Hypersurface Bohm-Dirac models

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We define a class of Lorentz-invariant Bohmian quantum models for N entangled but noninteracting Dirac particles. Lorentz invariance is achieved for these models through the incorporation of an additional dynamical space-time structure provided by a foliation of space time. These models can be regarded as the extension of Bohm's model for N Dirac particles, corresponding to the foliation into the equal-time hyperplanes for a distinguished Lorentz frame, to more general foliations. As with Bohm's model, there exists for these models an equivariant measure on the leaves of the foliation. This makes possible a simple statistical analysis of position correlations analogous to the equilibrium analysis for (the nonrelativistic) Bohmian mechanics. [S1050-2947(99)07910-X]

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I. INTRODUCTION

Among the different approaches to resolving the conceptual problems of quantum theory, Bohm's approach is perhaps the simplest. In a nutshell, it consists in adding the most basic dynamical variables, obeying additional evolution equations, to the description of a quantum system provided by its wave function ψ . For nonrelativistic quantum theory the additional variables are the positions of the particles, which evolve according to a "guiding equation" naturally suggested by the Schrödinger evolution. This theoryusually called Bohmian mechanics or the pilot-wave theory-is well understood. It has been analyzed, and its connection with the predictions of orthodox quantum theory explained, in the original papers of Bohm [1] as well as in later works (see, e.g., [2-4]). One of the main problems remaining for the Bohmian (or any other) approach is to find a satisfactory relativistic quantum theory, a theory that is fully Lorentz-invariant while avoiding the profound conceptual difficulties of orthodox quantum theory.

In his original papers, Bohm had an outline for a "Bohmian" field theory, with fields on space time as the additional variables. A year later he proposed a "Bohmian" model for one Dirac particle [5], which was subsequently extended by Bohm and coworkers to N Dirac particles [6]. For this N-particle model the additional variables are, as in Bohmian mechanics, the positions \mathbf{Q}_k , $k=1,\ldots,N$ of the particles. However, in contrast with Bohmian mechanics, the guiding equation for this theory

$$\frac{d\mathbf{Q}_{k}}{dt} = \frac{\psi^{\dagger} \boldsymbol{\alpha}_{k} \psi}{\psi^{\dagger} \psi} \tag{1}$$

is ultralocal on configuration space: The right-hand side of

Eq. (1) depends only upon the value of ψ at the positions of the particles and not upon spatial derivatives of ψ there. Here $\psi = \psi(\mathbf{q}_1, \ldots, \mathbf{q}_N, t)$, taking values in the *N*-particle spin space $(\mathbb{C}^4)^{\otimes N}$, solves the *N*-particle Dirac equation ($\hbar = c = 1$),

$$i\frac{\partial\psi}{\partial t} = \sum_{k=1}^{N} \left[-i\,\boldsymbol{\alpha}_{k} \cdot \boldsymbol{\nabla}_{k} - e\,\boldsymbol{\alpha}_{k} \cdot \mathbf{A}(\mathbf{q}_{k},t) + e\,\Phi(\mathbf{q}_{k},t) + \boldsymbol{\beta}_{k}m \right]\psi,$$
(2)

where $\alpha_k = (\alpha_k^1, \alpha_k^2, \alpha_k^3)$, $\alpha_k^i = I \otimes \cdots \otimes I \otimes \alpha^i \otimes I \otimes \cdots \otimes I$, with the *i*th Dirac α matrix α^i at the *k*th of the *N* places, and β_k is defined analogously. Φ and **A** are external electromagnetic potentials. (We may, of course, consider particledependent masses m_k , charges e_k , and external potentials Φ_k and \mathbf{A}_k , but for simplicity we shall not do so.) We shall call this model the Bohm-Dirac (BD) model. Just as with Bohm's proposal for a field theory, the BD model requires for its formulation the specification of a distinguished frame of reference—in terms of which the actual configuration $(\mathbf{Q}_1, \ldots, \mathbf{Q}_N)$ and the generic configuration $(\mathbf{q}_1, \ldots, \mathbf{q}_N)$ at time *t* is defined—and, in fact, the model is not Lorentzinvariant if N > 1 [6].

However, for N=1 this model is Lorentz-invariant, and may be formulated in a covariant way: Writing $X=X(\tau)$ for the space-time point along a trajectory, with (scalar) parametrization τ , the guiding equation may be written as

$$\frac{dX}{d\tau} = j \equiv \bar{\psi}\gamma\psi \tag{3}$$

with ψ satisfying the Dirac equation

$$(i\gamma \cdot \partial - e\gamma \cdot A - m)\psi = 0, \qquad (4)$$

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where $\gamma \cdot \partial \equiv \gamma^{\mu} \partial_{\mu}$ and $\gamma \cdot A \equiv \gamma^{\mu} A_{\mu}(x)$. Note that the righthand side of Eq. (3), the Dirac current $j = j^{\mu} \equiv \bar{\psi} \gamma^{\mu} \psi$, is the simplest four-vector that can be constructed from the Dirac spinor ψ .

Note also that the parameter τ has no intrinsic physical significance, so that Eq. (3) is equivalent to

$$\frac{dX}{d\tau} = aj$$

with arbitrary positive scalar field a = a(x). It is not the field of four-vectors *j* (having direction and length) that determines the particle motion, but rather the field of directions defined by *j*. In other words, the law for the particle motion could be formulated in a purely geometrical manner as the condition that the Dirac current *j* at every point along the trajectory be tangent to the trajectory at that point.

Because the Dirac current is timelike and divergence-free, $^{\rm l}$

$$\partial \cdot j = 0$$
,

there is a dynamically distinguished probability distribution on the set of particle paths $X(\tau)$ arising from Eq. (3). Any distribution on this space of paths can be defined by specifying for the path the crossing probability for some given equal-time surface Σ_0 in some Lorentz frame. (By this crossing probability we mean the distribution of the point through which the path crosses Σ_0 , which is the same thing as the probability distribution for the position of the particle in this frame at the given time.) The distinguished distribution is then defined by the crossing probability for Σ_0 given by ϱ $=j^0=\psi^{\dagger}\psi$ on Σ_0 (with ψ suitably normalized), which can be written in a covariant manner as $j \cdot n$ where *n* is the futureoriented unit normal to the surface. For this distribution the crossing probability for any other equal-time surface will also be given by $j \cdot n$, both for the original frame and any other Lorentz frame. We may roughly summarize the situation by saying that for the distinguished probability distribution, quantum equilibrium holds in all Lorentz frames at all times, with the quantum equilibrium distribution given by $\rho = \psi^{\dagger} \psi.$

More generally, the crossing probability for any spacelike hypersurface Σ will also be given by $j \cdot n$, with n = n(x) the future-oriented unit normal field to Σ . Moreover, for any oriented hypersurface Σ , the crossing measure (a signed measure that need not be normalized), which describes the expected number of signed crossings through area elements of Σ , with negatively oriented crossings counted negatively, is, for the distinguished distribution, also given by $j \cdot n$, with n = n(x) now the positively oriented unit normal field to Σ .² The *N*-particle BD model (1) also has a dynamically distinguished probability distribution on paths. As a consequence of Eq. (2) $\varrho = \psi^{\dagger} \psi$ satisfies, in the Lorentz frame in which the dynamics is defined, the continuity equation

$$\frac{\partial \boldsymbol{\varrho}}{\partial t} + \sum_{k=1}^{N} \boldsymbol{\nabla}_{k} \cdot \mathbf{J}_{k} = 0, \qquad (5)$$

where

$$\mathbf{J}_{k} = \boldsymbol{\varrho} \, \mathbf{v}_{k} = \boldsymbol{\psi}^{\dagger} \, \boldsymbol{\alpha}_{k} \, \boldsymbol{\psi}. \tag{6}$$

Thus, if the joint probability distribution for the positions of the N particles is given $\rho = \psi^{\dagger} \psi$ at some time $t = t_0$, then, for the corresponding distribution on paths, it will be given by $\rho = \psi^{\dagger} \psi$ at all times t. However, even for this distinguished distribution, quantum equilibrium will not, in general, hold in other Lorentz frames: The joint distribution of crossings of equal-time surfaces for other frames will in general not be given by $\psi'^{\dagger}\psi'$ (where ψ' is the wave function in the relevant Lorentz frame) [7,8]. Nonetheless, Bohm and coworkers have argued that the observational content of this model is as Lorentz-invariant as the covariant formalism of relativistic quantum theory: Since the predictions for results of measurements for this model can be regarded as reflected in the configuration of various devices and registers-and hence can be derived from probabilities for positions given by $\rho = \psi^{\dagger} \psi$ —at a common time in the distinguished frame, these predictions must agree with those of the usual interpretation. Thus, no violation of Lorentz invariance can be detected in experiments [6]. (In particular, the identity of the distinguished Lorentz frame cannot be ascertained by means of any possible observation.)

Lorentz invariance is, however, a delicate issue. Indeed, any theory can be made trivially Lorentz-invariant (or invariant under any other space-time symmetry), even on the microscopic level, by the incorporation of suitable additional structure [8]. For this reason Bell has stressed that one should consider what he has called "serious Lorentz invariance," a notion, however, that is extremely difficult to make precise in an adequate way [2]. Lacking a general criterion, we may nonetheless begin to get a handle on "serious Lorentz invariance" by analyzing some specific models. If the models involve additional structure, then whether or not we have serious Lorentz invariance will depend, of course, upon the detailed nature of this structure.

In [8] we have considered a model for which the additional structure for a system of N (noninteracting) Dirac particles is provided by a global synchronization among the particles: The trajectories of the particles are such that each one of them at some given space-time point is tangent to a vector field determined, given the wave function, by that point and those points along the trajectories of the other particles with which that point has been "synchronized." This additional synchronization structure is defined implicitly by the equation of motion and the model is not amenable to a statistical analysis in any obvious way. In other words, this model is not statistically transparent (see Sec. IV of [8]). Nonetheless, even this model provides a counterexample to the widely held belief that a Lorentz-invariant Bohmian

¹The claims in this and the next paragraph follow directly from the application of the divergence theorem (or Stokes' theorem) to an infinitesimally thin tube of paths between Σ_0 (see below) and the relevant hypersurface Σ .

²In this regard it is perhaps worth noting the following: In Minkowski space there is a natural duality between divergence-free vector fields and closed three-forms. Such a vector field defines a "deterministic" random path, whose "law" is given directly by the vector field, as in Eq. (3), and whose statistics are governed by the dual three-form, in the manner just described.

theory for many particles is impossible (unless only product states are allowed). In this regard, see also the local model of Squires [9].

In this paper we shall analyze a statistically transparent counterexample, the "hypersurface Bohm-Dirac (HBD) model." The basic idea was proposed in [10] in the context of bosonic quantum field theory: In addition to the wave function and field variables, a distinguished foliation of space time—a new element of geometrical structure defining simultaneity surfaces—is suggested as an additional dynamical variable of the theory. These surfaces need not be hyperplanes. The defining (Lorentz-invariant) equations of the theory should describe the evolution of the wave function, the field variables, and the simultaneity surfaces. For a careful philosophical discussion of how this may be compatible with some appropriate notion of relativity, even if the simultaneity surfaces should turn out to be unobservable, see Maudlin [11].

Here we shall consider such a theory, not for fields but for N (noninteracting) Dirac particles. We shall discuss an as yet incomplete hypersurface Bohm-Dirac model: The law for the evolution of the foliation is not specified beyond the requirement that it not involve the positions of the particles. We present no hypothesis concerning the origin of the foliation, but have in mind that the foliation should ultimately be governed by a Lorentz-invariant law; one that may, for example, involve the N-particle wave function. (For definiteness we shall give some very tentative and less than compelling examples of laws for the foliation in Sec. IV.) However, we show in Sec. III A that regardless of how the foliation is determined, the dynamics of the HBD model preserves the quantum equilibrium distribution on the leaves of the foliation. Thus the model is amenable to the same sort of statistical analysis as for nonrelativistic Bohmian mechanics. This is discussed briefly in Sec. III B.

II. THE HYPERSURFACE BOHM-DIRAC MODEL

A general foliation \mathcal{F} of codimension one on Minkowski space M can approximately be thought of as a partition of Minto three-dimensional hypersurfaces. These hypersurfaces are the leaves of the foliation. The simplest way to obtain a foliation is by a smooth function $f: M \to \mathbb{R}$ without critical points, i.e., $df \neq 0$ everywhere. The level sets $f^{-1}(s)$ are smooth hypersurfaces and form a foliation of M. With the one-form df_x , which vanishes on the tangent space of the hypersurface through $x \in M$, we may associate by the Lorentz metric the normal vector field $\partial f(x)$. If this is timelike everywhere, and thus the foliation hypersurfaces spacelike, we may normalize $\partial f(x)$ to obtain a unit normal vector field n(x) associated with the foliation \mathcal{F} .

We shall consider in this paper only spacelike foliations, i.e., foliations by spacelike hypersurfaces. While obviously different *f*'s may generate the same foliation \mathcal{F} , the futureoriented unit normal vector field *n* is uniquely determined by \mathcal{F} . When does a vector field v(x) determine a foliation \mathcal{F} such that for all $x \in M$, v(x) is normal to the tangent space of the foliation hypersurface through *x*? If we denote by *V* the one-form associated with *v* by the Lorentz metric, then, by Frobenius' theorem, the necessary and sufficient condition is that *V* be completely integrable, $V \land dV = 0$. Apart from the foliation, the other dynamical variables of the hypersurface Bohm-Dirac model are the usual ones: the wave function ψ , here for N Dirac particles, and the N path, the N-tuple of (everywhere either timelike or lightlike) space-time paths, which describes the trajectories of the N Dirac particles. Covariant laws for these dynamical variables suggest themselves when we write those of the Bohm-Dirac model, defined by Eqs. (1) and (2), in a coordinate-free, i.e., covariant manner.

To achieve this we consider first of all the ψ function in the multitime formalism: For *N* Dirac particles the wavefunction $\psi = \psi(x_1, x_2, \dots, x_N)$, $x_k \in M$, takes values in the *N*-particle spin space $(\mathbb{C}^4)^{\otimes N}$ and satisfies *N* Dirac equations

$$[i\gamma_k \cdot \partial_k - e\gamma_k \cdot A(x_k) - m]\psi = 0, \qquad (7)$$

 $k=1, \ldots, N$. Here, $\gamma_k = I \otimes \cdots \otimes I \otimes \gamma \otimes I \otimes \cdots \otimes I$, with γ at the *k*th of the *N* places, and *A* is an external electromagnetic potential. [Just as with Eq. (2), we may, of course, consider particle-dependent masses m_k , charges e_k , and external potentials A_k .] The system of equations (7) is a covariant version of Eq. (2); in this multitime form the Lorentz invariance of the law for ψ is manifest [6].³ The *N* Dirac particles are coupled by the common wave-function ψ . If this is entangled, we have nonlocal correlations between the *N* particles, despite the fact that the particles are noninteracting.

We shall now develop the guiding law for the N path. Note that the numerator of the right-hand side of Eq. (1) is given by a current j_k ,

$$j_k = \overline{\psi} \gamma_1^0 \dots \gamma_k \dots \gamma_N^0 \psi,$$

which involves matrix elements of an operator having as factors the 0 component γ^0 of a four-vector for all but the *k*th particle. Therefore, j_k can be expressed in a covariant manner by replacing γ_k^0 in the above expression with $\gamma_k \cdot n$, where *n* is the future-oriented unit normal to the t = const hyperplanes,

$$j_k = \overline{\psi}(\gamma_1 \cdot n) \dots \gamma_k \dots (\gamma_N \cdot n) \psi.$$
(8)

Moreover, the denominator of the right-hand side of Eq. (1) can be expressed covariantly as $j_k \cdot n$. Then, the covariant velocity of the *k*th particle—with respect to the time of a Lorentz frame with *n* as time axis—is

$$\frac{dX_k}{dt} = \frac{j_k}{j_k \cdot n}.$$
(9)

Since $j_k \cdot n = \overline{\psi}(\gamma_1 \cdot n) \dots (\gamma_N \cdot n)\psi$ is independent of *k*, we may reparametrize the paths with a parameter *s* so related to *t* that $t'(s) = j_k \cdot n$ to obtain

$$\frac{dX_k}{ds} = j_k \,. \tag{10}$$

³Note that in the single-time form (2) we can easily add an explicit interaction potential $V(\mathbf{q}_1, \ldots, \mathbf{q}_N, t)$ for the *N* Dirac particles, while in the multitime form this is impossible.

More generally, by further reparametrization we may obtain $dX_k/d\tau = aj_k$, where *a* is any positive scalar field. The physical particle dynamics—i.e., the *N* space-time paths defined by the equations of motion (and initial conditions)—is invariant under reparametrization.

A manifestly "parametrization-invariant" formulation of the dynamics—that is, such that a time parameter plays no role—is easily obtained: The space-time paths for the *N* particles are constrained by the currents j_k by requiring that the path for the *k*th particle at the point x_k be tangent to the current j_k evaluated at x_k and at the intersection points of the paths of the N-1 other particles with the t=const-hyperplane Σ_t containing x_k . If we denote by $X_k(\Sigma_t)$ the intersection point of the path X_k with the hyperplane Σ_t , and by $\dot{X}_k(\Sigma_t)$ a tangent of (or the tangent line to) the path X_k at $X_k(\Sigma_t)$, we may write the law for the *N* path as

$$\dot{X}_k(\Sigma_t) \parallel j_k(X_1(\Sigma_t), \dots, X_N(\Sigma_t)), \tag{11}$$

using the symbol \parallel for "is parallel to." In this geometric formulation the Bohm-Dirac dynamics depends upon the Lorentz frame only via its associated foliation into simultaneity hypersurfaces Σ_t , and thus naturally extends to an arbitrary foliation \mathcal{F} of Minkowski space-time *M* by curved spacelike hypersurfaces.⁴

Given such a foliation \mathcal{F} and $\Sigma \in \mathcal{F}$, let $X_k(\Sigma)$ be the intersection of the path X_k with Σ ,⁵ and let $\dot{X}_k(\Sigma)$ be a tangent of (or the tangent line to) the path X_k at $X_k(\Sigma)$. The law of the *N* path $X = (X_1, \ldots, X_N)$ for the hypersurface Bohm-Dirac model is defined by the currents j_k naturally extending Eq. (8),

$$j_k = \overline{\psi}(\gamma_1 \cdot n_1) \dots \gamma_k \dots (\gamma_N \cdot n_N) \psi, \qquad (12)$$

where $n_1 \equiv n(x_1), \ldots, n_N \equiv n(x_N)$, with *n* the future-oriented unit normal vector field associated with \mathcal{F} , via the HBD tangency condition (see also Fig. 1)

$$\dot{X}_k(\Sigma) \parallel j_k(X_1(\Sigma), \dots, X_N(\Sigma)).$$
(13)

(By considering the action of a suitable Lorentz transformation on $\gamma^0 \gamma \cdot n$ for arbitrary timelike unit vector *n* [transforming *n* to (1,0,0,0)], one sees that $\gamma^0 \gamma \cdot n$ is a positive operator in spin space \mathbb{C}^4 . Hence, $(\gamma_1^0 \gamma_1 \cdot n_1) \dots (\gamma_k^0 \gamma_k \cdot n) \dots (\gamma_N^0 \gamma_N \cdot n_N)$ is also positive, i.e., $j_k \cdot n \ge 0$ with "="



FIG. 1. Geometrical formulation of the dynamics for a system of three particles: For each particle the path of that particle, say, particle 1 at x_1 , must be tangent to the four-vector j_1 which is determined by (i) the intersections x_2 and x_3 of the trajectories of the other two particles with the hypersurface Σ containing x_1 , (ii) the future-oriented unit normals n_2 and n_3 at these points, and (iii) the wave function of the system evaluated at x_1 , x_2 , and x_3 : $j_1 = \overline{\psi}(x_1, x_2, x_3) \gamma_1(\gamma_2 \cdot n_2)(\gamma_3 \cdot n_3) \psi(x_1, x_2, x_3)$.

only if $\psi = 0$. This means that where it is nonzero, j_k is future-oriented and, like the path X_k , nowhere spacelike.)

We may also write down the equations of motion in the parametrized form analogous to Eq. (9) or Eq. (10). To do so it is convenient to label the hypersurfaces of the foliation using a function $f: M \to \mathbb{R}$ that generates the foliation as described above, and use this hypersurface labeling as the parameter for the particle trajectories—so that $X_k(s)$ is on the hypersurface $f^{-1}(s)$. From the geometrical characterization of the dynamics (13) we know that dX_k/ds is parallel to $j_k(X_1(s), \ldots, X_N(s))$, and the scale factor required to ensure $f(X_k(s)) = s$ for all k and s is easily seen to be $1/(\partial f \cdot j_k)$. Therefore,

$$\frac{dX_k}{ds} = \frac{j_k(X_1(s), \dots, X_N(s))}{\partial f(X_k(s)) \cdot j_k(X_1(s), \dots, X_N(s))}.$$
 (14)

For a flat foliation we may choose a Lorentz frame such that the foliation hyperplanes are the $x^0 = \text{const}$ planes, i.e., $f(x) = x^0$ for all x. Then $n = \partial f = (1,0,0,0)$ and Eq. (14) reduces to the Bohm-Dirac law (1).

III. STATISTICAL ANALYSIS OF THE HBD MODEL

A. Quantum equilibrium

We shall show now that for the hypersurface Bohm-Dirac model with foliation \mathcal{F} , there is a distinguished probability measure on *N* paths *X* satisfying the HBD tangency condition (13), one for which the distribution of hypersurface crossings

⁴This is in marked contrast with the parametrized dynamics such as given by Eq. (9) or Eq. (10), which need not extend in anything like the same form to a general foliation since the parametrized paths generated by the dynamics need not, in general, respect the foliation.

⁵Note that the paths X_k comprising an N path, since they are nowhere spacelike, can intersect Σ at most once. This is the main reason why it is important that the foliation \mathcal{F} be spacelike. Of course, also from the physical point of view a synchronization along spacelike hypersurfaces yields a picture which perhaps makes most sense. We shall assume without further ado global existence: that a fragment of an N path locally satisfying the HBD tangency condition, see Eq. (13), can be continued in such a manner that each of its paths X_k intersects every $\Sigma \in \mathcal{F}$.

 $X_1(\Sigma), \ldots, X_N(\Sigma)$ for $\Sigma \in \mathcal{F}$ depends only upon ψ restricted to Σ (or, more precisely, to Σ^N) for ψ satisfying Eq. (7). We shall say that such a distinguished measure, as well as the corresponding hypersurface crossing distribution, is *equivariant*, defining *quantum equilibrium*. The physical significance of the hypersurfaces $\Sigma \in \mathcal{F}$ is thus twofold: They serve [via Eq. (13)] to define the motion of the particles and, for a quantum equilibrium *N* path, it is "on these hypersurfaces" that manifestly the "particles are in quantum equilibrium."

The natural candidate for the equivariant crossing probability density ϱ of the HBD model is given by the obvious covariant extension of the equivariant density $\psi^{\dagger}\psi$ $(=\bar{\psi}\gamma_1^0\ldots\gamma_N^0\psi)$ of the BD model:

$$\varrho = \overline{\psi}(\gamma_1 \cdot n_1) \dots (\gamma_N \cdot n_N)\psi. \tag{15}$$

To see that this is in fact equivariant, note the following: In view of Eq. (12), (i) $\rho = j_k \cdot n_k$ and

$$j_k \cdot n_k$$
 is independent of k . (16)

Furthermore, (ii) the currents j_k are divergence-free:

$$\partial_k \cdot j_k = 0, \tag{17}$$

which follows immediately from Eq. (12) using the Dirac equation (7) and its adjoint. These two properties of the currents, Eqs. (16) and (17), are the key ingredients for the proof of the equivariance of ϱ . For any current satisfying Eqs. (16) and (17), for the particle dynamics defined by Eq. (13), $\varrho = j_k \cdot n_k$ is an equivariant probability density for crossings of the leaves of the foliation.⁶

The proof of this assertion consists of two steps: First, we determine how an arbitrary probability density R on crossings of a foliation hypersurface Σ evolves under the dynamics (13), i.e., we formulate the continuity equation of the hypersurface dynamics. In the second step, we show that $R = \rho$ solves the continuity equation. It then follows that if the probability distribution of the "positions of the N particles" on $\Sigma \in \mathcal{F}$ is given by ρ restricted to Σ , then for any other hypersurface $\Sigma' \in \mathcal{F}$, the probability distribution of the "positions of the N particles" on Σ' , which emerges by transport according to the dynamics (13), is given by ρ restricted to Σ' . Thus, ρ is equivariant.



FIG. 2. Motion of two particles in one space dimension from hypersurface Σ to Σ' : space-time view. We have indicated the positions of the primed points x'_k obtained from x_k via displacement from Σ to Σ' in the normal direction, and the images $\delta x'_k$ of the regions δx_k under this correspondence. The point on Σ' to which particle k moves when starting at $x_k \in \Sigma$ is given (to leading order) by $x_k + v_k \delta \tau_k$ with $v_k = j_k/(j_k \cdot n_k)$, where $\delta \tau_k$ is the Minkowski distance between x_k and x'_k .

Consider thus two infinitesimally close hypersurfaces Σ and Σ' belonging to the foliation \mathcal{F} . The probability distribution of the positions of the *N* particles on Σ is given by a density $R_{\Sigma}: \Sigma^N \to \mathbb{R}$ such that

Prob(particle *i* crosses Σ in δx_i , i = 1, ..., N) = $R_{\Sigma}(x_1, ..., x_N) \, \delta x_1 \dots \, \delta x_N$.

By δx we denote simultaneously an infinitesimal region on Σ around *x* and its area (i.e., three-volume). Now we compare R_{Σ} evaluated at $(x_1, \ldots, x_N) \in \Sigma^N$ with $R_{\Sigma'}$ evaluated at $(x'_1, \ldots, x'_N) \in (\Sigma')^N$, where $x' \in \Sigma'$ is obtained from $x \in \Sigma$ via displacement from Σ to Σ' in the normal direction, see Fig. 2. Let $\delta x'$ be the area of the image of the region δx under this correspondence. (Since the projection of the Lorentz metric on Σ' need not agree with the image, under $x \mapsto x'$, of its projection on Σ , δx and $\delta x'$ need not agree.)

Recall from elementary physics that a continuity equation such as Eq. (5) is an expression of a local conservation law that on the infinitesimal level can be stated as follows: The difference between the probability densities R_{Σ} on Σ^N and $R_{\Sigma'}$ on $(\Sigma')^N$ (with Σ' infinitesimally later than Σ) is accounted for by the flux through the lateral sides—to which the hypersurface normals are tangent—of the configurationspace-time box between $\delta x_1 \times \cdots \times \delta x_N \subset \Sigma^N$ and the corresponding set of (primed) points in $(\Sigma')^N$, see Fig. 3;

$$R_{\Sigma'}(x_1', \dots, x_N') \,\delta x_1' \dots \,\delta x_N' - R_{\Sigma}(x_1, \dots, x_N) \,\delta x_1 \dots \,\delta x_N$$

$$= -\sum_{k=1}^N \,\delta x_1 \dots \,\widehat{\delta x_k} \dots \,\delta x_N \int_{\partial(\delta x_k)} (R_{\Sigma} v_k)(x_1, \dots, x_{k-1}, y, x_{k+1}, \dots, x_N) \cdot (u_k \delta \tau)(y) dS_k, \tag{18}$$

where the hat on δx_k indicates that this term should be omitted from the product. Here, y is the integration variable on $\partial(\delta x_k)$, the (two-dimensional) boundary of δx_k regarded as a region in Σ , dS_k is the area element of $\partial(\delta x_k)$, u_k is the outward unit normal vector field in Σ to $\partial(\delta x_k)$, $\delta \tau(y)$ is the Minkowski distance between $y \in \Sigma$ and the corresponding $y' \in \Sigma'$ [so that $y' = y + \delta \tau(y)n(y)$], and

$$v_k = \frac{j_k}{j_k \cdot n_k} \tag{19}$$

is the covariant velocity of the *k*th particle relative to Σ , see Fig. 2.

⁶In contrast, the current $j_k = \overline{\psi} \gamma_k \psi$ we considered in [8] satisfies Eq. (17) but not Eq. (16).



FIG. 3. Conservation of probability for a system of two particles in one space dimension: configuration-space-time view with, for simplicity, the hypersurfaces drawn straightened out. [Note that the figure fails to convey the fact—displayed in Fig. 2—that the areas δx_k and $\delta x'_k$ may differ, and that also $\delta \tau(y_k)$ may differ from $\delta \tau(\bar{y}_k)$, where y_k and \bar{y}_k are the boundary points of δx_k .] The change of the probability of particle 1 being in δx_1 and particle 2 being in δx_2 from hypersurface Σ to Σ' is accounted for by the single-particle fluxes through the lateral sides of the configuration space-time box between $\delta x_1 \times \delta x_2 \subset \Sigma^2$ and the corresponding set of primed points on $(\Sigma')^2$, i.e.,

$$R_{\Sigma'}(x_1', x_2') \delta x_1' \delta x_2' - R_{\Sigma}(x_1, x_2) \delta x_1 \delta x_2$$

= -[(R_{\Sigma} v_1)(\bar{y_1}, x_2) \cdot (u_1 \delta \tau)(\bar{y_1}) + (R_{\Sigma} v_1)(y_1, x_2) \cdot (u_1 \delta \tau)(y_1)] \delta x_2
-[(R_{\Sigma} v_2)(x_1, \bar{y_2}) \cdot (u_2 \delta \tau)(\bar{y_2}) + (R_{\Sigma} v_2)(x_1, y_2) \cdot (u_2 \delta \tau)(y_2)] \delta x_1

Equation (18) is the natural extension of this formula to N particles in Minkowski space.

Equation (18) is the continuity equation for the HBD model in the ''infinitesimally integrated form.'' It is valid for any hypersurface dynamics defined by Eq. (13), regardless of whether the currents j_k satisfy Eqs. (16) and (17). However, as we shall now show, if the currents do satisfy Eqs. (16) and (17), then $R_{\Sigma} = \varrho |_{\Sigma} = (j_k \cdot n_k)|_{\Sigma}$ satisfies Eq. (18). Since⁷

$$\varrho(x'_1, \dots, x'_N) \,\delta x'_1 \dots \,\delta x'_N - \varrho(x_1, \dots, x_N) \,\delta x_1 \dots \,\delta x_N$$

= $\varrho(x'_1, \dots, x'_N) \,\delta x'_1 \dots \,\delta x'_N$
- $\varrho(x_1, x'_2, \dots, x'_N) \,\delta x_1 \,\delta x'_2 \dots \,\delta x'_N$
+ $\varrho(x_1, x'_2, \dots, x'_N) \,\delta x_1 \,\delta x'_2 \dots \,\delta x'_N$
- $\varrho(x_1, x_2, x'_3, \dots, x'_N) \,\delta x_1 \,\delta x_2 \,\delta x'_3 \dots \,\delta x'_N$

$$+\dots+\varrho(x_1,\dots,x_{N-1},x'_N)\,\delta x_1\dots\,\delta x_{N-1}\,\delta x'_N$$
$$-\varrho(x_1,\dots,x_N)\,\delta x_1\dots\,\delta x_N,\qquad(20)$$

we obtain in this case for the left-hand side of Eq. (18) (to leading order),

$$\sum_{k=1}^{N} \delta x_1 \dots \widehat{\delta x_k} \dots \delta x_N [j_k(x_1, \dots, x'_k, \dots, x_N) \cdot n(x'_k) \delta x'_k - j_k(x_1, \dots, x_k, \dots, x_N) \cdot n(x_k) \delta x_k],$$
(21)

while the integrand on the right-hand side of Eq. (18) becomes $(j_k \cdot u_k) \, \delta \tau \, dS_k$. Thus, subtracting the right-hand side of Eq. (18) from Eq. (21), we obtain (to leading order) the sum over k of the integral of j_k over the (outward-oriented) boundary of the space-time region above δx_k between Σ and Σ' . But since j_k is divergence-free [Eq. (17)], each such term, and hence the sum, vanishes. Thus Eq. (18) is satisfied, establishing the equivariance of Q.

We may also write the continuity equation (18) in a purely local form: Writing

$$\delta R_{\Sigma}(x_1,\ldots,x_N) = R'_{\Sigma'}(x'_1,\ldots,x'_N) - R_{\Sigma}(x_1,\ldots,x_N),$$

where

$$R'_{\Sigma'}(x'_1, \dots, x'_N) \,\delta x_1 \dots \,\delta x_N$$

= $R_{\Sigma'}(x'_1, \dots, x'_N) \,\delta x'_1 \dots \,\delta x'_N,$ (22)

and applying Gauss' theorem to the right-hand side of Eq. (18),

$$\int_{\partial(\delta x_k)} R_{\Sigma} v_k \cdot u_k \, \delta \tau \, dS_k = \operatorname{div}_k^{\Sigma} [R_{\Sigma} v_k^{\Sigma} \, \delta \tau(x_k)] \, \delta x_k,$$

where div $_{k}^{\Sigma}$ is the divergence with respect to the *k*th coordinate x_{k} on the Riemannian manifold Σ and v_{k}^{Σ} is the projection of v_{k} on Σ , yields

$$\delta R_{\Sigma} + \sum_{k=1}^{N} \operatorname{div}_{k}^{\Sigma} (R_{\Sigma} v_{k}^{\Sigma} \delta \tau_{k}) = 0, \qquad (23)$$

where $\delta \tau_k \equiv \delta \tau(x_k)$.

Using this form we may also check the equivariance of ϱ . To do so, we first "smoothly" label the hypersurfaces of the foliation \mathcal{F} by a parameter $s \in \mathbb{R}$, increasing in the future direction, that may be called a "time parameter," in terms of which Eq. (23) becomes a standard differential equation. The function $f: \mathcal{M} \to \mathbb{R}$ that maps any point $x \in \mathcal{M}$ to the label *s* of the hypersurface Σ_s to which *x* belongs generates the foliation in the manner described in Sec. II. In particular, ∂f $= \|\partial f\| n$, where *n* is the future-oriented unit normal vector field of \mathcal{F} . With $\delta s = \|\partial f_k\| \delta \tau_k$, where $\partial f_k \equiv \partial f(x_k)$, we get from Eq. (19) that

$$v_k \delta \tau_k = \frac{j_k}{j_k \cdot \partial f_k} \delta s \equiv \hat{v}_k \delta s, \qquad (24)$$

with $\hat{v}_k = dX_k/ds$ the velocity of the *k*th particle in the parametrized formulation of the dynamics (14).

⁷Note that this decomposition is possible because ϱ is defined on M^N (with *M* Minkowski space), in contrast with an arbitrary $R = (R_{\Sigma})_{\Sigma \in \mathcal{F}}$, defined only for *N*-tuples belonging to Σ^N for some $\Sigma \in \mathcal{F}$, for which, therefore, such a decomposition is impossible.

Consider now a coordinate system adapted to our parametrized foliation Σ_s : one coordinate is clearly given by *s*, and on one foliation hypersurface we introduce an (arbitrary) coordinate system *p*, which is transported to the other foliation hypersurfaces by the flow along the normal field, yielding the system of coordinates (s,p), allowing us to write x = (s,p) for $x \in M$. Then $x_k = (s_k, p_k) \in \Sigma_s \Leftrightarrow s_k = s$, and the relation between x = (s,p) and x' = (s',p') from Fig. 2 becomes p = p'. Let δp be the volume element defined by the *p* coordinates and let $\delta x = g(p,s) \delta p$. In these adapted coordinates the continuity equation (23) assumes, using Eqs. (22) and (24), the more standard form

$$\frac{1}{g_1 \cdots g_N} \frac{\partial (g_1 \cdots g_N R_s)}{\partial s} + \sum_{k=1}^N \operatorname{div}_k^{\Sigma_s} (R_s \hat{v}_k^{\Sigma_s}) = 0,$$
(25)

with $R_s(p_1, \ldots, p_N) = R_{\Sigma_s}((s, p_1), \ldots, (s, p_N))$ and $g_k = g(s, p_k), k = 1, \ldots, N$ (and where $\hat{v}_k^{\Sigma_s}$ is the projection of \hat{v}_k on Σ_s).⁸

For $R_s = \varrho_s$, Eq. (20) is what lies behind the usual (implication of the) chain rule

$$\frac{1}{g_1 \dots g_N} \frac{\partial (g_1 \dots g_N \varrho_s)}{\partial s} = \sum_{k=1}^N \frac{1}{g_k} \frac{\partial [g(s_k, p_k) \varrho(s_1, p_1, \dots, s_k, p_k)]}{\partial s_k} \bigg|_{s_k = s}.$$
 (26)

Splitting the four-divergence into pieces corresponding to variations orthogonal to and variations within Σ_s , we obtain

$$\operatorname{div} j = \|\partial f\| \left(\frac{1}{g} \frac{\partial}{\partial s} (gj^0) + \operatorname{div}^{\Sigma_s} (\|\partial f\|^{-1} j^{\Sigma_s}) \right),$$

where j^0 is the normal component of j, $j^0 = j \cdot n$. Setting $j = j_k$ and using div $j_k = 0$ [Eq. (17)] we then find with Eq. (24) that

$$\frac{1}{g_k} \frac{\partial (g_k \varrho)}{\partial s_k} + \operatorname{div}_k^{\Sigma_s} (\varrho \, \hat{v}_k^{\Sigma_s}) = 0$$

for all *k*. Therefore, in view of Eq. (26), summation over *k* establishes that $R_s = \rho_s$ satisfies the HBD continuity equation (25).

B. Comparison with quantum mechanics

The statistical analysis of the hypersurface Bohm-Dirac model can be based on the assumption that the probability distribution on *N* paths is given by the equivariant density ρ [Eq. (15)] on some simultaneity surface Σ belonging to the foliation \mathcal{F} . Then, by equivariance, the statistical predictions of the HBD model (i.e., the crossing probabilities) agree with

the quantum predictions for positions for any hypersurface in \mathcal{F} . But what can be said about the statistical predictions concerning a hypersurface which is not part of a member of \mathcal{F} ?

For one particle the situation is very simple: From the geometrical formulation of the HBD model (Sec. II) it follows immediately that the HBD model for one particle is foliation-independent, and, in fact, is the usual one-particle Bohm-Dirac theory given by Eqs. (3) and (4), with current $j = \bar{\psi} \gamma \psi$. Thus, in this case the statistical predictions of the model agree with the quantum predictions for position along *any* hypersurface.

The situation is analogous for *N* independent particles: If the wave-function ψ is a product wave-function, $\psi = \psi_1(x_1) \cdots \psi_N(x_N)$, then it follows from the multitime Dirac equation (7) that ψ_k satisfies the usual one-particle Dirac equation. Furthermore, the path of the *k*th particle is tangent to the one-particle current $\overline{\psi}_k \gamma \psi_k$ and thus independent of the paths of the other particles. Moreover, ϱ is the product of the corresponding one-particle distributions. Therefore, a product wave function indeed generates a foliation-independent motion, the motion of *N* independent Bohm-Dirac particles, and we thus have agreement with all the quantum position distributions in this case.

In the general case the situation is more subtle: If the *N*-particle wave function is entangled, it will not, in general, be the case that the distributions of crossings of hypersurfaces not belonging to the foliation agree with the corresponding quantum position distributions [7,8] (which, in fact, may be incompatible with the crossing statistics for any trajectory model whatsoever). However, this disagreement does not entail violations of the quantum predictions, as has been discussed for the case of the multitime translation-invariant Bohmian theory in [8]. In fact, insofar as results of measurement are concerned, the predictions of our model are the same as those of orthodox quantum theory, for positions or any other quantum observables, regardless of whether or not these observables refer to a common hypersurface belonging to \mathcal{F} .⁹

This is because the outcomes of all quantum measurements can ultimately be reduced to the orientations of instrument pointers, counter readings, or the ink distribution of computer printouts, if necessary brought forward in time to a common hypersurface in \mathcal{F} , or even to a single common location, for which agreement is assured. Nonetheless, this situation may seem paradoxical if we forget the nonpassive character of measurement in quantum mechanics. The point is that for Bohmian quantum theory, measurement can affect even distant systems, so that the resulting positions—and hence their subsequently measured values—are different from what they would have been had no measurement occurred.

⁸This evolution equation depends upon g only through the areaexpansion factor arising from the normal flow between hypersurfaces, and thus does not really depend upon the choice of coordinates on the hypersurfaces.

⁹This conclusion requires the rather dubious assumption that the relevant measurements can be understood in terms of noninteracting Dirac particles. However, in order to talk coherently about the quantum predictions for a model, it must be possible to understand measurement processes in terms of that model. The remarks we are making here would also be appropriate for the more realistic models for which this would be true.

IV. PERSPECTIVE

We have presented a hypersurface Bohm-Dirac model for *N* entangled but noninteracting Dirac particles. This model is a covariant extension of the Bohm-Dirac model, which involves a foliation by equal-time (flat) hypersurfaces, to arbitrarily shaped (smooth) hypersurfaces. How natural is this model?

When looking for a relativistic extension of nonrelativistic Bohmian mechanics one inevitably encounters two central, very different problems: that such an extension must involve a mechanism for nonlocal interactions between the particles, and that quantum equilibrium cannot hold in all Lorentz frames. For both of these problems the additional space-time structure provided by a foliation yields the most obvious solution: The motion of each particle at a point $x \in M$ depends upon the paths of the other particles via the points at which they intersect the leaf of the foliation containing x, and we have an equivariant density on the leaves of the foliation.

And the simplest way to achieve this, in a covariant manner, for a Dirac wave-function ψ , is via the current [Eq. (12)]: Form the natural tensor $\overline{\psi}\gamma_1 \dots \gamma_N \psi$, evaluated at x and the other intersection points, and contract in the slots corresponding to the other particles with the N-1 unit normals to the hypersurface at the corresponding points, to obtain the divergence-free four-vector j_k , the tangent to the trajectory at x. Thus, the dynamics of the HBD model is the simplest Lorentz-invariant dynamics compatible with the structure at hand, namely, the Dirac wave function and the foliation. Furthermore, the simultaneous normal component $Q = j_k \cdot n_k$ is an equivariant density on the leaves of the foliation.

It should be stressed, however, that the Lorentz invariance of the HBD model is—in Bell's sense—"serious" only if the foliation is regarded as an additional objective *dynamical* (in contrast to absolute) structure in the theory (and in the world, if the theory is to describe the world). It is this structure that is the innovation of what has been proposed here and in [10], not the model per se, which is indeed a rather straightforward covariant extension of the BD model.

However, in this paper we shall not try to find a "serious" law for the foliation \mathcal{F} or, what amounts to the same

thing, its normal vector field *n*. As a toy example, however, the foliation law could be given by an autonomous equation for *n*, such as $\partial_{\nu} n^{\mu} = 0$. Another class of toy examples involves a vector field *n* constructed from the wave function $\psi(x_1, \ldots, x_N)$: Consider the space-time vector fields $v_{kl}^{\mu}(x) = (\bar{\psi} \gamma_k^{\mu} \psi)(\widehat{x_1}, \ldots, \widehat{x_{l-1}}, x, \widehat{x_{l+1}}, \ldots, \widehat{x_N})$, where $(\widehat{x_1}, \ldots, \widehat{x_N})$ is a point fixed in a Lorentz-invariant way, for example, as a maximum of $\bar{\psi} \psi$. [Simply considering $v_k^{\mu}(x)$ $= (\bar{\psi} \gamma_k^{\mu} \psi)(x, \ldots, x)$ is not a good idea, since this will be zero for antisymmetric (fermion) wave functions.] Now one may set *n* equal to the integrable¹⁰ part of some v_{kl} .

A further possibility, which may be more serious, is to have, in addition to the particle degrees of freedom, an independent quantum field ϕ_{μ} that determines the foliation. Assume that for any quantum state Φ of the field, $(\Phi, \phi_{\mu} \Phi)$ is timelike and completely integrable. Then for any state Ψ of the particle-field system, set $n_{\mu} = (\Psi, \phi_{\mu} \Psi)$. Suppose that the particle and the field degrees of freedom are both dynamically and statistically independent, i.e., that there is neither quantum interaction nor entanglement between these degrees of freedom, so that in particular the full wave function $\Psi = \psi \otimes \Phi$. Then we may define the foliation by the normal field n_{μ} . The ϕ field can be regarded as very roughly analogous to a Higgs field, producing a kind of spontaneous symmetry breaking, where by the choice of Φ a particular foliation is determined, and relativistic invariance is thereby broken.

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¹⁰For an arbitrary vector field $v^{\mu}(x)$, the Fourier-transformed $\hat{v}^{\mu}(k)$ may be split into $\hat{v}^{\mu}_{\parallel}(k) = \hat{v}^{\nu}(k)k_{\nu}k^{\mu}/(k_{\lambda}k^{\lambda})$ and $\hat{v}^{\mu}_{\perp}(k) = \hat{v}^{\mu}(k) - \hat{v}^{\mu}_{\parallel}(k)$. The inverse Fourier-transformed $v^{\mu}_{\parallel}(x)$ satisfies the integrability condition $\partial_{\mu}v_{\parallel\nu} - \partial_{\nu}v_{\parallel\mu} = 0$.

- [1] D. Bohm, Phys. Rev. 85, 166 (1952); 85, 180 (1952).
- [2] J. S. Bell, Speakable and Unspeakable in Quantum Mechanics (Cambridge University Press, Cambridge, 1987).
- [3] D. Dürr, S. Goldstein, and N. Zanghì, J. Stat. Phys. 67, 843 (1992).
- [4] J. T. Cushing, A. Fine, and S. Goldstein, *Bohmian Mechanics and Quantum Theory: An Appraisal* (Kluwer Academic, Dordrecht, 1996).
- [5] D. Bohm, Prog. Theor. Phys. 9, 273 (1953).
- [6] D. Bohm and B.J. Hiley, The Undivided Universe: An Ontological Interpretation of Quantum Theory (Routledge, London,

1993).

- [7] T. M. Samols, J. Stat. Phys. 80, 793 (1995).
- [8] K. Berndl, D. Dürr, S. Goldstein, and N. Zanghì, Phys. Rev. A 53, 2062 (1996).
- [9] E. Squires, Phys. Lett. A 178, 22 (1993).
- [10] D. Dürr, S. Goldstein, and N. Zanghì, in *Stochastic Processes, Physics, and Geometry*, edited by S. Albeverio, G. Casati, U. Cattaneo, D. Merlini, and R. Moresi (World Scientific, Singapore, 1990), p. 374.
- [11] T. Maudlin, in *Bohmian Mechanics and Quantum Theory: An Appraisal* (Ref. [4]), p. 285.