# **Instability of superfluid Fermi gases induced by a rotonlike density mode in optical lattices**

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(Received 21 April 2009; published 29 December 2009)

We study the stability of superfluid Fermi gases in deep optical lattices in the BCS–Bose-Einstein condensation (BEC) crossover at zero temperature. Within the tight-binding attractive Hubbard model, we calculate the spectrum of the low-energy Anderson-Bogoliubov (AB) mode as well as the single-particle excitations in the presence of superfluid flow in order to determine the critical velocities. To obtain the spectrum of the AB mode, we calculate the density response function in the generalized random-phase approximation applying the Green's function formalism developed by Côté and Griffin to the Hubbard model. We find that the spectrum of the AB mode is separated from the particle-hole continuum having the characteristic rotonlike minimum at short wavelength due to the strong charge-density-wave fluctuations. The energy of the rotonlike minimum decreases with increasing the lattice velocity and it reaches zero at the critical velocity which is smaller than the pair-breaking velocity. This indicates that the superfluid state is energetically unstable due to the spontaneous emission of the short-wavelength rotonlike excitations of the AB mode instead due to pair breaking. We determine the critical velocities as functions of the interaction strength across the BCS-BEC crossover regime.

DOI: [10.1103/PhysRevA.80.063627](http://dx.doi.org/10.1103/PhysRevA.80.063627)

PACS number(s): 03.75.Ss, 37.10.Jk, 71.10.Fd

## **I. INTRODUCTION**

The recent realization of superfluidity in Fermi gases  $[1-6]$  $[1-6]$  $[1-6]$  has opened a new research frontier in ultracold atoms [[7](#page-12-2)]. A great experimental advantage of this system is the ability in controlling atomic interactions using a Feshbach resonance  $[8]$  $[8]$  $[8]$ . This allows us to access the crossover between the Bardeen-Cooper-Shrieffer (BCS)-type superfluidity and Bose-Einstein condensation (BEC) of bound molecules, which is referred to as the BCS–BEC crossover [[9–](#page-12-4)[16](#page-13-0)]. The study of superfluid Fermi gases in the BCS-BEC crossover is expected to offer new insights into the phenomena of superfluidity and superconductivity, which can be applied in various fields such as condensed-matter physics, nuclear physics, and particle physics.

One of the most dramatic features of a superfluid system is the dissipationless superfluid flow  $[17]$  $[17]$  $[17]$ . In particular, critical velocities of superfluid flow have attracted much interest in various systems such as superfluid  ${}^{4}$ He [[18](#page-13-2)], superfluid  ${}^{3}$ He [[19](#page-13-3)], and atomic Bose-Einstein condensates [[20](#page-13-4)[,21](#page-13-5)]. It is well known that the underlying mechanisms for the instability of dissipationless flow are different in the BCS and BEC regions in a uniform system. Namely, the instability of BCStype superfluids is considered to be dominated by Cooper pair breaking  $\lceil 22 \rceil$  $\lceil 22 \rceil$  $\lceil 22 \rceil$ , whereas the instability of Bose superfluids is induced by spontaneous emission of phonon excitations  $\lceil 23 \rceil$  $\lceil 23 \rceil$  $\lceil 23 \rceil$ . It is of interest to study how the mechanism of the instability in superfluid Fermi gases changes in the BCS-BEC crossover.

Recently, Miller *et al.* investigated experimentally the stability of superfluid flow in Fermi gases in *shallow* onedimensional (1D) optical lattices across the BCS-BEC crossover  $[24]$  $[24]$  $[24]$ . They measured superfluid critical velocities, at which the number of condensed atoms starts to decrease, by moving the optical lattice potential through the atomic cloud for different values of interatomic interaction and lattice depth  $[24]$  $[24]$  $[24]$ . The measured critical velocities showed a crossover behavior between the BCS and BEC regimes taking a maximum value at the crossover regime  $[24]$  $[24]$  $[24]$ . Critical velocities in superfluid Fermi gases in the BCS-BEC crossover have been also addressed theoretically in several papers  $[25-27]$  $[25-27]$  $[25-27]$ . The observed crossover behavior of the critical velocities has been predicted in Refs.  $[25,26]$  $[25,26]$  $[25,26]$  $[25,26]$ . However, most of the theoretical papers are limited within a uniform system  $\lceil 25 \rceil$  $\lceil 25 \rceil$  $\lceil 25 \rceil$  or a system in the presence of a single potential barrier  $[26]$  $[26]$  $[26]$ , which cannot be directly compared to the experiment using optical lattices in Ref. [[24](#page-13-8)]. In Ref. [[27](#page-13-10)], sound propagation in superfluid Fermi gases in optical lattices has been studied using the hydrodynamic approximation. However, microscopic calculation of the critical velocities of superfluid Fermi gases in optical lattices has not been worked out yet.

In this paper, we study the stability and critical velocities of superfluid Fermi gases in *deep* one-dimensional, twodimensional (2D), and three-dimensional (3D) optical lattices in the BCS-BEC crossover at zero temperature. We apply the generalized random-phase approximation (GRPA) developed by Côté and Griffin  $[28]$  $[28]$  $[28]$  to the attractive tightbinding Hubbard model in order to calculate the excitation spectra in the presence of a moving optical lattice. For the stability of Fermi gases in the BCS-BEC crossover, two kinds of excitations play crucial roles. One is the singleparticle excitation which arises when Cooper pairs are broken. The other is the collective density-fluctuation mode, the so-called Anderson-Bogoliubov (AB) mode [[29,](#page-13-13)[30](#page-13-14)]. In a uniform system, the single-particle excitation induces the instability of superfluid flow in the BCS regime, while phonon excitation of the Bogoliubov mode, which corresponds to the

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AB mode in the BCS regime, induces the instability in the BEC regime  $[24–26]$  $[24–26]$  $[24–26]$  $[24–26]$ . We find that in deep 1D, 2D, and 3D optical lattices, the excitation spectrum of the AB mode has a characteristic rotonlike structure and lies below the particlehole continuum due to the strong charge-density-wave (CDW) fluctuation. The energy of the rotonlike minimum decreases with increasing the superfluid velocity and it reaches zero before the particle-hole continuum does, i.e., before pair-breaking occurs. As a result, in contrast to the uniform case, the instability of superfluid flow in 1D, 2D, and 3D optical lattices is induced by the rotonlike excitations of the AB mode rather than by pair breakings. We calculate the critical velocities at which spontaneous emission of the rotonlike excitations occurs as functions of the interaction strength in the entire BCS-BEC crossover regime.

This paper is organized as follows. In Sec. [II,](#page-1-0) we present the model and formalism. We introduce the tight-binding Hubbard model and the Green's function formalism for the GRPA. In Sec. [III,](#page-5-0) we present the results for the stability and the critical velocities of superfluid Fermi gases in 1D, 2D, and 3D optical lattices. We calculate the excitation spectra and determine critical velocities as functions of the attractive interaction. We summarize our results in Sec. [IV.](#page-11-0)

## **II. MODEL AND FORMALISM**

<span id="page-1-0"></span>In this section, we summarize the Green's function formalism applied to an attractive Hubbard model. This is necessary for the calculation of response functions in GRPA. The excitation spectra of collective modes can be obtained as the poles of the response functions. Since our major interest is in the AB mode, we calculate the density response function assuming an external field coupled with density. To discuss the stability of superfluid states, we extend the previous work for the ground state  $\left[31,32\right]$  $\left[31,32\right]$  $\left[31,32\right]$  $\left[31,32\right]$  to the current-carrying states.

### **A. Green's function formalism**

We consider two-component atomic superfluid Fermi gases with equal populations loaded into optical lattices. We suppose that the optical lattice potential is moving with a constant velocity  $-\mathbf{v}$  in the laboratory frame. If the velocity of the lattice potential does not exceed the critical velocity, the Fermi gas remains stable in the laboratory frame due to its superfluidity. This situation can be described equivalently in the frame fixed with respect to the lattice potential as a superfluid Fermi gas flowing with a constant quasimomentum 2*mv*, where *m* is the mass of a fermion. In the following, we describe the system in the frame fixed with respect to the lattice potential. Namely, we assume a time-independent lattice potential and a supercurrent with the quasimomentum 2*mv*.

We assume that the optical lattice potential is sufficiently deep so that the tight-binding approximation is valid. Thus, the system can be described by a single-band Hubbard model as (we set  $\hbar = k_B = 1$ )

<span id="page-1-3"></span>
$$
H = -J\sum_{\langle i,j\rangle,\sigma} (c_{i\sigma}^{\dagger}c_{j\sigma} + \text{H.c.}) + U\sum_{i} c_{i\uparrow}^{\dagger} c_{i\downarrow}^{\dagger}c_{i\downarrow}c_{i\uparrow} - \mu\sum_{i,\sigma} c_{i\sigma}^{\dagger}c_{i\sigma},
$$
\n(1)

where  $c_{j\sigma}$  is the annihilation operator of a fermion on the *j*th site with pseudospin  $\sigma = \uparrow, \downarrow$ . Here, *J* is the nearest-neighbor hopping energy,  $U$  is the on-site interaction energy, and  $\mu$  is the chemical potential. We assume an attractive interaction between atoms  $(U<0)$ .

In order to calculate the density response function, we introduce a *fictitious* time-dependent external field  $P_j(t)$ which is coupled with the density. The Hamiltonian with the external field is given by

$$
K(t) = H + V(t),\tag{2}
$$

$$
V(t) = \sum_{j} P_j(t)n_j,
$$
\n(3)

where  $n_j \equiv \sum_{\sigma} c_{j\sigma}^{\dagger} c_{j\sigma}$  is the number operator. The density re sponse function is obtained by taking a functional derivative of the single-particle Green's function by the external field. This will be carried out in Sec. [II C.](#page-4-0)

We use the imaginary time Green's function technique [[33](#page-13-17)]. The Heisenberg representations of the annihilation and creation operators in the imaginary time  $\tau$  are defined as

$$
c_{j\sigma}(\tau) = \exp\left(\int_0^{\tau} d\tau' K(\tau')\right) c_{j\sigma} \exp\left(-\int_0^{\tau} d\tau' K(\tau')\right), \tag{4}
$$

$$
c_{j\sigma}^{\dagger}(\tau) = \exp\left(\int_0^{\tau} d\tau' K(\tau')\right) c_{j\sigma}^{\dagger} \exp\left(-\int_0^{\tau} d\tau' K(\tau')\right). \tag{5}
$$

We introduce the normal and anomalous single-particle Green's functions, respectively, as  $[34]$  $[34]$  $[34]$ 

$$
G_{ij,\sigma}(\tau,\tau') = -\langle T(c_{i\sigma}(\tau)c_{j\sigma}^{\dagger}(\tau'))\rangle, \tag{6}
$$

$$
F_{ij}(\tau,\tau') = -\langle T(c_{i\uparrow}(\tau)c_{j\downarrow}(\tau'))\rangle, \tag{7}
$$

where  $T(\cdots)$  represents the time-ordering operator with respect to  $\tau$ . Using the Nambu representation [[35](#page-13-19)] with

$$
\Psi_j(\tau) = \begin{pmatrix} c_{j\uparrow}(\tau) \\ c_{j\downarrow}^{\dagger}(\tau) \end{pmatrix}; \quad \Psi_j^{\dagger}(\tau) = (c_{j\uparrow}^{\dagger}(\tau), c_{j\downarrow}(\tau)), \tag{8}
$$

<span id="page-1-1"></span>the single-particle Green's function can be written in the matrix form as

$$
\hat{G}_{ij}(\tau,\tau') \equiv -\langle T(\Psi_i(\tau)\Psi_j^{\dagger}(\tau'))\rangle \n= \begin{pmatrix} G_{ij,\uparrow}(\tau,\tau') & F_{ij}(\tau,\tau') \\ F_{ij}^*(\tau,\tau') & -G_{ji,\downarrow}(\tau',\tau) \end{pmatrix} .
$$
\n(9)

We note that in the absence of the external field, the Green's function at equal sites and imaginary times  $\hat{G}_{jj}(\tau,\tau)$  is given by

$$
\hat{G}_{jj}(\tau,\tau) \equiv \lim_{\tau' \to \tau+0} \hat{G}_{jj}(\tau,\tau') \tag{10}
$$

$$
= \begin{pmatrix} \langle n_{j\uparrow} \rangle & \langle m_{j} \rangle \\ \langle m_{j} \rangle^* & 1 - \langle n_{j\downarrow} \rangle \end{pmatrix}, \tag{11}
$$

<span id="page-1-2"></span>where  $n_{j\sigma} \equiv c_{j\sigma}^{\dagger} c_{j\sigma}$  and  $m_j \equiv c_{j\downarrow} c_{j\uparrow}$  is the pair annihilation operator which is related to the wave function of Cooper pairs, as we discuss below.

From the equations of motion for  $c_{j\sigma}(\tau)$  and  $c_{j\sigma}^{\dagger}(\tau)$ , we obtain the equation for the matrix Green's function in Eq.  $(9)$  $(9)$  $(9)$ as

<span id="page-2-0"></span>
$$
\left(-\frac{\partial}{\partial \tau} + \mu \hat{\sigma}_3\right) \hat{G}_{ij}(\tau, \tau') + 2J \sum_{\langle l,m \rangle} \delta_{i,l} \hat{\sigma}_3 \hat{G}_{mj}(\tau, \tau')\n= \delta_{i,j} \delta(\tau - \tau') - U \hat{\sigma}_3 \langle T(n_i(\tau) \Psi_i(\tau) \Psi_j^{\dagger}(\tau')) \rangle\n+ P_i(\tau) \hat{\sigma}_3 \hat{G}_{ij}(\tau, \tau'), \qquad (12)
$$

where  $\hat{\sigma}_3$  is the Pauli matrix

$$
\hat{\sigma}_3 = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} . \tag{13}
$$

<span id="page-2-1"></span>The noninteracting Green's function  $\hat{G}_{ij}^0(\tau, \tau')$  satisfies

$$
\left(-\frac{\partial}{\partial \tau} + \mu \hat{\sigma}_3\right) \hat{G}_{ij}^0(\tau, \tau') + 2J \sum_{\langle l,m \rangle} \delta_{i,l} \hat{\sigma}_3 \hat{G}_{mj}(\tau, \tau')\n= \delta_{i,j} \delta(\tau - \tau').
$$
\n(14)

From Eqs.  $(12)$  $(12)$  $(12)$  and  $(14)$  $(14)$  $(14)$ , the Green's function satisfies the Dyson equation

<span id="page-2-2"></span>
$$
\hat{G}_{ij}(\tau,\tau') = \hat{G}_{ij}^{0}(\tau,\tau') + \sum_{l,m} \int_{0}^{\beta} d\tau_{1} \int_{0}^{\beta} d\tau_{2} \hat{G}_{il}^{0}(\tau,\tau_{1})
$$
\n
$$
\times \hat{\Sigma}_{lm}(\tau_{1},\tau_{2}) \hat{G}_{mj}(\tau_{2},\tau')
$$
\n
$$
+ \sum_{l} \int_{0}^{\beta} d\tau_{1} \hat{G}_{il}^{0}(\tau,\tau_{1}) P_{l}(\tau_{1}) \hat{\sigma}_{3} \hat{G}_{lj}(\tau_{1},\tau'), \qquad (15)
$$

where  $\beta = 1/T$  and *T* is the temperature. In Eq. ([15](#page-2-2)), the self-energy  $\hat{\Sigma}_{ij}(\tau, \tau')$  is given by

<span id="page-2-4"></span>
$$
\hat{\Sigma}_{ij}(\tau,\tau') = -U\hat{\sigma}_3 \sum_{l} \int_0^{\beta} d\tau_1 \langle T(n_i(\tau)\Psi_i(\tau)\Psi_i^{\dagger}(\tau_1)) \rangle
$$

$$
\times [\hat{G}_{lj}(\tau_1,\tau')]^{-1}.
$$
 (16)

Here, we introduced the inverse matrix Green's function  $[\hat{G}_{ij}(\tau, \tau')]^{-1}$ , which satisfies

<span id="page-2-3"></span>
$$
\sum_{l} \int_{0}^{\beta} d\tau_{1} \hat{G}_{il}(\tau, \tau_{1}) [\hat{G}_{lj}(\tau_{1}, \tau')]^{-1} = \delta_{i,j} \delta(\tau - \tau'). \quad (17)
$$

Using Eqs.  $(15)$  $(15)$  $(15)$  and  $(17)$  $(17)$  $(17)$  can be simplified as

$$
[\hat{G}_{ij}(\tau,\tau')]^{-1} = [\hat{G}_{ij}^{0}(\tau,\tau')]^{-1} - \hat{\Sigma}_{ij}(\tau,\tau') - P_i(\tau)\hat{\sigma}_3\delta_{i,j}\delta(\tau-\tau').
$$
\n(18)

To calculate the self-energy in Eq.  $(16)$  $(16)$  $(16)$ , we use the Hartree-Fock-Gor'kov (HFG) approximation [[28](#page-13-12)[,34](#page-13-18)]

<span id="page-2-7"></span>
$$
\langle T(n_l(\tau_2)\Psi_i(\tau)\Psi_j^{\dagger}(\tau_1))\rangle
$$
  

$$
\approx -\langle n_l(\tau_2)\rangle \hat{G}_{ij}(\tau,\tau_1) + \hat{G}_{il}(\tau,\tau_2)\hat{\sigma}_3 \hat{G}_{lj}(\tau_2,\tau_1). \quad (19)
$$

Thus, the self-energy in the HFG approximation is given by

$$
\hat{\Sigma}_{ij}(\tau,\tau') \simeq \hat{\Sigma}_{ij}^{\text{HFG}}(\tau,\tau')
$$
  
=  $U[\langle n_i(\tau) \rangle \hat{\sigma}_3 - \hat{\sigma}_3 \hat{G}_{ii}(\tau,\tau) \hat{\sigma}_3] \delta_{i,j} \delta(\tau - \tau').$  (20)

In the presence of supercurrent with velocity  $v$ , Cooper pairs are Bose-condensed into the state with the center-of-mass quasimomentum  $q = 2mv$ . Since the anomalous Green's function  $F_{jj}(\tau, \tau)$  can be regarded as the wave function of Cooper pairs  $\left[\dot{3}4\right]$ , it can be written as

$$
F_{jj}(\tau,\tau) = -\langle m_j \rangle = \frac{\Delta_v}{|U|} \exp(2im\mathbf{v} \cdot \mathbf{r}_j), \tag{21}
$$

<span id="page-2-5"></span>where  $\Delta_v$  is the superfluid gap and  $r_j$  is the location of the *j*th site. The exponential factor on the right-hand side of Eq.  $(21)$  $(21)$  $(21)$ describes the supercurrent with quasimomentum 2*mv*.

In the presence of the supercurrent, the normal and anomalous Green's functions can be written as

$$
G_{ij,\sigma}(\tau,\tau') = \exp(imv \cdot r_{ij})\tilde{G}_{ij,\sigma}(\tau,\tau'),\tag{22}
$$

$$
F_{ij}(\tau, \tau') = \exp(2im\boldsymbol{v} \cdot \boldsymbol{R}_{ij}) \widetilde{F}_{ij}(\tau, \tau'), \qquad (23)
$$

<span id="page-2-8"></span>respectively, where  $r_{ij} \equiv r_i - r_j$  is the relative coordinate and  $R_{ij} \equiv (r_i + r_j)/2$  is the center-of-mass coordinate of the Cooper pair. Here,  $\tilde{G}_{ij,\sigma}(\tau,\tau')$  and  $\tilde{F}_{ij}(\tau,\tau')$  are functions of  $r_{ij}$ .

To eliminate the phase factors associated with the supercurrent, it is convenient to introduce an operator  $\tilde{\Psi}_j(\tau)$  and a matrix Green's function  $\hat{G}_{ij}(\tau, \tau')$  as

$$
\widetilde{\Psi}_j(\tau) = \begin{pmatrix} c_{j\uparrow}(\tau) \exp(-im\boldsymbol{v} \cdot \boldsymbol{r}_j) \\ c_{j\downarrow}^{\dagger}(\tau) \exp(im\boldsymbol{v} \cdot \boldsymbol{r}_j) \end{pmatrix} = \hat{\gamma}_j \Psi_j(\tau), \qquad (24)
$$

<span id="page-2-6"></span>
$$
\hat{\tilde{G}}_{ij}(\tau,\tau') = -\langle T(\tilde{\Psi}_i(\tau)\tilde{\Psi}_j^{\dagger}(\tau'))\rangle = \hat{\gamma}_i \hat{G}_{ij}(\tau,\tau')\hat{\gamma}_j^*
$$
\n
$$
= \begin{pmatrix}\n\tilde{G}_{ij,1}(\tau,\tau') & \tilde{F}_{ij}(\tau,\tau') \\
\tilde{F}_{ij}(\tau,\tau') & -\tilde{G}_{ji,1}(\tau',\tau)\n\end{pmatrix},
$$
\n(25)

where the matrix  $\hat{\gamma}_i$  for the unitary transformation between  $\hat{G}$ and  $\hat{\tilde{G}}$  is given by

$$
\hat{\gamma}_j = \begin{pmatrix} \exp(-im\mathbf{v} \cdot \mathbf{r}_j) & 0\\ 0 & \exp(im\mathbf{v} \cdot \mathbf{r}_j \end{pmatrix} . \tag{26}
$$

Using Eq.  $(11)$  $(11)$  $(11)$  and  $(25)$  $(25)$  $(25)$ , in the absence of the external field,  $\hat{G}_{jj}(\tau,\tau)$  reduces to

$$
\hat{G}_{jj}(\tau,\tau) = \begin{pmatrix} \langle n_{j\uparrow} \rangle & \Delta_v / |U| \\ \Delta_v^* / |U| & 1 - \langle n_{j\downarrow} \rangle \end{pmatrix}.
$$
 (27)

In the following sections, we derive equations for  $\hat{G}$ .

## **B. Equilibrium Green's function**

In this section, we calculate the equilibrium Green's function in the absence of the external field  $P_j(t)$  within the HFG approximation introduced in Eq. ([19](#page-2-7)). Since  $\hat{G}_{ij}(\tau, \tau')$  in Eq.

 $(25)$  $(25)$  $(25)$  is a function of  $r_{ij}$ , we define the Fourier transform of the Green's function as

<span id="page-3-0"></span>
$$
\hat{\tilde{G}}_{ij}(\tau,\tau') = \frac{1}{\beta M} \sum_{k,\omega_n} \exp[i\mathbf{k} \cdot \mathbf{r}_{ij} - i\omega_n(\tau - \tau')] \hat{\tilde{G}}_k(i\omega_n),
$$
\n(28)

where *M* is the number of lattice sites and  $\omega_n \equiv (2n)^2$  $+1\pi/\beta$  is the Fermi Matsubara frequency. We note that  $\hat{G}_{ij}(\tau, \tau')$  cannot be expanded as Eq. ([28](#page-3-0)) because the phase factor in Eq.  $(23)$  $(23)$  $(23)$  describing the supercurrent depends on the center-of-mass coordinate *Rij*.

From Eq. ([15](#page-2-2)), we obtain the Dyson equation in Fourier space in the absence of the external field as

$$
\hat{\tilde{G}}_k(i\omega_n) = \hat{\tilde{G}}_k^0(i\omega_n) + \hat{\tilde{G}}_k^0(i\omega_n)\hat{\tilde{\Sigma}}_k(i\omega_n)\hat{\tilde{G}}_k(i\omega_n). \tag{29}
$$

<span id="page-3-1"></span>Here,  $\hat{\Sigma}_k(i\omega_n)$  is the Fourier transform of  $\hat{\Sigma}_{ij}(\tau,\tau')$  $\equiv \hat{\gamma}_i \hat{\Sigma}_{ij}(\tau, \tau') \hat{\gamma}_j^*$ . In Eq. ([29](#page-3-1)), the unperturbed Green's function  $\hat{G}_{k}^{0}(i\omega_{n})$  is given by

$$
\hat{G}_k^0(i\omega_n) = \begin{pmatrix} \frac{1}{i\omega_n - \xi_{k+mv}} & 0 \\ 0 & \frac{1}{i\omega_n + \xi_{k-mv}} \end{pmatrix},
$$
(30)

where  $\xi_k = 2J\Sigma_v(1 - \cos k_v d) - \mu$  is the kinetic energy,  $\nu$  is the index for spatial dimension, and *d* is the lattice constant. From Eq.  $(19)$  $(19)$  $(19)$ , we obtain the self-energy in Eq.  $(29)$  $(29)$  $(29)$  as

$$
\hat{\Sigma}_k(i\omega_n) = \begin{pmatrix} 0 & \Delta_v \\ \Delta_v^* & 0 \end{pmatrix}.
$$
 (31)

<span id="page-3-2"></span>In deriving Eq.  $(31)$  $(31)$  $(31)$ , we shifted the chemical potential by the Hartree-Fock energy *nU*/2, where *n* is the average number of atoms per site.

<span id="page-3-7"></span>By solving Eq.  $(29)$  $(29)$  $(29)$  with the self-energy in Eq.  $(31)$  $(31)$  $(31)$ , we obtain the single-particle Green's function as

$$
\hat{\tilde{G}}_k(i\omega_n) = \frac{\hat{A}_k}{i\omega_n - E_k^+} + \frac{\hat{B}_k}{i\omega_n - E_k^-},
$$
\n(32)

where

$$
\hat{A}_k = \begin{pmatrix} u_k^2 & u_k v_k^* \\ u_k v_k & |v_k|^2 \end{pmatrix},
$$
\n(33)

$$
\hat{B}_k = \begin{pmatrix} |v_k|^2 & -u_k v_k^* \\ -u_k v_k & u_k^2 \end{pmatrix},
$$
\n(34)

$$
u_k^2 = \frac{1}{2} \left( 1 + \frac{\overline{\xi}_k}{\mathcal{E}_k} \right),\tag{35}
$$

$$
|v_k|^2 = \frac{1}{2} \left( 1 - \frac{\overline{\xi}_k}{\mathcal{E}_k} \right),\tag{36}
$$

<span id="page-3-4"></span>

FIG. 1. Single-particle excitation spectra  $E_k^+$  (upper curves) and  $E_k^-$  (lower curves) in 3D optical lattices when the superfluid velocity  $|v|$  is 0 (solid line),  $0.2/md$  (dashed line),  $0.4/md$  (dash-dotted line), and  $|v_{pb}|$  (dotted line). Here, the superfluid flows along the  $(\pi,\pi,\pi)$ direction and  $|v_{pb}|$ =0.628/*md*. We set *U*/*J*=−6.0, *n*=0.5, and *k<sub>x</sub>*  $=k_y = k_z$ .

$$
u_k v_k^* = \frac{\Delta_v}{2\mathcal{E}_k}.\tag{37}
$$

<span id="page-3-3"></span>Here, the single-particle excitation energy is given by

$$
E_k^{\pm} = \eta_k \pm \mathcal{E}_k,\tag{38}
$$

where  $\mathcal{E}_k = \sqrt{\xi_k^2 + |\Delta_v|^2}, \ \bar{\xi}_k = (\xi_{k+mv} + \xi_{k-mv})/2, \text{ and } \eta_k = (\xi_{k+mv}$  $-\xi_{k-m\nu}$ )/2. Equation ([38](#page-3-3)) explicitly shows that  $E_k^{\pm}$  depend on the superfluid velocity *v*. The single-particle excitation spec-trum in Eq. ([38](#page-3-3)) is shown in Fig. [1](#page-3-4) for different superfluid velocities. In Fig. [1,](#page-3-4) the energy gap becomes smaller as  $|v|$ increases. When the energy gap reaches zero, pair breaking occurs, i.e.,  $v = v_{\text{pb}}$ , where  $v_{\text{pb}}$  is the pair-breaking velocity  $\lceil 36 \rceil$  $\lceil 36 \rceil$  $\lceil 36 \rceil$ .

<span id="page-3-5"></span>We determine the superfluid gap  $\Delta_n$  and the chemical potential  $\mu$  by solving self-consistently the number equation

$$
n = \frac{2}{M} \sum_{k} \left[ (u_k^2 - |v_k|^2) f(E_k^+) + |v_k|^2 \right] \tag{39}
$$

<span id="page-3-6"></span>and the gap equation

$$
\Delta_v = -\frac{U}{M} \sum_k u_k v_k^* [1 - 2f(E_k^+)].
$$
\n(40)

Equations  $(39)$  $(39)$  $(39)$  and  $(40)$  $(40)$  $(40)$  are obtained from the diagonal and off-diagonal elements of the Green's function in Eq. ([32](#page-3-7)). Here,  $f(\varepsilon) \equiv 1/[\exp(\beta \varepsilon) + 1]$  is the Fermi distribution function. This scheme of solving Eqs.  $(39)$  $(39)$  $(39)$  and  $(40)$  $(40)$  $(40)$  selfconsistently interpolates the weak-coupling BCS limit and strong-coupling BEC limit at low temperature when the fluctuation effect due to pairs with finite center-of-mass momenta can be neglected  $[10]$  $[10]$  $[10]$ . Throughout the work including the calculation of  $\Delta$ <sub>v</sub> and  $\mu$ , we assume *T*=0.

In Eq. ([40](#page-3-6)),  $\Delta_{v}$  depends on superfluid velocity  $v$  via  $u_{k}$ ,  $v_k$ , and  $E_k^+$ . To explicitly show this, we plot  $\Delta_v$  in 3D at *T*  $= 0$  as a function of  $|v| \le |v_{pb}|$  in Fig. [2.](#page-4-1) Note that when

<span id="page-4-1"></span>

FIG. 2. Superfluid gap  $\Delta_{v}$  as a function of  $|v|$  in 3D optical lattices when the superfluid flows along the  $(\pi,\pi,\pi)$  direction. Here,  $|\Delta_{v=0}| = 1.92J$  and  $|v_{pb}| = 0.628/md$ . We set  $U/J = -6.0$  and *n*  $= 0.5.$ 

 $E_k^+ = 0$  at  $|v| = |v_{pb}|$ , the superfluid gap does not vanish  $(\Delta_v)$  $\neq$  0), but the superfluid state is destabilized due to pair breaking.

### **C. Response function**

<span id="page-4-0"></span>In this section, we calculate the density response function. The density response function can be derived by taking a functional derivative of the density by the external field as

$$
\chi_{ij}(\tau,\tau') = \frac{\delta \langle n_i(\tau) \rangle}{\delta P_j(\tau')}.
$$
\n(41)

<span id="page-4-8"></span><span id="page-4-2"></span>Here, we introduce the three-point correlation function  $\hat{L}_{ijl}(\tau,\tau',\tau_1)$  as

$$
\hat{L}_{ijl}(\tau,\tau',\tau_1) \equiv \frac{\partial \hat{\overline{G}}_{ij}(\tau,\tau')}{\delta P_l(\tau_1)},\tag{42}
$$

where  $\hat{G}_{ij}(\tau, \tau') \equiv \sigma_3 \hat{G}_{ij}(\tau, \tau')$ . Using Eq. ([42](#page-4-2)), the density response function can be written as

<span id="page-4-3"></span>
$$
\chi_{ij}(\tau,\tau') = -\langle T(\delta n_i(\tau)\delta n_j(\tau'))\rangle = L_{ij}^{11}(\tau,\tau') + L_{ij}^{22}(\tau,\tau'),
$$
\n(43)

where  $\delta n_i(\tau) \equiv n_i(\tau) - \langle n_i(\tau) \rangle$  is the density-fluctuation operator and

$$
\hat{L}_{ij}(\tau,\tau') \equiv \lim_{\tau_1 \to \tau^+} \hat{L}_{ij}(\tau,\tau_1,\tau'). \tag{44}
$$

In deriving Eq.  $(43)$  $(43)$  $(43)$ , we used the functional differentiation of the single-particle Green's function by the external field as

<span id="page-4-4"></span>
$$
\frac{\partial \hat{\tilde{G}}_{ij}(\tau,\tau')}{\delta P_l(\tau_1)} = \langle T(\tilde{\Psi}_i(\tau)\tilde{\Psi}_j^{\dagger}(\tau')n_l(\tau_1))\rangle + \hat{\tilde{G}}_{ij}(\tau,\tau')\langle n_l(\tau_1)\rangle.
$$
\n(45)

We note that Eq.  $(45)$  $(45)$  $(45)$  is valid within the linear-response regime.

We derive the equation of motion for the three-point correlation function  $\hat{L}$ . Differentiating Eq. ([17](#page-2-3)) with respect to the external field, we obtain

$$
\frac{\delta \hat{G}_{ij}(\tau,\tau')}{\delta P_l(\tau_1)} = \sum_{m,n} \int_0^\beta d\tau_2 \int_0^\beta d\tau_3 \hat{G}_{im}(\tau,\tau_2)
$$
  
 
$$
\times \frac{\hat{\delta \Sigma}_{mn}(\tau_2,\tau_3)}{\delta P_l(\tau_1)} \hat{G}_{nj}(\tau_3,\tau')
$$
  
+ 
$$
\sum_m \int_0^\beta d\tau_2 \hat{\tilde{G}}_{im}(\tau,\tau_2) \frac{\delta P_m(\tau_2)}{\delta P_l(\tau_1)} \hat{\sigma}_3 \hat{\tilde{G}}_{mj}(\tau_2,\tau').
$$
(46)

Using the HFG approximation in Eq.  $(19)$  $(19)$  $(19)$ , the three-point correlation function satisfies

*ˆ*

<span id="page-4-5"></span>
$$
\hat{L}_{ijl}(\tau, \tau', \tau_1) = \hat{L}_{ijl}^0(\tau, \tau', \tau_1) \n+ U \sum_{m} \int_0^{\beta} d\tau_2 \hat{G}_{im}(\tau, \tau_2) \hat{G}_{mj}(\tau_2, \tau') \chi_{ml}(\tau_2, \tau_1) \n- U \sum_{m} \int_0^{\beta} d\tau_2 \hat{G}_{im}(\tau, \tau_2) \hat{L}_{ml}(\tau_2, \tau') \hat{G}_{mj}(\tau_2, \tau_1),
$$
\n(47)

where the lowest-order correlation function is given by  $\hat{L}_{ijl}^0(\tau, \tau', \tau_1) = \hat{G}_{il}(\tau, \tau_1) \hat{G}_{lj}(\tau_1, \tau')$ . Thus, the density response function can be obtained by solving Eq.  $(47)$  $(47)$  $(47)$  which is referred to as the GRPA equation  $\lceil 28 \rceil$  $\lceil 28 \rceil$  $\lceil 28 \rceil$ .

To see the diagrammatic structure of Eq.  $(47)$  $(47)$  $(47)$  more clearly, it is useful to rewrite Eq.  $(47)$  $(47)$  $(47)$  in terms of the irreducible correlation function

<span id="page-4-6"></span>
$$
\hat{L}_{ijl}(\tau,\tau',\tau_{1}) = \hat{L}_{ijl}^{0}(\tau,\tau',\tau_{1}) - U \sum_{m} \int_{0}^{\beta} d\tau_{2} \hat{\overline{G}}_{im}(\tau,\tau_{2}) \hat{\overline{L}}_{ml}(\tau_{2},\tau_{1}) \hat{\overline{G}}_{mj}(\tau_{2},\tau').
$$
\n(48)

Using Eq.  $(48)$  $(48)$  $(48)$ , Eq.  $(47)$  $(47)$  $(47)$  reduces to

<span id="page-4-7"></span>
$$
\hat{L}_{ijl}(\tau, \tau', \tau_1) = \hat{\bar{L}}_{ijl}(\tau, \tau', \tau_1) \n+ U \sum_{m} \int_0^{\beta} d\tau_2 \hat{\bar{L}}_{ijm}(\tau, \tau', \tau_2) \chi_{ml}(\tau_2, \tau_1).
$$
\n(49)

It is clear from Eq. ([48](#page-4-6)) that  $\hat{L}$  includes the ladder diagrams. On the other hand, Eq.  $(49)$  $(49)$  $(49)$  includes the bubble diagrams which lead to the random-phase approximation (RPA)  $[28]$  $[28]$  $[28]$ . In a homogeneous system, the contribution from the bubble diagrams for an attractive interaction can be neglected  $\lceil 28 \rceil$  $\lceil 28 \rceil$  $\lceil 28 \rceil$ . However, in our lattice system, since the bubble diagrams induce the instability due to the CDW fluctuation, it is crucial for the analysis of the stability of the system to keep the bubble diagrams in Eq.  $(49)$  $(49)$  $(49)$ . We compare the excitation spectra with and without the contribution from the bubble diagrams and give a detailed discussion of the effects of the CDW fluctuations on the stability of the system in Sec. [III.](#page-5-0)

We solve Eqs.  $(48)$  $(48)$  $(48)$  and  $(49)$  $(49)$  $(49)$  to calculate the density response function in Eq.  $(41)$  $(41)$  $(41)$ . We define the Fourier transform of  $\hat{L}_{ij}(\tau,\tau')$  as

$$
\hat{L}_{ij}(\tau,\tau') = \frac{1}{\beta M} \sum_{q,\Omega_n} \exp[iq \cdot r_{ij} - i\Omega_n(\tau - \tau')] \hat{L}_q(i\Omega_n),
$$
\n(50)

where  $\Omega_n = 2n\pi/\beta$  is the Bose Matsubara frequency. From Eq. ([43](#page-4-3)), the Fourier component of the density response function is given by  $\chi_q(i\Omega_n) = L_q^{11}(i\Omega_n) + L_q^{22}(i\Omega_n)$ .

The Fourier transforms of Eqs.  $(48)$  $(48)$  $(48)$  and  $(49)$  $(49)$  $(49)$  are represented as

<span id="page-5-1"></span>
$$
\hat{\overline{L}}_q(i\Omega_n) = \hat{L}_q^0(i\Omega_n) - \frac{U}{\beta M} \sum_{k,\omega_n} \hat{\overline{G}}_k(i\omega_n) \hat{\overline{L}}_q(i\Omega_n) \hat{\overline{G}}_{k-q}(i\omega_n - i\Omega_n)
$$
\n(51)

<span id="page-5-2"></span>and

$$
\hat{L}_q(i\Omega_n) = \hat{\overline{L}}_q(i\Omega_n) + U\hat{\overline{L}}_q(i\Omega_n)\chi_q(i\Omega_n),
$$
\n(52)

respectively, where

$$
\hat{L}_q^0(i\Omega_n) = \frac{1}{\beta M} \sum_{k,\omega_n} \hat{\bar{G}}_k(i\omega_n) \hat{\bar{G}}_{k-q}(i\omega_n - i\Omega_n).
$$
 (53)

In order to rewrite Eqs.  $(51)$  $(51)$  $(51)$  and  $(52)$  $(52)$  $(52)$  in simpler forms, we define a column vector  $\mathcal{L}_{q}(i\Omega_{n})$  as

$$
\mathcal{L}_q(i\Omega_n) \equiv \begin{pmatrix} L^{11} \\ L^{12} \\ L^{21} \\ L^{22} \end{pmatrix} . \tag{54}
$$

Here, we have used the notation  $L^{\mu\nu} \equiv L_q^{\mu\nu} (i\Omega_n)$ . The same notation is adapted for  $\overline{L}_q(i\Omega_n)$  and  $\mathcal{L}_q^0(i\Omega_n)$ . In addition, we define a  $4 \times 4$  matrix  $\hat{\mathcal{D}}$  as

<span id="page-5-9"></span>
$$
\hat{\mathcal{D}}_q(i\Omega_n) \equiv \begin{pmatrix} D^{1111} & D^{1121} & D^{1211} & D^{1221} \\ D^{1112} & D^{1122} & D^{1212} & D^{1222} \\ D^{2111} & D^{2121} & D^{2211} & D^{2221} \\ D^{2112} & D^{2122} & D^{2212} & D^{2222} \end{pmatrix}, \qquad (55)
$$

<span id="page-5-3"></span>where

$$
D_q^{\mu\nu\rho\lambda}(i\Omega_n) \equiv \frac{1}{\beta M} \sum_{k,\omega_n} \bar{G}_{k+q}^{\mu\nu}(i\omega_n + i\Omega_n) \bar{G}_k^{\rho\lambda}(i\omega_n). \quad (56)
$$

<span id="page-5-4"></span>From Eq. ([56](#page-5-3)),  $\hat{D}$  describes a single bubble diagram of a particle-hole excitation. Thus, Eqs.  $(51)$  $(51)$  $(51)$  and  $(52)$  $(52)$  $(52)$  can be written in the matrix forms as

$$
\overline{\mathcal{L}}_q(i\Omega_n) = \mathcal{L}_q^0(i\Omega_n) - U\hat{\mathcal{D}}_q(i\Omega_n)\overline{\mathcal{L}}_q(i\Omega_n),\tag{57}
$$

$$
\mathcal{L}_q(i\Omega_n) = \overline{\mathcal{L}}_q(i\Omega_n) + U\overline{\mathcal{L}}_q(i\Omega_n)\chi_q(i\Omega_n),\tag{58}
$$

respectively. Solving Eq.  $(57)$  $(57)$  $(57)$ , we obtain

$$
\overline{\mathcal{L}}_q(i\Omega_n) = [\hat{I} + U\hat{\mathcal{D}}_q(i\Omega_n)]^{-1} \mathcal{L}_q^0(i\Omega_n).
$$
 (59)

<span id="page-5-7"></span>Here,  $\hat{I}$  is the  $4 \times 4$  unit matrix. After the analytic continuation  $i\Omega_n \rightarrow \omega + i\delta$  (we take the limit  $\delta \rightarrow +0$  after the calculation), we obtain the density response function as

$$
\chi_q(\omega) = \frac{[\overline{\mathcal{L}}_q(\omega)]_1 + [\overline{\mathcal{L}}_q(\omega)]_4}{1 - U[(\overline{\mathcal{L}}_q(\omega))_1 + (\overline{\mathcal{L}}_q(\omega))_4]}.
$$
(60)

<span id="page-5-5"></span>The excitation spectrum of the AB mode is obtained from the pole of Eq.  $(60)$  $(60)$  $(60)$ .

If we only take into account the ladder diagrams,  $\chi_q(\omega)$ reduces to

$$
\chi_q^{\mathcal{L}}(\omega) = [\overline{\mathcal{L}}_q(\omega)]_1 + [\overline{\mathcal{L}}_q(\omega)]_4. \tag{61}
$$

<span id="page-5-6"></span>Equation ([61](#page-5-6)) shows that the pole of  $\chi^L(\omega)$  coincides with that of  $\overline{L}_q(\omega)$  which is obtained from the condition |*I*  $+ U \hat{\mathcal{D}}_{q}(\omega) | = 0$  in Eq. ([59](#page-5-7)).

We calculate the single bubble diagram  $\hat{\mathcal{D}}_q(\omega)$  in Eq. ([56](#page-5-3)). Substituting Eq.  $(32)$  $(32)$  $(32)$  into Eq.  $(56)$  $(56)$  $(56)$ , we obtain

<span id="page-5-8"></span>
$$
D_{q}^{\mu\nu\rho\lambda}(\omega) = \frac{1}{M} \sum_{k} \left[ \overline{A}_{k+q}^{\mu\nu} \overline{A}_{k}^{\rho\lambda} \frac{f(E_{k}^{+}) - f(E_{k+q}^{+})}{\omega + E_{k}^{+} - E_{k+q}^{+} + i \delta} - \overline{B}_{k-q}^{\mu\nu} \overline{B}_{k}^{\rho\lambda} \frac{f(E_{k}^{+}) - f(E_{k-q}^{+})}{\omega - E_{k}^{+} + E_{k-q}^{+} + i \delta} + \overline{A}_{k-q}^{\mu\nu} \overline{B}_{k}^{\rho\lambda} \frac{1 - f(E_{k}^{+}) - f(E_{-k+q}^{+})}{\omega - E_{k}^{+} - E_{-k+q}^{+} + i \delta} - \overline{B}_{k+q}^{\mu\nu} \overline{A}_{k}^{\rho\lambda} \frac{1 - f(E_{k}^{+}) - f(E_{-k-q}^{+})}{\omega + E_{k}^{+} + E_{-k-q}^{+} + i \delta} \right], \qquad (62)
$$

where  $\hat{\vec{A}} = \hat{\sigma}_3 \hat{A}$  and  $\hat{\vec{B}} = \hat{\sigma}_3 \hat{B}$ . In deriving Eq. ([62](#page-5-8)), we used  $E_{-k}^-$  =− $E_k^+$ . We assume  $E_k^+$ >0 because we are interested in the stability of Fermi gases before the pair breaking sets in, i.e.,  $|v| < |v_{\text{pb}}|$ . At *T*=0, the first and second terms in Eq. ([62](#page-5-8)) vanish from this condition. The density response functions in Eqs.  $(60)$  $(60)$  $(60)$  and  $(61)$  $(61)$  $(61)$  are calculated by using Eqs.  $(59)$  $(59)$  $(59)$  and  $(62)$  $(62)$  $(62)$ .

From Eq. ([62](#page-5-8)), the spectrum of the particle-hole excitation is given by

$$
\omega_q^{\text{ph}}(k) = E_k^+ + E_{-k+q}^+.
$$
 (63)

For fixed *q*,  $\omega_q^{\text{ph}}(k)$  makes a continuum for different *k* as shown in Fig. [3.](#page-6-0) The upper and lower boundaries of the particle-hole continuum are given by  $\min_k[\omega_q^{\text{ph}}(k)]$  and  $\max_k[\omega_q^{\text{ph}}(k)]$ , respectively.

## **III. RESULTS**

<span id="page-5-0"></span>In this section, by calculating the excitation spectra of the AB mode and the single-particle excitation, we discuss the stability of superfluid flow and determine the critical velocities of superfluid Fermi gases in 1D, 2D, and 3D optical lattices. For this purpose, we calculate the dynamic structure factor  $S_q(\omega) = -\text{Im}[\chi_q(\omega)]/\pi$ , which describes the response of the system to density perturbations with momentum *q* and frequency  $\omega$ . The dynamic structure factor can be directly

<span id="page-6-0"></span>

FIG. 3. Dynamic structure factor  $S_q(\omega)$  in 1D optical lattices. The upper gray region and the lower curve correspond to the particle-hole continuum and the AB mode spectrum, respectively. The dashed line is the lower boundary of the single-particle continuum. We set  $n=0.5$  (quarter filling),  $U=-2.0J$ , and  $v=0$ . The superfluid gap and chemical potential are  $|\Delta_n|$ =0.409*J*and  $\mu$ = 0.624*J*. Here, in numerical calculation,  $\delta$  is set to be small but finite  $(1.0 \times 10^{-4})$  so that the peak of the AB mode spectrum has a small finite width.

measured in experiments by using Bragg spectroscopy  $[37,38]$  $[37,38]$  $[37,38]$  $[37,38]$ .

Since the GRPA used in this paper is based on a meanfield approximation, it is more reliable for higher dimensions. Nevertheless, calculations of the excitation spectra in the simplest situation of 1D can be useful for understanding the essence of the physics governing the critical velocity of superfluid fermions in a lattice. Hence, we first discuss the excitation spectra and the stability of superfluid flow in 1D lattices. We note that mean-field theories have been widely used to qualitatively describe excitations of trapped atomic gases even in 1D because the finite size  $(\sim 100d)$  specific to cold atom systems excludes long-wavelength phase fluctuations that destroy the long-range superfluid order  $[39-41]$  $[39-41]$  $[39-41]$ .

In Fig. [3,](#page-6-0) we show the dynamic structure factor  $S_q(\omega)$  in 1D optical lattices to illustrate the basic properties of the excitation spectra. One clearly sees that the AB mode spectrum lies below the particle-hole continuum. In addition, the AB mode spectrum has a characteristic structure with local minima at short wavelengths which is similar to the roton spectrum in superfluid  ${}^{4}$ He [[42](#page-13-25)]. Then, it is expected that as the superfluid velocity increases, the energy of one of the rotonlike minima decreases and it reaches zero before the lower boundary of the particle-hole continuum does. This indicates that the instability may be induced by the rotonlike excitations of the AB mode rather than by the single-particle excitations because the single-particle excitations start to have negative energies when the particle-hole continuum reaches zero energy  $\lceil 25 \rceil$  $\lceil 25 \rceil$  $\lceil 25 \rceil$ . Indeed, we will show that this is the case in all of 1D, 2D, and 3D optical lattices in the remainder of this section.

We note that this AB mode-induced instability does not occur in superfluid Fermi gases in uniform 3D systems. It was found that the instability of superfluid Fermi gases in uniform 3D systems is induced by pair breaking  $[24,25]$  $[24,25]$  $[24,25]$  $[24,25]$ which is associated with the appearance of single-particle excitations with negative energies. In this case, the AB mode spectrum merges into the particle-hole continuum in contrast with the behavior in Fig. [3](#page-6-0) where the AB mode spectrum is separated from the particle-hole continuum. As a result, the particle-hole continuum reaches zero energy before the phonon part of the AB mode spectrum starts to have negative energy as the superfluid velocity increases. This leads to the instability induced by single-particle excitations, i.e., by pair breaking. The rotonlike structure of the AB mode spectrum also appears in a uniform 1D system  $\lceil 43 \rceil$  $\lceil 43 \rceil$  $\lceil 43 \rceil$  and a 2D lattice system  $[44, 45]$  $[44, 45]$  $[44, 45]$  $[44, 45]$  $[44, 45]$ .

In the following, we first calculate the pair-breaking velocity  $v_{\text{pb}}$ . We next discuss the behavior of the AB mode spectrum in 1D, 2D, and 3D optical lattices by calculating the dynamic structure factor  $S_q(\omega)$ . To discuss the excitation spectra, we only show the lower boundary of the particlehole continuum and the peak of the AB mode spectrum in  $S_q(\omega)$  because other details are not necessary for determining the critical velocities.

#### **A. Pair-breaking velocity**

In this section, we calculate the pair-breaking velocity  $v_{\text{pb}}$ , which can be analytically obtained from the condition that the lower boundary of the particle-hole continuum reaches zero energy, i.e.,  $\min_k[\omega_q^{\text{ph}}(k)] = 0$  [[25](#page-13-9)]. In 1D case, the pair-breaking velocity is given by  $[36]$  $[36]$  $[36]$ 

$$
v_{\rm pb} = \frac{1}{md} \sin^{-1} \left( \frac{|\Delta_v|}{\sqrt{\mu (4J - \mu)}} \right). \tag{64}
$$

<span id="page-6-3"></span><span id="page-6-1"></span>In 2D case, when the supercurrent is flowing in the  $(\pi,\pi)$ and  $(\pi, 0)$  directions, the pair-breaking velocities are calculated as

$$
|\mathbf{v}_{\rm pb}| = \frac{\sqrt{2}}{md} \sin^{-1} \left( \frac{|\Delta_v|}{\sqrt{\mu (8J - \mu)}} \right)
$$
 (65)

and

$$
|\mathbf{v}_{\rm pb}| = \begin{cases} \frac{1}{md} \sin^{-1} \left( \frac{|\Delta_v|}{\sqrt{\mu (4J - \mu)}} \right), & (\mu < 2J) \\ \frac{1}{md} \sin^{-1} \left( \frac{|\Delta_v|}{2J} \right), & (\mu \ge 2J), \end{cases}
$$
(66)

respectively. We address the stability of superfluid states in these two cases in Sec. [III C.](#page-8-0) In 3D case, when the supercurrent is flowing in the  $(\pi,\pi,\pi)$  direction, the pair-breaking velocities are calculated as

$$
|\mathbf{v}_{\rm pb}| = \frac{\sqrt{3}}{md} \sin^{-1} \left( \frac{|\Delta_v|}{\sqrt{\mu (12J - \mu)}} \right). \tag{67}
$$

<span id="page-6-2"></span>Since the order parameter and chemical potential depend on the superfluid velocity  $v$ , we must determine  $v_{\text{pb}}$  by solving Eq.  $(64)$  $(64)$  $(64)$ – $(67)$  $(67)$  $(67)$  self-consistently with Eqs.  $(39)$  $(39)$  $(39)$  and  $(40)$  $(40)$  $(40)$ . In the BCS limit ( $|U| \ll J$ ),  $v_{\text{pb}}$  approaches zero because  $|\Delta_v|$ becomes small. On the other hand, if |U| is so large that  $|\Delta_{\bf{v}}|$ is larger than the denominator in sin<sup>-1</sup> in Eqs.  $(64)$  $(64)$  $(64)$ – $(67)$  $(67)$  $(67)$ ,  $v_{\text{pb}}$ is not definable.

<span id="page-7-0"></span>

FIG. 4. Excitation spectra in 1D optical lattices for (a) currentfree  $(v=0)$  and (b) current-carrying  $(v=0.21/md)$  cases. Solid line and dashed line represent the spectrum of the AB mode which corresponds to the  $\delta$ -function peak of  $S_q(\omega)$  and the lower boundary of the particle-hole continuum, respectively. We set  $n=0.5$  (quarter filling) and  $U = -2.0J$ . The superfluid gap and chemical potential are (a)  $|\Delta_v|$ =0.409*J* and  $\mu$ =0.624*J*, and (b)  $|\Delta_v|$ =0.420*J* and  $\mu$  $= 0.655J$ .

## **B. Stability in 1D optical lattices**

In this section, we study the stability of superfluid Fermi gases in 1D optical lattices. In Fig. [4,](#page-7-0) we show the dynamic structure factor  $S_q(\omega)$  in 1D optical lattices. It is clearly seen in Fig. [4](#page-7-0) that the AB mode has a gapless and linear dispersion in the long-wavelength limit ( $|qd| \le 1$ ), which is consistent with the fact that the AB mode is a Nambu-Goldstone mode  $\left[35\right]$  $\left[35\right]$  $\left[35\right]$ . We obtain the analytic form of the phononlike dispersion relation in the long-wavelength limit  $(|qd| \le 1)$  as

<span id="page-7-1"></span>
$$
\omega_q = d(2J - \mu)\tan(mvd)q
$$
  
+ 
$$
\frac{1}{\alpha}\sqrt{(1+\alpha)\{v_F^2\alpha^2 - |\Delta_v|^2[\alpha^2 + (1+\alpha)\tan^2(mvd)]\}}|q|,
$$
  
(68)

where  $v_F = 2Jd \sin(k_F d) \cos(mv d)$  $k_{\rm F}$ = $|\cos^{-1}\{(2J$  $(-\mu)[2J \cos(m\nu d)]^{-1}$ |/*d*,  $\alpha = N_0 U$ , and  $N_0 = d(\pi v_F)^{-1}$ . The details of the derivation of Eq.  $(68)$  $(68)$  $(68)$  are summarized in the Appendix. When  $v=0$ , Eq. ([68](#page-7-1)) in the BC<u>S limit</u>  $(|\Delta_v| \ll \mu)$ reduces to the well-known form  $\omega_a = v_F \sqrt{1 + N_0 U} |q|$ , which was first obtained by Anderson for a uniform 3D system [[29](#page-13-13)[,31](#page-13-15)] (the dispersion has the coefficient  $1/\sqrt{3}$  in 3D case). We note that Eq.  $(68)$  $(68)$  $(68)$  is different from the dispersion relation obtained in Ref.  $[27]$  $[27]$  $[27]$  by using the hydrodynamic and tightbinding approximations.

As we pointed out earlier, in Figs.  $4(a)$  $4(a)$  and  $4(b)$  $4(b)$  $4(b)$ , it is clearly seen that the excitation spectrum of the AB mode lies below the particle-hole continuum and the AB mode spectrum has rotonlike minima at  $|q| \approx 2k_F^0$   $[k_F^0 = n\pi/(2d)]$  is the Fermi wave number in a noninteracting 1D system]. As *v* increases, the whole spectrum leans toward the left side and the energy of the rotonlike minimum with  $q<0$  decreases as shown in Fig.  $4(b)$  $4(b)$ . As a result, at a certain velocity  $v_c$ , the rotonlike minimum reaches zero energy, but this occurs before the lower boundary of the particle-hole continuum does. According to the Landau criterion  $[23]$  $[23]$  $[23]$ , this indicates that the spontaneous emission of rotonlike excitations of the AB

<span id="page-7-2"></span>

FIG. 5. Critical velocity  $v_c$  (solid line) and pair-breaking velocity  $v_{\text{pb}}$  (dashed line) as functions of  $U/J$  in 1D lattices. We set *n*  $= 0.5$  (quarter filling). From Eq. ([64](#page-6-1)),  $v_{\text{pb}}$  approaches  $\pi/2md$  when  $|U|/J \rightarrow 4$  (filled circle). When  $|U|/J > 4$ , the pair-breaking velocity is not definable.

mode is induced when  $v \geq v_c$  before pair-breaking occurs at  $v_{\rm ph}$ . Thus, the critical velocity is given by  $v_c$  at which the superfluid flow is destabilized due to the spontaneous emission of rotonlike excitations of the AB mode. We note that the phonon part of the AB mode spectrum becomes negative at a certain velocity larger than  $v_{\text{pb}}$ . This means that phonon excitations of the AB mode are irrelevant to the critical velocities.

This instability driven by negative-energy rotonlike excitations of the AB mode corresponds to the energetic instability called Landau instability. Another type of instability of superfluid called dynamical instability was also proposed for Fermi gases in optical lattices  $[27,46]$  $[27,46]$  $[27,46]$  $[27,46]$ . Dynamical instability is associated with the appearance of complex energy excitations which was first observed in Bose condensates in optical lattices  $\lceil 47 \rceil$  $\lceil 47 \rceil$  $\lceil 47 \rceil$ . One can distinguish the dynamical instability from the Landau instability by identifying the quasimomentum  $q$  of the excitations causing the instability. If the instability is caused by the excitations at the boundary of the first Brillouin zone, e.g.,  $q_x = \pm \pi/d$  or  $q_y = \pm \pi/d$  in 2D, it is the dynamical instability because these excitations inevitably couple with their antiphonon branches  $(48,49)$  $(48,49)$  $(48,49)$  $(48,49)$ . Indeed, we will see that in 2D lattices, the dynamical instability due to the AB mode at  $q = (\pm \pi/d, \pm \pi/d)$  can occur near the half filling or in the BEC region.

In Fig. [5,](#page-7-2) we show  $v_c$  and  $v_{pb}$  as functions of  $U/J$  in the BCS-BEC crossover region. One clearly sees that  $v_c$  is smaller than  $v_{\text{pb}}$ . We confirmed that  $v_c$  is smaller than  $v_{\text{pb}}$  in the entire BCS region  $(-1 \le U/J \le 0)$ . Thus, the instability is always induced by the rotonlike excitations in these regions. The difference between  $v_c$  and  $v_{pb}$  increases with increasing  $|U|/J$  when one approaches the BEC regime. In addition, both  $v_c$  and  $v_{pb}$  grow monotonically with increasing the interaction  $|U|/\tilde{J}$ .

When  $|U|/J \ge 1$  (BEC region), the size of the Cooper pairs becomes smaller than the lattice spacing and each Cooper pair forms a tightly bound molecular boson. In this region, the Hubbard model of Eq.  $(1)$  $(1)$  $(1)$  is mapped onto a hardcore Bose-Hubbard model with nearest-neighbor repulsive

<span id="page-8-1"></span>

FIG. 6. Excitation spectra in 1D optical lattices for (a) currentfree  $(v=0)$  and (b) current-carrying  $(v=0.12/md)$  cases. Solid line, dash-dotted line, and dashed line represent the spectrum of the AB mode obtained from Eq.  $(60)$  $(60)$  $(60)$ , Eq.  $(61)$  $(61)$  $(61)$ , and the lower boundary of the single-particle excitation continuum, respectively. We set *n* = 0.9 and *U*=−2.0*J*. The superfluid gap and chemical potential are calculated as (a)  $|\Delta_v|$ =0.345*J* and  $\mu$ =1.69*J*, and (b)  $|\Delta_v|$ =0.350*J* and  $\mu = 1.70J$ .

interactions (or equivalently the spin- $\frac{1}{2}$  *XXZ* model) [ $29,44,46,50$  $29,44,46,50$  $29,44,46,50$  $29,44,46,50$ ]. Since it is well known that any kinds of meanfield theory completely fail to describe hardcore bosons in 1D, our GRPA is also invalid in the BEC limit in 1D. Hence, we postpone the discussion of the BEC limit to the next section, where we will show results in 2D.

To discuss the origin of the rotonlike minima of the AB spectrum, we show the dynamic structure factor when *n*  $= 0.9$  in Fig. [6.](#page-8-1) Compared to the AB mode spectrum when  $n = 0.5$  in Fig. [4,](#page-7-0) the rotonlike minima have lower energies than those in Fig. [4.](#page-7-0) It turns out that as one approaches half filling  $(n=1)$ , the energy of the rotonlike minima becomes smaller. As is well known, the fluctuation due to the formation of charge-density-wave (CDW) order is enhanced near half filling in lattice fermion systems  $[51]$  $[51]$  $[51]$ . Thus, the CDW fluctuation leads to the rotonlike minima in the AB mode spectrum. At half filling, the rotonlike minima reach zero energy even in the current-free case  $(v=0)$  and the superfluid ground state becomes unstable due to the formation of CDW order. We show the critical velocity  $v_c$  when  $n=0.9$  as a function of  $U/J$  in Fig. [7.](#page-8-2) The critical velocity when  $n=0.9$ in Fig. [7](#page-8-2) is smaller than that when  $n=0.5$  $n=0.5$  in Fig. 5 due to the strong CDW fluctuation. We notice that  $v_c$  becomes almost constant below a certain value of interaction  $(U/J \le -2)$ , which reflects the fact that the energy difference between the lower boundary of the particle-hole continuum and the rotonlike minimum becomes large as |U| increases. This indicates that the CDW fluctuation is enhanced below this value of interaction.

To support the above consideration on the origin of the rotonlike minimum, we compare the AB mode spectra calculated by Eqs.  $(60)$  and  $(61)$  in Fig. [6.](#page-8-1) As discussed in Sec.  $II C$ , Eq. ([61](#page-5-6)) includes only the contribution from the ladder diagrams, while Eq.  $(60)$  $(60)$  $(60)$  includes the contributions both from ladder and RPA-type bubble diagrams. In Fig. [6,](#page-8-1) one clearly sees that the AB mode spectrum calculated by Eq.  $(60)$  $(60)$  $(60)$  lies below the one calculated by Eq.  $(61)$  $(61)$  $(61)$  which actually lies close to the lower boundary of the particle-hole con-

<span id="page-8-2"></span>

FIG. 7. Critical velocity  $v_c$  (solid line) and pair-breaking velocity  $v_{\rm pb}$  (dashed line) as functions of  $U/J$  in 1D lattices near half filling. We set  $n = 0.9$ .

tinuum. Since the RPA-type bubble diagrams include the effect of the CDW fluctuation  $[51]$  $[51]$  $[51]$ , this behavior of the AB mode spectrum is consistent with the above consideration that the CDW fluctuation leads to the rotonlike structure of the AB mode spectrum.

#### **C. Stability in 2D optical lattices**

<span id="page-8-0"></span>In this section, we discuss the stability of Fermi gases in 2D optical lattices. Here, we restrict ourselves to two characteristic situations where the superfluid flows along the  $(\pi, \pi)$  or  $(\pi, 0)$  directions (see Fig. [8](#page-8-3)) in order to see the effects of CDW fluctuations on the stability of superfluid Fermi gases.

First, we discuss the case when the superfluid flows along the  $(\pi, \pi)$  direction. We assume the superfluid velocity *v*  $= (v, v) / \sqrt{2}$ , where  $v = |v|$ . We calculate the dynamic structure factor  $S_q(\omega)$  only when  $q_x = q_y$  because superfluid flow is expected to be most unstable for excitations with momenta in the opposite direction to the flow.

<span id="page-8-3"></span>

FIG. 8. Schematic picture of 2D optical lattices. Arrows indicate the superfluid velocity in (a)  $(\pi, \pi)$  and (b)  $(\pi, 0)$  directions. Each circle represents a lattice site.

<span id="page-9-0"></span>

FIG. 9. Excitation spectra in 2D optical lattices when the superfluid flows along the  $(\pi, \pi)$  direction. Solid and dashed lines represent the spectrum of the AB mode and the lower boundary of the particle-hole continuum, respectively. The superfluid velocity is  $[(a)$ and (c)]  $v=0$ , (b)  $v=0.467/md$ , and (d)  $v=0.6502/md$ . We set *n*  $= 0.5$ ,  $q_x = q_y$ , [(a) and (b)]  $U = -4.5J$ , and [(c) and (d)]  $U = -12.0J$ . The superfluid gap and chemical potential are calculated as (a)  $|\Delta_{v}| = 1.33J$  and  $\mu = 2.23J$ , (b)  $|\Delta_{v}| = 1.39J$  and  $\mu = 2.29J$ , (c)  $|\Delta_{v}|$  $= 4.92J$  and  $\mu = 0.687J$ , and (d)  $|\Delta_v| = 4.97J$  and  $\mu = 0.746J$ .

In Figs.  $9(a)$  $9(a)$  and  $9(b)$ , we show the excitation spectra when  $n=0.5$  and  $q_x=q_y$ . When  $v=0$  [see Fig. [9](#page-9-0)(a)], the rotonlike structure is slightly seen in the AB mode spectrum. As *v* increases, the rotonlike structure becomes remarkable and one of the rotonlike minima goes down. At a certain velocity  $v_c$  smaller than the pair-breaking velocity  $v_{\text{pb}}$ , the energy of the rotonlike minimum reaches zero [see Fig.  $9(b)$  $9(b)$ ]. Thus, as 1D case, the critical velocity is given by  $v_c$  at which the instability due to spontaneous emission of rotonlike excitations of the AB mode sets in.

In Figs.  $9(c)$  $9(c)$  and  $9(d)$ , we show the excitation spectra in the BEC region  $(U=-12J)$  [[52](#page-13-35)]. There we see that the rotonlike minima of the AB mode are present also in this region. The critical velocity in the BEC region is also determined by the rotonlike excitations [see Fig.  $9(d)$  $9(d)$ ]. Since the rotonlike minima are shifted to  $q_x = q_y = \pm \pi/d$ , the instability caused by the rotonlike excitations is the dynamical instability  $[53]$  $[53]$  $[53]$ . The shift of the rotonlike minima to the edge of the Brillouin zone can be understood as follows. As mentioned before, in the BEC region the Hubbard model can be reduced to a hardcore Bose-Hubbard model with nearest-neighbor repulsive interactions. The nearest-neighbor repulsion enhances density wave fluctuations with the wave vector *k*  $=(\pi/d, \pi/d)$  [[54](#page-13-37)], leading to the rotonlike minimum at *q*  $=(\pi/d, \pi/d)$ . Thus, the shift of the roton minima means that as one approaches the BEC region, the origin of the roton minima changes from the nesting effect of the Fermi surface

<span id="page-9-1"></span>

FIG. 10. Critical velocity  $v_c$  (solid line) and pair-breaking velocity  $v_{\rm pb}$  (dashed line) as functions of  $U/J$  in 2D lattices when the superfluid flows along the  $(\pi, \pi)$  direction. We set (a)  $n=0.5$  and (b)  $n=0.8$ . From Eq. ([65](#page-6-3)),  $v_{\text{pb}}$  approaches  $\sqrt{2}\pi/2md$  when  $|U|/J$  $\rightarrow$  8 (filled circle). When  $|U|/J$  > 8, the pair-breaking velocity is not definable.

to the nearest-neighbor interactions between molecular bosons. Notice that in the limit of the low filling  $(n\rightarrow 0)$ , the roton minima of the AB mode do not survive any longer  $[44]$  $[44]$  $[44]$ and the critical velocity is determined by the longwavelength part (phonon branch) of the AB mode.

In Fig. [10,](#page-9-1) we show the critical velocity  $v_c$  and the pairbreaking velocity  $v_{\text{pb}}$  as functions of  $U/J$ .  $v_{\text{c}}$  and  $v_{\text{pb}}$  show qualitatively the same behavior as in 1D case. Namely,  $v_c$  is smaller than  $v_{\text{pb}}$  and they increase monotonically with increasing the interaction strength  $|U|/J$ .

Near half filling, the rotonlike structure when  $v=0$  becomes more remarkable due to strong CDW fluctuation [see Fig.  $11(a)$  $11(a)$ ]. Since it is seen in Fig.  $11(b)$  that the instability is driven by the rotonlike excitations with  $q = (\pm \pi/d, \pm \pi/d)$ , it is the dynamical instability. As a result of the strong CDW fluctuation, the critical velocity near half filling is smaller than the one at quarter filling, as shown in Fig.  $10(b)$  $10(b)$ . As in 1D case, the CDW fluctuation is strongly enhanced below a certain value of interaction  $(U/J \le -3)$  in Fig. [10](#page-9-1)(b). As a

<span id="page-9-2"></span>

FIG. 11. Excitation spectra in 2D optical lattices when the superfluid flows along the  $(\pi, \pi)$  direction. Solid and dashed lines represent the spectrum of the AB mode and the lower boundary of the particle-hole continuum, respectively. The superfluid velocities are (a)  $v=0$  and (b)  $v=0.260/md$ . The filling is  $n=0.8$ . We set *U*  $=-4.5J$  and  $q_x = q_y$ . The superfluid gap and chemical potential are calculated as (a)  $|\Delta_v|$ =1.60*J* and  $\mu$ =3.32*J*, and (b)  $|\Delta_v|$ =1.62*J* and  $\mu = 3.33J$ .

<span id="page-10-0"></span>

FIG. 12. Excitation spectra in 2D optical lattices when the superfluid flows along the  $(\pi, 0)$  direction for (a) current-free  $(v=0)$ and (b) current-carrying  $(v=0.585/md)$  cases. Solid and dashed lines represent the spectrum of the AB mode and the lower boundary of the particle-hole continuum, respectively. We set  $n=0.5$ , *U*  $=-4.5J$ , and  $q_y=0$ . The superfluid gap and chemical potential are calculated as (a)  $|\Delta_v|$ =1.33*J* and  $\mu$ =2.23*J*, and (b)  $|\Delta_v|$ =1.41*J* and  $\mu = 2.31J$ .

result,  $v_c$  is almost constant below this value of interaction.

Next, we discuss the case when the supercurrent flows along the  $(\pi, 0)$  direction in 2D optical lattices (see Fig. [8](#page-8-3)). We assume the superfluid velocity as  $v = (v, 0)$ . The excitation spectra when  $n=0.5$  is shown in Fig. [12.](#page-10-0) Here, we assume  $q_y=0$  from the same reason for the  $(\pi,\pi)$  case. The behavior of the energy spectra is qualitatively the same as the previous cases, i.e., the AB mode spectrum has the rotonlike structure and lies below the particle-hole continuum. As the superfluid velocity increases, the AB mode spectrum is pushed down and the energy of the rotonlike minimum decreases. The instability sets in at  $v_c$  when the energy of the rotonlike minimum becomes zero.

In Fig. [13,](#page-10-1) we show the critical velocity  $v_c$  and pairbreaking velocity  $v_{\text{pb}}$  as functions of  $U/J$ . They also show qualitatively the same behavior as the previous cases. Com-paring Figs. [10](#page-9-1) and [13,](#page-10-1) we find that  $v_c$  in the  $(\pi, \pi)$  case is smaller than the one in the  $(\pi, 0)$  case. This is because the nesting effect of the Fermi surface occurs in the  $(\pi, \pi)$  di-

<span id="page-10-1"></span>

FIG. 13. Critical velocity  $v_c$  (solid line) and pair-breaking velocity  $v_{\rm pb}$  (dashed line) as functions of  $U/J$  in 2D lattices when the superfluid flows along the  $(\pi, 0)$  direction. We set  $n = 0.5$ .

<span id="page-10-2"></span>

FIG. 14. Excitation spectra in 3D optical lattices when the superfluid flows along the  $(\pi,\pi,\pi)$  direction for [(a) and (c)]  $v=0$ , (b)  $v = 0.5969/md$ , and (d)  $v = 0.7894/md$ . Solid and dashed lines represent the spectrum of the AB mode and the lower boundary of the particle-hole continuum, respectively. We set  $n=0.5$ ,  $q_x=q_y$  $=q_z$ , [(a) and (b)]  $U=-6.0J$ , and [(c) and (d)]  $U=-14.0J$ . The superfluid gap and chemical potential are calculated as (a)  $|\Delta_v|$  $= 1.92J$  and  $\mu = 3.78J$ , (b)  $|\Delta_v| = 1.98J$  and  $\mu = 3.85J$ , (c)  $|\Delta_v|$  $= 5.70J$  and  $\mu = 2.11J$ , and (d)  $|\Delta_{\nu}| = 5.77J$  and  $\mu = 2.18J$ .

rection so that it is enhanced by the supercurrent in parallel to this direction.

We note that near the half filling, the AB mode spectrum with  $q_y \neq 0$  [i.e., not in the  $(\pi, 0)$  direction] may reach zero before that in the  $(\pi, 0)$  direction does, even when the superfluid flows in the  $(\pi, 0)$  direction. This is also due to the strong nesting effect in the  $(\pi, \pi)$  direction. However, even in this case, our main conclusion remains unchanged. Namely, the instability is induced by the AB mode excitations. This effect was also pointed out in the BEC regime  $[46]$  $[46]$  $[46]$ .

#### **D. Stability in 3D optical lattices**

Let us finally discuss the stability of Fermi gases in 3D optical lattices. We restrict ourselves to the situation where the superfluid flows along the  $(\pi,\pi,\pi)$  direction. We show the excitation spectra in Fig. [14](#page-10-2) for  $q_x = q_y = q_z$  and the critical velocity and the pair-breaking velocity as functions of *U*/*J* in Fig. [15.](#page-11-1) It is clearly seen that the AB mode lies well below the single-particle continuum and that in the entire region of the attractive interaction, the roton part of the AB mode reaches zero before the single-particle continuum does. This leads to the conclusion that the critical velocity of superfluid Fermi gases in a deep lattice is determined by the roton part of the AB mode except in the low filling limit, regardless of the dimensionality of the system.

<span id="page-11-1"></span>

FIG. 15. Critical velocity  $v_c$  (solid line) and pair-breaking velocity  $v_{\rm pb}$  (dashed line) as functions of  $U/J$  in 3D lattices when the superfluid flows along the  $(\pi,\pi,\pi)$  direction. We set *n*=0.5. From Eq. ([67](#page-6-2)),  $v_{\text{pb}}$  approaches  $\sqrt{3}\pi/2md$  when  $|U|/J \rightarrow 12$  (filled circle). When  $|U|/J$  > 12, the pair-breaking velocity is not definable.

## **IV. CONCLUSION**

<span id="page-11-0"></span>In conclusion, we have studied the stability of superfluid Fermi gases in 1D, 2D, and 3D optical lattices at *T*= 0. By applying the GRPA Green's function formalism developed by Côté and Griffin  $[28]$  $[28]$  $[28]$  to the attractive Hubbard model, we calculated the excitation spectra of the AB mode as well as the single-particle excitation in the presence of superfluid flow. We found that the AB mode spectrum has the characteristic rotonlike structure being separated from the particlehole continuum due to the strong CDW fluctuation. The energy of the rotonlike minimum decreases as the superfluid velocity increases and it reaches zero at the critical velocity before pair-breaking occurs. This indicates that the instability of superfluid flow in 1D, 2D, and 3D optical lattices is induced by the spontaneous emission of the rotonlike excitations of the AB mode. We calculated the critical velocity  $v_c$ as functions of *U*/*J* and confirmed that it is smaller than the pair-breaking velocity  $v_{\text{pb}}$  in the BCS and BCS-BEC crossover region. We also found that the CDW instability is strongly enhanced near half filling which leads to the suppression of the critical velocity when the attractive interaction is large.

Finally, we remark that our results are valid for superfluid Fermi gases in *deep* optical lattices because we employed the tight-binding Hubbard model. From this reason, our results for 1D optical lattices cannot be directly compared to the experiment with shallow optical lattices in Ref.  $[24]$  $[24]$  $[24]$ . However, superfluid Fermi gases have been already achieved in deep optical lattices in Ref.  $[55]$  $[55]$  $[55]$ . Our theoretical predictions in this paper may be verified if a superfluid Fermi gas is prepared in a moving deep optical lattice.

### **ACKNOWLEDGMENTS**

We wish to thank K. Kamide, K. Osawa, M. Takahashi, T. Kimura, T. Nikuni, and S. Kurihara for useful discussions and comments. We acknowledge Y. Ohashi and M. Tsubota for valuable comments. I.D. and D.Y. are supported by a Grant-in-Aid from JSPS.

#### **APPENDIX: AB PHONON SPECTRUM IN 1D**

Here, we give a detailed derivation of the AB mode spectrum in  $1D$  in the long-wavelength limit in Eq.  $(68)$  $(68)$  $(68)$ . The matrix  $\hat{\mathcal{D}}_q(i\Omega_n \to \omega)$  in Eq. ([55](#page-5-9)) is given as

$$
\hat{\mathcal{D}}_q(\omega) = \begin{pmatrix} a+b & d+e & f-g & c \\ -d-e & -h-i & -c & -f-g \\ -f+g & -c & i-h & -d+e \\ c & f+g & d-e & a-b \end{pmatrix}, \quad \text{(A1)}
$$

<span id="page-11-2"></span>where

$$
a = \frac{1}{M} \sum_{k} \frac{\mathcal{E}\mathcal{E}' - \overline{\xi}' \overline{\xi}}{2\mathcal{E}\mathcal{E}'} \frac{\mathcal{E} + \mathcal{E}'}{(\omega - \eta' + \eta)^2 - (\mathcal{E} + \mathcal{E}')^2}, \quad (A2)
$$

<span id="page-11-4"></span>
$$
b = \frac{1}{M} \sum_{k} \frac{\mathcal{E}\overline{\xi'} - \mathcal{E}'\overline{\xi}}{2\mathcal{E}\mathcal{E}'} \frac{\omega - \eta' + \eta}{(\omega - \eta' + \eta)^2 - (\mathcal{E} + \mathcal{E}')^2}, \quad \text{(A3)}
$$

$$
c = \frac{1}{M} \sum_{k} \frac{\Delta_v^2}{2\mathcal{E}\mathcal{E}'} \frac{\mathcal{E} + \mathcal{E}'}{(\omega - \eta' + \eta)^2 - (\mathcal{E} + \mathcal{E}')^2},\tag{A4}
$$

$$
d = \frac{1}{M} \sum_{k} \frac{\Delta_v}{2\mathcal{E}} \frac{\omega - \eta' + \eta}{(\omega - \eta' + \eta)^2 - (\mathcal{E} + \mathcal{E}')^2},
$$
 (A5)

$$
e = \frac{1}{M} \sum_{k} \frac{\Delta_{v} \overline{\xi}'}{2\mathcal{E} \mathcal{E}'} \frac{\mathcal{E} + \mathcal{E}'}{(\omega - \eta' + \eta)^2 - (\mathcal{E} + \mathcal{E}')^2},
$$
 (A6)

$$
f = \frac{1}{M} \sum_{k} \frac{\Delta_v}{2\mathcal{E}'} \frac{\omega - \eta' + \eta}{(\omega - \eta' + \eta)^2 - (\mathcal{E} + \mathcal{E}')^2},\tag{A7}
$$

$$
g = \frac{1}{M} \sum_{k} \frac{\Delta_{v} \overline{\xi}}{2\mathcal{E} \mathcal{E}'} \frac{\mathcal{E} + \mathcal{E}'}{(\omega - \eta' + \eta)^{2} - (\mathcal{E} + \mathcal{E}')^{2}},
$$
 (A8)

$$
h = \frac{1}{M} \sum_{k} \frac{\mathcal{E} \mathcal{E}' + \overline{\xi}' \overline{\xi}}{2\mathcal{E} \mathcal{E}'} \frac{\mathcal{E} + \mathcal{E}'}{(\omega - \eta' + \eta)^2 - (\mathcal{E} + \mathcal{E}')^2}, \quad \text{(A9)}
$$

<span id="page-11-3"></span>
$$
i = \frac{1}{M} \sum_{k} \frac{\mathcal{E}\overline{\xi}^{\prime} + \mathcal{E}^{\prime}\overline{\xi}}{2\mathcal{E}\mathcal{E}^{\prime}} \frac{\omega - \eta^{\prime} + \eta}{(\omega - \eta^{\prime} + \eta)^2 - (\mathcal{E} + \mathcal{E}^{\prime})^2}.
$$
 (A10)

Here, we used the notations  $\mathcal{E} = \mathcal{E}_k$ ,  $\mathcal{E}' = \mathcal{E}_{k+q}$ ,  $\overline{\xi} = \overline{\xi}_k$ ,  $\overline{\xi}'$  $\vec{a} = \vec{\xi}_{k+q}, \ \eta = \eta_k$ , and  $\eta' = \eta_{k+q}$ . For simplicity, we took  $\Delta_v$  to be real. We calculate the matrix elements  $(A2)$  $(A2)$  $(A2)$ – $(A10)$  $(A10)$  $(A10)$  in the limit of  $qd \leq 1$  and  $\omega / J \leq 1$ . Expanding Eq. ([A2](#page-11-2)) to the second order in  $q$  and  $\omega$ ,  $a$  is obtained as

INSTABILITY OF SUPERFLUID FERMI GASES INDUCED ...

$$
a \approx \frac{\Delta_v^2}{4M} \sum_k \left[ -\frac{1}{\mathcal{E}^3} + qd \frac{3J\bar{\xi} \sin(kd)\cos(mvd)}{\mathcal{E}^5} - \omega^2 \frac{1}{4\mathcal{E}^5} \right]
$$
  
+  $q d\omega \frac{J \cos(kd)\sin(mvd)}{\mathcal{E}^5}$   
+  $(qd)^2 \left( \frac{J^2 \Delta_v^2 \sin^2(kd)\cos^2(mvd)}{\mathcal{E}^7} + \frac{J^2 \cos^2(kd)\sin^2(mvd)}{\mathcal{E}^5} + \frac{3J\bar{\xi} \cos(kd)\cos(mvd)}{2\mathcal{E}^5} \right) \right].$  (A11)

<span id="page-12-6"></span>By carrying out the integration over  $k$ , we obtain

$$
a \approx -\frac{N_0}{2} - \frac{N_0 \omega^2}{12 \Delta_v^2} + q d\omega \frac{N_0 J \cos(k_{\rm F} d) \sin(mv d)}{3 \Delta_v^2}
$$

$$
- (qd)^2 N_0 \frac{(3 \Delta_v^2 - 16 J \mu + 4 \mu^2) \cos(2mv d)}{48 \Delta_v^2 \cos^2(mv d)}
$$

$$
- (qd)^2 N_0 \frac{\Delta_v^2 - 16 J^2 \sin^4(mv d)}{48 \Delta_v^2 \cos^2(mv d)}.
$$
(A12)

Here, we have used the standard approximation,  $\frac{1}{2\pi}\int_{-\pi/4}^{\pi/4} d\vec{k} F(\vec{k}) \rightarrow N_0 \int_{-\infty}^{\infty} d\vec{\xi} F(\vec{\xi}),$  where  $F(\vec{k})$  is an arbitrary function. By calculating Eqs.  $(A3)$  $(A3)$  $(A3)$ – $(A10)$  $(A10)$  $(A10)$  in the same way as Eq.  $(A2)$  $(A2)$  $(A2)$ , we obtain

$$
b \simeq 0,\tag{A13}
$$

$$
c \simeq -\frac{N_0}{2} - \frac{N_0 \omega^2}{12\Delta_v^2} + q d\omega \frac{N_0 J \cos(k_{\rm F} d) \sin(mv d)}{3\Delta_v^2} + (qd)^2 \frac{2J(4\mu - J) + 8J^2 \sin^4(mv d)}{24\Delta_v^2} - (qd)^2 \frac{(16J^2 + \Delta_v^2) \sin^2(mv d)}{24\Delta_v^2},
$$
 (A14)

$$
d \simeq f \simeq -\frac{N_0}{4\Delta_v} [\omega - 2q dJ \cos(k_{\rm F} d)\sin(mv d)], \quad \text{(A15)}
$$

$$
e \simeq -qd\omega \frac{\tan(m\upsilon d)}{12\Delta_{\upsilon}} + (qd)^{2} \frac{(2J - \mu)[3 + 2\tan^{2}(m\upsilon)]}{24\Delta_{\upsilon}},
$$
\n(A16)

$$
g \simeq -q d\omega \frac{\tan(m\omega d)}{12\Delta_v} - (qd)^2 \frac{(2J - \mu)[3 - 2\tan^2(m\nu)]}{24\Delta_v},
$$
\n(A17)

$$
h \approx \frac{1}{U} + \frac{N_0}{2} - \frac{N_0 \omega^2}{6\Delta_v^2} + q d\omega \frac{2N_0 J \cos(k_{\rm F} d) \sin(mv d)}{3\Delta_v^2} + (qd)^2 N_0 \frac{13\Delta_v^2 \sin^2(mv d) - 16J^2 \sin^4(mv d)}{24\Delta_v^2 \cos^2(mv d)} - (qd)^2 N_0 \frac{3\Delta_v^2 + 2J^2 - 8J\mu + 2\mu^2}{12\Delta_v^2 \cos^2(mv d)},
$$
\n(A18)

$$
i \simeq \frac{qd \tan(mvd)}{2} \left(N_0 + \frac{1}{U}\right). \tag{A19}
$$

<span id="page-12-7"></span>Note that when we calculate  $1/M\Sigma_k 1/\mathcal{E}$ , we have used the gap equation  $[Eq. (40)]$  $[Eq. (40)]$  $[Eq. (40)]$  to eliminate the divergence. Assuming that the AB mode has a linear dispersion relation in the long-wavelength limit  $(qd \leq 1)$  and calculating the pole of Eq.  $(60)$  $(60)$  $(60)$  by using Eqs.  $(A12)$  $(A12)$  $(A12)$ – $(A19)$  $(A19)$  $(A19)$ , we obtain Eq.  $(68)$  $(68)$  $(68)$ .

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