Spin dynamics in a spin-orbit-coupled Fermi gas

Stefan S. Natu^{*} and S. Das Sarma

Condensed Matter Theory Center and Joint Quantum Institute, Department of Physics, University of Maryland, College Park, Maryland 20742-4111, USA (Received 10 May 2013; revised manuscript received 22 July 2013; published 11 September 2013)

We study the dynamics of a nondegenerate, harmonically trapped Fermi gas following a sudden ramp of the spin-orbit coupling strength using a Boltzmann equation approach. In the absence of interactions and a Zeeman field, we solve the spin-orbit-coupled Boltzmann equation *analytically* and derive expressions for the phase-space and temporal dynamics of an arbitrary initial spin state. For a fully spin-polarized initial state, the total magnetization exhibits collapse and revival dynamics in time with a period set by the trapping potential. In real space, this corresponds to oscillations between a fully polarized state and a spin helix. To make predictions relevant to current experiments on spin-orbit-coupled Fermi gases, we then numerically study the dynamics in the presence of an additional momentum-independent Zeeman field. We find that the spin helix is robust for weak magnetic fields but disappears for stronger field strengths. Finally, we explore the spin dynamics in the presence of interactions and find that weak interactions *enhance* the amplitude of the spin helix.

DOI: 10.1103/PhysRevA.88.033613

PACS number(s): 67.85.-d, 75.70.Tj, 03.75.Kk

I. INTRODUCTION

The physics of spin-orbit coupling is at the heart of fundamental phenomena such as the spin Hall effect [1,2], as well as practical devices such as the spin transistor [3]. Key to these developments in the field of spintronics is the understanding of how parameters such as the spin-orbit coupling, interaction, disorder, and geometry separately, as well as collectively, influence spin dynamics [4,5]. The creation of low-temperature atomic and molecular spin-orbit-coupled Bose and Fermi gases has paved the way for studying this physics in a setting where these parameters are well characterized [6-13]. Furthermore, a tool largely unique to ultracold gases is the ability to induce the spin-orbit coupling dynamically, thereby enabling the study of out-of-equilibrium physics in these systems. In this paper, we solve the dynamics of a noninteracting nondegenerate Fermi gas following a sudden ramp of the spin-orbit-coupling strength. The resulting out-of-equilibrium dynamics is rich [8,14–16]: Coherence inherent in cold atomic gases, but almost always absent in solid state systems, leads to collapse and revival of the total magnetization [15]. In real space, this is manifested by the spontaneous appearance of a helical spin texture which is the analog of the persistent spin helix observed in two-dimensional electron gases [17,18].

A further advantage of cold atomic systems is the ability to use magnetic fields to control the interactions between the atomic states [19,20]. This, particularly in the presence of spin-orbit coupling and *s*-wave superfluidity, may enable experimentalists to realize novel states of matter with topological properties whose excitations exhibit non-Abelian statistics [21,22]. In addition to their fundamental importance, such topological states of matter may serve as a platform for faulttolerant quantum computation [23]. However, a key challenge in observing this physics in ultracold gases is the inherent difficulty in attaining the sufficiently low temperatures needed to realize these topological states. In contrast, the nonequilibrium dynamics we study here occurs at high temperatures and can be observed in *current* experiments. We numerically study the spin dynamics of a fully polarized, weakly interacting Fermi gas using a collisionless Boltzmann equation. We find largeamplitude spin waves analogous to those previously observed in dilute spin-polarized hydrogen [24–28] and more recently in ultracold Bose and Fermi gases [29-34]. Our numerical study of spin dynamics in the collisionless spin-orbit-coupled Fermi gas complements a recent, linear-response study on a homogeneous system by Tokatly and Sherman [14]. We point out that the ultralow temperature degenerate version of our system (with spin-orbit coupling and large Zeeman splitting) would manifest topological superfluidity in the presence of ordinary s-wave superfluidity induced by suitable Feshbach resonance [22].

The effects predicted in our work are unlikely to be of much experimental significance in solid-state spin-orbit-coupled systems because of the strong disorder and decoherence intrinsically present in these systems, as well as the ultrafast time scales for spin relaxation [35]. In principle, the physics we predict should be present in both atomic and molecular Fermi and Bose gases since it is an intrinsically high-temperature phenomenon, but in light of recent experiments [10–12], we discuss our theoretical details using the atomic Fermi gas as the representative system of study.

This paper is organized as follows. In Sec. II, we describe our system and derive the collisionless Boltzmann equation for a two-component Fermi gas in a two-dimensional (2D) harmonic trap in the presence of spin-orbit coupling. In the subsequent sections, we choose an initial state which is a stationary state of the Hamiltonian in the absence of spin-orbit coupling. We then drive the system out of equilibrium by suddenly turning on the spin-orbit coupling and study the resulting dynamics. In Sec. III, we solve the Boltzmann equation in the absence of interactions, and in Sec. IV, we consider the effect of interactions on the spin dynamics. We summarize our results in Sec. V.

1050-2947/2013/88(3)/033613(8)

^{*}snatu@umd.edu

II. SETUP

In the absence of spin-orbit coupling, the Hamiltonian for two hyperfine states of a Fermi gas can be expressed as a sum of single-particle and two-body interaction terms,

$$\mathcal{H} = \mathcal{H}_{\rm s} + \mathcal{H}_{\rm int},\tag{1}$$

where the single-particle Hamiltonian is composed of a kinetic term and a potential term arising from the external trapping potential,

$$\mathcal{H}_{\rm s} = \sum_{i=\uparrow,\downarrow} \int d^3 \mathbf{r} \psi_i^{\dagger} \left[-\frac{\hbar^2 \nabla_{\mathbf{r}}^2}{2m} + U(\mathbf{r}) \right] \psi_i,$$

where ψ_i denotes the fermionic annihilation operator for a particle in hyperfine state $\{\uparrow,\downarrow\}$ and mass *m*. Here $U(\mathbf{r})$ refers to the external trapping potential, which we assume to be cylindrically symmetric $U(\mathbf{r}) = \frac{1}{2}m\omega_r^2(x^2 + y^2) + \frac{1}{2}m\omega_z^2z^2$, where ω_r and ω_z are the trapping frequencies in the radial and longitudinal directions, respectively. As the relevant spin-orbit physics is 2D, we assume a quasi-2D, pancake geometry, obtained by tight confinement in the longitudinal (z) direction $(\omega_z \gg \omega_r)$. For simplicity, we assume that both atomic states experience identical trapping potentials, but all of our calculations can be readily extended to include more general trapping potentials realized in experiments.

At the ultracold temperatures realized in these experiments, the dominant contribution to scattering comes from the *s*-wave channel. The interaction Hamiltonian therefore takes the simple form of a contact interaction,

$$\mathcal{H}_{\rm int} = \frac{g}{2} \int d\mathbf{r} \psi_{\uparrow}^{\dagger}(\mathbf{r}) \psi_{\downarrow}^{\dagger}(\mathbf{r}) \psi_{\downarrow}(\mathbf{r}) \psi_{\uparrow}(\mathbf{r}), \qquad (2)$$

with interaction strength $g = 4\pi\hbar^2 a/m$, where *a* denotes the *s*-wave scattering length. A key advantage of the fermionic experiments [10–12] is the ability to use magnetic fields to tune the interactions between different internal states, via a Feshbach resonance [19,20]. Here we work in the weakly interacting regime, which can be realized by working near the zero crossing of the Feshbach resonance.

The spin-orbit Hamiltonian containing terms linear in momentum takes the form

$$\mathcal{H}_{\text{SOC}} = \alpha(\sigma_x p_y - \sigma_y p_x) + \beta(\sigma_x p_x - \sigma_y p_y), \qquad (3)$$

where the first term is the Rashba contribution and the second term is the Dresselhaus contribution, parametrized by the coupling constants α and β , respectively. Here σ_x , σ_y , and σ_z denote Pauli matrices.

In the ultracold-gas setting, spin-orbit coupling is generated by using a pair of Raman beams to drive transitions between two hyperfine states of an atom, while simultaneously imparting a momentum kick [6–13]. The resulting spin-orbit Hamiltonian takes the Rashba equal Dresshelhaus form $\alpha = \beta$, where the magnitude of α is determined by the wavelength of the Raman beams and the angle at which they intersect. The Raman beams also produce a momentum-independent Zeeman field, $\mathcal{H}_Z = -\hbar\Omega_R\sigma_z$, where Ω_R is the strength of the Raman coupling and is proportional to the intensity of the Raman lasers. As we show below, this term plays an important role in the dynamics. The scheme described here was first successfully demonstrated in experiments at NIST using bosonic 87 Rb [6,7]. Recently, a similar scheme was used to generate spin-orbit coupling in a Fermi gas of 6 Li and 40 K [10–12].

As we are primarily motivated by current experiments, we limit ourselves to the case $\alpha = \beta$ in Eq. (3). In this limit, the spin-orbit Hamiltonian can be diagonalized by independently rotating the momentum coordinate and performing a global spin rotation. We denote the diagonal basis as $\{\psi_+, \psi_-\}$. Throughout, we use both the diagonal basis $(\{+, -\})$ and the pseudospin basis $(\{\uparrow, \downarrow\})$ corresponding to the original hyperfine states, depending on the context. The generalization to $\alpha \neq \beta$ is straightforward within our formalism and will be the subject of a future work [36].

All of the work described here is in the *nondegenerate* limit. The only requirement on the temperature T is that it should be smaller than the detuning energy between the magnetic sublevels, so that the gas can be described by a two-level (pseudospin- $\frac{1}{2}$) system. This is readily accomplished as the splitting between hyperfine levels is much larger than the Fermi energy for typical densities [10].

The physics of the weakly interacting gas is dominated by coherent mean-field dynamics which occurs on a time scale $\tau_{\rm mf} \sim (an_0/m)^{-1}$, which is much faster than the time scale for energy exchanging collisions $\tau_{\rm coll} \sim (4\pi a^2 n_0 v)^{-1}$ (v is the characteristic velocity of the particles and n_0 is the density). For a trapped gas of $N \sim 10^5$ ⁶Li atoms at a temperature $T \sim 10^{-6}$ K, the collisionless limit ($\tau_{\rm mf} \ll \tau_{\rm coll}$) corresponds to scattering lengths $a \sim 10a_{\rm B}$, where $a_{\rm B}$ is the Bohr radius. The recent experiment of Cheuk *et al.* is already in this regime [11], while the experiment of Wang *et al.* [10] has stronger interactions of $a \sim 200a_{\rm B}$, which can be tuned near zero using a Feshbach resonance [20].

Mathematically, the weakly interacting gas can be described using a collisionless Boltzmann equation. Following Ref. [37], we use the Heisenberg equations for $\psi_{\sigma}(r,t)$ to derive the equations of motion for the spin dependent Wigner function

$$\begin{aligned} \widehat{\mathbf{F}} &= \begin{pmatrix} f_{\uparrow\uparrow}(\mathbf{p},\mathbf{R},t) & f_{\uparrow\downarrow}(\mathbf{p},\mathbf{R},t) \\ f_{\downarrow\uparrow}(\mathbf{p},\mathbf{R},t) & f_{\downarrow\downarrow}(\mathbf{p},\mathbf{R},t) \end{pmatrix}, \\ f_{\sigma\sigma'}(\mathbf{p},\mathbf{R},t) &= \int d\mathbf{r} e^{i\mathbf{p}\cdot\mathbf{r}} \left\langle \psi_{\sigma}^{\dagger} \left(\mathbf{R}-\frac{\mathbf{r}}{2},t\right) \psi_{\sigma'} \left(\mathbf{R}+\frac{\mathbf{r}}{2},t\right) \right\rangle, \end{aligned}$$
(4)

which is the quantum analog of the classical distribution function. Here **p** represents the momentum, $\mathbf{r} = \mathbf{r}_1 - \mathbf{r}_2$ is the relative coordinate, and $\mathbf{R} = \frac{\mathbf{r}_1 + \mathbf{r}_2}{2}$ is the center-of-mass coordinate.

The diagonal components of \overrightarrow{F} can be integrated in momentum to give the respective spin densities $n_{\sigma\sigma}(\mathbf{R},t) = \langle \psi_{\sigma}^{\dagger}(\mathbf{R},t)\psi_{\sigma}(\mathbf{R},t)\rangle = \int \frac{d\mathbf{p}}{(2\pi)^3} f_{\sigma\sigma}(\mathbf{p},\mathbf{R},t)$, while the off-diagonal components correspond to quantum coherences that are absent in a classical model of a spin- $\frac{1}{2}$ gas.

For the pancake geometry considered here, all the relevant dynamics is 2D. Assuming thermal equilibrium in the longitudinal direction, we decompose the Wigner function in the radial and axial directions as $f_{\sigma\sigma'}(\mathbf{p}, \mathbf{R}, t) = f_{\sigma\sigma'}(p_r, r, t)f(p_z, z)$, where $f(p_z, z) = e^{-\beta(p_z^2/2m + \frac{1}{2}m\omega_z^2z^2)}$, where $\beta = 1/k_BT$.

We obtain a 2D density $n_{\sigma\sigma'}^{2D}(r,t) = \int dp_r f_{\sigma\sigma'}(p_r,r,t) \int dp_z f(p_z,z)$ by integrating out the longitudinal coordinate. Averaging over the *z* direction, we obtain an effective quasi 2D Boltzmann equation,

$$\partial_{t} \overrightarrow{\mathbf{F}} + \frac{\mathbf{p}}{m} \cdot \nabla_{\mathbf{r}} \overrightarrow{\mathbf{F}} - \nabla_{\mathbf{r}} U \nabla_{\mathbf{p}} \overrightarrow{\mathbf{F}} \\ = i [\overrightarrow{\mathbf{V}}, \overrightarrow{\mathbf{F}}] + \frac{1}{2} \{ \nabla_{\mathbf{r}} \overrightarrow{\mathbf{V}}, \nabla_{\mathbf{p}} \overrightarrow{\mathbf{F}} \} + i \alpha p_{+} [\sigma_{+}, \overrightarrow{\mathbf{F}}] \\ + \frac{\alpha}{2} \{ \sigma_{+}, \nabla_{r_{+}} \overleftarrow{\mathbf{F}} \},$$
(5)

where $p_+ = p_x + p_y$, $\sigma_+ = \sigma_x + \sigma_y$, and the interaction potential \overleftrightarrow{V} is [31]

$$\overleftrightarrow{\mathbf{V}} = \begin{pmatrix} g^{2\mathrm{D}}n_{\downarrow\downarrow} - \hbar\Omega_R & -g^{2\mathrm{D}}n_{\uparrow\downarrow} \\ -g^{2\mathrm{D}}n_{\downarrow\uparrow} & g^{2\mathrm{D}}n_{\uparrow\uparrow} + \hbar\Omega_R \end{pmatrix}, \qquad (6)$$

where g^{2D} is an effective 2D interaction strength given by $g^{2D} = 2\sqrt{\pi}\hbar^2 a/(m\Lambda_{\text{th}})$, where $\Lambda_{\text{th}} = \sqrt{2\pi\hbar^2/mk_BT}$. The diagonal components in the interaction matrix arise from forward scattering (Hartree) while the off-diagonal terms arise from exchange interactions (Fock). Commutators and anticommutators are denoted by [,] and {,} respectively.

The single-particle limit of Eq. (5) was derived by Mishchenko and Halperin [5], who used this approach to study the transport properties of a 2D electron gas. Here we generalize the Boltzmann equation to include the effect of \overrightarrow{V} on the phase-space and spin-space evolution of the Wigner function.

In general, Eq. (5) is a nonlinear matrix equation which has to be solved numerically. Below, we assume an initial state which is a stationary state of the Hamiltonian in the absence of spin-orbit coupling ($\alpha = \Omega_R = 0$). We then suddenly turn on the Raman coupling and investigate the resulting out-ofequilibrium dynamics.

III. NONINTERACTING LIMIT

A. $\Omega_R = 0$

The noninteracting limit in ultracold Fermi gases is achieved by working at the zero crossing of a Feshbach resonance. As the magnetic fields corresponding to the zero crossing of the interactions are typically small [11], we do not expect the Raman couplings to deviate appreciably from their zero field values [38]. We first consider the case where upon switching on the Raman coupling the spin-orbit coupling (α) is nonzero, but the Zeeman term $\Omega_R = 0$. While this scenario does not correctly model the present experiments, in this limit the Boltzmann equation is exactly solvable for an arbitrary initial spin state, thus serving as a conceptual starting point. We remark that although we consider only the nondegenerate limit here, our results can be readily generalized to temperatures below the Fermi temperature.

Mathematically, the spin-orbit-coupled problem in the absence of a Zeeman field can be mapped to a single-particle problem in the absence of spin-orbit coupling by a *local* gauge transformation of the fermionic basis wave functions [14]. Here we solve the problem by performing a *global* rotation of the spin and momentum coordinates, which maps the spin-orbit-coupled problem in the absence of a Zeeman field

to a single-particle problem in the presence of a momentumdependent magnetic field.

In order to proceed, we introduce dimensionless position and momentum coordinates $\tilde{r} = r/r_{\text{trap}}$ and $\tilde{p} = p\sqrt{2\pi}\hbar/\Lambda_{\text{th}}$, where $r_{\text{trap}} = \sqrt{\hbar/m\omega_r}$ is the characteristic length scale of motion in the trap and and $\Lambda_{\text{th}} = \sqrt{2\pi\hbar^2/mk_BT}$ is the thermal deBroglie wavelength. We normalize time in units of the radial trapping frequency $\tilde{t} = t/\omega_r$. We also introduce a parameter $\eta = \sqrt{\hbar\omega_r/k_BT}$. As the spin-orbit Hamiltonian only couples to momentum in the $p_x + p_y$ direction, it suffices to consider the evolution of the distribution in the $p_+ = p_x + p_y$ and $r_+ = x + y$ directions of phase space.

Consider an arbitrary initial spin state given by the Wigner distribution function, $\overleftarrow{\mathbf{F}}(\widetilde{p}_+,\widetilde{r}_+,t=0) = e^{-\frac{1}{4}(\widetilde{p}_+^2+\eta^2r_+^2)}\overleftarrow{f}$, where \overleftarrow{f} is a 2 × 2 matrix corresponding to the initial spin state. We omit the p_-,r_- directions for now as they have no dynamics. In the absence of spin-orbit coupling or interactions ($\alpha = \Omega_R = g^{2D} = 0$), the initial state is stationary.

Next, we express the Boltzmann equation in the diagonal basis by performing the *global* transformation $\overleftarrow{F} \rightarrow B^{\dagger} \overleftarrow{f} B$, where the unitary matrix

$$B = \begin{pmatrix} \frac{1}{\sqrt{2}} & \frac{-1+i}{2} \\ \frac{1+i}{2} & \frac{1}{\sqrt{2}} \end{pmatrix}$$

rotates $\sigma_x + \sigma_y$ to $\sqrt{2}\sigma_z$. In the rotated basis, the spin + and - components evolve independently.

The dynamics in phase space can now be solved by making an ansatz for the diagonal components of the rotated spindensity matrix,

$$F_{++/--}(\tilde{p}_{+},\tilde{r}_{+},\tilde{t}) = A_{+/-}e^{-\frac{1}{4}\{[\tilde{p}_{+}\mp a(\tilde{t})]^{2} + \eta^{2}[\tilde{r}_{+}\mp b(\tilde{t})]^{2}\}},$$
 (7)

where a(0) = b(0) = 0. The coefficients $A_{+/-}$ are the diagonal matrix elements of the spin density matrix \overleftarrow{f} after rotation into the $\{+,-\}$ basis.

Substituting the ansatz of Eq. (7) into Eq. (5) we find $a(\tilde{t}) = \tilde{\alpha}\eta[\cos(\tilde{t}) - 1]$ and $b(\tilde{t}) = \tilde{\alpha}\sin(\tilde{t})$, where we have introduced a dimensionless, spin-orbit-coupling constant $\tilde{\alpha} = \sqrt{2\alpha}\sqrt{m/\hbar\omega_r}$, which parametrizes the strength of the spin-orbit coupling relative to the trapping potential. Thus, the rotated spin densities simply perform oscillations in real and momentum space with an amplitude set by the strength of the spin-orbit interaction and period set by the trap frequency.

Similarly, one can solve for the dynamics of the off-diagonal components to find

$$F_{+-}(\tilde{p}_{+},\tilde{r}_{+},\tilde{t})$$

$$= A_{+-}e^{\frac{-4\tilde{\alpha}^{2}[1-\cos(\tilde{t})]}{\eta^{2}}}e^{-\frac{1}{2}([\tilde{p}_{+}-2i\tilde{\alpha}/\eta\sin(\tilde{t})]^{2}+\eta^{2}[\tilde{r}_{+}-2i\tilde{\alpha}/\eta^{2}[1-\cos(\tilde{t})]]^{2})},$$
(8)

where A_{+-} is the off-diagonal matrix element of the spindensity matrix after rotation into the \pm basis. The dynamics of F_{-+} is obtained by replacing $\tilde{\alpha} \rightarrow -\tilde{\alpha}$ in Eq. (8). Unlike the diagonal components, the magnitudes of the off-diagonal components are *not* conserved and oscillate in time.

The corresponding spin densities are found by integrating the above expressions for the Wigner functions in momentum space. By rotating the diagonal basis back into the hyperfine basis $\{\uparrow,\downarrow\}$, one obtains the dynamics of an arbitrary initial spin state.

To illustrate the role of quantum coherence, we consider the dynamics of a fully polarized initial state corresponding to all particles in the \uparrow state. We study the dynamics of the longitudinal magnetization density and the total magnetization:

$$m_{z}(\mathbf{r},t) = \int d\mathbf{p}[f_{\uparrow\uparrow}(\mathbf{p},\mathbf{r},t) - f_{\downarrow\downarrow}(\mathbf{p},\mathbf{r},t)],$$

$$M(t) = \int d\mathbf{r}m_{z}(r,t).$$
(9)

The zero-temperature dynamics of the total magnetization for this initial state was considered previously by Stanescu, Zhang, and Galitski [15]. By exactly solving for the quantum dynamics in a trap, they demonstrated that the total magnetization exhibits collapse and revival dynamics and produced analytic formulas for the total magnetization in weak spin-orbit coupling limit.

Here we show that similar dynamics also occurs in the nondegenerate gas, which is much more readily accessible in experiments. Furthermore, we obtain analytic expressions for the total magnetization for *arbitrary* values of the spin-orbit coupling.

Rotating the spin-polarized state to the diagonal basis, one finds that the density matrix has both diagonal and off-diagonal matrix elements ($A_+ = A_- = 1/2$ and $A_{+-} = -\frac{1-i}{2\sqrt{2}}$), whose dynamics is given by Eqs. (7) and (8). Transforming back to the hyperfine basis and integrating over momentum, the longitudinal magnetization density takes the form:

$$m_{z}(\tilde{r}_{+},\tilde{t}) \sim e^{-\frac{1}{2}\eta^{2}\tilde{r}_{+}^{2} - \frac{\tilde{\alpha}^{2}}{\eta^{2}}[1 - \cos(2\tilde{t})]} \cos\{2\tilde{r}_{+}\tilde{\alpha}[1 - \cos(\tilde{t})]\}$$
(10)

and the total magnetization is

$$M(\tilde{t}) \sim \sqrt{\frac{2\pi}{\eta}} e^{-\frac{4\tilde{a}^2}{\eta^2}[1-\cos(\tilde{t})]},\tag{11}$$

where we have ignored an overall normalization factor resulting from integration over momentum. The expression for the transverse magnetization density is rather cumbersome, but the total transverse magnetization remains zero at all times.

From Eq. (11), it is clear that the total longitudinal magnetization exhibits collapse and revival dynamics in time with a period which depends *only* on the trapping potential and is completely independent of the temperature or the spin-orbit coupling $\tilde{\alpha} \ll 1$, our expression reads $M \sim 1 - 4\tilde{\alpha}^2/\eta^2[1 - \cos(\tilde{t})]$ [15]. In this limit, the magnetization exhibits sinusoidal oscillations with an amplitude which is proportional to $8\tilde{\alpha}^2/\eta^2$. For strong spin-orbit coupling, $\tilde{\alpha} \gg 1$, the magnetization becomes strongly peaked near $t = 2\pi n/\omega_r$, where *n* is an integer, and decays exponentially, away from these points.

The collapse and revival of the total magnetization is a trap effect. In a homogeneous system, the momentum-dependent spin-orbit magnetic field will simply cause the spins to dephase, particles with different momenta will precess at different rates, and the total magnetization will go to zero irreversibly on a time scale set by the spin-orbit coupling strength [14]. It is also important to emphasize that the collapse and revival in the total magnetization described above has a different origin from what is observed in the experiments of Wang *et al.* [10]. We discuss this in more detail later.

To understand the origin of the magnetization oscillations, we now turn to the dynamics of the longitudinal and transverse magnetization density following the ramp. In a trapped geometry, from Eq. (10), we find that the longitudinal magnetization density exhibits periodic oscillations in space and in time. The temporal oscillations have a period of $t = 2\pi/\omega_r$, while at $t = \pi/\omega_r$ the spatial oscillations have a characteristic wavelength of $\lambda_{\rm sh} = \pi r_{\rm trap}/(2\tilde{\alpha})$ where $r_{\rm trap} = \sqrt{\hbar/m\omega_r}$.

In Fig. 1 we plot the magnetization density normalized to the initial magnetization at the center [m(r = 0, t = 0)] as a function of time for the parameters above. Brighter colors indicate positive magnetization while darker colors indicate negative values of m_z . We choose a rather weak spin-orbit coupling strength in order to enhance the wavelength of the spatial oscillations at $t = \pi/\omega_r$. The transverse components of the spin are indicated by arrows whose lengths correspond to the magnitude of the spin vector in the x-y plane. At t = 0, all spins are pointing in the \uparrow direction indicated by the bright region in the density plot. Over time a transverse component develops and a spin helix emerges. At $t = \pi/\omega_r$, the spin



FIG. 1. (Color online) Time evolution of the longitudinal (density plot) and transverse magnetization (arrows) in the $r_{+} = x + y$ direction, following a sudden ramp of the spin-orbit coupling strength. The Zeeman term is set to zero here ($\Omega_R = 0$). Brighter colors indicate positive magnetization (\uparrow) and darker colors indicate negative magnetization (\downarrow) . The arrows indicate the direction of the magnetization in the x-y plane, and the length of the arrows indicate the magnitude of the transverse spin normalized to the total spin. At t = 0 all the spins are pointing in the z direction (all atoms in the \uparrow state). Spin-orbit coupling causes the atoms to precess in time, and at half the trapping period a spin helix is produced. The wavelength of the helix at $t = \pi/\omega_r$ depends only on the spin-orbit interaction and is $\lambda_{\rm sh} = \pi r_{\rm trap}/2\tilde{\alpha}$, where $r_{\rm trap} = \sqrt{\hbar/m\omega_r}$. To clearly illustrate this effect, we choose a weak spin-orbit coupling of $\tilde{\alpha} = \sqrt{2\alpha} \sqrt{m/\hbar\omega_r} = 0.125$ and $\eta = \sqrt{\hbar\omega/k_BT} = 0.25$ in these simulations. The initial state is recovered after $t = 2\pi/\omega_r$.

oscillations have maximum amplitude, with a wavelength of $\Lambda_{\rm sh}/r_{\rm trap} = \pi/2\tilde{\alpha}$.

Energy-conserving dynamics in phase space implies that the momentum of a particle is linked to its position. Moreover, the spin of an atom is linked to its momentum via the spin-orbit coupling. The spin-orbit Hamiltonian shifts the minimum of the dispersion to finite momenta. In a harmonically confined system, isoenergy contours are circles in phase space that are shifted to finite momenta due to the spin-orbit coupling. A wave packet polarized in the \uparrow direction centered at r = p = 0will follow the isoenergy contours in phase space, while simultaneously rotating in spin space. At $t = \pi/\omega_r$ the atomic wave packet is centered around r = 0 in real space, and in order to conserve the total energy, the distribution will be centered around $\tilde{p}_{\pm} = \pm 2\tilde{\alpha}\eta$ in momentum space. As a result, atoms with opposite momenta precess in opposite directions in spin space, producing a spin helix. At $t = 2\pi/\omega_r$, the atoms return to their original distribution in real and momentum space, and the initial state is recovered. The spin helix has a smaller net magnetization as compared to the fully polarized initial state, thus explaining the oscillations in the net magnetization.

Experimentally, the spin density can be imaged *in situ* using phase-contrast imaging [39,40]. The wavelength of the spin helix is set by the spin-orbit interaction which is determined by the wavelength of the Raman beams. For the parameters used in the experiment of Cheuk *et al.* [11], $\Lambda_{sh} \sim 0.5 \,\mu$ m, which may be below the experimental resolution. However, the wavelength can be increased by decreasing the spin-orbit coupling strength. At present, experiments on spin-orbit-coupled cold gases suffer from extremely short lifetimes (<500 ms) due to the large heating rates resulting from inelastic light scattering from the Raman beams [41]. For typical trapping potentials, the time scale for the appearance of the spin helix is $t = \pi/\omega_r \sim 50$ ms, so the short lifetimes may not be a major limitation in observing the spin helix.

B. $\Omega_R \neq 0$

We now turn to the experimentally relevant case where upon suddenly switching on the Raman beams, the atoms also experience a constant, momentum-independent magnetic field, parametrized by a dimensionless parameter $\tilde{B} = \Omega_R/\omega_r$. In this limit, the problem cannot be solved analytically as the Zeeman and spin-orbit terms in Eq. (5) do *not* commute. Instead, we numerically integrate the Boltzmann equation on a 4D grid in phase space. We choose a 20 × 20 lattice in **R** and **p** with a spatial resolution of $\delta r = 2r_{\text{trap}}$, where $r_{\text{trap}} = \sqrt{\hbar/m\omega_r}$ is the characteristic length scale of motion in the trap, and momentum resolution of $\delta p = 0.62\pi/\Lambda_{\text{th}}$. The integration is done using a split-step method that conserves total particle number and the total energy to high accuracy for sufficiently small time steps.

For simplicity we choose a fully polarized initial state, which is stationary in the absence of interactions or spin-orbit coupling. We then consider two limits upon switching on the Raman beams: $\tilde{B} \sim \tilde{\alpha}$ and $\tilde{B} > \tilde{\alpha}$. The two parameters can be controlled independently as the magnitude of $\tilde{\alpha}$ is set by the wavelength of the Raman beams, and Ω_R is set by the laser intensity. The resulting dynamics of the total magnetization as well as the magnetization density is plotted in Fig. 2.



FIG. 2. (Color online) (Left) Time evolution of the total magnetization in the system [see Eq. (9)] normalized to the total particle number for three different values of the dimensionless Zeeman coupling $\tilde{B} = \Omega_R/\omega_r$. In each case, we fix $\tilde{\alpha} = \sqrt{2\alpha}\sqrt{m/\hbar\omega_r} = 0.25$. From top to bottom: (black) $\tilde{B} = 0$; (blue) $\tilde{B} = 0.25$; (red) $\tilde{B} = 1$. (Right) Magnetization density along the $r_+ = x + y$ direction for the same values of the spin-orbit coupling strength and Zeeman field as in the left figure at fixed time $t = \pi/\omega_r$. The magnetization densities are normalized to the total central density denoted $n = n(r = 0, t = 0) = \int d\mathbf{p} [f_{\uparrow\uparrow}(\mathbf{p}, 0, 0) + f_{\downarrow\downarrow}(\mathbf{p}, 0, 0)]$. Solid curves represent the magnetization density in the *z* direction [Eq. (9)], while the dashed curves indicate the magnetization density in the *x* direction: $m_x(\mathbf{r}, t = \pi/\omega_r) = \int d\mathbf{p} [f_{\uparrow\downarrow}(\mathbf{p}, \mathbf{r}, \pi/\omega_r) + f_{\downarrow\uparrow}(\mathbf{p}, \mathbf{r}, \pi/\omega_r)]$.

As shown in Fig. 2, the addition of a Zeeman term causes the magnetic field oscillations to decay over time. At long times, the total magnetization acquires a new steady-state value which is smaller than 1. As the strength of the Zeeman field is increased, the fully polarized initial state becomes increasingly stable, and the total magnetization remains close to 1 at long times with small oscillations.

The dynamics can be understood as follows. In the presence of a Zeeman field, each spin precesses about a new magnetic field, which is the sum of the spin-orbit magnetic field and the Zeeman field and is tilted away from the *x*-*y* plane by an angle $\sin(\theta_p) = \tilde{B}/\sqrt{\tilde{B}^2 + (\tilde{\alpha}\eta\tilde{p}_+)^2}$. As the initial state has a Gaussian distribution of atoms with different momenta with spins pointing in the *z* direction, atoms with momenta greater than $p > \tilde{p}_{crit} = \tilde{B}/\tilde{\alpha}\eta$ will primarily experience the spin-orbit magnetic field and precess about the *x*-*y* plane, while atoms with momenta $p < p_{crit}$ will predominantly see the Zeeman field and precess about the *z* axis.

In a thermal gas, the characteristic width of the momentum distribution is set by $p_{\text{th}} = \sqrt{2\pi\hbar}/\Lambda_{\text{th}}$. If the Zeeman field is weak compared to the spin-orbit strength ($\alpha p_{\text{th}} \gg \hbar \Omega_R$), the majority of the atoms still experience the spin-orbit magnetic field, and the spin-density wave is preserved (as shown in the blue curves in Fig. 2) for the first few oscillations. On longer time scales the magnetic field affects the spin precession of even the atoms with $p \gg p_{\text{crit}}$ and the spin-density wave disappears. On very long times the system settles into a new steady state with a lower net magnetization.

On the other hand, if the Zeeman field is strong ($\alpha p_{\text{th}} \ll \hbar \Omega_R$), the majority of the atoms experience a net magnetic field which is only slightly tilted away from the *z* axis. As the initial state is fully polarized, the total magnetization remains close to 1 at all times with rapid oscillations whose frequency is given by Ω_R . The transverse spin density remains small at all times (see the dashed red curve in Fig. 2) and no spin helix is found.

The spin helix is the result of the interplay between dynamics in spin space and dynamics in phase space. Observing the spin helix therefore requires that the time scales for spin dynamics be comparable to the time scales for dynamics in phase space.

In the presence of a large Zeeman field, as in the experiments of Wang *et al.* [10] ($E_Z \sim \text{kHz} \gg \hbar \omega_r \sim 100 \text{ Hz}$), the time scale for spin dynamics is set by the Zeeman term, which is much too short for any dynamics to occur in phase space. Thus, while the total magnetization shows collapse and revivals (on a time scale set by $t \sim 2\pi/E_Z \sim 1$ ms), the magnetization density in real space is unaffected. By contrast, the collapse and revival dynamics discussed in Sec. III A occurs because the initial state is out of equilibrium in *phase space*, when the Raman coupling is switched on. As a result, the magnetization oscillates on a time scale set by $t \sim 2\pi/\omega_r \sim 10$ ms, and a spin spiral appears.

IV. INTERACTIONS

We now consider the effect of weak interactions on the dynamics discussed above. As discussed previously, we work in the the collisionless limit (sometimes referred to as the Knudsen regime), which is valid as long as $\tau_{mf} \ll \tau_{coll}$.

Spin waves in dilute, quantum gases were studied theoretically by Bashkin and others [24–27] and observed in experiments on spin-polarized hydrogen [28]. More recently, Du *et al.* [29] explored similar physics in a weakly interacting, thermal gas of ⁶Li atoms. Here we explore spin waves in a collisionless thermal gas with spin-orbit coupling.

To compare with the results of Sec. III, we consider a fully polarized initial state which is the stationary state of the Boltzmann equation in the absence of spin-orbit coupling or interactions. We then simultaneously switch on the interactions, and the spin-orbit coupling. We define a dimensionless interaction strength $\tilde{g} = g^{2D}n/\hbar\omega_r$, where *n* is the total initial density at the trap center $n = n(r = 0, t = 0) = \int d\mathbf{p}[f_{\uparrow\uparrow}(\mathbf{p},0,0) + f_{\downarrow\downarrow}(\mathbf{p},0,0)]$. For typical trapping potentials used in experiments, $\omega_r \sim 2\pi \times 100$ Hz, the Knudsen regime corresponds to $\tilde{g} \sim 0.1$ [29]. We choose $\tilde{g} = 0.25$. Furthermore, as shown in Sec. III B, a large Zeeman field ($\tilde{B} \gg \tilde{\alpha}$) stabilizes the spin-polarized state and produces little spin dynamics. Hence, we consider $\tilde{B} = \tilde{\alpha} = 0.25$.

In Fig. 3 we plot the total magnetization density along the r_+ direction and the total net magnetization as a function of time. As is apparent from the figure, the inclusion of interaction drives large magnetization oscillations on a characteristic time scale much larger than $t = \pi/\omega_r$. This time scale grows as the interaction strength is increased. In real space, the magnetization minimum is manifested as a largeamplitude longitudinal spin wave. The panels on the right show the magnetization density in the longitudinal and transverse



FIG. 3. (Color online) (Top left) Density plot showing the magnetization density along r_+ [see Eq. (10)], normalized to the total initial central density as a function of time. The parameters are $\tilde{g} = \tilde{B} = \tilde{\alpha} = 0.25$. (Recall that $\tilde{g} = g^{2D}n/\hbar\omega_r$, $\tilde{B} = \Omega_R/\omega_r$, and $\tilde{\alpha} = \sqrt{2\alpha} \sqrt{m/\hbar\omega_r}$.) Brighter colors indicate positive magnetization, while darker colors indicate negative magnetization. (Bottom left) Total magnetization in the entire system as a function of time for three different values of \tilde{g} . In each plot, $\tilde{B} = \tilde{\alpha} = 0.25$. From top to bottom $\tilde{g} = 0$ (dashed), $\tilde{g} = 0.25$ (solid), and $\tilde{g} = 0.5$ (dotted). (Panels on right) Longitudinal (solid) and transverse (dashed) magnetization densities in the r_{+} direction at different times. The top panel shows the spin helix without interactions for $\tilde{B} = \tilde{\alpha} = 0.25$ (same as the central panel in Fig. 2) for comparison. The central and bottom panels are snapshots of the longitudinal and transverse magnetization at time $t = 2.5\pi/\omega_r$ and $t = 5\pi/\omega_r$ for the same parameters as in the top left.

directions (solid and dashed, respectively) at two different times $t = 2.5\pi/\omega_r$ (red) and $t = 5\pi/\omega_r$ (blue) for parameters corresponding to the density plot. The spin densities for the spin helix [cf. Fig. 2 (right) center panel] for the same values of \tilde{B} and $\tilde{\alpha}$, but $\tilde{g} = 0$, are shown in the top panel for comparison. The amplitude of the spin wave seen here is much larger than that observed in previous studies of the spin- $\frac{1}{2}$ gas in the absence of spin-orbit coupling [29].

Interactions alter the physics of the noninteracting gas in two crucial ways. Forward scattering modifies the isoenergy contours in phase space: Spin \uparrow atoms experience a mean field proportional to the density of the \downarrow atoms and vice versa. More importantly, due to the exchange interaction, when two atoms collide, they precess about the common axis of their total spin. As argued by Lhuillier and Laloë, it is this effect that gives rise to spin waves in the collisionless gas [25]. In the experiments of Du *et al.* [29], the initial state was polarized along the *x* direction, and spin dynamics was the result of the negligible difference in the trapping potentials experienced by the spin \uparrow and \downarrow atoms. Consequently, the amplitude of the resulting spin wave was very small, $m_z(\mathbf{r})/n_0 \ll 1$ [29].

By contrast, in the present case the initial state is polarized along z. Absent spin-orbit coupling, this state has *no* dynamics as it is stationary with respect to the Zeeman field and

PHYSICAL REVIEW A 88, 033613 (2013)

the interaction term. Spin-orbit coupling, however, causes the spins to precess about the x-y plane with a precession rate proportional to the momentum. Over time, atoms with different momenta precess at different rates and can collide with one another. The atomic collisions subsequently lead to a dynamical spin segregation in real space [25]. Unlike the experiments of Du *et al.* [29], the initial spin precession in our system occurs due to spin-orbit coupling, which is not a small effect. Consequently, the amplitude of the resulting spin wave is also larger than what was found in Ref. [29].

We remark that the amplitude of the total magnetization oscillations, and the associated spin texture in real space, depends nonmonotonically on the strength of the Zeeman field. For $\tilde{B} \ll \tilde{\alpha}$, and weak interactions $\tilde{g} \ll 1$, the magnetization oscillations are only slightly larger than those studied in Sec. III A, and the real space magnetization density resembles the spin helix shown in Fig. 3 (top right). The magnetization oscillations reach a maximum at some \tilde{B}_{crit} , which depends on \tilde{g} and $\tilde{\alpha}$. Upon further increase of \tilde{B} , the spin \uparrow state becomes stable and there is no spin dynamics.

V. SUMMARY AND CONCLUSIONS

In summary, we studied the spin dynamics in a weakly interacting, nondegenerate Fermi gas following a sudden ramp of the spin-orbit coupling strength. In the absence of interactions and a Zeeman field, we produced analytic expressions for the total magnetization and the magnetization density for arbitrary values of the spin-orbit coupling, trap frequency, and temperature. We argued that a fully polarized initial state will give rise to a spin helix on a time scale set by the trapping period with a wavelength that depends on the ratio of the spin-orbit interaction to the trap frequency. For the high-temperature gas, we generalized the analytic results obtained by Stanescu *et al.* [15] to arbitrary spin-orbit coupling strength. We then numerically studied the spin dynamics in the presence of interactions and a Zeeman field, highlighting the role played both separately and collectively by these terms. For weak Zeeman fields the spin helix is preserved but it disappears for very large Zeeman fields, as the fully polarized initial state becomes increasingly stable. In the presence of interactions, however, the dynamics is more complicated. Interactions tend to *enhance* the amplitude of the spin oscillations. We explain this effect as a dynamical spin segregation driven by exchange.

Finally, we briefly comment on the role of Fermi statistics in our calculations. In principle, all the results discussed in this paper apply equally well to bosons. The spin helix discussed in Sec. III is a single-particle effect and should also occur in a bosonic system. Importantly, it is a *dynamical* effect and should be contrasted with the spin stripe phase predicted to occur in equilibrium spin-orbit-coupled Bose-Einstein condensates by Wang *et al.* [42] and Ho and Zhang [43].

To conclude, our work shows that a seemingly one-particle term in the Hamiltonian (i.e., spin-orbit coupling), when coupled with coherent nonequilibrium dynamics of the Fermi gas, could manifest rather nonobvious and complex (and more importantly, observable) behavior in cold atoms and molecules even at relatively high temperature where the quantum degeneracy of the system is not of any key importance. The dynamical physics we predict here is unlikely to manifest itself in any solid-state systems (and can only be seen in atomic gases) although the basic concept of spin-orbit coupling originates in solids, as discussed in detail in Refs. [1,2].

ACKNOWLEDGMENTS

It is a pleasure to thank Juraj Radic for his insights and his careful reading of the manuscript. We also gratefully acknowledge discussions with Victor Galitski, I. V. Tokatly, and E. Y. Sherman. This work is supported by JQI-NSF-PFC, AFOSR-MURI, and ARO-MURI.

- [1] J. E. Hirsch, Phys. Rev. Lett. 83, 1834 (1999).
- [2] I. Zutic, J. Fabian, and S. Das Sarma, Rev. Mod. Phys. 76, 323 (2004).
- [3] S. Datta and B. Das, Appl. Phys. Lett. 56, 665 (1990); J.
 Wunderlich *et al.*, Science 330, 1801 (2010).
- [4] M. I. D'yakonov *et al.*, Sov. Phys. JETP **63**, 655 (1986);
 E. G. Mishchenko, A. V. Shytov, and B. I. Halperin, Phys. Rev. Lett. **93**, 226602 (2004); A. A. Burkov, A. S. Nunez, and A. H. MacDonald, Phys. Rev. B **70**, 155308 (2004);
 D. D. Awschwalom and M. E. Flatté, Nat. Phys. **3**, 153 (2007); B. A. Bernevig and J. Hu, Phys. Rev. B **78**, 245123 (2008).
- [5] E. G. Mishchenko and B. I. Halperin, Phys. Rev. B 68, 045317 (2003).
- [6] Y. K. Lin, K. Jimenez-Garcia, and I. B. Spielman, Nature (London) 471, 83 (2011).
- [7] Y.-J. Lin, R. L. Compton, A. R. Perry, W. D. Phillips, J. V. Porto, and I. B. Spielman, Phys. Rev. Lett. **102**, 130401 (2009).

- [8] J. Y. Zhang, S. C. Ji, Z. Chen, L. Zhang, Z. D. Du, B. Yan, G. S. Pan, B. Zhao, Y. J. Deng, H. Zhai, S. Chen, and J. W. Pan, Phys. Rev. Lett. 109, 115301 (2012).
- [9] J.-Yi. Zhang, S.-C. Ji, L. Zhang, Z.-D. Du, W. Zheng, Y.-J. Deng, H. Zhai, S. Chen, and J.-W. Pan, arXiv:1305.7054.
- [10] P. Wang, Z-Q. Yu, Z. Fu, J. Miao, L. Huang, S. Chai, H. Zhai, and J. Zhang, Phys. Rev. Lett. **109**, 095301 (2012).
- [11] L. W. Cheuk, A. T. Sommer, Z. Hadzibabic, T. Yefsah, W. S. Bakr, and M. W. Zwierlein, Phys. Rev. Lett. 109, 095302 (2012).
- [12] R. A. Williams, M. C. Beeler, L. J. LeBlanc, K. Jimenez-Garcia, and I. B. Spielman, Phys. Rev. Lett. 111, 095301 (2013).
- [13] Z. Fu, L. Huang, Z. Meng, P. Wang, L. Zhang, S. Zhang, H. Zhai, P. Zhang, and J. Zhang, arXiv:1306.4568.
- [14] I. V. Tokatly and E. Y. Sherman, Phys. Rev. A 87, 041602(R) (2013).
- [15] T. D. Stanescu, C. Zhang, and V. Galitski, Phys. Rev. Lett. 99, 110403 (2007).
- [16] J. Larson, B. M. Anderson, and A. Altland, Phys. Rev. A 87, 013624 (2013).

- [17] B. A. Bernevig, J. Orenstein, and S.-C. Zhang, Phys. Rev. Lett. 97, 236601 (2006).
- [18] J. D. Koralek et al., Nature (London) 458, 610 (2009).
- [19] I. Bloch, J. Dalibard, and W. Zwerger, Rev. Mod. Phys. 80, 885 (2008).
- [20] W. Ketterle and M. W. Zwierlein, Soc. Ital. Phys. 31, 247 (2008).
- [21] J. D. Sau, R. Sensarma, S. Powell, I. B. Spielman, and S. Das Sarma, Phys. Rev. B 83, 140510(R) (2011).
- [22] C. Zhang, S. Tewari, R. M. Lutchyn, and S. Das Sarma, Phys. Rev. Lett. 101, 160401 (2008).
- [23] A. Kitaev, Ann. Phys. 303, 2 (2003).
- [24] E. P. Bashkin, JETP Lett. **33**, 8 (1981).
- [25] C. Lhuillier and F. Laloë, J. Phys (Paris) 43, 197 (1982); 43, 225 (1982).
- [26] W. J. Gully and W. J. Mullin, Phys. Rev. Lett. 52, 1810 (1984).
- [27] B. Castaing, Physica B 126, 212 (1984).
- [28] B. R. Johnson, J. S. Denker, N. Bigelow, L. P. Levy, J. H. Freed, and D. M. Lee, Phys. Rev. Lett. 52, 1508 (1984).
- [29] X. Du, L. Luo, B. Clancy, and J. E. Thomas, Phys. Rev. Lett. 101, 150401 (2008).
- [30] H. J. Lewandowski, D. M. Harber, D. L. Whitaker, and E. A. Cornell, Phys. Rev. Lett. 88, 070403 (2002).

- [31] S. S. Natu and E. J. Mueller, Phys. Rev. A 79, 051601(R) (2009).
- [32] U. Ebling, A. Eckardt, and M. Lewenstein, Phys. Rev. A 84, 063607 (2011).
- [33] F. Piéchon, J. N. Fuchs, and F. Laloë, Phys. Rev. Lett. 102, 215301 (2009).
- [34] M. Ö. Oktel and L. S. Levitov, Phys. Rev. Lett. **88**, 230403 (2002).
- [35] M. I. Dyakonov and V. I. Perel, Zh. Eksp. Teor. Fiz 60, 1954 (1971) [Sov. Phys. JETP 33, 1053 (1971)].
- [36] J. Radic, S. S. Natu, and V. Galitski (unpublished).
- [37] *Quantum Statistical Mechanics*, edited by L. P. Kadanoff and G. Baym (Benjamin, New York, 1962).
- [38] R. Wei and E. J. Mueller, Phys. Rev. A 87, 042514 (2013).
- [39] Y. Shin, M. W. Zwierlein, C. H. Schunck, A. Schirotzek, and W. Ketterle, Phys. Rev. Lett. 97, 030401 (2006).
- [40] L. E. Sadler, J. M. Higbie, S. R. Leslie, M. Vengalattore, and D. M. Stamper-Kurn, Phys. Rev. Lett. 98, 110401 (2007).
- [41] I. B. Spielman, Phys. Rev. A **79**, 063613 (2009).
- [42] C. Wang, C. Gao, C.-M. Jian, and H. Zhai, Phys. Rev. Lett. 105, 160403 (2010).
- [43] T.-L. Ho and S. Zhang, Phys. Rev. Lett. 107, 150403 (2011).