Predictions of laser-cooling temperatures for multilevel atoms in three-dimensional polarization-gradient fields

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We analyze the dynamics of atom-laser interactions for atoms having multiple, closely spaced, excited-state hyperfine manifolds. The system is treated fully quantum mechanically, including the atom's center-of-mass degree of freedom, and motion is described in a polarization gradient field created by a three-dimensional laser configuration. We develop the master equation describing this system, and then specialize it to the lowintensity limit by adiabatically eliminating the excited states. We show how this master equation can be simulated using the Monte Carlo wave function technique, and we provide details on the implementation of this procedure. Monte Carlo calculations of steady state atomic momentum distributions for two fermionic alkaline earth isotopes, ^{25}Mg and ^{87}Sr , interacting with a three-dimensional lin- \perp -lin laser configuration are presented, providing estimates of experimentally achievable laser-cooling temperatures.

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I. INTRODUCTION

The complex behavior that occurs when a multilevel atom interacts with polarization-gradient fields has been of interest for some time now. Sub-Doppler cooling $[1]$ occurs because of elaborate optical-pumping processes produced by laser light in atoms with a sublevel structure, as seen, for example, in the lin- \perp -lin and the $\sigma_+ - \sigma_-$ laser configurations. The semiclassical understanding of these interactions $[2-7]$ in one or more dimensions has led to a reasonably good qualitative understanding of the underlying mechanisms. Semiclassical analysis has even in some cases provided quantitative predictions of sub-Doppler laser cooling temperatures measured in experiments [7].

However, the most direct route to a quantitative understanding of atom-laser interactions is via a fully quantized master equation for the atom, in which the center-of-mass (CM) motion of the atom is taken into account quantum mechanically. This allows behavior at low laser intensities and low atomic velocities, the regime laser cooling strives to reach, to be described correctly. The drawback of solving such a master equation, however, is the large number of basis states required for the calculation, due to the additional momentum states. This problem becomes especially pronounced when attempting to model three-dimensional (3D) systems, where the state space grows as the cube of the number of one-dimensional momentum states needed.

The Monte Carlo wave-function (MCWF) technique, introduced in the early 1990s, has allowed significant progress to be made on the subject of atom-photon interactions in 3D as well as lower-dimensional calculations. The MCWF technique is a simulation procedure for the master equation that involves the propagation of single stochastic wave functions, rather than density operators, with random processes occurring at random intervals due to interactions with the photon field that cause spontaneous emission. It has been shown that this method is equivalent to the master equation in the limit of a large number of independent stochastic wave functions

[8]. The MCWF technique has been successfully utilized to calculate 3D sub-Doppler laser cooling temperatures for atoms with Zeeman degeneracy in the ground and excited states $[9]$.

The majority of the research done on laser cooling has involved essentially two-level systems, consisting of a ground state and an excited state, which may or may not contain degenerate sublevels. However, some investigations have explored atomic systems in which multiple distinct excited states come into play. In particular, the use of bichromatic laser fields [10,11] to cool three-level Λ systems have been extensively studied (see Refs. [12-14], for example).

In this paper we focus primarily on monochromatic laser cooling for atoms with multiple closely spaced hyperfine excited-state manifolds. Figure 1 provides a graphical illustration of this type of atomic configuration. This situation is of importance, for example, in alkaline-earth atoms with a nonzero nuclear magnetic moment. If the excited state manifolds are spaced in energy on the order or smaller than the excited state linewidth γ , coherences between these mani-

FIG. 1. Energy level diagram of an atom with multiple hyperfine manifolds. If the energy spacing of the excited-state manifolds are of the order or smaller than the natural linewidth of the transition, the usual sub-Doppler cooling transition $(F_g \leftrightarrow F_e = F_g + 1)$ is not isolated and the other manifolds must be taken into account.

folds become non-negligible, and can have a significant effect on the optical pumping processes required for sub-Doppler cooling and on the dynamics of the atom-photon interaction. Sub-Doppler laser cooling was experimentally identified in fermionic 87 Sr [15], despite significant spectral overlap in the excited state. At the time, it was hypothesized that the large ground-state degeneracy in 87 Sr (due to the large nuclear spin $I=9/2$) was somehow able to overcome the decrease in cooling due to the spectral overlap. Other systems with spectral overlap in the excited state are $\rm{^{39}K}$ [16], 7 Li [17], and the fermionic isotopes of Yb [18]. In $87Rb$, the effects of excited-state spectral overlap on the effectiveness of velocity-selective coherent population trapping have been explored, both experimentally and theoretically $[19]$. Our goal in the paper is to provide a detailed discussion of the theoretical techniques required to model such systems realistically. In a future publication, we plan to present comprehensive laser-cooling predictions for a variety of atoms.

The structure of this paper is as follows. In Sec. II, we develop the master equation for a laser-driven atom with multiple excited-state manifolds, and then specialize this equation to the low-intensity limit. In Sec. III, we introduce the MCWF technique and apply it to this low-intensity master equation. In Sec. V, we perform full Monte Carlo masterequation simulations for ²⁵Mg and ⁸⁷Sr atoms in a 3D lin- \perp lin laser configuration as an example of using this technique determine expected temperatures for these atoms in a laser cooling experiment. In Sec. VI, we conclude.

II. MASTER EQUATION IN THE LOW-INTENSITY LIMIT

In this section we develop the master equation describing a multilevel atom interacting with a coherent laser field and coupled to a vacuum photon field. It is this equation, with quantized atomic CM, that will provide an accurate description of atom-photon dynamics, and this master equation will provide the basis for the Monte Carlo simulations that will be discussed later.

The full Hamiltonian for the atom-laser system plus the radiation field is

$$
H = H_A + H_R + V_{A-L} + V_{A-R},\tag{1}
$$

where $H_A = \sum_i \hbar \omega_i P_i + P^2 / 2m$ is the bare atomic Hamiltonian, H_R is the vacuum radiation field Hamiltonian, and V_{A-L} and V_{A-R} are the atom-laser and atom-radiation field coupling terms, respectively. In the atomic Hamiltonian, P_i is a projection operator onto the *i*th internal excited-state manifold, $\hbar \omega_i$ is the energy of the *i*th excited-state manifold relative to the ground-state manifold, *P* is the atomic CM momentum operator, *m* is the atomic mass, and the sum runs over all excited-state manifolds. We have assumed in Eq. (1) that the effects of atom-laser and atom-radiation-field coupling are independent $\lceil 20 \rceil$.

We can view Eq. (1) in terms of system-reservoir interactions. The *system* consists of the atom, the laser, and their interaction. The system Hamiltonian is

The *reservoir* is the vacuum radiation field, having many more modes than the system. With the Markov approximation, along with a few other approximations, the master equation is then given by

$$
\dot{\sigma} = \frac{i}{\hbar} [\sigma, H_S] + \mathcal{L}_{\rm sp}[\sigma]. \tag{3}
$$

The operator σ is the system reduced density operator element, i.e., the reservoir degrees of freedom have been traced over, $\sigma = \text{Tr}_R \rho$. The remaining term, $\mathcal{L}_{\text{sp}}[\sigma]$, encompasses the interaction between the atom and the vacuum photon field, and provides for the phenomenon of spontaneous emission.

The relaxation operator due to spontaneous emission, which we derive in detail in the Appendix, is given by

$$
\mathcal{L}_{\rm sp}[\sigma] = \frac{3\gamma}{8\pi} \int d^2\Omega \sum_{\epsilon \perp \mathbf{k}} \sum_{i,j} e^{-i\mathbf{k} \cdot \mathbf{R}} (\epsilon^* \cdot \mathbf{A}^{(i)}) \sigma(\epsilon \cdot \mathbf{A}^{(j)^\dagger}) e^{i\mathbf{k} \cdot \mathbf{R}}
$$

$$
- \frac{3\gamma}{16\pi} \int d^2\Omega \sum_{\epsilon \perp \mathbf{k}} \sum_{i,j} [(\epsilon \cdot \mathbf{A}^{(i)^\dagger}) e^{i\mathbf{k} \cdot \mathbf{R}} e^{-i\mathbf{k} \cdot \mathbf{R}}
$$

$$
\times (\epsilon^* \cdot \mathbf{A}^{(j)}) \sigma + \sigma(\epsilon \cdot \mathbf{A}^{(i)^\dagger}) e^{i\mathbf{k} \cdot \mathbf{R}} e^{-i\mathbf{k} \cdot \mathbf{R}} (\epsilon^* \cdot \mathbf{A}^{(j)})], \tag{4}
$$

where $\mathbf{A}^{(i)^\dagger}$ and $\mathbf{A}^{(i)}$ are vector raising and lowering operators, respectively, between the ground state and the *i*th excited state, **R** is the atomic CM position, **k** is the direction of the photon emitted in the relaxation process, and γ is the decay rate of the exited states. The integral is performed over a solid angle in the vector **k** and the sum over $\epsilon \perp \mathbf{k}$ refers to the two polarization directions perpendicular to **k**. Note that here and throughout this paper, we assume that each of the excited-state hyperfine manifolds has the same lifetime τ $=\gamma^{-1}$. Expanding these vector operators in a basis of spherical unit vectors, $\hat{\epsilon}_{\pm 1} = \pm (\hat{x} \pm i\hat{y})/\sqrt{2}$ and $\hat{\epsilon}_0 = \hat{z}$, we have

$$
\mathbf{A}^{(i)} = \sum_{q=0,\pm 1} (-1)^q \hat{\epsilon}_{-q} A_q^{(i)}.
$$
 (5)

The spherical components of the vector operators are

$$
A_q^{(i)} = \sum_{M_g, M_{e_i}} \alpha_{F_g, F_{e_i}, M_g, M_{e_i}, J_g, J_e, I} |J_g I F_g M_g \rangle \langle J_e I F_{e_i} M_{e_i} | \quad (6)
$$

$$
A_{q}^{(i)^{\dagger}} = \sum_{M_{g},M_{e_{i}}} \alpha_{F_{g},F_{e_{i}^{*}}M_{g},M_{e_{i}^{*}}J_{g},J_{e},I} | J_{e}IF_{e_{i}}M_{e_{i}}\rangle\langle J_{g}IF_{g}M_{g}|, \quad (7)
$$

where

$$
\alpha_{F_g, F_{e_i}, M_g, M_{e_i}, J_g, J_e, I}
$$
\n
$$
= (-1)^{F_g + F_{e_i} + M_g + J_e + I} \sqrt{(2F_g + 1)(2F_{e_i} + 1)(2J_e + 1)}
$$
\n
$$
\times \begin{pmatrix} F_g & 1 & F_{e_i} \\ -M_g & M_g - M_{e_i} & M_{e_i} \end{pmatrix} \begin{Bmatrix} J_g & F_g & I \\ F_{e_i} & J_e & 1 \end{Bmatrix}.
$$
\n(8)

Equation (4) is written in a way that makes it explicit that it is in the Lindblad form $[21-23]$. As we will see later, it is important for the relaxation operator to be of this form in order to make use of the MCWF technique. Because the complex exponentials in the second line cancel each other,

the remaining integral over a solid angle can be evaluated, whereby Eq. (4) can be equivalently written as $[8]$

$$
\mathcal{L}_{\rm sp}[\sigma] = \frac{3\gamma}{8\pi} \int d^2\Omega \sum_{\epsilon \perp \mathbf{k}} \sum_{i,j} e^{-i\mathbf{k} \cdot \mathbf{R}} (\epsilon^* \cdot \mathbf{A}^{(i)}) \sigma(\epsilon \cdot \mathbf{A}^{(j)^\dagger}) e^{i\mathbf{k} \cdot \mathbf{R}}
$$

$$
- \frac{\gamma}{2} \sum_{i} [P_{e_i} \sigma + \sigma P_{e_i}]. \tag{9}
$$

We will now examine that the atom-laser interaction term, which is given in the electric-dipole approximation by

$$
V_{A-L}(\mathbf{R},t) = -\mathbf{D} \cdot \mathbf{E}_L(\mathbf{R},t),\tag{10}
$$

where $\mathbf{E}_L(\mathbf{R},t)$ is the electric field of the laser and **D** is the electric dipole operator. As usual, we treat the laser as a classical field, since it is a densely populated mode of the electric field. We can write the laser electric field in terms of its positive and negative frequency components, $\mathbf{E}_L(\mathbf{R},t)$ $= \mathbf{E}_L^{(+)}(\mathbf{R})e^{-i\omega t} + \text{c.c.,}$ and then expand into spherical components,

$$
\mathbf{E}_{L}^{(+)}(\mathbf{R}) = \frac{E_0}{2} \sum_{q=0,\pm 1} (-1)^q a_q(\mathbf{R}) \hat{\epsilon}_{-q},
$$
 (11)

where E_0 is the electric-field amplitude and $a_q(\mathbf{R})$ are the expansion coefficients. Making the rotating-wave approximation, so that

$$
V_{A-L}(\mathbf{R},t) = -\mathbf{D}^{(+)} \cdot \mathbf{E}_L^{(+)}(\mathbf{R})e^{-i\omega t} - \mathbf{D}^{(-)} \cdot \mathbf{E}_L^{(-)}(\mathbf{R})e^{i\omega t},\tag{12}
$$

where $\mathbf{D}^{(+)} = \sum_i P_{e_i} \mathbf{D} P_g$ and $\mathbf{D}^{(-)} = \sum_i P_g \mathbf{D} P_{e_i}$, we find

$$
V_{A-L} = -\frac{\Omega}{2} \sum_{i} \mathcal{D}_i(\mathbf{R}) e^{-i\omega t} + \text{H.c.}
$$
 (13)

In the previous equation we have defined the atom-laser raising operator,

$$
\mathcal{D}_i^{\dagger}(\mathbf{R}) = \sum_{q=0,\pm 1} a_q(\mathbf{R}) A_q^{(i)^{\dagger}}, \tag{14}
$$

and lowering operator,

$$
\mathcal{D}_i(\mathbf{R}) = \sum_{q=0,\pm 1} a_q^*(\mathbf{R}) A_q^{(i)},\tag{15}
$$

and introduced the "invariant" Rabi frequency,

$$
\Omega = \frac{E_0 \langle J_e || D || J_g \rangle}{\sqrt{2J_e + 1}},\tag{16}
$$

where $\langle J_e || D || J_g \rangle$ is the reduced dipole matrix element between the ground and excited states. This form of a Rabi frequency, defined in terms of the reduced matrix element between the $J = J_g$ ground state and the $J = J_e$ excited state is convenient because, in general, Rabi frequencies for transitions to different excited-state manifolds will not be the same.

Next, we observe that the second term in Eq. (9) is comprised of excited-state projection operators both pre- and post-multiplying the system density operator. Thus, it is clear that this term can be absorbed into the free-evolution commutator term in Eq. (3), allowing the master equation to be equivalently described by Hamiltonian evolution determined by an effective Hamiltonian H_{eff} , plus a term that is commonly called a *jump term*, and that cannot be written in the form of a commutator with the system density operator. We thus have

$$
\dot{\sigma} = -\frac{i}{\hbar} (H_{\text{eff}} \sigma - \sigma H_{\text{eff}}^{\dagger}) \n+ \frac{3\gamma}{8\pi} \int d^{2} \Omega \sum_{\epsilon \perp \mathbf{k}} \sum_{i,j} e^{-i\mathbf{k} \cdot \mathbf{R}} (\boldsymbol{\epsilon}^{*} \cdot \mathbf{A}^{(i)}) \sigma (\boldsymbol{\epsilon} \cdot \mathbf{A}^{(j)^{\dagger}}) e^{i\mathbf{k} \cdot \mathbf{R}},
$$
\n(17)

where the effective Hamiltonian H_{eff} is given by

$$
H_{\text{eff}} = \frac{P^2}{2m} - \sum_{i} \hbar \left(\delta_i + i \frac{\gamma}{2} \right) P_{e_i} + V_{A \cdot L}, \tag{18}
$$

where V_{A-L} is as given in Eq. (13). In obtaining Eqs. (17) and (18), we have made the usual rotating-frame transformation, which removes the free-evolution atomic Bohr frequencies from the problem. The more relevant frequencies are instead the laser detunings $\delta_i = \omega - \omega_i$ from the *i*th excited-state hyperfine manifold. The master equation given in Eq. (17) is fully general, but has been written in a form that will facilitate setting up a stochastic wave function simulation using the MCWF technique described later.

We would like to now specialize the master equation just discussed to the limit of low laser intensity. Specifically, this limit is valid when the saturation parameter for the atom in the *i*th excited-state hyperfine manifold,

$$
s_i = \frac{\Omega^2 / 2}{\delta_i^2 + (\gamma / 2)^2},\tag{19}
$$

is small, which occurs when the laser intensity is small or the laser detuning from the atomic transition is large. In this limit, the excited states are said to adiabatically follow the ground states. The excited states can then be eliminated from the equations of motion, resulting in a master equation in terms of only the ground-state sub-density-matrix,

$$
\sigma_{gg} = P_g \sigma P_g. \tag{20}
$$

In this limit, the master equation becomes (see Sec. 8.3.3 of Ref. [24])

$$
\dot{\sigma}_{gg} = -\frac{i}{\hbar} (h_{\text{eff}} \sigma_{gg} - \sigma_{gg} h_{\text{eff}}^{\dagger}) + \int d^2 \Omega \sum_{\epsilon \perp \mathbf{k}} \sum_{i,j} [\epsilon^* \cdot \mathbf{B}^{(i)}]
$$

$$
\times (\mathbf{R}, \mathbf{k})] \sigma_{gg} [\epsilon \cdot \mathbf{B}^{(i)\dagger} (\mathbf{R}, \mathbf{k})]. \tag{21}
$$

The new effective Hamiltonian is given by

$$
h_{\text{eff}} = \frac{P^2}{2m} + \sum_{i} \frac{s_i}{2} \hbar \left(\delta_i - i \frac{\gamma}{2} \right) \mathcal{D}^{(i)}(\mathbf{R}) \mathcal{D}^{(i)^\dagger}(\mathbf{R}). \tag{22}
$$

The new decay raising and lowering operators are given by

$$
B_q^{(i)^\dagger}(\mathbf{R}, \mathbf{k}) = \sqrt{\frac{3s_i \gamma}{8\pi}} A_q^{(i)^\dagger} e^{i\mathbf{k} \cdot \mathbf{R}} \mathcal{D}^{(i)}(\mathbf{R}), \tag{23}
$$

and

$$
B_q^{(i)}(\mathbf{R}, \mathbf{k}) = \sqrt{\frac{3s_i \gamma}{8\pi}} A_q^{(i)} e^{-i\mathbf{k} \cdot \mathbf{R}} \mathcal{D}^{(i)^\dagger}(\mathbf{R}).
$$
 (24)

Note that this new lowering (raising) operator contains two components: a raising (lowering) operator $\mathcal{D}^{(i)^\dagger}(\mathbf{R})$ [$\mathcal{D}^{(i)}$ \times (**R**)] between the ground state and the *i*th excited-state manifold due to the atom-laser interaction, and a lowering (raising) operator $A_q^{(i)}$ $\left(A_q^{(i)}\right)^{\dagger}e^{i\mathbf{k}\cdot\mathbf{R}}$ $\left(A_q^{(i)}\right)$ $\int_{a}^{(i)} e^{-i\mathbf{k} \cdot \mathbf{R}}$ of type *q* corresponding to coupling with the reservoir photon field via a photon with polarization *q*. Thus, the jump operator in the lowintensity equations describes a transition cycle of the atom involving coupling to both the laser and the reservoir photon field. Note also that this new operator and the effective-Hamiltonian term in the equation of motion are both proportional to the saturation parameter s_i , the perturbation parameter.

III. THE MONTE CARLO WAVE-FUNCTION TECHNIQUE

The MCWF $[8,9,25-29]$ technique is a means of interpreting a system-reservoir master equation—which describes the evolution of a density operator for a system interacting with a large external reservoir—as the evolution of an ensemble of individual wave functions, each undergoing random *quantum jumps*. The free evolution of the stochastic wave functions is determined by the effective Hamiltonian that we found in the previous section. The nature of the quantum jumps is determined by the leftover term in the master equation, which cannot be absorbed into the free-evolution commutator. The components of this leftover term are often called quantum-jump operators.

In the following, we will deal primarily with the master equation in the low-intensity limit, as developed in the previous section, although the methods could just as easily be applied to the arbitrary-intensity master equation. The lowintensity limit, however, provides a reduction in the number of internal atomic states required in the calculation, and this will be beneficial for performing calculations later. Furthermore, since the lowest temperatures are achieved for low laser intensities, such a specialization does not hinder our ability to calculate lower bounds of temperature.

Having already expressed the master equation in a form involving an effective Hamiltonian and a jump term in the previous section, the application of the MCWF technique is rather straightforward along the lines developed in the literature (see, in particular, Ref. [8]). For a single stochastic wave function, the procedure is as follows. First, set the wave function to an initial value. Then, numerically propagate the wave function for a time step δt according to the effective Hamiltonian H_{eff} only, from an initial value $|\psi(t)\rangle$ to a final value $|\psi^{(1)}(t+\delta t)\rangle$,

$$
|\psi^{(1)}(t+\delta t)\rangle = \left(1 - \frac{iH_{\rm eff}\delta t}{\hbar}\right)|\psi(t)\rangle.
$$
 (25)

Restrictions on the size of δt are given such that the firstorder truncation of the time-evolution operator in Eq. (25) is approximately valid. We note that H_{eff} is non-Hermitian by construction, as a result of the absorbing parts of the relaxation operator into the original (Hermitian) bare system Hamiltonian. Because of this, propagation with *H*_{eff} will not conserve the norm of the wave function when propagated to $|\psi^{(1)}(t+\delta t)\rangle$. The time step δt of the propagation must be chosen so that $\delta p \ll 1$ in the inner product,

$$
\langle \psi^{(1)}(t + \delta t) | \psi^{(1)}(t + \delta t) \rangle = 1 - \delta p. \tag{26}
$$

The quantity δp is the loss of norm resulting from propagating with H_{eff} for a time step δt , and is found to be

$$
\delta p = \delta t \langle \psi(t) | \sum_{i} \mathbf{B}^{(i)^\dagger}(\mathbf{R}, \mathbf{k}) \cdot \mathbf{B}^{(i)}(\mathbf{R}, \mathbf{k}) | \psi(t) \rangle
$$

\n
$$
= \delta t \langle \psi(t) | \sum_{i} \sum_{q=0, \pm 1} B^{(i)^\dagger}_{q}(\mathbf{R}, \mathbf{k}) B^{(i)}_{q}(\mathbf{R}, \mathbf{k}) | \psi(t) \rangle
$$

\n
$$
= \sum_{i} \sum_{q=0, \pm 1} \delta p_{i,q}.
$$
 (27)

The total loss of norm has been decomposed into individual elements each corresponding to a particular type of interaction with the reservoir (i.e., the *q* value of the interaction, or the excited state i involved). These individual contributions are given by

$$
\delta p_{i,q} = \delta t \langle \psi(t) | B_q^{(i)^\dagger}(\mathbf{R}, \mathbf{k}) B_q^{(i)}(\mathbf{R}, \mathbf{k}) | \psi(t) \rangle. \tag{28}
$$

We see that the loss of norm due to a given type of interaction with the reservoir is determined by the quantummechanical expectation value of the product of jump operators of this type of interaction. The loss of norm δp can also be interpreted as the probability for a quantum jump to occur.

After the wave function has been propagated as described above, and the values of $\delta p_{i,q}$ calculated, it must then be determined whether or not a quantum jump occurred. This is achieved by generating a pseudorandom number on a computer and comparing it to the value of the total jump probability δp . If the random number is less than δp , a quantum jump occurred, and if it is greater, no quantum jump occurred. If a quantum jump does occur, the type of quantum jump must also be calculated by comparing the random number with the individual subprobabilities $\delta p_{i,q}$ in the same manner.

If a quantum jump of type *q*, *i* occurs, we must apply the quantum jump lowering operator B_q^0 $\mathbf{R}^{(i)}$ (**R**, **k**) to the wave function from the beginning of the time step,

$$
|\psi(t+\delta t)\rangle = \sqrt{\frac{\delta t}{\delta p_{i,q}}} B_q^{(i)}(\mathbf{R}, \mathbf{k}) |\psi(t)\rangle.
$$
 (29)

The square-root factor in front of the lowering operator is necessary for renormalization. If no quantum jump occurs, then we simply renormalize the wave function.

FIG. 2. (Color online) An example of a characteristic MCWF stochastic trajectory. Shown is the average of the atomic CM energy over 500 independent stochastic wave functions, as a function of time, for a two-level atom in a 1D standing-wave laser field. The energy is given in units of the recoil energy $E_r = \hbar^2 k^2 / 2m$, and time is given in units of the inverse recoil frequency $\omega_r^{-1} = \hbar / E_r$. All wave functions are initialized in the ground state of the atom and localized in momentum space with zero momentum. The steady state, wherein the system fluctuates around an average value, is seen to be achieved after a transient relaxation period. Error bars indicate the variance in the data at each given time for the ensemble of 500 stochastic wave functions. An estimate of the steady-state atomic CM energy is obtained by performing a time average over all wave functions for all times after the relaxation regime. The error bar of such an average will be smaller than the error bars in the figure, which apply only to the data for a given time.

The resulting wave function is then used as the starting point for propagation over the next time step, and the procedure is repeated.

A good approximation of the true system density matrix is achieved by combining the trajectories of a number of independently propagated stochastic wave functions, each trajectory having a unique sequence of pseudorandom numbers. A thorough discussion of the statistical issues involved with the MCWF technique can be found in Ref. [8].) Once a suitable ensemble of stochastic wave function trajectories has been obtained, an estimate of the true expectation value of an operator is found by taking the ensemble average of the expectation value of that operator with respect to the stochastic wave functions. For example, an estimate of the average kinetic energy at a time *t* for a system for which *N* independent stochastic wave functions have been calculated is given by

$$
\langle E \rangle(t) = \frac{1}{N} \sum_{i=1}^{N} \left\langle \psi_i(t) \left| \frac{P^2}{2m} \right| \psi_i(t) \right\rangle, \tag{30}
$$

where $\psi_i(t)$ is the *i*th stochastic wave function, given at time *t*.

Figure 2 demonstrates a simple example of the application

of the MCWF technique, wherein the average kinetic energy is calculated for a two-level atom interacting with a one-dimensional standing-wave field. For this calculation, we have used a Rabi frequency of $\Omega = \gamma/2$ and a detuning of $\delta = -\gamma/2$, where γ is the decay rate of the upper to the lower atomic state, and we have set $\gamma = 400E_r$, where $E_r = \hbar^2 k^2 / 2m$ is the recoil energy. The atomic kinetic energy, averaged over 500 stochastic wave functions, each initialized to zero momentum, is plotted as a function of time, with error bars indicating the error in the ensemble average for a given time. The separation of the transient relaxation period from the steady state is clear, the steady state regime being characterized by fluctuations in the average energy about a mean. This noise is due to the finite number of stochastic wave functions being used, and if a greater number of wave functions were used, the amplitude of the fluctuations would be decreased. In the limit of an infinite number of wave functions, the true density-matrix solution of the master equation would be obtained. An estimate of the steady-state kinetic energy is found by time averaging the calculated data over the entire steady-state regime. Since this is a larger ensemble than the set of wave functions for a single time, the error of such an average will be smaller than the error bars shown in the figure.

IV. CALCULATIONS FOR 25Mg AND 87Sr

Our purpose in this section is to illustrate the application of the theory developed up to this point to a complicated system. We wish to quantitatively study the dynamics of particular atoms interacting with 3D polarization-gradient laser fields. The balance of the frictional cooling forces along with the diffusion experienced by the atom due to spontaneous emission and its interaction with the laser leads to a steady-state momentum distribution that determines the temperature of a gas of such atoms. In particular, we will study here the cooling of the fermionic isotopes of two alkaline-earth atoms, ^{25}Mg (nuclear spin $I = 5/2$, ${}^{1}S_{0} - {}^{1}P_{1}$ width $\gamma/2\pi = 81$ MHz $\Delta \omega_{13}/2\pi = 46$ MHz and $\Delta \omega_{23} / 2\pi = 27$ MHz, where we have assumed a hyperfine quadrupole parameter $B = 16 \text{ MHz}$ [30]) and 8^{37} Sr (*I*=9/2, $^{1}S_{0}$ - $^{1}P_{1}$ width $\gamma/2\pi$ =32 MHz, hyperfine splittings $\Delta \omega_{13} / 2\pi = 43$ MHz and $\Delta \omega_{23} / 2\pi = -17$ MHz). These atoms, having a nonzero nuclear magnetic moment, have degenerate (assuming zero magnetic field) Zeeman sublevels. These sublevels allow for the mechanism of sub-Doppler cooling in an appropriate laser configuration. Both ^{25}Mg and 87Sr exhibit significant excited-state spectral overlap, with $\Delta \omega_{13} / \gamma = 0.57$, $\Delta \omega_{23} / \gamma = 0.33$, and $\Delta \omega_{13} / \gamma = 1.3$, $\Delta \omega_{23}$ / γ =−0.53, respectively. We consider the 3D lin- \perp -lin laser configuration, consisting of a pair of opposing beams along each Cartesian axis, in which each beam is linearly polarized orthogonal to its opposing beam. Furthermore, for this calculation, we set to zero the relative phases of the three sets of laser pairs.

Having a nuclear spin of $I = 5/2$, the ¹S₀ state of ²⁵Mg results in a hyperfine ground state with six sublevels. Use of the low-intensity master equation given in Eq. (17) allows us to consider only these six internal states of the atom, since

the excited states have been adiabatically eliminated in this regime. However, as noted in Ref. [9], a momentum grid extending to 20*k* in each direction with a spacing of *k* would yield a density matrix with $(6 \times 41^3)^2 \approx 2 \times 10^{11}$ elements. A direct solution of this master equation is not numerically feasible, even without considering the further increases in matrix size necessary to describe the master equation relaxation operator in Liouville space $[31]$. On the other hand, the MCWF method only requires numerical propagation of individual wave functions, which would be represented by vectors with $6 \times 41^3 \approx 4 \times 10^5$ elements. If the number of independent stochastic wave functions required us to achieve satisfactory convergence for the calculation of a particular property of the system is not unreasonably large, the MCWF method provides a distinct advantage over a direct master-equation solution.

We follow the procedure outlined in Sec. III, working in the low-intensity limit in order to reduce the number of internal atomic states in the calculation, which increases the efficiency of calculation. Since laser cooling is most effective at low laser intensities, this turns out to be a useful regime in which to work, with the additional benefit that lower temperatures require a smaller number of atomic CM momentum states in the calculation. We must determine the effective Hamiltonian as given in Eq. (22) and the jump operators as given in Eqs. (23) and (24) for each atom, and for the particular laser field being considered.

We consider here the lin- \perp -lin laser configuration in three dimensions, with the relative phases of the beams set to zero. The positive-frequency component of the electric field is

$$
\mathbf{E}_{L}(\mathbf{R},t) = \frac{E_{0}}{2} [\hat{y}e^{ikX} + \hat{z}e^{-ikX} + \hat{z}e^{ikY} + \hat{x}e^{-ikY} + \hat{x}e^{ikZ} + \hat{y}e^{-ikZ}]
$$

$$
= \frac{E_{0}}{2} \sum_{q=0,\pm 1} (-1)^{q} a_{q}(\mathbf{R}) \hat{\epsilon}_{-q},
$$
(31)

with spherical coefficients

$$
a_{+1}(\mathbf{R}) = -\frac{1}{\sqrt{2}}(e^{-ikY} + e^{ikZ} + ie^{ikX} + ie^{-ikZ}), \qquad (32)
$$

$$
a_{-1}(\mathbf{R}) = +\frac{1}{\sqrt{2}}(e^{-ikY} + e^{ikZ} - ie^{ikX} - ie^{-ikZ}), \qquad (33)
$$

$$
a_0(\mathbf{R}) = e^{-ikX} + e^{ikY}.
$$
 (34)

With these coefficients, along with parameters appropriate to the particular atom under consideration, the atom-laser raising and lowering operators given in Eqs. (14) and (15) can be constructed. With knowledge of the effective Hamiltonian and the raising and lowering operators, we can then proceed with the MCWF procedure, as outlined.

Our example entails propagating 20 stochastic wave functions each for three different values of the light-shift parameter, $\hbar |\delta_3| s_3 / (2E_{\text{rec}}) = 10$, 20, and 30, for both ²⁵Mg $(I=5/2)$ and ⁸⁷Sr $(I=9/2)$. We consider only $\delta_3 = -5\gamma$. As in Fig. 2, we calculate the stochastic trajectories of the ensemble average (i.e., averaged over the 20 wave functions) kinetic energy for each atom as a function of time. We con-

FIG. 3. (Color online) Results for calculated ensemble-average energies (rms momentum squared) for 25 Mg and 87 Sr, as a function of the light-shift parameter $\hbar |\delta_3| s_3 / (2E_{\text{rec}})$. For comparison, also shown are the calculated energies for atoms with isolated transitions, $J_e = J_g + 1$, with $J_g = 1, 2, 3$, and 4, with detuning $\delta = -5\gamma$. See the text for a discussion.

tinue this propagation until the transient regime has been passed for some time, and use the time average over the steady-state ensemble-average kinetic energy to provide an estimate of the total average kinetic energy and the final error. The results are shown in Fig. 3, along with the energies for atoms with an isolated cooling transition for comparison, $J_e = J_e + 1$ with $J_e = 1, 2, 3$, and 4, with detuning $\delta = -5\gamma$, as first calculated by Castin and Mølmer in Ref. [9]. From this cursory analysis, we can see that 25 Mg should exhibit a sharp rise in temperature with increasing laser intensity, while ${}^{87}Sr$ will cool to sub-Doppler temperatures, even for higher intensities, as has been noted experimentally $[15]$.

Detailed calculations of this sort, for realistic atoms, are quite computationally expensive. For example, a single data point for the Mg and Sr calculations presented here required on the order of 200 h wall time for a 20 processor parallel code, running on a cluster of 2.4 GHz Intel Zeon processors. There remains work to be done, improving the numerical efficiency of our initial codes. Our goal in this paper has been to present our method and some illustrative results; in a future publication we plan to expand upon these initial results using improved, faster codes and present comprehensive predictions of laser cooling temperatures for a variety of atoms.

V. CONCLUSIONS

In conclusion, we have provided a detailed description of the fully quantum-mechanical master equation that describes an atom with multiple internal internal structure interacting with a 3D polarization-gradient laser field. We have shown how the spontaneous-emission relaxation operator is generalized for atoms of this type. The MCWF technique has been

applied to these equations of motion, providing a more efficient means of performing calculations for these systems compared to a full solution of the master equation. A few example calculations have been presented to illustrate the application of this theory to atomic systems interacting with laser configurations commonly used in experiments. After making improvements in the efficiency of our codes, we intend to expand upon this work in a future publication and provide a comprehensive survey of laser cooling calculations for atoms with a multilevel internal structure.

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APPENDIX: RELAXATION OPERATOR FOR AN ATOM WITH MULTIPLE EXCITED-STATE HYPERFINE MANIFOLDS

In this appendix, we outline the major steps in deriving the spontaneous-emission relaxation operator for an atom with multiple hyperfine excited-state manifolds. Detailed derivations of this sort, but including only a single excited state manifold, exist elsewhere in the literature (see, for example, Ref. [32]). Our intent here is to highlight the steps important in generalizing the previous work to include coherences between other exited states. We will work within the framework of the theory of system-reservoir interactions and follow the notation of Ref. [32].

The total Hamiltonian for an atom coupled to a vacuum radiation field is then given by $H = H_A + H_R + V_{A-R}$, where H_A and H_R are the atom and reservoir bare Hamiltonian, respectively, and V_{A-R} is the atom-reservoir coupling and is given in the the electric-dipole approximation as

$$
V_{A-R} = -\mathbf{D} \cdot \mathbf{E}(\mathbf{R}) = -\sum_{q=0,\pm 1} (-1)^q D_q E_{-q}(\mathbf{R}).
$$
 (A1)

Here, **D** is the electric dipole operator for the atom and **E** is the electric-field operator for the photon field, and we have expanded the interaction into its spherical components. To simplify the formalism, we will begin by ignoring the atomic CM and setting the position coordinate to be the origin, $R=0$. At the end we will then generalize the equations to include the CM degree of freedom.

In general, the total density operator ρ evolves according to the Liouville equation, $\dot{\rho}(t) = (i/\hbar)[\rho(t), H]$. Making the usual assumptions involved in deriving the master equation $[20,21,32]$, we arrive at an equation of motion for the reduced density operator of the system $\sigma = \text{Tr}_R \rho$,

$$
\dot{\sigma}(t) = \frac{i}{\hbar} [\sigma(t), H_A] - \frac{1}{\hbar^2} \int_0^\infty d\tau \sum_q (-1)^q
$$

$$
\times \{g_q(\tau) [D_q e^{-iH_A \tau/\hbar} D_{-q} e^{iH_A \tau/\hbar} \sigma(t)]
$$

$$
- e^{-iH_A \tau/\hbar} D_q e^{iH_A \tau/\hbar} \sigma(t) D_{-q}] + \text{H.c.}.
$$
 (A2)

In the previous equation, $g_q(\tau)$ is the two-time

gg	ge_1	ge_2	ge_3	
e_1 g $ e_1e_1$		e_1e_2	e_1e_3	
e_2 g $ e_2e_1 $		e_2e_2	e_2e_3	
e_3 g e_3 e_1		e_3e_2	$\text{e}_{\text{3}}\text{e}_{\text{3}}$	

FIG. 4. The partitioning of the density operator for an atom with multiple coupled excited-state manifolds, each potentially having multiple substates.

correlation function of the reservoir and is defined as $g_q(\tau) = \text{Tr}_R[\sigma_R \tilde{E}_q(\tau) \tilde{E}_q(0)]$, where the variables with overtildes are operators in the interaction representation, $\tilde{E}_q(t) = e^{iH_Rt/\hbar} E_q e^{-iH_Rt/\hbar}$. We assume that the reservoir is initially a vacuum, so that $\sigma_R = |0\rangle\langle 0|$. From this we can see that $g_q(\tau) = \sum_{\nu} |\langle \nu | E_q | 0 \rangle|^2 e^{-i\omega_{\nu} t}$, where the kets and bras refer to reservoir states. Note that $g(\tau)^* = g(-\tau)$. The correlation time of the reservoir τ_C is defined such that $g(\tau) \to 0$ for $\tau \gg \tau_C$.

In addition to the above approximations, we will also make the *secular approximation*, which requires that the equation of motion for each density-matrix element $\dot{\sigma}_{ii}$ have only terms involving density-matrix elements σ_{kl} on the right-hand side, such that $|\omega_{ij} - \omega_{kl}| \le \gamma$, where $\omega_{ij} \equiv \omega_i - \omega_j$ and where γ is the order of magnitude of the systemreservoir coupling. In the following, we will consider a system with a ground state coupled to multiple excited states that are separated in energy of the order of or smaller than γ . Thus, the ground-excited energy splitting $|\omega_{ge}| \ge \gamma$ will be a nonsecular frequency, while $\omega_{e_i e_j} \sim \gamma$ will be a secular frequency.

The particular atomic system that we are considering consists of a ground state with electronic angular momentum $J = J_g = 0$ and an excited state with $J = J_g = 1$. The electronic angular momentum is coupled to the nuclear spin quantum number *I*, resulting in a ground state with total angular momentum $F_g=I$, and three excited states with ${F_{e_i}} = {I-1, I, I+1}$. These assumptions are made for concreteness, but we note that this derivation can be easily extended to arbitrary angular momentum schemes. It is useful to decompose the system density operator, as illustrated in Fig. 4,

$$
\dot{\sigma}(t) = \dot{\sigma}_{gg}(t) + \sum_{i,j} \dot{\sigma}_{e_i e_j}(t) + \sum_i \left[\dot{\sigma}_{e_i g}(t) + \dot{\sigma}_{g e_i}(t) \right], \tag{A3}
$$

where $\sigma_{ij}(t) = P_i \sigma(t) P_j$; P_i is a projection operator onto the *i*th hyperfine manifold, $P_i = \sum_{M_i} |JIFM_i\rangle\langle JIFM_i|$; and M_i is the substate label for the *i*th manifold. Two relations that will be useful in the following are

$$
\sum_{M_g,M_{e_i}} \langle M_g | D_q | M_{e_i} \rangle | M_g \rangle \langle M_{e_i} | = A_q^{(i)} \frac{\langle J_g || D || J_e \rangle}{\sqrt{2J_e + 1}}, \quad (A4)
$$

$$
\sum_{M_g, M_{e_i}} (-1)^q \langle M_{e_i} | D_{-q} | M_g \rangle |M_{e_i} \rangle \langle M_g | = A_q^{(i)^\dagger} \frac{\langle J_g || D || J_e \rangle^\dagger}{\sqrt{2J_e + 1}},
$$
\n(A5)

where A_q^0 $\binom{i}{q}^{\dagger}$ and $A_q^{(i)}$ $\alpha_a^{(i)}$ are the atomic raising and lowering operators defined in $Eq. (6)$, and where we have made use of symmetry properties of the three-*J* and six-*J* symbols [33].

We focus on the equation for the ground-state subdensity-operator $\sigma_{gg}(t)$ in Eq. (A3). Beginning by taking matrix elements of Eq. (A2) between ground-state sublevel kets, we proceed as usual by eliminating terms that violate energy conservation (e.g., photon emission coupled to atomic excitation), and absorbing interaction-induced energy shifts into the energies of the internal atomic levels. The resulting equation of motion is

$$
\dot{\sigma}_{gg}(t) = \frac{\omega_0^3}{3\epsilon_0(2\pi)^3 c^3 \hbar} \frac{|\langle J_g || D || J_e \rangle|^2}{2J_e + 1} \sum_q \sum_{i,j} A_q^{(i)} \sigma(t) A_q^{(j)^\dagger}
$$
\n
$$
\approx \gamma \sum_q \sum_{i,j} A_q^{(i)} \sigma(t) A_q^{(j)^\dagger}.
$$
\n(A6)

We have assumed that the energy splittings between the ground state and the various excited states are all approximately equal, and accordingly have defined $\omega_0 = \omega_{e_i} - \omega_g$ for $i=1,2,3$. Equivalently, we have assumed that the decay rate for all of the excited-state manifolds is approximately equal, and defined as $\gamma = \gamma_{J_{e_i} \to J_g}$ for *i* = , 1, 2, 3. Note that the double sum over excited-state manifolds in Eq. (A6) will clearly result in intermanifold coherence effects in the equations of motion.

Regarding the energy shifts of the internal atomic states that arise due to the interaction with the reservoir states, it is important to mention a subtle feature not found in the simpler case of degenerate isolated excited states. Such energy shifts occur in the form of divergent principal-part integrals of virtual transition amplitudes $[20,32]$. For degenerate isolated manifolds, these diverging terms can be shown to cancel each other in the equations of motion. However, for the case of multiple, nondegenerate manifolds, these terms no longer cancel exactly, and pathological divergences related to reservoir-dressed internal atomic energy splittings remain. A thorough exploration of these terms is outside the scope of this paper, and for the present purposes, we ignore such diverging terms and absorb interaction-induced energy splittings into the defined energy levels of the atoms.

Working in the same manner as for the ground-ground sub-density-operator, we can find the equations of motion for the excited-state sub-density-operators,

$$
\dot{\sigma}_{e_i e_j}(t) = -i\omega_{e_i e_j} P_{e_i} \sigma(t) P_{e_j} - \frac{\gamma}{2} \sum_q \sum_{k,l} P_{e_i} [A_q^{(k)}{}^{\dagger} A_q^{(l)} \sigma(t) + \sigma(t) A_q^{(k)}{}^{\dagger} A_q^{(l)}] P_{e_j},
$$
\n(A7)

where we have added a trivial summation index that will be useful later when combining the various sub-density-matrix decay terms. Similarly, the equations of motion for the optical-coherence subdensity operators are

$$
\dot{\sigma}_{e_{i}g}(t) = -i\omega_{e_{i}g}P_{e_{i}}\sigma(t)P_{g} - \frac{\gamma}{2} \sum_{q} \sum_{k,l} P_{e_{i}A}^{(k)^\dagger} A_{q}^{(l)} \sigma(t)P_{g},
$$
\n(A8)

and $\dot{\sigma}_{ge_i}(t) = \dot{\sigma}_{e_i g}^{\dagger}(t)$.

Using Eq. $(A3)$, we can construct the equation of motion due to spontaneous emission for the full density operator,

$$
\dot{\sigma}(t) = \frac{i}{\hbar} [\sigma(t), H_A] + \gamma \sum_{q} \sum_{i,j} A_q^{(i)} \sigma(t) A_q^{(j)^\dagger} - \frac{\gamma}{2} \sum_{q} \sum_{i,j} [A_q^{(i)^\dagger} A_q^{(j)} \sigma(t) (t) A_q^{(i)^\dagger} A_q^{(j)}].
$$
 (A9)

Defining the spontaneous emission relaxation operator,

$$
\mathcal{L}_{\text{sp}}[\sigma] = \gamma \sum_{q} \sum_{i,j} A_q^{(i)} \sigma(t) A_q^{(j)^\dagger} - \frac{\gamma}{2} \sum_{q} \sum_{i,j} \left[A_q^{(i)^\dagger} A_q^{(j)} \sigma(t) + \sigma(t) A_q^{(i)^\dagger} A_q^{(j)} \right],
$$
\n(A10)

we can write the equation of motion as

$$
\dot{\sigma}(t) = \frac{i}{\hbar} [\sigma(t), H_A] + \mathcal{L}_{\rm sp}[\sigma].
$$
 (A11)

Including the atomic CM dependence that we have been ignoring since the beginning amounts to adding an integral over momentum states in three dimensions that should have been included when we inserted atomic projection operators. With this addition, the full relaxation operator takes the form shown in Eq. (4) .

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