Relativistic many-body calculations of transition rates for Ca⁺, Sr⁺, and Ba⁺

C. Guet and W.R. Johnson*

Département de Recherche Fondamentale, Service de Physique Atomique, Centre d'Etudes Nucléaires de Grenoble,

85X, 38041 Grenoble CEDEX, France

(Received 8 February 1990; revised manuscript received 6 March 1991)

Relativistic many-body theory is applied to determine amplitudes for $4p \rightarrow 4s$, $4p \rightarrow 3d$, and $3d \rightarrow 4s$ transitions in Ca⁺; $5p \rightarrow 5s$, $5p \rightarrow 4d$, and $4d \rightarrow 5s$ transitions in Sr⁺; and $6p \rightarrow 6s$, $6p \rightarrow 5d$, and $5d \rightarrow 6s$ transitions in Ba⁺. The calculations, which are carried out through third order in perturbation theory, give lifetimes for excited p states that agree with measured values to better than 10%. The agreement improves to the 2% level after semiempirical corrections are included. The lifetimes predicted for the metastable $4d_{3/2}$ and $4d_{5/2}$ states in Sr⁺ and for the $5d_{5/2}$ state in Ba⁺ also agree well with measurement, but the theoretical lifetime of the $5d_{3/2}$ state in Ba⁺ is a factor of 4 larger than the measured lifetime.

I. INTRODUCTION

In the present paper, relativistic many-body perturbation theory (MBPT) calculations of the lifetimes of lowlying excited states of the alkaline-earth-metal ions Ca⁺ Sr⁺, and Ba⁺ are compared with accurate experimental lifetimes as a test of the quality of theoretical methods for atomic systems with high nuclear charge. The lifetime of the $6p_{3/2}$ state of Ba⁺ [1], which has been measured to 0.3% accuracy, provides an especially valuable comparison standard. Lifetimes of the $6p_{1/2}$ state of Ba⁺ [2], the $5p_{1/2}$ and $5p_{3/2}$ states of Sr⁺ [2], and the $4p_{1/2}$ and $4p_{3/2}$ states of Ca⁺ [3], all of which have been measured to 1% accuracy, provide other useful comparison values. We carry out MBPT calculations through third order for the amplitudes of $p \rightarrow s$ and $p \rightarrow d$ transitions in the three alkaline-earth-metal ions, Ca⁺, Sr⁺, and Ba⁺. From these calculations, we can predict lifetimes of the lowest excited $p_{1/2}$ and $p_{3/2}$ states for each ion. Following the standard prescription of MBPT [4], we obtain theoretical values for the lifetimes of the excited p states that agree with experimental lifetimes to better than 10%.

It has been shown in theoretical studies of the $6p \rightarrow 6s$ transition amplitudes in cesium [5], that a semiempirical correction, in which the valence-electron self-energy operator is scaled to give the experimental correlation energy, leads to theoretical lifetimes that agree very closely with precise measurements. We make semiempirical corrections of this type in the present paper, and obtain lifetimes of the excited p states of Ca⁺, Sr⁺, and Ba⁺ that agree with measurement to better than 2%.

The techniques used to study the allowed E1 transitions $p \to s$ and $p \to d$ are also used to study the forbidden E2 transitions $d \to s$ in the alkaline-earth-metal ions. Here the experimental measurements [6–9], which are made in ion traps, are less precise. Interestingly, we obtain theoretical values that agree well with measurement for the lifetime of the $5d_{5/2}$ state of Ba⁺, and for the $4d_{3/2}$ and $4d_{5/2}$ states of Sr⁺; but the theoretical lifetime for the $5d_{3/2}$ state of Ba⁺ is found to be in serious disagreement with the measured lifetime [9].

Comparisons between ab initio calculations and precisely measured atomic properties of heavy atoms are particularly important in view of recent efforts to determine weak-interaction coupling constants from measured parity-nonconserving (PNC) amplitudes [10-12]. Measurements of the $7s \rightarrow 6s$ PNC amplitudes in cesium, accurate to 2%, have been made recently [10]. Both experimental and theoretical determinations of PNC amplitudes are needed in order to extract the weak coupling constants from such measurements; the accuracy which can be obtained for the coupling constants is limited by the accuracy of the theoretical transition amplitudes. Indeed, the need for accurate theoretical transition amplitudes to interpret experiments on weak interactions in heavy atoms has been an important stimulant to research on relativistic many-body theory [13-15]. To determine the accuracy of the theoretically calculated PNC amplitudes, one typically calculates amplitudes for allowed transitions and hyperfine constants of low-lying states and compares these with precise experiments. Thirdorder perturbation theory calculations of allowed transition amplitudes and hyperfine constants in cesium are found to agree with measurements to about 3% [4]; moreover, when the third-order self-energy operator is scaled to give the measured correlation energy, the agreement improves to better than 1% [5].

II. THEORY

Transition amplitudes in the present calculation are evaluated through third order in perturbation theory. The first-order calculations are independent-particle calculations made using Dirac-Hartree-Fock (DHF) wave functions. We solve the DHF equations for the closed ionic core and obtain a frozen-core Hartree-Fock (HF) potential $V_{\rm HF}$. This potential is used to construct a oneelectron Hamiltonian $h_{\rm HF}$. The valence DHF equations,

$$h_{\rm HF} \, v_a = \varepsilon_a v_a, \tag{1}$$

are then solved for the orbitals v_a and eigenvalues ε_a of states of interest. For example, Eq. (1) is solved to give the $6s_{1/2}$, $6p_{1/2}$, $6p_{3/2}$, $5d_{3/2}$, and $5d_{5/2}$ states in the case of Ba⁺. The DHF eigenvalue ε_a is the zeroth-order approximation to the ionization energy of the state a,

$$E_a^{(0)} = \varepsilon_a. \tag{2}$$

Once the valence orbitals have been calculated, the firstorder transition amplitudes are found by evaluating the reduced matrix elements of the dipole operator:

$$D_{ba}^{(1)} = \langle v_b || ez || v_a \rangle. \tag{3}$$

Second-order calculations give the core-polarization corrections to the first-order independent-particle amplitudes. The second-order transition amplitude is given by

$$D_{ba}^{(2)} = \sum_{c,n} \frac{D_{cn}^{(1)}(g_{bnac} - g_{bnca})}{\varepsilon_c - \varepsilon_n - \omega} + \sum_{c,n} \frac{(g_{bcan} - g_{bcna})D_{nc}^{(1)}}{\varepsilon_c - \varepsilon_n + \omega}.$$
(4)

In Eq. (4), the summation index c extends over all occupied core states and the index n extends over all positiveenergy excited states. The quantity ω is the excitation energy $\omega = \varepsilon_b - \varepsilon_a$. The quantities g_{abcd} in Eq. (4) are unsymmetrized two-particle Coulomb matrix elements:

$$g_{abcd} = \iint \frac{1}{|\mathbf{r}_1 - \mathbf{r}_2|} v_a^{\dagger}(\mathbf{r}_1) v_c(\mathbf{r}_1) v_b^{\dagger}(\mathbf{r}_2) v_d(\mathbf{r}_2) \times d^3 r_1 d^3 r_2.$$
(5)

We include the higher-order core-polarization corrections associated with the random-phase approximation (RPA) in our second-order amplitude. This is done by replacing the first-order amplitudes $D_{cn}^{(1)}$ and $D_{nc}^{(1)}$ on the right-hand side of Eq. (4) with the corresponding RPA amplitudes D_{cn}^{RPA} and D_{nc}^{RPA} obtained by solving the core RPA equations

$$D_{an}^{\text{RPA}} = D_{an}^{(1)} + \sum_{b,m} \frac{D_{bm}^{\text{RPA}}(g_{amnb} - g_{ambn})}{\varepsilon_b - \varepsilon_m - \omega} + \sum_{b,m} \frac{(g_{abnm} - g_{abmn})D_{mb}^{\text{RPA}}}{\varepsilon_b - \varepsilon_m + \omega}, \tag{6}$$

$$D_{na}^{\text{RPA}} = D_{na}^{(1)} + \sum_{b,m} \frac{D_{bm}^{\text{RPA}}(g_{nmab} - g_{nmba})}{\varepsilon_b - \varepsilon_m - \omega} + \sum_{b,m} \frac{(g_{nbam} - g_{nbma})D_{mb}^{\text{RPA}}}{\varepsilon_b - \varepsilon_m + \omega}.$$
 (7)

The present third-order corrections include only the dominant Brueckner-orbital terms. Third-order RPA terms are already included in the second-order amplitude, and the much smaller [4, 5] third-order structural radiation and normalization corrections are omitted in the present calculation. The third-order Bruecknerorbital corrections δv_a to the valence orbital v_a are found by solving the inhomogeneous Dirac equation

$$(h_{\rm HF} - \varepsilon_a)\delta v_a = (E_a^{(2)} - \Sigma_{\varepsilon_a}^{(2)})v_a, \qquad (8)$$

where $\Sigma_{\epsilon}^{(2)}$ is the second-order self-energy operator and $E_{a}^{(2)}$ is the second-order correction to the ionization energy of the state *a*. Equation (8) has a nontrivial solution since

$$E_a^{(2)} = \langle v_a | \Sigma_{\varepsilon_a}^{(2)} | v_a \rangle.$$
(9)

The physically acceptable solution to Eq. (8) is orthogonal to v_a . Once the Brueckner-orbital corrections δv_a are determined by solving Eq. (8) for the states of interest, the corresponding third-order corrections to the transition amplitudes $D_{ba}^{(3)}$ are given by

$$D_{ba}^{(3)} = \langle \delta v_b || ez || v_a \rangle + \langle v_b || ez || \delta v_a \rangle.$$
⁽¹⁰⁾

The theoretical amplitude obtained from third-order MBPT, D_{ba}^{theor} , is just the sum of the three terms discussed above,

$$D_{ba}^{\text{theor}} = D_{ba}^{(1)} + D_{ba}^{(2)} + D_{ba}^{(3)}.$$
 (11)

III. RESULTS OF CALCULATION

Ionization energies of the five states for each of the ions considered are listed in Table I. In the third column of the table we give the lowest-order DHF energies $E^{(0)}$; in the fourth column, we give the second-order correlation energy $E^{(2)}$; in the fifth column we give the corrected theoretical energy $E^{\text{theor}} = E^{(0)} + E^{(2)}$; and in the sixth column, we give the experimental ionization energy. For all states considered, the second-order correlation energy is seen to be too large. In the final column of the table, we give the factor by which one must multiply $E^{(2)}$ (or the self-energy operator $\Sigma_{\epsilon_a}^{(2)}$) in order that the second-order correlation energy agrees exactly with the experimental correlation energy $E^{\text{expt}} - E^{(0)}$. This scale factor is used below to adjust the correlation corrections to transition amplitudes.

In Table II, we present amplitudes and rates for the E1 transitions from the lowest p states of each of the three ions. In the third column of the table, we give the first-order DHF amplitudes $D^{(1)}$; and in the fourth column, we give the MBPT amplitudes D^{theor} , determined from Eq. (11). In the fifth column of Table II, we present semiempirical amplitudes D^{SE} obtained using scaled Brueckner orbitals. To calculate these semiempirical amplitudes, we make two major changes to the MBPT scheme described above.

(i) We include a class of higher-order corrections to the transition amplitudes by interchanging the second- and third-order calculations. To accomplish this interchange, the Brueckner orbitals are first calculated for each valence state a, and corrected valence orbitals w_a , given by

Ion	State	E ⁽⁰⁾	E ⁽²⁾	$E^{ theor}$	$E^{ ext{expt}}$	Scale factor
Ba ⁺	$6s_{1/2}$	-0.34327	-0.029 80	-0.37308	-0.36764	0.8178
	$5d_{3/2}$	-0.31046	-0.04126	-0.35172	-0.34543	0.8476
	$5d_{5/2}$	-0.30830	-0.03918	-0.34748	-0.34178	0.8545
	$6p_{1/2}$	-0.26092	-0.01650	-0.27742	-0.27532	0.8727
	$6p_{3/2}$	-0.25458	-0.01488	-0.26946	-0.26762	0.8763
Sr+	$5s_{1/2}$	-0.38292	-0.02546	-0.40839	-0.40535	0.8810
	$4d_{3/2}$	-0.30703	-0.03576	-0.34279	-0.33903	0.8949
	$4d_{5/2}$	-0.30638	-0.03494	-0.34132	-0.33775	0.8978
	$5p_{1/2}$	-0.28483	-0.01355	-0.29838	-0.29730	0.9203
	$5p_{3/2}$	-0.28171	-0.01293	-0.29463	-0.29365	0.9234
Ca ⁺	681/2	-0.41663	-0.02173	-0.43836	-0.43626	0.9034
	$5d_{3/2}$	-0.33087	-0.04681	-0.37768	-0.37407	0.9229
	$5d_{5/2}$	-0.33076	-0.04655	-0.37731	-0.37379	0.9244
	$5p_{1/2}$	-0.31000	-0.01217	-0.32217	-0.32148	0.9433
	$5p_{3/2}$	-0.30909	-0.01202	-0.32111	-0.32046	0.9459

TABLE I. Theoretical and experimental energies (a.u.) for alkaline-earth-metal ions.

$$w_a = \sqrt{1 - \lambda} \, v_a + \delta v_a, \tag{12}$$

$$\lambda = \int d^3 r \, \delta v_a^{\dagger}(\mathbf{r}) \delta v_a(\mathbf{r}), \qquad (13)$$

are constructed. We use the corrected valence orbitals w_a rather than DHF orbitals v_a to obtain a modified first-order amplitude

$$\tilde{D}_{ba}^{(1)} = \langle w_b \| ez \| w_a \rangle, \tag{14}$$

which is approximately the sum of $D_{ba}^{(1)}$ and $D_{ba}^{(3)}$. Corrected orbitals for the valence states a and b are also used instead of DHF orbitals in Eq. (4) to obtain a modified

second-order amplitude $\tilde{D}_{ba}^{(2)}$. The amplitudes $D_{cn}^{(1)}$ and $D_{nc}^{(1)}$ on the right-hand side of Eq. (4) are replaced by RPA amplitudes as described previously. The modified second-order amplitudes determined in this way include the fourth- and higher-order terms arising from interference between the RPA and Brueckner-orbital corrections.

(ii) The second-order self-energy operator (and consequently the second-order energy) is scaled by the factor given in the sixth column of Table I to give the experimental correlation energy. This energy scaling causes the Brueckner orbitals, which are determined from Eq. (8), to be scaled by precisely the same factor. These scaled Brueckner orbitals are then used to calculate the modified amplitudes $\tilde{D}_{ba}^{(1)}$ and $\tilde{D}_{ba}^{(2)}$ described in the preceding

TABLE II. Electric dipole transition amplitudes (a.u.) and transition rates (10^9 sec^{-1}) for alkaline-earth-metal ions.

Ion	Transition	D ⁽¹⁾	$D^{ theor}$	D^{SE}	A ^{SE}
Ba ⁺	$6p_{1/2} \rightarrow 6s_{1/2}$	3.891	3.273	3.300	0.091 78
	$6p_{1/2} \rightarrow 5d_{3/2}$	3.745	2.920	3.009	0.03342
	$6p_{3/2} \rightarrow 6s_{1/2}$	5.478	4.616	4.658	0.11625
	$6p_{3/2} \rightarrow 5d_{3/2}$	1.635	1.270	1.312	0.00435
	$6p_{3/2} \rightarrow 5d_{5/2}$	5.001	3.940	4.057	0.03595
Sr ⁺	$5p_{1/2} \rightarrow 5s_{1/2}$	3.485	3.052	3.060	0.126 51
	$5p_{1/2} \rightarrow 4d_{3/2}$	3.729	2.989	3.052	0.00725
	$5p_{3/2} \rightarrow 5s_{1/2}$	4.921	4.313	4.325	0.13962
	$5p_{3/2} \rightarrow 4d_{3/2}$	1.657	1.326	1.355	0.00092
	$5p_{3/2} \rightarrow 4d_{5/2}$	5.003	4.026	4.109	0.00776
Ca+	$4p_{1/2} \rightarrow 4s_{1/2}$	3.201	2.886	2.890	0.13526
	$4p_{1/2} \rightarrow 3d_{3/2}$	3.082	2.315	2.373	0.00877
	$4p_{3/2} \rightarrow 4s_{1/2}$	4.527	4.083	4.088	0.13895
	$4p_{3/2} \rightarrow 3d_{3/2}$	1.376	1.033	1.059	0.000 93
	$4p_{3/2} \rightarrow 3d_{5/2}$	4.135	3.108	3.186	0.00824

TABLE III. Lifetimes of excited np states for alkalineearth-metal ions.

Ion	State	$ au^{ ext{SE}}$ (ns)	$ au^{ ext{expt}}$ (ns)
Ba ⁺	$6p_{1/2}$	7.99	7.92 ± 0.08^{a}
	$6p_{3/2}$	6.39	6.31 ± 0.02^{b}
Sr ⁺	$5p_{1/2}$	7.48	$7.47\pm0.07^{ t a}$
	$5p_{3/2}$	6.74	6.69 ± 0.07^{a}
Ca^+	$4p_{1/2}$	6.94	$6.96 \pm 0.18^{\circ}$
	$4p_{3/2}$	6.75	$6.87 \pm 0.18^{\circ}$

^a See Gaillard et al., Ref. [2].

^b See Andrä, Ref. [1].

^c See Gosselin, Pinnington, and Ansbacher, Ref. [3].

paragraph (i). The resulting semiempirical amplitude

$$D_{ba}^{\rm SE} = \tilde{D}_{ba}^{(1)} + \tilde{D}_{ba}^{(2)} \tag{15}$$

is given in the fifth column of Table II.

The final column of Table II gives the transition rates A^{SE} , calculated using the semiempirical amplitudes D^{SE} and the measured energies.

In Table III, the lifetimes of the excited $p_{1/2}$ and $p_{3/2}$ states, determined from the semiempirical transition rates given in Table II, are compared with experimental measurements. For the $6p_{3/2}$ state of Ba⁺, the experimental lifetime is accurate to 0.3% [1]; for the $6p_{1/2}$ state of Ba⁺ and for the $5p_{1/2}$ and $5p_{3/2}$ states of Sr⁺, the lifetimes are measured to 1% accuracy [2]; and for the $4p_{1/2}$ and $4p_{3/2}$ states of Ca⁺, they are measured to 2.5% accuracy [3]. The largest disparity between the semiempirical and measured lifetimes is found for the $4p_{3/2}$ state of Ca⁺, where the difference is 1.8%. These results are consistent with previous many-body calculations of transition rates for cesium [5].

In Table IV, we present the results of MBPT calculations of the lifetime of the metastable (n-1)d states of the alkaline-earth-metal ions. These states decay to the *ns* ground states by E2 emission. The calculation of the E2 amplitudes follows the pattern outlined in the preceding paragraphs, replacing the electric dipole operator by the electric quadruple operator. The semiempirical cal-

TABLE IV. Lifetimes of excited (n - 1)d states for alkaline-earth-metal ions.

Ion	State	$ au^{ ext{SE}}$ (s)	$ au^{\text{expt}}$ (s)	$ au^{\mathrm{other}}$ (s)
Ba ⁺	$5d_{3/2} \\ 5d_{5/2}$	83.7 37.2	17.5 ± 4^{a} 47.0 ± 16^{b} , 32.0 ± 5^{c}	45.4 ^e , 72.1 ^f 19.0 ^e , 33.2 ^f
Sr ⁺	$4d_{3/2} \\ 4d_{5/2}$	0.454 0.405	$0.395 \pm 0.038^{ m d} \ 0.345 \pm 0.033^{ m d}$	0.257 ^e 0.209 ^e
Ca+	3d _{3/2} 3d _{5/2}	$\begin{array}{c} 1.271 \\ 1.236 \end{array}$		0.797 ^e 0.774 ^e

^a See Schneider and Werth, Ref. [9].

^b See Plumelle et al., Ref. [6].

^c See Nagourney, Dandberg, and Dehmelt, Ref. [7].

^d See Gerz, Hilberath, and Werth, Ref. [8].

^e Warner, Ref. [16].

^f Garstang and Hill, Ref. [17].

culations are consistent with the measurements, except for $5d_{3/2}$ state of Ba⁺, where the theoretical lifetime is four times larger than the measured lifetime. This large difference cannot be explained by two-photon decay or by M1 decay from the $5d_{3/2}$ state; both processes are highly suppressed compared to E2 transitions.

In summary, we have obtained theoretical lifetimes of low-lying p states of alkaline-earth-metal ions that agree with precisely measured lifetimes to better than 2% accuracy. These comparisons serve as a measure of the quality of similar theoretical calculations, in cases where benchmark experiments are unavailable. For the metastable $5d_{3/2}$ state of Ba⁺, we disagree seriously with experiment, and suggest that the measured lifetime is too small by a factor of about 4.

ACKNOWLEDGMENTS

The authors owe a debt of thanks to J. E. Hansen and to H. J. Andrä for useful discussions. This research was supported in part by National Science Foundation Grant No. PHY89-07258.

- * Present address: Department of Physics, University of Notre Dame, Notre Dame, IN 46556.
- H. J. Andrä, in *Beam-Foil Spectrocopy*, edited by I. A. Sellin and D. J. Pegg (Plenum, New York, 1976), Vol. 2, p. 835.
- M. Gaillard, H. J. Plöhn, H. J. Andrä, D. Kaiser, and H. H. Schulz, in *Beam-Foil Spectrocopy* (Ref. [1]), Vol. 2, p. 853.
- [3] R. N. Gosselin, E. H. Pinnington, and W. Ansbacher, Nucl. Instrum. Methods B 31, 305 (1988).
- [4] W. R. Johnson, M. Idrees, and J. Sapirstein, Phys. Rev. A 35, 3218 (1987).

- [5] V. A. Dzuba, V. V. Flambaum, A. Ya. Kraftmakher, and O. P. Sushkov, Phys. Lett. A 142, 373 (1989).
- [6] F. Plumelle, M. Desaintfuscien, J. L. Duchene, and C. Audoin, Opt. Commun. 34, 71 (1980).
- [7] W. Nagourney, J. Dandberg, and H. Dehmelt, Phys. Rev. Lett. 56, 2797 (1986).
- [8] Ch. Gerz, Th. Hilberath, and G. Werth, Z. Phys. D 5, 97 (1987).
- [9] R. Schneider and G. Werth, Z. Phys. A 293, 103 (1979).
- [10] M. C. Noecker, B. P. Masterson, and C. E. Wieman, Phys. Rev. Lett. 61, 310 (1988).
- [11] M. A. Bouchiat et al., Phys. Lett. 117B, 358 (1982);

134B, 463 (1984).

- [12] P. Drell and E. D. Commins, Phys. Rev. A 32, 2196 (1985); C. E. Tanner and E. D. Commins, Phys. Lett. 56, 332 (1986).
- [13] V. A. Dzuba, V. V. Flambaum, and O. P. Sushkov, Phys. Lett. A 141, 147 (1989); J. Phys. B 18, 597 (1985); 20, 3297 (1987).
- [14] W. R. Johnson, S. A. Blundell, Z. W. Liu, and J. Sapirstein, Phys. Rev. A 37, 1395 (1988).
- [15] A.-M. Mårtensson-Pendrill, J. Phys. (Paris) 46, 1949 (1985).
- [16] B. Warner, Mon. Not. R. Astron. Soc. 139, 115 (1968).
- [17] R. H. Garstang and S. J. Hill, Publ. Astron. Soc. Pac. 78, 70 (1966).