# Upper and Lower Bounds to Atomic and Molecular Properties. I. Be-Sequence Oscillator Strengths (Dipole-Length Formulation) for the 1s<sup>2</sup>2s<sup>21</sup>S - 1s<sup>2</sup>2s 2p<sup>1</sup>P Transition

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Wave functions, which include interelectron coordinates  $r_{ij}$  explicitly, are calculated for the  $1s^{2}2s^{2} {}^{1}S$ and  $1s^{2}2s 2p {}^{1}P$  states of Be I, C III, and O v. These wave functions are used to calculate oscillator strengths, including upper and lower bounds, for the lowest  ${}^{1}S {}^{-1}P$  transition. Interpolation techniques are used to make a graphical study of the oscillator-strength behavior along the isoelectronic sequence. Comparisons are made with previous experimental and theoretical results. The results of this study are oscillator strengths for the  $1s^{2}2s^{2} {}^{1}S \rightarrow 1s^{2}2s 2p {}^{1}P$  Be isoelectronic sequence with rigorous upper and lower bounds of (7-10)% and probable accuracy  $\leq 2\%$ .

## I. INTRODUCTION

Spectroscopic data for atoms and ions fall into a category of fundamental scientific material for which justification of measurement or calculation is scarcely necessary. The measurement of the wavelengths of spectral lines and the subsequent deduction of energy levels is usually not too difficult.<sup>1,2</sup> The study of the strengths (intensities) of spectral lines is, however, a much more formidable problem.<sup>3</sup> On the experimental side, absolute oscillator strengths, or f values, are typically determined with uncertainties of 10-20% or more.<sup>4</sup> On the theoretical side, Pekeris and co-workers appear to have established some helium f values to 1% or better,<sup>5</sup> and other transitions in the two-electron isoelectronic sequence have been determined to perhaps 5%.6,7 But results reliable to better than 10-20% have not been generally accessible in larger systems.<sup>8, 9</sup> In addition, the question of the reliability of theoretical predictions has become increasingly acute,<sup>10</sup> and there is a need for procedures which lead to upper and lower bounds for the properties of interest, so that rigorous error limits are attached to the error estimates. In view of the above, we have calculated f values for the  ${}^{1}S-{}^{1}P$ transition (by the dipole-length formulation) for some members of the beryllium isoelectronic sequence, and used the interpolation techniques (based on Rayleigh-Schrödinger perturbation theory) of Wiese and Weiss<sup>11</sup> to make a graphical study of the *f*-value behavior along the isoelectronic sequence. The calculations utilize the best wave functions (on any energy criterion) currently available.<sup>12, 13</sup> In addition, the procedure of Weinhold<sup>14, 15</sup> has been used to calculate upper and lower bounds to the computed f values. The results of this study are oscillator strengths for the  $1s^22s^2 S^2 S^2 P$  Be isoelectronic sequence

with rigorous upper and lower bounds of 7-10% and probable accuracy  $\leq 2\%$ .

#### II. THEORY

#### A. Wave Functions

La Paglia and Sinanouglu<sup>16</sup> have shown that for the resonance transition of the beryllium isoelectronic sequence  $(1s^22s^2 \, {}^{1}S-1s^22s2p \, {}^{1}P)$ , Hartree-Fock wave functions are unreliable and wave functions, including correlation effects, of considerable accuracy must be employed, not only for the ground state, but also for excited states. The procedure adapted here for incorporating correlations into the wave function is the method of Sims and Hagstrom,<sup>13</sup> in which a variational trial function is written as a linear combination of known many-electron functions,

$$\Psi = \sum_{k} C_{k} \Phi_{k} , \qquad (1)$$

and the configurations  $\Phi_k$  are themselves antisymmetrized, projected products of orbitals and interelectronic coordinates. By minimizing the energy with respect to the coefficients  $C_k$ , one is led to the usual matrix eigenvalue equation for the energy and the coefficients  $C_k$ . The eigenvalues are always upper bounds to the energy of the corresponding excited (or ground) state.<sup>17</sup> The eigenvalue equation is solved by Jacobi's method<sup>18</sup> after the overlap matrix has been triangulated and the Hamiltonian matrix transformed by the method of Michels, Van Dine, and Elliott.<sup>19</sup>

In Eq. (1), the  $\Phi_k$  are

$$\Phi_{k} = O(L^{2})O_{as}\left(\chi_{k} r_{ij}^{\nu_{k}} \prod_{s=1}^{4} \phi_{ks}(\tilde{\mathbf{r}}_{s})\right), \qquad (2)$$

where  $O(L^2)$  is an idempotent orbital-angularmomentum projection operator<sup>20</sup> and  $O_{as}$  is the

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projection operator that guarantees the antisymmetry of the wave function

$$O_{as} = (4!)^{-1} \sum_{P} (-1)^{p_{P}}, \qquad (3)$$

where the summation runs over all the 4! permutations P, with p being the parity of the corresponding permutation P. For this work, the spin function was taken as  $\chi_k = \chi = \frac{1}{2}(\alpha_1\beta_2 - \beta_1\alpha_2) \times (\alpha_3\beta_4 - \beta_3\alpha_4)$ , which has  $S = S_x = 0$ . In Eq. (2),  $\phi_{ks}(\mathbf{\bar{r}}_s)$  refers to the sth basis orbital in the *k*th configuration. The orbital basis consists of Slater-type orbitals (STO's) of the general form

$$\phi(\vec{\mathbf{r}}) = \frac{(2\zeta)^{n+1/2}}{[(2n)!]^{1/2}} r^{n-1} e^{-\zeta r} Y_{l,m} , \qquad (4)$$

where the set  $\{Y_{i,m}\}$  consists of normalized spherical harmonics in the Dirac phase convention.<sup>21</sup> In Eq. (2), the restriction of only one  $r_{ij}$  coordinate per term, first proposed by James and Coolidge,<sup>22</sup> has been retained. The basis set consists of s and p Slater-type orbitals (STO's) and powers of interelectronic coordinates:  $r_{ij}^{v}(v=0, 1, 2)$ . Previous calculations<sup>13</sup> have shown that this is an excellent basis set for beryllium.

#### **B.** Oscillator Strengths

The basic theory of electric dipole radiation in atomic systems is given by Condon and Shortley,<sup>21</sup> and we follow their terminology closely. Assuming Russell-Saunders (LS) coupling and the nonrelativistic many-electron Hamiltonian (Hartree atomic units)<sup>23</sup>

$$H = \sum_{i=1}^{N} \left( \frac{1}{2} \Delta_{i} - \frac{Z}{r_{i}} \right) + \sum_{i < j} \frac{1}{r_{ij}}, \qquad (5)$$

the energy states of an atom are characterized by the set of quantum members  $\gamma LSM_LM_S$ . Here  $\gamma$ denotes the electronic configuration, L the total angular momentum, S the total spin, and  $M_L$  and  $M_S$  are the projections of L and S, respectively, on the axis of quantization.<sup>24</sup> For an electric dipole transition connecting the terms (multiplets)  $\gamma LS$  and  $\gamma'L'S'$ , Condon and Shortley introduced the quantity  $(\gamma LS; \gamma'L'S')$ , called the absolute multiplet strength. It is defined in terms of the electric-dipole-moment operator

$$\vec{\mathbf{P}} = \sum_{i=1}^{N} \vec{\mathbf{r}}_{i}$$
by<sup>25</sup>

$$\begin{split} & \$(\gamma LS:\gamma'L'S') = \sum_{\mathbf{M}_{L},\mathbf{M}_{S}} \sum_{\mathbf{M}_{L'},\mathbf{M}_{S}'} |\langle \gamma LSM_{L}M_{S}|\vec{\mathbf{P}}| \\ & \times \gamma'L'S'M_{L'}M_{S'}\rangle|^{2} \end{split}$$
(6)

The f value or multiplet oscillator strength is related to \$ by

$$f(\gamma LS + \gamma'L'S') = \frac{2}{3} \frac{E(\gamma'L'S') - E(\gamma LS)}{(2L+1)(2S+1)} \, \&(\gamma LS; \gamma'L'S') \,.$$

$$(7)$$

The familiar electric dipole selection rules,  $\Delta S = 0$ ,  $\Delta M_s = 0$ ,  $\Delta L = 0$ ,  $\pm 1$ , and  $\Delta M_L = 0$ ,  $\pm 1$ , hold for the Hamiltonian we use. Moreover, given  $\gamma LS$  and  $\gamma'$ , L'(=L-1, L, L+1) S'(=S), as a consequence of the Wigner-Eckart theorem,<sup>26</sup> every nonvanishing matrix element of the operator  $\vec{P}$ can be expressed in terms of a single parameter  $\beta(\gamma LS; \gamma' L'S')$  multiplied by expressions involving only L,  $M_L$ , and  $M_L$ . The parameter is independent of the quantum numbers  $M_L$ ,  $M_{L'}$ ,  $M_s$ , and  $M_{s'}$ , the summation indices occurring in Eq. (6). Consequently, the expressions for the multiplet oscillator strength may be written as products of the parameter  $\beta(\gamma LS; \gamma'L'S')$  and a single algebraic function of L and S resulting from the summation. Finally, notice that only one nonvanishing matrix element need be evaluated to determine  $\beta(\gamma LS; \gamma' L'S')$ , so that instead of computing all the matrix elements occurring in the summations occurring in Eq. (6), a single matrix element suffices to find the multiplet oscillator strength for given  $\gamma LS$ ,  $\gamma' L'S'$ .

For the  $1s^22s^{2}{}^{1}S-1s^22s2p {}^{1}P$  transition, the  ${}^{1}S$  term consists of one state with  $S = M_S = L = M_L = 0$ . The  ${}^{1}P$  term consists of three states corresponding to  $S = M_S = 0$ , L' = 1, and  $M_{L'} = -1$ , 0, 1. We choose to express ( $\gamma LS \rightarrow \gamma' L'S'$ ) in terms of the matrix element corresponding to the  ${}^{1}P$  state with  $S = M_S = 0$ , L' = 1, and  $M_{L'} = 0$ . Then Eq. (6) becomes (unnormalized wave functions), from Goldberg,<sup>27</sup>

$$8(\gamma^{1}S;\gamma'^{1}P) = 3 \left| \frac{\int dx_{1} \cdots dx_{n} \psi_{\gamma}^{*}(^{1}S) \vec{P}\psi_{\gamma'}(^{1}P,M_{L'}=0)}{\left[ \langle \psi_{\gamma}(^{1}S) | \psi_{\gamma}(^{1}S) \rangle \langle \psi_{\gamma'}(^{1}P,M_{L}=0) | \psi_{\gamma'}(^{1}P,M_{L}=0) \rangle \right]^{1/2}} \right|^{2},$$

$$(8)$$

and we need only construct  $\psi_{\gamma}({}^{1}S)$  and  $\psi_{\gamma'}({}^{1}P, M_{L}, = 0)$  and evaluate Eq. (8) to compute 8.

It is generally known<sup>28</sup> that with  $\psi_{\gamma} LSM_LM_S$  as exact eigenfunctions of H [Eq. (5)], the commutation relations of H with  $\vec{P}$  (dipole-length operator) lead to two additional relations for  $(\gamma LS; \gamma' L'S')$ ,<sup>8</sup> which arise from the use of the dipole-velocity and dipole-acceleration operators given by

$$\vec{\nabla} = \sum_{i=1}^{N} \vec{\nabla}_i \tag{9}$$

and

$$\vec{\mathbf{A}} = \sum_{i=1}^{N} \frac{Ze^{2}\vec{\mathbf{r}}_{i}}{|\vec{\mathbf{r}}_{i}|^{3}}, \qquad (10)$$

respectively. These three formulas are equivalent when one uses exact eigenfunctions of H [Eq. (5)] or of an independent-electron central-field-model Hamiltonian. However, they disagree whenever the eigenfunctions of a central-field model are improved by including some correlation, even by only the Hartree-Fock (HF) procedure.<sup>29</sup> Starace<sup>30</sup> has shown that the length formula is the physically correct one for all methods, such as Hartree-Fock, which involve the diagonalization of an approximate, but nonlocal, Hamiltonian. For the procedure we are using, which involves the use of a local potential, the use of the three different operators will give different results. It has been generally argued<sup>8</sup> that, since the acceleration operator contains terms behaving as  $1/r^2$ , the dipoleacceleration form is most sensitive to the accuracy of the wave function in the immediate neighborhood of the nucleus, so that oscillator strengths computed with variationally stable approximate wave functions (emphasizing more distant regions) are not reliably given in terms of this operator. Since the dipole-length and dipole-velocity forms should agree as the wave function approaches the exact eigenfunction, some authors have taken the closeness of agreement of & computed with dipolelength and dipole-velocity operators as a measure of the accuracy of their computation of S, but there is now abundant evidence that this procedure is not very reliable.<sup>31</sup> In the present work, we compute upper and lower bounds to the operator and use these theoretical bounds as a measure of the accuracy of the f value; we compute f values in terms of the dipole-length operator as given by Eq. (8).<sup>32</sup>

To evaluate Eq. (8) we define

$$\begin{aligned} x_{0}(i) &= z_{i} = r_{i} \cos \theta_{i} = ar_{i} Y_{1,0}(i), \\ x_{+1}(i) &= -2^{-1/2} (x_{i} + i y_{i}) = ar_{i} Y_{1,1}(i), \\ x_{-1}(i) &= 2^{-1/2} (x_{i} - i y_{i}) = ar_{i} Y_{1,-1}(i), \end{aligned}$$
(11)

where  $a = (\frac{1}{3} 4\pi)^{1/2}$  and the  $Y_{l,m}$  are the same as in Eq. (4). Let

$$X_{0} = \sum_{i=1}^{N} x_{0}(i), \quad X_{-1} = \sum_{i} x_{-1}(i), \quad X_{1} = \sum_{i} x_{1}(i).$$
(12)

Then<sup>33</sup>

$$|\vec{\mathbf{P}}|^2 = |X_0|^2 + |X_{-1}|^2 + |X_1|^2.$$
(13)

Since  $x_j(i) = \pm a[2, 1, j]_0^i$  in terms of unnormalized STO's defined by

$$[n, l, m]_{S}^{l} = r_{i}^{n-1} e^{-\zeta r_{i}} Y_{l,m}(i), \qquad (14)$$

we obtain, for 8,

$$\delta(\gamma^{1}S; \gamma^{1}P) = (\frac{4}{3}\pi)3 \sum_{j=-1}^{1} \left| \int dx_{1} \cdots dx_{n} \psi_{\gamma}^{*}(^{1}S) \right| \\ \times \sum_{i=1}^{N} \left[ 2, 1, j \right]_{0}^{i} \psi_{\gamma} \cdot (^{1}P, M_{L'} = 0) \right|^{2}$$
(15)

$$= 4\pi (|\tilde{W}_{-1}|^2 + |\tilde{W}_0|^2 + |\tilde{W}_{+1}|^2)$$
  
=  $4\pi |\tilde{W}_0|^2$ , (16)

owing to symmetry. (*j* must equal 0 or the integral vanishes.) In writing Eq. (15), we have assumed that  $\psi_{\gamma}$  and  $\psi_{\gamma'}$  are normalized ( $\langle \psi_{\gamma} | \psi_{\gamma} \rangle = 1$ ), and  $|\tilde{W}_j|$  is defined by

$$|\tilde{W}_{j}| = \left| \int dx_{1} \cdots dx_{n} \psi_{\gamma}^{*}({}^{1}S) \sum_{i=1}^{N} [2, 1, j]_{0}^{i} \psi_{\gamma'}({}^{1}P, M_{L'} = 0) \right|$$
(17)

To evaluate Eq. (16), we use the fact that the wave functions employed to evaluate Eq. (16) are given by Eq. (1) and are linear combinations of antisymmetrized, projected products of orbitals and interelectronic coordinates. Thus

$$|\tilde{W}_{j}| = \sum_{K,L} C_{K} C_{L} |W_{j}|_{KL}, \qquad (18)$$

where

$$|\tilde{W}_{j}|_{KL} = \langle \Phi_{K} | \sum_{i=1}^{N} [2, 1, j]_{0}^{i} | \Phi_{L} \rangle.$$
(19)

In Eq. (19),  $\Phi_L$  is a linear combination of orbitals and interelectronic coordinates, so that Eq. (19) can be evaluated by expanding the product of  $[2, 1, j]_0^i$  and the appropriate STO of  $\Phi_L$  in terms of STO's by the formulas given in a previous paper.<sup>34</sup> Then the integrals required to evaluate Eq. (16) are similar in type to those required to calculate overlap matrix elements using these wave functions and can be evaluated as described later.

## C. Upper and Lower Bounds to Oscillator Strengths

Weinhold<sup>14</sup> has described a procedure for calculating rigorous upper and lower bounds to dipole strengths. If the exact transition moment is

$$W_{ab} = \langle \Psi_a | W_l | \Psi_b \rangle, \qquad (20)$$

where  $\Psi_a$  and  $\Psi_b$  are the exact wave functions for states *a* and *b*, respectively, and  $W_i$  is some par-

$$\tilde{W}_{ab} = \langle \psi_a | W_l | \psi_b \rangle, \qquad (21)$$

where  $\psi_a$  and  $\psi_b$  are the approximate wave functions for states *a* and *b*, respectively, is

$$W_{ab} \ge S_{a\pm} S_{b\pm} W_{ab} \pm \epsilon_{a+} S_{b\pm} \Delta_{ab} \pm \epsilon_{b+}$$

$$\times [\langle \Psi_a | W_1^2 | \Psi_a \rangle_+ - (S_{a-} \tilde{W}_{ab} - \epsilon_{a+} \Delta_{ab})^2]^{1/2}.$$
(22)

Here  $S_a \equiv \langle \psi_a | \Psi_a \rangle$ ,  $S_b \equiv \langle \psi_b | \Psi_b \rangle$ , i.e., the (positive) overlap integrals of the approximate wave functions  $\psi_a$ ,  $\psi_b$  with the true wave functions  $\Psi_a$ ,  $\Psi_b$ , and  $\epsilon_{a+}$ ,  $\epsilon_{b+}$  are simply

$$\epsilon_{a+} \equiv (1 - S_{a-}^2)^{1/2}, \quad \epsilon_{b+} \equiv (1 - S_{b-}^2)^{1/2},$$

and  $\Delta_{ab}$  is defined by

$$\Delta_{ab} \equiv \left( \left\langle \psi_b \left| W_1^2 \right| \psi_b \right\rangle - \tilde{W}_{ab}^2 \right)^{1/2} \,. \tag{23}$$

In Weinhold's scheme,  $\tilde{W}_{ab}$  and  $\langle \psi_b | W_l^2 | \psi_b \rangle$  are computed directly. Without loss of rigor,  $S_{a+}, S_{b+}$ may be taken to be 1, and for  $\Psi_a, \Psi_b$ , the lowest states of their respective symmetries, we can use the "Eckart criterion" to determine  $S_{a-}, S_{b-}$ ; e.g.,

$$S_{a-}^{2} = (E_{a}^{*} - \langle \psi_{a} | H | \psi_{a} \rangle)(E_{a}^{*} - E_{a}), \qquad (24)$$

where  $E_a$  and  $E_a^*$  are the ground- and excitedstate energy (with the same symmetry as the ground state) of the system with Hamiltonian H. The remaining unknown of relation (22) is the upper bound  $\langle \Psi_a | W_i^2 | \Psi_a \rangle_+$ . Weinhold presents several formulas for its treatment; we take

$$\langle \Psi_a | W_l^2 | \Psi_a \rangle \approx \langle \psi_a | W_l^2 | \psi \rangle, \qquad (25)$$

where  $\psi$  is our highly accurate "S" wave function. With  $|W_j|$ 's defined by Eq. (17), we have that

$$\$(\gamma^{1}S; \gamma^{1}P) = 4\pi |\tilde{W}_{0}|^{2} = 3 \tilde{W}_{ab}^{2} , \qquad (26)$$

so that

$$\left|\tilde{W}_{0}\right| = (3/4\pi)^{1/2} \tilde{W}_{ab} . \tag{27}$$

Relation (22) holds for  $|W_0|$  in place of  $W_{ab}$  if we compute  $\langle \Psi_a | W_i^2 | \Psi_a \rangle$  by relation (25) and multiply by  $3/4\pi$  to account for the difference between  $|\tilde{W}_0|$  and  $\tilde{W}_{ab}$  given by Eq. (27). Also, the only remaining unknown in the calculation of upper and lower bounds to  $|\tilde{W}_0|$  is  $\langle \psi_b | W_i^2 | \psi_b \rangle$ . With the aid of<sup>34</sup>

$$[2, 1, j]_{0}^{i} [2, 1, j]_{0}^{i} = \sum_{P_{a}=0}^{1} r_{i}^{2} [(2l_{1}+1)/4\pi]^{1/2} C^{l_{1}}(1, j; 1, j) Y_{l_{1},0} \quad (l_{1}=2-2P_{a})$$
$$= \sum_{P_{a}=0}^{1} [(2l_{1}+1)/4\pi]^{1/2} C^{l_{1}}(1, j; 1, j) [3, l_{1}, 0]_{0}^{i}, \qquad (28)$$

we obtain

$$\langle \psi_b | (3/4\pi) X_0 X_0 | \psi_b \rangle = \sum_{i=1}^{4} \sum_{\substack{P_a = 0 \\ i \neq k}}^{1} [(2l_1 + 1)/4\pi]^{1/2} C^{l_1}(1, 0; 1, 0) \langle \psi_b | [3, l_1, 0]_0^i | \psi_b \rangle$$

$$+ \sum_{\substack{i,k \\ i \neq k}} \langle \psi_b | [2, 1, 0]_0^i * [2, 1, 0]_0^k | \psi_b \rangle.$$

$$(29)$$

Here  $C^{L}(l^{1}, m^{1}; l, m)$  is the Condon-Shortley coefficient<sup>21</sup> defined by

The matrix elements in Eq. (29) are handled similarly to the way  $|W_j|$  was treated earlier.

# D. Trends of Oscillator Strengths Along the Be Isoelectronic Sequence

The use of interpolation techniques based on Rayleigh-Schrödinger perturbation theory to study systematic trends of f values along an isoelectronic sequence is due to Wiese and Weiss.<sup>11</sup> The fundamental idea is to scale all distances by the nuclear charge Z and then apply conventional perturbation theory, treating the interelectronic repulsion as the perturbation:

$$H = \sum_{i} \left( \frac{-\Delta_{i}(\rho)}{2} - \frac{1}{\rho_{i}} \right) + \frac{1}{Z} \sum_{i < j} \frac{1}{\rho_{ij}}$$
$$= H_{0} + \frac{1}{Z} V, \qquad (30)$$

where  $\rho = Zr$ .

Perturbation theory then gives for the wave function and energy,

$$\Psi = \Psi_0 + (1/Z)\Psi_1 + \cdots,$$
 (31)

$$E = E_0 Z^2 + E_1 Z + E_2 + \cdots .$$
 (32)

Here  $\Psi_0$  is just a configuration of hydrogenic orbitals,  $E_0$  is the sum over all electrons of hydrogenic energies, and the first-order energy is given by

$$E_1 = \langle \Psi_0 | V | \Psi_0 \rangle . \tag{33}$$

If there is a degeneracy in zeroth order (as in the transition we are studying) then the firstorder energy is found by diagonalizing the firstorder energy matrix which removes the degeneracy and redefines the zeroth-order wave function as a linear combination of the degenerate hydrogenic configurations. This charge-expansion approach of Wiese and Weiss has also been used to calculate f values.<sup>35</sup> Dalgarno and co-workers,<sup>36, 37</sup> working within this framework, have derived formulas for the multiplet strength as a function of Z by solving for the orbitals through first order and including the zeroth-order configuration mixing. The results of this method should be accurate for sufficiently large Z, and are presented in Fig. 1 (Z-expansion results) in extending the curves to  $Z^{-1}=0$ . The multiplet strength, from Eqs. (6) and (31), has the Z dependence

$$\mathbf{\$} = \mathbf{S}_0 / Z^2 + \mathbf{S}_1 / Z^3 + \cdots, \qquad (34)$$

where  $S_0$  is the multiplet strength obtained from hydrogenic functions and may well involve the first-order mixing. For the oscillator strength we are interested in, there is no change in principal quantum number so that the zeroth-order energy difference vanishes and the Z dependence of the f value is

$$f = f_1 / Z + f_2 / Z^2 + \cdots, (35)$$

so that the oscillator strength tends to zero for large Z.<sup>38</sup>

These considerations have several consequences. In the first place, notice that  $Z^{-1}$  is the natural parameter against which to plot f values in a graphical display of the isoelectronic sequence. The entire sequence is then compressed in the region between  $Z^{-1} = 0$  and the value of the neutral atom. Second, one can calculate relatively easily the way in which these curves start at the zero end, in this case, from the slope of the curve at the origin.<sup>39</sup> These considerations are implicit in the curve presented as Fig. 1.

#### **III. METHOD OF CALCULATION**

## A. Outline

The procedure we follow for calculating f values is to first compute <sup>1</sup>S and <sup>1</sup>P wave functions variationally, and then use these wave functions to compute oscillator strengths as described earlier. The method of calculation of <sup>1</sup>S wave functions for the Be isoelectronic sequence has been given previously,<sup>13</sup> and for the Be <sup>1</sup>S ground state we use in this work the first 105 terms of the 107-term wave function reported in that work. The <sup>1</sup>S wave functions for CIII and OV, the other two members of the Be isoelectronic sequence computed in this work, are 89-configuration truncations of the 107-term wave function with different orbital exponents; these will be presented later in this paper.

The <sup>1</sup>P wave functions used in this work differ from the <sup>1</sup>S wave functions in that they are derived from configurations containing three S-type STO's and one P-type STO in order to obtain an over-all P symmetry. The  $O_{as}$  projection treatment of spin, final Hamiltonian, and overlap matrix elements is similar to the treatment in Ref. 13, but the  $O(L^2)$  projection is different.

## B. $O(L^2)$ Projection

For the  $O(L^2)$  projection we use the formula of Löwdin<sup>40</sup>:

$$O_{L,M_{L}}^{L^{2}} = (2L+1) \frac{(L+M_{L})!}{(L-M_{L})!} \times \sum_{\nu=0}^{L_{\max}-L} (-1)^{\nu} \frac{L_{-}^{L-M_{L}+\nu}L_{+}^{L-M_{L}+\nu}}{\nu!(2L+\nu+1)!}, \quad (36)$$

where  $L_{+}$  and  $L_{-}$  are the (step-up and step-down) operators  $L_{+} = L_{x} + iL_{y}$ ,  $L_{-} = L_{x} - iL_{y}$ . For  $L = M_{L}$ =0, Eq. (36) reduces to

$$O_{1,0}^{L^{2}} = 3 \sum_{\nu=0}^{L_{\max}-L} (-1)^{\nu} \frac{L^{\nu+1}_{-} L^{\nu+1}_{*}}{\nu! (\nu+3)!}.$$
 (37)

TABLE I. Orbital exponents chosen for the wave functions of this work.

	Stat Be	ie I	с	ш	0 <b>v</b>		
	<sup>1</sup> S	<sup>1</sup> P	<sup>1</sup> S	<sup>1</sup> P	<sup>1</sup> S	<sup>1</sup> P	
ns <sup>a</sup>	3.6847 <sup>b</sup>	3,6903	5,6575	5.6734	7,6295	7.6564	
ns"	0.9562 <sup>b</sup>	1.0851	1.8316	1.7685	2,6903	2.6152	
np	6.2 <sup>c</sup>	•••	9.7 °	•••	13.2 <sup>c</sup>	•••	
np"	1.1 °	0.4897 (2p") 0.5600 (3p")	2.5 <sup>c</sup>	1.946 <sup>c</sup>	3.9°	2,998	

<sup>a</sup> Unless otherwise noted, values are taken from Ref. 44.

<sup>b</sup> Reference 43.

<sup>c</sup> Reference 41.

Config.		Coefficient of 53-term wave function						
No. n	Configuration	Be I C III O V						
1	$1s^22s''2p''$	4.764777	4,456 084	4.030 196				
2	$1s^2 2s'' 2p'' r_{10}$	$9.858159 \times 10^{-2}$	$7.125036 \times 10^{-2}$	$6.164211 \times 10^{-2}$				
3	$1s^2 2s'' 2h'' r_{12}^2$	$1.301.724 \times 10^{-1}$	9.009462 $\times 10^{-2}$	$6.036.654 \times 10^{-2}$				
3	$13 23 2p 7_{12}$	$7 136207 \times 10^{-1}$	$2.062.765 \times 10^{-1}$	$1.520.908 \times 10^{-1}$				
4	182828 2p	$4.777.091 \times 10^{-1}$	$-1.006.418 \times 10^{-1}$	$-5.247703 \times 10^{-2}$				
5	1s3s2s" 2p"	4.777 081 × 10	$-1.090410\times 10$	$-5.247755 \times 10^{-2}$				
6	2s <sup>2</sup> 2s" 2p"	-3.610 930 × 10 -	-7.524 009×10 -	-6.204 963 × 10 -				
?	1s <sup>4</sup> 3s" 3p"	-2.236168	$-1.254405 \times 10^{-1}$	-6.857813×10 "				
8	1s <sup>2</sup> 3s"2p"	-1.547299  imes 10	-7.897644	-5.826 508				
9	$1s^2 2s'' 2p'' r_{34}^2$	1.338971	$4.474569 \times 10^{-1}$	$3.809192 \times 10^{-1}$				
10	$1s^2 3s'' 2p'' r_{34}^2$	$-1.340863 \times 10$	-3.445850	-2.645 772				
11	$1s^23s''2t''r_{10}$	$1.356403 \times 10^{-1}$	$-2.627807 \times 10^{-2}$	$-2.208844 \times 10^{-2}$				
19	$1e^{2}3e''2b''r^{2}$	$-1.698274 \times 10^{-1}$	$-9.194345 \times 10^{-2}$	$-6.797355 \times 10^{-2}$				
10	$13 33 2p 1_{12}$		$-3.13434340 \times 10^{-1}$	$-0.502281 \times 10^{-2}$				
13	182828 3p	= 7.270 078 × 10	$-1.001710\times10$	-9.000 201 × 10				
14	2s <sup>2</sup> 2s" 3p"	3.076321×10	2.335176×10 °	2.420 712 × 10 =				
15	1s2s3s" 2p"	-1.254746	$-4.610471 \times 10^{-1}$	-3.643 951×10 <sup>-1</sup>				
16	1s <sup>2</sup> s"2p"r <sub>34</sub>	-3.009266	$-7.613410 \times 10^{-1}$	$-5.540680 \times 10^{-1}$				
17	1s <sup>2</sup> 3s"2 <b>p"r</b> <sub>34</sub>	$2.606178\! imes\!10$	7.453266	5.174 738				
18	1s4s2 <b>s"</b> 2 <b>p"</b>	$2.734655 \times 10^{-1}$	$2.896810  imes 10^{-1}$	$2.242952 \times 10^{-1}$				
19	1s <sup>2</sup> 1s"2 <i>p</i> "	$-6.549024 \times 10^{-1}$	$-9.442578\! imes\!10^{-1}$	$-9.468604 \times 10^{-1}$				
20	$1s^24s''2p''$	$2.725431 \times 10$	9.304 938	6.063 070				
01	1-29-49 - 44-	1 400 752 × 10-1	0.002.22010-3	6 218 000 × 10 <sup>-3</sup>				
21	$1s^{-}3s^{-}3p^{-}r_{12}$	-1.400753×10	9.903 329 × 10 °	6.318 000 × 10				
22	1s2s3s"3p"	$5.625126 \times 10^{-1}$	$1.336102 \times 10^{-6}$	$1.178410 \times 10^{-4}$				
23	1s3s3s"3p"	$2.336654 \times 10^{-1}$	$-3.549046 \times 10^{-2}$	$-2.944542 \times 10^{-2}$				
<b>24</b>	1s3s3s"2p"	$-2.688191 \times 10^{-1}$	$8.032144 \times 10^{-2}$	$4.233014 \times 10^{-2}$				
25	2 <i>s</i> ²3 <b>s″</b> 3 <i>p″</i>	$-2.609941 \times 10^{-1}$	$-2.122250 \times 10^{-2}$	$-2.245426 \times 10^{-2}$				
26	2s <sup>2</sup> 3s" 2p"	$2.401137 \times 10^{-1}$	2.114013×10 <sup>-2</sup>	$2.389507 \times 10^{-2}$				
27	$1s^2 4s'' 2p'' r_{34}$	$-4.511435 \times 10$	$-1.467124 \times 10$	-9.644 455				
28	$1s^2 4s'' 2 p'' r_{3}^2$	$2.469396 \times 10$	7.941 418	5.594 460				
29	1s2s2s"2p"r	5,713 677 $\times$ 10 <sup>-1</sup>	$5.249247 \times 10^{-1}$	$3.994919 \times 10^{-1}$				
30	$1s2s2s''2p''r_{12}^2$	$-3.613879 \times 10^{-2}$	$-6.697089 \times 10^{-3}$	$7.209994 \times 10^{-3}$				
			_1					
31	$1s3s2s''2p''r_{12}$	$-8.156346 \times 10^{-1}$	$-8.029208 \times 10^{-1}$	$-6.131649 \times 10^{-1}$				
32	$1s3s2s''2p''r_{12}^2$	$4.765640 \times 10^{-1}$	$4.269118 \times 10^{-1}$	$3.125237 \times 10^{-1}$				
33	$2s^2 2s'' 2p'' r_{12}^2$	$-2.554983 \times 10^{-2}$	$-2.725063 \times 10^{-2}$	$-2.186619 \times 10^{-2}$				
34	$1s4s2s''2p''r_{12}^2$	$-1.588584 \times 10^{-1}$	$-1.378832 \times 10^{-2}$	$-1.003420 \times 10^{-1}$				
35	1s <sup>2</sup> 5s" 2p"	$-2.103935 \times 10$	-5.652 667	-3.353 813				
36	$1s^25s''2p''r_{34}$	3.626329  imes 10	1,196395×10	7.502 105				
37	$1s^2 5s'' 2p'' r_2^3$	$-1.994613 \times 10$	-7,432 937	-4.957 139				
38	$1s^26s''2b''$	7.150 092	1,454 424	$7.532903 \times 10^{-1}$				
39	$1s^26s''2h''r$	-1 268 021 × 10	-3 235 979	-2 194 769				
40	$1s^{2}6s'' 2p'' r_{34}^{2}$	6.969 993	2.588 735	1.640 764				
		_4	4	_1				
41	$1s2s2s''2p''r_{34}$	$-8.271090 \times 10^{-1}$	$-2.441955 \times 10^{-1}$	$-1.862829 \times 10^{-1}$				
42	$1s2s2s''2p''r_{34}^2$	$7.409538 \times 10^{-1}$	$5.009553 \times 10^{-2}$	$4.410859 \times 10^{-2}$				
43	1s3s2s"2 <b>p"r</b> <sub>34</sub>	$-8.067233 \times 10^{-1}$	$-1.628441  imes 10^{-2}$	$-1.016370 \times 10^{-2}$				
44	$1s3s2s''2p''r_{34}^2$	$1.981111 \times 10^{-1}$	$1.100518 \times 10^{-2}$	$-9.469788 \times 10^{-4}$				
45	$1s^2 4s'' 2p'' r_{12}$	$1.044517\! imes10^{-2}$	$-1.572072  imes 10^{-5}$	$-5.212241 \times 10^{-3}$				
46	$1s^2 4s'' 2p'' r_{12}^2$	$-1.382374 \times 10^{-1}$	$-1.031250 \times 10^{-1}$	$-7.445267 \times 10^{-2}$				
47	$1s^2 3s'' 3 p'' r_{12}^2$	$3.872209 \times 10^{-2}$	$-2.453043 \times 10^{-3}$	-1.296 422 × 10 <sup>-3</sup>				
48	1s2s3s"2.0"r2	1.699 981	$5.061419 \times 10^{-1}$	3.692 899×10 <sup>-1</sup>				
49	$1s2s3s''2h''r^2$	-1,194 784	$-2.307126 \times 10^{-1}$	$-1.842234 \times 10^{-1}$				
50	$1s3s3s''2p''r_{34}^2$	$1.871062 \times 10^{-1}$	$1.606 919 \times 10^{-2}$	$1.873209 \times 10^{-2}$				
51	1s2s4s"2p"	$-2.428041 \times 10^{-2}$	$-7.975607  imes 10^{-2}$	$-5.431176  imes 10^{-2}$				
52	1s3s4s"2p"	$1.017389 \times 10^{-1}$	$9.991232 \times 10^{-2}$	6.226 619×10 <sup>-2</sup>				
53	$2s^2 4s'' 2p''$	$-1.686599 \times 10^{-2}$	$-1.303484 \times 10^{-2}$	$-1.057253 \times 10^{-2}$				

TABLE II. 53-configuration wave function for the  ${}^{1}\!P$  states of Be1, C III, and O v.

For the configurations we use (three S-type STO's and one P-type STO),  $L_{max}$ =1 and Eq. (37) becomes simply

$$O_{1,0}^{L^2} = (3/3!)L_{-}L_{+}; \quad O_{1,0}^{L^2}Y_{10} = \frac{1}{2}\sqrt{2}Y_{10}.$$
(38)

# C. Integral Treatment

To complete the evaluation of all matrix elements needed for this work, a scheme is needed to handle the two-, three-, and four-electron integrals of the form

$$\int f_1(\mathbf{\tilde{r}}_1) f_2(\mathbf{\tilde{r}}_2) f_3(\mathbf{\tilde{r}}_3) f_4(\mathbf{\tilde{r}}_4) \\ \times g_1(r_{12}) g_2(r_{23}) \cdots g_b(r_{34}) \, dv \,. \tag{39}$$

The treatment of these integrals has been described in detail in a previous publication.<sup>34</sup>

## IV. CHOICE OF PARAMETERS AND CONFIGURATIONS

The wave functions employed here for the  ${}^{1}S$  and  ${}^{1}P$  states of BeI, CIII, and OV are of the form [Eq. (1)]

$$\psi = \sum_{k} C_{k} \Phi_{k}, \qquad (40)$$

where  $\Phi_k$  is a properly antisymmetrized and projected product of  $\chi$ , STO's defined by Eq. (2), and an  $r_{ij}^{\nu}$  factor. The STO basis is, in the Weiss<sup>41</sup> notation, the basis set 1s, 2s, ..., ns, 2p, ..., npfor the inner shell; 1s'', 2s'', ..., ns'', 2p'', ..., np'' for the outer shell. The STO orbital exponents [ $\zeta$  in Eq. (4)] used in this work are given in Table I with references to the works from which they were taken. No attempt was made to optimize the exponents, since it was felt that costly insignificant improvements would have been made to the energies obtained.<sup>13</sup> However, Weinhold's scheme for computing upper and lower bounds does have the feature that the orbital exponents can be varied directly with respect to S and its bounds, so that experimentation with shorter-length wave functions may be justified.<sup>42</sup>

The 105-configuration and 89-configuration  ${}^{1}S$  wave functions used in this work were truncations of the 107-term wave function reported in Ref. 34, with exponents as listed in Table I.<sup>41, 43, 44</sup> The 53-configuration  ${}^{1}P$  wave functions used in this work are given in Table II, where under configuration is listed only the unique

$$r_{ij}^{v_k} \prod_{s=1}^4 \Phi_{ks}(\vec{r}_s)$$

part of  $\Phi_k$  as defined by Eq. (2).

#### V. RESULTS

The results of the calculations we have done are presented in Table III, where we tabulate  $|\tilde{W}_{0}|$ ,

TABLE III.	Computed	values fo	r Beı,	Сш,	and Ov (in	amu).
------------	----------	-----------	--------	-----	------------	-------

C	onfigs.	$^{1}S \rightarrow ^{1}P$ transition								Enormal	
¹S	<sup>1</sup> P	1 📆	8	f	$E(^{1}S)$	E	( <sup>1</sup> <b>P</b> )	$\Delta E_{\rm calc}$	<sup>1</sup> S	<sup>1</sup> P	
Be	I										
	$46 \times 18$	0.930 01	10.869	1.40587	-14.661 55	5 -14.4	443 17	0.21838	-14.6662	-14.47267	
	89×53	0.90973	10.400	1.3447	-14.66648	85 -14.4	472 457	0.19403			
	$105 \times 53$	0.90934	10.391	1,3436	-14.666 54	6 -14.	472 457	0.19409			
С	11										
	89×53	0.442 61	2.1494	0.76537	-36.53337	-36.0	062 83	0.47054	-36.534 03	-36.06767	
0 1	,										
	89×53	0.291 37	1.0668	0.514 62	-68.40957	-67.0	684 33	0.72524	-68.41211	-67.68856	
C	onfigs										
iS	<sup>1</sup> P	$\langle \psi_b \mid (3/4\pi) X$	$\left[ \left[ X_{0} \right] \psi_{b} \right]  angle$	S <sub>a</sub> _	S <sub>b</sub> _	$\Delta_{ab}$	W <sub>&lt;</sub>	W>	f<	f,	
Be	I										
	$46 \times 18$	5.126	75	0.98977	0.795 42	2.06442	0.49794	1.224 5	2 0.4028	2,4313	
	89×53	4.517	69	0.99975 <sup>a</sup>	0.99862	1.920 95	0.86549	0.9526	1 1.2171	1.47758	
	105  imes 53	4.517	69	0.99986ª	0.99862	1.92114	0.87636	6 0.940 <b>9</b>	8 1.2479	1.438 69	
С	п										
	89×53	0.717	64	0.99975 <sup>a</sup>	0.996 65 <sup>a</sup>	0.72231	0.42507	0.4586	0.705 91	0.821 73	
0 1	,										
	89×53	0.298	96	0.999 53 <sup>a</sup>	0.999 01 <sup>a</sup>	0.46267	0.27675	<b>0.305</b> 5	9 0.46425	0.556 06	

<sup>a</sup> These values were computed using Weinberger's formula, Ref. 45.

the absolute multiplet strength S, the f values, term energies, the relative term energy  $\Delta E = E({}^{1}P) - E({}^{1}S)$ , and various quantities related to the calculation of upper and lower bounds to f. In evaluating  $S_{a-}$  and  $S_{b-}$ , either Eq. (24) was used or Weinberger's formula<sup>45</sup>

$$S^{2} \ge \frac{E_{n+1} - J_{0}}{E_{n+1} - E_{0}} \prod_{\nu=1}^{n} \frac{E_{\nu} - J_{0}}{E_{\nu} - E_{0}} \frac{J_{\nu} - E_{0}}{J_{\nu} - J_{0}},$$
 (41)

where  $E_0$ ,  $E_1$ , ...,  $E_{n+1}$  are the n+2 lowest eigenvalues of H and the  $J_i$  are the corresponding Rayleigh-Ritz estimates  $\langle \phi_i | H | \phi_i \rangle$ . Here the  $\phi_i$  are approximations to the exact eigenfunctions and it is assummed that

$$E_0 \leq J_0 \leq E_1 \leq J_1 \leq \cdots \leq E_n \leq J_n \leq E_{n+1}. \tag{42}$$

To use either Eq. (24) or (41), one must know either the nonrelativistic energies  $E_n$  or their lower bounds. For Be, the 107-term energy of Ref. 13 was obtained after examining 145 terms; the maximum energy contribution of the 38 terms which were dropped at various stages of the calculation was 0.000 072 amu. This gives us an "exact" nonrelativistic energy estimate of -14.666 62 amu for the Be ground state. To obtain the nonrelativistic energy estimate for the other states, we add the experimentally determined relative term energies to the ground-state estimate. The nonrelativistic energy estimates for CIII and OV states were obtained in the same manner, with the <sup>1</sup>S ground-state nonrelativistic energy estimates being taken from Weiss.<sup>41</sup>

For BeI, calculations were done on (i) the 46term <sup>1</sup>S wave function obtained by dropping all terms from the 105-term <sup>1</sup>S wave function with  $E_n - E_{m-1} < 0.001$  amu and the 18-term <sup>1</sup>P wave function obtained from the first 18 terms of the 53-term <sup>1</sup>P wave function; (ii) the 89-term <sup>1</sup>S wave function obtained by dropping all terms from the 105-term <sup>1</sup>S wave function with  $E_n - E_{n-1} < 0.0001$ amu and the 53-term <sup>1</sup>P wave function; and (iii) the 105-term <sup>1</sup>S wave function and the 53-term <sup>1</sup>P wave function. For calculations (i) and (ii), the term in brackets in Eq. (23) is small,<sup>46</sup> so that, approximately.

$$W_{ab} \stackrel{\leq}{\geq} S_{a\pm} S_{b\pm} \tilde{W}_{ab} \pm \epsilon_{a+} S_{b\pm} \Delta_{ab} . \tag{43}$$

Since in going from the 89-configuration <sup>1</sup>S wave function to the 105-configuration <sup>1</sup>S wave function,  $|\tilde{W}_0|, f$ , and  $\Delta_{ab}$  change by <0.2%, it would appear that going to even more accurate <sup>1</sup>S wave functions would merely tighten the theoretical upper and lower bounds, without changing the f value (to  $\leq 2\%$ ). Therefore, we feel that the computed fvalues are probably within 1% of the true values. Arguments may be advanced in support of this

TABLE IV.  $\Delta E = E({}^{1}P) - E({}^{1}S)$  for Be I, C III, and O v (in amu).

	Δ.	E
	Calc.	Obs.
Be 1	0.19409	0.193 95
Сш	0.47054	0.46636
O V	0.72524	0.723 55

conclusion.

(a) Extremely good wave functions on an energy criterion were employed for both the  ${}^{1}S$  and  ${}^{1}P$ states. The 53-term  ${}^{1}P$  wave function gave an energy of -14.472 57 amu, as compared with our exact nonrelativistic energy estimate of -14.47267 amu. This estimate involves an assumption that relativistic effects are the same for  ${}^{1}S$  and  ${}^{1}P$ states. Tatewaki et al.44 argue that the relativistic correction for a 2p electron in the L shell rather than a 2s electron is <0.0001 amu for Be. However, Verhaegen and Moser<sup>47</sup> place the relativistic correction at <0.001 amu, so that the  ${}^{1}P$  wave function may not be as accurate as claimed. In any event, for calculations (ii) and (iii) the accuracy of the  ${}^{1}P$  wave function is not critical, since in Eq. (43) it is only  $\epsilon_{a+}$  and not  $\epsilon_{b+}$  which is critical for accurate bounds.

(b) We obtain extremely good values for the difference in term energies  $\Delta E = E({}^{1}P) - E({}^{1}S)$ . This could be fortuitous, i.e., the energy difference is more accurate than either  $E({}^{1}S)$  or  $E({}^{1}P)$ . The calculated and experimental energy differences are tabulated in Table IV for BeI, CIII, and OV.

Taking 2% accuracy as a guide, we notice that at every step of the calculation the bounds do not reflect the true accuracy of the calculation. The bounds calculations are particularly bad in the  $46 \times 18$  case, which leads to the conjecture that the bounds formulas require trial wave functions which closely approximate the exact wave function



FIG. 1. F value as a function of 1/Z for various elements.

		Ве і		Сш		(	) v	
		HF	CI	HF	CI	HF	CI	
Burke <i>et al</i> .ª	length <i>l</i>	1.774	1.4237	1.0746	0 <b>.79</b> 30	0.7055	0.5250	
	velocity v	1.0132	1.3862	0.5627	0.8047	0.3039	0.4922	
Weiss <sup>b</sup>	l	1.813	1.410	1,112	0.770	0.741	0.517	
	v	0.951	1.455	0.537	0.843	0.353	0.570	
This work	l		1.3436		0.76537		0.51462	
		NIV	7					
Nussbaumer <sup>c</sup>	l	0.64						

TABLE V. Recent theoretical calculations of oscillator strengths.

<sup>a</sup> Reference 53.

<sup>b</sup>A. W. Weiss (private communication).

or else the bounds may be substantially different from the exact result and thus not very useful.<sup>48</sup> However, it should be pointed out that no attempt has been made to optimize the basis orbital exponents with respect to S and its bounds.

For CIII and OV we have employed 89-configuration wave functions, since the BeI calculation showed that the f value doesn't change much in going to a 105-term expansion. The rigorous bounds obtained are on the same level of accuracy as the BeI case, in which (i) there are more interleaving J's [Eq. (42)] to employ in the Weinberger estimates [Eq. (41)] and (ii)  $\Delta_{ab}$  is much smaller.

In Fig.  $1^{49-51}$  we graph our results as discussed earlier. The results for BeI, CIII, and OV are extrapolated to  $Z^{-1} = 0$  using the Z-expansion results of Weiss as a guide.<sup>39</sup> A few experimental results are included for comparison. The experimental results are uniformly too low and in most cases the error estimates appear unrealistic. Wiese<sup>52</sup> has conjectured that the experimental failure is due to cascading effects in the beam-foil experiments.

A few recent theoretical calculations, which include configuration interaction (CI), have been done. The f values of these calculations are shown in Table V along with our best values. These previous calculations treat the  $1s^2$  core as being fixed, and yield results which appear to be too high. In addition, the agreement of length and velocity is seen not to be an accurate measure of the error in the calculations. (Burke *et al.*<sup>53</sup> report a calculation for BeI in which length and velocity results give 1.3863 and 1.3860, respectively, although they recommend the values in Table V.) Thus, neglect of core (and intershell) correlation in this case (lowest  ${}^{1}S-{}^{1}P$  transition for the Be <sup>c</sup> Reference 33.

isoelectronic sequence) appears to give results in error by approximately 5-10%, at least for Be1.

#### **VI. CONCLUSIONS**

On the basis of the results presented here,<sup>54</sup> it seems reasonable to conclude that the introduction of  $r_{ij}$  coordinates for an atomic wave function with N>3 can lead to reliable f values (~2% accuracy). It also appears that to obtain reliable upper and lower bounds to oscillator strengths, wave functions of considerable accuracy must be employed. A level of accuracy for less-exact calculations which neglect core correlations has been established for the Be isoelectronic sequence. In addition, recommended values for the entire Be-isoelectronic-sequence  $1s^22s^2$   ${}^{1}S-1s^22s2p^{1}P$  transition have been presented (Fig. 1).

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Note added in manuscript. Figure 1 depicts the experimental status as of 1971. The experimental status as of 1973 is depicted in Fig. 2 of Ref. 55. The basic picture doesn't change; all the beam-foil measurements seem to be consistently low. The experiment that seems to be in the best shape is that of Lawrence and Savage, Ref. 50, which is a phase-shift measurement. We thank Dr. A. W. Weiss for bringing this reference to our attention.

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- <sup>23</sup>Here Z is the nuclear charge and  $-(1/2)\Delta_i$  is the operator corresponding to the kinetic energy of electron *i*. Here and in the remainder of this work we use Hartree atomic units, in which the unit of energy is chosen as  $m_g e^4/\hbar^2 = 1$  amu (of energy), with  $m_e$  being the mass of the electron. Note here that one could use reduced Hartree atomic units as in Ref. 13, but then the conversion factors from experimental to theoretical values would be different for different Z. The primary results of the calculation (energy eigenvalues, absolute multiplet strengths) are the same in either system of units.
- <sup>24</sup>Actually the configuration designation is rigorous only for Russell-Saunders (*LS*) coupling in a central field. We retain the terminology here to conform to standard spectroscopic notation with the understanding that by  $1s^22s^{21}S$  and  $1s^22s2p^{1}P$  we mean that lowest terms of those respective symmetries. For a generalization of the multiplet strength form with interacting configurations see, e.g., R. N. Zare, J. Chem. Phys. <u>47</u>, 3561 (1967).
- <sup>25</sup>Note here that we are interested in computing the f value for a transition from one spectroscopic *term* to another. Spectroscopic terms consist of *levels* that

differ in the total angular momentum J. Each level is (2J+1) degenerate in the absence of magnetic fields. The levels of different J belonging to the same term have different energies owing to spin-orbit interaction; neglecting spin-orbit interaction they are degenerate. Thus in the absence of magnetic fields and neglecting spin-orbit interactions [as use of the Hamiltonian Eq. (5) implies], states of the term are degenerate. To get the multiplet oscillator strength we sum over all the levels of each term and all the states of each level. This sum is  $\sum_{J} \sum_{M_J} = \sum_{J} (2J+1) = (2L+1)(2S+1)$ . In LS coupling, neglecting spin-orbit interactions, we can equivalently take  $\sum_{M_L} \sum_{M_S} M_S$ , as in Eq. (6). The energy difference  $E(\gamma'L'S') - E(\gamma LS)$  is the average energy difference of states or the weighted energy difference of levels. Since the energy eigenvalues obtained by use of the Hamiltonian given by Eq. (5) depend only on  $\gamma, L, S$ , a computed term separation  $\Delta E = E(\gamma' L'S')$  $-E(\gamma LS)$  can be obtained by merely substracting the two appropriate energy values obtained with the wave functions we use.

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- <sup>32</sup>The length formula was chosen over the velocity formula because of the relative ease of computing the matrix elements. For transitions without change in principal quantum number, Crossley [Ref. (3)] argues that this is the correct formula to choose. However, it should be of interest to see whether length and velocity forms are converging upon the same value with the wave functions employed in this work, and the appropriate calculations are planned [J. S. Sims (unpublished)].
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