# Anomalous density for Bose gases at finite temperature 

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#### Abstract

We analyze the behavior of the anomalous density as function of the radial distance at different temperatures in a variational framework. We show that the temperature dependence of the anomalous density agrees with the Hartree-Fock-Bogoliubov (HFB) calculations. Comparisons between the normal and anomalous fractions at low temperature show that the latter remains higher and, consequently, the neglect of the anomalous density may destabilize the condensate. These results are compatible with those of Yukalov. Surprisingly, the study of the anomalous density in terms of the interaction parameter shows that the dip in the central density is destroyed for sufficiently weak interactions. We explain this effect.


DOI: 10.1103/PhysRevA.84.043633
PACS number(s): 67.85.Jk, 03.75.Hh, 21.60.Fw, 31.15.xt

## I. INTRODUCTION

The investigation of the Bose-Einstein condensation (BEC) phenomenon at finite temperature has attracted a great number of physicists in the recent years. The traditional theories to describe Bose-Einstein condensation at finite temperature are based on the Bogoliubov quasiparticle approach, developed originally for a spatially homogeneous Bose-condensed gas at $T \rightarrow 0$ [1] and employed by Lee and Yang [2] at finite temperatures. The generalization of the Bogoliubov method for spatially inhomogeneous systems has been described by De Gennes [3]. The main idea of Bogoliubov is to separate out the condensate part in the field operator $\Psi(\vec{r})=\Phi(\vec{r})+\bar{\Psi}(\vec{r})$. The system may be split into two subcomponents; namely, the condensate described by its density, $n_{c}(r)=|\Phi(r)|^{2}$ and the thermal cloud $\tilde{n}\left(\vec{r}, \vec{r}^{\prime}\right)=\left\langle\bar{\Psi}^{\dagger}(\vec{r}) \bar{\Psi}\left(\vec{r}^{\prime}\right)\right\rangle$. These two components have been intensively studied both theoretically and experimentally [4-34]. Similarly, this approximation motivates the definition of an additional mean-field contribution $\tilde{m}\left(\vec{r}, \vec{r}^{\prime}\right)=$ $\left\langle\bar{\Psi}(\vec{r}) \bar{\Psi}\left(\vec{r}^{\prime}\right)\right\rangle$. This is often referred to as the pair anomalous average and bears its name from the fact that there is an unequal number of creation and annihilation operators being averaged over. An analogous correlation plays a dominant role in the BCS theory of superconductivity [3], where fermionic atoms pair up to form the so-called Cooper pairs. In the case of Bose-Einstein condensation of neutral bosonic atoms, where the condensate mean field $\Phi(r)$ is the dominant parameter, the anomalous average plays a minor role for sufficiently repulsive interaction atoms in the condensate. However, this contribution becomes crucial in the presence of attractive interactions and molecular BECs [35].

Among the theoretical investigations of the anomalous density, we can cite in particular those of Hutchinson et al. [36] and Giorgini [37] based on the mean-field HFB-Bogoliubovde Gennes (BdG) approximation. To go beyond the mean field, Fedichev et al. [38] and Proukakis et al. [39,40] developed a finite-temperature perturbation theory using an HFB basis. Another kind of approach has been developed by Griffin [41] based on the Green's function method to derive an equation for the condensate and its fluctuations. Yukalov [42] adopted a

[^0]quite different approach by using the notion of representative statistical ensembles. Recently, Wright et al. [43,44] have used the classical-field trajectories of the projected Gross-Pitaevskii equation.

An interesting alternative nonperturbative and nonclassicalfield approach to the finite-temperature Bose gas is provided by the so-called time-dependent variational principle. This principle was proposed by Balian and Vénéroni (BV) a long time ago [45]. The BV variational principle has been applied to various quantum problems including heavy-ion reactions [46], quantum fields out of equilibrium [47], and attempts to go beyond the Gaussian approximation for fermion systems [48]. Therefore, the BV principle has been used to provide the best approximation to the generating functional for two- and multitime correlation functions of a set of bosonic and fermionic observables [49-52]. More recently, it was used to derive a set of equations governing the dynamics of trapped Bose gases [53,54]. The point is that this principle uses the notion of least-biased state, which is the best ansatz compatible with the constraints imposed on the system. For our purposes, we use a Gaussian density operator. The main difference between our approach and the earlier variational treatments is that, in our variational theory, we do not minimize only the expectation values of a single operator such as the free energy in the variational HF and HFB approximation or the thermodynamics potential as is done in the variational approach of Bijlsma and Stoof [24]. Conversely, our variational theory is based on the minimization of an action in addition to a Gaussian variational ansatz. The action to minimize involves two types of variational objects: one related to the observables of interest and the other that is akin to a density matrix [45,52,53]. This leads to a set of coupled time-dependent mean-field equations for the condensate, the noncondensate, and the anomalous average. We call this approach "time-dependent Hartree-Fock-Bogoliubov" (TDHFB). The TDHFB equations that we derive in this paper are quite general and fully consistent because they do not require any simplifying assumption for the thermal cloud or the anomalous density. They may provide in this sense a kind of generalization to the previously discussed approximations. Moreover, what is important in the TDHFB approach for Bose systems is that there have been no assumptions on weak interactions. Therefore, the theory is valid even for strong interactions.

In fact, in many approximations, the anomalous density is always neglected. This is based in the general claim that its contributions are negligibly small. However, Yukalov [42] pointed out that, even if it is the case, it is an essential ingredient for the consistency of the formalism. We will go further by using our TDHFB formalism to show that the contribution of the anomalous density is not as small as claimed. On the other hand, the anomalous density profiles seem to have no structure at the center of the trap for weak interactions. This is in contradiction with what was found in the literature [36,43], where the HFB-BdG approximation and the classical-field trajectories of the projected Gross-Pitaevskii equation were used, and where these densities are found to have a "dip."

The paper is organized as follows: In Sec. II, we review the main steps used to derive the TDHFB equations from the Balian-Vénéroni variational principle. In Sec. III, the TDHFB equations are applied to a trapped Bose gas to derive a coupled dynamics of the condensate, the noncondensate, and the anomalous densities. We then restrict ourselves to the local densities and discuss the properties of the underlying equations, their relevance to the finite-temperature case, as well as their relations with other known approximations such as the mean-field Bogoliubov-de Gennes equations. In Sec. IV, we study the behavior of the TDFHB equations in homogeneous Bose gases; in particular, we discuss some properties of the normal and anomalous densities as functions of the temperature. Section V is devoted to presenting the static equations and the physical boundary conditions relevant to the trapped Bose gas. We analyze the profiles of the anomalous density for different temperatures both for large and small values of the interaction parameter. We confront our results with the HFB-BdG predictions. Furthermore, we focus on the comparison between the normal and anomalous fractions at low, intermediate, and high temperatures. Our conclusions are drawn in Sec. 6.

## II. THE VARIATIONAL TDHFB EQUATIONS

The time-dependent variational principle of Balian and Vénéroni requires first the choice of a trial density operator. In our case, we will consider a Gaussian time-dependent density operator. This ansatz, which belongs to the class of the generalized coherent states, allows us to perform the calculations since there exists Wick's theorem, while retaining some fundamental aspects such as the pairing between atoms.

The Gaussian density operator $D(t)$ is completely characterized by the partition function $Z(t)=\operatorname{Tr} D(t)$, the one-boson field expectation value $\langle\Psi\rangle(\vec{r}, t) \equiv \Phi(\vec{r}, t)=$ $\operatorname{Tr} \Psi(\vec{r}) D(t) / Z(t)$ and the single-particle density matrix $\rho\left(\vec{r}, \vec{r}^{\prime}, t\right)$ defined as

$$
\rho\left(\vec{r}, \vec{r}^{\prime}, t\right)=\left(\begin{array}{cc}
\left\langle\bar{\Psi}^{\dagger}\left(\vec{r}^{\prime}\right) \bar{\Psi}(\vec{r})\right\rangle & -\left\langle\bar{\Psi}\left(\vec{r}^{\prime}\right) \bar{\Psi}(\vec{r})\right\rangle  \tag{2.1}\\
\left\langle\bar{\Psi}^{\dagger}\left(\vec{r}^{\prime}\right) \bar{\Psi}^{\dagger}(\vec{r})\right\rangle & -\left\langle\bar{\Psi}(\vec{r}) \bar{\Psi}^{\dagger}\left(\vec{r}^{\prime}\right)\right\rangle
\end{array}\right) .
$$

In the preceding definitions, $\Psi(\vec{r})$ and $\Psi^{\dagger}(\vec{r})$ are the boson destruction and creation field operators (in the Schrödinger representation), respectively, satisfying the usual canonical commutation rules

$$
\begin{equation*}
\left[\Psi(\vec{r}), \Psi^{\dagger}\left(\vec{r}^{\prime}\right)\right]=\delta\left(\vec{r}-\vec{r}^{\prime}\right), \tag{2.2}
\end{equation*}
$$

and

$$
\bar{\Psi}(\vec{r})=\Psi(\vec{r})-\langle\Psi(\vec{r})\rangle
$$

Upon introducing these variational parameters into the BV principle, one obtains dynamical equations for the expectation values of the one- and two-boson field operators:

$$
\begin{gather*}
i \frac{\partial\langle\Psi\rangle}{\partial t}(\vec{r}, t)=\frac{\partial E}{\partial\left\langle\Psi^{\dagger}\right\rangle(\vec{r}, t)}, \quad i \frac{\partial\left\langle\Psi^{\dagger}\right\rangle}{\partial t}(\vec{r}, t)=-\frac{\partial E}{\partial\langle\Psi\rangle(\vec{r}, t)} \\
i \frac{\partial \rho}{\partial t}\left(\vec{r}, \vec{r}^{\prime}, t\right)=-2\left[\rho, \frac{\partial E}{\partial \rho}\right]\left(\vec{r}, \vec{r}^{\prime}, t\right) \tag{2.3}
\end{gather*}
$$

where $E=\langle H\rangle$ is the mean-field energy. We may notice at this point that the system (2.3) is closed and does not require any further ingredients. The truncation of the full hierarchy is no longer brutally performed but rather obtained by softly restricting the full Hilbert space to the single-particle one.

One of the most noticeable properties of the TDHFB equations (2.3) is the unitary evolution of the single-particle density matrix $\rho$, which means that the eigenvalues of $\rho$ are conserved. This implies, in particular, the conservation of the von Neumann entropy $S=-\operatorname{Tr}(D) \ln D$ and the fact that an initially pure state, satisfying $\rho(\rho \ln +1)=0$, remains pure during the evolution. This property also leads to conservation of the Heisenberg parameter [52]

$$
\begin{align*}
I\left(\vec{r}, \vec{r}^{\prime}\right)= & \int d \vec{r}^{\prime \prime}\left[\left\langle\bar{\Psi}^{\dagger}(\vec{r}) \bar{\Psi}\left(\vec{r}^{\prime \prime}\right)\right\rangle\left\langle\bar{\Psi}\left(\vec{r}^{\prime \prime}\right) \bar{\Psi}^{\dagger}\left(\vec{r}^{\prime}\right)\right\rangle\right. \\
& \left.-\left\langle\bar{\Psi}^{\dagger}(\vec{r}) \bar{\Psi}^{\dagger}\left(\vec{r}^{\prime \prime}\right)\right\rangle\left\langle\bar{\Psi}\left(\vec{r}^{\prime \prime}\right) \bar{\Psi}\left(\vec{r}^{\prime}\right)\right\rangle\right] . \tag{2.4}
\end{align*}
$$

## III. APPLICATION OF THE TDHFB FORMALISM TO TRAPPED BOSE GASES

Let us apply the previous equations (2.3) to a system of trapped bosons interacting via a two-body potential. The grand canonical Hamiltonian may be written in the form

$$
\begin{align*}
H= & \int d \vec{r} \Psi^{\dagger}(\vec{r})\left[-\frac{\hbar^{2}}{2 M} \Delta+V_{\mathrm{ext}}(\vec{r})-\mu\right] \Psi(\vec{r}) \\
& +\frac{1}{2} \int d \vec{r} d \vec{r}^{\prime} \Psi^{\dagger}(\vec{r}) \Psi^{\dagger}\left(\vec{r}^{\prime}\right) V\left(\vec{r}, \vec{r}^{\prime}\right) \Psi\left(\vec{r}^{\prime}\right) \Psi(\vec{r}) \tag{3.1}
\end{align*}
$$

where $V\left(\vec{r}, \vec{r}^{\prime}\right)$ is the interaction potential, $V_{\mathrm{ext}}(\vec{r})$ is the external confining field, and $\mu$ is the chemical potential. For the sake of clarity, we will omit the time dependence whenever they are obvious. Next, we introduce the order parameter $\Phi(\vec{r})=$ $\langle\Psi(\vec{r})\rangle$ and the nonlocal densities

$$
\begin{gather*}
\tilde{n}\left(\vec{r}, \vec{r}^{\prime}\right) \equiv \tilde{n}^{*}\left(\vec{r}, \vec{r}^{\prime}\right)=\left\langle\Psi^{\dagger}(\vec{r}) \Psi\left(\vec{r}^{\prime}\right)\right\rangle-\Phi^{*}(\vec{r}) \Phi\left(\vec{r}^{\prime}\right), \\
\tilde{m}\left(\vec{r}, \vec{r}^{\prime}\right) \equiv \tilde{m}\left(\vec{r}^{\prime}, \vec{r}\right)=\left\langle\Psi(\vec{r}) \Psi\left(\vec{r}^{\prime}\right)\right\rangle-\Phi(\vec{r}) \Phi\left(\vec{r}^{\prime}\right) \tag{3.2}
\end{gather*}
$$

where we note that $\tilde{n}(\vec{r}, \vec{r}) \equiv \tilde{n}(\vec{r})$ and $\tilde{m}(\vec{r}, \vec{r}) \equiv \tilde{m}(\vec{r})$ are respectively the noncondensate and the anomalous densities.

The energy may be readily computed to yield

$$
\begin{align*}
E= & \int d \vec{r} h^{\mathrm{sp}}(\vec{r})\left[\tilde{n}(\vec{r}, \vec{r})+\Phi(\vec{r}) \Phi^{*}(\vec{r})\right]+\int d \vec{r} d \vec{r}^{\prime} V\left(\vec{r}, \vec{r}^{\prime}\right)|\Phi(\vec{r})|^{2}\left|\Phi\left(\vec{r}^{\prime}\right)\right|^{2} \\
& +\frac{1}{2} \int d \vec{r} d \vec{r}^{\prime} V\left(\vec{r}, \vec{r}^{\prime}\right)\left[\tilde{m}^{*}\left(\vec{r}, \vec{r}^{\prime}\right) \tilde{m}\left(\vec{r}, \vec{r}^{\prime}\right)+\tilde{n}\left(\vec{r}, \vec{r}^{\prime}\right) \tilde{n}\left(\vec{r}^{\prime}, \vec{r}\right)+\tilde{n}(\vec{r}, \vec{r}) \tilde{n}\left(\vec{r}^{\prime}, \vec{r}^{\prime}\right)\right] \\
& +\frac{1}{2} \int d \vec{r} d \vec{r}^{\prime} V\left(\vec{r}, \vec{r}^{\prime}\right)\left[\tilde{n}\left(\vec{r}, \vec{r}^{\prime}\right) \Phi(\vec{r}) \Phi^{*}\left(\vec{r}^{\prime}\right)+\tilde{n}\left(\vec{r}^{\prime}, \vec{r}\right) \Phi^{*}(\vec{r}) \Phi\left(\vec{r}^{\prime}\right)+\tilde{n}(\vec{r}, \vec{r}) \Phi\left(\vec{r}^{\prime}\right) \Phi\left(\vec{r}^{\prime}\right)+\tilde{n}\left(\vec{r}^{\prime}, \vec{r}^{\prime}\right) \Phi(\vec{r}) \Phi(\vec{r})\right] \\
& +\frac{1}{2} \int d \vec{r} d \vec{r}^{\prime} V\left(\vec{r}, \vec{r}^{\prime}\right)\left[\tilde{m}^{*}\left(\vec{r}, \vec{r}^{\prime}\right) \Phi(\vec{r}) \Phi\left(\vec{r}^{\prime}\right)+\tilde{m}\left(\vec{r}, \vec{r}^{\prime}\right) \Phi^{*}(\vec{r}) \Phi^{*}\left(\vec{r}^{\prime}\right)\right] . \tag{3.3}
\end{align*}
$$

In Eq. (3.3), $h^{\mathrm{sp}}=-\frac{\hbar^{2}}{2 M} \Delta+V_{\mathrm{ext}}(\vec{r})-\mu$ is the single particle Hamiltonian.

Now one inserts the expression (3.3) in the general equations of motion (2.3) to get the explicit form of the TDHFB equations for a trapped Bose gas:

$$
\begin{align*}
i \hbar \dot{\Phi}(\vec{r})= & h^{\mathrm{sp}}(\vec{r}) \Phi(\vec{r})+\int d \vec{r}^{\prime} V\left(\vec{r}, \vec{r}^{\prime}\right)\left[\left|\Phi\left(\vec{r}^{\prime}\right)\right|^{2} \Phi(\vec{r})+\Phi^{*}\left(\vec{r}^{\prime}\right) \tilde{m}\left(\vec{r}, \vec{r}^{\prime}\right)+\Phi\left(\vec{r}^{\prime}\right) \tilde{n}\left(\vec{r}, \vec{r}^{\prime}\right)+\Phi(\vec{r}) \tilde{n}\left(\vec{r}^{\prime}, \vec{r}^{\prime}\right)\right],  \tag{3.4a}\\
i \hbar \dot{\tilde{n}}\left(\vec{r}^{\prime}, \vec{r}^{\prime}\right)= & {\left[h^{\left.\operatorname{sp}(\vec{r})-h^{\operatorname{sp}}\left(\vec{r}^{\prime}\right)\right] \tilde{n}\left(\vec{r}, \vec{r}^{\prime}\right)+\int d \vec{r}^{\prime \prime} V\left(\vec{r}^{\prime}, \vec{r}^{\prime \prime}\right)\left[a\left(\vec{r}^{\prime \prime}, \vec{r}^{\prime}\right) \tilde{n}\left(\vec{r}, \vec{r}^{\prime \prime}\right)+a\left(\vec{r}^{\prime \prime}, \vec{r}^{\prime \prime}\right) \tilde{n}\left(\vec{r}, \vec{r}^{\prime}\right)+b\left(\vec{r}^{\prime}, \vec{r}^{\prime \prime}\right) \tilde{m}\left(\vec{r}^{\prime \prime}, \vec{r}\right)\right]}\right.} \\
& -\int d \vec{r}^{\prime \prime} V\left(\vec{r}^{\prime}, \vec{r}^{\prime \prime}\right)\left[a\left(\vec{r}, \vec{r}^{\prime \prime}\right) \tilde{n}\left(\vec{r}^{\prime \prime}, \vec{r}\right)+a\left(\vec{r}^{\prime \prime}, \vec{r}^{\prime \prime}\right) \tilde{n}\left(\vec{r}, \vec{r}^{\prime}\right)+b\left(\vec{r}^{\prime}, \vec{r}^{\prime}\right) \tilde{m}^{\prime}\left(\vec{r}^{\prime \prime}, \vec{r}^{\prime}\right)\right],  \tag{3.4b}\\
i \hbar \dot{\tilde{m}\left(\vec{r}, \vec{r}^{\prime}\right)=} & {\left[h^{\left.\mathrm{sp}(\vec{r})+h^{\mathrm{sp}}\left(\vec{r}^{\prime}\right)\right] \tilde{m}\left(\vec{r}, \vec{r}^{\prime}\right)+\int d \vec{r}^{\prime \prime} V\left(\vec{r}^{\prime}, \vec{r}^{\prime \prime}\right)\left\{a\left(\vec{r}^{\prime \prime}, \vec{r}^{\prime}\right) \tilde{m}\left(\vec{r}, \vec{r}^{\prime \prime}\right)+a\left(\vec{r}^{\prime \prime}, \vec{r}^{\prime \prime}\right) \tilde{m}\left(\vec{r}, \vec{r}^{\prime}\right)+b\left(\vec{r}^{\prime}, \vec{r}^{\prime}\right)\left[\tilde{n}^{*}\left(\vec{r}, \vec{r}^{\prime \prime}\right)+\delta\left(\vec{r}-\vec{r}^{\prime \prime}\right)\right]\right\}}\right.} \\
& +\int d \vec{r}^{\prime \prime} V\left(\vec{r}^{\prime}, \vec{r}^{\prime \prime}\right)\left[a\left(\vec{r}^{\prime \prime}, \vec{r}\right) \tilde{m}\left(\vec{r}^{\prime}, \vec{r}^{\prime \prime}\right)+a\left(\vec{r}^{\prime \prime}, \vec{r}^{\prime \prime}\right) \tilde{m}\left(\vec{r}, \vec{r}^{\prime}\right)+b\left(\vec{r}, \vec{r}^{\prime \prime}\right) \tilde{n}\left(\vec{r}^{\prime \prime}, \vec{r}^{\prime}\right)\right] . \tag{3.4c}
\end{align*}
$$

In the Eqs. (3.4), the dots denote time derivatives and we have introduced the quantities

$$
\begin{align*}
& a\left(\vec{r}, \vec{r}^{\prime}\right)=\tilde{n}\left(\vec{r}, \vec{r}^{\prime}\right)+\Phi^{*}(\vec{r}) \Phi\left(\vec{r}^{\prime}\right) \\
& b\left(\vec{r}, \vec{r}^{\prime}\right)=\tilde{m}\left(\vec{r}, \vec{r}^{\prime}\right)+\Phi(\vec{r}) \Phi\left(\vec{r}^{\prime}\right) \tag{3.5}
\end{align*}
$$

It is worth noticing that similar equations have been derived elsewhere using quite different approaches. For instance, Stoof [24] used a variational plus perturbative effective action, Proukakis [35,39] used truncation of the Heisenberg equations, and Chernyak et al. [55] used the generalized coherent state representation. The latter approach yields equations very close to ours, but the authors did not pursue further their analysis.

For contact potential $V\left(\vec{r}, \vec{r}^{\prime}\right)=g \delta\left(\vec{r}-\vec{r}^{\prime}\right)$, where $g$ is related to the $s$-wave scattering length $a$ by $g=4 \pi \hbar^{2} a / M$. This implies that the integrations in Eqs (3.4) are removed:

$$
\begin{equation*}
i \hbar \dot{\Phi}(r)=\left[h^{\mathrm{sp}}+g n_{c}(r)+2 g \tilde{n}(r)\right] \Phi(r)+g \tilde{m}(r) \Phi^{*}(r), \tag{3.6a}
\end{equation*}
$$

The equations of motion for $\tilde{n}\left(r, r^{\prime}\right)$ and $\tilde{m}\left(r, r^{\prime}\right)$ may be written in the compact form

$$
\begin{equation*}
i \hbar \frac{d \rho}{d t}=\mathfrak{s} \rho-\rho \mathfrak{s}^{+} \tag{3.6b}
\end{equation*}
$$

where we have defined the $2 \times 2$ matrices

$$
\begin{aligned}
& \Im\left(r, r^{\prime}\right)=\left(\begin{array}{lc}
h\left(r, r^{\prime}\right) & \Delta\left(r^{\prime}, r^{\prime}\right) \\
-\Delta^{*}(r, r) & -h^{*}\left(r, r^{\prime}\right)
\end{array}\right), \\
& \rho\left(r, r^{\prime}\right)=\left(\begin{array}{ll}
\tilde{n}\left(r, r^{\prime}\right) & \tilde{m}\left(r, r^{\prime}\right) \\
\tilde{m}^{*}\left(r, r^{\prime}\right) & \tilde{n}^{*}\left(r, r^{\prime}\right)+1
\end{array}\right)
\end{aligned}
$$

and

$$
h\left(r, r^{\prime}\right)=h^{\mathrm{sp}}(r)+2 g a\left(r^{\prime}, r^{\prime}\right), \quad \Delta(r, r)=g b(r, r)
$$

Equations (3.6) constitute the TDHFB equations for the contact interaction potential approximation in real space.

It is well known that, in the HFB theory, the issues of the ultraviolet (uv) divergence of the anomalous density arise from the zero-point occupation of quasiparticle modes [36]. In our case, the diverging term appears in Eq. (3.4c) and becomes highly nontrivial when one considers the contact potential. This is precisely the term which leads to uv divergences in the anomalous density in HFB theory [36]. Having identified the origin of this divergence, we can habitually eliminate it from the problem, by regularizing the anomalous average. This is achieved by following the method of Refs. [56,57].

Equations (3.6) remain complicated even for the contact potential where the noncondensate and anomalous densities are nonlocal functions of two spatial points $\tilde{n}\left(r, r^{\prime}\right)$ and $\tilde{m}\left(r, r^{\prime}\right)$. To proceed further and to investigate the behavior of
the various density profiles both for homogeneous and trapped gases, we consider in this article only the quantities for $r=r^{\prime}$ since these are the most physically accessible.

## IV. ANOMALOUS DENSITY FOR HOMOGENEOUS BOSE GAS

In the uniform case $\left[V_{\text {ext }}(r)=0\right]$ and for a thermal distribution at equilibrium, the relation (2.4) may rewritten as

$$
\begin{equation*}
I_{p}=\left(2 \tilde{n}_{p}+1\right)^{2}-4\left|\tilde{m}_{p}\right|^{2}=\operatorname{coth}^{2}\left[\varepsilon_{p} /(2 k T)\right], \tag{4.1}
\end{equation*}
$$

where $\varepsilon_{p}$ is the Bogoliubov energy spectrum defined by the expression

$$
\begin{equation*}
\varepsilon_{p}=\left[\left(E_{p}+2 g n-\mu\right)^{2}-g^{2}\left(n_{c}+\tilde{m}\right)^{2}\right]^{1 / 2} \tag{4.2}
\end{equation*}
$$

with $E_{p}=p^{2} /(2 m)$ being the energy of a free particle.
Note that the expression (4.2) can be derived by computing the random-phase approximation (RPA) modes from (3.6a). It is well known that, to satisfy the Goldstone or the HugenholtzPines [58] theorem, the spectrum (4.2) should be gapless in the long-wavelength limit. This is indeed satisfied provided

$$
\begin{equation*}
\mu=g(n+\tilde{n}-\tilde{m}), \tag{4.3}
\end{equation*}
$$

where $n=n_{c}+\tilde{n}$ is the total density.
Moreover, at zero temperature, the relation (4.1) becomes

$$
\begin{equation*}
\left|\tilde{m}_{p}\right|^{2}=\tilde{n}_{p}\left(\tilde{n}_{p}+1\right) . \tag{4.4}
\end{equation*}
$$

Equation (4.4) constitutes an explicit relationship between the normal and the anomalous densities at zero temperature and indicates that the anomalous density and the thermal cloud density are of the same order of magnitude at low temperatures, which leads to the fact that neglecting $\tilde{m}$ while maintaining $\tilde{n}$ is a quite hazardous approximation. Finally, as a technical remark, note that the dependence of $\tilde{n}$ and $\tilde{m}$ at zero temperature allows us to eliminate the noncondensate density from the TDHFB equations, thereby reducing the dimensionality of the problem and simplifying the numerical solution.

A straightforward calculation using Eq. (4.1) leads to a novel form of the normal and anomalous densities as a function of $\sqrt{I_{p}}$ :

$$
\begin{align*}
& \tilde{n}_{p}=\frac{1}{2}\left[\frac{E_{p}+g\left(n_{c}+\tilde{m}\right)}{\varepsilon_{p}} \sqrt{I_{p}}-1\right], \\
& \tilde{m}_{p}=-\frac{g\left(n_{c}+\tilde{m}\right)}{2 \varepsilon_{p}} \sqrt{I_{p}} \tag{4.5}
\end{align*}
$$

where $\sqrt{I_{p}}=\operatorname{coth}\left[\varepsilon_{p} /(2 k T)\right]$.
It is worth noticing that Eqs. (4.5) together with (3.6a) constitute the generalized HFB equations at finite temperature. This shows that our formalism can be reproduced easily by the full HFB equations both at finite and zero temperatures [see Eqs. (4.6)].

It is convenient now to analyze the asymptotic behavior of Eqs. (4.5) with respect to temperature.

At low temperature $\tilde{n}_{p}$ and $\tilde{m}_{p}$ behave as

$$
\begin{align*}
& \tilde{n}_{p}=\frac{1}{2}\left[\frac{E_{p}+g\left(n_{c}+\tilde{m}\right)}{\varepsilon_{p}}-1\right] \quad(T \rightarrow 0) \\
& \tilde{m}_{p}=-\frac{g\left(n_{c}+\tilde{m}\right)}{2 \varepsilon_{p}} \quad(T \rightarrow 0) \tag{4.6}
\end{align*}
$$

At high temperature we can use the asymptotic form $\sqrt{I_{p}} \cong$ $2 k T / \varepsilon_{p}$. Then, $\tilde{n}_{p}$ and $\tilde{m}_{p}$ take the form

$$
\begin{align*}
& \tilde{n}_{p}=\frac{E_{p}+g\left(n_{c}+\tilde{m}\right)}{\varepsilon_{p}}\left(\frac{k T}{\varepsilon_{p}}\right)-\frac{1}{2} \quad\left(T \rightarrow T_{c}\right) \\
& \tilde{m}_{p}=-\frac{g\left(n_{c}+\tilde{m}\right)}{\varepsilon_{p}}\left(\frac{k T}{\varepsilon_{p}}\right) \quad\left(T \rightarrow T_{c}\right) \tag{4.7}
\end{align*}
$$

To illustrate comparatively the behavior of the normal and anomalous densities as functions of temperature for a uniform Bose gas, it is useful to introduce the


FIG. 1. Absolute value of the anomalous density $\tilde{m}(\eta)$ (solid line) and the noncondensate density $\tilde{n}(\eta)$ (dashed line) vs the temperature $\tau$ for different values of $\eta$.
dimensionless variables $\eta=\varepsilon_{p} /\left[g\left(n_{c}+\tilde{m}\right)\right]$ and $\tau=$ $k T /\left[g\left(n_{c}+\tilde{m}\right)\right]$. Equations (4.5) become

$$
\begin{align*}
& \tilde{n}(\eta)=\frac{\sqrt{1+\eta^{2}}}{2 \eta} \sqrt{I(\eta)}-\frac{1}{2} \\
& \tilde{m}(\eta)=-\frac{1}{2 \eta} \sqrt{I(\eta)} \tag{4.8}
\end{align*}
$$

where $\sqrt{I(\eta)}=\operatorname{coth}[\eta /(2 \tau)]$.
In Fig. 1 we show that the absolute value of the anomalous density is always larger than the noncondensed value for small values of $\eta=0.1$ at low temperature. Moreover, the absolute value of the anomalous density is comparable with the noncondensed density for $\eta=1(|\tilde{m}| \approx \tilde{n}$ for $\tau \approx 1.2)$. However, at high temperatures, the anomalous density becomes much smaller than the noncondensed density. As can be seen from these figures, the behavior of the normal and anomalous densities found here is in good agreement with the recent theoretical results of Refs. [59-61].

Finally, we can check that omitting the anomalous density, while keeping the normal one, is mathematically inappropriate; this is clearly shown in (4.4). From these facts, it is understandable now that neglecting the anomalous density at low temperature for a homogeneous Bose gas as one habitually does in the literature is principally an unjustified approximation. Let us turn now to analyze the situation for a trapped Bose gas.

## V. ANOMALOUS DENSITY FOR TRAPPED BOSE GAS

The static TDHFB equations are obtained by setting to zero the right-hand sides of Eqs. (3.6) in their local form. For numerical purposes, it is convenient to start our treatment with the dimensionless form of the set (3.6). Let us consider a spherical trap with frequency $\omega, V_{\text {ext }}(r)=\frac{1}{2} m \omega^{2} r^{2}$ and use the harmonic oscillator length $a_{\mathrm{HO}}=\sqrt{\hbar /(m \omega)}$, as well as $a_{\mathrm{HO}}^{-3}$ and $\hbar \omega$ as units of length, density, and energy respectively. The dimensionless radial distance is $q=r / a_{\text {но }}$. The dimensionless condensed, noncondensed, and anomalous densities
are, respectively, $\hat{n}_{c}=a_{\mathrm{HO}}^{3} n_{c}, \hat{\tilde{n}}=a_{\mathrm{HO}}^{3} \tilde{n}$, and $\hat{\tilde{m}}=a_{\mathrm{HO}}^{3} \tilde{m}$. Therefore, $\hat{n}=\hat{n}_{c}+\hat{\tilde{n}}$ is the dimensionless total density. The static TDHFB equations can be solved numerically. The numerical method divides into two parts: The first part consists of finding the solutions of the static TDHFB equations that satisfies the boundary conditions summarized as follows: since the full wave function and the normal and anomalous averages must vanish as $q \rightarrow \infty$, the nonlinear term inside Eq. (3.6a) becomes negligible compared to the other terms. Therefore, the two equations of $\hat{\Phi}$ and $\hat{\tilde{m}}$ have the same approximate form $\hat{\Phi} \approx \hat{\tilde{m}} \approx q^{-3 / 2} \exp \left(-q^{2} / 4\right)$ for $q \rightarrow \infty$. The second part is used to propagate the solutions of the static TDHFB equations with an adaptable numerical method.

To illustrate our finite temperature formalism, we consider the ${ }^{87} \mathrm{Rb}$ gas with the following parameters $[30,62]: a=$ $5.82 \times 10^{-9} \mathrm{~m}, a_{\mathrm{H} 0}=7.62 \times 10^{-7} \mathrm{~m}$ and $\hbar \omega=1.32 \times 10^{-31}$ J. A convenient dimensionless parameter describing the effective strength of the interactions is $\gamma=N a / a_{\mathrm{H} 0}[20,63]$. We begin by plotting the anomalous density as functions of the radial distance for several values of temperatures for $N=20000$ atoms, i.e., $\gamma=153$.

Figure 2(a) depicts the anomalous density for varying condensation fraction and for $\gamma=153$. We notice that, by decreasing $N_{c} / N, \tilde{m}$ begins to increase in absolute value then decreases when $N_{c} / N$ approaches $50 \%$. This overall behavior has also been obtained in $[36,43]$. The anomalous density remains real and negative whatever the temperature and position. The presence of effective repulsive atomic interactions between the atoms, and thus also between the condensate and the thermal cloud, leads to the appearance of a local dip in the anomalous density at the center of the trap.

In Fig. 2(b) we show that, for a small value of the interaction parameter $(\gamma=1.53)$, the shape of the anomalous density is surprisingly modified. In particular, we observe that the dip in the neighborhood of the center of the trap disappears and the curve takes a Gaussian form. The central anomalous density is lowered for weak interactions. Such a result can be justified by the effect of interactions (i.e., for a small number of particles or a small value of $\gamma$ ). The interactions between the atoms


FIG. 2. Anomalous density profiles for various condensate fractions.


FIG. 3. Normal and anomalous fractions vs reduced temperature for $\gamma=153$.
of the condensate and the thermal cloud are lowered, which leads automatically to the correlations becoming weak which is why the anomalous density remains small and takes a Gaussian shape together with the thermal and the condensate density. We note that the effect of interactions on the anomalous density was studied theoretically earlier by many authors, but no one looked at how the shape of this quantity varies as a function of the interactions.

We now turn our attention to analyze more obviously the temperature-dependent behavior of the noncondensate and the anomalous fractions defined, respectively, as $\tilde{n} / N$ and $|\tilde{M} / N|$ where $\tilde{N}=\int d \vec{r} \tilde{n}(r)$ and $\tilde{M}=\int d \vec{r} \tilde{m}(r)$ are the integrated values of the noncondensate and the anomalous densities.

In Fig. 3, we plot the noncondensate and the absolute value of anomalous fractions as function of the reduce temperature $T / T_{c}$ for $\gamma=153$ (where we follow the method outlined in [36]). It can be seen that $\tilde{N} / N$ is increasing significantly with increasing temperature, which is why $\tilde{N}$ reaches a maximum value close to the BEC transition temperature. Furthermore, it is quite interesting to observe that the absolute value of the anomalous fraction $|\tilde{M} / N|$ becomes small as the temperature approaches zero or $T_{c}$. It reaches a maximum at intermediate temperatures ( $T \approx 0.5 T_{c}$ ).

On the other hand, a careful observation of Fig. 3 shows that $|\tilde{M} / N|$ is larger than $\tilde{N} / N$ at low temperature $\left(T \leqslant 0.5 T_{c}\right)$, which is justified mathematically by our relation (4.4). This result is in disagreement with what is given in the literature [36]. In fact, the authors of this Ref. [36] were still based on the idea that the anomalous fraction is always smaller than the normal fraction for any range of temperature, even at low temperatures, without any physical or mathematical justifications. Second, when the temperature approaches its critical value, the fraction of noncondensed particles approaches unity. That's why at $T \cong T_{c}$ the anomalous fraction simultaneously with the condensate fraction become much smaller than the normal fraction. Such a feature is well understood and explained in the literature $[36,42,43,61]$, although the absolute value of the anomalous and normal fractions are of the same order
at intermediate temperatures. It is worth noticing that this behavior persists even for $\gamma=1.53$ with a small change in the amplitudes of these fractions. Hence, we may infer from these results that the anomalous fraction plays a central role at low temperatures. It is therefore highly unlikely to neglect it for $T<T_{c}$.

It is well known that the anomalous density is proportional to the $s$-wave scattering length (i.e., the interactions between atoms). In fact, $\tilde{m}\left(\vec{r}, \vec{r}^{\prime}\right)=\left\langle\bar{\Psi}(\vec{r}) \bar{\Psi}\left(\vec{r}^{\prime}\right)\right\rangle=0$ for the ideal Bose gas. Indeed, the anomalous average appears in all calculations for a Bose system with broken gauge symmetry. Its importance will be highlighted if we compare the expressions, derived with and without taking into account the anomalous averages, with measured quantities, thus measuring the contributions from the anomalous averages. In some cases, the difference is drastic. For instance, the BEC transition is of first order when the anomalous average is not taken into account, while it is second order when we take it into account. Another example is the fact that the compressibility becomes infinite, implying that the system is unstable if the anomalous averages are absent. Moreover, the superfluid transition does not occur if the anomalous averages are omitted [61].

Since the presence of anomalous average means that particles in the system are correlated, it is important for the current experiments with ultracold atoms to find new methods of characterizing the correlated many-body states, such as systems near the Feshbach resonance, rotating condensates, atoms in optical lattices, and low-dimensional systems [64]. An interesting question to ask is whether the anomalous averages can be measured experimentally. In their recent papers, Polkovnikov et al. $[65,66]$ argued that this can be done by using interference experiments with two independent condensates for any geometry. Nevertheless this technique is not applicable for a single condensate as in our case where there are correlations between the atoms of the condensate and the thermal cloud. To this point, as far as we know, there is no experimental technique to directly measure the anomalous density itself. Experimental data about the anomalous density remain a great challenge for experimentalists.

## VI. CONCLUSIONS

By using a Gaussian density operator, we derive from the time-dependent Balian-Vénéroni variational principle a set of coupled equations of motion for a self-interacting trapped Bose gas. These equations govern in a self-consistent way the dynamics of the condensate, the thermal cloud, and the anomalous average. Our time-dependent Hartree-Fock-Bogoliubov (TDHFB) equations generalize in a natural way many of the famous approximations found in the literature such as the Bogoliubov, the Gross-Pitaevskii [67,68], the Popov [69], the Beliaev [70], the Bogoliubov-de Gennes [36], Zaremba-Nikuni-Griffin (ZNG) equations [22], and others.

The comparison between the normal and anomalous densities at different temperatures is analyzed for a homogeneous Bose gas. This analysis shows the significance of the anomalous density compared to the normal density at low temperature [42].

To better understand the advantages of our approach and owing to the importance of accounting for many-body effects, we analyzed the behavior of the anomalous density for trapped Bose gases. We solve numerically the static TDHFB equations in the local limit and for a contact potential. The outcomes of our numerical explorations are numerous. First of all, the numerical resolution of our equations is relatively easy and is not as time-consuming as the HFB-BdG calculations, especially for large atom numbers. For instance, the latter cannot be handled correctly as soon as $N \sim 10^{4}$ to $10^{5}$. By contrast, there are no such limitations in our case. We recover a well-known theoretical prediction of HFB-BdG [36] since $\tilde{m}$ increases with the temperature and then decreases as one approaches the transition. Furthermore, we found that the dip of the anomalous density is destroyed for sufficiently weak interactions. Moreover, we show that, at low temperatures, the anomalous fraction is larger than the noncondensate fraction.

The former necessarily plays a major role in the BoseEinstein condensation phenomenon. Any approach neglecting the anomalous fraction at low temperatures will inevitably lead to inconsistencies.

The dynamics of the anomalous density will be the goal of our next work with the aim of understanding how this quantity evolves in time. Furthermore, to examine more carefully our TDHFB formalism we will focus in the future on the behavior of the anomalous density in dipolar Bose-Einstein condensates at finite temperature both in the static and the dynamic cases.

## ACKNOWLEDGMENTS

We would like to thank J. Dalibard and A. Polkovnikov for many useful comments about this work. Special thanks to G. Shlyapnikov, V. I. Yukalov, J. Walraven, R. Balian, and M. Vénéroni.
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