

Relativistic corrections of $m\alpha^6$ order to the rovibrational spectrum of H_2^+ and HD^+ molecular ions

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The major goal of high-precision studies of rovibrational states in the hydrogen molecular ions is to provide an alternative way for improving the electron-to-proton mass ratio, or the atomic mass of electron. By now the complete set of relativistic and radiative corrections have been obtained for a wide range of rovibrational states of H_2^+ and HD^+ up to order $R_\infty\alpha^4$. In this work we complete calculations of various contributions to the $R_\infty\alpha^4$ order by computing the relativistic corrections to the binding energy of electron.

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In recent years several laser spectroscopy experiments have been proposed [1,2] for high precision measurements of the vibrational spectrum of the hydrogen molecular ions H_2^+ and HD^+ . These experiments have metrological interest and are aimed at a sup-ppb precision. In order to improve the present accuracy of the electron-to-proton mass ratio [3] the uncertainty of the spectroscopic data (as well as of the theoretical calculations of the spectra to compare to) should be below 1 part per billion (1 ppb). To meet these stringent requirements, the theoretical calculations should achieve at least a level of 10 kHz (or $\sim 10^{-11}$ in atomic units).

While the variational calculations of the nonrelativistic energies have reached a numerical precision of $10^{-15} - 10^{-30}$ a.u. [4–10], the radiative and relativistic corrections have not been presented in the literature with required accuracy. Only recently high-precision variational calculations for the ro-vibrational states in the range of the total orbital momentum $L=0-4$ and vibrational quantum number $v=0-4$ for the H_2^+ and HD^+ molecular ions along with relativistic and radiative corrections of orders $R_\infty\alpha^2$, $R_\infty\alpha^2(m/M)$, $R_\infty\alpha^3$, $R_\infty\alpha^3(m/M)$, and, partially, $R_\infty\alpha^4$ have been obtained [11]. Here m is a mass of an electron and M is a mass scale characteristic for molecular nuclei.

The total $R_\infty\alpha^4$ order contribution to the energy in a non-recoil limit consist of the radiative one- and two-loop corrections and the relativistic correction for the Dirac electron. The spin-orbit contribution to the hyperfine structure, which arises from the anomalous magnetic moment, has been considered separately for H_2^+ in Ref. [12] and for HD^+ in Ref. [13]. The radiative corrections are known in an analytic form (see, for example, Refs. [14,15]).

The major aim of the present work is to calculate relativistic contribution due to a bound electron to the energies of rovibrational states of H_2^+ and HD^+ in the same range of total orbital angular momentum L and vibrational quantum number v as in Ref. [11]. The atomic units ($\hbar=e=m_e=1$) are used throughout.

I. RADIATIVE CORRECTIONS OF ORDER $R_\infty\alpha^4$

For a given relative accuracy of $\sim 10^{-10} - 10^{-11}$ recoil corrections of orders $R_\infty\alpha^4(m/M)$ and higher are small and may be neglected. That allows us to reduce calculation of higher

order corrections for the Coulomb three-body system to the problem of a bound electron in an external field.

The radiative corrections of order $R_\infty\alpha^4$ in the external field approximation can be expressed as follows [14,15]:

$$\begin{aligned} E_{\text{se}}^{(4)} &= \alpha^4 \frac{4\pi}{m_e^2} \left(\frac{139}{128} - \frac{1}{2} \ln 2 \right) \langle Z_1^2 \delta(\mathbf{r}_1) + Z_2^2 \delta(\mathbf{r}_2) \rangle, \\ E_{\text{anom}}^{(4)} &= \alpha^2 \frac{\pi}{m_e^2} \left[\left(\frac{\alpha}{\pi} \right)^2 \left(\frac{197}{144} + \frac{\pi^2}{12} - \frac{\pi^2}{2} \ln 2 + \frac{3}{4} \zeta(3) \right) \right] \\ &\quad \times \langle Z_1 \delta(\mathbf{r}_1) + Z_2 \delta(\mathbf{r}_2) \rangle, \\ E_{\text{vp}}^{(4)} &= \frac{4\alpha^3}{3m_e^2} \left[\frac{5\pi\alpha}{64} \right] \langle Z_1^2 \delta(\mathbf{r}_1) + Z_2^2 \delta(\mathbf{r}_2) \rangle, \\ E_{2\text{loop}}^{(4)} &= \frac{\alpha^4}{m_e^2 \pi} \left[-\frac{6131}{1296} - \frac{49\pi^2}{108} + 2\pi^2 \ln 2 - 3\zeta(3) \right] \\ &\quad \times \langle Z_1 \delta(\mathbf{r}_1) + Z_2 \delta(\mathbf{r}_2) \rangle. \end{aligned} \quad (1)$$

The last equation includes both Dirac form factor and polarization operator contributions.

II. RELATIVISTIC CORRECTIONS OF ORDER $R_\infty\alpha^4$

The most problematic contribution of $R_\infty\alpha^4$ order is the relativistic correction for a Dirac electron. It can be obtain within the adiabatic two-center approximation as follows (for details, see Ref. [16]).

We start from the nonrelativistic Schrödinger equation with the Hamiltonian

$$H_0 = \frac{p^2}{2m_e} + V, \quad V = -\frac{Z_1}{r_1} - \frac{Z_2}{r_2}. \quad (2)$$

The total contribution to the energy of a bound electron at the $R_\infty\alpha^4 \sim m_e c^2 \alpha^6$ order is defined by

$$\Delta E^{(6)} = \langle H_B Q (E_0 - H_0)^{-1} Q H_B \rangle + \langle H^{(6)} \rangle. \quad (3)$$

Here $H^{(6)}$ is the effective Hamiltonian for the interaction of an electron with the external field of two centers in this order, which can be expressed in the form

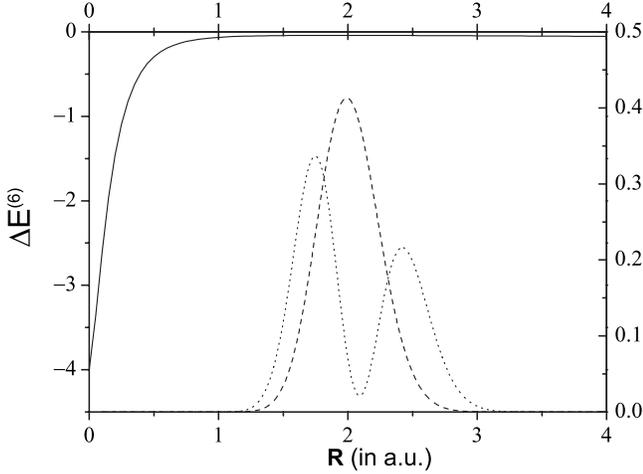


FIG. 1. Adiabatic “effective” potentials for the relativistic $m\alpha^6$ order correction for H_2^+ molecular ion ($Z_1=Z_2=1$). Energies are in (atomic units) $\times \alpha^4$.

$$H^{(6)} = \frac{p^6}{16m_e^5} + \frac{(\mathcal{E}_1 + \mathcal{E}_2)^2}{8m_e^3} - \frac{3\pi}{16m_e^4} \{p^2[\rho_1 + \rho_2] + [\rho_1 + \rho_2]p^2\} + \frac{5}{128m_e^4} (p^4V + Vp^4) - \frac{5}{64m_e^4} (p^2Vp^2),$$

$$\mathcal{E}_i = -Z_i \mathbf{r}_i / r_i^3, \quad \rho_i = Z_i \delta(\mathbf{r}_i). \quad (4)$$

H_B is the Breit-Pauli interaction

$$H_B = -\frac{p^4}{8m_e^3} + \frac{\pi}{2m_e^2} [Z_1 \delta(\mathbf{r}_1) + Z_2 \delta(\mathbf{r}_2)] + \left(Z_1 \frac{[\mathbf{r}_1 \times \mathbf{p}]}{2m_e^2 r_1^3} + Z_2 \frac{[\mathbf{r}_2 \times \mathbf{p}]}{2m_e^2 r_2^3} \right) \mathbf{s}. \quad (5)$$

Both terms in Eq. (3) are divergent. In order to remove the infinities a transformation to the second order term can be applied which separates a divergent part

$$H'_B = H_B + (H_0 - E_0)U + U(H_0 - E_0),$$

$$\begin{aligned} \langle H_B Q(E_0 - H_0)^{-1} Q H_B \rangle \\ = \langle H'_B Q(E_0 - H_0)^{-1} Q H'_B \rangle + \langle U H_B + H_B U \rangle \\ - 2 \langle U \rangle \langle H_B \rangle + \langle U(H_0 - E_0)U \rangle \end{aligned} \quad (6)$$

with $U = \frac{1}{4m_e} [Z_1/r_1 + Z_2/r_2] = -\frac{1}{4m_e} V$.

The last three terms of the second expression in Eq. (6) can be recast in a form of a new effective interaction

$$\begin{aligned} H^{(6)} &= (U H_B + H_B U) - 2U \langle H_B \rangle - U(E_0 - H_0)U \\ &= \frac{p^4 V + V p^4}{32m_e^4} - \frac{\pi V [\rho_1 + \rho_2]}{4m_e^3} + \frac{(\mathcal{E}_1 + \mathcal{E}_2)^2}{32m_e^3} + \frac{V}{2m_e} \langle H_B \rangle. \end{aligned} \quad (7)$$

Taking into account that Ψ_0 is a solution of the Schrödinger equation $H_0 \Psi_0 = E_0 \Psi_0$, one may obtain from the above the following finite expression [16]

TABLE I. Relativistic corrections of $R_\infty \alpha^4$ order [in units $\alpha^4 \times (1 \text{ a.u.})$], H_2^+ .

	$v=0$	$v=1$	$v=2$	$v=3$	$v=4$
$L=0$	-0.042097	-0.042908	-0.043786	-0.044732	-0.045729
$L=1$	-0.042100	-0.042912	-0.043792	-0.044740	-0.045738
$L=2$	-0.042107	-0.042922	-0.043805	-0.044757	-0.045756
$L=3$	-0.042117	-0.042938	-0.043825	-0.044782	-0.045783
$L=4$	-0.042133	-0.042959	-0.043854	-0.044818	-0.045820

$$\begin{aligned} \Delta E^{(6)} &= \langle H'_B Q(E_0 - H_0)^{-1} Q H'_B \rangle + \langle H^{(6)} \rangle + \langle H^{(6)} \rangle \\ &= \langle H'_B Q(E_0 - H_0)^{-1} Q H'_B \rangle + \frac{3E_0 \langle V^2 \rangle}{4m_e^2} - \frac{5E_0^2 \langle V \rangle}{4m_e^2} \\ &\quad - \frac{3\pi E_0 \langle (\rho_1 + \rho_2) \rangle}{4m_e^3} + \frac{\langle \mathbf{p} V^2 \mathbf{p} \rangle}{8m_e^3} + \frac{\langle V \rangle \langle H_B \rangle}{2m_e} + \frac{E_0^3}{2m_e^2}. \end{aligned} \quad (8)$$

This new expression can be now calculated numerically, since all the terms are finite.

The “effective” potentials of $\Delta E^{(6)}(R)$ have been obtained for different bond lengths in Ref. [16]. Results are shown in Fig. 1.

Averaging them over the radial wave function of particular state one may get corresponding contribution to the energy of that state of order $R_\infty \alpha^4$. Results of numerical calculation of the relativistic corrections at this order are presented in Tables I and II. For the transition frequency this adiabatic approach provides about three significant digits.

III. HIGHER ORDER RADIATIVE CORRECTIONS

The electron ground-state wave function to a good extent may be approximated by $\psi_e(\mathbf{r}_e) = C[\psi_{1s}(\mathbf{r}_1) + \psi_{1s}(\mathbf{r}_2)]$, where ψ_{1s} is the hydrogen ground-state wave function. So, the most important $R_\infty \alpha^5$ order contributions can be evaluated using this approximate wave function and the expressions

$$\begin{aligned} E_{\text{se}}^{(5)} &= \alpha^5 \sum_{i=1,2} \left\{ \frac{Z_i^3}{m_e^2} \left[-\ln^2 \frac{1}{(Z_i \alpha)^2} + A_{61} \ln \frac{1}{(Z_i \alpha)^2} + A_{60} \right] \right. \\ &\quad \left. \times \langle \delta(\mathbf{r}_i) \rangle \right\}, \end{aligned}$$

TABLE II. Relativistic corrections of $R_\infty \alpha^4$ order [in units $\alpha^4 \times (1 \text{ a.u.})$], HD^+ .

	$v=0$	$v=1$	$v=2$	$v=3$	$v=4$
$L=0$	-0.042043	-0.042738	-0.043483	-0.044278	-0.045126
$L=1$	-0.042045	-0.042741	-0.043487	-0.044284	-0.045132
$L=2$	-0.042050	-0.042748	-0.043496	-0.044295	-0.045146
$L=3$	-0.042058	-0.042759	-0.043510	-0.044312	-0.045167
$L=4$	-0.042069	-0.042773	-0.043529	-0.044336	-0.045195

TABLE III. Summary of contributions to the $(v=0, L=0) \rightarrow (v'=1, L'=0)$ transition frequency (in MHz). The nonrelativistic energies are obtained with the CODATA02 [20] recommended values.

	H_2^+	HD^+
ΔE_{nr}	65 687 511.0686	57 349 439.9717
ΔE_{α^2}	1091.041(03)	958.152(03)
ΔE_{α^3}	-276.544(02)	-242.125(02)
ΔE_{α^4}	-1.997	-1.748
ΔE_{α^5}	0.120(23)	0.105(19)
ΔE_{tot}	65 688 323.688(25)	57 350 154.355(21)

$$E_{2\text{loop}}^{(5)} = \frac{\alpha^5}{\pi m_e^2} [B_{50}] \langle Z_1^2 \delta(\mathbf{r}_1) + Z_2^2 \delta(\mathbf{r}_2) \rangle, \quad (9)$$

where the constants A_{61} , A_{60} , and B_{50} are taken equal to the constants of the $1s$ state of the hydrogen atom $A_{61}=5.419 \dots$ [17], $A_{60}=-30.924 \dots$ [18], and $B_{50}=-21.556 \dots$ [19]. We note that the leading contribution ($R_\infty \alpha^5 \ln^2 \alpha$) is exact.

IV. RESULTS AND CONCLUSION

Various contributions to the frequency interval of the fundamental transition are summarized in Table III. Uncertainty in orders $R_\infty \alpha^2$ and $R_\infty \alpha^3$ are primarily due to numerical uncertainty in calculation of leading order terms such as $\langle \mathbf{p}^4 \rangle$ in the Breit-Pauli Hamiltonian or the Bethe logarithm $\beta(L, v)$ (see Refs. [21,22] for the details), and can be improved by more extensive calculations. We estimate uncertainty due to

finite size of nuclei as $\sim 3 \times 10^{-4}$ MHz for these transitions. So, the latter is so far negligible for the rovibrational spectroscopy. For the contribution of order $R_\infty \alpha^5$ the error bars are determined by the total contribution of the terms with coefficients A and B in Eq. (9).

Recently, the $(v, L): (0, 2) \rightarrow (4, 3)$ rovibrational transition for the HD^+ ion has been precisely measured in the experiment at the Düsseldorf university [23]. Comparison with the theoretical calculation demonstrates a very good agreement:

$$E_{\text{expt}} = 214\,978\,560.6(5) \text{ MHz},$$

$$E_{\text{theor}} = 214\,978\,560.88(7) \text{ MHz}.$$

In conclusion, the relativistic corrections of order $R_\infty \alpha^4$ allow us to reduce the relative accuracy of the fundamental transition frequency in H_2^+ to about 3×10^{-9} or 0.3 ppb. Further improvement we expect to achieve by numerical estimate of coefficients A_{61} , A_{60} , and B_{50} from Eq. (9) using the two-center adiabatic (or external field) approximation. That may reduce the final uncertainty by a factor of 5–10 and the relative uncertainty to less than 10^{-10} . Eventually, it will make real the main goal of our studies: improving of the m_p/m_e mass ratio from the rovibrational spectroscopy of H_2^+ and HD^+ .

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- [1] B. Grémaud, D. Delande, and N. Billy, *J. Phys. B* **31**, 383 (1998).
[2] S. Schiller and C. Lämmerzahl, *Phys. Rev. A* **68**, 053406 (2003).
[3] D. L. Farnham, R. S. Van Dyck, and P. B. Schwinberg, *Phys. Rev. Lett.* **75**, 3598 (1995); T. Beier, H. Häffner, N. Hermanspahn, S. G. Karshenboim, H.-J. Kluge, W. Quint, S. Stahl, J. Verdú, and G. Werth, *ibid.* **88**, 011603 (2001).
[4] L. Hilico, N. Billy, B. Grémaud, and D. Delande, *Eur. Phys. J. D* **12**, 449 (2000); J.-Ph. Karr, S. Kilic, and L. Hilico, *J. Phys. B* **38**, 853 (2005).
[5] V. I. Korobov, *Phys. Rev. A* **61**, 064503 (2000).
[6] D. H. Bailey and A. M. Frolov, *J. Phys. B* **35**, 4287 (2002).
[7] Z.-C. Yan, J.-Y. Zhang, and Y. Li, *Phys. Rev. A* **67**, 062504 (2003).
[8] M. M. Cassar and G. W. F. Drake, *J. Phys. B* **37**, 2485 (2004).
[9] S. Schiller and V. Korobov, *Phys. Rev. A* **71**, 032505 (2005).
[10] H. Li, J. Wu, B.-L. Zhou, J.-M. Zhu, and Z.-C. Yan, *Phys. Rev. A* **75**, 012504 (2007).
[11] V. I. Korobov, *Phys. Rev. A* **74**, 052506 (2006).
[12] V. I. Korobov, L. Hilico, and J.-Ph. Karr, *Phys. Rev. A* **74**, 040502(R) (2006).
[13] D. Bakalov, V. I. Korobov, and S. Schiller, *Phys. Rev. Lett.* **97**, 243001 (2006).
[14] J. R. Sapirstein and D. R. Yennie, in *Quantum Electrodynamics*, edited by T. Kinoshita (World Scientific, Singapore, 1990).
[15] M. I. Eides, H. Grotch, and V. A. Shelyuto, *Phys. Rep.* **342**, 63 (2001).
[16] V. I. Korobov and Ts. Tsogbayar, *J. Phys. B* **40**, 2661 (2007).
[17] A. J. Layzer, *Phys. Rev. Lett.* **4**, 580 (1960).
[18] K. Pachucki, *Ann. Phys. (N.Y.)* **226**, 1 (1993).
[19] K. Pachucki, *Phys. Rev. Lett.* **72**, 3154 (1994); M. I. Eides and V. A. Shelyuto, *Phys. Rev. A* **52**, 954 (1995).
[20] P. J. Mohr and B. N. Taylor, *Rev. Mod. Phys.* **77**, 1 (2005).
[21] V. I. Korobov, *Phys. Rev. A* **70**, 012505 (2004).
[22] V. I. Korobov, *Phys. Rev. A* **73**, 024502 (2006).
[23] B. Roth, J. C. J. Koelemeij, H. Daerr, and S. Schiller, *Phys. Rev. A* **74**, 040501(R) (2006); J. C. J. Koelemeij, B. Roth, A. Wicht, I. Ernsting, and S. Schiller, *Phys. Rev. Lett.* **98**, 173002 (2007).