Combined Configuration-Interaction-Hylleraas-Type Wave-Function Study of the Ground State of the Beryllium Atom

James S. Sims[†] and Stanley Hagstrom

Chemistry Department, Indiana University, Bloomington, Indiana 47401

(Received ll March 1971)

A method is proposed for the accurate determination of atomic wave functions and energies by the explicit introduction of interelectronic coordinates into a configuration-interaction wave function. This is accomplished by choosing the configurations in the wave function to be antisymmetrized projected products of one-electron functions with powers of interelectronic coordinates. A 107-configuration wave function for the ground state of the beryllium atom was constructed from a basis set consisting of s and b Slater-type orbitals and powers of interelectronic coordinates: r_{ij}^v ($v=0, 1, 2$). The energy obtained from this wave function $(E = -14,66654 \text{ a.u.})$ is an upper bound to the "exact" nonrelativistic energy of this state and it is believed to be within 0.0002 a.u. of the "exact" value. The advantages that the present method offers for extending accurate Hylleraas-method calculations to atomic systems with $N > 3$ are discussed.

I. INTRODUCTION

The quantum-mechanical description of atoms and molecules involves the solution of the timeindependent Schrödinger equation

 (1) $\mathcal{H}\Psi = E\Psi$

for the wave function $\Psi = \Psi(X_1, X_2, X_3, \ldots, X_N)$, where $X_i = (\mathbf{r}_i, \xi_i)$ is the combined space-spin coordinate for electron i . In the case of light atoms, where magnetic interactions can be neglected, $\mathcal X$ is the nonrelativistic many-electron Hamiltonian, which in atomic units¹ is

$$
3C = \sum_{i=1}^{N} \left(-\frac{1}{2} \Delta_i - \frac{Z}{r_i} \right) + \sum_{i < j} \frac{1}{r_{ij}} \quad , \tag{2}
$$

where Ξ is the nuclear charge and $-\frac{1}{2}\Delta_i$ is the operator corresponding to the kinetic energy of electron i.

One solution to Eq. (1) is that obtained by the Hartree-Fock (HF) method, which has proven useful for evaluating electronic properties described by many one-particle operators (e.g., static electric moment and electric or magnetic polarizabilities).² For closed-shell systems, the conventional (or restricted) HF solution to Eq. (1) is a single energy-optimized Slater determinant made up of symmetry-adapted spin orbitals (one-electron functions which are the products of spatial and spin functions). For most open-shell cases, more than one determinant is necessary to ensure that the wave function is an eigenfunction of L^2 and S^2 (again in the convention: or restricted HF method).³

Although the conventional HF approximation yields⁴ over 99% of the total electronic energies of first-rom atoms, it nevertheless fails to be accurate enough for the prediction of two- (or N -) electron properties. Thus, for example, the conventional HF method fails to predict the low-energy cross sections of electron-atom scattering. ' The accurate computation of quantities like transition probabilities, 6 electron affinities of atoms and ions, 7,8 hyperfine-structure constants, 9 and photoionization cross sections for atoms, 10 atomic ions, 11 and molecules¹² requires solutions of Eq. (1) more accurate than the HF solution.

The two standard variational¹³ methods for improving the HF solution were first introduced by Hylleraas $^{\rm 14,15}$ in a series of papers on heliumlik systems. They are (i) the configuration-interaction (CI) method, 14,16 in which the wave function Ψ is expanded in terms of antisymmetrized and pro-
iected products of spin orbitals,¹⁷ and (ii) the jected products of spin orbitals, 17 and (ii) the Hylleraas (Hy) method,^{15,18} in which the interelectronic coordinates $r_{\boldsymbol{ij}}$ are explicitly included in the terms of the wave function. There are theoretical grounds^{17,18} for thinking that both the CI and the Hy methods are general methods capable of yielding variational solutions which converge to the exact solution of Eq. (1) with any desired degree of accuracy if a sufficient number of terms are included.

On the basis of calculations on atomic systems with atomic number $N = 2^{19-2l}$ and $N = 3$, ∞ , ∞ it has generally been held^{18, 23} that for any N, CI expansions converge much more slowly than Hy-method expansions. On the other hand, (i) the computational difficulties associated with extending the Hy procedure to atomic systems with $N > 3$ have been considered so formidable that few calculations of this type have been attempted; (ii) it has been argued 24 that the success of the Hy method for $N= 2$ and $N= 3$ cannot be extrapolated to heavier

 $\overline{4}$

908

atoms; and (iii) recently A. V. Bunge, 25 C. F. Bunge²⁶ and co-workers, 27 and Schaefer and $Harris^{8, 28}$ have developed methods for refining the CI method to such a degree that there appears to be no more real theoretical or practical difficulties involved in doing CI calculations. 25

In view of the above, it seems desirable to examine whether indeed the success of the Hy method can be extrapolated to atomic systems with $N > 3$.²⁹ To accomplish this, an attempt has been made to formulate a method of explicitly introducing interelectronic coordinates r_{ij} into the wave function, which will be computationally practical for atomic calculations with $N > 3$. This has been done by a method which introduces r_{ij} coordinates into the CI method; i.e., the configurations in the wave function Ψ are chosen to be antisymmetrized projected products of s or p Slater-type orbitals $(STO's)^{30}$ with powers of the interelectronic coordinates r_{ij} . This form of the wave function has the following advantages over the expansions in interparticle coordinates^{31,32} which characterize all the Hy-method calculations for atoms with $N > 3$ to date^{22,33-35}: (i) Products of one-electron functions and at most one r_{ij} power can be used to represent computationally difficult terms of the standard Hytype wave functions such as "unlinked" products (e.g., $r_{12}r_{34}^2$) in beryllium. (ii) Use of \bar{p} STO's simplifies the problem of handling nonzero angular momentum states, which arises when one chooses a standard Hy-type wave function. (iii) Important configurations from a CI calclulation in an sp STO basis can be used directly in the wave function, with obvious computational advantages.³⁶

The calculations have been done on the beryllium atom because previous applications of the Hy method to beryllium³³⁻³⁵ have not exposed the success of the Hy method for $N > 3$, and a very refined CI calculation^{28} exists for comparison. Also, the mathematical difficulties involved in introducing r_{ij} -dependent terms into wave functions for atoms with an arbitrary number of electrons occur in the beryllium atom. Thus this calculation is a test of the general applicability (for nonrelativistic atomic calculations) of the computational and theoretical techniques developed.

II. WAVE FUNCTION

The variational solution to Eq. (1) that we are seeking is of the form

$$
\Psi_{\text{Be}}(X_1, X_2, X_3, X_4) = \sum_{K} C_K \Phi_K (X_1, X_2, X_3, X_4) ,
$$
\n(3)

where

$$
\Phi_K = O(L^2) O_{\text{as}} \left(\chi_K \gamma_{ij}^{\nu_K} \prod_{s=1}^4 \varphi_{Ks} (\vec{r}_s) \right) . \tag{4}
$$

In Eq. (4), $O(L^2)$ is an idempotent orbital angular

momentum projection operator 37 and $O_{\sf as}$ is the projection operator which guarantees the antisymmetry of the wave function

$$
O_{\text{as}} = (41)^{-1} \sum_{p} (-1)^{p} P , \qquad (5)
$$

where the summation runs over all the 4! permutations P , with p being the parity of the corresponding permutation P. Also, χ_K is one of the two linearly independent spin functions $\chi_1 = \frac{1}{2} (\alpha_1 \beta_2 - \beta_1 \alpha_2)$ $\times (\alpha_3\beta_4 - \beta_3\alpha_4)$ and $\chi_2 = \frac{1}{2} (\alpha_1\beta_4 - \beta_1\alpha_4) (\alpha_2\beta_3 - \beta_2\alpha_3)$ which span the four-electron spin space with $S=S_{\epsilon}$ = 0. In Eq. (4), $\varphi_{Ks}(\vec{r}_s)$ refers to the sth basis orbital in the Kth configuration. The orbital basis consists of STO's of the general form

$$
\varphi(\vec{r}) = \frac{(2\xi)^{n+1/2}}{[(2n)!]^{1/2}} r^{n-1} e^{-\xi r} Y_{l,m} \quad , \tag{6}
$$

where the set $\{Y_{l,m}\}$ consists of normalized spher ical harmonics in the Dirac phase convention.³⁸ ical harmonics in the Dirac phase convention.³⁸

The approach that we have followed in seeking a variational solution to Eq. (1) of the form given by Eq. (3) differs from attempts by other authors to use a correlation factor^{15, 32, 39, 40} or to combine a correlation factor with the more systematic approach of the CI method. $17,41$ The approach that we have followed may be characterized as one which allows for the utilization of a pair correlation factor r_{ij} in those configurations which make the greatest contribution in a CI expansion, in contrast with the attempts to use an average factor for the whole atom (correlation factor).

In Eq. (4) the restriction of only one r_{ij} coordinate per term, which was first proposed by James and Coolidge, ⁴² has been retained. A standard Hy-type expansion in interparticle coordinates for $N > 3$ gives rise to two different types of products of interelectronic coordinates: (i) the "linked" type, in which at least one electron subscript appears in two or more r_{ij} 's and (ii) the "unlinked" type, in which there are no repeated subscripts. Linked products (e.g., $r_{12}r_{23}$) do not appear very important energetically; evidence of this is the extremely accurate lithium calculation of Larsson²² in which the one term with more than one $r_{\boldsymbol{ij}}$ coordinat lowered the energy by only 7×10^{-6} a.u. It is felt that important unlinked products of r_{ij} coordinates can be represented within an $s\, p$ STO, r_{ij}^v basis by a single r_{ij}^v factor and excited orbitals from the STO basis. 43 For these reasons, the limitation of only one r_{ij} coordinate per term is not felt to be a severe limitation for the wave function of Eq. (4).

There is some indication of how well one might expect to do with the restricted basis: The $s\,p$ STO basis $r_{ij}^v(v=0, 1, 2)$ has recently been given by Bunge. 44 In his paper on the He atom, Bunge examines the six-term wave function of Hylleraas^{15(a)} (Ψ_{6Hy}) in terms of angular energy limits. He finds that the sp-limit energy error of Ψ_{6Hy} (when compared with his estimate of the "exact" $s\,p$ limit) is identical to the total energy error of Ψ_{6Hy} to six decimal places (a.u.). He then concludes that one should be able to compute helium wave functions with energy errors smaller than 10^{-6} a.u. by taking the direct product of a suitable $s\,p$ basis and the Ψ_{6Hy} (an s basis and r_{12}^v , $v = 0, 1, 2$). While there are still questions concerning the validity of extrapolating the successes of calculations on helium to other atomic systems, 29 this analysis clearly indicates that such a restricted basis might be an excellent one for states of first-row atoms.

III. METHOD OF CALCULATION

A. $O(L^2)$ Projection

Since $O(L^2)$ is a product of terms involving the operator L^2 , ³⁷ the commutation of $O(L^2)$ and r_{ij}^k
follows from the commutation of L^2 and r_{ij}^k . Since $O(L^2)$ and O_{as} commute, the $O(L^2)$ projection in Eq. (4) reduces to a projection on the four-orbital Hartree product

$$
\prod_{s=1}^4 \varphi_{Ks}(\vec{r}_s).
$$

The results of this projection have been tabulated by Bunge⁴⁵; the Bunge explicit formulas were used in this calculation.⁴⁶

B. Choice of Spin Function

The linearly independent spin functions chosen in this work are the valence bond (VB) spin functions, ⁴⁷ corresponding to $S = S$ _r=0:

$$
\begin{aligned} \chi_1 &= \frac{1}{2} \left(\alpha_1 \beta_2 - \beta_1 \alpha_2 \right) \left(\alpha_3 \beta_4 - \beta_3 \alpha_4 \right) \,, \\ \chi_2 &= \frac{1}{2} \left(\alpha_1 \beta_4 - \beta_1 \alpha_4 \right) \left(\alpha_2 \beta_3 - \beta_2 \alpha_3 \right) \,. \end{aligned} \tag{7}
$$

The two VB functions are related by a permutation; specifically,

$$
P_{13} = \binom{1\,234}{3\,214}
$$

has the effect of taking χ_2 into χ_1 :

$$
P_{13} \chi_2 = \frac{1}{2} \begin{pmatrix} 1234 \\ 3214 \end{pmatrix}
$$

\n
$$
\times (\alpha_1 \alpha_2 \beta_3 \beta_4 - \alpha_1 \beta_2 \alpha_3 \beta_4 - \beta_1 \alpha_2 \beta_3 \alpha_4 + \beta_1 \beta_2 \alpha_3 \alpha_4)
$$

\n
$$
= -\frac{1}{2} (\alpha_1 \beta_2 \alpha_3 \beta_4 - \alpha_1 \beta_2 \beta_3 \alpha_4 - \beta_1 \alpha_2 \alpha_3 \beta_4 + \beta_1 \alpha_2 \beta_3 \alpha_4)
$$

\n
$$
= -\chi_1 .
$$
 (8)

It should be noted that a term (configuration) formed from the second spin function χ_2 and a particular choice of r_{ij} factor and Hartree product

$$
F_K(\vec{r}_1, \vec{r}_2, \vec{r}_3, \vec{r}_4) = r_{ij}^{v_H} \prod_{s=1}^4 \varphi_{Ks}(\vec{r}_s) , \qquad (9)
$$

is equivalent to the term formed from the first spin function χ_1 and the choice of r_{ij} factor and Hartree

product $P_{13}F_{K}$. This can be seen from the relationships

$$
|\Phi_K| = |O(L^2)O_{as}\{\chi_2 F_K\}| = |O(L^2)O_{as}\{\chi_1 P_{13} F_K\}|,
$$
\n(10)

which follow from Eq. (8) and an identity for the antisymmetrizer $O_{as} = (-1)^p P^r P^{\sigma}$ (P^r permutes only the space coordinates and P^{σ} permutes only the spin coordinates). Thus it is possible to converge upon the exact wave function by employing only one spin function.

In practice one usually chooses F_K such that the energy lowering is appreciable and such that the correct localization of the electron orbitals is given (through the values of inner- and outer-shell orbital exponents). The configuration formed from F_K and χ_2 , or, equivalently, the configuration formed from $P_{13}F_K$ and χ_1 , would not be expected to give a good energy lowering. Thus, when a few configurations containing the second spin function were tried, their effect was negligible (to 10^{-6} a.u.). These observations and results parallel those of Larsson²² for the ²S ground state of lithium, although he found that the second spin function made important adjustments to the Fermi contact term. Since the emphasis here is on obtaining a good energy, the final 107-configuration wave function resulted from only terms with the spin function χ_1 .

The method used to generate VB spin functions and to evaluate matrix elements over VB spin functions was developed by E. A. Scarzafava. 48 The explicit formula for the Hamiltonian matrix element after integration over spin coordinates is $(\chi_K = \chi_L)$ $= \chi_1$), 46

$$
H_{KL} = \langle \Phi_K | \mathcal{H} | \Phi_L \rangle
$$

= (1/41)(O(L²)A^{\alpha}A^{\beta}S₁^TF_K | \mathcal{H} | S₁^TF_L \rangle , (11)

where $A^{\alpha}A^{\beta} = (1 - P_{13}^r) (1 - P_{24}^r)$ is a product of spacepair antisymmetrizers (un-normalized), and S_1^r $=\frac{1}{2}(1+P_{12}^r)(1+P_{34}^r)$ is a product of commuting spacepair symmetrizers (normalized).

C. Integral Treatment

To complete the evaluation of the Hamiltonian matrix element H_{KL} , a scheme is needed to handle the two-, three-, and four-electron integrals of the form

$$
\int f_1(\mathbf{\tilde{r}}_1) f_2(\mathbf{\tilde{r}}_2) f_3(\mathbf{\tilde{r}}_3) f_4(\mathbf{\tilde{r}}_4) g_1(r_{12})
$$

× $g_2 (r_{23}) \cdot \cdot \cdot g_k (r_{34}) dv$. (12)

A good scheme for the evaluation of three-electron integrals of this form containing nonspherically symmetrical STQ's has been formulated by Ohrn and Nordling.⁴⁹ Their scheme has been extended in this work to the treatment of four-electron integrals. Also, the Ohrn and Nordling scheme for the evaluation of their V and W auxiliary integrals involves

TABLE I. 11-configuration wave function for the ground state of atomic beryllium with the energy for various basis-set orbital exponents.

^aFor a definition of the notation used, see Sec. IV.

infinite summations; in this work, McKoy's⁵⁰ closed-form formulas for certain of these auxiliary integrals were used. In some instances McKoy's formulas were found to lead to significant differencing errors; a stable computational procedure was found for this region by using Euler's method⁵¹ to transform McKoy's formulas.⁵² We intend to describe our integral treatment in more detail in another paper.⁵²

IV. CHOICE OF PARAMETERS AND CONFIGURATIONS

The wave function employed here for the ¹S ground state of beryllium is of the form $Eq. (3)$

$$
\Psi_I = \sum_K C_K^I \, \Phi_K \,, \tag{13}
$$

where Φ_K is a properly antisymmetrized and projected product of χ_1 , STO's defined by Eq. (6), and an $r_{ij}^{v_K}$ factor. By minimizing the energy with respect to the coefficients C_K^I , one is led to the usual secular equation with the Ith lowest root being an upper bound to the exact energy of the respective state. The secular equation is solved by Jacobi's method⁵³ after the overlap matrix has been triangulated and the Hamiltonian matrix transformed by the method of Michels, Van Dine, and Elliott.⁵⁴

In order to determine the orbital exponents of the STO basis, three 11-configuration calculations were performed. The wave function is given in Table I. where under configuration is listed only the unique

$$
r_{ij}^{v_K} \prod_{s=1}^4 \varphi_{Ks} (\vec{\bf r}_s)
$$

part of Φ_K ($\chi_K = \chi_1$) as defined by Eq. (4). The STO
basis is, in the Weiss notation, ⁵⁵ the basis set 1s, 2s, ..., ns; $1s''$, ..., ns''. The first basis-set orbital exponents $\xi_s = 3.6847$, $\xi_{s''} = 0.9562$, 56,57 were chosen to repeat the 11-configuration calculation of Karl.³⁵ Next two other sets of orbital exponents were tried: the set $\xi_s = 3.6848$, $\xi_s = 0.9560^{58}$ and the set $\xi_s = 3.7$, $\xi_{s''} = 0.9$. Neither case did as well as

the first set (see Table I). Therefore, $\xi_s = 3.6847$ and $\xi_{s''}=0.9562$ were chosen for the remaining calculations.

The effect of not having done an exhaustive minimization at this point is not felt to be significant because other investigators have found that the orbital exponents have very little dependence on the presence or absence of an r_i , coordinate. ^{22, 23} In addition, configurations containing inner-shell ns' orbitals $(\xi_s, = 7, 8)$ and some of the "open-shell" configurations of Gentner and Burke³⁴ (ζ_{s} = 2.3 in outershell orbitals) were added to larger calculations and found to give insignificant energy improvements.

Next, the various 12-configuration calculations displayed in Table II were performed in order to (i) determine ξ_{b} and (ii) test how well the r_{34}^2 terms can represent the well-known $2s-2p$ near degeneracy in beryllium. Since the Weiss ξ_{b} ["] = 1.1 proved superior to $\xi_{\mu\nu} = 0.9562$, in the remaining calculations ξ_{b} = 1.1 was used and the Weiss value ξ_{b} = 6.2 was used for the K shell. Also it should be noted that neither $1s^22s^{\prime\prime}\,^2r_{34}^2$ nor $1s^21s^{\prime\prime}\,^2r_{34}^2$ does nearly as well as $1s^22p''^2$.

Next, configurations related to the best of the 12configuration wave function and also some new configurations were added to the 12-configuration wave function; those which gave only a small energy improvement were dropped. The process was continued until almost all conceivable combinations were tried, resulting in a final 107-configuration wave function. The terms in the best 107-configuration wave function were grouped by type and by the magnitude of their coefficients; configurations were then successively removed one by one until the $1s^22s''^2$ configuration was reached.

V. RESULTS

The final 107-configuration wave function is displayed in Table III, along with the energy $E(n)$ for the n -configuration wave function (configurations) $1-n$) and the estimate of the energy contribution of each configuration obtained from $E(n) - E(n-1)$.

The 1-configuration wave function is an approximation to the HF function and the energy of the 1configuration wave function, $E = -14.55674$ a.u. is close to the HF energy of -14.57302 a.u.⁵⁸ The 22-configuration energy of -14.588316 a.u. is not

TABLE II. 12-configuration wave functions and energies for the ground state of atomic beryllium.

Energy (a, u)	5'	Wave function
-14.6465	\cdots	
-14.6480	\cdots	$\frac{\Phi_{11}+1s^2\,2s^{\prime\prime}{}^2\,r_{34}^2}{\Phi_{11}+1s^2\,1s^{\prime\prime}{}^2\,r_{34}^2}$
-14.6514	0.9562	$\Phi_{11} + 1s^2 2p''$
-14.6529	1.1	$\Phi_{11} + 1s^2 2p''$

^a From Table I; $\xi_s = 3.6847$, $\xi_{s'} = 0.9562$.

 $\hat{\boldsymbol{\lambda}}$

 $- \, 14$, $660 \, 003 \, 54$

 $- \, 0.000 \, 011 \, 27$

TABLE III. 107-configuration wave function for the ¹S ground state of atomic beryllium, with the energy contribution of each configuration.

 \equiv

		-111.1 www.cnutu				
Configuration		Energy $E(n)$ of n -configuration	Energy contribution $E(n)-E(n-1),$	Coefficient 107 -term (unnormalized)		
No. n	Configuration ^a	function (a.u.)	(a.u.)	wave function		
57	$2s^2$ 1s'' r_{12}^2	$-14,66003305$	-0.00002951	$-6.703544743\times10^{-3}$		
58	$\begin{array}{c}\n2s1s \\ 2s'')\n\end{array}\n\quad\n\begin{array}{c}\n2s''\n\end{array}\n\quad\n\begin{array}{c}\n\ddot{r}_{34} \\ 3s1s \\ 2s'''\n\end{array}\n\quad\n\begin{array}{c}\n\ddot{r}_{34}^2\n\end{array}$	$-14,66005460$	-0.00002155	$-1.300546033\times10^{-2}$		
59		$-14,66007522$	-0.00002062	8.192479333 \times 10 ⁻³		
60	$4s1s$ 2s'' r_{34}^2	-14.66011178	-0.00003656	$-5.401901555\times10^{-3}$		
61	$1s^2$ 3s'' r_{34}^2	-14.66011195	-0.00000017	$-7.906469502\times10^{-3}$		
62	$1s23s''2r34$	-14.66011250	-0.00000055	$-1.398060670\times10^{-2}$		
63	$1s^2$ 2s" r_{13}	$-14,66120865$	-0.00109615	-2.429212061		
64	$1s^2$ 1s $^{\prime\prime}$ 2s $^{\prime\prime}$ r_{13}	-14.66243593	-0.00122728	1.010 637 493		
65	$1s^2$ 3s" 2s" r_{13}	-14.66243602	-0.00000009	2.333 595 641 \times 10 ¹		
66	$1s^2$ 4s''2s" r_{13}	-14.66245219	-0.00001617	-8.317840025×10^{1}		
67	$1s^2$ 5s $''$ 2s $''$ r_{13}	-14.66246382	-0.00001163	1.226970702×10^{2}		
68	$1s^2$ 6s" $2s$ " r_{13}	-14.66246459	-0.00000077	-6.688201230×10^{1}		
69	$1s^2$ 2s'' r_{13}^2	$-14,66249434$	-0.00002975	-1.010450088×10^{1}		
70	$1s^2$ 3s $^{\prime\prime}$ 2s $^{\prime\prime}$ r_{13}^2	-14.66250885	-0.00001451	3.919 367 631 \times 10 ¹		
71	$1s^2$ 4s" $2s" r_{13}^2$	$-14,66252967$	-0.00002082	-5.921306844×10^{1}		
72	$1s^2$ 5s $''2s''r_{13}^2$	-14.66259020	-0.00006053	3.269302918×10^{1}		
73	$1s^2$ 3s''2s'' r_{24}^2	-14.66264872	-0.00005852	7.076308541×10^{-1}		
74	$1s^2$ 4s'' 2s'' r_{24}^2	-14.66269166	-0.00004294	$-7.641050527\times10^{-2}$		
75	$1s^2$ 5s" $2s" r_{24}^2$	-14.66270798	-0.00001632	$-7.032753091\times10^{-2}$		
76	$2s1s$ $2s$ $\prime\prime$ ² r_{13}^2	-14.66276858	-0.00006060	$-2.701382819\times10^{-1}$		
77	$3s1s$ $2s$ \prime r_{13}^2	-14.66280060	-0.00003202	3.183369817×10^{-2}		
78	$1s^2$ 3s $\prime\prime$ \prime \prime \prime \prime \prime	-14.66281688	-0.00001628	$-2.636753100\times10^{-1}$		
79	$2s1s$ $2s$ $\frac{12}{r_{13}}$	-14.66306710	-0.00025022	8.671072552×10^{-1}		
80	$\frac{3s}{1s^2}\frac{2s''^2}{1s''^2}\frac{r_{13}}{r_{13}^2}$	-14.66306728	-0.00000018	$-1.042006362\times10^{-1}$		
81		$-14,66349843$	-0.00043115	2.772584514×10^{-2}		
82	$1s^2$ 3s'' r_{13}	-14.66349884	-0.00000041	$-1.188093091\times10^{-1}$ 1.206853851×10^{-1}		
83	$1s^2 4s''^2 r_{13}$	-14.66350213	-0.00000329	7.058385812×10^{-3}		
84	$2s^2 2s^{\prime\prime}{}^2 r_{13}$ $1s^2$ 1s" r_{13}	$-14,66350296$	-0.00000083 -0.00001083	$-2.828054601\times10^{-1}$		
85		-14.66351379	-0.00000237	$-8.479834642\times10^{-1}$		
86	$1s^2$ 3s''2s'' r_{24} $1s^2$ 4s''2s'' r_{24}	-14 , 663 516 16	-0.00002154	5.482149284 \times 10 ⁻²		
87 88	$1s^2$ 5s" 2s" r_{24}	-14.66353770 $-14,66354585$	-0.00000815	8.709 265 657 \times 10 ⁻²		
89	$2s1s$ $2s$ n^2 r_{24}	-14.66356673	-0.00002088	$-1.054300751\times10^{-1}$		
90	$1s^2$ 6s" $2s" r_{24}$	$-14,66356759$	-0.00000086	1.684944651×10^{-2}		
91	$2s1s$ 2s'' ² r_{24}	-14.66357262	-0.00000503	4.183593501 \times 10 ⁻²		
92	$2s1s$ 3s $"2s"$ r_{24}^2	$-14,66358303$	-0.00001041	1.314304853×10^{-2}		
93	$1s^2 3p''^2$	-14.66358311	-0.00000008	5.218659990×10^{-1}		
94	$1s^2 2p''^2$	$\!-\!14.663\,630\,11$	-0.00004700	$-1.409783443\times10^{-1}$		
95	$1s^2 3p'' 2p''$	-14.66367888	-0.00004877	$-1.954639743\times10^{-1}$		
96	$1s^2 2p''^2 r_{12}^2$	-14.66491282	-0.00123394	1.815973512×10^{-2}		
97	$1s^2 2p''^2 r_{12}$	-14.666 097 10	-0.00118428	$-6.389885865\times10^{-2}$		
98	2s1s 2p''	-14.66630086	-0.00020376	6.631 045 101×10^{-2}		
99	3s1s 2p''	-14 , 666 463 91	-0.00016305	$-2.992109286\times10^{-2}$		
100	$2p^2$ $2s$ \prime 2	-14.66647428	-0.00001037	7.725719831×10^{-3}		
101	$2p^2$ 2s'' r_{12}	$-14,66651325$	-0.00003897	$-7.914756166\times10^{-3}$		
102	1s $2p2p$ " $2s$ "	$-14.666\,513\,84$	-0.00000059	$-1.033981466\times10^{-2}$		
103	1s3b3b''2s''	$\,$ $ 14$ $\,$ 666 514 72	-0.00000088	$-7.282806240\times10^{-3}$		
104	$1s 2p^2 2s''$	$\!-14$, 666 539 25	-0.00002453	2.172123776×10^{-3}		
105	$1s 2p''^2 2s''$	– 14.666 544 84	-0.00000559	$-1.771692106\times10^{-3}$		
106	3s1s3s''2s''	-14 , $666\,546\,47$	-0.00000162	9.255 210 743×10^{-2}		
107	$3s1s$ $4s$ $^{\prime\prime}2s^{\prime\prime}$	-14.66654658	-0.00000012	$-3.268371201\times10^{-2}$		

TABLE III. (Continued)

^aFor a definition of the notation used, see Sec. IV.

TABLE IV. Comparison of various ab initio^a variational calculations on the ground state of the beryllium atom.

 a By ab initio is meant [L. C. Allen and A. M. Karo, Rev. Mod. Phys. $32, 275$ (1960)] use of the exact nonrelativistic Hamiltonian, consideration of all electrons simultaneously, and evaluation of all integrals as accurately as necessary.

 b N. Sabelli and J. Hinze, J. Chem. Phys. 50 , 684 $(1969).$

 $\text{``Thought to be the most reliable estimate (see Sec. V).}$

as good as the Weiss²⁰ value of -14.59110 a.u. for his more extensive s-basis set. This is because (i) the Weiss configurations containing s' and s'' orbitals were added to more extensive wave functions containing powers of r_{12} and r_{34} . In these calculations, the effect of configurations containing s' and s'" orbitals was observed to be negligible (to 10^{-6} a.u.). (ii) The 107-configuration wave function can be simplified without appreciably effecting the energy; thus, for example, a 94-configuration truncation of the 107-configuration wave function was found with an energy of $E = -14.666505$ a.u.

The energy results obtained with a 107-term wave function are tabulated in Table IV and compared with a number of ab initio variational calculaparcu while a number of ab *intro* variational carea not been able to match the success of the Weiss²⁰ 55-configuration calculation, whereas in the present work we have found a 59-term Hy-type wave function with an energy improvement of 0. 002 35 a. u. over the Weiss wave function. The 92-term wave function contained in the final 107-term wave function is of the type employed in all the previous Hy-method calculations for atoms with $N \geq 3$. ^{22,33–35} Comparing it with the 59-term wave function indicates that the convergence is slow towards the end of the conventional Hy-type expansion. In contrast, inclusion of the p STO's, which Bunge has shown give an excellent basis (when combined with a conventional Hy basis) for the K shell, and which for small calculations (Table III) are more important than r_{34}^2 terms in the L shell, apparently speeds the convergence.

Inspection of Table III shows that the $2s-2p$ degeneracy has been represented to a great degree by terms containing r_{34}^2 . Apparently the importance of the p STO basis for this calculation is primarily due to the configurations 1s² 2p'^{'2} r_{12} and 1s² 2p'^{'2} r_{12}^2 , which are the four-excitations arising from a correlated electron-pair wave function Ω

$$
\Omega = O_{as} \left\{ K(1, 2) \ L(3, 4) \right\},\tag{14}
$$

but without the strong orthogonality condition. Their relative importance can be explained by the fact that the $2p''$ configuration alone contributes ~90% of the L-shell correlation energy^{59,60} and configurations containing r_{12} and r_{12}^2 are very important in the K shell.

The final 107-term wave function obtained after examining approximately 145 terms formed from an s p STO, r_{ij}^v (v=0, 1, 2) basis and keeping the most important, yields an energy of —14. 66654 a. u. This compares favorably with the Bunge²⁶ 180-term CI wave function obtained after examining approximately 1000 terms (Table IV).

The "exact" energy for the nonrelativisitic Hamiltonian in Eq. (2) has been estimated by several authors^{59,61,62} from experimental data by making corrections for relativistic effects. Only the lates estimate of this type, ⁶¹ based upon the nonrelativi: tic corrections of Hartmann and Clementi⁶³ computed with a HF wave function, has been included in Table IV.

In contrast, Bunge²⁶ has estimated the exact energy by studying patterns of convergence based upon his 180-term Be wave function and his work on Be^{**}.⁴⁴ Since this is an extrapolation from a highly accurate direct solution of Eq. (1), with an error which is estimated to be not greater than 0. 0003 a.u., $E = -14.66639$ a.u. is taken as an upper bound to the exact nonrelativistic energy. Adding to this the error estimate G. 0003 a. u. yields a lower bound of -14.6667 a.u., and our final wave function yields an energy within 0. 0002 a.u. $(\approx 0.005 \text{ eV})$ of the Bunge exact energy.

In building up the wave function, terms which contributed only slightly were dropped and the energy loss on dropping them was computed. In this way

38 terms were dropped; their inclusion in a 145 term wave function could lower the energy no more than 0.000072 a.u. (the sum of the energy losses on dropping these terms at various stages of the computation). This tends to indicate that our wave function is converging to an exact energy within the limits of the Bunge estimate.

VI. CONCLUSIONS

On the basis of the results presented here, it seems reasonable to conclude that the introduction of r_{i} , coordinates into an atomic wave function with $N > 3$ is computationally practical. The basis consisting of an s p STO basis and r_{ij}^v (v = 0, 1, 2) is indeed excellent for beryllium, yielding an upper bound to the ground state of greater accuracy than can be practically obtained by CI expansions at present. (Compare the s $p \, dfg$ energy limit of Bunge, 26 $E = -14.66598 a.u.$

The inclusion of p STO's gives more flexibility to the wave function and leads to simpler expansions for the KL shells. But the real significance of in-

*Research supported in part by grants from the U.S. Air Force Office of Scientific Research and the National Science Foundation.

~National Institute of Health Predoctoral Fellow, 1966- 1968. Present address: Mail Stop 245-4 NASA/Ames Research Center, Moffett Field, Calif. 94035.

¹The atomic unit of energy is chosen as $\mu e^{4}/\hbar^{2}=1$ a.u. (of energy), where μ is the reduced mass of the electron: $\mu = m_e m_N/(m_e + m_N)$.

²For fairly recent summaries (with references) of HF theory see, for example, R. K. Nesbet, in Advances in Chemical Physics, edited by I. Prigogine (Interscience, New York, 1965), Vol. 9; F. L. Pilar, Elementary Quan tum Chemistry, (McGraw-Hill, New York, 1968); or S. Fraga and G. Malli, Many Electron Systems: Properties and Interactions (W. B. Saunders, Philadelphia, 1968).

 ${}^{3}C$. C. J. Roothaan and P. Bagus, in Methods in Computational Physics, edited by B. Alder (Academic, New York, 1963), Vol. 2.

 $4J.$ C. Slater, Quantum Theory of Atomic Structure (McGraw-Hill, New York, 1960), Vol. 2.

 5 K. Smith, R. J. W. Henry, and P. G. Burke, Phys. Bev. 157, 51 (1967).

 $6A.$ W. Weiss, Phys. Rev. 162, 71 (1967); J. Chem. Phys. 47, 3573 (1967).

 7 A. W. Weiss, Phys. Rev. 166, 70 (1968).

 8 H. F. Schaefer and F. E. Harris, Phys. Rev. 170, 108 (1968); Phys. Bev. Letters 21, 1561 (1968).

⁹H. F. Schaefer, R. A. Klemm, and F. E. Harris, Phys. Rev. 176, 49 (1968); 181, 137 (1969).

¹⁰R. J. W. Henry, J. Chem. Phys. 48, 3635 (1968); Planetary Space Sci. 15 , 1747 (1967); R. J. W. Henry and L. Lipsky, Phys. Bev. 153, 51 (1967).

 11 M. B. Hidalgo, Astrophys. J. $153, 981$ (1968).

¹²A general approach to the treatment of the hydrogen molecule including configuration interaction is being

cluding the p STO's appears to be in dealing with the four-excitations arising from a correlated electronpair wave function $\{K(1, 2), L(3, 4)\}$. Thus it has been found that the terms $1s^{2}2p''^{2}r_{12}$ and $1s^{2}2p''^{2}r_{12}^{2}$ are important four-excitations. In a standard Hy-method calculation the effect of these terms would have to be represented by configurations containing products of r_{ij} coordinates (such as $r_{12}r_{34}^2$ and $r_{12}^2r_{34}^2$), which have proven computationally difficult to deal with.

In conclusion, the present method appears to offer distinct advantages for extending accurate Hymethod calculations to atomic systems with $N > 3$.

ACKNOWLEDGMENTS

One of us (J.S.S.) acknowledges the National Aeronautics and Space Administration for providing funds for the use of a CDC-6600 computer at the Lawrence Radiation Laboratory, Berkeley, California, which was indispensable in the completion of this research. The authors wish to thank Dr. Carlos Bunge for the communication of his results prior to publication.

developed in this laboratory by J. R. Rumble, Theoretical Chemistry Group Technical Note No. 002, Indiana University, 1969 (unpublished).

 13 For a comparative study of atomic variational wave functions with references, see J. N. Silverman and G. H. Brigman, Bev. Mod. Phys. 39, 228 (1967).

 $^{14}E.$ A. Hylleraas, Z. Physik 48, 469 (1928).

 $15(a)$ E. A. Hylleraas, Z. Physik 54 , 347 (1929); (b)

 $\frac{60}{16}$, 624 (1930); (c) $\frac{65}{16}$, 209 (1930).
 $\frac{16}{16}$ S. F. Boys, Proc. Roy. Soc. (London) <u>A201</u>, 125 (1950).

 $^{17}P.$ O. Lowdin, Phys. Rev. $97, 1474$ (1955); Rev.

Mod. Phys. 32, 328 (1960). See also P. O. Lowdin, in

Advances in Chemical Physics, edited by I. Prigogine (Interscience, New York, 1959), Vol. 2.

 18 L. Szasz, Phys. Rev. 126 , 169 (1962).

i9D. H. Tycho, L. H. Thomas, and K. M. King, Phys. Bev. 109, 369 (1958).

 20 A. W. Weiss, Phys. Rev. 122, 1826 (1961).

²¹C. L. Pekeris, Phys. Rev. 115 , 1216 (1959).

²²S. Larsson, Phys. Rev. $169, 49$ (1968).

23J. N. Silverman and G. H. Brigman, Rev. Mod. Phys. 39, 228 (1967). This reference contains an extensive bibliography of CI and Hy studies.

 24 T. L. Gilbert, Rev. Mod. Phys. $35, 491$ (1963).

 $25A$. V. Bunge, Ph. D. dissertation (University of Florida, 1968) (unpublished) .

 26 C. F. Bunge, Phys. Rev. 168, 92 (1968).

²⁷C. F. Bunge and E. M. A. Peixoto, Phys. Rev. A $\frac{1}{1}$, 1277 (1970); C. S. Causey and C. F. Bunge (unpublished); D. Munch and C. F. Bunge (unpublished).

 28 H. F. Schaefer and F. E. Harris, Phys. Rev. 167, 67 {1968).

 29 The alternative possibility is that expansions in powers of r_i , r_j , and r_i are not particularly convenient and are possibly slowly convergent for pairs of electrons (I-shell and intexshell pairs) less tightly bound than the K-shell

pair of electrons, or that cumbersome products of powers of r_{ij} coordinates will be required.

 30 The STO notation that we use [see Eq. (6) in the text] and some of the STO's themselves are from A. W. Weiss, Ref. 20.

 31 Interparticle coordinates are the magnitudes of the distances between each pair of particles in an atomic or molecular system. For an atom, these are the interelectronic distances r_{ij} and the electron-nuclear distances r_i . See, e.g. , Ref. 32.

 $32A$. A. Frost, Theoret. Chim. Acta 1, 36 (1962).

33L. Szasz and J. Byrne, Phys. Rev. 158, ³⁴ (1967).

 $34R$. F. Gentner and E. A. Burke, Phys. Rev. 176, 63 (1968) .

 35 J. Karl, Ph.D. dissertation (University of Wisconsin, 1966) (unpublished) .

 36 L. Szasz (Ref. 18) has presented a general theory of correlated wave functions from which our proposed wave function can be obtained by expanding some of his N-electron correlation functions by the CI method, but in applications [Phys. Letters 3 , 263 (1963) and Ref. 33] he uses Hy-type wave functions employing only powers of r_i , r_j , and r_{ij} .

 $37P.$ O. Lowdin, Rev. Mod. Phys. 36 , 966 (1964).

 ^{38}E . U. Condon and G. H. Shortley, The Theory of Atomic Spectra (Cambridge U. P. , Cambridge, England, 1963).

 $39C.$ C. J. Roothaan and A. W. Weiss, Rev. Mod. Phys. 32, 194 (1960).

 40 S. F. Boys and N. C. Handy, Proc. Roy. Soc. (London) A309, 209 (1969); A310, 43 (1969); A310, 63 (1969). ^{41}P . O. Lowdin and L. Redei, Phys. Rev. 114, 752 (1959).

 42 H. M. James and A. S. Coolidge, Phys. Rev. 49 , 688 (1936).

 43 O. Sinanoglu [J. Chem. Phys. 36, 705 (1962)] has pointed out that four-excitations in Be are mainly "unlinked clusters, "which can be accounted for in the wave function employed here by terms like $1s^22p^{\prime\prime\,2}\,r_{12}$ (notation as in Sec. IV). The formalism employed here differs from Sinanoğlu's by being a nonorthogonal formalism; nevertheless, it is felt that Sinanoglu's arguments for the relative importance of "unlinked" products and the relative unimportance of "linked" products of pairs are valid in considering the importance of configurations containing products of interelectronic coordinates.

 44 C. F. Bunge, Theoret. Chim. Acta 16 , 126 (1970).

⁴⁵C. F. Bunge, Ph.D. dissertation (University of Florida, 1966 , (unpublished).

 46 Further details have been described in J. S. Sims.

Ph. D. dissertation (Indiana University, 1969) (unpublished). 47 H. Eyring, J. Walter, and G. Kimball, Quantum

Chemistry (Wiley, New York, 1944), Chap. 13; G. Rumer, Gott. Nachr. Math. Phys. Kl. 377 (1932).

 ^{48}E . A. Scarzafava, Ph.D. dissertation (Indiana University, 1969) (unpublished).

 $49Y.$ Ohrn and J. Nordling, J. Chem. Phys. 39, 1864 (1963).

 ^{50}V . McKoy, J. Chem. Phys. 42 , 2959 (1965).

 51 T. J. Bromwich, An Introduction to the Theory of Infinite Series (Macmillan, London, 1926).

 52 J. S. Sims and S. A. Hagstrom, J. Chem. Phys. (to be published).

 $53A$ FAP routine written by F. P. Prosser and references to the method fincluding the original reference, C. G. J. Jacobi, J. Reine Angew. Math. 30, ⁵¹ (1846)] are available as Program 3, Quantum Chemistry Program Exchange, Indiana University. The FAP routine has been rewritten in . FORTRAN by E. A. Scarzafava, Quantum Chemistry Group, Indiana University.

 54 H. H. Michels, C. P. Van Dine, and P. Elliott, Program 97, Quantum Chemistry Program Exchange, Indiana University (unpublished) .

 55 Priming on functions of the same symmetry refers to different values of ξ ; e.g., 1s, 2s, 1s', 1s' means a set of STO's with orbital exponents of $\zeta_{1s} = \zeta_{2s} \neq \zeta_{1s} \neq \zeta_{1s}$. The second s function is denoted s ' since it corresponds closely to the Weiss s' orbital. See Ref. 20.

56E. Clementi and D. L. Raimondi, J. Chem. Phys.

38, 2686 (1963). C. C. J. Boothaan, Laboratory for Molecular Structure and Specta Technical Report No. 1955, University of Chicago, 1955 (unpublished).

 58 C. C. J. Roothaan, L. M. Sachs, and A. W. Weiss, Rev. Mod. Phys. 32, ¹⁸⁶ (1960); E. Clementi, IBM J. Res. Develop. 9, 2 (1965) and suppl. (1965).

 59 R. E. Watson, Phys. Rev. 119, 170 (1960).

 60 J. Linderberg and H. Shull, J. Mol. Spectry. 5, 1 (1960).

 61 K. J. Miller and K. Ruedenberg, J. Chem. Phys. 48, 3414 (1968); 48, 3447 (1968); 48, 3450 (1968).

⁶²R. K. Nesbet, Phys. Rev. 155, 51 (1967).

3H. Hartmann and E. Clementi, Phys. Rev. 133, A1295 (1964).