Cold three-body collisions in hydrogen–hydrogen–alkali-metal atomic systems

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We have studied hydrogen-hydrogen-alkali three-body systems in the adiabatic hyperspherical representation. For the spin-stretched case, there exists a single *X*H molecular state when *X* is one of the bosonic alkali atoms: 7 Li, 23 Na, 39 K, 87 Rb, or 133 Cs. As a result, the only recombination process is the one that leads to formation of *XH* molecules, $H + H + X \rightarrow XH + H$, and such molecules will be stable against vibrational relaxation. We have calculated the collision rates for recombination and collision-induced dissociation as well as the elastic cross sections for H + *X*H collisions up to a temperature of 0.5 K, including the partial wave contributions from $J^{\Pi} = 0^+$ to 5⁻. We have also found that there is just one three-body bound state for such systems for $J^{\Pi} = 0^+$ and no bound states for higher angular momenta.

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I. INTRODUCTION

In the last decade, studies of three-body collisional processes have attracted tremendous attention due to their great relevance for the rapidly growing field of cold and ultracold atomic gases [\[1\]](#page-8-0). In such systems, three-body recombination and elastic and inelastic atom-molecule collisions are of particular interest.

Three-body recombination is a scattering process where three free particles collide, with two of them binding to form a molecular state, converting the binding energy into the relative kinetic energy of the atom and molecule produced. Three-body recombination, therefore, is important generically as it can shed light on binding in nuclear and chemical reactions. In ultracold atomic gas experiments, three-body recombination can lead to huge losses near a Feshbach resonance [\[2–4\]](#page-8-0) and has been studied extensively to understand the lifetime and the stability of the gas samples $[5-10]$.

Elastic atom-molecule collisions are crucial for determining the dynamics of ultracold atom-molecule mixtures at the meanfield level, and inelastic atom-molecule collisions have a big impact on the lifetime of Feshbach molecules in such systems [\[11–13\]](#page-8-0). Furthermore, in the regime of large two-body *s*-wave scattering length *a*, achieved near a Feshbach resonance, threebody collisional processes show universal scaling behavior with *a* as a result of Efimov physics $[6-8,14-16]$. These universal aspects have been observed experimentally in recent years [\[17–23\]](#page-8-0) and confirm our understanding of three-body universal properties [\[16\]](#page-8-0).

Ultracold three-body collisions, however, are only universal when the dynamics are predominantly determined by the long-range behavior of the atom-atom scattering wave function. When a system behaves universally, the complicated atom-atom interaction can thus be replaced with a much simpler model designed to reproduce the long-range wave function. It is in this context that the adiabatic hyperspherical representation has been applied to the calculation of three-body recombination. While it has proven very useful for getting

deeper insights into this process, calculating recombination for chemically important species using realistic interactions requires substantial further technical development, but does not pose fundamental difficulties.

Nevertheless, there are a few realistic systems that are sufficiently simple that recombination calculations are possible with the tools already available. For instance, recombination of helium atoms, for which there is a single 4 He₂ ro-vibrational bound state, has been studied within the adiabatic hyperspherical representation [\[24,25\]](#page-8-0). Parker *et al.* have studied the Ne + Ne + H system, calculating the $J^{\Pi} = 0⁺$ partial wave contribution to recombination and collision-induced dissociation rates [\[26\]](#page-8-0). The studies done by Suno *et al.*for three helium atoms $[24,25]$ and for He + He + alkali systems $[27]$ have included higher partial wave contributions in order to calculate the total recombination rate up to temperatures of at least 10 mK. These results are relevant to the buffer gas cooling technique used in cold and ultracold experiments [\[28,29\]](#page-8-0), where three-body recombination can lead to dramatic losses.

In this paper, we study three-body processes involving two hydrogen atoms and one alkali atom. Knowledge of these processes could benefit future ultracold atomic-gas experiments. The diatomic molecules produced by three-body recombination, for instance, are heteronuclear and thus have a permanent dipole moment. Such molecules have recently been the focus of much attention [\[30–34\]](#page-8-0). Moreover, our results for the light-alkali-atom systems at higher temperatures approach the regime relevant to the evolution of interstellar gases [\[35\]](#page-8-0).

If all atoms are spin-stretched, these systems are amenable to calculation since the $H + H$ interaction has no bound state and $X + H$ has only a single *s*-wave bound state for all the alkali species *X* considered here $[36-39]$: ⁷Li, ²³Na, ³⁹K, $87Rb$, and $133Cs$. In thermal alkali-hydrogen mixtures, the recombination rate K_3 for the process $H + H + X \rightarrow XH + H$ is related to the density of the alkali atoms by

$$
\frac{dn_X}{dt} = -K_3 n_H^2 n_X,\tag{1}
$$

where n_H and n_X are the densities of the H atoms and alkali atoms, respectively. Note that there is a 2! reduction in the rate if the hydrogen atoms are in a condensate [\[40\]](#page-8-0). The three-body calculations are thus simplified by having

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only one recombination channel, but the presence of sharp avoided crossings in the three-body potentials makes the calculations a challenge in the adiabatic representation. These sharp crossings are between different families of adiabatic potentials corresponding to $H + H$ and $X + H$.

We will use atomic units throughout unless specified otherwise. It is also convenient to convert energies to temperature units by dividing by Boltzmann's constant k_B (i.e., 3.17×10^{-6} a.u. = 1 K).

II. METHOD

After separating the center-of-mass motion, the relative motion of the three particles can be represented by the mass-scaled Jacobi vectors ρ_{12} and $\rho_{1,23}$ [\[41\]](#page-8-0):

$$
\rho_{12} = (r_1 - r_2)/d, \tag{2}
$$

$$
\rho_{12,3} = d\left(r_3 - \frac{r_1 + r_2}{2}\right),
$$
\n(3)

where r_1 , r_2 , and r_3 are the laboratory-frame position vectors of the two hydrogen atoms with mass m_H and the alkali atom with mass m_X , respectively. In the above equations, the mass scaling factor *d* is given by

$$
d^{2} = \frac{m_{X}}{\mu} \frac{2m_{H}}{2m_{H} + m_{X}}.
$$
 (4)

The three-body reduced mass is defined as follows to preserve the phase-space volume element [\[41\]](#page-8-0):

$$
\mu = \sqrt{\frac{m_H^2 m_X}{2m_H + m_X}}.\tag{5}
$$

In the adiabatic hyperspherical representation the hyperradius *R*, $R^2 = \rho_{12}^2 + \rho_{12,3}^2$, is the only coordinate with the dimensions of length and represents the overall size of the three-body system. The remaining degrees of freedom, the hyperangles, are represented collectively by Ω . We use body-frame Delves' coordinates [\[42\]](#page-8-0) such that $\Omega \equiv (\phi, \theta, \alpha, \beta, \gamma)$, with

$$
\phi = \tan^{-1}\left(\frac{\rho_{12,3}}{\rho_{12}}\right), \quad 0 \leqslant \phi \leqslant \pi/2, \tag{6}
$$

and θ the angle between the vectors ρ_{12} and $\rho_{1,23}$ such that $0 \le \theta$ $\theta \leq \pi$. The remaining hyperangles are the three Euler angles *α*, $β$, and $γ$ describing the rotation of the plane containing the three particles. As a result, the interparticle distances r_{ij} are determined in terms of the internal coordinates (R, θ, ϕ) only:

$$
r_{12} = Rd\cos\phi,\tag{7}
$$

$$
r_{23} = R \left(\frac{d^2}{4} \cos^2 \phi + \frac{1}{d^2} \sin^2 \phi + \frac{1}{2} \sin 2\phi \cos \theta \right)^{1/2}, \quad (8)
$$

$$
r_{31} = R \left(\frac{d^2}{4} \cos^2 \phi + \frac{1}{d^2} \sin^2 \phi - \frac{1}{2} \sin 2\phi \cos \theta \right)^{1/2}.
$$
 (9)

This definition of the hyperangles facilitates the symmetrization of the wave function under exchange of the two H atoms.

After rescaling the three-body wave function Ψ as $\psi =$ $R^{5/2}\Psi$, the three-body Schrödinger equation takes the form

$$
\left[-\frac{1}{2\mu} \frac{\partial^2}{\partial R^2} + \frac{\Lambda^2}{2\mu R^2} + V(R,\Omega) \right] \psi = E \psi, \qquad (10)
$$

where $V(R, \Omega)$ includes all the interactions and Λ^2 is the hyperangular momentum operator, defined by taking ρ_{12} to be the quantization axis for the body-fixed frame [\[42\]](#page-8-0) and expressed as

$$
\Lambda^2 = T_0 + T_1 + T_2 - 1/4,\tag{11}
$$

with

$$
T_0 = \frac{\partial^2}{\partial \phi^2} - \frac{4}{\sin^2 2\phi} \frac{1}{\sin \theta} \frac{\partial}{\partial \theta} \left(\sin \theta \frac{\partial}{\partial \theta} \right), \quad (12)
$$

$$
T_1 = \frac{4}{\sin^2 2\phi} \frac{1}{\sin^2 \theta} J_z^2 - \frac{1}{\cos^2 \phi} \left(2J_z^2 - J^2 \right), \quad (13)
$$

$$
T_2 = \frac{1}{\cos^2 \phi} \left(2i J_y \frac{\partial}{\partial \theta} + 2 \cot \theta J_x J_z \right). \tag{14}
$$

The components $J \equiv (J_x, J_y, J_z)$ are the total orbital angular momentum operator projected on the body-frame axes.

Since we assume the atoms to be spin-stretched (i.e., in the total spin state with the largest-magnitude spin projection), the relevant Born-Oppenheimer potential surface is the lowestquartet surface. We approximate this surface as a pairwise sum of ${}^{3}\Sigma_{u}$ two-body potentials:

$$
V(R, \Omega) = v_{HH}(r_{12}) + v_{XH}(r_{23}) + v_{XH}(r_{31}).
$$
 (15)

The two-body potentials $v_{HH}(r)$ and $v_{XH}(r)$ are shown in Fig. 1. At small distances, these potentials are determined from *ab initio* calculations [\[36,38,39\]](#page-8-0) while their long-range behavior is determined by the usual dispersion potentials [\[37,38\]](#page-8-0). Without including the fine and hyperfine interactions, these two-body potentials are expected to be quite accurate. For instance, the ab *initio* data for v_{HH} are claimed to have eight digits of accuracy $[36]$, and v_{XH} are claimed to have an absolute accuracy ranging from 10^{-6} – 10^{-7} a.u. depending on the species [\[38\]](#page-8-0). The dispersion potentials are believed to have an error of a few percent [\[37,38\]](#page-8-0).

FIG. 1. (Color online) The ³ Σ_u potentials for H + H and *X* + H. Among all these combinations, only $X + H$ systems have a single, weakly bound molecular state.

Note that, for three atoms, a nonadditive three-body term should be included in $V(R, \Omega)$. This three-body term depends on the spatial configuration of the three atoms and can be significant for certain configurations. The only fully quantum mechanical recombination calculation for a realistic system that has so far included the three-body term found its effect to be negligible [\[25\]](#page-8-0). But that work treated He, and He is not very polarizable compared to H or the other alkalis. The impact of the three-body term for the present systems is thus expected to be correspondingly larger. Unfortunately, neither the full three-body surface nor the three-body term is available for the quartet state of $H + H + X$. The importance of the three-body term for the systems we are investigating, though, can be estimated qualitatively by looking at the available three-body terms for the quartet surface of identical alkali atoms [\[43,44\]](#page-8-0), as they have similar electronic structure. The three-body term is most significant when all the atoms are close together. For Li, the three-body potential can make the minimum of the total potential about four times deeper than the pair-wise-sum potential [\[44\]](#page-8-0). For heavier alkali atoms, the three-body term can change the potential minimum by a factor of 1.2–1.5 [\[43\]](#page-8-0). When the atoms are far apart, a three-body dispersion interaction should also be included. The contribution of this interaction, however, is much smaller. For Li, the three-body dispersion interaction is a few percent of the pair-wise sum potential, and even smaller for heavier alkali atoms [\[44\]](#page-8-0). We thus expect that our results will change dramatically when the three-body interactions are included. We know, for instance, that the numerical value of the rate can change over a broad range for model problems [\[45\]](#page-8-0). However, our calculation for $H + H + X$ systems can give a sense of the order of magnitude for the three-body observables and serve as a starting point for the study of three-body interactions in such three-body systems. We further note that including a three-body term or a full three-body surface to improve the accuracy of our numerical results poses no particular problem for our approach [\[45\]](#page-8-0).

In order to solve Eq. (10) , we first expand the three-body wave function as

$$
\psi = \sum_{\nu=0}^{\infty} F_{\nu E}(R) \Phi_{\nu}(R; \Omega), \tag{16}
$$

where $\Phi_{\nu}(R; \Omega)$ are the channel functions obtained as the solutions of the adiabatic equation

$$
\left[\frac{\Lambda^2}{2\mu R^2} + V(R,\Omega)\right]\Phi_\nu = U_\nu(R)\Phi_\nu.
$$
 (17)

We solve this equation as a function of *R*, and its eigenvalues are the adiabatic potentials $U_{\nu}(R)$. Therefore, upon substitution of ψ , Eq. [\(10\)](#page-1-0) reduces to a set of coupled ordinary differential equations:

$$
\left[-\frac{1}{2\mu}\frac{d^2}{dR^2} + U_{\nu}(R)\right]F_{\nu E}(R) - \frac{1}{2\mu}\sum_{\nu'}\left[P_{\nu\nu'}(R)\frac{d}{dR} + \frac{d}{dR}P_{\nu\nu'}(R) + Q_{\nu\nu'}(R)\right]F_{\nu' E}(R) = EF_{\nu E}(R),\tag{18}
$$

with nonadiabatic couplings $P_{\nu\nu}$ and $Q_{\nu\nu}$ given by

$$
P_{\nu\nu'}(R) = \langle \langle \Phi_{\nu} | \frac{d}{dR} | \Phi_{\nu'} \rangle \rangle, \tag{19}
$$

$$
Q_{\nu\nu'}(R) = -\left\langle \left| \frac{d\Phi_{\nu}}{dR} \right| \frac{d\Phi_{\nu'}}{dR} \right\rangle \right\rangle. \tag{20}
$$

Here, the double brackets denote integration over only the hyperangular degrees of freedom.

In our calculations, the biggest computational burden comes from solving the five-dimensional adiabatic equation (17). To facilitate its solution, we separate out the external degrees of freedom (α, β, γ) and simultaneously obtain eigenstates of total orbital angular momentum and parity by further expanding the adiabatic wave functions Φ_{ν} on the basis of symmetrized Wigner *D* functions [\[42\]](#page-8-0):

$$
\Phi_{\nu}(R;\Omega) = \sum_{K=0}^{J} u_{\nu K}(R;\theta,\phi) \tilde{D}_{KM}^{J\Pi}(\alpha,\beta,\gamma), \qquad (21)
$$

where Π denotes the total parity of the system and K and M denote the projection of the total orbital angular momentum on the body-fixed and space-fixed *z* axes, respectively. Since molecules *X*H have only a single *s*-wave state, recombination only happens for the parity-favored case [i.e., when $\Pi =$ $(-1)^{J}$].

Although we are solving for the motion of the nuclei, we must require that the total wave function, including the electronic degrees of freedom, is antisymmetric under exchange of the two protons. For the electronic symmetry we are considering, exchanging protons introduces a sign change in the electronic wave function. Thus, the nuclear part of the wave function must be symmetric under proton exchange.

Since we neglect hyperfine interactions, we can couple the two protons' spin to give a total spin $I = 0, 1$ and consider their contribution to recombination independently. For $I = 1$, the spin wave function is symmetric under exchange, requiring the spatial wave function to also be symmetric. Such spatial symmetry leads to nonvanishing K_3 at ultracold temperatures [\[46\]](#page-8-0). For $I = 0$, however, the spin wave function is antisymmetric under exchange of the two protons, and the spatial wave function is therefore also antisymmetric. For this symmetry, K_3 vanishes in the zero-temperature limit $[46]$. In the present work, we calculate K_3 for $I = 1$, which is the dominant recombination process for ultracold temperatures. For more general cases where the $I = 1$ state is not preferentially prepared, our results give only a partial contribution to K_3 for temperatures beyond the ultracold regime.

We built the exchange symmetry of the two protons into the boundary conditions of the body-frame components $u_{\nu K}$. Permuting the two protons only affects the hyperangles Ω :

$$
P_{12}\tilde{D}_{KM}^{J\Pi} = \Pi(-1)^K \tilde{D}_{KM}^{J\Pi},\tag{22}
$$

$$
P_{12}\theta = \pi - \theta. \tag{23}
$$

For even parity, the permutation requirements can be equivalently expressed as u_{vK} being symmetric about $\theta = \pi/2$ for even *K* and antisymmetric for odd *K*. For odd parity, u_{vK} should be antisymmetric for even *K* and symmetric for odd *K*. Imposing these boundary conditions, we need only solve Eq. (17) in the range $0 \le \theta \le \pi/2$.

Asymptotically (i.e., as $R \to \infty$), the adiabatic potentials with the diagonal couplings included are determined by the energies of the break-up components. For the atom-molecule channel, the potentials behave like

$$
W_0(R) = U_0(R) - \frac{1}{2\mu} Q_{0,0} \to E_{XH} + \frac{l(l+1)}{2\mu R^2}, \quad (24)
$$

where the partial angular momentum *l* is the relative orbital angular momentum between the atom and the molecule. Since all of the *X*H systems have only an *s*-wave bound state, $l = J$. For the three-body break-up channels, the potentials behave like

$$
W_{\nu}(R) \to \frac{\lambda(\lambda+4) + 15/4}{2\mu R^2}.
$$
 (25)

The values of λ are nonnegative integers determined by J^{Π} and the identical particle symmetry [\[46\]](#page-8-0).

Accurate numerical calculations of the three-body observables depend largely on the accuracy of the adiabatic potentials and channel functions [Eqs. [\(17\)](#page-2-0) and [\(21\)](#page-2-0)] and, ultimately, on the nonadiabatic couplings [Eqs. [\(19\)](#page-2-0) and [\(20\)](#page-2-0)]. By expanding the body-frame components $u_{\nu K}(R; \theta, \phi)$ on a two-dimensional, direct-product B-spline basis [\[47\]](#page-8-0), we obtain accurate potentials and couplings up to $R \approx 2000$ a.u. Beyond this distance, we extrapolate the potentials using the known asymptotic expansions [\[48\]](#page-8-0). Typically, a (θ, ϕ) mesh of 60×250 gives eigenvalues converged to at least eight digits. We have found that, due to the sharp avoided crossings occurring at small *R*, a hyperradial grid of about 3000 points is necessary to accurately resolve most of the abrupt changes in the nonadiabatic couplings. Many sharper crossings remain, though, that must be traced individually.

III. THREE-BODY SCATTERING OBSERVABLES

It is well known $[16]$ that when the scattering length *a* greatly exceeds the characteristic range of the two-body interaction, three-body scattering observables are dramatically affected. For the systems we consider here, the long-range part of the two-body interaction is the van der Waals potential $-C_6/r_{ij}^6$. Therefore, they are characterized by the van der Waals length $l_{\text{vdW}} = (2\mu_{ij} C_6)^{1/4}$ [\[16\]](#page-8-0), where μ_{ij} is the twobody reduced mass. In Table I, we list the bound-state energies, the scattering lengths, and the van der Waals lengths for all of the two-body potentials we used. Notice that none of the scattering lengths are substantially larger than the van der Waals lengths and thus the condition for universal behavior $(|a| \gg l_{\text{vdW}})$ is not fulfilled. As a result, we do not expect to observe universal physics for these systems.

After obtaining the potentials and couplings, we solve the hyperradial equation [\(18\)](#page-2-0) using finite elements as described in Ref. [\[49\]](#page-8-0). For recombination processes, the total recombination rate K_3 is the sum over all the partial wave contributions $K_3^{J\Pi}$ [\[24,25\]](#page-8-0):

$$
K_3 = \sum_{J,\Pi} K_3^{J\Pi} = 2! \sum_{J,\Pi} \sum_i \frac{32(2J+1)\pi^2}{\mu k^4} |S_{f \leftarrow i}^{J\Pi}|^2, \quad (26)
$$

where $k = \sqrt{2\mu E}$ and $S_{f \leftarrow i}^{\text{J}\Pi}$ is the scattering matrix element from the initial three-body continuum channel to the final

TABLE I. The two-body bound state energy E_{XH} , scattering length *a*, and van der Waals length l_{vdW} for the H + H and X + H interactions.

	E_{XH} (a.u.)	a (a.u.)	l_{vdW} (a.u.)
$H + H$		1.557	10.45
$Li + H$	-1.268×10^{-7}	63.71	21.50
$Na + H$	-3.376×10^{-7}	43.26	22.58
$K + H$	-7.360×10^{-7}	34.72	25.12
$Rb + H$	-2.446×10^{-7}	50.24	25.92
$Cs + H$	-1.784×10^{-7}	56.85	27.18

atom-molecule channel. From the asymptotic form of the three-body entrance channel, the threshold behavior of K_3^{J} is determined by the smallest λ for that symmetry [\[46\]](#page-8-0) such that:

$$
K_3^{J\Pi} \propto E^{\lambda_{\min}^{J\Pi}}.\tag{27}
$$

In our calculations, we have included the lowest six partial-waves J^{Π} , implying that $\lambda_{\min}^{J\Pi} = 0, 1, 2, 3, 4$, and 5, respectively.

Collision-induced dissociation $H + XH \rightarrow X + H + H$ is the time-reversed process of three-body recombination. The dissociation rate D_3 is defined as $[24]$

$$
D_3 = \sum_{J,\Pi} D_3^{J\Pi} = \sum_{J,\Pi} \sum_f \frac{(2J+1)\pi^2}{\mu_{12,3} k_{12,3}} \left| S_{f\leftarrow i}^{J\Pi} \right|^2, \qquad (28)
$$

where $k_{12,3} = \sqrt{2\mu_{12,3}(E - E_{XH})}$ and $\mu_{12,3} = m_H(m_H + E_{HH})$ m_X)/(2 $m_H + m_X$) is the reduced mass between the H atom and the *X*H molecule.

Note that the channels that the indices i and f refer to are reversed from those in K_3 . Since the *S*-matrix is unitary, *D*³ can be readily calculated once the *S*-matrix elements for K_3 are known. Near the three-body breakup threshold where collision-induced dissociation becomes energetically possible, $D_3^{\text{J}\,\Pi}$ behaves like [\[46\]](#page-8-0)

$$
D_3^{J\Pi} \propto E^{\lambda_{\min}^{J\Pi}+2}.\tag{29}
$$

For atom-molecule collisions, the elastic cross section is [\[25\]](#page-8-0)

$$
\sigma_2 = \sum_{J,\Pi} \sigma_2^{J\Pi} = \sum_{J,\Pi} \frac{(2J+1)\pi}{k_{12,3}^2} \left| S_{0\leftarrow 0}^{J\Pi} - 1 \right|^2. \tag{30}
$$

The threshold behavior of $\sigma_2^{\text{J}\text{T}}$, in contrast to recombination, is determined solely by *J* and follows the standard Wigner threshold law,

$$
\sigma_2^{J\Pi} \sim (E - E_{X\text{H}})^{2J} \,. \tag{31}
$$

The calculation of scattering solutions to Eq. [\(18\)](#page-2-0) are complicated by the sharp avoided crossings in the adiabatic potentials: we typically use 5×10^4 hyperradial elements distributed as $R_i \propto i^3$ from $R = 10$ a.u. to $R \approx 2000$ a.u. In the asymptotic region ($R > 2000$ a.u.), the density of elements is fixed to eight elements per shortest de Broglie wavelength. To calculate the scattering observables, we match the numerical solutions to the asymptotic analytical solutions at $R = 10^5$ a.u.

for recombination, and at $R = 5 \times 10^3$ a.u. for atom-molecule collisions. The convergence of the scattering observables with respect to the number of adiabatic channels is also dramatically affected by the sharp avoided crossings. Even the threshold behavior for $K_3^{J\Pi}$ and $\sigma_2^{J\Pi}$ requires a fairly large number of adiabatic channels for convergence, which we take to be from 12 to 25 for all the calculations. The resulting K_3 and σ_2 are converged to at least two digits for all partial waves, and the three-body bound-state energies are converged to three digits.

In our calculations, we have included $J^{\Pi} = 0^{+}$, 1⁻, 2⁺, 3−, 4+, and 5[−] for the convergence of the total rates and cross sections at high energies. The overall convergence of the total total rates and cross sections are converged to two digits for $E < 200$ mK and one digit for 200 mK $< E < 500$ mK.

IV. RESULTS

A. Three-body recombination rates

Since the adiabatic hyperspherical potentials $U_{\nu}(R)$ are important in understanding the underlying three-body physics involved in the scattering processes, we first discuss their behavior. We see that the avoided crossings become sharper as we go to heavier alkali atoms. As an illustration, in Fig. 2 we show the lowest six adiabatic potentials $U_{\nu}(R)$ with $J^{\overline{\Pi}} = 0^{+}$

FIG. 2. (Color online) The adiabatic three-body potentials for (a) $H + H + Li$, and (b) $H + H + Cs$. The insets show the avoided crossings within the energy range we have considered for the rates and cross sections.

FIG. 3. (Color online) The nonadiabatic couplings for $H + H +$ Cs with $J^{\Pi} = 0^+$. Note that $Q_{1,2}$ has been multiplied by 5000 to have a magnitude similar to $P_{1,2}$. The index ν for the adiabatic potentials starts from 0.

for $H + H + Li$ and $H + H + Cs$. The potentials for higher partial waves behave similarly but become more repulsive as *J* increases.

To demonstrate the effects of sharp avoided crossings on the adiabatic potentials on the nonadiabatic couplings, we show a key crossing in the insets of Fig. 2. Figure 3 shows the corresponding couplings $P_{\nu\nu'}$ and $Q_{\nu\nu'}$ for $H + H + Cs$ with $J^{\Pi} = 0^+$. It can be seen that, when a sharp avoided crossing occurs between two potential curves, the *Pνν* and $Q_{\nu\nu}$ coupling those curves show sharp spikes at that *R*. These couplings must be carefully traced out with a dense hyperradial grid in order to obtain an accurate solution of Eq. [\(18\)](#page-2-0).

In our calculations, the convergence of K_3 depends critically on the behavior of the adiabatic potentials. We have found that, except for $H + H + Li$, all the systems have sharp avoided crossings below 1 Kelvin. To calculate K_3 for energies above the crossings like the one shown in the inset of Fig. $2(b)$, both of the potentials involved in the crossing need to be included to avoid spurious resonances.

The three-body recombination rates for different alkali species *X* are shown in Figs. $4(a) - 4(f)$. The corresponding data are available in electronic form [\[50\]](#page-8-0). Generally, the three-body recombination rates are dominated by the $J^{\Pi} = 0⁺$ contribution at ultracold energies, and by the $J^{\overline{\Pi}} = 1^-$ and 2^+ contributions near the highest energies we have calculated. It can be seen that the total three-body recombination rates for $H + H + Na$, $H + H + Rb$, and $H + H + Cs$ behave similarly and that the rates for $H + H + Li$ and $H + H + K$ recombination behave differently. In particular, the $J^{\Pi} = 0^{+}$ partial rates for $H + H + K$ recombination are much smaller than the rates for other systems near the zero-energy threshold. The total rates for $H + H + K$ recombination are then dominated by the $J^{\Pi} = 1⁻$ partial wave contribution for a large energy range from about 0.5 mK to 50 mK. Interestingly, we have observed that the threshold values of the recombination rates for different *X* are ordered by the magnitude of their nonadiabatic couplings P_{01} and Q_{01} at large hyperradii $R > 200$ a.u. for 0^+ . This suggests that, for the present cases, 0^+ recombination is dominated by inelastic transitions from the lowest continuum channel to the atom-molecule channel at large distances.

FIG. 4. (Color online) The total three-body recombination rate K_3 and the partial rates $K_3^{J\Pi}$ for $H + H + X \rightarrow H + XH$, where *X* is (a) Li, (b) Na, (c) K, (d) Rb, and (e) Cs. The thermally averaged total recombination rates are shown in (f).

In thermal gases, it is crucial to consider the thermal distribution of the collisional energies when calculating a collision rate. Assuming a Boltzmann distribution, we have performed a thermal average of the energy-dependent rates. The thermally averaged recombination rate $\langle K_3 \rangle$ are given by [\[51\]](#page-8-0)

$$
\langle K_3 \rangle = \frac{1}{2(k_{\rm B}T)^3} \int_0^\infty K_3 E^2 e^{-E/(k_{\rm B}T)} dE. \tag{32}
$$

The results are shown in Fig. $4(f)$. To perform the thermal average, we extrapolate K_3 from the lowest energy we have calculated to zero energy using the known threshold behavior. Since we could not similarly extrapolate to infinite energy for the integral in Eq. (32), the integral was limited to the energies we could calculate. Consequently, the thermally averaged rates are converged to more than one digit only for temperatures below 100 mK. It can be seen that the energy dependence of the rates is largely preserved. Furthermore, the thermally averaged rates for these systems lie close to each other when the temperature is beyond 10 mK. In Table II , we list the values of K_3 for the processes $H + H + X \rightarrow XH + H$ in the zero-energy limit for reference.

B. Collision-induced dissociation rates

Using the simple relation between K_3 and D_3 , we have also calculated D_3 for the same range of E . In Fig. [5,](#page-6-0) we show the thermally averaged collision-induced dissociation rate $\langle D_3 \rangle$ as

TABLE II. The zero-energy limit of the three-body recombination rates K_3 .

	K_3 (cm ⁶ /s)
$H + H + Li$	1.3×10^{-29}
$H + H + Na$	6.1×10^{-31}
$H + H + K$	2.3×10^{-32}
$H + H + Rb$	9.7×10^{-31}
$H + H + Cs$	2.5×10^{-30}

a function of the temperature of the $XH + H$ mixture, where $\langle D_3 \rangle$ is given by [\[49\]](#page-8-0)

$$
\langle D_3 \rangle = \frac{2}{\sqrt{\pi}} \frac{1}{(k_{\rm B}T)^{3/2}} \int_0^\infty D_3 E^{1/2} e^{-E/(k_{\rm B}T)} dE. \tag{33}
$$

The energy E in the integrand is relative to the $XH + H$ threshold. That is, it is the $XH + H$ scattering energy, and this must be taken into account when evaluating $D_3(E)$. It is interesting to note that, although dissociation is allowed only when the collision energy exceeds the molecular binding energy, Fig. 5 shows that in a thermal gas dissociation can occur for temperatures well below the dissociation threshold. In fact, because we know the threshold behavior from Eq. [\(29\)](#page-3-0) to be $D_3 \propto (E - E_{XH})^2$ for $E \ge E_{XH}$ (and zero below E_{XH}), we can explicitly calculate $\langle D_3 \rangle$ below threshold:

$$
\langle D_3 \rangle \propto 2\sqrt{x}(15x - 2) e^{-1/x}
$$

 $+ \sqrt{\pi}[4 + 3x(5x - 4) \text{erfc}(1/\sqrt{x})],$ (34)

with $x = k_B T/E_{XH}$. By contrast, $\langle K_3 \rangle$ and $\langle \sigma_2 \rangle$ have the same threshold behavior as the energy-dependent quantities (with *E* replaced by $k_B T$). This formula for $\langle D_3 \rangle$ is likely valid only for temperatures below E_{XH} since the tail of the thermal distribution starts sampling energies outside the threshold regime for higher temperatures, making our assumption for the behavior of D_3 invalid.

FIG. 5. (Color online) The thermally averaged total collisioninduced dissociation rate $\langle D_3 \rangle$ for $H + XH \rightarrow X + H + H$. For all species, the rates are shown up to 100 mK beyond the three-body break-up threshold, which is indicated by the vertical dashed line.

FIG. 6. (Color online) The total atom-molecule elastic cross section σ_2 and the partial cross sections $\sigma_2^{\prime\,\Pi}$ for H + KH \rightarrow H + KH.

C. Atom-molecule elastic cross sections

As a representative example, we plot the total and partial cross sections for elastic collisions between H and KH in Fig. 6. The $J^{\Pi} = 0^{+}$ partial wave contribution dominates for collisional energies below 100 mK, beyond which the J^{Π} = 1[−] contribution becomes dominant. The $J^Π = 1[−]$ contribution has a pronounced minimum near 20 mK, but this feature has only a negligible effect on the total cross section. The partial atom-molecule elastic cross sections for the other alkali species are not shown, as their energy dependence is qualitatively the same as shown for $H + KH$. Instead, we show in Fig. 7 the thermally averaged cross sections for all the alkali species. The thermally averaged cross section can be derived from the thermally averaged elastic scattering rate and is given by [\[49\]](#page-8-0)

$$
\langle \sigma_2 \rangle = \frac{1}{(k_{\rm B}T)^2} \int_0^\infty \sigma_2 E e^{-E/(k_{\rm B}T)} dE. \tag{35}
$$

The total elastic cross sections for all alkali species are converged to two digits for all energies. All these data are available in electronic form [\[50\]](#page-8-0).

In Table [III,](#page-7-0) we list the values of the elastic cross section extrapolated to zero temperature as well as the corresponding values for the atom-molecule scattering length a_{H+XH} for all the alkali species. Both quantities increase with the respective values of the two-body scattering lengths (see Table [I\)](#page-3-0) or,

FIG. 7. (Color online) The thermally averaged total atommolecule elastic cross section $\langle \sigma_2 \rangle$ for H + *X*H \rightarrow H + *X*H.

TABLE III. The atom-molecule zero temperature elastic cross section and scattering length between H and *X*H.

	σ_2 (cm ²)	$4\pi a^2$ (cm ²)	a_{H+XH} (a.u.)
$H + LiH$	2.2×10^{-12}	1.4×10^{-12}	80
$H + NaH$	8.2×10^{-13}	6.6×10^{-13}	48
$H + KH$	4.6×10^{-13}	4.2×10^{-13}	36
$H + RbH$	1.3×10^{-12}	8.9×10^{-13}	60
$H + CsH$	1.6×10^{-12}	1.1×10^{-12}	68

equivalently, the size of the molecular state. In fact, we can get an order-of-magnitude estimate for the elastic cross section by simply using $\sigma_2 \approx 4\pi a^2$. These estimated values are also shown in Table III. The rough agreement implies that, for our simple pair-wise-sum potential, the zero-energy elastic cross section is mainly determined by the size of the *X*H molecule. By extension, the atom-molecule scattering length can be approximated by the $X + H$ scattering length at the same level of approximation.

D. Three-body bound-state energies

To complete our study of these systems, we calculate the three-body bound states. The three-body energy spectra for the $H + H +$ alkali systems are very simple due to their weakly-interacting nature. In our calculations, we have found only one triatomic vibrational bound state for $J^{\Pi} = 0⁺$ for all the systems. No bound levels are found for higher angular momenta. The three-body binding energies, relative to the atom-molecule break-up threshold, are listed in Table IV.

To get a sense of the sizes of the triatomic molecules, we have also calculated the expectation values of the interatomic distances $\langle r_{XH} \rangle$ and $\langle r_{HH} \rangle$, given by

$$
\langle r_{\rm AH} \rangle = \sum_{\nu,\nu'} \int_0^\infty F_\nu(R) F_{\nu'}(R) \langle \langle \Phi_\nu | r_{\rm AH} | \Phi_{\nu'} \rangle \rangle \, dR,\qquad(36)
$$

where *A* represents *X* or H. From these bond lengths, we can also calculate the bond angle at the *X* atom and find them to be consistently around $100°$ for all species. All of this geometrical information is included in Table IV. From the small binding energies and large bond lengths, it is clear that these are very floppy states, as is expected for van der Waal's molecules. We expect, though, that the inclusion of three-body terms in the interaction potential will tend to bind these states more strongly, reducing the bond lengths correspondingly. The three-body term may further tend to increase the bond angle toward a linear configuration. The three-body term might even be sufficient to bind additional states, at least for some species.

TABLE IV. The 0^+ triatomic bound-state energies, expectation values of interatomic distances, and bond angles.

	E_{XH_2} (a.u.)	$\langle r_{XH} \rangle$ (a.u.)	$\langle r_{\rm HH} \rangle$ (a.u.)	Bond Angle
LiH ₂	9.02×10^{-8}	43	65	98°
NaH ₂	2.58×10^{-7}	30	46	100°
KH_{2}	6.24×10^{-7}	25	37	95°
RbH ₂	1.95×10^{-7}	34	52	100°
CsH ₂	1.43×10^{-7}	38	59	102°

FIG. 8. (Color online) The $J^{\Pi} = 0^{+}$ hyperradial channel function $F_0(R)$ for the triatomic bound states.

Finally, we have verified numerically that the triatomic binding energies are indeed predominantly determined by the lowest adiabatic channel $\nu = 0$. Specifically, the channel probability $\int |F_v(R)|^2 dR$ for $v = 0$ is beyond 99% for all the systems. The channel functions $F_0(R)$ are shown in Fig. 8. It is interesting to note that, except for $KH₂$, all the triatomic states have a large hyperradial extent, reaching values up to a few hundred atomic units, consistent with the bond lengths listed in Table IV.

V. SUMMARY

In this paper, we have studied three-body scattering and the bound state spectra for two hydrogen atoms and one alkali atom using a fully quantum mechanical approach. Solving the threebody Schrödinger equation in the adiabatic hyperspherical representation, we have calculated the three-body recombination rates and atom-molecule elastic cross sections for all the alkali species for temperatures up to 0.5 Kelvin. The biggest uncertainty in our calculations by far is the interaction potential. Nevertheless, we expect that our results give a correct order-of-magnitude estimate of the three-body scattering observables. For three-body recombination, the lowest three partial waves dominate the total recombination rates in the energy range we have calculated. For the elastic atom-molecule collisions, the cross sections are dominated by a single-partial wave contribution for the energy range in our calculations, which is $J^{\Pi} = 0^+$ at lower energies and $J^{\Pi} = 1^-$ at higher energies. The bound-state spectra are very simple, with only one ro-vibrational three-body state for each of the alkali species.

Finally, the difficulty of sharply avoided crossings we met at small hyperradius raises an alert for doing adiabatic calculations for realistic systems. The complicated short-range three-body dynamics can give rise to rapidly varying behavior in the adiabatic potentials and the nonadiabatic couplings, which makes the adiabatic calculations much harder and less reliable. For such cases, a diabatic representation of some sort will become necessary, especially for small distances [\[52,53\]](#page-8-0). Besides this technical improvement, the calculations can be made more realistic through the inclusion of full three-body potential surfaces should they become available. Explicitly

accounting for fine and hyperfine interactions would further improve the results.

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- [1] C. Chin, R. Grimm, P. Julienne, and E. Tiesinga, [Rev. Mod.](http://dx.doi.org/10.1103/RevModPhys.82.1225) Phys. **82**[, 1225 \(2010\).](http://dx.doi.org/10.1103/RevModPhys.82.1225)
- [2] C. A. Stan, M. W. Zwierlein, C. H. Schunck, S. M. F. Raupach, and W. Ketterle, Phys. Rev. Lett. **93**[, 143001 \(2004\).](http://dx.doi.org/10.1103/PhysRevLett.93.143001)
- [3] S. Inouye, J. Goldwin, M. L. Olsen, C. Ticknor, J. L. Bohn, and D. S. Jin, Phys. Rev. Lett. **93**[, 183201 \(2004\).](http://dx.doi.org/10.1103/PhysRevLett.93.183201)
- [4] T. Weber, J. Herbig, M. Mark, H.-C. Nägerl, and R. Grimm, Phys. Rev. Lett. **91**[, 123201 \(2003\).](http://dx.doi.org/10.1103/PhysRevLett.91.123201)
- [5] P. O. Fedichev, M. W. Reynolds, and G. V. Shlyapnikov, [Phys.](http://dx.doi.org/10.1103/PhysRevLett.77.2921) Rev. Lett. **77**[, 2921 \(1996\).](http://dx.doi.org/10.1103/PhysRevLett.77.2921)
- [6] E. Nielsen and J. H. Macek, [Phys. Rev. Lett.](http://dx.doi.org/10.1103/PhysRevLett.83.1566) **83**, 1566 (1999).
- [7] B. D. Esry, C. H. Greene, and J. P. Burke, [Phys. Rev. Lett.](http://dx.doi.org/10.1103/PhysRevLett.83.1751) **83**, [1751 \(1999\).](http://dx.doi.org/10.1103/PhysRevLett.83.1751)
- [8] P. F. Bedaque, E. Braaten, and H.-W. Hammer, [Phys. Rev. Lett.](http://dx.doi.org/10.1103/PhysRevLett.85.908) **85**[, 908 \(2000\).](http://dx.doi.org/10.1103/PhysRevLett.85.908)
- [9] T. Weber, Jens Herbig, M. Mark, H.-C. Nägerl, and R. Grimm, Phys. Rev. Lett. **91**[, 123201 \(2003\).](http://dx.doi.org/10.1103/PhysRevLett.91.123201)
- [10] G. Roati, M. Zaccanti, C. D'Errico, J. Catani, M. Modugno, A. Simoni, M. Inguscio, and G. Modugno, [Phys. Rev. Lett.](http://dx.doi.org/10.1103/PhysRevLett.99.010403) **99**, [010403 \(2007\).](http://dx.doi.org/10.1103/PhysRevLett.99.010403)
- [11] S. Jochim, M. Bartenstein, A. Altmeyer, G. Hendl, S. Riedl, C. Chin, J. Hecker Denschlag, and R. Grimm, [Science](http://dx.doi.org/10.1126/science.1093280) **302**, [2101 \(2003\).](http://dx.doi.org/10.1126/science.1093280)
- [12] C. Chin, T. Kraemer, M. Mark, J. Herbig, P. Waldburger, H.-C. Nägerl, and R. Grimm, *Phys. Rev. Lett.* **94**[, 123201 \(2005\).](http://dx.doi.org/10.1103/PhysRevLett.94.123201)
- [13] C. A. Regal, M. Greiner, and D. S. Jin, [Phys. Rev. Lett.](http://dx.doi.org/10.1103/PhysRevLett.92.040403) **92**, [040403 \(2004\).](http://dx.doi.org/10.1103/PhysRevLett.92.040403)
- [14] V. Efimov, [Phys. Lett. B](http://dx.doi.org/10.1016/0370-2693(70)90349-7) **33**, 563 (1970).
- [15] E. Braaten and H.-W. Hammer, [Phys. Rev. Lett.](http://dx.doi.org/10.1103/PhysRevLett.87.160407) **87**, 160407 [\(2001\).](http://dx.doi.org/10.1103/PhysRevLett.87.160407)
- [16] E. Braaten and H.-W. Hammer, Phys. Rep. **428**[, 259 \(2006\).](http://dx.doi.org/10.1016/j.physrep.2006.03.001)
- [17] T. Kraemer, M. Mark, P. Waldburger, J. G. Danzl, C. Chin, B. Engeser, A. D. Lange, K. Pilch, A. Jaakkola, H.-C. Nägerl, and R. Grimm, [Nature \(London\)](http://dx.doi.org/10.1038/nature04626) **440**, 315 (2006).
- [18] S. Knoop, F. Ferlaino, M. Mark, M. Berninger, H. Sch, H.-C. Näger, and R. Grimm, Nat. Phys. 5[, 227 \(2009\).](http://dx.doi.org/10.1038/nphys1203)
- [19] T. B. Ottenstein, T. Lompe, M. Kohnen, A. N. Wenz, and S. Jochim, Phys. Rev. Lett. **101**[, 203202 \(2008\).](http://dx.doi.org/10.1103/PhysRevLett.101.203202)
- [20] J. H. Huckans, J. R. Williams, E. L. Hazlett, R. W. Stites, and K. M. O'Hara, Phys. Rev. Lett. **102**[, 165302 \(2009\).](http://dx.doi.org/10.1103/PhysRevLett.102.165302)
- [21] G. Barontini, C. Weber, F. Rabatti, J. Catani, G. Thalhammer, M. Inguscio, and F. Minardi, [Phys. Rev. Lett.](http://dx.doi.org/10.1103/PhysRevLett.103.043201) **103**, 043201 [\(2009\).](http://dx.doi.org/10.1103/PhysRevLett.103.043201)
- [22] M. Zaccanti, B. Deissler, M. Fattori, M. Jona-Lasinio, G. Roati, M. Inguscio, and G. Modugno, Nat. Phys. **5**[, 586 \(2009\).](http://dx.doi.org/10.1038/nphys1334)
- [23] N. Gross, Z. Shotan, S. Kokkelmans, and L. Khaykovich, *[Phys.](http://dx.doi.org/10.1103/PhysRevLett.103.163202)* Rev. Lett. **103**[, 163202 \(2009\).](http://dx.doi.org/10.1103/PhysRevLett.103.163202)
- [24] H. Suno, B. D. Esry, C. H. Greene, and J. P. Burke Jr., *[Phys.](http://dx.doi.org/10.1103/PhysRevA.65.042725)* Rev. A **65**[, 042725 \(2002\).](http://dx.doi.org/10.1103/PhysRevA.65.042725)
- [25] H. Suno and B. D. Esry, Phys. Rev. A **78**[, 062701 \(2008\).](http://dx.doi.org/10.1103/PhysRevA.78.062701)
- [26] G. A. Parker, R. B. Walker, B. K. Kendrick, and R. T. Pack, [J. Chem. Phys.](http://dx.doi.org/10.1063/1.1503313) **117**, 6083 (2002).

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- [27] H. Suno and B. D. Esry, Phys. Rev. A **80**[, 062702 \(2009\).](http://dx.doi.org/10.1103/PhysRevA.80.062702)
- [28] S. C. Doret, C. B. Connolly, W. Ketterle, and J. M. Doyle, *[Phys.](http://dx.doi.org/10.1103/PhysRevLett.103.103005)* Rev. Lett. **103**[, 103005 \(2009\).](http://dx.doi.org/10.1103/PhysRevLett.103.103005)
- [29] W. Campbell and J. Doyle,*Cold Molecules: Theory, Experiment, Applications* (Taylor and Francis, Boca Raton, 2009), p. 473.
- [30] J. M. Sage, S. Sainis, T. Bergeman, and D. DeMille, *[Phys. Rev.](http://dx.doi.org/10.1103/PhysRevLett.94.203001)* Lett. **94**[, 203001 \(2005\).](http://dx.doi.org/10.1103/PhysRevLett.94.203001)
- [31] K.-K. Ni, S. Ospelkaus, M. H. G. de Miranda, A. Pe'er, B. Neyenhuis, J. J. Zirbel, S. Kotochigova, P. S. Julienne, D. S. Jin, and J. Ye, Science **322**[, 231 \(2008\).](http://dx.doi.org/10.1126/science.1163861)
- [32] S. Ospelkaus, K.-K. Ni, G. Quéméner, B. Neyenhuis, D. Wang, M. H. G. de Miranda, J. L. Bohn, J. Ye, and D. S. Jin, [Phys. Rev.](http://dx.doi.org/10.1103/PhysRevLett.104.030402) Lett. **104**[, 030402 \(2010\).](http://dx.doi.org/10.1103/PhysRevLett.104.030402)
- [33] K. Aikawa, D. Akamatsu, M. Hayashi, K. Oasa, J. Kobayashi, P. Naidon, T. Kishimoto, M. Ueda, and S. Inouye, [Phys. Rev.](http://dx.doi.org/10.1103/PhysRevLett.105.203001) Lett. **105**[, 203001 \(2010\).](http://dx.doi.org/10.1103/PhysRevLett.105.203001)
- [34] J. Deiglmayr, A. Grochola, M. Repp, K. Mörtlbauer, C. Glück, J. Lange, O. Dulieu, R. Wester, and M. Weidemüller, *[Phys. Rev.](http://dx.doi.org/10.1103/PhysRevLett.101.133004)* Lett. **101**[, 133004 \(2008\).](http://dx.doi.org/10.1103/PhysRevLett.101.133004)
- [35] E. F. von Dishoeck, *The Physics and Chemistry of Interstellar Molecular Clouds* (Springer, Berlin, 1995).
- [36] W. Kolos and L. Wolniewicz, [J. Chem. Phys.](http://dx.doi.org/10.1063/1.1697142) **43**, 2429 (1965).
- [37] J. Mitroy and M. W. J. Bromley, [Phys. Rev. A](http://dx.doi.org/10.1103/PhysRevA.71.032709) **71**, 032709 [\(2005\).](http://dx.doi.org/10.1103/PhysRevA.71.032709)
- [38] N. Geum, G.-H. Jeung, A. Derevianko, R. Côté, and A. Dalgarno, [J. Chem. Phys.](http://dx.doi.org/10.1063/1.1388044) **115**, 5984 (2001).
- [39] R. Côté (private communication).
- [40] Y. Kagan, B. V. Svistunov, and G. V. Shlyapnikov, JETP Lett. **42**, 209 (1985).
- [41] L. M. Delves, Nucl. Phys. **9**, 391 (1958); **20**[, 275 \(1960\).](http://dx.doi.org/10.1016/0029-5582(60)90174-7)
- [42] C. D. Lin, [Phys. Rep.](http://dx.doi.org/10.1016/0370-1573(94)00094-J) **257**, 1 (1995).
- [43] P. Soldán, M. T. Cvitaš, and J. M. Hutson, *[Phys. Rev. A](http://dx.doi.org/10.1103/PhysRevA.67.054702) 67*, [054702 \(2003\).](http://dx.doi.org/10.1103/PhysRevA.67.054702)
- [44] M. Marinescu and A. F. Starace, Phys. Rev. A **55**[, 2067 \(1997\).](http://dx.doi.org/10.1103/PhysRevA.55.2067)
- [45] J. P. D'Incao, C. H. Greene, and B. D. Esry, [J. Phys. B](http://dx.doi.org/10.1088/0953-4075/42/4/044016) **42**, [044016 \(2009\).](http://dx.doi.org/10.1088/0953-4075/42/4/044016)
- [46] B. D. Esry, C. H. Greene, and H. Suno, [Phys. Rev. A](http://dx.doi.org/10.1103/PhysRevA.65.010705) **65**, 010705 [\(2001\).](http://dx.doi.org/10.1103/PhysRevA.65.010705)
- [47] B. D. Esry, Ph.D. thesis, University of Colorado, Boulder, 1997, App. C.
- [48] E. Nielsen, D. V. Fedorov, A. S. Jensen, and E. Garrido, *[Phys.](http://dx.doi.org/10.1016/S0370-1573(00)00107-1)* Rep. **347**[, 373 \(2001\).](http://dx.doi.org/10.1016/S0370-1573(00)00107-1)
- [49] J. P. Burke, Ph.D. thesis, University of Colorado, Boulder, 1999.
- [50] See supplemental material at [\[http://link.aps.org/supplemental](http://link.aps.org/supplemental/10.1103/PhysRevA.83.032703) $/10.1103$ /PhysRevA.83.032703] for numerical data of K_3 and *σ*2.
- [51] H. Suno, B. D. Esry, and C. H. Greene, [Phys. Rev. Lett.](http://dx.doi.org/10.1103/PhysRevLett.90.053202) **90**, [053202 \(2003\);](http://dx.doi.org/10.1103/PhysRevLett.90.053202) J. P. D'Incao, H. Suno, and B. D. Esry, *[ibid.](http://dx.doi.org/10.1103/PhysRevLett.93.123201)* **93**, [123201 \(2004\).](http://dx.doi.org/10.1103/PhysRevLett.93.123201)
- [52] B. D. Esry and H. R. Sadeghpour, [Phys. Rev. A](http://dx.doi.org/10.1103/PhysRevA.68.042706) **68**, 042706 [\(2003\).](http://dx.doi.org/10.1103/PhysRevA.68.042706)
- [53] Y. Wang and B. D. Esry (to be published).