Overlapping sequence of doubly excited series in the vuv spectrum of argon

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Observations of overlapping sequences of doubly excited resonances in the absorption spectrum of argon are reported at improved resolution, exhibiting more detail than earlier studies. Their appearance is compared with the doubly excited manifolds recently reported in the spectrum of He by Domke et al. [Phys. Rev. Lett. 66, 1306 (1991)] and by Zubek et al. [J. Phys. B 24, L342 (1991)]. The salient differences are pointed out and their origin is explained. The simplest case of overlapping channels has been accounted for by Lane's theory [J. Phys. B 17, 2213 (1984)] for one broad intruder in a Rydberg series of autoionizing resonances. Numerous interesting examples of the q-reversal effect are uncovered.

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Double excitations appear in the photoabsorption spectra of atoms as families of series, which become more numerous and more crowded as the double-ionization threshold is approached. In He, the double-ionization threshold stands well above all single excitations, so that, although the doubly excited spectrum contains much detailed structure, its overall appearance is quite regular, as demonstrated by recent data (Domke et al. [I] and Zubek et al. [2]).

As one moves to heavier rare gases, there are many reasons for which this overall regularity can be lost. First, the manifold of doubly excited channels moves down closer to the single excitations, which themselves become more numerous owing to excitations from inner shells. Thus, single and double excitations are no longer well separated in energy. This has two consequences: first, there is usually an enhancement of double-excitation spectra at the expense of the single excitations, and second, both classes of spectra lose some of their regularity.

Another effect to be considered is the appearance of multiplet structure in the parent ion. Indeed, helium is the only rare gas for which the parent ion is hydrogenic, so that, with increasing excitation energy, doubly excited manifolds converge on successive, initially well-separated thresholds of different principal quantum number N of the parent ion. In all other rare gases, multiplet structure due to the open $p⁴$ subshell will be present even for the lowest excitations. Thus, in all other rare gases, overlapping resonances occur even amongst the lowest doubleexcitation channels, which are relatively unperturbed in He.

A further complication is that, as the onset of orbital contraction in the first long periods is approached, the doubly excited spectra involve both *ns* and *nd* electrons, whose behavior in homologous sequences exhibits the

systematic trends characteristic of orbital collapse (Connerade [3]). A case of particular interest in this respect is Ar, because it lies just before the first transition sequence.

A combination of high experimental resolution and broad coverage of a wide spectral range, such as can only be achieved with synchrotron radiation, is required in order to study doubly excited manifolds of Rydberg series. In the present paper, we report observations over the full energy range of the parent-ion terms $3s^23p^4(^1S,{}^3P,$ and $1D$)4s, 4p, and 3d. Interestingly, not all possible manifolds are observed: the spectrum is dominated by a small number of channels.

Experiments were performed at the 2.5-GeV electron synchrotron of the Physikalisches Institut in Bonn, using a high-resolution 3-m normal-incidence spectrograph equipped with a 5000 line/mm holographic grating. Details of the apparatus and techniques have previously been published (Baig et al. [4—6]). Wavelength calibration was obtained by comparison with the sharp lines of the same spectrum as previously recorded by Baig and Ohno [7] and Madden et al. [8]. The absolute accuracy of the observed spectral lines is ± 0.008 Å achieved by using the third-order Chebychev-polynomial fit. The spectra recorded on the photographic plates were digitized with a computer-controlled microdensitometer. The transmission of the blackened photographic plate was measured in steps of 5 μ m using a slit width of 10 μ m at the photomultiplier. One photographic plate of 20-cm length covers a spectral range of roughly 120 A. The digitized spectrum therefore consists of about 40000 data points. As the response curve of the photographic emulsion is nonlinear, the plotted profile is not proportional to the absorption cross section. Relative intensities are, however, correct when lines on the same exposure are compared which lie on the linear part of the response curve.

FIG. 1. The absorption spectrum of Ar I, showing the complete spectrum covering the spectral region between 320 and 400 Å. The identifications for the resonance between 370 and 400 \AA are taken from Madden *et al.* [8].

The present spectrum provides much improved detail as compared with the earlier work of Madden et al. [8]. The experimental spectra we have obtained are shown in Figs. 1-3. From our observations and the series limits listed in the tables of Moore [9], fairly constant quantum defects for the different channels have been extracted, which allow quite reliable groupings into Rydberg series to be achieved, once account is also taken of the pronounced intensity perturbations which the broadest

lowest members of each channel produce amongst the higher members of the next channel down in energy, a situation strongly reminiscent of the interactions found by Domke et al. [1] between the $N = 5$ and 6 doubly excited channels of He.

The 3s inner-shell excitation spectra

$$
3s^23p^6(^1S_0) \longrightarrow 3s(^6S_{1/2})3p^6np^{-1}P_1,
$$

exhibiting window-type autoionizing resonances, were re-

FIG. 2. The absorption spectrum of Ar I, showing the double-excited Rydberg series terminating to four high-lying limits (see text for details). The intensity variations are due to the low-lying members of the series belonging to the other channels, particularly around 367, 360, and 354 A.

TABLE I. Broad resonances series; the leading member was picked up by looking at the interesting variation around 366 A. $3s²3p⁶ S₀ \rightarrow 3s²3p⁴(¹S₀)4s(²S_{1/2})np¹P$

nl	Wavelength (Ă)	Wave number (cm^{-1})	n^*
4p	366.2^a	273075	2.324
5p	352.2^a	283930	3.351
6p	346.428 ^b	288 660.3	4.365
7p	344.202	290 527.1	5.310
8p	342.858	291 666.0	6.314
9p	342.034	292 368.6	7.316
10p	341.496	292829.2	8.309
11p	341.098	293 170.9	9.378
12p	340.842	293 391.1	10.334
Limit		294418.56	

'Overlapping.

Very weak.

ported by Baig and Ohno [7]. In the present work, the following two-electron transitions are considered:

$$
3s23p6(1S0) \Rightarrow 3s23p4(3P, 1D2, 1S0)3dnp, nf
$$

\n
$$
\Rightarrow 3s23p4(3P, 1D2, 1S0)4snp
$$

\n
$$
\Rightarrow 3s23p4(3P, 1D2, 1S0)4pns, nd.
$$

A remarkable phenomenon is the following: although, within the energy range presently studied, there are 24 terms of the $3s^23p^44s$, 4p, and 3d complex, which can in principle produce a vast number of accessible series limits of different J values; Rydberg sequences to only nine distinguishable thresholds have been observed. Again, one can perhaps dram a comparison with He, for which unexpected simplicity of structure is found in the doubly excited spectrum (Fano [10]).

More specifically, we observe one series attaching to the term $3s^{2}3p^{4}({}^{1}S_{0})4s({}^{2}S_{1/2})$, which must be an np series with $J=1$ and is the only possible combination with the ground state for this term of the parent ion. The data for this series are tabulated in Table I. For the $3s²3p⁴(¹D₂)4s²D$ term, we only observe one nf series to

TABLE II. Window-type autoionizing resonance series, $3s^23p^6{}^1S_0\rightarrow 3s^23p^4{}(^1D)4s(^2D_{3/2})nf^1P_1.$

nl	Wavelength (Ă)	Wave number $\rm (cm^{-1})$	n^*	,,, 3d 4d
4f	371.843	268 930.7	4.017	5d
5f	368.705	271219.6	4.932	6d
6f	366.746	272 668.3	5.986	7d
7f	365.667	273472.9	6.971	8d
8f	364.960 ^a	274 002.7	7.968	9d
9f	364.483	274 361.2	8.951	10d
10f	364.133	274 625.0	9.961	11d
11f	363.858	274 832.5	11.052	12d
Limit		275 730.88		Limi

'Overlapping.

TABLE III. This series shows anomalous intensity behavior at high *n* values ($11 \le n \le 22$) which is due to mixing with the autoionizing resonance around 377 Å. Madden et al. [8] have assigned this resonance as $({}^{1}D_{2})3d({}^{2}F_{5/2})4p$. $3s^23p^6{}^1S_0 \rightarrow 3s^23p^4{}^3P_0)4s(^2P_{3/2})np(^1P_{11}.$

nl	Wavelength (\check{A})	Wave number (cm^{-1})	$\boldsymbol{n}^{\,*}$
5p	392.726	254 630.5	3.199
6p	386.375	258815.9	4.097
7p	382.872	261 183.9	5.130
8p	381.068	262 420.4	6.116
9р	380.034	263 134.4	7.031
10p	379.275	263 661.0	8.050
11p	378.786	264 001.3	9.005
12p	378.410	264 263.6	10.030
13p	378.137	264454.5	11.042
14p	377.898	264 621.7	12.237
15p	377.727	264 741.5	13.380
16p	377.622	264815.1	14.264
17p	377.550	264 865.6	14.983
18p	377.444	264 940.0	16.272
19p	377.384	264 982.1	17.168
20p	377.328	265021.4	18.153
21p	377.274	265059.4	19.286
22p	377.241	265 082.5	20.090
23p	377.206	265 107.1	21.066
24p	377.172	265 131.0	22.164
Limit		265 354.41	

the (${}^2D_{3/2}$) limit, although an *np* electron attaching to either the ${}^{2}D_{3/2}$ or ${}^{2}D_{5/2}$ limit could result in accessible levels. The tentative level assignments are based on the observed quantum-defect values listed in Table II.

Turning now to the $3s^23p^4(3P)$ -based limits, we again find one np series to the limit involving a 4s electron, namely $3s^23p^4(^3P_0)4s(^2P_{3/2})$, which is listed in Table III. For the limits involving a $4p$ electron, two series are found, one nd series to $3s^2\frac{3p^4(3p)}{4p^2}$ (see Table

TABLE IV. Window-type autoionizing resonances series; assignment as $ns\left[\frac{5}{2}\right]$ series is discarded as is does not combin with the ground-state ${}^{1}S_{0}$. The 4d ${}^{1}P_{1}$ member lies at the threshold of $3s^23p^4(^3P_1)3d(^2D_{5/2})nf$ series. $3s^23p^6{}^1S_0$ \rightarrow 3s²3n⁴(³P)4n(²D_{5in})nd¹P

nl	Wavelength (Ă)	Wave number (cm^{-1})	n^*
3d	371.843	268 930.7	2.547
4d	360.86^a	277 114.3	3.546
5d	356.381	280 598.6	4.575
6d	354.339	282 215.6	5.502
7d	352.994	283 291.0	6.560
8d	352.211	283920.7	7.559
9d	351.674	284 354.3	8.591
10d	351.315	284 644.8	9.578
11d	351.055	284855.7	10.553
12d	350.859	285014.8	11.524
Limit		285841.10	

TABLE V. Asymmetric autoionizing resonance series, $3s²3p⁶1S₀ \rightarrow 3s²3p⁴(³P)4p⁴P_{5/2})nd¹P₁$.

nl	Wavelength $\tilde{\textbf{A}}$	Wave number (cm^{-1})	n^*	nl
3d	372.07	268 768.1	2.960	4f
4d	364.562	274 301.8	3.963	5f
5d	361.213	276845.0	4.969	6f
6d	359.475	278 183.5	5.944	7 f
7d	358.389	279 026.4	6.963	8f
8d	357.705	279 560.0	7.965	9f
9d	357.237	279 926.2	8.971	10f
10d	356.905	280186.6	9.973	11f
11d	356.656	280382.2	10.996	
				Limit
Limit		281 289.79		^a Overlappin

IV), one *nd* series to $3s^23p^4(^3P)4p(^4P_{5/2})$, listed in Table **V** and none to any $J_c = \frac{1}{2}$ or $\frac{3}{2}$ limits. We attribute the series to the excitation of an nd electron because the lowest effective quantum number is nearly equal to 3.

For the limits involving a 3d electron, we find series to four limits: $3s^23p^4(^3P)3d(^2D_{5/2}, {}^{2}F_{5/2}, {}^{2}P_{3/2},$ and $P_{1/2}$; again, there is one series to each limit. Wave lengths of all these observed series are listed in Tables VI—IX, respectively. An interesting observation is that for a Rydberg series attaching to high angular momentum of the core $({}^2D_{5/2}$ or ${}^2F_{5/2}$), the effective quantum numbers are nearly hydrogenic (see Tables VI and VII), whereas for lower angular momentum of the core $({}^2P_{3/2})$ or ${}^{2}P_{1/2}$), the effective quantum numbers are as one would expect for a penetrating np series (see Tables VIII and IX).

It is interesting that we observe series to all possible terms involving a 4s electron, which indicates that the excitation processes from the $3s²3p⁶1S₀$ ground state favor doubly excited configurations involving 4s over 4p or 3d. In short, we are finding far fewer series than the selection rules allow, but it is not obvious what new rules should be used to predict which are the strong double excitations.

TABLE VI. Weak window-type autoionizing resonance series, $3s^23p^6{}^1S_0 \rightarrow 3s^23p^4{}^3P)3d(^2D_{5/2})nf^1P_1$.

nl	Wavelength (\tilde{A})	Wave number (cm^{-1})	n^*
4f	368.279	271 533.3	4.058
5f	365.179	273838.3	5.017
6f	363.396	275 181.9	6.032
7f	362.327	275993.8	7.056
8f	361.637	276 520.4	8.087
9f	361.213	276845.0	9.005
10 f	360.862	277 114.3	10.062
11f	360.603	277 313.3	11.136
12f	360.430	277446.4	12.082
13f	360.291	277 553.5	13.047
Limit		278 198.08	

TABLE VII. Window-type autoionizing resonance series, $3s²3p⁶$ ¹S₀ \rightarrow $3s²3p⁴(³P)3d(²F_{5/2})nf¹P₁$.

Wavelength (Å)	Wave number (cm^{-1})	n^*
371.279	269 339.2	3.973
367.975	271757.6	4.920
366.082	273 162.9	5.924
364.960^a	274 002.7	6.926
364.224	274 556.3	7.956
363.738	274923.2	8.960
363.396	275 181.9	9.951
363.146	275 371.4	10.929
	276 290.08	

TABLE VIII. Series $3s^23p^6{}^1S_0 \rightarrow 3s^23p^4{}^3P\cdot 3d\cdot ({}^2P_{3/2})np^1P_1$.

nl	Wavelength (\check{A})	Wave number $\rm (cm^{-1})$	n^*
4p	396.7	252080	2.302
5p	381.358	262 220.8	3.224
6p	374.723^a	266863.8	4.307
7p	371.843	268 930.7	5.339
8p	370.319	270037.5	6.326
9p	369.313	270 773.0	7.395
10p	368.705	271 219.6	8.387
11p	368.279	271 533.3	9.383
12p	367.975	271757.6	10.361
13p	367.742	271929.8	11.362
14p	367.570	272057.0	12.322
15p	367.427	272 162.9	13.338
16p	367.310	272 249.6	14.387
17p	367.224	272 313.4	15.339
Limit		272 779.74	

'Very weak.

TABLE IX. Series $3s^23p^6{}^1S_0 \rightarrow 3s^23p^4(^3P)3d(^2P_{1/2})np^1P_1$.

	Wavelength	Wave number	
nl	$(\check{\mathbf{A}})$	$\rm (cm^{-1})$	n^*
4p	393.065	254410.9	2.511
5p	380.001 ^a	263 157.2	3.559
6p	375.373 ^a	266401.7	4.500
7p	372.866	268 192.9	5.500
8p	371.378	269 267.5	6.556
9p	370.574	269851.6	7.465
10p	369.934	270318.5	8.547
11p	369.520	270621.4	9.565
12p	369.223	270839.0	10.572
13p	368.994	271007.1	11.613
14p	368.848	271 114.4	12.464
15p	368.705	271219.6	13.510
Limit		271820.80	

^aVery weak.

A remarkable effect in the spectrum is the approximate alternation of the asymmetry parameters ^q between successive series as the highest observed parent-ion state $3s²3p⁴(¹S₀)$ 4s is approached from below. Each series can be characterized by a given average q value. Thus, for example, the highest energy $3s^23p^4(^1S_0)4s(^2S_1/2)np$ series possesses a large effective q value, corresponding to nearly asymmetric absorption peaks. The next down in energy, $3s^23p^4(^3p)4p(^2D_{5/2})nd$ series, possesses a small mean value of q, characteristic of nearly symmetric windows. The next one down, $3s^23p^4(^3P)4p(^4P_{5/2})nd$ series, has negative q , with a pronounced q reversal (Connerade [11], Baig and Connerade [12] and Connerade et al. [13]), giving positive q for the lower member. This q reversal is especially interesting as an example of a case in which the perturber is actually a window resonance, for which q is nearly zero. Within the perturbed series converging on $3s^23p^4(^3P)4p(^4P_{5/2})$, there is the further remarkable effect that the higher members are broad while the lower members are sharp. One could seek to explain this as resulting from an increase in the autoionization rate due to the presence of a new continuum above the $3s²3p⁴(³P)3d(²D_{5/2})$ threshold, but it seems more likely, in view of the change in symmetry just referred to and the absence of similar broadenings as the other thresholds in the spectrum are crossed, that the effect is in fact a fiuctuation in width of the type discussed by Connerade and Lane [14].

Indeed, the approach of Lane [15] is a useful framework within which to describe the kind of interactions between autoionizing resonances which are observed in the present spectrum. Thus, the lowest-energy series (where the interactions are simplest) has been fitted to the relation (Lane [15])

$$
\sigma\!=\!\sigma_0(1\!-\!\rho_n^2)\!+\!\sigma_0\!\rho_n^2\frac{(q_n\!+\!\epsilon_n^{})^2}{1\!+\!\epsilon_n^2}\ ,
$$

where

$$
\varepsilon_n = \frac{[E - E_n - F_n(E)]}{\frac{1}{2} \Gamma_n},
$$
\n
$$
q_n = \frac{(1 + \varepsilon_{nB}^2)Z - (s + q_B^{-1}) + \varepsilon_{nB} (1 - s q_B^{-1})}{(1 + \varepsilon_{nB}^2)u + \varepsilon_{nB} (s + q_B^{-1}) + (1 - s q_B^{-1})},
$$
\n
$$
\rho_n = \frac{(1 + \varepsilon_{nB}^2)u + \varepsilon_{nB} (s + q_B^{-1}) + (1 - s q_B^{-1})}{\{(\Gamma_n / \Gamma_n^0)(1 + \varepsilon_{nB}^2)\left[W^2(1 + \varepsilon_{nB}^2) + (1 + 2\varepsilon_{nB} q_B^{-1} - q_B^{-2})\right]\}^{1/2}}
$$
\n
$$
F_n = \Gamma_{n\frac{1}{2}t}^{0.1}t^2(1 + \varepsilon_{nB}^2)^{-1}[(1 - s^2)\varepsilon_{nB} - 2s].
$$

FIG. 3. The absorption spectrum of argon showing an enlarged portion of the spectrum between 360 and 375 A. The Rydberg series converging to five limits are marked.

TABLE X. Values of parameters as obtained from our fit. Index B indicates values for the broad underlying state. The width of the Rydberg-series members is calculated after $\Gamma_n^0 = \Gamma^0 / (n^*)^3$.

Parameter	Value	
$E_{\rm lim}$	265325.26 cm ⁻¹	
	1.898	
$\frac{\mu}{\Gamma^0}$	76014.12 cm ⁻¹	
\boldsymbol{u}	-17.52	
S	1.262×10^{-8}	
t	9.229×10^{-5}	
W	21.52	
z	2.00	
σ_0	44 976.764	
E_B	264807 cm ⁻¹	
Γ_B	2372 cm^{-1}	
q_B	-0.047	

This is valid in the presence of several continua for narrow resonances and one broad state, by summing over a finite sequence of resonances and assuming their unperturbed oscillator strengths decrease as $1/n³$. The result of the fit is shown in Fig. 4 whereas the parameters obtained are listed in Table X.

It is interesting to compare our conclusions with the photoelectron-satellite studies by Wills et al. [16], which show pronounced resonance structure in the relative intensities of the satellites which arise by decay into the parent ionic states. Unfortunately, the band pass of the primary photon excitation in the photoelectron study is only 75 meV. Thus, the series in their spectra can only be followed up to the fourth members whereas, in the most favorable cases, we follow series up to the 19th member. For this reason, most of the complex interactions we describe are not visible in the photoelectron data.

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- [1] M. Domke, C. Xue, A. Puschmann, T. Mandel, E. Hudson, D. A. Shirley, G. Kaindl, C. H. Greene, H. R. Sadeghpour, and H. Petersen, Phys. Rev. Lett. 66, 1306 (1991).
- [2] M. Zubek, C. Dawber, R. I. Hall, L. A. Avaldi, K. Ellis, and G. C. King, J. Phys. B24, L342 (1991).
- [3] J. P. Connerade, J. Phys. B 24, 605 (1991).
- [4] M. A. Baig, J. P. Connerade, C. Mayhew, G. Noeldeke, and M. Seaton, J. Phys. B 17, L383 (1984).
- [5] M. A. Baig and J. P. Connerade, J. Phys. B 17, ¹⁷⁸⁵ (1984).
- [6] M. A. Baig, A. Rashid, I. Ahmad, M. Rafi, J. P. Connerade, and J. Hormes, J. Phys. B23, 3489 (1990).
- [7] M. A. Baig and M. Ohno, Z. Phys. D 3, 369 (1986).
- [8]R. P. Madden, D. L. Ederer, and J. Codling, Phys. Rev. ^A 177, 136 (1969).

FIG. 4. Densitometer trace of the series converging to the $3s²3p⁴(³P)4s(²P_{3/2})$ limit. The intensity variation between $n = 11, \ldots, 22$ is due to the $({}^{1}D_2)3d({}^{2}F_{5/2})4p$ perturber. The upper part of the figure is the fitted profile for an autoionizing Rydberg series interacting with a broad feature using Lane's formalism. The parameters are given in the text.

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- [9] C. E. Moore, Atomic Energy Levels, Natl. Bur. Stand. Ref. Data Ser., Natl. Bur. Stand. (U.S.) Circ. No. 35 (U.S. GPO, Washington, DC, 1971), Vol. I.
- [10] U. Fano, Rep. Proc. Phys. 46, 97 (1983).
- [11] J. P. Connerade, Proc. R. Soc. London Ser. A 362, 361 (1978).
- [12]M. A. Baig and J. P. Connerade, J. Phys. B 18, 3487 (1985).
- [13] J. P. Connerade, A. M. Lane, and M. A. Baig, J. Phys. B 1S, 3507 (1985).
- [14] J. P. Connerade and A. M. Lane, J. Phys. B 20, 1757 (1987).
- [15] A. M. Lane, J. Phys. B 17, 2213 (1984).
- [16] A. A. Wills, A. A. Cafolla, F. I. Currell, J. Corner, A. Svensson, and M. A. MacDonald, J. Phys. B 22, 3217 (1989).