## Leading-order relativistic corrections to the hydrogen molecular ion

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We present the improved calculations over the results of Korobov [Phys. Rev. A 74, 052506 (2006)] for the energy levels and hyperfine structure of the hydrogen molecular ion  $H_2^+$ , including the relativistic corrections of orders  $\alpha^2$  Ry and  $\alpha^2 m_e/m_p$  Ry, using variationally constructed basis sets in Hylleraas coordinates. In particular, the computational uncertainty of 3 kHz in Korobov's calculation for the  $\alpha^2$  Ry correction to the two-photon transition frequency  $(L=0, v=0) \rightarrow (L=0, v=1)$  is eliminated.

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I. INTRODUCTION

In recent years,  $H_2^+$  and its isotopes were cooled to a few mK sympathetically with laser-cooled beryllium ions by Schiller and co-workers [1,2] aiming at a spectroscopic accuracy of  $10^{-10}$  level. Meanwhile, experiments on high-resolution spectroscopy for the two-photon transition  $(L, v = 0) \rightarrow (L, v = 1)$  in  $H_2^+$  were proposed by Karr *et al.* [3,4] to achieve a few kHz accuracy. The two-photon transition between the lowest rovibrational states of  $H_2^+$  can be used to derive an improved value of the electron-proton mass ratio  $m_e/m_p$  [5,6]. Schiller and Korobov [7] studied in detail the theoretical dependence of the rovibrational transition frequency of  $H_2^+$  on  $m_e/m_p$ .

The CODATA recommended values [8] of the fundamental constants for  $m_e/m_p$  is accurate to about five parts in 10<sup>10</sup>. In order to obtain an improved value for  $m_e/m_p$  by 1 order of magnitude in precision, the theory and experiment must be accurate at least to a few parts in  $10^{12}$  or better [2,9,10]. Thus, relativistic and QED corrections of orders  $\alpha^5$  Ry and  $\alpha^6$  Ry should be included in the calculations, which in turn requires very high-precision nonrelativistic wave functions. Korobov [11] evaluated the nonrelativistic rovibrational energy levels  $\tilde{H}_2^+$  to a precision of  $10^{-15}$  for v=0-4 and L =0-4. After that, the relativistic and radiative corrections up to  $\alpha^6$  Ry for H<sub>2</sub><sup>+</sup> and HD<sup>+</sup> have been calculated [11–14]. Relativistic corrections of order  $\alpha^6(m/M)$  Ry to the hyperfine structure have also been estimated as well [15,16]. However, these results have not been confirmed independently. Furthermore, the computational uncertainty in the lowestorder relativistic correction of  $\alpha^2$  Ry is about 3 kHz, which is at the level of 40 parts in  $10^{12}$  to the total transition frequency of  $(0,0) \rightarrow (0,1)$  [11,14]. Up to now, the most precise nonrelativistic energies for the low-lying states of H<sub>2</sub><sup>+</sup> were determined variationally by Li et al. [10] and Hajikata et al. [17] using a completely different approach from [11]. As an example, the ground-state energy of  $H_2^+$  has been calculated to 30 significant digits. The purpose of this Brief Report is to report our calculations for the leading relativistic corrections to the low-lying rovibrational states in  $H_2^+$ , which provides an independent verification of Korobov's calculations. At the same time, much-improved results for the energy levels and hyperfine structure will be presented. Atomic units  $(\hbar = e = m_e = 1)$  are used throughout this Brief Report, unless otherwise stated. The proton-electron mass ratio is chosen to be 1836.152 672 61(85) [18].

#### **II. NONRELATIVISTIC WAVE FUNCTIONS**

Consider the hydrogen molecular ion  $H_2^+$ . After eliminating the center of mass degrees of freedom, the eigenvalue problem for the nonrelativistic Hamiltonian  $H_0$  becomes [10]

$$H_0 \Psi = E_0 \Psi, \tag{1}$$

with

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$$H_0 = \lambda_1 \nabla_{r_1}^2 + \lambda_2 \nabla_{r_2}^2 + \lambda_{12} \nabla_{r_1} \cdot \nabla_{r_2} + V.$$
<sup>(2)</sup>

In the above,  $\mathbf{r}_1$  and  $\mathbf{r}_2$  are, respectively, the position vectors of the electron and one proton, relative to the other proton situated at the origin,  $\lambda_1 = -(1+m_p)/(2m_p)$ ,  $\lambda_2 = \lambda_{12} = -1/m_p$ ,  $V = -1/r_1 + 1/r_2 - 1/r_{12}$  is the coulomb interaction among the three particles, and  $\mathbf{r}_{12} = \mathbf{r}_1 - \mathbf{r}_2$ . The energy eigenvalue problem for  $H_0$  is solved variationally in the following Hylleraas basis set:

$$\phi_{ijk}(\mathbf{r}_1, \mathbf{r}_2) = r_1^i r_2^j r_{12}^k e^{-\alpha r_1 - \beta r_2} \mathcal{Y}_{l_1 l_2}^{LM}(\mathbf{r}_1, \mathbf{r}_2), \qquad (3)$$

where  $\mathcal{Y}_{l_1 l_2}^{LM}(\mathbf{r}_1, \mathbf{r}_2)$  is the vector coupled product of spherical harmonics for the electron and the proton. More details on the construction of basis set for  $H_2^+$  can be found in [10]. The basic type of integrals required in the calculation of matrix elements can be evaluated analytically [19] using Perkins' expansion for  $r_{12}^c$ . Singularities of singular integrals that appear in the evaluation of Breit interaction can be canceled out according to the procedure in [20].

### **III. LEADING-ORDER RELATIVISTIC CORRECTIONS**

The leading-order relativistic corrections of order  $\alpha^2$  from the Breit-Pauli Hamiltonian are well established and may be found in [11,21]. The complete spin-independent contributions are

$$H_{\rm spin-indep} = \alpha^2 H_{\rm Breit} + H_{\rm nuc}, \qquad (4)$$

$$H_{\text{Breit}} = H_{\text{rc}} + E_{\text{kin}} + H_{\text{tr-ph}} + H_{\text{Darwin}}, \qquad (5)$$

$$H_{\rm rc} = -\frac{\nabla_{r_1}^4}{8} + \frac{4\pi}{8} [\delta(\mathbf{r}_1) + \delta(\mathbf{r}_{12})], \qquad (6)$$

$$H_{\rm kin} = 2 \left[ -\frac{1}{8m_p^3} \nabla_{r_2}^4 \right],$$
 (7)

$$H_{\text{tr-ph}} = \frac{1}{2m_p r_1^3} [r_1^2 \nabla_{r_1} \cdot (\nabla_{r_1} + \nabla_{r_2}) + \mathbf{r}_1 \mathbf{r}_1 : (\nabla_{r_1} + \nabla_{r_2}) \nabla_{r_1}] - \frac{1}{2m_p^2 r_2^3} [r_2^2 \nabla_{r_2} \cdot (\nabla_{r_1} + \nabla_{r_2}) + \mathbf{r}_2 \mathbf{r}_2 : (\nabla_{r_1} + \nabla_{r_2}) \nabla_{r_2}]$$

$$-\frac{1}{2m_p r_{12}^3} [r_{12}^2 \nabla_{r_1} \cdot \nabla_{r_2} + \mathbf{r}_{12} \mathbf{r}_{12} : \nabla_{r_1} \nabla_{r_2}], \qquad (8)$$

$$H_{\text{Darwin}} = \frac{1}{8m_p^2} [\delta(\mathbf{r}_1) + \delta(\mathbf{r}_{12})], \qquad (9)$$

$$H_{\rm nuc} = \frac{2\pi (R_p/a_0)^2}{3} [\delta(\mathbf{r}_1) + \delta(\mathbf{r}_{12})].$$
(10)

In the above,  $H_{\rm rc}$  is the major relativistic contribution of order  $\alpha^2$  from the bound electron,  $H_{\rm kin}$  is the correction due to the relativistic kinetic energy of the two protons,  $H_{\rm tr-ph}$  is from the transverse photon exchange,  $H_{\rm Darwin}$  is the Darwin term, and  $H_{\rm nuc}$  is the leading-order correction from the finite nuclear charge distribution, where  $R_p$ =0.8768(69) fm [8] is the proton root-mean-square charge radius and  $a_0$  is the Bohr radius. The global operator method, proposed by Drachman [22], is used to calculate some singular operators such as  $\delta(\mathbf{r}_i)$  and  $\nabla_{r.}^4$ .

# IV. LEADING-ORDER RELATIVISTIC CORRECTIONS TO HYPERFINE STRUCTURE

The leading-order relativistic corrections to hyperfine structure of  $H_2^+$  are described by the spin-dependent terms of

TABLE I. Convergence of  $\langle \nabla_{r_1}^4 \rangle$  and  $\langle \nabla_{r_2}^4 \rangle$  in the state (0,0). N denotes the size of basis set.

| N       | $\langle \nabla^4_{r_1} \rangle$ | $\langle \nabla^4_{r_2} \rangle$ |
|---------|----------------------------------|----------------------------------|
| 6084    | 6.2856600593122250               | 79.7976493122364                 |
| 6795    | 6.2856600593122253               | 79.7976493122401                 |
| 7560    | 6.2856600593122214               | 79.7976493122417                 |
| 8381    | 6.2856600593122213               | 79.7976493122422                 |
| 9260    | 6.2856600593122207               | 79.7976493122427                 |
| Extrap. | 6.2856600593122200(7)            | 79.7976493122432(5)              |

Breit-Pauli Hamiltonian that sums up pairwise interactions of the three particles. The electron-proton and proton-proton spin-orbit interactions are expressed as

$$\begin{split} V_{\text{so-pe}} &= \alpha^2 \Bigg[ -\frac{1+2\kappa_e}{2} \bigg( \frac{1}{r_1^3} (i\mathbf{r}_1 \times \nabla_{r_1}) + \frac{1}{r_{12}^3} (i\mathbf{r}_{12} \times \nabla_{r_1}) \bigg) \cdot \mathbf{s}_e \\ &+ \frac{1+\kappa_e}{m_p} \bigg( \frac{1}{r_1^3} (i\mathbf{r}_1 \times \nabla_{r_1}) + \frac{1}{r_1^3} (i\mathbf{r}_1 \times \nabla_{r_2}) \\ &+ \frac{1}{r_{12}^3} (i\mathbf{r}_{12} \times \nabla_{r_2}) \bigg) \cdot \mathbf{s}_e + \frac{1+2\kappa_p}{m_p^2} \bigg( \frac{1}{r_1^3} (i\mathbf{r}_1 \times \nabla_{r_1}) \cdot \mathbf{I}_0 \\ &+ \frac{1}{r_1^3} (i\mathbf{r}_1 \times \nabla_{r_2}) \cdot \mathbf{I}_0 + \frac{1}{r_{12}^3} (i\mathbf{r}_{12} \times \nabla_{r_2}) \cdot \mathbf{I}_2 \bigg) \\ &- \frac{1+\kappa_p}{2m_p} \bigg( \frac{1}{r_1^3} (i\mathbf{r}_1 \times \nabla_{r_1}) \cdot \mathbf{I}_0 + \frac{1}{r_{12}^3} (i\mathbf{r}_{12} \times \nabla_{r_1}) \cdot \mathbf{I}_2 \bigg) \bigg], \end{split}$$
(11)

$$V_{\text{so-pp}} = \alpha^{2} \left[ \frac{1 + 2\kappa_{p}}{2m_{p}^{2}} \left( \frac{1}{r_{2}^{3}} (i\mathbf{r}_{2} \times \boldsymbol{\nabla}_{r_{1}}) \cdot \mathbf{I}_{0} + \frac{1}{r_{2}^{3}} (i\mathbf{r}_{2} \times \boldsymbol{\nabla}_{r_{2}}) \cdot \mathbf{I}_{0} \right. \\ \left. + \frac{1}{r_{2}^{3}} (i\mathbf{r}_{2} \times \boldsymbol{\nabla}_{r_{2}}) \cdot \mathbf{I}_{2} \right) + \frac{\mu_{p}}{m_{p}^{2}} \left( \frac{1}{r_{2}^{3}} (i\mathbf{r}_{2} \times \boldsymbol{\nabla}_{r_{1}}) \cdot \mathbf{I}_{2} \right. \\ \left. + \frac{1}{r_{2}^{3}} (i\mathbf{r}_{2} \times \boldsymbol{\nabla}_{r_{2}}) \cdot \mathbf{I}_{2} + \frac{1}{r_{2}^{3}} (i\mathbf{r}_{2} \times \boldsymbol{\nabla}_{r_{2}}) \cdot \mathbf{I}_{0} \right) \right], \qquad (12)$$

TABLE II. Expectation values of various operators for some rovibrational states of  $H_2^+$  and comparison to Korobov's work [11].

| (L,v)                                       | (0,0)                    | (1,0)                    | (0,1)                    |
|---|--------------------------|--------------------------|--------------------------|
| $\overline{\langle \nabla_{r_1}^4 \rangle}$ | 6.28566006               | 6.27803905               | 6.12451981               |
| This work                                   | 6.2856600593122200(7)    | 6.278039037386287(5)     | 6.124519807747879(3)     |
| $\langle \nabla_{r_2}^4 \rangle$            | 79.7976                  | 85.0505                  | 334.898                  |
| This work                                   | 79.7976493122432(5)      | 85.050455613540(5)       | 334.89830192549(4)       |
| $\langle \delta(\mathbf{r}_1) \rangle$      | 0.206736476              | 0.206491321              | 0.201310665              |
| This work                                   | 0.2067364762888147850(1) | 0.206491320158817300(3)  | 0.2013106647020826675(1) |
| $R_{pe}$                                    | 1.17012                  | 1.16881                  | 1.14081                  |
| This work                                   | 1.170117625033754694(1)  | 1.16881866381904511(8)   | 1.1408052271300473(2)    |
| $R_{pp}$                                    | 4.60193                  | 4.83433                  | 12.8961                  |
| This work                                   | 4.601934312504856498(2)  | 4.8343364736703696570(6) | 12.89614649485648(2)     |

TABLE III. Numerical values for reduced matrix elements of orbital parts of spin-dependent operators in the state (1,0).

| Operator                 | Value                               |  |  |
|--------------------------|-------------------------------------|--|--|
| $\overline{S_{1}^{(1)}}$ | $0.2704494968631(4) \times 10^{-4}$ |  |  |
| $S_{2}^{(2)}$            | -0.2996739237238803(7)              |  |  |
| $S_{2}^{(1)}$            | -0.2968096504450887(5)              |  |  |
| $S_{1}^{(2)}$            | $0.139939978335(3) \times 10^{-4}$  |  |  |
| $S_{1}^{(12)}$           | $0.270449496868(5) \times 10^{-4}$  |  |  |
| $S_2^{(12)}$             | 0.296782605495398(5)                |  |  |
| $T_{pe}$                 | 0.21539521185521(1)                 |  |  |
| $T_{pp}$                 | 0.3283262297065680565838(1)         |  |  |

where  $\mu_e = -(1 + \kappa_e)$  is the magnetic moment of the electron in Bohr magneton,  $\mu_p = 1 + \kappa_p$  is the magnetic moment of the proton in nuclear magneton, and  $\mathbf{I}_0$ ,  $\mathbf{I}_2$ , and  $\mathbf{s}_e$  are the spin operators for the two nuclei and the electron, respectively. The electron-proton and proton-proton spin-spin tensor interactions are expressed as

$$V_{\rm ss-ep} = \alpha^2 \frac{\mu_e \mu_p}{m_p} \left\{ \frac{r_1^2(\mathbf{s}_e \cdot \mathbf{I}_0) - 3(\mathbf{r}_1 \cdot \mathbf{s}_e)(\mathbf{r}_1 \cdot \mathbf{I}_0)}{r_1^5} + \frac{r_{12}^2(\mathbf{s}_e \cdot \mathbf{I}_2) - 3(\mathbf{r}_{12} \cdot \mathbf{s}_e)(\mathbf{r}_{12} \cdot \mathbf{I}_2)}{r_{12}^5} - \frac{8\pi}{3} [\delta(\mathbf{r}_1)(\mathbf{s}_e \cdot \mathbf{I}_0) + \delta(\mathbf{r}_{12})(\mathbf{s}_e \cdot \mathbf{I}_2)] \right\}, \quad (13)$$

$$V_{\text{ss-pp}} = \alpha^2 \frac{\mu_p^2}{m_p^2} \left[ \frac{r_2^2 (\mathbf{I}_0 \cdot \mathbf{I}_2) - 3(\mathbf{r}_2 \cdot \mathbf{I}_0)(\mathbf{r}_2 \cdot \mathbf{I}_2)}{r_2^5} - \frac{8\pi}{3} \delta(\mathbf{r}_2) (\mathbf{I}_0 \cdot \mathbf{I}_2) \right].$$
(14)

The nuclear spin-spin tensor interaction cannot be neglected and was first considered in [16].

## **V. CALCULATIONS AND RESULTS**

Table I presents a convergence study for  $\langle \nabla_{r_1}^4 \rangle$  and  $\langle \nabla_{r_2}^4 \rangle$  in (0,0) evaluated using the global operator method. Numerical values for various spin-independent operators of the Breit interaction are listed in Table II, together with a comparison

TABLE IV. Contributions to the transition frequency  $(0,0) \rightarrow (0,1)$  in MHz.

|                           | This work        | Korobov                    |
|---------------------------|------------------|----------------------------|
| $\Delta E_{\rm nr}$       | 65687511.0686(4) | 65687511.0686 <sup>a</sup> |
| $\Delta E_{\text{Breit}}$ | 1091.081355(2)   | 1091.081(03) <sup>b</sup>  |
| $\Delta E_{\rm nuc}$      | -0.0410(6)       | -0.0410(3) <sup>b</sup>    |

<sup>a</sup>Table VI of Ref. [11].

<sup>b</sup>Obtained using data of Table IV in Ref. [11].

to Korobov's work [11]. Table III contains the numerical values for various spin-dependent operators in the state (1,0). The notations for orbit-orbit operators are as follows:

$$R_{pe} = -\left\langle \frac{1}{r_1^3} [r_1^2 \nabla_{r_1} \cdot (\nabla_{r_1} + \nabla_{r_2}) + \mathbf{r}_1 \mathbf{r}_1 : (\nabla_{r_1} + \nabla_{r_2}) \nabla_{r_1}] \right\rangle,$$
(15)

and the definition for  $R_{pp}$  is obtained from  $R_{pe}$  by  $1 \leftrightarrow 2$ . The notations for the reduced matrix elements of orbital parts of spin-dependent operators are

$$S_{j}^{(i)} = \left\langle \left\| \frac{1}{r_{i}^{3}} (i\mathbf{r}_{i} \times \nabla_{r_{j}}) \right\| \right\rangle, \quad S_{j}^{(12)} = \left\langle \left\| \frac{1}{r_{12}^{3}} (i\mathbf{r}_{12} \times \nabla_{r_{j}}) \right\| \right\rangle,$$
$$T_{pe} = \left\langle \left\| [\mathbf{r}_{1} \otimes \mathbf{r}_{1}]^{(2)} \right\| \right\rangle, \quad T_{pp} = \left\langle \left\| [\mathbf{r}_{2} \otimes \mathbf{r}_{2}]^{(2)} \right\| \right\rangle, \quad (16)$$

where i, j=1, 2 and the superscript 2 stands for a second-order tensor.

Table IV lists the contributions to the two-photon transition frequency  $(0,0) \rightarrow (0,1)$  from the nonrelativistic energy, the Breit interaction, and the nuclear charge distribution. In the table, the uncertainty of 0.4 kHz in  $\Delta E_{nr}$  is due entirely to the uncertainty in the Rydberg constant [8].  $\Delta E_{\text{Breit}}$  is the leading relativistic correction of order  $\alpha^2$ . In Korobov's calculations [11,14], the uncertainty in  $\Delta E_{\text{Breit}}$  is estimated to be 3 kHz that mainly comes from the computational uncertainty in  $\langle \nabla_{r_1}^4 \rangle$ . By contrast, the uncertainty of 2 Hz in our value is due to the uncertainty in the fine-structure constant  $\alpha$ . Additional, the uncertainty due to proton-electron mass ratio [16]in Breit interactions is about 0.003 Hz. The largest uncertainty of 0.6 kHz (about ten parts in 10<sup>12</sup> to the total transition frequency) is from the proton root-mean-square charge radius, which is at the level of 0.8% accuracy. A muonic hydrogen Lamb-shift experiment is currently underway at

TABLE V. Hyperfine structure of  $H_2^+$  for the state (1,0), in MHz.

| $\Delta E_{ m hfs}$ |                | $[C_1^\pm,C_3^\pm]$ |                             |                       |
|---------------------|----------------|---------------------|-----------------------------|-----------------------|
| $(\widetilde{F},J)$ | This work      | Ref. [15]           | This work                   | Ref. [[4]].           |
| (1/2,3/2)           | -930.433201(8) | -930.4332           | [-0.999878130, 0.015611707] | [-0.999878, 0.015612] |
| (1/2,1/2)           | -910.757885(8) | -910.7579           | [-0.999243469, 0.038890742] | [-0.999243, 0.038891] |
| (3/2,5/2)           | 474.106321(4)  | 474.1063            | [0, 1]                      | [0, 1]                |
| (3/2,3/2)           | 481.953402(4)  | 481.9534            | [0.015611707, 0.999878130]  | [0.015612, 0.999878]  |
| (3/2,1/2)           | 385.398521(4)  | 385.3985            | [0.038890742, 0.999243469]  | [0.038891, 0.999243]  |

TABLE VI. Numerical values of the hyperfine coefficients of the effective Hamiltonian for the state (1,0), in MHz.

| Coeffi-<br>cient | This work        | Ref. [15] | Ref. [24]     | Ref. [25] <sup>a</sup> |
|------------------|------------------|-----------|---------------|------------------------|
| $b_F$            | 922.991798(8)    | 922.992   | 922.990       | 922.940(20)            |
| c <sub>e</sub>   | 42.41631862(4)   | 42.4163   | 42.421        | 42.348(29)             |
| $c_I$            | -0.0416729322(4) | -0.04168  | $-0.0417^{b}$ |                        |
| $d_1$            | 128.489751(1)    | 128.490   | 128.482       | 128.259(26)            |
| $d_2$            | -0.297559002(5)  | -0.2975   |               |                        |

<sup>a</sup>Experimental results.

<sup>b</sup>Reference [26].

the Paul Scherrer Institute [23] with a purpose of reducing the uncertainty of  $R_p$  to 0.1%. Thus a further reduction of uncertainty in  $\Delta E_{\text{nuc}}$  to a few parts in 10<sup>12</sup> is feasible in the near future.

The leading-order relativistic corrections to the hyperfine structure are calculated in the angular coupling scheme

$$\mathbf{I} = \mathbf{I}_0 + \mathbf{I}_2, \quad \mathbf{F} = \mathbf{I} + \mathbf{s}_e, \quad \mathbf{J} = \mathbf{L} + \mathbf{F}.$$

The possible values of *F* and *J* are listed in [4]. It should be pointed out that *F* is an approximate quantum number for the case when *L* is odd. Thus, for an *L*=1 state, *I* can only be 1 and the eigenstates of  $J=L\pm\frac{1}{2}$  are the linear combination of F=1/2 and F=3/2 hyperfine states [see Eq. (20) in [4]]

$$\left| v, L, s_{e}, I, \widetilde{F}, J = L \pm \frac{1}{2} \right\rangle = C_{1}^{\pm} \left| v, L, \frac{1}{2}, 1, \frac{1}{2}, L \pm \frac{1}{2} \right\rangle$$
$$+ C_{3}^{\pm} \left| v, L, \frac{1}{2}, 1, \frac{3}{2}, L \pm \frac{1}{2} \right\rangle.$$
(17)

Since the mixing between F=1/2 and F=3/2 states is quite weak, one can label the states by the dominant *F*, denoted by  $\tilde{F}$ . Here only the hyperfine structure of (1,0) state is considered. Our results are listed in Table V. The uncertainties in our calculations are mainly due to the proton magnetic moment. One can see that good agreement with [4,15] is obtained.

The effective spin Hamiltonian is used according to [15]

$$H_{\text{eff}} = b_F (\mathbf{I} \cdot \mathbf{s}_e) + c_e (\mathbf{L} \cdot \mathbf{s}_e) + c_I (\mathbf{L} \cdot \mathbf{I}) + \frac{d_1}{(2L-1)(2L+3)} \left(\frac{2}{3} \mathbf{L}^2 (\mathbf{I} \cdot \mathbf{s}_e) - [(\mathbf{L} \cdot \mathbf{I})(\mathbf{L} \cdot \mathbf{s}_e) + (\mathbf{L} \cdot \mathbf{s}_e)(\mathbf{L} \cdot \mathbf{I})]\right) + \frac{d_2}{(2L-1)(2L+3)} \left(\frac{1}{3} \mathbf{L}^2 \mathbf{I}^2 - \frac{1}{2} (\mathbf{L} \cdot \mathbf{I}) - (\mathbf{L} \cdot \mathbf{I})^2\right).$$
(18)

Numerical values of the coefficients contained in Eq. (18) are listed in Table VI. Uncertainties in the coefficients are mainly caused by the uncertainty in  $\mu_p$  [8], while uncertainties from the operators in Table III are negligible. It is worthwhile to emphasize that the 10 Hz difference in  $c_I$  with Korobov's calculation [15] in Table VI is related to the discrepancy in  $S_1^{(2)}$ , where Korobov's value [27] is about  $10^{-16}$  whereas ours is  $10^{-5}$ . However, it will not introduce noticeable effect on the final results of the hyperfine splittings at the level of kHz [16].

In conclusion, we have presented an independent calculation of the leading-order relativistic corrections to the (0,0), (0,1), and (1,0) states of  $H_2^+$ , including the hyperfine structure. The computational uncertainties that exist in previous work have now been eliminated and all uncertainties are due entirely to the fundamental physical constants.

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