

## Evolution from BCS to BEC superfluidity in the presence of spin-orbit coupling

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We discuss the evolution from BCS to Bose-Einstein condensate (BEC) superfluids in the presence of spin-orbit coupling for a balanced mixture of ultracold fermions. The dependence of several thermodynamic properties, such as chemical potential, order parameter, pressure, entropy, isothermal compressibility, and spin susceptibility tensor on the spin-orbit coupling and interaction parameter at low temperatures are analyzed. We studied both equal Rashba and Dresselhaus (ERD) and the Rashba-only (RO) spin-orbit coupling. Comparisons between the two cases reveal several striking differences in the corresponding thermodynamic quantities. Finally, we propose measuring the isothermal compressibility and spin susceptibility as a way of detecting the effects of the spin-orbit coupling.

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Superfluidity is a ubiquitous phenomenon that is encountered in nearly every area of physics, including condensed-matter physics, nuclear physics, astrophysics, and atomic and molecular physics. Superflow results from strong correlations between particles, which for any given interacting Fermi system could not be controlled externally until the recent advent of ultracold atoms. In standard condensed matter, there is a continuous search for new charged superfluids (superconductors) since the type and strength of interactions cannot be tuned even within the same class of materials. In the cases of nuclear matter and neutron stars, the tunability of interactions is extremely difficult. However, the situation is much more favorable for ultracold Fermi atoms, where the ability to control interactions between particles, via Feshbach resonances, has been demonstrated in experimental studies of the so-called crossover from BCS to Bose-Einstein condensate (BEC) superfluidity.

Further control of interactions is now possible through newly developed experimental techniques that allow the production of fictitious magnetic fields which couple to neutral bosonic atoms [1,2]. These fictitious magnetic fields are generated through an all-optical process but produce real effects like the creation of vortices in the superfluid state of bosons. Furthermore, artificial spin-orbit coupling has also been produced in neutral bosonic systems [3] where the strength of the coupling can be controlled optically. In principle, the same techniques can be applied to ultracold fermions [3,4], which, when coupled with the control over the interaction using Feshbach resonances, allows for the exploration of superfluidity not only as a function of interactions but also as a function of fictitious magnetic fields [6] or as a function of spin-orbit coupling as discussed here. An introduction to the effects of controllable fictitious magnetic and spin-orbit fields can now be found in the literature [7].

It is in anticipation of experiments involving spin-orbit coupling in fermionic atoms such as  ${}^6\text{Li}$ ,  ${}^{40}\text{K}$ ,  ${}^{171}\text{Yb}$ , and  ${}^{173}\text{Yb}$  that we discuss here the evolution from BCS to BEC superfluidity in the presence of controllable spin-orbit couplings for balanced fermions in three dimensions. We investigate spin-orbit effects with Dresselhaus [8] and Rashba [9] terms and analyze several thermodynamic quantities including the order parameter, chemical potential, thermodynamic poten-

tial, entropy, pressure, isothermal compressibility, and spin susceptibility tensor as a function of spin-orbit coupling and interaction parameter at low temperatures.

*Hamiltonian.* To address the problem of the evolution from BCS to BEC superfluidity in the presence of spin-orbit fields for balanced or imbalanced Fermi-Fermi mixtures, we start with the generic Hamiltonian density

$$\mathcal{H}(\mathbf{r}) = \mathcal{H}_0(\mathbf{r}) + \mathcal{H}_I(\mathbf{r}). \quad (1)$$

The single-particle Hamiltonian density is

$$\mathcal{H}_0(\mathbf{r}) = \sum_{\alpha\beta} \psi_{\alpha}^{\dagger}(\mathbf{r}) [\hat{K}_{\alpha} \delta_{\alpha\beta} - h_i(\mathbf{r}) \sigma_{i,\alpha\beta}] \psi_{\beta}(\mathbf{r}), \quad (2)$$

where  $\hat{K}_{\alpha} = -\nabla^2/(2m_{\alpha}) - \mu_{\alpha}$  is the kinetic energy in reference to the chemical potential  $\mu_{\alpha}$ , and  $h_i(\mathbf{r})$  is the spin-orbit field along the  $i$  direction ( $\alpha = \uparrow, \downarrow$ ,  $i = x, y, z$ ). The interaction term is  $\mathcal{H}_I(\mathbf{r}) = -g \psi_{\uparrow}^{\dagger}(\mathbf{r}) \psi_{\downarrow}^{\dagger}(\mathbf{r}) \psi_{\downarrow}(\mathbf{r}) \psi_{\uparrow}(\mathbf{r})$ , where  $g$  is a contact interaction. In this Rapid Communication, we set  $\hbar = k_B = 1$ .

*Effective action.* The partition function at temperature  $T$  is  $Z = \int \mathcal{D}[\psi, \psi^{\dagger}] \exp(-S[\psi, \psi^{\dagger}])$  with action

$$S[\psi, \psi^{\dagger}] = \int d\tau d\mathbf{r} \left[ \sum_{\alpha} \psi_{\alpha}^{\dagger}(\mathbf{r}) \frac{\partial}{\partial \tau} \psi_{\alpha}(\mathbf{r}) + \mathcal{H}(\mathbf{r}, \tau) \right]. \quad (3)$$

Using the standard Hubbard-Stratanovich transformation that introduces the pairing field  $\Delta(\mathbf{r}, \tau) = g \langle \psi_{\downarrow}(\mathbf{r}, \tau) \psi_{\uparrow}(\mathbf{r}, \tau) \rangle$ , we can write the intermediate action  $S_{\text{int}}[\psi, \psi^{\dagger}, \Delta, \Delta^{\dagger}] = S_{\text{no}}[\psi, \psi^{\dagger}] + S_1[\psi, \psi^{\dagger}, \Delta, \Delta^{\dagger}]$ , where the no-interaction action is

$$S_{\text{no}}[\psi, \psi^{\dagger}] = \int d\tau d\mathbf{r} \left[ \sum_{\alpha} \psi_{\alpha}^{\dagger}(\mathbf{r}) \frac{\partial}{\partial \tau} \psi_{\alpha}(\mathbf{r}) + \mathcal{H}_0(\mathbf{r}, \tau) \right]$$

and the action due to the auxiliary field is

$$S_1 = \int d\tau d\mathbf{r} \left[ \frac{|\Delta(\mathbf{r}, \tau)|^2}{g} - \Delta \psi_{\uparrow}^{\dagger} \psi_{\downarrow}^{\dagger} - \Delta^{\dagger} \psi_{\downarrow} \psi_{\uparrow} \right].$$

Using the four-dimensional vector representation  $\Psi^{\dagger}(\mathbf{r}, \tau) = \{\psi_{\uparrow}^{\dagger}, \psi_{\downarrow}^{\dagger}, \psi_{\uparrow}, \psi_{\downarrow}\}$ , the intermediate action becomes

$$S_{\text{int}} = \int d\tau d\mathbf{r} \left[ \frac{|\Delta(\mathbf{r}, \tau)|^2}{g} + \frac{1}{2} \Psi^{\dagger} \mathbf{M} \Psi + \frac{1}{2} (\tilde{K}_{\uparrow} + \tilde{K}_{\downarrow}) \right].$$

The  $4 \times 4$  matrix  $\mathbf{M}$  is

$$\mathbf{M} = \begin{pmatrix} \partial_\tau + \tilde{K}_\uparrow & -h_\perp & 0 & -\Delta \\ -h_\perp^* & \partial_\tau + \tilde{K}_\downarrow & \Delta & 0 \\ 0 & \Delta^* & \partial_\tau - \tilde{K}_\uparrow & h_\perp^* \\ -\Delta^* & 0 & h_\perp & \partial_\tau - \tilde{K}_\downarrow \end{pmatrix}, \quad (4)$$

where  $h_\perp = h_x - ih_y$  corresponds to the transverse component of the spin-orbit field,  $h_z$  corresponds to the parallel component with respect to the quantization axis  $z$ ,  $\tilde{K}_\uparrow = \hat{K}_\uparrow - h_z$ , and  $\tilde{K}_\downarrow = \hat{K}_\downarrow + h_z$ . Integration over the fields  $\Psi$  and  $\Psi^\dagger$  leads to the effective action

$$S_{\text{eff}} = \int d\tau d\mathbf{r} \left[ \frac{|\Delta(\mathbf{r}, \tau)|^2}{g} - \frac{T}{2V} \ln \det \frac{\mathbf{M}}{T} + \tilde{K}_+ \delta(\mathbf{r} - \mathbf{r}') \right], \quad (5)$$

where  $\tilde{K}_+ = (\tilde{K}_\uparrow + \tilde{K}_\downarrow)/2$  is the average kinetic energy and  $V$  is the volume of the system.

*Saddle-point approximation.* To proceed, we use the saddle-point approximation  $\Delta(\mathbf{r}, \tau) = \Delta_0 + \eta(\mathbf{r}, \tau)$  and separate the matrix  $\mathbf{M}$  into two parts. The first one is the saddle point matrix  $\mathbf{M}_0$ , where the transformation  $\Delta(\mathbf{r}, \tau) \rightarrow \Delta_0$  takes  $\mathbf{M} \rightarrow \mathbf{M}_0$ . The second one is the fluctuation matrix  $\mathbf{M}_F = \mathbf{M} - \mathbf{M}_0$ , which depends only on  $\eta(\mathbf{r}, \tau)$  and its Hermitian conjugate.

Using the saddle-point approach, we write the effective action as  $S_{\text{eff}} = S_0 + S_F$ , where

$$S_0 = \int d\tau d\mathbf{r} \left[ \frac{|\Delta_0|^2}{g} - \frac{T}{2V} \ln \det \frac{\mathbf{M}_0}{T} + \tilde{K}_+ \delta(\mathbf{r} - \mathbf{r}') \right]$$

is the saddle-point action and

$$S_F = \int d\tau d\mathbf{r} \left[ \frac{|\eta(\mathbf{r}, \tau)|^2}{g} + \mathcal{L} - \frac{T}{2V} \ln \det (\mathbf{1} + \mathbf{M}_0^{-1} \mathbf{M}_F) \right]$$

is the fluctuation action for all orders in the fluctuation field, with  $\mathcal{L} = [\Delta_0 \eta^*(\mathbf{r}, \tau) + \Delta_0^* \eta(\mathbf{r}, \tau)]/g$ . The effects of fluctuations at both zero temperature and near the critical temperature are discussed later.

A transformation to the momentum-frequency coordinates  $(\mathbf{k}, i\omega_n)$ , where  $\omega_n = (2n + 1)\pi T$ , leads to

$$S_0 = \frac{V}{T} \frac{|\Delta_0|^2}{g} - \frac{1}{2} \sum_{j, \mathbf{k}, i\omega_n} \ln \left[ \frac{i\omega_n - E_j(\mathbf{k})}{T} \right] + \sum_{\mathbf{k}} \frac{\tilde{K}_+}{T},$$

where  $E_j(\mathbf{k})$  are the eigenvalues of the matrix

$$\mathbf{H}_0 = \begin{pmatrix} \tilde{K}_\uparrow(\mathbf{k}) & -h_\perp(\mathbf{k}) & 0 & -\Delta_0 \\ -h_\perp^*(\mathbf{k}) & \tilde{K}_\downarrow(\mathbf{k}) & \Delta_0 & 0 \\ 0 & \Delta_0^* & -\tilde{K}_\uparrow(-\mathbf{k}) & h_\perp^*(-\mathbf{k}) \\ -\Delta_0^* & 0 & h_\perp(-\mathbf{k}) & -\tilde{K}_\downarrow(-\mathbf{k}) \end{pmatrix}, \quad (6)$$

which describes the Hamiltonian of the elementary excitations in the four-dimensional vector basis  $\Psi^\dagger(\mathbf{k}) = \{\psi_\uparrow^\dagger(\mathbf{k}), \psi_\downarrow^\dagger(\mathbf{k}), \psi_\uparrow^\dagger(-\mathbf{k}), \psi_\downarrow^\dagger(-\mathbf{k})\}$  defined in momentum space. The spin-orbit field is  $\mathbf{h}_\perp(\mathbf{k}) = \mathbf{h}_R(\mathbf{k}) + \mathbf{h}_D(\mathbf{k})$ , where the first term is of the Rashba-type  $\mathbf{h}_R(\mathbf{k}) = v_R(-k_y \hat{x} + k_x \hat{y})$  and the second is of the Dresselhaus-type  $\mathbf{h}_D(\mathbf{k}) = v_D(k_y \hat{x} + k_x \hat{y})$ . We assume, without loss of generality, that  $v_R > 0$  and  $v_D > 0$ . The magnitude of the transverse field is

then  $h_\perp(\mathbf{k}) = \sqrt{(v_D - v_R)^2 k_y^2 + (v_D + v_R)^2 k_x^2}$ . In the limiting cases of Rashba-only (RO) with  $v_D = 0$  and of equal Rashba-Dresselhaus (ERD) couplings with  $v_R = v_D = v/2$ , the transverse fields are  $h_\perp(\mathbf{k}) = v_R \sqrt{k_x^2 + k_y^2}$  ( $v_R > 0$ ) and  $h_\perp(\mathbf{k}) = v|k_x|$  ( $v > 0$ ), respectively.

*Order parameter and number equations.* The saddle-point thermodynamic potential  $\Omega_0 = T S_0$  is obtained by integrating out the fermions, leading to

$$\Omega_0 = V \frac{|\Delta_0|^2}{g} - \frac{T}{2} \sum_{\mathbf{k}, j} \ln \{1 + \exp[-E_j(\mathbf{k})/T]\} + \sum_{\mathbf{k}} \tilde{K}_+,$$

with  $\tilde{K}_+ = [\tilde{K}_\uparrow(-\mathbf{k}) + \tilde{K}_\downarrow(-\mathbf{k})]/2$ . The order parameter is determined via the minimization of  $\Omega_0$  with respect to  $|\Delta_0|^2$ , leading to

$$\frac{V}{g} = -\frac{1}{2} \sum_{\mathbf{k}, j} n_F[E_j(\mathbf{k})] \frac{\partial E_j(\mathbf{k})}{\partial |\Delta_0|^2}, \quad (7)$$

where  $n_F[E_j(\mathbf{k})] = 1/\{\exp[E_j(\mathbf{k})/T] + 1\}$  is the Fermi function for energy  $E_j(\mathbf{k})$ . We replace the contact interaction  $g$  by the scattering length  $a_s$  through the relation  $1/g = -m_+/(4\pi a_s) + (1/V) \sum_{\mathbf{k}} [1/(2\epsilon_{\mathbf{k},+})]$ , where  $m_+ = 2m_\downarrow m_\uparrow / (m_\downarrow + m_\uparrow)$  is twice the reduced mass,  $\epsilon_{\mathbf{k},\alpha} = k^2/(2m_\alpha)$  are the kinetic energies, and  $\epsilon_{\mathbf{k},+} = [\epsilon_{\mathbf{k},\uparrow} + \epsilon_{\mathbf{k},\downarrow}]/2$ . The number of particles at the saddle point is obtained by  $N_\alpha = -\partial \Omega_0 / \partial \mu_\alpha$ , leading to

$$N_\alpha = \frac{1}{2} \sum_{\mathbf{k}} \left[ 1 - \sum_j n_F[E_j(\mathbf{k})] \frac{\partial E_j(\mathbf{k})}{\partial \mu_\alpha} \right]. \quad (8)$$

The self-consistent relations shown in Eqs. (7) and (8) are general for arbitrary mass and population imbalances. However, next we particularize our discussion to the case of a balanced system with equal masses.

*Balanced populations.* In the case of mass and population balanced systems, the four eigenvalues of the matrix  $\mathbf{H}_0$  are  $E_1(\mathbf{k}) = \sqrt{[\epsilon_1(\mathbf{k})]^2 + |\Delta_0|^2}$ ,  $E_2(\mathbf{k}) = \sqrt{[\epsilon_2(\mathbf{k})]^2 + |\Delta_0|^2}$ ,  $E_3(\mathbf{k}) = -E_1(\mathbf{k})$ , and  $E_4(\mathbf{k}) = -E_2(\mathbf{k})$ . Here, the auxiliary energies are  $\epsilon_1(\mathbf{k}) = \xi(\mathbf{k}) + h_\perp(\mathbf{k})$  and  $\epsilon_2(\mathbf{k}) = \xi(\mathbf{k}) - h_\perp(\mathbf{k})$ . The corresponding order parameter equation at the saddle-point level is

$$\frac{V}{g} = \frac{1}{2} \sum_{\mathbf{k}} \left[ \frac{X_1(\mathbf{k})}{2E_1(\mathbf{k})} + \frac{X_2(\mathbf{k})}{2E_2(\mathbf{k})} \right], \quad (9)$$

where  $X_m(\mathbf{k}) = \tanh[E_m(\mathbf{k})/2T]$  ( $m = 1, 2$ ). Since the mixture of equal mass fermions is balanced, the chemical potentials are the same  $\mu_\uparrow = \mu_\downarrow = \mu$ , and the associated number equation is  $N = -\partial \Omega / \partial \mu$  that reduces to

$$N = \sum_{\mathbf{k}} \left[ 1 - \frac{X_1(\mathbf{k})}{2E_1(\mathbf{k})} \epsilon_1(\mathbf{k}) - \frac{X_2(\mathbf{k})}{2E_2(\mathbf{k})} \epsilon_2(\mathbf{k}) \right]. \quad (10)$$

In Fig. 1, we show the zero-temperature behavior of  $|\Delta_0|$  and  $\mu$  as a function of  $1/(k_F a_s)$  for various values of spin-orbit coupling in the ERD and RO cases. In the ERD case, the order parameter  $|\Delta_0|$  is independent of  $v$ , and the chemical potential  $\mu(v)$  is simply  $\mu(v) = \mu(0) - mv^2/2$ , since the transverse field  $h_\perp(\mathbf{k}) = v|k_x|$  can be eliminated by momentum shifts

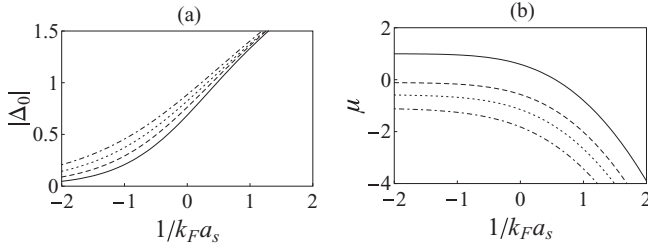


FIG. 1. Order parameter  $|\Delta_0|$  and chemical potential  $\mu$  (in units of the Fermi energy  $\epsilon_F$ ) as a function of interaction parameter  $1/(k_F a_s)$  for different spin-orbit couplings  $v_R/v_F = 0$  (solid), 0.8 (dashed), 1.0 (dotted), and 1.2 (dot-dashed) at  $T = 0$  in the RO case. For the ERD case  $|\Delta_0|$  corresponds to the solid line in (a), while  $\mu$  corresponds to the solid line in (b) shifted by  $-mv^2/2$ . Here  $v_F = k_F/m$  is the Fermi velocity and the Fermi system is balanced.

along the  $x$  direction, effectively gauging away spin-orbit effects in the *charge or momentum* sector. This symmetry also implies that the critical temperature  $T_c$  as a function of  $1/(k_F a_s)$  for finite  $v$  is the same as that for  $v = 0$ . However, in the RO case, shifts in momentum cannot gauge away the spin-orbit coupling, and  $|\Delta_0|$  increases with increasing  $v_R$ , while  $\mu$  decreases as  $v_R$  increases, exhibiting the same tendency as in the ERD case. In the BCS regime, the increase of  $|\Delta_0|$  with  $v_R$  also leads to an increase of  $T_c$  with increasing  $v_R$ .

**Momentum distribution and excitation spectrum.** The momentum distribution  $n(\mathbf{k})$  is obtained from Eq. (10) using the definition  $N = \sum_{\mathbf{k}} n(\mathbf{k})$ . At fixed momentum component  $k_z = 0$  and fixed interaction strength, the momentum distribution  $n(\mathbf{k})$  shifts continuously with increasing spin-orbit coupling in the BCS [ $1/(k_F a_s) \ll -1$ ] or unitarity regimes [ $1/(k_F a_s) \rightarrow 0$ ]. For zero spin-orbit coupling,  $n(\mathbf{k})$  is that of a superfluid degenerate Fermi system with identical single-particle bands  $\xi(\mathbf{k})$  and has a nearly flat momentum distribution until the Fermi momentum is reached. However, as the spin-orbit coupling is turned on, nonidentical single-particle bands  $\xi_{\uparrow}(\mathbf{k}) = \xi(\mathbf{k}) - h_{\perp}(\mathbf{k})$  and  $\xi_{\downarrow}(\mathbf{k}) = \xi(\mathbf{k}) + h_{\perp}(\mathbf{k})$  emerge in the helicity basis [5]  $|\mathbf{k}\uparrow\rangle, |\mathbf{k}\downarrow\rangle$  and produce a double structure with a reasonably flat momentum distribution centered around finite momenta in the  $k_x$ - $k_y$  plane. In the BEC regime [ $1/(k_F a_s) \gg 1$ ], the momentum distributions for weak and strong spin-orbit coupling broadens substantially due to the loss of degeneracy in the Fermi system when the chemical potential goes below the minima of the helicity bands and becomes large and negative. Even though there is a substantial change in the momentum distribution as a function of the spin-orbit coupling, we notice that the excitation energies  $E_1(\mathbf{k})$  and  $E_2(\mathbf{k})$  are always gapped for all values of the interaction parameter  $1/(k_F a_s)$  or the spin-orbit field  $h_{\perp}(\mathbf{k})$ , immediately suggesting that thermodynamic properties, which depend on the excitation energies, evolve smoothly from the BCS to the BEC regime in the balanced case for fixed values of spin-orbit coupling. The omnipresence of a gap in the excitation spectrum shows that the evolution from BCS to BEC superfluidity at finite spin-orbit coupling for balanced systems is a crossover. The situation is different for imbalanced systems, where gapless regions emerge in the excitation spectrum and topological phase transitions occur,

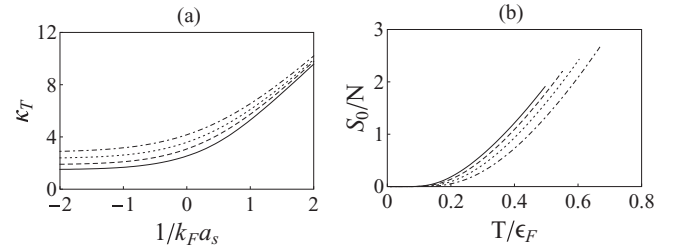


FIG. 2. (a) Compressibility  $\kappa_T$  [in units of  $1/(n\epsilon_F)$ ] as a function of interaction parameter  $1/(k_F a_s)$  at  $T = 0$  and (b) entropy per particle  $S_0/N$  as a function of temperature  $T$  (in units of  $\epsilon_F$ ) at unitarity in the RO case, for  $v_R/v_F = 0$  (solid), 0.8 (dashed), 1.0 (dotted), and 1.2 (dot-dashed). Here,  $n = N/V$  is the total particle density.

so long as the system is stable [10,11]. A thermodynamic signature of this crossover for balanced systems is seen in the isothermal compressibility discussed next.

**Isothermal compressibility.** An important thermodynamic property, which can now be measured experimentally using the fluctuation-dissipation theorem, is the isothermal compressibility

$$\kappa_T = -\frac{1}{V} \left( \frac{\partial P}{\partial V} \right)_T = \frac{V}{N^2} \left( \frac{\partial N}{\partial \mu} \right)_T. \quad (11)$$

As shown in Fig. 2(a), for the RO case, the isothermal compressibility  $\kappa_T$  at fixed interaction parameter  $1/(k_F a_s)$  increases with increasing spin-orbit coupling  $v_R$  as the Fermi system becomes less degenerate, reducing the Pauli pressure, and thus more compressible. However, in the ERD case, the isothermal compressibility for fixed interaction parameter does not change with increasing spin-orbit coupling  $v$ . In this high-symmetry situation, the momentum shift in the energy spectrum and the accompanied shift in the chemical potential do not affect the degeneracy of the Fermi system or the Pauli pressure, leading to an isothermal compressibility which is independent of the spin-orbit coupling  $v$ .

**Equation of state and entropy.** Since the thermodynamic potential  $\Omega = -PV$ , the saddle-point pressure is  $P_0(T, \mu_\alpha) = -\Omega_0/V$ , which can be shown to be always positive for arbitrary spin-orbit coupling. The general trend of the pressure for fixed interaction parameter (from the BCS to the unitarity regimes) is to decrease with increasing spin-orbit coupling for both ERD and RO cases. The situation in the BEC regime requires the inclusion of quantum fluctuations to recover the corresponding Lee-Yang corrections in the presence of spin-orbit effects. The entropy is then calculated from  $S = -(\partial \Omega / \partial T)_{V, \mu_\alpha}$ . In Fig. 2(b), we show the saddle-point entropy  $S_0$  for the RO case at unitarity. For fixed  $T$ ,  $S_0$  decreases with increasing spin-orbit coupling due to the stabilization of superfluidity by the spin-orbit field.

**Spin susceptibility tensor.** A rotation of the matrix  $\mathbf{H}_0$  into the helicity basis  $|\mathbf{k}\uparrow\rangle, |\mathbf{k}\downarrow\rangle$  introduces order parameters  $\Delta_{0,\uparrow\uparrow}(\mathbf{k})$  and  $\Delta_{0,\downarrow\downarrow}(\mathbf{k})$ , which are controlled by the spin-orbit coupling. The emergence of the triplet component affects dramatically the spin susceptibility of the system. Using standard linear response theory [12], the Pauli uniform spin

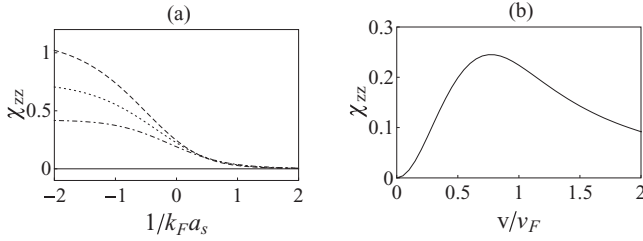


FIG. 3. (a) Pauli spin susceptibility  $\chi_{zz}$  (in units of  $\mu_B^2 n / \epsilon_F$ ) as a function of  $1/(k_F a_s)$  at  $T = 0$  in the ERD case for  $v_R/v_F = 0$  (solid), 0.8 (dashed), 1.0 (dotted), and 1.2 (dot-dashed). (b) Pauli spin susceptibility  $\chi_{zz}$  as a function of  $v/v_F$  at  $T = 0$  at unitarity in the ERD case.

susceptibility tensor [13] per unit volume is

$$\chi_{ij} = -\frac{\mu_B^2}{V} \sum_{\mathbf{k}} [a_{ij}(\mathbf{k}) - b_{ij}(\mathbf{k})], \quad (12)$$

where the spin-spin correlations in the single-particle channel are  $a_{ij}(\mathbf{k}) = \sum_{i\omega} \text{Tr}[\sigma_i \mathbf{G}(\mathbf{k}, i\omega) \sigma_j \mathbf{G}(\mathbf{k}, i\omega)]$  and in the pair (anomalous) channel are  $b_{ij}(\mathbf{k}) = \sum_{i\omega} \text{Tr}[\sigma_i \mathbf{F}(\mathbf{k}, i\omega) \sigma_j^T \mathbf{F}^\dagger(\mathbf{k}, i\omega)]$ . The matrices  $\mathbf{G}$  and  $\mathbf{F}$  are the block matrices appearing in the inverse of  $\mathbf{M}$  defined in Eq. (4),

$$\tilde{\mathbf{M}}^{-1}(\mathbf{k}, i\omega) = \begin{pmatrix} \mathbf{G} & \mathbf{F} \\ \mathbf{F}^\dagger & \mathbf{G} \end{pmatrix}.$$

In Fig. 3(a), we show plots of  $\chi_{zz}$  for the ERD case at  $T = 0$  as a function of  $1/(k_F a_s)$  for various values of spin-orbit coupling, and the behavior of  $\chi_{zz}$  for the RO case is qualitatively similar. In Fig. 3(b), we show  $\chi_{zz}$  versus  $v$  in the unitary limit  $1/(k_F a_s) = 0$ . The maximum in  $\chi_{zz}$  corresponds to the largest possible spin response. For small and large  $v$ ,  $\chi_{zz}$  is small.

In the ERD case  $\chi_{zz} = \chi_{xx} \neq \chi_{yy}$ , and in the zero-temperature limit  $\chi_{yy}(T \rightarrow 0) = 0$ , while  $\chi_{zz} = \chi_{xx}$  remains finite for nonzero spin-orbit coupling. In the RO case  $\chi_{zz} \neq \chi_{xx} = \chi_{yy}$ , and in the  $T \rightarrow 0$  limit  $\chi_{xx}(T \rightarrow 0) = \chi_{yy}(T \rightarrow 0) = \chi_{zz}(T \rightarrow 0)/2$ . Lastly, for  $h_\perp(\mathbf{k}) = 0$  (no spin-orbit coupling) the spin susceptibility tensor becomes  $\chi_{ij} = \chi \delta_{ij}$ , where the scalar  $\chi = [\mu_B^2 / (2VT)] \sum_{\mathbf{k}} \text{sech}^2[\sqrt{\xi_{\mathbf{k}}^2 + |\Delta_0|^2} / (2T)]$  is the Yoshida function, which vanishes at zero temperature, that is,  $\chi(T \rightarrow 0) = 0$ . The existence of nonzero spin response even at  $T = 0$  is a direct measure of the induced triplet component of the order parameter due to the presence of spin-orbit coupling, since a pure singlet superfluid at  $T = 0$  must have zero-spin susceptibility when all fermions are paired into a zero-spin state.

*Conclusions.* We have studied the effects of spin-orbit coupling in the evolution from BCS to BEC superfluidity at low temperatures for balanced populations. We discussed effects of spin-orbit coupling on thermodynamic properties including the order parameter, chemical potential, pressure, entropy, isothermal compressibility, and spin susceptibility tensor. Finally, we also proposed a way to detect experimentally the effects of spin-orbit coupling by measuring the isothermal compressibility and the spin susceptibility.

*Note added.* Recently, we became aware of additional papers [14–16] that discuss the effects of spin-orbit fields during the evolution from BCS to BEC superfluidity for balanced fermions. While these papers focus on the Rashba spin-orbit coupling only, we also discuss the case of equal Rashba-Dresselhaus coupling and compute the spin-orbit dependence of several thermodynamic quantities including the entropy, isothermal compressibility, and spin susceptibility tensor.

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