

Solution of the relativistic Dirac-Woods-Saxon problem

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The Dirac equation is written for the special case of a spinor in a relativistic potential with the even and odd components related by a constraint, and solved exactly with the even component chosen to be the Woods-Saxon potential. The corresponding radial wave functions for the two-component spinor are obtained in terms of the hypergeometric function, and the energy spectrum of the bound states is obtained as a solution to a given equation with boundary constraints in which the nonrelativistic limit reproduces the usual Woods-Saxon potential.

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The Woods-Saxon potential plays an essential role in microscopic fields since it can be used to describe the interaction of a neutron with a heavy nucleus. Although the nonrelativistic Schrödinger equation with this potential has been solved for S states [1] and the single-particle motion in atomic nuclei has been explained quite well, the relativistic effects for a particle under the action of this potential are more important, especially for a strong-coupling system. The relativistic Coulomb and oscillator potential problems, including their bound-state spectra and wave functions, have already been established for a long time [2–9], and their nonrelativistic limits reproduce the usual Schrödinger-Coulomb and Schrödinger-oscillator solutions, respectively. Recently, Kennedy proposed a two-component approach to solving the one-dimensional Dirac equation, and obtained the scattering and bound-state solutions for the Woods-Saxon potential [10]. However, we are still not aware of any solutions of the three-dimensional Dirac equation with the Woods-Saxon potential, which may be more important in the field of nuclear structure. Fortunately, Alhaidari has just put forward an approach to the three-dimensional Dirac equation [11], and solved a class of shape-invariant potentials that includes Dirac-Rosen-Morse, Dirac-Eckart, Dirac-Pöschl-Teller, and Dirac-Scarff potentials [12], and presented the relativistic bound-state spectra and spinor wave functions. Following the procedure used in [11], we make an attempt, in this paper, to solve the three-dimensional Woods-Saxon potential problem. Suppose a spinor particle populates a relativistic potential with the even and odd components $V(r)$ and $W(r)$, and the even component is chosen to be the spherically symmetric Woods-Saxon potential. We solve the Dirac equation by employing the same strategy as that used by Alhaidari in the Dirac-Morse problem, namely, by adding a radial term to the odd part of the Dirac operator, and applying a unitary transformation to the Dirac equation such that the resulting second-order differential equation becomes Schrödinger-like, and solvable.

For simplicity, atomic units ($m = e = \hbar = 1$) are used, and the corresponding speed of light $c = \alpha^{-1}$ is taken. It is well known that the Hamiltonian for a charged Dirac particle in a four-component electromagnetic potential (A_0, \vec{A}) can be written as

$$H = \begin{pmatrix} 1 + \alpha A_0 & -i\alpha\vec{\sigma} \cdot \vec{\nabla} + \alpha\vec{\sigma} \cdot \vec{A} \\ -i\alpha\vec{\sigma} \cdot \vec{\nabla} + \alpha\vec{\sigma} \cdot \vec{A} & -1 + \alpha A_0 \end{pmatrix}, \quad (1)$$

where α is the fine structure constant and $\vec{\sigma}$ are the three 2×2 Pauli spin matrices. Replacing the two off-diagonal terms $\alpha\vec{\sigma} \cdot \vec{A}$ in the Hamiltonian of Eq. (1) by $\pm i\alpha\vec{\sigma} \cdot \vec{A}$ and writing (A_0, \vec{A}) as $(\alpha V(r), \hat{r}W(r))$, we obtain the following two-component radial Dirac equation:

$$\begin{pmatrix} 1 + \alpha^2 V(r) & \alpha \left[\frac{\kappa}{r} + W(r) - \frac{d}{dr} \right] \\ \alpha \left[\frac{\kappa}{r} + W(r) + \frac{d}{dr} \right] & -1 + \alpha^2 V(r) \end{pmatrix} = \varepsilon \begin{pmatrix} g(r) \\ f(r) \end{pmatrix}, \quad (2)$$

where $V(r)$ and $W(r)$ are the even and odd components of the relativistic potential, κ is the coupling parameter defined as $\kappa = \pm(j + \frac{1}{2})$ for $l = j \pm \frac{1}{2}$, and ε is the relativistic energy. This equation gives two coupled first-order differential equations for the two radial spinor components. By eliminating the lower component we obtain a second-order differential equation for the upper. Unfortunately, the resulting equation may turn out to be not Schrödinger-like, i.e., it may contain first-order derivatives. However, a general local unitary transformation may be applied to eliminate the first-order derivative as follows:

$$r = q(x)$$

and

$$\begin{pmatrix} g(r) \\ f(r) \end{pmatrix} = \begin{pmatrix} \cos[\rho(x)] & \sin[\rho(x)] \\ -\sin[\rho(x)] & \cos[\rho(x)] \end{pmatrix} \begin{pmatrix} \phi(x) \\ \theta(x) \end{pmatrix}. \quad (3)$$

The stated requirement gives the following constraint:

$$\frac{dq}{dx} \left[-\alpha^2 V + \cos(2\rho) + \alpha \sin(2\rho)(W + \kappa/q) + \alpha \frac{d\rho/dx}{dq/dx} + \varepsilon \right] = \text{const} \equiv \eta \neq 0. \quad (4)$$

This transformation and the resulting constraint are the relativistic analog of a point canonical transformation in nonrelativistic quantum mechanics. Here, we consider the case of a global unitary transformation defined by $q(x) = x$ and $d\rho/dx = 0$. Substituting these in the constraint Eq. (4) yields

$$W(r) = \frac{\alpha}{S} V(r) - \frac{\kappa}{r},$$

$$\eta = C + \varepsilon, \tag{5}$$

where $S = \sin(2\rho)$ and $C = \cos(2\rho)$. This maps the radial Dirac equation (2) into the following:

$$\begin{pmatrix} C + 2\alpha^2 V & \alpha \left[-\frac{S}{\alpha} + \frac{\alpha C}{S} V - \frac{d}{dr} \right] \\ \alpha \left[-\frac{S}{\alpha} + \frac{\alpha C}{S} V + \frac{d}{dr} \right] & -C \end{pmatrix} \times \begin{pmatrix} \phi(r) \\ \theta(r) \end{pmatrix} = \varepsilon \begin{pmatrix} \phi(r) \\ \theta(r) \end{pmatrix}, \tag{6}$$

which in turn gives an equation for the lower spinor component in terms of the upper:

$$\theta(r) = \frac{\alpha}{C + \varepsilon} \left[-\frac{S}{\alpha} + \frac{\alpha C}{S} V + \frac{d}{dr} \right] \phi(r), \tag{7}$$

while the differential equation for the upper component reads

$$\left[-\frac{d^2}{dr^2} + \frac{\alpha^2}{T^2} V^2 + 2\varepsilon V - \frac{\alpha}{T} \frac{dV}{dr} - \frac{\varepsilon^2 - 1}{\alpha^2} \right] \phi(r) = 0, \tag{8}$$

where $T = S/C = \tan(2\rho)$.

Having given this brief review of the method proposed by Alhaidari [11,12], let us consider the Dirac-Woods-Saxon potential problem. In Eq. (8), the independent even potential component $V(r)$ is chosen to be the spherically symmetric Woods-Saxon potential:

$$V(r) = -\frac{V_0}{1 + e^{(r-R)/b}} \quad (V_0 > 0) \tag{9}$$

with $b \ll R$ (R is the nuclear radius, and b is the range of the potential). Consequently, we obtain the following second-order differential equation for the upper spinor component:

$$\left\{ -\frac{d^2}{dr^2} - \frac{\lambda}{b^2} \frac{-\lambda + e^{(r-R)/b}}{[1 + e^{(r-R)/b}]^2} - \frac{2\varepsilon V_0}{1 + e^{(r-R)/b}} - \frac{\varepsilon^2 - 1}{\alpha^2} \right\} \phi(r) = 0, \tag{10}$$

where $\lambda = \alpha b V_0 / T$ is a dimensionless parameter. In order to reproduce the nonrelativistic Woods-Saxon problem, we take the nonrelativistic limit ($\alpha \rightarrow 0, \varepsilon \approx 1 + \alpha^2 E$), where E is the nonrelativistic energy. The above equation is then reduced to the following wave equation:

$$\left\{ -\frac{d^2}{dr^2} - \frac{\lambda}{b^2} \frac{-\lambda + e^{(r-R)/b}}{[1 + e^{(r-R)/b}]^2} - \frac{2V_0}{1 + e^{(r-R)/b}} - 2E \right\} \phi(r) = 0, \tag{11}$$

which is in the form of the Woods-Saxon problem only for $\lambda = -1$. In that case, Eq. (11) becomes the Schrödinger equation for the potential $V(r) = -V_0^{\text{NR}} [1 + e^{(r-R)/b}]^{-1}$, where $V_0^{\text{NR}} = V_0 - 1/2b^2$. The energy spectrum for the S -wave non-

relativistic Woods-Saxon problem is already known (see, for example, Ref. [1]). In order to obtain the relativistic energy spectrum easily and directly, we rewrite the relativistic Eq. (10), for $\lambda = -1$, as

$$\left[-\frac{d^2}{dr^2} - 2 \frac{\varepsilon V_0 - 1/2b^2}{1 + e^{(r-R)/b}} - \frac{\varepsilon^2 - 1}{\alpha^2} \right] \phi(r) = 0. \tag{12}$$

Comparing this with the nonrelativistic Eq. (11) for $\lambda = -1$, we arrive at the following parameter correspondence map:

$$\begin{aligned} b &\rightarrow b, \quad R \rightarrow R, \\ V^{\text{NR}} &\rightarrow \varepsilon V_0 - 1/2b^2, \\ E &\rightarrow (\varepsilon^2 - 1)/2\alpha^2. \end{aligned} \tag{13}$$

Using this parameter map we obtain immediately the relativistic energy spectrum

$$\begin{aligned} \arg \Gamma(2i\gamma) - 2 \arg \Gamma(\mu + i\gamma) - \tan^{-1} \frac{\gamma}{\mu} + \frac{\gamma R}{b} \\ = (n + 1/2)\pi, \quad n = 0, \pm 1, \pm 2, \dots, \end{aligned} \tag{14}$$

and the upper spinor component $\phi(r)$

$$\begin{aligned} \phi_n(r) = e^{i\gamma(r-R)/b} [1 + e^{(r-R)/b}]^{-\mu - i\gamma} F(\mu + i\gamma, \mu + i\gamma \\ + 1, 2\mu + 1; [1 + e^{(r-R)/b}]^{-1}) = 0, \end{aligned} \tag{15}$$

where

$$\mu = b \sqrt{(1 - \varepsilon^2)/\alpha^2}$$

and

$$\gamma = \sqrt{(\varepsilon^2 - 1)b^2/\alpha^2 + 2b^2\varepsilon V_0 - 1}.$$

The lower spinor component can also be derived from Eq. (7) as

$$\begin{aligned} \theta_n(r) = \frac{\alpha}{C + \varepsilon_n} \left\{ \left[-\frac{S}{\alpha} - \frac{\mu}{b} + \left(\frac{\mu + i\gamma}{b} - \frac{V_0 \alpha}{T} \right) \right] \right. \\ \times \frac{1}{1 + e^{(r-R)/b}} \left. \right\} \phi_n(r) - \frac{1}{b} \left(\mu - \frac{-1 + \nu^2}{2\mu + 1} + i\gamma \right) \\ \times e^{(1+i\gamma)(r-R)/b} [1 + e^{(r-R)/b}]^{-\mu - 2 - i\gamma} \\ \times F\left(\mu + i\gamma + 1, \mu + i\gamma + 2, 2\mu + 2; \frac{1}{1 + e^{(r-R)/b}} \right). \end{aligned} \tag{16}$$

Moreover, we can also give the solution of the relativistic Dirac-Woods-Saxon problem with a general λ as the deformation parameter. In order to do that, a new variable $x = [1 + e^{(r-R)/b}]^{-1}$ is introduced. Now, the equation of the upper spinor component has been transformed into a second-order differential equation as follows:

$$\begin{aligned} x^2(1-x)^2 \frac{d^2 \phi}{dx^2} + x(1-x)(1-2x) \frac{d\phi}{dx} - [\mu^2 + \lambda(\lambda + 1)x^2 \\ - (\lambda + \nu^2)x] \phi = 0, \end{aligned} \tag{17}$$

where the parameters $\mu^2 = b^2(1 - \varepsilon)^2/\alpha^2$, $\nu^2 = 2b^2\varepsilon V_0$, and $\lambda = \alpha b V_0/T$. Bound-state solutions of Eq. (17) satisfy the boundary conditions $\phi(0) = 0$ ($r \rightarrow \infty$) and $\phi(1) = 0$ ($r \rightarrow 0$). Using the trial function

$$\phi(x) = x^\mu (1-x)^\delta \psi(x) \quad (18)$$

with

$$\delta = \sqrt{\mu^2 - \nu^2 + \lambda^2}, \quad (19)$$

we find that Eq. (17) can be turned into the hypergeometric differential equation

$$x(1-x)\psi''(x) + [2\mu + 1 - (2\mu + 2\delta + 2)x]\psi'(x) - (\mu^2 + \delta^2 - \lambda^2 + 2\mu\delta + \mu + \delta - \lambda)\psi(x) = 0. \quad (20)$$

Taking into account the boundary condition $\phi(0) = 0$, we get

$$\phi(x) = x^\mu (1-x)^\delta F(\mu + \lambda + 1 + \delta, \mu - \lambda + \delta, 2\mu + 1; x). \quad (21)$$

In order to describe the behavior of the formula (21) in the vicinity of $x = 1$, we transform $\phi(x)$ into the following form:

$$\begin{aligned} \phi(x) &= x^\mu (1-x)^\delta \frac{\Gamma(2\mu + 1)\Gamma(-2\delta)}{\Gamma(\mu - \lambda - \delta)\Gamma(\mu + \lambda + 1 - \delta)} \\ &\quad \times F(\mu + \lambda + 1 + \delta, \mu - \lambda + \delta, 1 + 2\delta; 1-x) + x^\mu \\ &\quad \times (1-x)^{-\delta} \frac{\Gamma(2\mu + 1)\Gamma(2\delta)}{\Gamma(\mu + \lambda + 1 + \delta)\Gamma(\mu - \lambda + \delta)} \\ &\quad \times F(\mu - \lambda - \delta, \mu + \lambda + 1 - \delta, 1 - 2\delta; 1-x). \quad (22) \end{aligned}$$

Near the origin ($x \rightarrow 1, r = 0$) the wave function behaves as

$$\begin{aligned} \phi(x) &\sim (1-x)^\delta \frac{\Gamma(2\mu + 1)\Gamma(-2\delta)}{\Gamma(\mu - \lambda - \delta)\Gamma(\mu + \lambda + 1 - \delta)} \\ &\quad + (1-x)^{-\delta} \frac{\Gamma(2\mu + 1)\Gamma(2\delta)}{\Gamma(\mu + \lambda + 1 + \delta)\Gamma(\mu - \lambda + \delta)}. \quad (23) \end{aligned}$$

To discuss formula (23) we note that $\mu^2 - \nu^2 + \lambda^2 < 0$, so that according to Eq. (19) δ turns out to be imaginary:

$$\delta = i\gamma, \quad \gamma = \sqrt{\nu^2 - \mu^2 - \lambda^2}. \quad (24)$$

We then write

$$\begin{aligned} \phi(x) &\sim \frac{\Gamma(2\mu + 1)\Gamma(-2i\gamma)}{\Gamma(\mu - \lambda - i\gamma)\Gamma(\mu + \lambda + 1 - i\gamma)} \left[(1-x)^{i\gamma} \right. \\ &\quad + \frac{\Gamma(2i\gamma)\Gamma(\mu - \lambda - i\gamma)\Gamma(\mu + \lambda + 1 - i\gamma)}{\Gamma(-2i\gamma)\Gamma(\mu - \lambda + i\gamma)\Gamma(\mu + \lambda + 1 + i\gamma)} \\ &\quad \left. \times (1-x)^{-i\gamma} \right]. \quad (25) \end{aligned}$$

In the neighborhood of $r = 0$, $1 - x = e^{-R/b}$ approximately, so we have

$$\begin{aligned} \phi(x) &\sim \frac{\Gamma(2\mu + 1)\Gamma(-2i\gamma)}{\Gamma(\mu - \lambda - i\gamma)\Gamma(\mu + \lambda + 1 - i\gamma)} \left[e^{-i\gamma R/b} \right. \\ &\quad + \frac{\Gamma(2i\gamma)\Gamma(\mu - \lambda - i\gamma)\Gamma(\mu + \lambda + 1 - i\gamma)}{\Gamma(-2i\gamma)\Gamma(\mu - \lambda + i\gamma)\Gamma(\mu + \lambda + 1 + i\gamma)} e^{i\gamma R/b} \left. \right]. \quad (26) \end{aligned}$$

The boundary condition $\phi(1) = 0$ ($r \rightarrow 0$) leads to

$$\begin{aligned} \frac{\Gamma(2i\gamma)\Gamma(\mu - \lambda - i\gamma)\Gamma(\mu + \lambda + 1 - i\gamma)}{\Gamma(-2i\gamma)\Gamma(\mu - \lambda + i\gamma)\Gamma(\mu + \lambda + 1 + i\gamma)} e^{i\gamma R/b} + e^{-i\gamma R/b} \\ = 0, \quad (27) \end{aligned}$$

or

$$\frac{\Gamma(2i\gamma)\Gamma(\mu - \lambda - i\gamma)\Gamma(\mu + \lambda + 1 - i\gamma)}{\Gamma(-2i\gamma)\Gamma(\mu - \lambda + i\gamma)\Gamma(\mu + \lambda + 1 + i\gamma)} e^{2i\gamma R/b} = -1. \quad (28)$$

Thus we obtain the quantum condition

$$\exp\{2i[\arg \Gamma(2i\gamma) - \arg \Gamma(\mu - \lambda + i\gamma) - \arg \Gamma(\mu + \lambda + 1 + i\gamma) + \gamma R/b]\} = -1, \quad (29)$$

that is,

$$\begin{aligned} \arg \Gamma(2i\gamma) - \arg \Gamma(\mu - \lambda + i\gamma) - \arg \Gamma(\mu + \lambda + 1 + i\gamma) \\ + \gamma R/b = \left(n + \frac{1}{2}\right)\pi, \quad (n = 0, \pm 1, \pm 2, \dots) \quad (30) \end{aligned}$$

or

$$\begin{aligned} \arg \Gamma\left(2i \frac{b}{\alpha} \sqrt{\varepsilon_n^2 - 1 + 2V_0\varepsilon_n\alpha^2 - V_0^2\alpha^4/T^2}\right) - \arg \Gamma\left(\frac{b}{\alpha} \sqrt{1 - \varepsilon_n^2} - b\alpha V_0/T + i \frac{b}{\alpha} \sqrt{\varepsilon_n^2 - 1 + 2V_0\varepsilon_n\alpha^2 - V_0^2\alpha^4/T^2}\right) \\ - \arg \Gamma\left(\frac{b}{\alpha} \sqrt{1 - \varepsilon_n^2} + b\alpha V_0/T + 1 + i \frac{b}{\alpha} \sqrt{\varepsilon_n^2 - 1 + 2V_0\varepsilon_n\alpha^2 - V_0^2\alpha^4/T^2}\right) + \frac{R}{\alpha} \sqrt{\varepsilon_n^2 - 1 + 2V_0\varepsilon_n\alpha^2 - V_0^2\alpha^4/T^2} \\ = \left(n + \frac{1}{2}\right)\pi, \quad (31) \end{aligned}$$

and the upper spinor component is given by

$$\phi_n(r) = e^{i\gamma(r-R)/b} [1 + e^{(r-R)/b}]^{-\mu - i\gamma} F\left(\mu + \lambda + 1 + i\gamma, \mu - \lambda + i\gamma, 2\mu + 1; \frac{1}{1 + e^{(r-R)/b}}\right). \quad (32)$$

With Eq. (7), the lower spinor component is also obtained as

$$\theta_n(r) = \frac{\alpha}{C + \varepsilon_n} \left\{ \left[-\frac{S}{\alpha} - \frac{\mu}{b} + \left(\frac{\mu + i\gamma}{b} - \frac{V_0\alpha}{T} \right) \frac{1}{1 + e^{(r-R)/b}} \right] \phi_n(r) - \frac{1}{b} \left(\mu - \frac{\lambda + \nu^2}{2\mu + 1} + i\gamma \right) e^{(1+i\gamma)(r-R)/b} [1 + e^{(r-R)/b}]^{-\mu - 2 - i\gamma} \right. \\ \left. \times F\left(\mu + \lambda + 2 + i\gamma, \mu - \lambda + 1 + i\gamma, 2\mu + 2; \frac{1}{1 + e^{(r-R)/b}}\right) \right\}. \quad (33)$$

It can be easily seen that Eqs. (30), (32), and (33) give the same results as Eqs. (14)–(16) with $\lambda = -1$.

In conclusion, we have obtained the exact solution of the Dirac equation for the Woods-Saxon potential and presented the explicit form of the spinor wave function. The relativistic bound-state spectrum is also obtained through an equation that may be very useful to describe single-particle motion in nuclei, because spin-orbit coupling is involved automatically there.

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