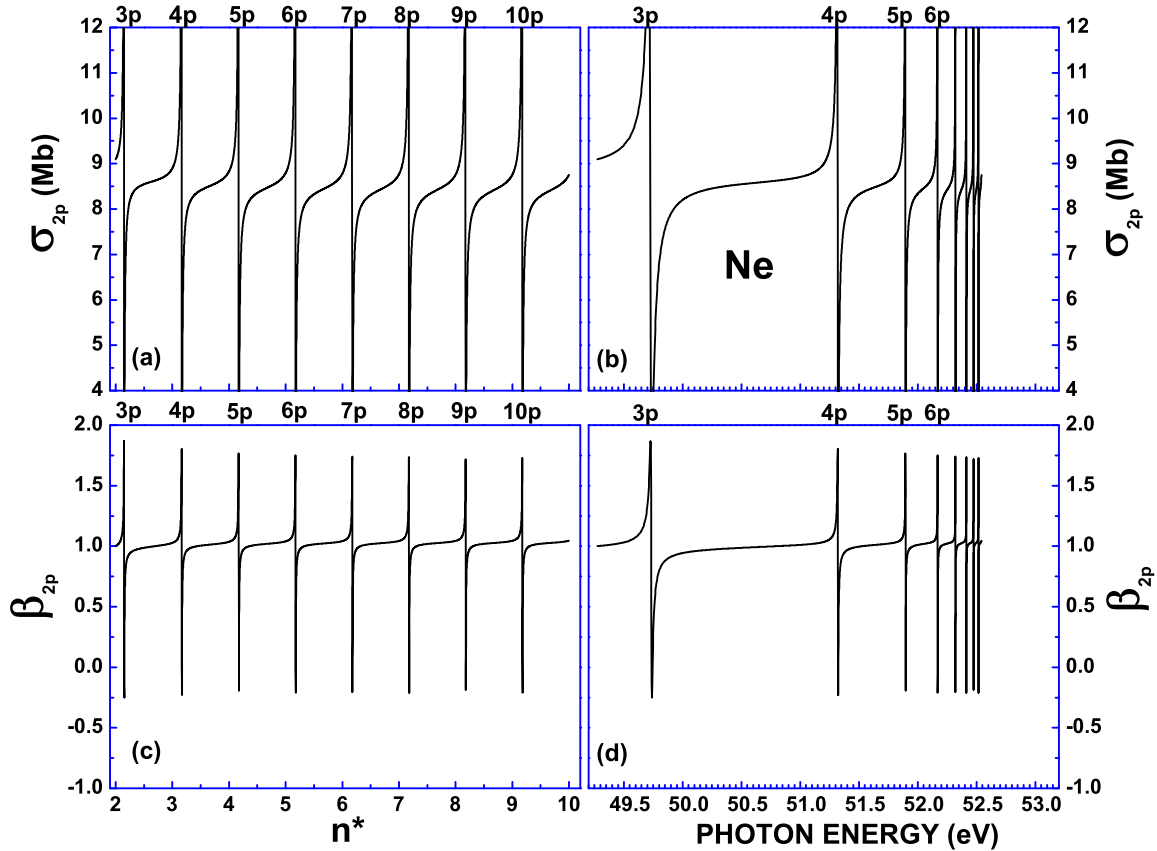


FIG. 3. (Color online) Total  $2p$  cross sections and (weighted average) angular distribution asymmetry parameters for Ne in the region of the  $2s \rightarrow np$  resonances vs effective quantum number  $n^*$  and photon energy. The resonances are labeled.

the cross section as a function of effective quantum number, the cross section as a function of the photon energy,  $\beta$  as a function of the effective quantum number, and  $\beta$  as a function of the photon energy.

The autoionization resonances in the  $2p$  cross section shown in Figs. 3 to 6 are seen to be periodic functions of the effective quantum number  $n^*$ ; when plotted against energy, they converge to their respective  $2s$  thresholds. The quantum defects determined for Ne and for a few members of the isoelectronic ions are given in Table II, along with results using other theoretical methods and/or experimental techniques. It is seen from Table II that the present results are in quite good agreement with experimental and theoretical values reported in the literature. For example, the quantum defect for the  $^1P$  resonance in atomic neon found from this work is about 0.84, which is in close agreement with the available theory (0.842) and experiment [0.832(6)] [19]. The background cross section is about 8.8 Mb, in good agreement with the available experimental and theoretical data [19,45]. The background value of the  $\beta$  parameter for Ne, as seen from Fig. 3, is roughly unity in this energy range. Based upon the general expression for  $\beta$ , this value is indicative of the  $p \rightarrow d$  matrix element being much larger than the  $p \rightarrow s$  channel [50–52]. This is one illustration of the detailed physics of the particular photoionization process that can be gleaned from studies of the photoelectron angular distribution, information which cannot be obtained from the cross section itself.

TABLE II. Quantum defects  $\mu_{np}$  for the  $2s \rightarrow np \ ^1P$  resonances of the Ne isoelectronic sequence. The present work gives the asymptotic value calculated, while the experimental and previous theories give the range reported.

Atom or ion	Present work	Experiment	Previous theory
Ne	$\sim 0.84$	$\sim 0.832(6)^a$ $\sim 0.829\text{--}0.89^b$ $\sim 0.85\text{--}0.88^d$	$\sim 0.842^a$ $\sim 0.79\text{--}0.81^c$
Na <sup>+</sup>	$\sim 0.66$	$\sim 0.68\text{--}0.78^d$	$\sim 0.31\text{--}0.6^e$
Mg <sup>2+</sup>	$\sim 0.55$	$\sim 0.57^d$	$\sim 0.48\text{--}0.56^f$
Al <sup>3+</sup>	$\sim 0.48$	$\sim 0.497^d$ ( $^1P_1$ )	$\sim 0.44755\text{--}0.48047^g$
Si <sup>4+</sup>	$\sim 0.43$		$\sim 0.27\text{--}0.28^i$
Ar <sup>8+</sup>	$\sim 0.28$	$\sim 0.286\text{--}0.33^h$	$\sim 0.27889^j$ ( $^1P_1$ )
Fe <sup>16+</sup>	$\sim 0.17$		$\sim 0.17562^j$ ( $^1P_1$ )
Kr <sup>26+</sup>	$\sim 0.116$		$\sim 0.12524^j$ ( $^1P_1$ )
Xe <sup>44+</sup>	$\sim 0.074$		$\sim 0.09489^j$ ( $^1P_1$ )

<sup>a</sup>Reference [19].

<sup>b</sup>Reference [16].

<sup>c</sup>Reference [46].

<sup>d</sup>Reference [11].

<sup>e</sup>Reference [47].

<sup>f</sup>Reference [48].

<sup>g</sup>Reference [49].

<sup>h</sup>Reference [23].

<sup>i</sup>Reference [25].

<sup>j</sup>Reference [29].

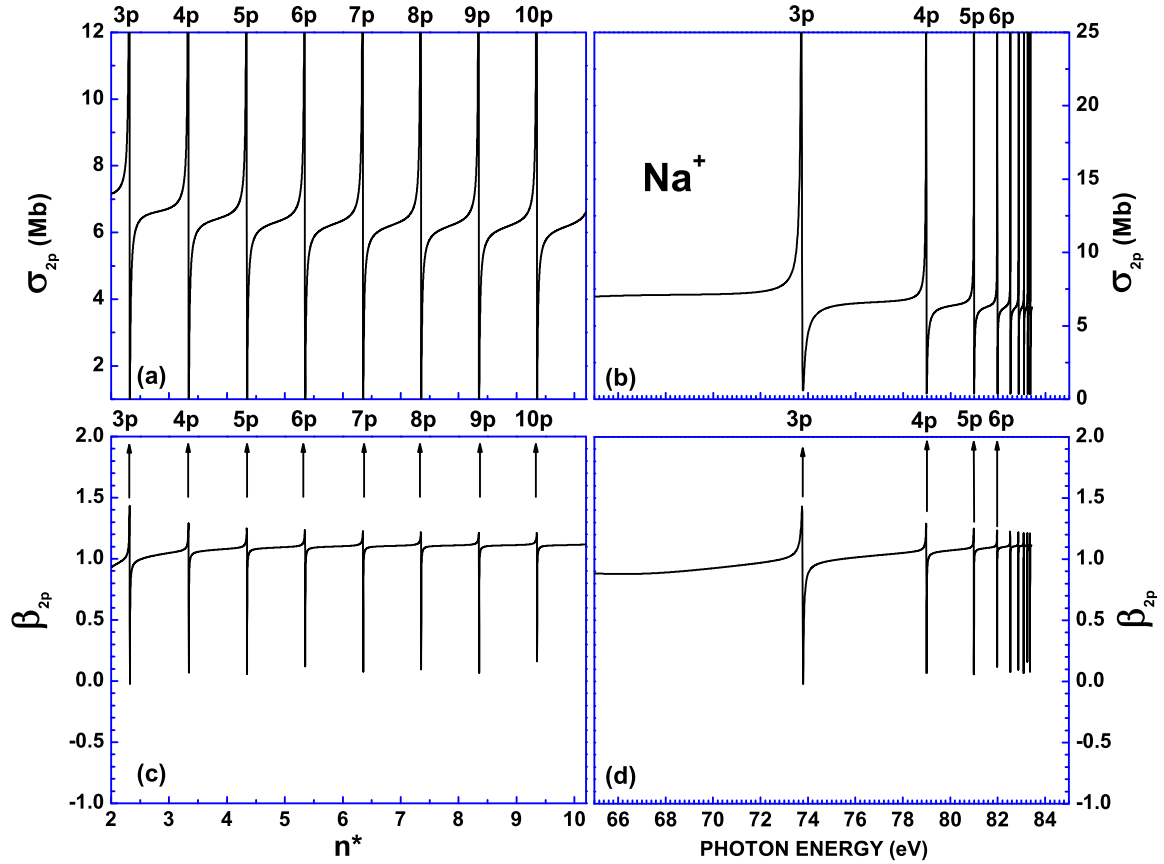


FIG. 4. (Color online) Total  $2p$  cross sections and (weighted average) angular distribution asymmetry parameters for  $\text{Na}^+$  in the region of the  $2s \rightarrow np$  resonances vs effective quantum number  $n^*$  and photon energy. The resonances are labeled.

For  $\text{Na}^+$  (Fig. 4), the background cross section, in the resonance range, is somewhat smaller than that of Ne and decreases over the range from about 7.5 to roughly 6 Mb. The resonance shapes differ somewhat as well. The background value of the  $\beta$  parameter is still unity, indicating that, in this case too, the  $p \rightarrow d$  transition dominates in this region. For  $\text{Al}^{3+}$ , the background cross section is about 5.27 Mb above the  $2p_{1/2}$  threshold, and it gradually decreases to about 3.57 Mb in the vicinity of the  $2s$  threshold. The cross-section values for  $\text{Al}^{3+}$  are in close agreement with the reported values [30,31]. The shapes of the resonances in the angular distribution parameter show marked differences from the Ne case. These trends are continued in  $\text{Si}^{4+}$  (Fig. 5), for which the background cross section varies from about 3.5 to about 2.5 Mb over the resonance range and the resonance shapes start to differ markedly. The background  $\beta$  is seen to rise significantly above the value of unity, thus indicating that the  $p \rightarrow d$  channel is not so dominant in this case. Finally, for  $\text{Ar}^{8+}$  (Fig. 6), the background cross section is significantly smaller, decreasing from 1.4 to 1.1 Mb over the range, and the background value of  $\beta$  is as large as 1.4, thereby showing the decreasing dominance of the  $p \rightarrow d$  channel with increasing  $Z$  along the isoelectronic sequence.

Concerning quantum defects, it is known that, in the limit of infinite  $Z$ , all quantum defects approach zero because interelectron interactions become irrelevant compared to the interaction with the nuclear charge in that limit. It is of interest, however, to investigate the detailed behavior of the quantum

defects as a function of  $Z$  along an isoelectronic sequence. At low  $Z$ , which is governed by nonrelativistic physics, the quantum defect of a resonance is simply  $n - n^*$ , where  $n$  is the principal quantum number and  $n^*$  is the effective quantum number. But at higher  $Z$ , where relativistic physics and  $jj$  coupling hold sway, the relationship is not so simple; in that case, the quantum defect is given by [33,35,53]

$$\mu_{nlj} = n - n^* + (j + 1) - [(j + 1)^2 - (z\alpha)^2]^{1/2}, \quad (4)$$

where  $z$  is the asymptotic charge seen by the photoelectron and  $\alpha$  is the fine-structure constant. Our RRPA results show that the lowest values of  $Z$  (or  $z$ ), the  $2s \rightarrow np$  transitions, which result in final states  $2s2p^6np$ , are  $LS$  coupled,  $^1P_1$  or  $^3P_1$ , with the amplitude of the latter being extremely small. But with increasing  $Z$ , the coupling rapidly moves away from  $LS$  and towards  $jj$  coupling, and by  $Z = 40$ , the final states are fully  $jj$  coupled,  $2s_{1/2}2p^6np_{1/2}$  and  $2s_{1/2}2p^6np_{3/2}$ , both coupled to  $J = 1$ ; this is required for dipole transitions from a  $J = 0$  initial state.

The calculated asymptotic quantum defects (actually calculated for  $n = 20$ ) over the whole isoelectronic sequence are shown in Fig. 7, where the approach to zero, at the higher  $Z$  values, is evident. It is also clear that, as relativistic effects become important, the  $np$  quantum defects depend upon  $j$ , and the  $p_{3/2}$  quantum defects approach zero more quickly than the  $p_{1/2}$  quantum defects. This occurs because the  $np_{1/2}$  electrons are more compact than their  $np_{3/2}$  counterparts, so more of their density is in the interior non-Coulomb region of

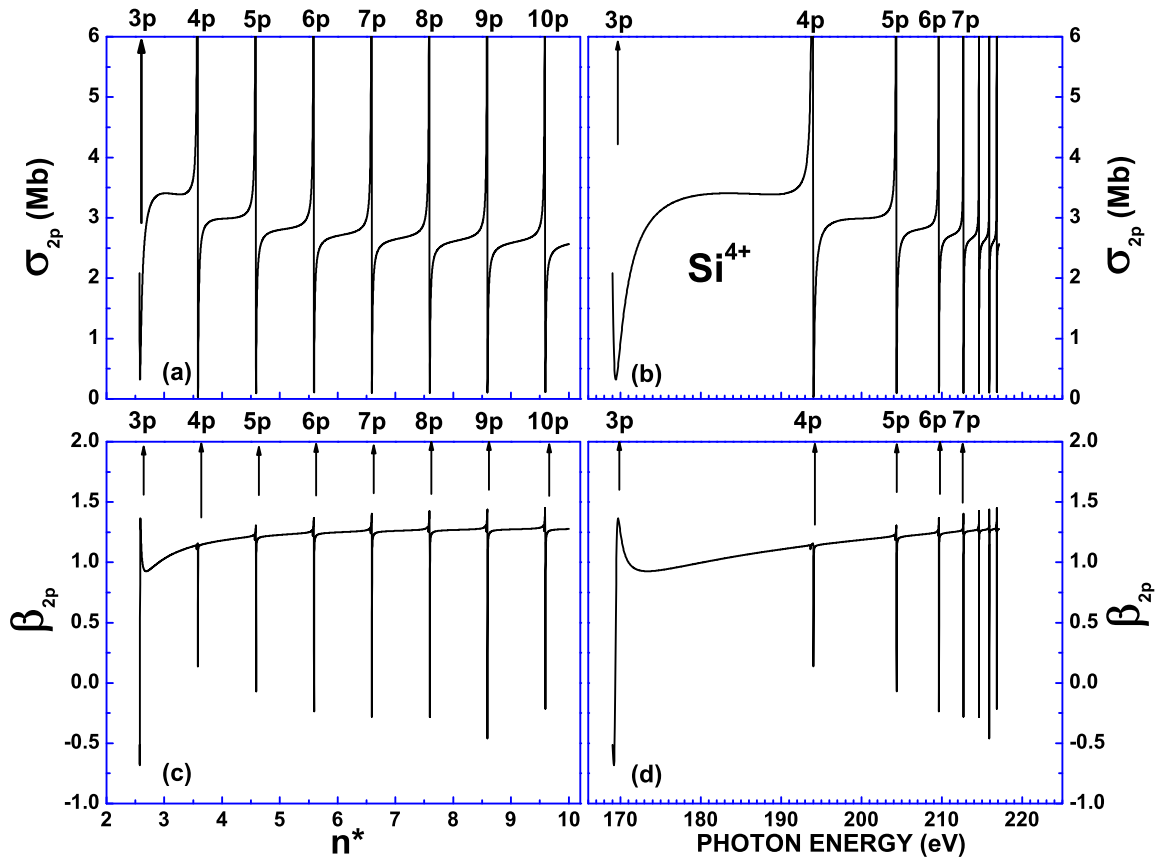


FIG. 5. (Color online) Total  $2p$  cross sections and (weighted average) angular distribution asymmetry parameters for  $\text{Si}^{4+}$  in the region of the  $2s \rightarrow np$  resonances vs effective quantum number  $n^*$  and photon energy. The resonances are labeled.

the potential, which is where the quantum defect is generated. In the low- $Z$  region, the top curve (larger quantum defects) represents the  $^3P$  resonances, which are more tightly bound than the  $^1P$  resonances owing to the fact that exchange interaction is attractive in the triplet case and repulsive in the singlet states.

A recent theoretical study of Ne-like  $\text{Cu}^{19+}$  found the asymptotic quantum defects to be about 0.28 [54], in considerable disagreement with the present results of about 0.18, as seen in Fig. 7. Now, it can be shown on general grounds that the asymptotic quantum defects must decrease with increasing  $Z$ . However, as seen in Table II, the asymptotic quantum defect for  $\text{Ar}^{8+}$  is about 0.28, which is corroborated by an independent calculation [25]. Thus, it appears likely that the results of Ref. [54] are quantitatively in error. Settling this question is of astrophysical importance.

To compare theoretical cross sections with existing experimental data, we have applied a Gaussian convolution to our results for the autoionization resonance profiles with appropriate widths to simulate the measurements. In addition, the theoretical results have been shifted in energy to align the  $2s$  thresholds with experiment by the amounts indicated in Table I. To give some idea of the effect(s) of the convolutions, Fig. 8 shows the convoluted theoretical  $2s \rightarrow np$  resonances in atomic neon along with the unconvoluted resonances. The comparison shows clearly that the convoluted resonances are wider and have a smaller amplitude than their unconvoluted counterparts; these characteristics *must* go together since the

convolution preserves oscillator strength. The convolution does not affect the position of the resonances, nor does it affect the background cross section. All the experimental results are for the relative cross section, so they were put on an absolute scale by normalizing experiment to theory in the region of the nonresonant continuum where the theoretical results are likely to be excellent. Using this procedure for the experimental cross sections, the  $2s \rightarrow 3p$ ,  $2s \rightarrow 4p$ , and  $2s \rightarrow 5p$  resonances in the neon atom, convoluted at 18 meV FWHM, as quoted in Ref. [55], are shown in Fig. 9, along with the experimental data. From the comparison, it is evident that there is excellent quantitative agreement between theory and experiment for the  $4p$  and  $5p$  resonances. The agreement is not as good for the  $3p$  resonance. This is most likely due to the fact that the quantum-defect parameters experience their greatest variation with energy in this energy region, so that the interpolation required by the RMQDT methodology is mostly at risk here. However, as indicated earlier, the accuracy will improve with increasing  $Z$ ; that is, the situation for the Ne  $2s \rightarrow 3p$  resonance is the worst-case scenario.

Although data exist for  $\text{Na}^+$  [20], they are not in a form that could be digitized and compared herein. But there is accessible experimental data for  $\text{Ar}^{8+}$ , specifically in the region of the  $2s \rightarrow 5p$  resonance [23], which is the lowest resonance in the continuum for  $\text{Ar}^{8+}$ ; the  $2s \rightarrow 4p$  resonance is just below threshold, as discussed earlier, and the  $2s \rightarrow 3p$  resonance is well below threshold. A comparison with our calculated data is given in Fig. 10, which shows clearly that agreement

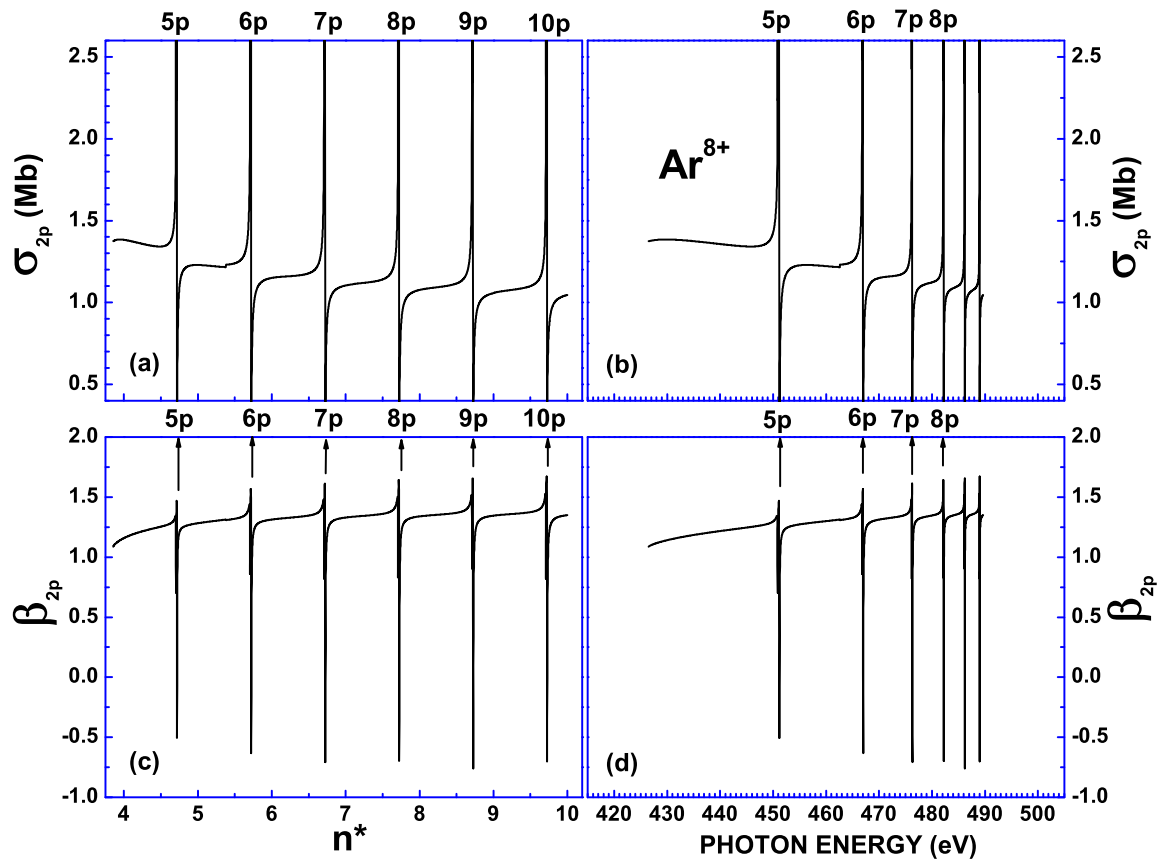


FIG. 6. (Color online) Total  $2p$  cross sections and (weighted average) angular distribution asymmetry parameters for  $\text{Ar}^{8+}$  in the region of the  $2s \rightarrow np$  resonances vs effective quantum number  $n^*$  and photon energy. The resonances are labeled.

is quite good, despite the rather significant scatter in the experimental points. The results are in good agreement with the experiment when convoluted with a Gaussian with a FWHM of 100 meV. Based upon these comparisons with experiment, we can be reasonably confident that the predictions of the present

calculations should be reasonably accurate for the  $2s \rightarrow np$  resonances over the entire isoelectronic sequence above the lowest member of the series in each case.

Now resonances can be characterized by parameters determined by their position, width, and shape, which are known

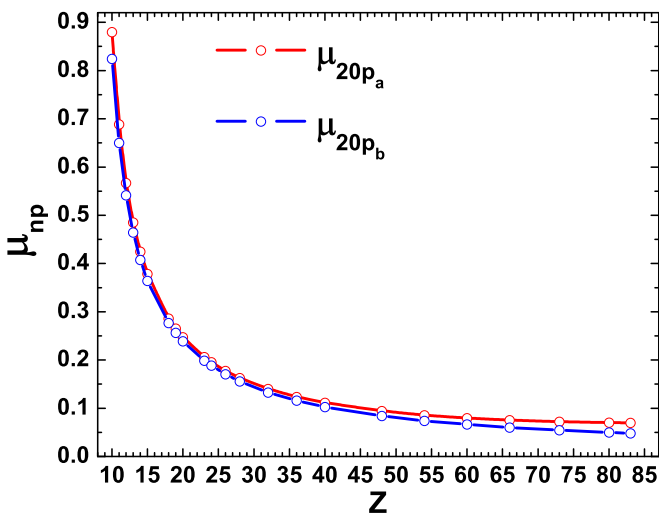


FIG. 7. (Color online) Quantum defects of the Ne isoelectronic sequence as a function of nuclear charge  $Z$ . For low  $Z$ , the top and bottom curves represent the  $^3P$  and  $^1P$  resonances, respectively, while for high  $Z$ , they represent the  $p_{1/2}$  and  $p_{3/2}$  excitations, respectively.

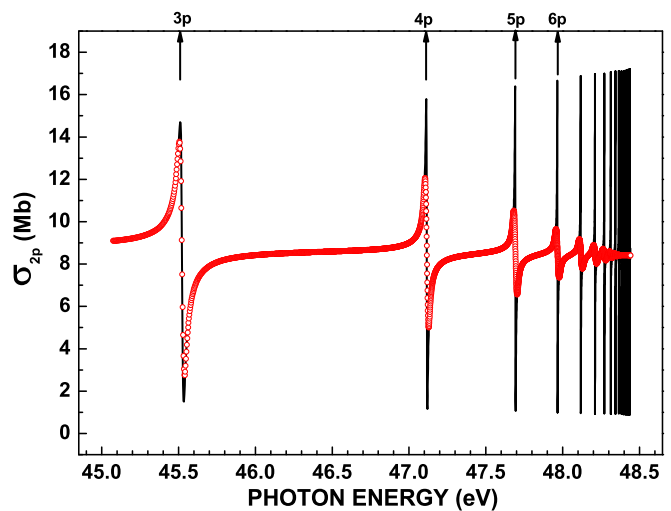


FIG. 8. (Color online) Total  $2p$  photoionization cross section in the region of the  $2s \rightarrow np$  resonances in atomic Ne obtained using RMQDT (solid line) and the cross section convoluted with FWHM = 18 meV (open circles).

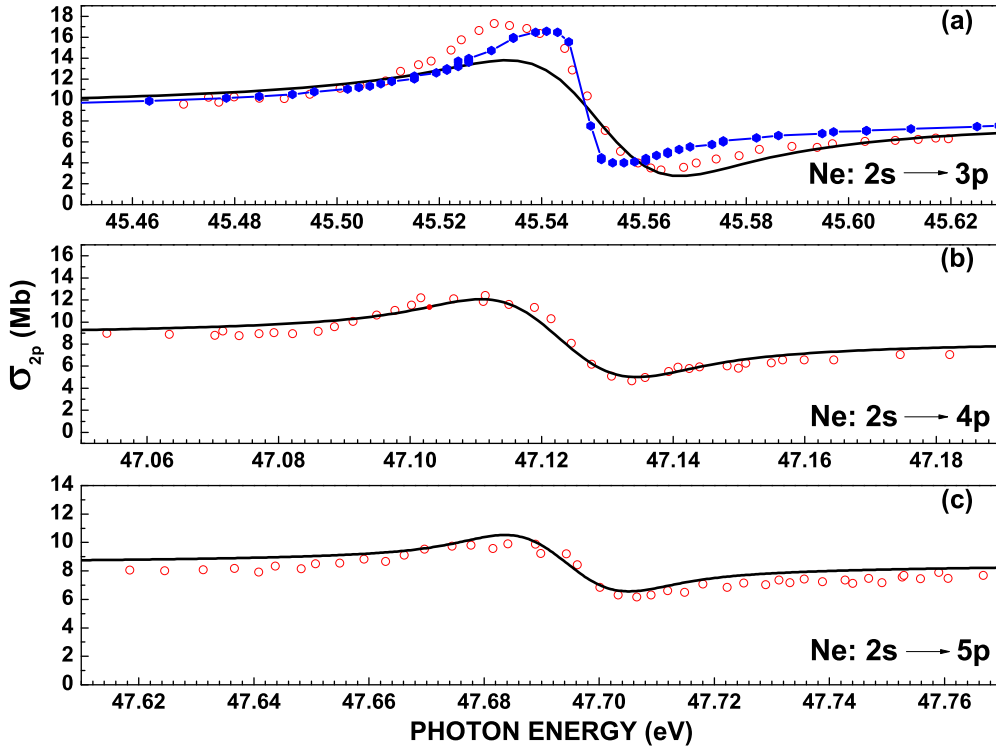


FIG. 9. (Color online) Total  $2p$  cross section for Ne, convoluted at  $\text{FWHM} = 18$  meV, in the region of the (a)  $2s \rightarrow 3p$ , (b)  $2s \rightarrow 4p$ , and (c)  $2s \rightarrow 5p$  resonances (solid lines). Also shown are the experimental results of Ref. [55] (open circles) and Ref. [19] (connected solid hexagons).

as Fano parameters [6,7,16]. Using these parameters, the photoionization cross section in a resonance region is given by

$$\sigma(E) = \sigma_0(E) \left[ (1 - \rho^2) + \rho^2 \left( \frac{(q + \epsilon)^2}{(1 + \epsilon^2)} \right) \right], \quad (5)$$

where  $\epsilon = \frac{(E - E_r)}{(\Gamma/2)}$ ,  $E_r$  is the resonance energy,  $\Gamma$  is the width,  $q$  is the shape parameter,  $\rho^2$  is the correlation coefficient,

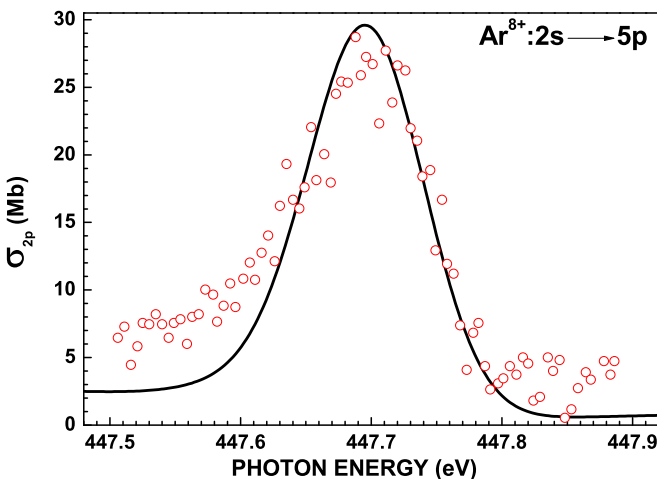


FIG. 10. (Color online) Total calculated  $2p$  cross section in the neighborhood of the  $2s \rightarrow 5p$  autoionization resonance in  $\text{Ar}^{8+}$ , convoluted with a 100 meV FWHM instrumental width (solid line), along with the experimental results of Ref. [23] (open circles).

and  $\sigma_0(E)$  is the background cross section; the detailed cross section in a resonance region can be specified completely using these parameters.

Our calculated Fano parameters are presented in Table III for cases where experimental data exist, specifically for  $2s \rightarrow 3p$ ,  $2s \rightarrow 4p$ ,  $2s \rightarrow 5p$ , and  $2s \rightarrow 6p$  in Ne,  $2s \rightarrow 3p$  in  $\text{Na}^+$  and  $\text{Mg}^{2+}$ , and for the  $2s \rightarrow 5p$  transition in  $\text{Ar}^{8+}$ . Table III shows relatively good agreement between the present results and experiment in all cases shown, except for the  $2s \rightarrow 5p$  transition in  $\text{Ar}^{8+}$ . First of all, there is a non-negligible relativistic splitting between the two  $5p$  resonances; however, owing to experimental instrumental width and the fact that one  $5p$  resonance is much weaker than the other, this does not show up experimentally, as seen in Fig. 10. Second, as discussed previously, the lowest member of each series is not quite so accurate when using RMQDT, and this case is the lowest member of the resonance series for  $\text{Ar}^{8+}$ . Third, and most important, a very fine energy mesh is required to properly characterize a resonance. In this case, the experimental energy mesh was 5 meV [23]. On the theoretical side, Ref. [25] used a mesh of 0.001 Ry (13.6 meV), Ref. [59] used a mesh of about 0.0015 Ry (about 20 meV), and  $\Delta n^* = 0.001$  (24.2 meV; present work) in this energy region. We studied how different energy meshes yield different cross-section profiles for  $\text{Ar}^{8+}$ , with  $\Delta n^* = 0.0001$  (2.42 meV),  $\Delta n^* = 0.001$  (24.2 meV), and  $\Delta n^* = 0.005$  (121 meV). We find that wider meshes yield more erroneous values, and further scrutiny of these resonances, both theoretical and experimental, is warranted. However, the oscillator strength of these resonances is preserved since the mesh splitting employed in the present work is smaller than the instrumental width.



TABLE III. Fano parameters for various resonances compared to experiment and previous theory. Experimental results are italicized. Uncertainties (where quoted) are given in parentheses.

Atom or ion: Resonance	Resonance energy $E_r$ (eV)	Width $\Gamma$ (meV)	Shape parameter $q$	Correlation coefficient $\rho^2$
Ne: $2s \rightarrow 3p$	49.725 <sup>a</sup>	13 <sup>a</sup>	-1.4 <sup>a</sup>	0.77 <sup>a</sup>
Ne ( $2s2p^63p$ )	45.5442 <sup>b</sup>	16(2) <sup>b</sup>	-1.58(1) <sup>c</sup>	0.75(5) <sup>c</sup>
	45.53397 <sup>b</sup>	34.9 <sup>b</sup>	-1.59(1) <sup>c</sup>	0.72 <sup>c</sup>
	45.557 <sup>c</sup>	13.2(10) <sup>c</sup>	-1.53(1) <sup>c</sup>	0.73 <sup>c</sup>
	45.546(8) <sup>d</sup>	18.6(10) <sup>c</sup>	-1.6(2) <sup>d</sup>	0.70(7) <sup>d</sup>
	46.253 <sup>e</sup>	13(2) <sup>d</sup>	-3.69 <sup>e</sup>	0.514 <sup>e</sup>
	45.5655 <sup>f</sup>	13.9 <sup>e</sup>	-0.34 <sup>g</sup>	0.93 <sup>g</sup>
	45.538 <sup>h</sup>	11.4 <sup>f</sup>	-1.16 <sup>g</sup>	0.91 <sup>g</sup>
		11.7 <sup>g</sup>	-1.6 <sup>g</sup>	0.76 <sup>g</sup>
		12.1 <sup>g</sup>	-1.3 <sup>g</sup>	0.77 <sup>g</sup>
		31.8 <sup>h</sup>	-1.32 <sup>h</sup>	
Ne: $2s \rightarrow 4p$	51.318 <sup>a</sup>	7 <sup>a</sup>	-1.35 <sup>a</sup>	0.63 <sup>a</sup>
Ne ( $2s2p^64p$ )	47.1193(50) <sup>b</sup>	6.65 <sup>b</sup>	-1.47(1) <sup>c</sup>	0.78(11) <sup>c</sup>
	47.11092 <sup>b</sup>	5.7(10) <sup>c</sup>	-1.88 <sup>c</sup>	0.72 <sup>c</sup>
	47.111 <sup>c</sup>	4.3 <sup>c</sup>	-1.82 <sup>c</sup>	0.73 <sup>c</sup>
	47.121(5) <sup>d</sup>	4.5(1.5) <sup>d</sup>	-1.6(3) <sup>d</sup>	0.70(7) <sup>d</sup>
	47.397 <sup>e</sup>	3.86 <sup>e</sup>	-3.95 <sup>e</sup>	0.505 <sup>e</sup>
	47.1278 <sup>f</sup>	5.28 <sup>f</sup>	-1.75 <sup>g</sup>	0.76 <sup>g</sup>
		3.8 <sup>g</sup>	-1.46 <sup>g</sup>	0.77 <sup>g</sup>
Ne: $2s \rightarrow 5p$	51.894 <sup>a</sup>	3 <sup>a</sup>	-1.15 <sup>a</sup>	0.71 <sup>a</sup>
Ne ( $2s2p^65p$ )	47.6952(15) <sup>b</sup>	2.47 <sup>b</sup>	-1.46(5) <sup>c</sup>	0.6(2) <sup>c</sup>
	47.69182 <sup>b</sup>	3.6(18) <sup>c</sup>	-1.9 <sup>c</sup>	0.74 <sup>c</sup>
	47.687 <sup>c</sup>	1.8 <sup>c</sup>	-1.87 <sup>c</sup>	0.75 <sup>c</sup>
	47.692(5) <sup>d</sup>	2(1) <sup>d</sup>	-1.6(5) <sup>d</sup>	0.70(14) <sup>d</sup>
	47.814 <sup>e</sup>	1.62 <sup>e</sup>	-4.05 <sup>e</sup>	0.502 <sup>e</sup>
	47.6975 <sup>f</sup>	2.61 <sup>f</sup>		
Ne: $2s \rightarrow 6p$	52.168 <sup>a</sup>	1.58 <sup>a</sup>	-1.04 <sup>a</sup>	0.89 <sup>a</sup>
Ne ( $2s2p^66p$ )	47.9650(30) <sup>b</sup>	1.28 <sup>b</sup>		
	47.96708 <sup>b</sup>	1.44 <sup>f</sup>		
	47.9690 <sup>f</sup>			
Na <sup>+</sup> : $2s \rightarrow 3p$	73.746 <sup>a</sup>	60 <sup>a</sup>	-1.8 <sup>a</sup>	0.9 <sup>a</sup>
Na <sup>+</sup> ( $2s2p^63p$ )	69.95 <sup>i</sup>			
Mg <sup>2+</sup> : $2s \rightarrow 3p$	101.525 <sup>a</sup>	90 <sup>a</sup>	-2.3 <sup>a</sup>	0.975 <sup>a</sup>
Mg <sup>2+</sup> ( $2s2p^63p$ )	98.2 <sup>j</sup>			
	98.22 <sup>k</sup>			
Ar <sup>8+</sup> : $2s \rightarrow 5p$	451.123 <sup>a</sup>	22.395 <sup>a</sup>	-7.6 <sup>a</sup>	0.866 <sup>a</sup>
Ar <sup>8+</sup> ( $2s2p^65p$ )	447.71(10) <sup>l</sup>	30(5) <sup>l</sup>	-6(1) <sup>l</sup>	
	447.54(30) <sup>l</sup>	29.01 <sup>m</sup>		
	447.726 <sup>m</sup>	25.8 <sup>n</sup>		
	447.33 <sup>n</sup>			

<sup>a</sup>Present results (RRPA+RMQDT, L form).

<sup>b</sup>Reference [19].

<sup>c</sup>Reference [55].

<sup>d</sup>Reference [16].

<sup>e</sup>Reference [56].

<sup>f</sup>Reference [46].

<sup>g</sup>Reference [43].

<sup>h</sup>Reference [57].

<sup>i</sup>Reference [58].

<sup>j</sup>Reference [47].

<sup>k</sup>Reference [11].

<sup>l</sup>Reference [23].

<sup>m</sup>Reference [25].

<sup>n</sup>Reference [29].

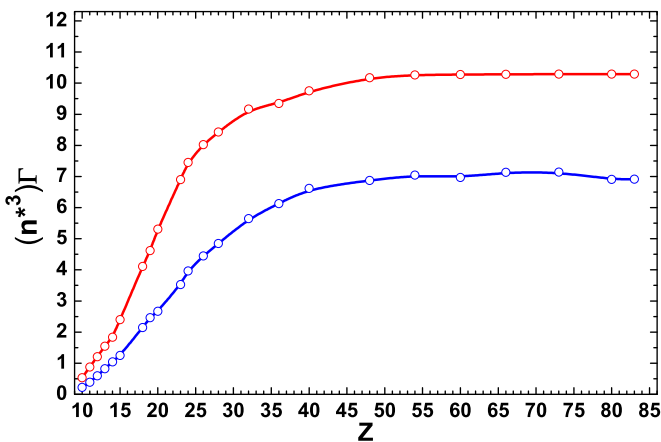
TABLE IV. Width  $\Gamma$ , effective quantum number  $n^*$ , and the product  $(n^*)^3\Gamma$  for  $20p_a$  and  $20p_b$  resonances for the ions in the Ne isoelectronic sequence.

$Z$	$n^*(20p_a)$	$n^*(20p_b)$	$\Gamma(20p_a)$ $\times 10^{(-4)} \text{ (eV)}$	$\Gamma(20p_b)$ $\times 10^{(-4)} \text{ (eV)}$	$[n^*(20p_a)]^3$ $\times \Gamma(20p_a)$	$[n^*(20p_b)]^3$ $\times \Gamma(20p_b)$
10	19.12034	19.17585	0.760	0.315	0.531252	0.222113
11	19.31165	19.34998	1.222	0.533	0.880095	0.386161
12	19.43293	19.45852	1.65	0.802	1.210874	0.590886
13	19.51473	19.53571	2.085	1.109	1.549508	0.826836
14	19.57532	19.59233	2.45	1.39	1.837777	1.045377
15	19.62166	19.63592	3.17	1.65	2.394785	1.249216
18	19.71376	19.72323	5.365	2.8	4.110344	2.148286
19	19.73477	19.74353	6.01	3.195	4.619242	2.458925
20	19.75263	19.76093	6.89	3.46	5.309994	2.669920
23	19.79384	19.80116	8.9	4.54	6.902083	3.524745
24	19.80451	19.81166	9.6	5.1	7.456990	3.965818
26	19.82263	19.82955	10.3	5.7	8.022709	4.444403
28	19.83746	19.84431	10.8	6.2	8.431055	4.845066
32	19.86009	19.86726	11.7	7.2	9.164937	5.646072
36	19.87651	19.88423	11.9	7.8	9.344743	6.132265
40	19.88884	19.89724	12.4	8.4	9.755512	6.616949
48	19.90587	19.91596	12.9	8.7	10.174971	6.872630
54	19.91428	19.92586	13	8.9	10.266849	7.041112
60	19.92037	19.93365	13	8.8	10.276271	6.970167
66	19.92466	19.93985	13	9	10.282912	7.135233
73	19.92802	19.94562	13	9	10.288115	7.141430
80	19.92977	19.95031	13	8.7	10.290825	6.908252
83	19.93003	19.95212	13	8.7	10.291228	6.910133

In any case, it is well known from quantum-defect theory [34,35] that the widths  $\Gamma$  of the higher Rydberg states decrease as  $(n^*)^3$  so that the product  $(n^*)^3\Gamma$  approaches a constant as  $n$  becomes large. Thus, knowledge of this constant for any given case, along with the asymptotic quantum defect, allows one to obtain the width of any state in the Rydberg series, except perhaps the few lowest. How this constant varies over the isoelectronic sequence is also of interest. To investigate this, calculations have been performed for the  $n = 20$  states of various members of the isoelectronic sequence and the widths obtained. As mentioned previously, the two

resonances for each  $n$  in the Ne isoelectronic sequence are essentially  $LS$  at low  $Z$  and  $jj$  at high  $Z$  and some mixture at intermediate- $Z$  values. Thus, the  $(n^*)^3\Gamma$  values for the two resonances for the ions in the Ne isoelectronic sequence are studied in the present work by labeling them as  $p_a$  and  $p_b$  resonances and not as singlet and triplet or as  $p_{1/2}$  and  $p_{3/2}$ .

The results of the calculations are shown in Table IV and Fig. 11. Both Table IV and the curves in Fig. 11 show clearly that the widths increase as a function of  $Z$  in the lower- $Z$  range and then level off for intermediate and high  $Z$ . Near the lower edge of the  $Z$  range, the top curve represents  $^3P$ , and the bottom curve represents the  $^1P$  states; that is, the  $^3P$  resonances are wider than the  $^1P$  resonances. At high  $Z$ , the top curve is for  $np_{1/2}$ , and the bottom curve is for  $np_{3/2}$ ;  $np_{1/2}$  states are broader than their  $p_{3/2}$  counterparts. For intermediate  $Z$ , we cannot easily characterize the resonances since they are neither  $LS$  nor  $jj$  but something in between and different for each  $Z$ . Table IV shows that the ratio of the widths of the two resonances with the same  $n$  is about 2.5 at low  $Z$ , and this drops to about 1.5 at high  $Z$ .


 FIG. 11. (Color online) Behavior of  $(n^*)^3\Gamma$  for the  $20p_a$  (top curve) and  $20p_b$  (bottom curve) resonances in the Ne isoelectronic sequence.

#### IV. SUMMARY AND CONCLUSIONS

The  $2s \rightarrow np$  resonances for a number of members of the Ne isoelectronic sequence have been investigated using RRPA and RMQDT. Asymptotic quantum defects and widths were obtained for a number of members of the sequence, and the results were presented and analyzed. Since these quantities behave smoothly as a function of  $Z$ , the data for any member of the sequence not calculated directly can be

obtained by interpolation. For each  $Z$ , there are two series of  $2s \rightarrow np$  transitions, which lead to  $1s^2 2s 2p^6 np$   $J = 1$  resonances that are distinguished by two different angular momentum couplings. At low  $Z$ , these  $1s^2 2s 2p^6 np$   $J = 1$  resonances can be characterized as  $^1P_1$  (optically allowed) and  $^3P_1$  (optically forbidden); it was seen that the transition to the optically allowed resonance state was much stronger at low  $Z$ , as expected. At high  $Z$ , the coupling is  $jj$ , and the resonance states are characterized as  $[[1s^2 2s 2p^6]_{1/2} np_{1/2}]$   $J = 1$ , and  $[[1s^2 2s 2p^6]_{1/2} np_{3/2}]$   $J = 1$ ;  $jj$  coupling is reached by approximately  $Z = 40$ . At intermediate  $Z$ , the coupling differs at each  $Z$  as we pass from  $LS$  to  $jj$  coupling, with increasing  $Z$ . Generally good agreement was found with experiment and previous theory, although certain discrepancies were noted, e.g., with a previous calculation of the quantum defect of these resonances for  $\text{Cu}^{19+}$  [54]. Since the quantum defects along this isoelectronic series are monotone decreasing, the

comparison showed clearly that the results of the earlier calculation were inaccurate since their values were larger than even  $\text{Ar}^{8+}$ , for which the values were confirmed by experiment; this shows the utility of performing calculations of isoelectronic sequences, which allow the results for individual  $Z$  to be placed in perspective.

Finally, it was demonstrated that, to get an accurate characterization of a resonance, a rather fine energy mesh must be used. Results were shown for three different energy meshes which gave rather different pictures of the  $2s \rightarrow 5p$  resonance in  $\text{Ar}^{8+}$ .

#### ACKNOWLEDGMENTS

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