¹I. Freund, Opt. Commun. (to be published); a brief account of some aspects of this process is given here. ²Coherent three-quantum processes are discussed by (a) I.

Freund and B. F. Levine, Phys. Rev. Lett. 23, 854 (1969); Phys. Rev. Lett. 25, 1241 (1970); (b) P. Eisenberger and S. L. McCall, Phys. Rev. Lett. 26, 684 (1971); Phys. Rev. A 3, 1145 (1971); (c) S. S. Jha and J. W. F. Woo, Phys. Rev. B 5, 4210 (1972); J. W. F. Woo and S. S. Jha, Phys. Rev. B

6, 4081 (1972); S. S. Jha and J. W. F. Woo, Nuovo Cimento B 10, 229 (1972).

³Incoherent three-quantum processes are discussed by B. F. Levine and I. Freund, Opt. Commun. **3**, 197 (1971).

- ⁴Resonant three-quantum processes are discussed by I. Freund and B. F. Levine, Opt. Commun. **3**, 101 (1971).
- ⁵(a) G. Breit and E. Teller, Astrophys. J. **91**, 215 (1940); (b) J. Shapiro and G. Breit, Phys. Rev. **113**, 179 (1959); (c) L. Spitzer and J. L. Greenstein, Astrophys. J. **114**, 407 (1951); (d) W. Zernik, Phys. Rev. **132**, 320 (1963); Phys. Rev. **133**, A117 (1964); Phys. Rev. **135**, A51 (1964); (e) B. Honig and J. Jortner, J. Chem. Phys. **47**, 3698 (1967); (f) S. Klarsfeld, Phys. Lett. **30**, 382 (1969); (g) P. Braunlich and P. Lambropoulos, Phys. Rev. Lett. **25**, 135 (1970); Phys. Rev. Lett. **25**, 986 (1970); Phys. Rev. A **5**, 1013 (1972); (h) W. R. Johnson, Phys. Rev. Lett. **29**, 1123 (1972); (i) J. M. Wessner, D. K. Anderson, and R. T. Robiscoe, Phys. Rev. Lett. **29**, 1126
- (1972); (j) H. R. Reiss, Phys. Rev. A 6, 817 (1972); Phys.
- Rev. A 1, 803 (1970); Phys. Rev. D 4, 3533 (1971).

⁶(a) R. Novick, in *Physics of One- and Two-Electron Atoms*, edited by F. Bopp and H. Kleinpoppen (North-Holland, Amsterdam, 1969), pp. 296-325; (b) M. Lipeles, R. Novick, and N. Tolk, Phys. Rev. Lett. 15, 690 (1965); (c) C. J. Artura, N. Tolk, and R. Novick, Astrophys. J. Lett. 157, L181 (1969); (d) C. A. Kocher, J. E. Clendenin, and R. Novick, Phys. Rev. Lett. 29, 615 (1972); (e) M. H. Prior, Phys. Rev. Lett. 29, 615 (1972); (e) M. H. Prior, Phys. Rev. Lett. 29, 611 (1972); (f) R. C. Elton, L. J. Palumbo, and H. R. Griem, Phys. Rev. Lett. 20, 783 (1968); (g) R. Marrus and R. W. Schmieder, Phys. Rev. A 5, 1160 (1972), and references to earlier work by these authors contained herein.
⁷O. Bely, J. Phys. B 1, 718 (1968).
⁸M. Göppert-Mayer, Ann. Phys. (Leipzig) 9, 273 (1931).

⁹M. Ianuzzi, Phys. Rev. Lett. 13, 371 (1964).

¹⁰N. Bloembergen, Nonlinear Optics (Benjamin, New York, 1965), Chaps. 2-8.

¹¹W. L. Peticulas, R. Morris, and K. E. Rieckhoff, J. Chem. Phys. **42**, 4164 (1965).

¹²I. Freund (unpublished).

¹³P. A. M. Dirac, The Principles of Quantum Mechanics (Oxford U.P., London, 1958), Sec. 63.

¹⁴International Tables for X-Ray Crystallography (Kynoch, Birmingham, England, 1965), Vol. III, Sec. 3.2.

¹⁵R. W. James, *The Optical Principles of the Diffraction of* X-Rays (Cornell U.P., Ithaca, N.Y., 1965), Chap. IV.

¹⁶P. Lambroulos, Phys. Rev. 168, 1418 (1969).

PHYSICAL REVIEW A

VOLUME 7, NUMBER 6

JUNE 1973

Auger Spectra of Carbon and Argon Following Ionization by Equal-Velocity α Particles and Deuterons^{*}

R. L. Watson

Department of Chemistry and Cyclotron Institute, Texas A&M University, College Station, Texas 77843

L. H. Toburen

Battelle, Pacific Northwest Laboratories, Richland, Washington 99352 (Received 26 December 1972)

The spectra of carbon K- and argon L-shell Auger electrons resulting from ionization produced by 0.15-0.50-MeV/amu α particles and deuterons incident on a thin gas cell have been measured. Absolute ionization cross sections were obtained from the Auger-electron yields and, in addition, the yields obtained with equal-velocity projectiles were used to test the projectile z dependence of the ionization process. Intense satellite structure was observed in the Auger spectra and these features were attributed to Auger transitions from multiple vacancy states formed during the collision.

I. INTRODUCTION

Several measurements have been reported^{1,2} in which comparisons of x-ray yields resulting from equal-velocity α -particle and deuteron or proton bombardments have been used to test the theoretical prediction^{3,4} that K-shell ionization cross sections are directly proportional to the square of the projectile charge. Significant deviations from the expected dependence have been deduced on the assumption that the fluorescence yields (which relate x-ray production cross sections to ionization cross sections) are the same for a given target element, regardless of the type of projectile being used to produce the ionization. It has recently been discovered, however, that even in collisions involving relatively light ions, such as α particles and protons, appreciable *L*-shell ionization frequently accompanies the production of a *K*-shell vacancy.⁵⁻⁷ Since the fluorescence yields for multiply ionized states are expected to differ from those for singly ionized states, the assumption that the fluorescence yield is independent of the exciting particle could lead to erroneous results, especially for light elements where the effects of multiple ionization may result in a relatively large change in fluorescence yield.⁸ Furthermore, x-ray yield measurements are subject to errors resulting from the critical absorption of x rays emitted from multiply ionized atoms in which the binding-energy increases are sufficient to shift the x rays above an absorption edge.^{6,8}

The above difficulties relating to the use of measured x-ray yields to deduce ionization cross sections are not encountered in converting measured Auger-electron yields to ionization cross sections for low-Z elements where fluorescence yields are very small. The present work was undertaken primarily for the purpose of testing the projectile Z dependence of K-shell ionization cross sections for carbon and *L*-shell ionization cross sections for argon by comparing the Auger yields obtained in bombardments with equal-velocity α particles and deuterons. Absolute ionization cross sections were also measured and are presented herein. Additional information relating to multiple ionization by α particles and deuterons was obtained from an analysis of the satellite structure appearing in the measured Auger spectra.

II. EXPERIMENTAL METHOD

 α -particle and deuteron beams with energies of 0.150-0.500 MeV/amu were obtained by means of a Van de Graaff generator. Since the rf ion source used in the Van de Graaff accelerator has

a very low efficiency for the production of He⁺⁺ ions, the α -particle beam was produced by the gas stripping of an accelerated He⁺ beam. The accelerated He⁺ beam entered a gas-stripping cell after magnetic identification, and a second analyzing magnet was used following the stripping cell to separate the α particles from other products of charge transfer. This method of magnetic separation with intermediate stripping is also effective in removing unwanted beam contamination, such as H_2^+ ions, which might otherwise be associated with the α -particles beam. The gas-stripping cell used in this work was approximately 7 m long and was operated at a pressure of about 2×10^{-5} torr. These conditions produced optimum stripping of the primary beam with minimal loss in beam intensity owing to gas scattering.

The deuteron beam was produced by accelerating D⁺ and using the same magnetic analysis described above for α particles. The gas-stripping region was evacuated to approximately 1×10^{-6} torr when D⁺ beams were used. Energy analysis of electrons ejected from D⁺ collisions with the target gas under study showed no evidence of electrons associated with H₂⁺ breakup. This result indicated that H₂⁺ contamination of the D⁺ beam, if any, was negligible.

The basic experimental technique for the measurement of electron-emission cross sections has



FIG. 1. Carbon K-shell Auger-electron spectra for deuterons and α particles incident on C_2H_6 . The dashed lines labeled A, A', and A'' indicate the expected positions of the normal 1s-2p2p, 1s-2s2p, and 1s-2s2s Auger transitions, respectively. The other dashed lines indicate the calculated positions of some of the possible satellite transitions listed in Table III.

been described in detail previously.⁹ The incident beam is collimated and directed through a differentially pumped target-gas cell. Electrons ejected by ion-atom or ion-molecule interactions pass through a slit in the target cell, are energy analyzed by a cylindrical-mirror electrostatic analyzer, and are detected by a continuous channel electron multiplier. The energy resolution of the electrostatic analyzer used in the present work was 3.5% full width at half-maximum (FWHM). The energy calibration of the electrostatic analyzer was accomplished by comparison of the Auger spectra of ethane, argon, and neon with those measured by Moddeman, ¹⁰ Carlson and Krause, ¹¹ and Edwards and Rudd, ¹² respectively. The energy scale used in our work should be accurate to better than 1.0%. Research-grade target gases were used with impurities of less than 4 parts in 10^4 .

III. DATA ANALYSIS AND RESULTS

A. Ionization Cross Sections

The spectra obtained in the present measurements are shown in Figs. 1 and 2. Auger yields were determined from the measured spectra by integrating the areas under the Auger-electron peaks.

Before these peak integrations were carried out, backgrounds based on the electron continua on each side of the Auger peaks were subtracted. After correcting the Auger yields for solid angle, electron absorption, and analyzer transmission (as discussed in Ref. 9), ionization cross sections were computed on the basis of isotropic emission of Auger electrons. K-shell Auger-electron emission resulting from K-shell vacancy production by proton impact has previously been shown to be isotropic.⁹ Nonisotropic angular distributions are theoretically possible for Auger transitions following the creation of a vacancy having quantum number $j > \frac{1}{2}$, ^{13,14} and a small anisotropy has, in fact, been observed by Cleff and Mehlhorn¹⁵ for the $L_{3}M_{2,3}M_{2,3}$ ('S₀) transition resulting from electron impact ionization of argon. Volz and Rudd,¹⁶ however, have measured angular distributions for several L_3MM transitions in argon excited by 300keV protons, and did not find any significant deviations from isotropy. As has already been noted, radiative transitions fill less than 0.5% of the primary carbon K-shell and argon L-shell vacancies, and hence correction for this competitive mode of decay is not required.

The cross sections obtained in the present study



FIG. 2. L-shell Auger-electron spectra for deuterons and α particles incident on argon. The dashed lines labeled A, A', and A'' indicate the expected positions of the normal 2p-3p3p, 2p-3s3p, and 2p-3s3s Auger transitions, respectively.

7

TABLE I. Cross sections for carbon K-shell and argon L-shell ionization by α particles and deuterons (in units of 10^{-18} cm²).^a

E/M	E/M Carbon (K shell)		Argon (L shell)	
(MeV/amu)	α particles	deuterons	α particles	deuterons
0.150	1.57	0.35		
0.250	3.34	0.66	7.61	1.96
0.375	5.20	1.05	12.5	2.73
0.500	4.96	1.10	15.0	3.18

^aAbsolute errors are estimated to be $\pm 20\%$.

for carbon K-shell ionization and argon L-shell ionization by equal-velocity α particles and deuterons are listed in Table I. The uncertainties in the absolute values are estimated to be $\pm 20\%$.

Presented in Table II are the ratios $N_{\alpha}/4N_{d}$, where N_{α} and N_{d} are the Auger yields obtained with equal-velocity α particles and deuterons, respectively. Because the Auger yields for α particles and deuterons were measured under nearly identical conditions, these yield ratios are subject to far fewer uncertainties than are the absolute cross sections.

Since the number of particles passing through the target was determined from the integrated current produced in a Faraday cup, any projectiles which were not fully stripped of their electrons would cause the measured Auger cross sections to appear larger than they actually are. The switching magnet located approximately 3 m upstream from the target chamber ensured that the projectiles were fully stripped at that point, but, during their flight from the switching magnet to the Faraday cup, charge-changing collisions could occur with residual gas molecules in the beam line and target-gas molecules in the target chamber. Calculations based upon charge-changing cross sections given by Allison¹⁷ indicated that this effect was important only for the 600-keV α -particle measurements, where it was found that under the conditions of the present experiment approximately 15% of the He^{**} beam ended up as He^* in the Faraday cup. As a result of this correction (which has been included in the values listed in Tables I and II), the 150-keV/amu Auger-electron yield ratio is considerably more uncertain than the rest.

B. Satellite Structure

An examination of the carbon K-shell Auger spectra shown in Fig. 1 reveals various structural changes in the gross features which appear to depend both on the type and energy of the projectile. In particular, the main peak is observed to be shifted about 4 eV down in energy in the spectra obtained with α particles as compared to the spectra obtained with deuterons. The differences in the structural features of the Auger spectra for the two types of particles become increasingly apparent as the bombarding energy decreases.

In comparing the features of the argon L-shell Auger spectra shown in Fig. 2, considerably more pronounced differences in structure are found. Especially striking is the abrupt reduction of the intensity of the high-energy peak in going from deuterons to α particles.

Effects similar to those described above have also been observed recently by Stolterfoht¹⁸ in Kshell Auger spectra obtained by proton impact on N_2 , CO_2 , and CH_4 . He attributes the changes in Auger structure to satellite peaks associated with double-vacancy production during the collision. This interpretation is consistent with evidence provided by x-ray measurements^{5,19} which show that simultaneous K- plus L-shell ionization does indeed occur with substantial probability in light-ion collisions with intermediate- and low-Z atoms.

Referring back to Figs. 1 and 2, the normal transitions (i. e., the usual transitions associated with an initial *single*-vacancy configuration) are labeled with the letters A, A', and A''. These letters correspond to the 1s - 2p2p, 1s - 2s2p, and 1s - 2s2s transitions, respectively, for carbon and to the 2p - 3p3p, 2p - 3s3p, and 2p - 3s3s transitions, respectively, for argon. The energies of the normal Auger transitions were taken from Moddeman¹⁰ and Siegbahn *et al.*²⁰ in the case of carbon (CH₄ and C₂H₆), and from Carlson and Krause¹¹ in the case of argon.

In an effort to identify the various structural features of the measured Auger spectra as satellites associated with specific Auger transitions, Hartree-Fock-Slater calculations were carried out with the program of Herman and Skillman.²¹ Energies of the *KLL* Auger transitions for carbon atoms and the *LMM* transitions for argon atoms, both in various multiply ionized initial configurations, were obtained from the differences in calculated total energies for initial and final states. The absolute transition energies calculated for free carbon atoms were not expected to accurately represent the measured Auger satellites for molecular C_2H_6 . The energy *differences* between the normal transitions and their corresponding multi-

TABLE II. Auger-electron yield ratios for equal-velocity α particles and deuterons incident on carbon and argon.

E/M	$N\alpha/4Nd$		
(MeV/amu)	Carbon (K shell)	Argon (L shell)	
0.150	1.13 ± 0.12	1	
0.250	1.25 ± 0.07	0.96 ± 0.05	
0.375	1.24 ± 0.07	1.14 ± 0.06	
0.500	$\textbf{1.13} \pm \textbf{0.06}$	$\textbf{1.18} \pm \textbf{0.06}$	

		Energy (eV)		
		Measured ^a	Calcul	lated ^b
	Transition	E	ΔE	E
A	$1s-2p2p(1s^{-1})$	250	0	250
в	$(1s^{-1}, 2s^{-1})$		-10	240
С	$(1s^{-1}, 2s^{-2})$		-16	234
	$(1s^{-2})$		22	272
A'	$1_{s-2s2p}(1_{s}^{-1})$	237	0	237
	$(1s^{-1}, 2s^{-1})$		- 9	228
	$(1s^{-1}, 2p^{-1})$		- 9	228
	$(1s^{-1}, 2s^{-1}, 2p^{-1})$		-14	223
A″	$1s - 2s 2s (1s^{-1})$	230	0	230
	$(1s^{-1}, 2p^{-1})$		- 9	221
	$(1s^{-1}, 2p^{-2})$		-13	217

TABLE III. Auger energies for *KLL* transitions in multiply ionized carbon atoms.

^aMeasured energies are for CH₄ from Refs. 10 and 20. ^b ΔE is calculated energy difference between satellite transition and corresponding normal transition. *E* is transition energy obtained by adding ΔE to the measured energy of the appropriate normal transition.

ple-ionization satellites, however, should be nearly the same for C_2H_6 molecules as for free carbon atoms. The satellite energies were estimated by subtracting the calculated energy differences from the measured energies of the appropriate normal transitions. The results of the calculations for carbon are summarized in Table III. In this table the initial-state configuration is identified by indicating the number of vacancies in each level be-

TABLE IV. Auger energies for *LMM* transitions in multiply ionized argon atoms.

		Energy (eV)		
		Measured ^a	Calculated ^b	
	Transition	E	ΔE	E
A	$2p - 3p 3p (2p^{-1})$	206	0	206
в	$(2p^{-1}, 3p^{-1})$		- 9	197
С	$(2p^{-1}, 3p^{-2})$		- 20	186
	$(2p^{-1}, 3p^{-3})$		- 32	174
	$(2p^{-2})$		10	216
Α′	$2p-3s3p(2p^{-1})$	191	0	191
D	$(2p^{-1}, 3p^{-1})$		-12	179
Е	$(2p^{-1}, 3p^{-2})$	*	- 23	168
	$(2p^{-1}, 3p^{-3})$		- 35	156
A"	$2p-3s3s(2p^{-1})$	180	0	180
F	$(2p^{-1}, 3p^{-1})$		-12	168
G	$(2p^{-1}, 3p^{-2})$		-24	156
	$(2p^{-1}, 3p^{-3})$		- 36	144

^aMeasured energies from Ref. 11.

^b ΔE is the calculated energy difference between the satellite transition and the corresponding normal transition. *E* is transition energy obtained by adding ΔE to the measured energy of the appropriate normal transition. yond those in the ground-state configuration. The letters preceding some of the transitions in Table I are used to identify the positions of these lines in Fig. 1.

The absolute transition energies calculated for the normal LMM transitions in argon were within 1 or 2 eV of the measured energies as given by Carlson and Krause.¹¹ The results of the calculations for the argon satellite transitions are summarized in Table IV. Additional analysis of the argon spectra was accomplished by fitting a series of Gaussian peaks having widths determined by the experimental spectrometer resolution to each spectrum. This procedure enabled a fairly accurate determination of the area of the peak corresponding to the normal 2p-3p3p transition (peak A in Fig. 2). Using this peak area and the intensities of the other "normal" transitions relative to the intensity of the 2p-3p3p transition as obtained from Carlson and Krause,¹¹ the normal Auger spectrum was subtracted from each of the measured spectra. The result of this analysis is shown in Fig. 3 for the Auger spectrum obtained with 2-MeV α particles. The letters in this figure indicate the expected locations of a few of the satellite transitions listed in Table IV. In this manner the total yields of Auger electrons associated with satellite transitions could be estimated. The ratios of the residual Auger yields (after subtrac-



FIG. 3. Total and residual *L*-shell Auger-electron spectra for 2.0-MeV α particles incident on argon. The dashed lines indicate the calculated positions of some of the possible satellite transitions listed in Table IV.

E/M	$N_{R'}$	N_T
(MeV/amu)	Deuterons	α particles
0.250	0.40	0.83
0.375	0.38	0.74
0.500	0.36	0.67

TABLE V. Residual-to-total L-shell Auger-electron yield ratios for argon.

1858

tion of the normal spectra) to the total Auger yields are given in Table V.

IV. DISCUSSION

A. Ionization Cross Sections

The K-shell ionization cross sections for carbon obtained in the present work are compared with the results of other thin-target measurements in Fig. 4. They are plotted as a function of the velocity ratio v_p/\overline{v}_K , where v_p is the projectile velocity and \overline{v}_K is the average K-shell electron velocity $[\overline{v}_K = (2U_K/m)^{1/2}$ and U_K is the K-shell bind-

ing energy]. The cross sections given by Harrison *et al.*²² were deduced from x-ray measurements for protons on CH₄. Those given by Toburen^{23,24} and by Stolterfoht¹⁸ were derived from Auger-electron measurements. Toburen's values are averages for protons incident on CH₄, C₂H₆, and C₂H₂, whereas Stolterfoht's values are for protons incident on CH₄. Although different molecular environments affect the Auger yields to some extent, a previous study of carbon Auger yields obtained for different molecular forms of carbon showed the magnitude of this effect to be of the order of 5% or less for CH₄ and C₂H₆.²⁴

The cross sections for deuterons incident on C_2H_6 obtained in the present work overlap quite well with the proton values obtained by the other investigators except at higher velocity ratios, where the present results along with the previous results of Toburen are about 20% higher than those of Stolterfoht. The cross sections obtained with α particles (divided by z^2) are systematically higher than the cross sections for deuterons throughout the range of velocity ratios measured.



FIG. 4. Comparison of carbon K-shell ionization cross sections measured for thin targets with the predictions of various theories. The K-shell ionization cross section divided by the square of the projectile charge is plotted vs the projectile to K-shell electron velocity ratio where $\bar{v}_K = (2u_K/m)^{1/2}$ (u_K is the K-shell binding energy, 283 eV).

The solid and dashed curves in Fig. 4 show the predictions of the Born-approximation theory from tabulations given by Khandelwal et al.²⁵ and the predictions of the classical binary-encounter theory as presented by Garcia, ²⁶ respectively. Similar comparisons have already been discussed in some detail by the above-mentioned experimenters, and so we restrict ourselves here to only a few brief observations regarding the over-all trends. In general, both theoretical approaches appear to give satisfactory agreement with the experimental results for protons above a velocity ratio of approximately 0.5. For 0.5 < $v_p/\overline{v}_K < 2.0$ it is not clear from the experimental data which of the two theoretical descriptions gives the best results, but in any case the differences between them are only about 25% at most in this velocity range.

Below $v_{\rm p}/\bar{v}_{\rm K} \simeq 0.5$, both theories appear to overestimate the cross sections, with the binary-encounter theory overestimating by the smaller amount. Brandt et al.²⁷ have attributed the discrepancies between experimental cross sections and those calculated in the Born approximation at low velocity ratios to an increased binding of the K-shell electrons resulting from the close proximity of the projectile during the collision. Using the prescription given by Brandt, ²⁸ we have corrected the Born-approximation cross sections for this binding effect and the results are shown by the dot-dashed curve in Fig. 4. (The Coulomb correction also discussed by Brandt²⁸ is negligible over this range of velocity ratios.) It is seen that this curve falls considerably below the experimental points at high velocity ratios, but it joins the

1859



FIG. 5. Comparison of the argon *L*-shell ionization cross sections with the predictions of the binary-encounter and Born-approximation theories. The *L*-shell ionization cross section divided by the square of the projectile charge is plotted vs the projectile to *L*-shell electron velocity ratio where $\overline{v}_L = (2\overline{u}_L/m)^{1/2} (\overline{u}_L$ is the average *L*-shell binding energy, 264 eV).

experimental points of Harrison *et al.*²² at low velocity ratios. This underestimation of experimental results at the higher velocity ratios has been noted by Brandt,²⁸ and he and his co-workers¹ have developed a possible explanation for this in terms of a polarization effect.

The results of the L-shell ionization cross-section measurements for argon are shown in Fig. 5 along with the predictions of the Born-approximation and binary-encounter theories. It is apparent that fairly large discrepancies exist between the experimental and theoretical L-shell ionization cross sections at low velocity ratios. The experimental results appear to be converging with the Born-approximation curve at high velocity ratios.

B. Projectile Z Dependence

As was stated in Sec. I, one of the principal objectives of this study was to test the projectile Zdependence in the light-element region using Augerelectron yields. By comparing Auger yields instead of x-ray yields, as has been done in all previous studies, ^{1,2} uncertainties associated with the fluorescence yields are eliminated. A comparison of the Auger yields for carbon and argon in the form of a plot of the cross-section ratios $\sigma_{\alpha}/4\sigma_{d}$ as a function of the projectile energy per amu is shown in Fig. 6. If the predicted z^2 dependence were to be upheld, the experimental points would lie along the dashed line at $\sigma_{\alpha}/4\sigma_d = 1$. Rather large deviations are observed from a z^2 dependence for both K-shell ionization of carbon and L-shell ionization of argon.

To be consistent with the results of the abovementioned x-ray measurements, the z^2 deviation should exhibit a crossover behavior in which the ratios are smaller than unity at low projectile velocities, rise above unity at higher velocities, and, finally, decrease back toward unity at the highest velocities. The present results do indeed follow this trend. The theory of Basbas *et al.*¹ contends that the crossover phenomenon occurs as a result of the exact cancellation of the binding effect (which acts to decrease the ionization cross section) and the polarization effect (which acts to increase the ionization cross section). This cancellation is expected to occur for *n*-shell ionization when

$$1/q_0 a_n \simeq 1 , \qquad (1)$$

where $\hbar q_0$ is the minimum momentum transfer for *n*-shell ionization given by

$$\hbar q_0 = U_n / v_p , \qquad (2)$$

and a_n is the Bohr radius of the *n* shell (corrected for inner-electron screening):

$$a_n = n^2 \hbar^2 / m Z_n e^2 . \tag{3}$$

In the above equations, U_n , v_p , m, and Z_n are the *n*-shell binding energy, projectile velocity, electron mass, and screened nuclear charge, respectively. Equation (3) can be rewritten in terms of the Bohr velocity of an *n*-shell electron (corrected for inner-electron screening) v_n to give

$$a_n = n^3 v_n \hbar / 2Z_n^2 R , \qquad (4)$$

where R is the Rydberg constant. Combining Eqs. (1), (2), and (4) results in the following expression for the crossover-point condition:

$$v_{p}/v_{n} = \frac{1}{2}n(n^{2}U_{n}/Z_{n}^{2}R) = \frac{1}{2}n\theta_{n}$$
 (5)

The crossover-point condition expressed in terms of the velocity ratio used throughout this paper



FIG. 6. Ionization crosssection ratio plotted vs the projectile energy per amu for carbon K shell and argon L shell. The behavior of this ratio for an exact z^2 dependence is shown by the dashed line.

 $[v_p/\overline{v}_n \text{ where } \overline{v}_n \equiv (2U_n/m)^{1/2}]$ is obtained by noting that

$$v_n / \overline{v}_n = \frac{(2Z_n^2 R / n^2 m)^{1/2}}{(2U_n / m)^{1/2}}$$
$$= \theta_n^{-1/2} \quad . \tag{6}$$

Substituting this expression into Eq. (5), the final result is

$$v_p/\overline{v}_n = \frac{1}{2} n \theta_n^{1/2} . \tag{7}$$

The crossover point given by Eq. (7) occurs at $v_p/\overline{v}_R = 0.401$ for carbon (E/M = 83 keV/amu) and at $v_p/\overline{v}_L = 0.636$ (E/M = 194 keV/amu) for argon. Unfortunately, the large uncertainty in the 150-keV/amu point for carbon does not permit a conclusive experimental determination of the cross-over velocity ratio. Based upon the present data, about all one can say in this regard is that a cross-over should occur somewhere around $v_p/\overline{v}_K = 0.44$ (E/M = 100 keV/amu). In the case of argon, it is noted that the observed crossover velocity ratio is 0.75 (E/M = 270 keV/amu), which is 18% larger than the velocity ratio given by Eq. (7).

In summary, we have shown that large deviations from a z^2 dependence occur for carbon K-shell ionization and argon L-shell ionization, and that these deviations are not a result of variations in fluorescence yields. Although the present results are unable to accurately establish the crossover velocity ratio for carbon, it is apparent that the general trend of the projectile z dependence of carbon Kshell ionization cross sections is similar to that observed for argon L-shell ionization and to the trends deduced previously in K x-ray measurements with Cl, Ca, Ti, and Ni targets.²

C. Satellite Structure

In Sec. III B it was noted that evidence of satellite structure associated with Auger transitions from atoms which had undergone multiple ionization during the collision process appears in the measured carbon *KLL* and argon *LMM* Auger-electron spectra. In this section we concern ourselves with identifying the origins of the various structural features seen in these Auger spectra.

Carbon. Unfortunately, in the case of the KLL Auger spectra for carbon (Fig. 1), the broad widths of the various lines preclude very conclusive statements regarding satellite identities on the basis of our data alone. These linewidths are not solely a consequence of our experimental resolution but are characteristic of C_2H_6 , even in very-high-resolution spectra.²⁰ In comparing our spectra for C_2H_6 with Stolterfoht's spectra for CH_4 , however, it becomes apparent that the shift in the main peak energy in going from deuterons to α particles results from a substantial increase in the intensity of the satellite peak appearing at an energy of approximately 243 eV (peak B') in Stolterfoht's spectra. Referring back to the calculated satellite peak energies in Table III, the most likely candidate appears to be the 1s - 2p 2p satellite transition associated with an initial double-vacancy-state configuration of $1s^{-1}2s^{-1}$ and having a calculated energy of 241 eV. The 1s-2p2p satellite transition for an initial vacancy-state configuration $1s^{-1}2p^{-1}$ would not be possible for a free carbon atom since only two 2p electrons are initially available. However, since we are dealing with carbon atoms bound in a molecule (C_2H_6), in effect, six 2p electrons are available and a 1s - 2p 2p satellite transition for an initial-state configuration of $1s^{-1}2p^{-1}$ would be expected to occur. In such a case this satellite line would occur at approximately the same energy as the one originating from a $1s^{-1}2s^{-1}$ initial configuration and may, in fact, be the dominant transition contributing to the observed satellite peak. The estimated position of the transition associated with an initial $1s^{-1}2s^{-1}$ or $1s^{-1}2p^{-1}$ vacancy is indicated by the dashed line labeled B in Fig. 1 and, as can be seen in the spectra obtained with 1.0- and 0.6-MeV α particles, coincides nicely with the observed structure.

Also, plainly visible in the Auger spectrum obtained with 0.6-MeV α particles (in Fig. 1) is a peak at approximately 234 eV. The energy of this peak corresponds to the energy calculated for the 1s-2p2p satellite transition associated with an initial triple-vacancy-state configuration of $1s^{-1}2s^{-2}$ (or $1s^{-1}2p^{-2}$). The position of this transition is indicated by the dashed line labeled C in Fig. 1.

We conclude, then, that the dominant features contributing to our Auger spectra for carbon (besides the "normal" Auger transitions) are associated with 1s-2p2p transitions arising from initial states having one and two vacancies in the 2s or 2plevel in addition to a 1s vacancy. The relative intensities of these satellite transitions increase both as the projectile charge increases and as the projectile velocity decreases.

Argon. The residual Auger-electron spectrum (after subtraction of the "normal" spectrum) for 2.0-MeV α particles on argon (Fig. 3) shows significant structural features at 195, 183, 168, and 152 eV. Comparing these energies with the calculated energies for argon *LMM* satellites in Table IV, it is found that these energies correspond most closely to 2p-3p3p, 2p-3s3p, and 2p-3s3s satellite transitions associated with initial-state configurations having one and two vacancies in 3s or 3p levels in addition to a 2p vacancy. (Although Table IV only gives transition energies for 3p vacancies, they are the same within 2 eV for 3s vacancies.) The calculated positions of these transitions are indicated by the dashed lines in Fig. 3. In this figure it is seen that the calculated energy for transition G is somewhat higher than the structural feature appearing at 152 eV but, for the most part, the other indicated satellite energies coincide fairly well with the residual Auger structure.

If one makes the assumption that multiple-vacancy production in light-ion collisions is simply a result of simultaneous collisions with individual electrons during the ions' passage through the atom then one is led to expect that the cross section for the production of multiple vacancies is proportional to the product of the cross sections for producing each vacancy individually.⁵ Thus if σ_n is the total cross section for *n*-shell ionization (irrespective of what other ionization is produced simultaneously) and if σ_n , is defined similarly for the n' shell, then it follows that the total cross section for simultaneously producing at least one *n*-shell and one n'shell vacancy is

$$\sigma_{nn'} \simeq c \sigma_n \sigma_{n'} , \qquad (8)$$

where c is a proportionality constant, or

$$\frac{\sigma_{nn'}}{\sigma_n} \propto \sigma_{n'} , \qquad (9)$$

The above result has been tested by Knudson, Burkhalter, and Nagel⁵ for simultaneous K- plus L-shell ionization produced by protons incident on Al, and by Li, Watson, and Hansen¹⁹ for simultaneous K- plus L-shell ionization produced by deuterons and α particles incident on Ca, Ti, and Fe. The former group found that the projectile velocity dependence of the K- plus L-shell cross-section ratio was correctly given by the L-shell ionization cross-section velocity dependence as calculated in the Born approximation. The latter group compared the same ratio with the L-shell crosssection velocity dependence predicted by the classical theory of Gryzinski⁴ and found similar agreement.

*Work supported by the U.S. Atomic Energy Commission. ¹George Basbas, Werner Brandt, Roman Laubert, Anthony Ratowski, and Arthur Schwarzschild, Phys. Rev. Lett. 27, 171 (1971); G. Basbas, W. Brandt, and R. Laubert, Phys. Lett. 34, 277 (1971).

- ²C. W. Lewis, R. L. Watson, and J. B. Natowitz, Phys. Rev. A 5, 1773 (1972).
- ³E. Merzbacher and H. W. Lewis, in Encyclopedia of Physics, edited by S. Flügge (Springer, Berlin, 1957), Vol. 34, p. 166. ⁴Michal Gryziński, Phys. Rev. 138, A336 (1964).

⁵A. R. Knudson, P. G. Burkhalter, and D. J. Nagel, in Pro-

ceedings of the International Conference on Inner-Shell Ionization Phenomena, Atlanta, 1972, Conf.-720404 (USAEC, Oak Ridge, Tenn., 1973), Vol. 3, p. 1675.

- ⁶J. S. Hansen, T. K. Li, and R. L. Watson, in Ref. 5, p. 1682.
- ⁷C. F. Moore, M. Senglaub, B. Johnson, and P. Richard, Phys. Lett. 40, 107 (1972).
- ⁸L. H. Toburen and F. P. Larkins, Phys. Rev. A 6, 2035 (1972).

In the present case of argon, Eq. (9) predicts that the *L*-plus *M*-shell to *L*-shell ionization cross-section ratio should follow the velocity dependence of the M-shell cross section. The ratios of the residual-to-total Auger yields given in Table V are estimates of the above cross-section ratio. We have compared the projectile velocity dependence of these ratios with the *M*-shell ionization cross-section velocity dependence as predicted by the Gryzinski model and find satisfactory agreement. It should be noted that this simple picture of multiple ionization predicts that σ_{LM}/σ_L should approximately follow a z^2 dependence on projectile charge. In Table V it is seen that the present results for deuterons and α particles only differ by about a factor of 2 instead of the factor of 4 required by a z^2 dependence.

In summary then, we have shown that most of the dominant features of the LMM Auger-electron spectra resulting from light-ion collisions with argon atoms are explainable in terms of satellite transitions associated with initial states having one or more vacancies in 3s or 3p levels. This conclusion was reached on the basis of (a) the correspondence between the observed structure and the calculated satellite transition energies, and (b) the agreement between the observed velocity dependence of the residual Auger yield with that predicted by the simple independent-collision picture of multiple ionization.

ACKNOWLEDGMENTS

The authors would like to thank L. L. Nichols, W. A. Glass, J. E. Choate, and C. A. Ratcliffe for their help in overcoming the many problems encountered in obtaining α particles with the positive-ion Van de Graaff accelerator. One of us (RLW) gratefully acknowledges travel support provided by the Northwest College and University Association for Science.

⁹L. H. Toburen, Phys. Rev. A 3, 216 (1971).

¹⁰W. I. Moddeman, Oak Ridge National Laboratory Report

- 17, 1079 (1966). ¹²A. K. Edwards and M. E. Rudd, Phys. Rev. 170, 140 (1968).
 - ¹³W. Mehlhorn, Phys. Lett. 26, 166 (1968).
- ¹⁴S. Flügge, W. Mehlhorn, and V. Scmidt, Phys. Rev. Lett. **29,** 7 (1972).
 - ¹⁵B. Cleff and W. Mehlhorn, Phys. Lett. 37, 3 (1971).
 - ¹⁶D. J. Volz and M. E. Rudd, Phys. Rev. A 2, 1395 (1970).
 - ¹⁷S. K. Allison, Rev. Mod. Phys. 30, 1137 (1958).
 - ¹⁸N. Stolterfoht, in Ref. 5, Vol. 2, p. 1043.
- ¹⁹T. K. Li, R. L. Watson, and J. S. Hansen, Phys. Rev. A (to be published). ²⁰K. Siegbahn, C. Nordling, G. Johansson, J. Hedman, P. F.
- Heden, K. Hamrin, U. Gelius, T. Bergmark, L. O. Werme, R. Manne, and Y. Baer, ESCA Applied to Free Molecules (North-Holland, Amsterdam, 1969), p. 103.

No. ORNL-TM-3012 (1970), p. 125 (unpublished). ¹¹T. A. Carlson and M. O. Krause, Phys. Rev. Lett.

²¹F. Herman and S. Skillman, Atomic Structure Calculations (Prentice-Hall, Englewood Cliffs, N.J., 1963).

 22 K. G. Harrison, H. Tamara, and F. J. deHeer, Phys. Rev. A (to be published).

²³L. H. Toburen, in Ref. 5, Vol. 2, p. 979.

²⁴L. H. Toburen, Phys. Rev. A 5, 2482 (1972).

PHYSICAL REVIEW A

VOLUME 7, NUMBER 6

(1966).

At. Data 1, 103 (1969).

JUNE 1973

Error Bounds on Dynamic Polarizability and Oscillator Strength*

P. T. Gee and K. T. Tang

Department of Physics, Pacific Lutheran University, Tacoma, Washington 98447 (Received 26 January 1973)

Two uniformly convergent series are introduced to give bounded estimates of the dynamic polarizability in the region below as well as above the first resonance line. Upper and lower bounds on the oscillator strength can also be obtained from these approximations. The series are simple and rapidly convergent.

The frequency-dependent (dynamic) polarizability $\alpha(\omega)$, being important in many physical and chemical considerations, has been subjected to intensive study in recent years.^{1,2} Here we wish to introduce two approximations for $\alpha(\omega)$. One is uniformly convergent to it from above and the other from below. They have the advantages of being simple and rapidly convergent. Furthermore, the upper and lower bounds of the oscillator strength can also be obtained from these approximations.

The dynamic polarizability is given by the Kramers-Heisenberg dispersion formula

$$\alpha(\omega) = \sum_{i} \frac{f_{i}}{\epsilon_{i}^{2} - \omega^{2}} , \qquad (1)$$

where f_i is the electric dipole oscillator strength for the transition from the ground state to state *i* with excitation energy ϵ_i . The summation sign is understood to include integration over the continuum and atomic units are used for all quantities. Expanding Eq. (1) in a Taylor series about the origin, we obtain the Cauchy equation

$$\alpha(\omega) = \sum_{j=0}^{\infty} \mu_j \omega^{2j} , \qquad (2)$$

with the Cauchy moments μ_j given by the oscillator-strength sums S(k) defined as

$$\mu_{j} = S(-2j-2) = \sum_{i} \frac{f_{i}}{\epsilon_{i}^{2j+2}} \quad . \tag{3}$$

The radius of convergence is ϵ_1 . The first few Cauchy moments can be determined either experimentally or theoretically. To approximate Eq. (1) with finite number of Cauchy moments, we introduce the following:

$$\alpha_{a}^{n}(\omega) = \frac{1}{\epsilon_{1}^{2} - \omega^{2}} \sum_{k=0}^{n} a_{2k} \omega^{2k} , \qquad (4)$$

²⁵G. S. Khandelwal, B. H. Choi, and E. Merzbacher,

²⁷W. Brandt, R. Laubert, and I. Sellin, Phys. Rev. 151, 56

²⁶J. D. Garcia, Phys. Rev. A 1, 280 (1970).

²⁸W. Brandt, in Ref. 5, Vol. 2, p. 948.

with

$$a_{2k} = \begin{cases} S(-2)\epsilon_1^2, & k=0\\ S(-2k-2)\epsilon_1^2 - S(-2k), & k\neq 0 \end{cases}$$
(5)

and

$$\alpha_{b}^{n}(\omega) = \frac{1}{(\epsilon_{1}^{2} - \omega^{2})(\epsilon_{2}^{2} - \omega^{2})} \sum_{k=0}^{n>0} b_{2k} \omega^{2k} , \qquad (6)$$

with

$$b_{2k} = \begin{cases} S(-2)\epsilon_1^2 \epsilon_2^2, & k = 0 \\ S(-4)\epsilon_1^2 \epsilon_2^2 - (\epsilon_1^2 + \epsilon_2^2)S(-2), & k = 1 \\ S(-2k-2)\epsilon_1^2 \epsilon_2^2 - (\epsilon_1^2 + \epsilon_2^2)S(-2k) \\ + S(-2k+2), & k > 1 \end{cases}$$
(7)

To study their relationship with $\alpha(\omega)$, we examine the difference functions³

$$g_a^n(\omega) = \alpha_a^n(\omega) - \alpha(\omega)$$
(8)

and

$$g_b^n(\omega) = \alpha_b^n(\omega) - \alpha(\omega) .$$
(9)

Substituting Eqs. (4) and (1) into Eq. (8) and using Eqs. (5) and (3) to simplify the expression,⁴ we find

$$g_a^n(\omega) = \frac{\omega^{2n+2}}{\epsilon_1^2 - \omega^2} \sum_i \frac{(\epsilon_i^2 - \epsilon_1^2)f_i}{\epsilon_i^{2n+2}(\epsilon_i^2 - \omega^2)} \quad . \tag{10}$$

Similarly, from Eq. (9), we have

2 2

$$g_{b}^{n}(\omega) = -\frac{\omega^{2n+2}}{(\epsilon_{1}^{2} - \omega^{2})(\epsilon_{2}^{2} - \omega^{2})} \times \sum_{i} \frac{(\epsilon_{i}^{2} - \epsilon_{1}^{2})(\epsilon_{i}^{2} - \epsilon_{2}^{2})f_{i}}{\epsilon_{i}^{2n+2}(\epsilon_{i}^{2} - \omega^{2})} \quad . \quad (11)$$