

# Single-photon ionization of helium from 4.5 to 12 keV by Compton scattering and the photoelectric effect

D. V. Morgan and R. J. Bartlett

*Los Alamos National Laboratory, Los Alamos, New Mexico 87545*

(Received 22 September 1998; revised manuscript received 11 January 1999)

We have measured the ratio of the cross sections for double-to-single ionization in helium for several monoenergetic photon energies between 4.5 and 12 keV using time-of-flight ion charge state spectroscopy. In this energy range, both the photoelectric effect and inelastic (Compton) scattering contribute significantly to the total cross section. The ionization states caused by Compton scattering were distinguished from those caused by the photoelectric effect by the different recoil energies of the helium ion associated with the two processes. The ratios of the double-to-single ionization cross sections of helium for the photoelectric effect ( $R_p$ ) and for Compton scattering ( $R_c$ ) are given, and compared with previous measurements and theoretical calculations. The measured value for  $R_c$  at 12 keV is  $(1.21 \pm 0.27)\%$ , which agrees well with the theoretical calculations of Andersson and Burgdörfer [Phys. Rev. A **50**, R2810 (1994)]. [S1050-2947(99)05905-3]

PACS number(s): 32.80.Cy, 32.80.Fb

## I. INTRODUCTION

Helium is the simplest neutral atomic system in which electron-electron correlation effects can be studied. The existence of doubly ionized helium from the interaction with a single photon results directly from these correlations for both photoabsorption and inelastic (Compton) scattering. However, the final states associated with double ionization in these two processes differ considerably. For photoabsorption, there is a high probability that the photon energy is transferred almost entirely to the primary photoelectron [1], while for Compton scattering most of the initial x-ray energy is given to the scattered x ray [2], hence the ejected electrons are comparatively slow. At x-ray energies between 4.5 and 12 keV, the strongly energy-dependent photoionization cross section,  $\sigma_p$ , falls rapidly, whereas the Compton scattering cross section,  $\sigma_c$ , is a slowly increasing function of energy [3,4]. At approximately 6.3 keV,  $\sigma_p$  and  $\sigma_c$  are equal, therefore we call this energy range the “crossover” region. Theoretical calculations for the double-to-single ionization ratio in the crossover region, because of the intrinsic difference between Compton scattering and the photoelectric effect, are performed independently. Samson *et al.* [5] have shown that it is necessary to distinguish between these two effects to make a meaningful comparison between theory and experiment at energies where both effects contribute significantly. Here, we report measurements for the double-to-single ionization ratio using monoenergetic photons at energies below the crossover which distinguish between Compton scattering and the photoelectric effect.

## II. EXPERIMENT

The experiments performed at photon energies of 4.5 and 5.5 keV were conducted at the Los Alamos National Laboratory beamline X8A at the National Synchrotron Light Source, which has been described previously [6,7]. This beamline provides monoenergetic x rays using a Si(111) monochromator with a bandpass of approximately 2 eV. The

time-of-flight (TOF) spectrometer used in this experiment has been described in detail previously [6–8]. Helium atoms are injected into the “extraction region” by an effusive gas nozzle. The ambient helium pressure was maintained at  $2.7 \times 10^{-6}$  torr, following evacuation to a base pressure of  $\sim 10^{-9}$  torr. The monochromatic synchrotron beam passes through the extraction region of the TOF spectrometer, in which  $\text{He}^+$  and  $\text{He}^{2+}$  ions accumulate. A pulsed field extracts the ions and accelerates them into a field-free drift tube. The ions are then detected by a dual microchannel plate (MCP) biased at  $-5.46$  kV. The field pulse and the signal from the MCP serve as the start and stop pulses, respectively, for the time-of-flight measurements. The TOF spectrometer axis was aligned parallel to the polarization of the x-ray beam. Because the photoionization process preferentially produces electrons along the polarization axis, the recoil velocity of the photoions either increases or decreases the time of flight relative to an ion produced with little or no recoil (Compton ionization process). Thus, the time-of-flight spectrum has three peaks for both  $\text{He}^+$  and  $\text{He}^{2+}$ : the early and late time peaks are produced by photoions, and the central peak is produced by Compton ions.

Our pulsed-field type of TOF apparatus has collection efficiencies that depend on the speed and trajectory of the recoil ion. This is largely because the rapidly moving ions escape from the interaction region prior to application of the pulsed extraction electric field. In order to determine the collection efficiencies for both the photoelectric and Compton ionization processes, we have employed a Monte Carlo simulation of our TOF spectrometer. The recoil velocities of the different helium ion states are one input to the Monte Carlo simulation. Other inputs include the beam spot size, the pulse widths, pulse period, and amplitudes of the ion extraction electric field, and the TOF spectrometer dimensions. The collection efficiencies for both single ( $\eta_c^+$ ) and double ( $\eta_c^{2+}$ ) ionization from Compton scattering are much higher than the corresponding collection efficiencies for photoionization ( $\eta_p^+$  and  $\eta_p^{2+}$ ), because of the large recoil

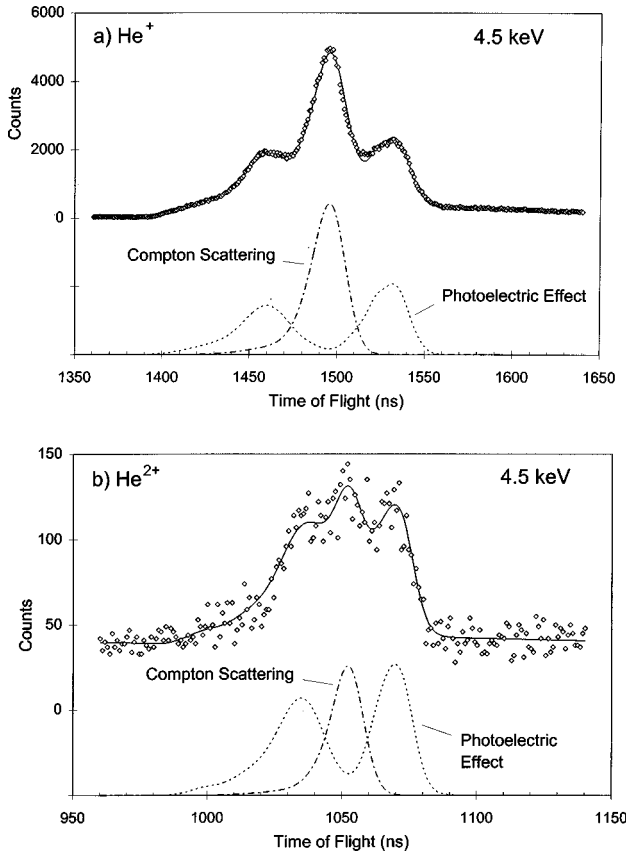


FIG. 1. TOF spectrum ( $\diamond$ ) and least-squares fit to the data (solid line) at 4.5 keV for (a)  $\text{He}^+$  and (b)  $\text{He}^{2+}$ . The Monte Carlo calculated efficiency curves for Compton scattering (dot-dashed line) and the photoelectric effect (dotted line) are also shown.

velocities associated with the photoionization process. Because of the narrower divergence of the  $\text{He}^{2+}$  ions with respect to the axis of the TOF spectrometer,  $\eta_c^{2+}$  is slightly higher than  $\eta_c^+$ . For our calculation of  $\eta_p^+$  and  $\eta_p^{2+}$ , we have assumed that the directional distribution probability of ejected photoions is proportional to  $\cos^2 \theta$ , where  $\theta$  is the angle of the photoion trajectory with respect to the TOF spectrometer axis [9].

The experiment to determine  $R_c$  at 12 keV was performed on the Los Alamos National Laboratory beamline X8C at the National Synchrotron Light Source [10], which provides fo-

TABLE I. Comparison of photoelectric and Compton scattering cross sections in the crossover region.

Photon energy (keV)	$\sigma_c^+/\sigma_p^+$		$\sigma_I/\sigma_p$
	This work	Hino <i>et al.</i> <sup>a</sup>	Viegele <sup>b</sup>
4.0		0.14	0.15
4.5	0.21		0.24 <sup>c</sup>
5.0			0.38
5.5	0.43		0.56 <sup>c</sup>
6.0		0.81	0.83
12.0		10.94	10.83

<sup>a</sup>Reference [11].

<sup>b</sup>Reference [3];  $\sigma_I$  is the incoherent scattering cross section.

<sup>c</sup>Reference [3], determined by cubic interpolation.

TABLE II. Values for  $R_c$  and  $R_p$  obtained in this work.

Photon energy (keV)	$R_c$ (%)	$R_p$ (%)
4.5	$1.07 \pm 0.49$	$2.00 \pm 0.24$
5.5	$0.72 \pm 0.29$	$2.04 \pm 0.37$
12.0	$1.21 \pm 0.27$	

cused monochromatic x rays for energies up to 20 keV, with a bandwidth of approximately 4 eV at 12 keV. The count rate for  $\text{He}^+$  was obtained by subtracting the background counts from the time-of-flight spectrum, then integrating the spectrum over the single ionization flight time. In this energy range, the Compton cross section is larger than the photoionization cross section, and the TOF spectrometer collection efficiency, as determined by the Monte Carlo simulation, is much larger for the Compton scattering process than for the photoionization process. Although both processes contribute to the  $\text{He}^+$  signal, we estimate that the photoabsorption component of the experimental spectrum is only about 1%. The corresponding total collection efficiency for the TOF spectrometer for  $\text{He}^{2+}$  Compton ions was virtually identical to the collection efficiency for  $\text{He}^+$  ions.

### III. RESULTS

Time-of-flight spectra for  $\text{He}^{2+}$  and  $\text{He}^+$  ions are shown in Fig. 1 for 4.5 keV photons. We observe good agreement between the overall shape of the efficiency curves calculated by the Monte Carlo program and the singly ionized helium TOF spectrum. The experimental values for  $\sigma_c^+/\sigma_p^+$  were obtained by a least-squares fit of the amplitude and time-shift parameters of the calculated efficiency curves to the singly ionized peaks in the TOF data. To our knowledge, these are the only experimental values for  $\sigma_c^+/\sigma_p^+$  in the crossover region obtained with monoenergetic x rays. The experimen-

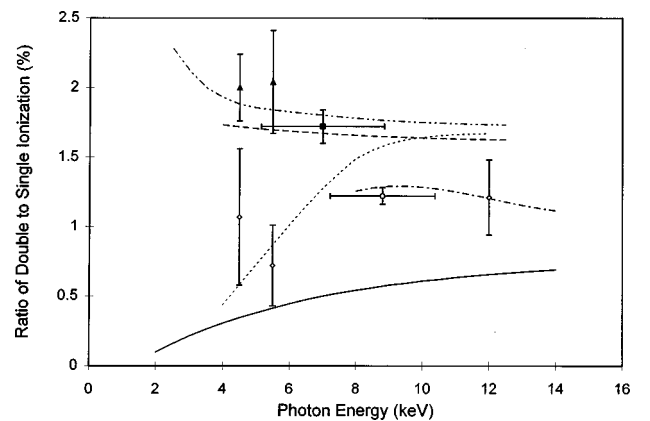


FIG. 2. Ratio of double-to-single helium ionization from Compton scattering ( $R_c$ ) and from the photoelectric effect ( $R_p$ ). Present experimental results for  $R_c$  ( $\diamond$ ) and  $R_p$  ( $\blacktriangle$ ),  $R_c = (1.22 \pm 0.06)\%$  at  $8.8^{+1.5}_{-1.65}$  keV ( $\circ$ ), and  $R_p = (1.72 \pm 0.12)\%$  at  $7.0^{+2.1}_{-1.6}$  keV ( $\blacksquare$ ), obtained by Spielberger *et al.* [9]. Also shown are the calculations of  $R_c$  by Suric *et al.* [14] (solid line), Andersson and Burgdörfer [13] (dot-dashed line), and Hino *et al.* [11] (dotted line), as well as the calculations of  $R_p$  by Andersson and Burgdörfer [20] (two dot-dashed line) and Hino *et al.* [1] (dashed line).

TABLE III. Values for  $R_c$  at 12.0 keV and at the high-energy limit.

$\hbar\omega$ (keV)	$R_c$ (%)							
	Experiment			Theory				
	This work	Spielberger <i>et al.</i> <sup>a</sup>	Wehlitz <i>et al.</i> <sup>b</sup>	Andersson <i>et al.</i> <sup>c</sup>	Suric <i>et al.</i> <sup>d</sup>	Hino <i>et al.</i> <sup>e</sup>	Kornberg <i>et al.</i> <sup>f</sup>	Amusia and Mikhailov <sup>g</sup>
12.0	1.21±0.27			1.2	0.66	1.7		
$\hbar\omega \rightarrow \infty$		0.84 <sup>+0.08</sup> <sub>-0.11</sub>	1.25±0.30	0.84	0.8	1.6	0.8	1.68

<sup>a</sup>Reference [17], measured at  $\hbar\omega = 58$  keV.

<sup>b</sup>Reference [18], measured at  $\hbar\omega = 57$  keV.

<sup>c</sup>Reference [13].

<sup>d</sup>Reference [14].

<sup>e</sup>Reference [11].

<sup>f</sup>Reference [15].

<sup>g</sup>Reference [16].

tal values for  $\sigma_c^+/\sigma_p^+$  at 4.5 and 5.5 keV are compared with the ratio of the incoherent scattering cross section to the photoelectric cross section of Viegeler [3] and with the values for  $\sigma_c^+/\sigma_p^+$  given by Hino *et al.* [11] in Table I. Our results fall somewhat below the values ascribed by Viegeler [3], however this may be caused in part by the inclusion of nonionizing inelastic (Raman) scattering events in the attenuation coefficient calculations which do not contribute to the Compton ionization process [12]. We do not believe significant systematic errors caused by our TOF spectrometer efficiency calculations exist in our measurements of  $R_c$  and  $R_p$  because both  $\eta_p^{2+}/\eta_p^+$  and  $\eta_c^{2+}/\eta_c^+$  are close to unity. The very low number of  $\text{He}^{2+}$  ions produced by Compton scattering observed at 4.5 keV and the reduced peak separation caused by the small photoion recoil velocity at this energy make the uncertainty in our measurement of  $R_c$  at 4.5 keV rather large. Our experimental values for  $R_c$  and  $R_p$  and their estimated uncertainties at 4.5, 5.5, and 12.0 keV are given in Table II. The values for  $R_c$  and  $R_p$  obtained in this experiment are compared with experimental and theoretical results for photon energies in the crossover region in Fig. 2.

#### IV. DISCUSSION

Several calculations of the double-to-single ionization ratio for Compton scattering,  $R_c = \sigma_c^{2+}/\sigma_c^+$ , have been performed from 4 keV to the high-energy (nonrelativistic) limit [2,11,13–16]. These results for 12 keV and the high-energy limit are summarized in Table III. Measurements of  $R_c$  by Spielberger *et al.* [17] at 58 keV give a value of  $R_c = (0.84^{+0.08}_{-0.11})\%$ , which favors the lower calculated values in Table III, and, in particular, those of Andersson and Burgdörfer [13]. Wehlitz *et al.* [18] have measured a higher value,  $R_c = (1.25 \pm 0.30)\%$  at 57 keV, but with a relatively large error bar. This measurement does not favor any of the calculations in Table III. Our value of  $(1.21 \pm 0.27)\%$  at 12 keV, even with the large error bar, agrees only with the calculations of Andersson and Burgdörfer [13]. At 4.5 and 5.5

keV our results for  $R_c$  are consistent with the calculations of Hino *et al.* [11], but somewhat higher than the values calculated by Suric *et al.* [14].

For the photoelectric case, theoretical studies [1,19,20] generally agree, with the exception of Drukarev [21], that the double-to-single ratio for the photoelectric effect,  $R_p = \sigma_p^{2+}/\sigma_p^+$ , decreases asymptotically to a value of approximately 1.67 in the high-energy limit. Recently, Spielberger *et al.* [9] have obtained values for  $R_c$  and  $R_p$  in the crossover region using cold target ion momentum spectroscopy (COLTRIMS), which is quite similar to the method used here and described above. Using broadband synchrotron radiation, these authors have separated the components of the photoelectric effect from Compton scattering to obtain values for both  $R_c$  and  $R_p$ . They have measured  $R_c = (1.22 \pm 0.06)\%$  at a mean energy of 8.8 keV and  $R_p = (1.72 \pm 0.12)\%$  at a mean energy of 7.0 keV for photoabsorption. This is the highest energy for which an experimental value for  $R_p$  has been obtained. Our monoenergetic values for  $R_p$  at 4.5 and 5.5 keV, shown in Fig. 2, are consistent with both the theoretical calculations and the experimental work of Spielberger *et al.* [16].

#### V. CONCLUSION

This paper reports monoenergetic x-ray measurements of  $R_c$  and  $R_p$  as well as experimental values for  $\sigma_c^+/\sigma_p^+$  for helium in the crossover region. In addition, our measurement of  $R_c$  at 12.0 keV and that of Spielberger *et al.* [17] at 58 keV are in agreement with the calculations of Andersson and Burgdörfer [13], and together favor this calculation over all others reported.

#### ACKNOWLEDGMENTS

This work was supported by the U.S. Department of Energy through Los Alamos National Laboratory. We would like to thank M. Sagurton for his contributions to this work.

- [1] K. Hino, T. Ishihara, F. Shizumi, N. Toshima, and J. H. McGuire, *Phys. Rev. A* **48**, 1271 (1993).
- [2] P. M. Bergstrom, K. Hino, and J. H. Macek, *Phys. Rev. A* **51**, 3044 (1995).
- [3] W. T. Viegele, *At. Data* **5**, 51 (1973).
- [4] J. A. R. Samson, *Phys. Rev. Lett.* **72**, 3329 (1994).
- [5] J. A. R. Samson, C. H. Greene, and R. J. Bartlett, *Phys. Rev. Lett.* **71**, 201 (1993).
- [6] M. Sagurton, R. J. Bartlett, J. A. R. Samson, Z. X. He, and D. Morgan, *Phys. Rev. A* **52**, 2829 (1995).
- [7] D. V. Morgan, R. J. Barlett, and M. Sagurton, *Phys. Rev. A* **51**, 2939 (1995).
- [8] D. V. Morgan, M. Sagurton, and R. J. Bartlett, *Phys. Rev. A* **55**, 1113 (1997).
- [9] L. Spielberger, O. Jagutzki, B. Krässig, U. Meyer, Kh. Khayyat, V. Mergel, Th. Tschentscher, Th. Buslaps, H. Bräuning, R. Dörner, T. Vogt, M. Achler, J. Ullrich, D. S. Gemmell, and H. Schmidt-Böcking, *Phys. Rev. Lett.* **74**, 4615 (1995).
- [10] R. W. Alkire, M. Sagurton, F. D. Michaud, W. J. Trela, R. J. Bartlett, and R. Rothe, *Nucl. Instrum. Methods Phys. Res. A* **352**, 535 (1995).
- [11] K. Hino, P. M. Bergstrom, and J. H. Macek, *Phys. Rev. Lett.* **72**, 1620 (1994).
- [12] A. Sommerfeld, *Phys. Rev.* **50**, 38 (1936).
- [13] L. Andersson and J. Burgdörfer, *Phys. Rev. A* **50**, R2810 (1994).
- [14] T. Suric and K. Pisk, *Phys. Rev. Lett.* **73**, 790 (1994).
- [15] M. A. Kornberg and J. E. Miraglia, *Phys. Rev. A* **53**, R3709 (1996).
- [16] M. Ya. Amusia and A. I. Mikhailov, *Phys. Rev. Lett.* **71**, 50 (1993).
- [17] L. Spielberger, O. Jagutzki, B. Krässig, U. Meyer, Kh. Khayyat, V. Mergel, Th. Tschentscher, Th. Buslaps, H. Bräuning, R. Dörner, T. Vogt, M. Achler, J. Ullrich, D. S. Gemmell, and H. Schmidt-Böcking, *Phys. Rev. Lett.* **76**, 4685 (1996).
- [18] R. Wehlitz, R. Hentges, G. Prümper, A. Farhat, T. Buslaps, N. Berrah, J. C. Levin, I. A. Sellin, and U. Becker, *Phys. Rev. A* **53**, R3720 (1996).
- [19] A. Dalgarno and H. R. Sadeghpour, *Phys. Rev. A* **46**, R3591 (1992).
- [20] L. R. Andersson and J. Burgdörfer, *Phys. Rev. Lett.* **71**, 50 (1993).
- [21] E. G. Drukarev, *Phys. Rev. A* **51**, R2684 (1995).