PHYSICAL REVIEW A GENERAL PHYSICS

THIRD SERIES, VOLUME 37, NUMBER 7

APRIL 1, 1988

Variational calculations for ⁴He I: Improved energies for singlet and triplet $n D$ and $n F$ levels ($n = 3-8$)

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Hylleraas method variational calculations with up to 458 expansion terms are reported for the $n^{1}D$, $n^{3}D$ terms and for the $n^{1}F$, $n^{3}F$ terms ($n = 4-8$) of neutral helium. Convergence arguments are presented to obtain new estimates of the exact nonrelativistic energies of these terms. The reliaare presented to obtain new estimates of the exact nonrelativistic energies of these terms. The relia-
bility of the estimates ranges from 3×10^{-10} a.u. in the worst case to 10^{-11} a.u. in the best case. The nonrelativistic singlet and triplet energies are combined with previously calculated values for relativistic and other small energy contributions to evaluate the total ionization energies for the $n D$ and n F levels. Comparisons of the n ¹D and n³D energies with other calculations and available data yield improved values for relativistic contributions. The calculated relativistic $n F$ ionization energies agree with values based on core-polarization theoretical $n G$ and $n H$ energies within deviations of 3 MHz $(5\times10^{-11}$ a.u.) for $n=5$ to 1 MHz $(1.5\times10^{-11}$ a.u.) for $n=7,8$. The calculated n F finestructure separations generally agree with experiment for $n = 5-8$ within uncertainties of the order of 1 to 0.1 MHz. Predicted values given for the 4 F fine-structure separations and the 4^1F ionization energy are probably accurate within a few MHz.

I. INTRODUCTION

Recent improvements in experimental wavelength $\frac{1}{4}$ measurements of He optical transitions¹⁻⁴ have led to improved values for a number of lsnl energy levels $(n = 2-6)$. The accuracies of these experimental results have sparked renewed interest in the theoretical determination of these energy values. Comparisons of experimental and theoretical results provide important tests of ihe theory for this basic two-electron atom, and certain combinations of experimental and theoretical results have yielded increasingly more accurate ionization energies for 1snl levels $(n \geq 2).$ ⁵⁻⁹

To obtain energy values theoretically at least as accurate as the experimental results requires accurate systematic variational calculations employing hundreds of terms produced from a correlated basis set, such as the pioneering work of Pekeris and co-workers.¹⁰⁻¹² Pekeris and co-workers treated only low-lying S and P levels of He, presumably because in their method a new procedure is required for each diferent angular momentum. In a previous paper,¹³ we have performed calculations on low-lying ${}^{1}D$ and ${}^{1}F$ levels in the spirit of the definitive calculations by Pekeris and co-workers on low-lying ${}^{1}P$ levels. However our previous, approach suffered the same problem as the previous S and P calculations in that the numerical convergence of the calculations gets progressively poorer for higher and higher levels

In recent papers, Kono and Hattori^{14, 15, 6,7} have show: that a careful choice of basis functions greatly improves the convergence and permits very accurate energy level calculations to be performed for essentially any level. In this work we have used Kono and Hattori's method of basis set selection to obtain new estimates of the nonrelativistic energies of the $n^{1}D$, $n^{3}D$ terms ($n = 3-8$) and of the $n^{1}F$, $n^{3}F$ terms (n = 4-8). The D results extend Kono and Hattori's results to greater precision. The ${}^{1}F$ results extend the results of Sims, Parmer, and Reese to greater precision, and the ${}^{3}F$ results are on a level of accuracy never before achieved. We have combined these new results with available calculated values for the relativistic and other small contributions to obtain new predictions for $n D$ and $n F$ ionization energies. These ionization energies are tested for consistency with $n G$ and n H ionization energies calculated from core-polarization theory and are used to test the accuracy of the Breit-Pauli fine-structure theory as applied to $1snf$ configurations.

II. WAVE FUNCTIONS

The method used in the calculations is the combined configuration interaction (CI)-Hylleraas (Hy) method, originally proposed by Sims and Hagstrom.¹⁶ This

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method has been shown to be capable of giving very accurate values for the energies, oscillator strengths, and polarizabilities of two-, three-, and four-electron systems (Ref. 13 and references therein), and its utility for Xelectron systems up to and including $N = 10$ has been demonstrated by Clary and Handy.¹⁷ The CI-Hy method wave functions can be written in a form which is conventional for the two-electron Hy method calcula $tions^{10-12,18}$

$$
\Psi(\mathbf{r}_1, \mathbf{r}_2) = \sum_K C_K \Phi_K(\mathbf{r}_1, \mathbf{r}_2) , \qquad (1)
$$

where the terms Φ_K are specifically of the form

$$
\Phi_K(\mathbf{r}_1, \mathbf{r}_2) = (1 \pm P_{12}) r_1^i r_2^j r_{12}^k
$$

$$
\times Y_0^0(1) Y_i^j(2) \exp(-\xi r_1 - \eta r_2) . \tag{2}
$$

 P_{12} is the operator which permutes electrons 1 and 2, Y_i is the $(m = l)$ spherical harmonic chosen to give $S(l = 0)$, $P(l=1)$, $D(l=2)$, or $F(l=3)$ symmetry, and the plus sign is for singlet levels, the minus sign for triplet levels. In addition to terms of the form of Eq. (2), for D and F levels terms of the form

$$
\Phi_K(\mathbf{r}_1, \mathbf{r}_2) = (1 \pm P_{12}) r_1^i r_2^j r_{12}^k Y_1^1(1) Y_{l-1}^{l-1}(2)
$$

× $\exp(-\xi r_1 - \eta r_2)$ (3)

are required to form a complete set of D and F angular momentum functions.^{14,19}

Following Kono and Hattori we employ two sets of basis functions, differing in the nonlinear parameters ξ and η , to accelerate convergence. The first set is chosen to have the correct asymptotic form as $r_1, r_2 \rightarrow \infty$. Hence, for the first set (which we refer to as the ξ set), ξ and η are fixed at the values Z and $(-2E - Z^2)^{1/2}$, respectively, where E is an estimate of the energy eigenvalue of the level in question and Z is the nuclear charge. Terms are chosen for this set by taking all terms for which $i = 0, 1$; $k = 0, 1$; and $L \le j \le N+4$, where N and L are quantum numbers for the state in question. This choice of terms gives a rough approximation to the true wave function using a relatively small number of terms but is slowly convergent. To guard against errors due to using a fixed number of ξ terms, we have included more ξ terms in our wave functions than Kono and Hattori.²⁰ For the second set²⁰ (referred to as the ζ set), $\xi = \eta = \zeta$, where ζ is the (single) nonlinear parameter to be optimized. Terms are chosen for this set by taking all terms for which

$$
i + j + k + |i - j| (1 - \delta_{k0}) \le \nu
$$
, (4)

where δ_{k0} is the Kronecker δ . The nonlinear parameter was optimized for each wave function essentially by determining whether a change in sign of the derivative of $E(\zeta)$ has occurred over the interval ζ to $\zeta + 2\epsilon$ (ϵ is an arbitrary small interval). This requires solving the eigenvalue equation for a series of ζ points taken at intervals of 0.1 (when $\zeta \ge 1$) or 0.05 (when $\zeta < 1$) in atomic units (a.u.). In order to carry out convergence studies, calculations were performed for $v=8$, 10, 12, and 14. Following Pekeris and co-workers, the eigenvalues have been extrapolated, using the formula $¹¹$ </sup>

$$
E_{\text{extrapolated}} = E_1 + \frac{(E_1 - E_0)(E_2 - E_1)}{2E_1 - E_0 - E_2}, \qquad (5)
$$

where E_0 , E_1 , and E_2 are the values at $v= 10$, 12, and 14, respectively.

III. NONRELATIVISTIC ENERGIES-RESULTS AND DISCUSSIQN

The calculations reported here were carried out at the National Bureau of Standards on a Cyber 205 in double precision (\sim 29 digits).

A. $1D, 3D$ terms

The results of our ${}^{1}D$ and ${}^{3}D$ calculations are reported in Tables I and II. These calculated values are rigorous upper bounds to the exact nonrelativistic energies. In these tables, we have also included our estimates of the exact nonrelativistic energies of these terms, obtained using Eq. (5). Our estimates are the values labeled "Extrapolated." These results show a definite convergence of the eigenvalues as the number of expansion terms is increased. Note that the convergence is not degraded for higher principal quantum numbers. In Table III we compare these results with the best previous calculations of these energy eigenvalues, the calculations of Kono and Hattori,^{6,7} and the calculations of Sims, Parmer, and Reese.¹

Our calculations essentially take the calculations of Kono and Hattori one step further to $v=14$. We increased the value of v (a) to obtain better eigenvalues and (b) to test the extrapolation formula used by both Kono and Hattori and us. Our calculated results are consistent with Kono and Hattori's results for all ν , but our estimates of the exact nonrelativistic energies of these terms are in all cases higher than those of Kono and Hattori. Thus it appears that the extrapolation formula overshoots the exact nonrelativistic energies (our wave functions are more complete than those of Kono and Hattori). This suggests that using the difference between calculated and estimated energies as a measure of the uncertainty in the estimated energies should be a conservative procedure. Comparing our present calculations with our earlier results makes it evident that the previously calculated energies got progressively worse as the principal quantum number increased. In contrast, the accuracy of the present results increases with increasing n . The worsening of the energy eigenvalues as n increased in the previous work (which we took note of in the uncertainty in our energy estimates) was due primarily to two factors: (1) the lack of optimization of the nonlinear parameters employed in the calculation and (2} the choice of terms employed in the wave functions. The comparison sugemployed in the wave functions. The comparison sug-
gests that the addition of the term $|i-j|$ $(1-\delta_{k0})$ in Eq. (4) significantly improves the energy convergence of the wave functions, especially for higher n (we previously employed $i + j + k \leq v$ as the selection criterion). This term eliminates some expansion terms with nonzero k and eminiates some expansion terms with nonzero κ and large $|i-j|$, a desirable elimination since where large

 $|i-j|$ terms are important (electrons far apart), the r_i dependency is small. Using the difference between calculated and estimated energies as a measure of the uncertainty of the estimated exact energies, we estimate the uncertainties for the D terms reported in Tables I and II to certainties for the *D* terms reported in Tables I and II to
vary from 3×10^{-10} a.u. in the worst case to 10^{-11} a.u. in the best case.

B. ${}^{1}F, {}^{3}F$ terms

The results of our ${}^{1}F$ and ${}^{3}F$ calculations (rigorous upper bounds to the exact nonrelativistic energies) are reported in Tables IV and V. In these tables we have again also included our estimates (the values labeled "Extrapolated") of the exact nonrelativistic energies of these terms, obtained using Eq. (5). A definite convergence of the eigenvalues is again evident as the number of expansion terms is increased. The convergence is not degraded for higher principal quantum numbers, but appears to be about the same for all principal quantum numbers.

In Table VI we compare these results with the best previous calculations, the calculations of Godefroid and Verhaegen,²¹ the calculations of Brown,²² and those of Sims, Parmer, and Reese.¹³ Comparing our present ${}^{1}F$ results with those of Ref. 13, we note again that in the previous work the calculated energies got worse as the principal quantum number increased, a defect the present calculations correct. Our present results are consistent with the previous results. Using the difference between calculated and estimated energies as a measure of the uncertainty in the estimated energies, we estimate the uncertainties for the F terms reported in Tables IV and V to tainties for the *F* terms reported in Tables IV and V
vary from 6×10^{-11} a.u. in the worst case to 3×10^{-11} a.u. in the best case.

IV. ENERGIES FOR ⁿ D AND ⁿ ELEVELS INCLUDING RELATIVISTIC AND OTHER CONTRIBUTIONS

A. Exchange energies, theoretical 1sn/ level structures

The nonrelativistic singlet-triplet separations n^3D $n^{1}D$ and $n^{3}F-n^{1}F$ from our results yield very accurate values of the corresponding exchange energies K . For example, the difference between our final calculated value for the nonrelativistic 3^3D-3^1D separation and the extrapolated value, 1.557638×10^{-5} a.u., is only 2.4×10 a.u. We take this separation as $2K = 1.557638(30)$ \times 10⁻⁵ a.u. and include the corresponding value in MHz units under "Present calculation" in Table VII. (In the remainder of this paper ionization energies, or term

TABLE I. Values of nonrelativistic energies for ¹D terms of He I. N_{ξ} denotes the number of ξ expansion terms, N_{ζ} the number of ζ expansion terms, N_{tot} the total number of terms in the wave function, and ζ_{opt} the optimum value of ζ . Energies are in atomic units.

Term	\mathcal{V}	N_{ξ}	N_{ζ}	$N_{\rm tot}$	$\zeta_{\rm opt}$	Energy
3^1D	8	24	86	110	1.05	-2.05562072171
	10	24	162	186	1.20	-2.05562073214
	12	24	271	295	1.20	-2.05562073256
	14	24	414	438	1.20	-2.05562073279
				Extrapolated		-2.05562073307
4^1D	8	28	86	114	1.05	-2.03127983697
	10	28	162	190	1.20	-2.03127984454
	12	28	271	299	1.20	-2.03127984568
	14	28	414	442	1.10	-2.03127984603
				Extrapolated		-2.03127984619
5^1D	8	32	86	118	1.05	-2.02001582918
	10	32	162	194	1.10	-2.02001583488
	12	32	271	303	1.20	-2.02001583591
	14	32	414	446	1.20	-2.02001583609
				Extrapolated		-2.02001583613
6^1D	$\bf 8$	36	86	122	1.05	-2.01389822202
	10	36	162	198	1.10	-2.01389822646
	12	36	271	307	1.20	-2.01389822726
	14	36	414	450	1.20	-2.01389822739
				Extrapolated		-2.01389822742
7^1D	$\,8\,$	40	86	126	0.95	-2.01021002426
	10	40	162	202	1.10	-2.01021002771
	12	40	271	311	1.20	-2.01021002834
	14	40	414	454	1.20	-2.01021002843
				Extrapolated		-2.01021002845
8^1D	$\bf 8$	44	86	130	0.95	-2.00781650944
	10	44	162	206	1.10	-2.00781651198
	12	44	271	315	1.20	-2.00781651248
	14	44	414	458	1.20	-2.00781651255
				Extrapolated		-2.00781651256

values, are taken as positive quantities; the sign of any energy contribution ls positive or negative according to whether it increases or decreases the ionization energy.)

The values of the n^3F-n^1F separations from the energies in Tables IV and V converge more rapidly than the individual $n^{1}F$ and $n^{3}F$ energies; we estimate uncertain-
ties less than 2×10^{-11} a.u. for these separations as ob ties less than 2×10^{-11} a.u. for these separations as obtained from the final estimated individual energies. The corresponding values of these separations in MHz units are given in Table VII for $n = 4-8$.

Correlation contributions to the 1snd exchange energies are, of course, effectively included to a very high order in our variational singlet-triplet separations. Chang and Poe²³ calculated exchange energies for 1snd and $1snf$ configurations using Brueckner-Goldstone perturbation theory through second-order contributions. Their values, included under "Other calculations" in Table VII, have errors smaller than 1% for the 1snd configurations, but their $1snf$ exchange energies are too small by 10 to 15%. The nonrelativistic n^3D-n^1D separations obtained from Kono and Hattori's extrapolated energies⁷ agree with our Notion and Hattori's extrapolated energies agree with values within $< 10^{-4}$ cm⁻¹ (3 MHz) for all $n = 3-8$.

We can now use our results to obtain improved theoretical descriptions of the lsnd and lsnf level structures. The Hamiltonian of the Breit equation in the Pauli approximation leads to three "magnetic" interactions and a relativistic contribution, denoted here by Δ_r , in addition to the exchange energy.^{24,25} We also include the mass-polarization energy ϵ_M and, for 1snd terms, the Lamb shift Δ_L . Cok and Lundeen²⁵ have given a convenient tabulation of the 4×4 lsnl energy matrix using LS-coupling basis states. Since the relativistic, masspolarization, and Lamb-shift interactions are diagonal in the approximation of interest, the separations of the four lsnl levels are affected only by the differences of these interactions for the triplet and singlet states. We denote these differences by $\Delta_{r,s,t}$, $\epsilon_{M,s,t}$, and $\Delta_{L,s,t}$; thus, for examthese differences by $\Delta_{r, st}$, $\epsilon_{M, st}$, and Δ_{I}
ple, $\Delta_{r, st}$ (1snd) $\equiv \Delta_{r}(n^3D) - \Delta_{r}(n^1D)$.

B. $n \binom{3}{2}$ - $n \binom{1}{2}$ separations and semiempirical values for the relativistic contributions

In their treatment of the $1s3d$ configuration for 4 He I and ³He I, Sansonetti and Martin⁴ used the experimental data to evaluate several energy contributions, including the separation E_{st} . The latter quantity is defined as the triplet-singlet separation given by the energy matrix with all three magnetic interactions set equal to zero:

$$
E_{st} = 2K + \Delta_{r,st} + \epsilon_{M,st} + \Delta_{L,st} \tag{6}
$$

TABLE II. Values of nonrelativistic energies for ³D terms of He I. N_{ξ} denotes the number of ξ expansion terms, N_S the number of ζ expansion terms, N_{tot} the total number of terms in the wave function, and ζ_{opt} the optimum value of ζ . Energies are in atomic units.

Term	$\pmb{\nu}$	N_{ξ}	N_{ζ}	N_{tot}	$\xi_{\rm opt}$	Energy
$3^{3}D$	$\bf 8$	24	70	94	0.90	-2.05563629853
	10	24	137	161	1.10	-2.05563630757
	12	24	236	260	1.20	-2.05563630917
	14	24	369	393	1.20	-2.05563630941
				Extrapolated		-2.05563630945
4^3D	8	28	70	98	0.95	-2.03128883832
	10	28	137	165	1.10	-2.03128884575
	12	28	236	264	1.20	-2.03128884706
	14	28	369	397	1.20	-2.03128884732
				Extrapolated		-2.03128884738
5^3D	8	32	70	102	0.95	-2.02002102062
	10	32	137	169	1.10	-2.02002102625
	12	32	236	268	1.20	-2.02002102723
	14	32	369	401	1.20	-2.02002102737
				Extrapolated		-2.02002102739
6 ³ D	$\bf 8$	36	70	106	1.00	-2.01390141005
	10	36	137	173	1.10	-2.01390141455
	12	36	236	272	1.20	-2.01390141531
	14	36	369	405	1.20	-2.01390141541
				Extrapolated		-2.01390141543
7 ³ D	8	40	70	110	0.95	-2.01021210204
	10	40	137	177	1.10	-2.01021210525
	12	40	236	276	1.20	-2.01021210585
	14	40	369	409	1.20	-2.01021210593
				Extrapolated		-2.01021210594
8^3D	8	44	70	114	0.95	-2.00781793166
	10	44	137	181	1.10	-2.00781793416
	12	44	236	280	1.20	-2.00781793463
	14	44	369	413	1.20	-2.00781793469
				Extrapolated		-2.00781793470

Term	Authors	Calculated	Exact (estimated)
3^1D	Sims et al.	-2.0556207317	-2.05562074
	Kono and Hattori	-2.055620732	-2.055620734
	Present work	-2.05562073279	-2.05562073307
4 1D	Sims et al.	-2.0312798371	-2.03127986
	Present work	-2.03127984603	-2.03127984619
5^1D	Sims et al.	-2.0200157081	-2.020017
	Kono and Hattori	-2.020015836	-2.020015837
	Present work	-2.02001583609	-2.02001583613
8^1D	Kono and Hattori	-2.0078165125	-2.0078165127
	Present work	-2.00781651255	-2.00781651256
$3^{3}D$	Kono and Hattori	-2.055636308	-2.055636310
	Present work	-2.05563630941	-2.05563630945
$5~^3D$	Kono and Hattori	-2.020021027	-2.020021028
	Present work	-2.02002102737	-2.02002102739
$8~^3D$	Kono and Hattori	-2.0078179346	-2.0078179348
	Present work	-2.00781793469	-2.00781793470

TABLE III. Selected nonrelativistic energy estimates for ${}^{1}D$ and ${}^{3}D$ terms of He I. Energies are in atomic units.

Sansonetti and Martin used an energy matrix including E_{st} and the three magnetic interactions but omitting the $\Delta_{r, st}$ and $\Delta_{L, st}$ contributions. They assumed an experimental value for the ofF-diagonal spin-orbit matrix element²⁶ and evaluated the other two magnetic interactions and E_{st} by parametrically adjusting their values to obtain optimal agreement of the resulting calculated level separations with the experimental separations. The neglect of the $\Delta_{r, st}$ and $\Delta_{L, st}$ contributions did not significant affect the values obtained for E_{st} and for the two diagonal magnetic interactions; the effect of the omitted contributions was mainly absorbed by the exchange-energy parameter.

We have evaluated the $\Delta_{r, st}$ energies for ⁴He 1s3d and

Term	$\boldsymbol{\nu}$	N_{ξ}	N_{ζ}	N_{tot}	$\zeta_{\rm opt}$	Energy
$4~{}^{1}F$	$\bf 8$	24	66	90	0.75	-2.03125513887
	10	24	140	164	0.80	-2.03125514363
	12	24	254	278	0.85	-2.03125514420
	14	24	414	438	0.95	-2.03125514434
				Extrapolated		-2.03125514439
$5~{}^1F$	8	28	66	94	0.75	-2.02000293056
	10	28	140	168	0.80	-2.02000293623
	12	28	254	282	0.85	-2.02000293693
	14	28	414	442	0.95	-2.02000293714
				Extrapolated		-2.02000293714
$6~{}^1F$	8	32	66	98	0.75	-2.01389067728
	10	32	140	172	0.80	-2.01389068291
	12	32	254	286	0.85	-2.01389068360
	14	32	414	446	0.95	-2.01389068376
				Extrapolated		-2.01389068381
$7~{}^{1}F$	8	36	66	102	0.75	-2.01020524204
	10	36	140	176	0.80	-2.01020524727
	12	36	254	290	0.85	-2.01020524788
	14	36	414	450	0.95	-2.01020524803
				Extrapolated		-2.01020524808
$8~^1F$	8	40	66	106	0.75	-2.00781329172
	10	40	140	180	0.80	-2.00781329640
	12	40	254	294	0.85	-2.00781329695
	14	40	414	454	0.95	-2.00781329707
				Extrapolated		-2.00781329710

TABLE IV. Values of nonrelativistic energies for ¹F terms of He I. N_{ξ} denotes the number of ξ expansion terms, N_{ζ} the number of ζ expansion terms, N_{tot} the total number of terms in the wave function, and ζ_{opt} the optimum value of ζ . Energies are in atomic units.

Energy Term $\boldsymbol{\nu}$ N_ξ N_{ζ} $N_{\rm tot}$ &opt 90 0.75 -2.031 255 163 27
-2.031 255 167 83 4^3F 8 24 66 10 24 140 164 0.75 278 0.80 —2.⁰³¹ ²⁵⁵ ¹⁶⁸ ²⁴ —2.⁰³¹ ²⁵⁵ ¹⁶⁸ ³⁶ 12 24 254 24 438 0.95 14 414 $-2.031 255 168 41$
 $-2.020 002 951 48$
 $-2.020 002 956 75$
 $-2.020 002 957 32$
 $-2.020 002 957 37$
 $-2.013 890 691 88$ Extrapolated 5^3F 8 94 0.75 28 66 10 28 140 168 0.80 282 0.90 12 28 254 14 28 414 442 0.95 Extrapolated 6^3F 8 32 98 0.75 66 $-2.013\,890\,697\,69$
 $-2.013\,890\,698\,16$
 $-2.013\,890\,698\,36$
 $-2.010\,205\,253\,42$ 10 32 140 172 0.80 32 254 286 0.85 12 32 446 0.95 14 414 Extrapolated $7~{}^{3}F$ 8 36 66 102 0.75 10 36 140 176 0.80 $-2.010\,205\,257\,88$
 $-2.010\,205\,258\,23$
 $-2.010\,205\,258\,39$
 $-2.007\,813\,300\,35$
 $-2.007\,813\,304\,12$
 $-2.007\,813\,304\,42$ 12 36 254 290 0.85 14 36 414 450 0.95 Extrapolated $8~^3F$ 8 40 66 106 0.75 180 10 40 140 0.80 12 40 254 294 0.85 —2.⁰⁰⁷ ⁸¹³ ³⁰⁴ ⁵¹ —2.⁰⁰⁷ ⁸¹³ ³⁰⁴ ⁵⁵ 454 0.95 14 40 414 Extrapolated

TABLE V. Values of nonrelativistic energies for ³F terms of He I. N_{ξ} denotes the number of ξ expansion terms, N_c the number of ζ expansion terms, N_{tot} the total number of terms in the wave function, and ζ_{opt} the optimum value of ζ . Energies are in atomic units.

³He 1s3d using Eq. (6) and the previously derived E_{st} values. The relevant data are given in Table VIII. In addition to our value for the exchange energy, we used calculated $\epsilon_{M, st}$ and $\Delta_{L, st}$ contributions from Kono and Hattori.⁷ The estimated errors of the latter quantities are too small to contribute significantly to the uncertainties of the resulting $\Delta_{r,st}$ values. We have included the ³He data

in Table VIII mainly to show the consistency of the $\Delta_{r, st}$ values obtained for 4 He and 3 He. This consistency depends on the fact that the calculated $\epsilon_{M,st}$ ⁽⁴He) value used in the table (-247.6 MHz) falls well within the range of values -243 ± 18 MHz required to fit the difference in the E_{st} values⁴ for ⁴He and ³He.

Kono and Hattori⁷ evaluated the relativistic contribu-

TABLE VI. Selected nonrelativistic energy estimates for 1F and 3F terms on He I. Energies are in atomic units.

Term	Authors	Calculated	Exact (estimated)
4 ¹ F	Godefroid and Verhaegen	-2.031255	
	Brown	-2.03125428	-2.031255
	Sims et al.	-2.0312551441	-2.031255145
	Present work	-2.03125514434	-2.03125514439
5 ¹ F	Brown	-2.02000164	-2.020004
	Sims et al.	-2.0200029316	-2.02000295
	Present work	-2.02000293709	-2.02000293714
6^1F	Brown	-2.01388950	-2.013889
	Sims et al.	-2.0138906649	-2.0138907
	Present work	-2.01389068376	-2.01389068381
$4^{3}F$	Godefroid and Verhaegen	-2.031255	
	Brown	-2.03125432	-2.031255
	Present work	-2.03125516836	-2.03125516841
5^3F	Brown	-2.02000168	-2.020004
	Present work	-2.02000295732	-2.02000295737
6^3F	Brown	-2.01388952	-2.013889
	Present work	-2.01389069830	-2.01389069836

TABLE VII. Nonrelativistic singlet-triplet separations for the 1s3d and 1snf ($n = 4-8$) configurations. These separations are equal to twice the exchange energies $2K(1snl)$. Units are MHz.

Separation	2K Present calculation	2K Other calculations	
$3^{3}D-3^{1}D$	102 473.6(20)	$102475(15)^{a}$	
		102000 ^b	
$4^{3}F - 4^{1}F$	158.02(13)	137 ^b	
$5^{3}F - 5^{1}F$	133.09(13)	119 ^b	
$6^3F - 6^1F$	95.72(13)	86 ^b	
$7^{3}F - 7^{1}F$	67.83(13)	61 ^b	
$8^{3}F - 8^{1}F$	49.01(13)	44 ^b	

'Kono and Hattori, Ref. 7.

^bChang and Poe, Ref. 23.

tions using their variationally calculated wave functions and the formula:

$$
\Delta_r = \alpha^2 \left[-\frac{Z^4}{8} + \frac{1}{4} \langle \mathbf{p}_1^4 \rangle - \pi Z \langle \delta(\mathbf{r}_1) \rangle - \pi \langle \delta(\mathbf{r}_{12}) \rangle \right] + \frac{\alpha^2}{2} \left\langle \frac{\mathbf{p}_1 \cdot \mathbf{p}_2}{r_{12}} + \frac{\mathbf{r}_{12} \cdot (\mathbf{r}_{12} \cdot \mathbf{p}_1) \mathbf{p}_2}{r_{12}^3} \right\rangle, \tag{7}
$$

where α is the fine-structure constant. Table IX includes their $\Delta_r({}^3D)$, $\Delta_r({}^1D)$, and $\Delta_{r,st}$ values for the 1s3d and ls7d configurations. The total calculated ionization energies $T(^3D)$ and $T(^1D)$ are also listed for these configurations.

The apparently more accurate $\Delta_{r,st}$ (1s3d) value derived semiempirically in Table VIII and given under "Present results" in Table IX indicates an error of about 16 MHz in Kono and Hattori's $\Delta_{r,st}$ value. This error is

TABLE VIII. Energy differences pertinent to semiempirical evaluation of the relativistic contribution to the ls3d singlettriplet separation. The E_{st} value represents the ${}^{3}D-{}^{1}D$ separation in the absence of the "magnetic" interactions. The E_{st} values for 4 He and 3 He 1s3d were obtained previously by fitting the experimental level structures (Refs. 4 and 26). Calculated values are given for the exchange $(2K)$, mass-polarization $(\epsilon_{M,st})$, and Lamb-shift ($\Delta_{L, st}$) contributions to E_{st} . The relativistic contribution $\Delta_{r,st}$ was evaluated by subtracting the other contributions from the E_{st} value. Units are MHz.

'Reference 4.

Reference 26.

'Present result.

^dValue for ³He is 4.6 MHz smaller than for ⁴He due to the difference of the reduced electron mass for the two isotopes (normal mass effect).

'Reference 7.

^fValue for ³He is equal to 1.3272 ϵ_M ⁽⁴He), the numerical factor being the ratio of the 4 He and 3 He nuclear masses.

also essentially equal to the difference between the experialso essentially equal to the difference between the experimental $3³D-3¹D$ separation,⁴ 102 233(3) MHz, and Kono and Hattori's calculated value including all contributions, 102217(24) MHz, since Kono and Hattori's calculated nonrelativistic $3³D-3¹D$ separation agrees with our value within about 1 MHz (Table VII). As shown previously,⁹ Kono and Hattori's calculated $n \, {}^{3}D \cdot n \, {}^{1}D$ separations including all contributions are systematically smaller than the experimental values by amounts varying approxi-

TABLE IX. Relativistic contributions and total ionization energies for the 1s3d and 1s7d terms. The differences of the Δ_r ⁽³D) and Δ_r ⁽¹D) relativistic contributions are labeled $\Delta_{r, st}$. The total ionization energies $T({}^3D)$ and $T({}^1D)$ under Kono and Hattori are the sums of all contributions calculated in Ref. 7, including the Lamb shifts. The $T({}^3D)$ and $T({}^1D)$ energies under Present results were obtained by combining the adjusted Δ , values in this column with the present nonrelativistic ionization energies (Tables I and II) and using the other contributions from Ref. 7. The energies in cm⁻¹ units are based on a value of 109 722.273 50(7) cm^{-1} for the ⁴He Rydberg constant (Ref. 27).

	Kono and Hattori	Present results
1s3d		
$\Delta_r({}^3D)$	833(9) MHz	$841(6) \text{ MHz}^{\text{a}}$
$\Delta_r(^1D)$	$878(9)$ MHz	$870(6)$ MHz ^a
$\Delta_{r,st}$	$-45(13)$ MHz	$-29(4)$ MHz
$T(^3D)$	$12\,209.1126(6)$ cm ⁻¹	$12\,209.112\,67(28)$ cm ⁻¹
$T({}^{1}D)$	$12\,205.7030(5)$ cm ⁻¹	$12\,205.702\,56(25)$ cm ⁻¹
1s7d		
$\Delta_r({}^3D)$	$135.8(3)$ MHz	$137.2(9)$ MHz ^a
$\Delta_r({}^1D)$	$141.8(3)$ MHz	140.4(9) $MHza$
$\Delta_{r,st}$	$-6.0(4)$ MHz	$-3.2(8)$ MHz
$T(^3D)$	$2240.99533(10)$ cm ⁻¹	2 240.995 33(4) cm ⁻¹
$T({}^{1}D)$	$2240.54066(10)$ cm ⁻¹	$2240.54057(4)$ cm ⁻¹

^aValue obtained by semiempirical adjustment of Δ , value from Ref. 7; see text.

mately as $140/n^2$ MHz for $n = 3-8$. Applying the procedure outlined above to the data and calculated energies for the higher 1snd configurations, we obtain results similar to those for 1s3d: the systematic errors of Kono and Hattori's n^3D-n^1D separations are apparently mainly due to errors in their calculated Δ , contributions.

Although we have not calculated the relativistic and other small contributions using our wave functions, we have combined the available data to deduce improved values for the *n D* relativistic contributions. The $\Delta_r(^3D)$ and $\Delta_r({}^1D)$ values given under "Present results" in Table IX were obtained by increasing Kono and Hattori's Δ_r ³D) values by 70/n² MHz and decreasing their $\Delta_r({}^1D)$ values by the same amounts. The $T({}^3D)$ and $T({}^{1}D)$ ionization energies were then evaluated by adding these "adjusted" Δ , values and all other small contributions⁷ to our nonrelativistic ionization energies.²⁷ The $1s7d$ results are given as an example of this procedure for the higher *n* values. The $\Delta_r(^3D)$ and $\Delta_r(^1D)$ values obtained for $n = 3-8$ by adjustment of Kono and Hattori's values in this manner are supported by the agreement, within small deviations, of the resulting $T(^3D)$ and $T({}^{1}D)$ energies with the ${}^{3}D$ and ${}^{1}D$ ionization energies adopted in Ref. 9; the latter values for $n = 5-8$ are ultimately based on core-polarization theoretical $n G$ and n H term values²⁸ with uncertainties significantly smaller than the uncertainties of Kono and Hattori's Δ , values. The adjusted Δ , values and total n D ionization energies for $n = 4, 5, 6, 8$ may be easily obtained as above. The resulting total ionization energies for $n = 5-8$ agree with the values adopted in Ref. 9 within a maximum deviation of 7×10^{-5} cm⁻¹ (2 MHz) for $T(5^1D)$.

As noted above, Kono and Hattori's⁷ extrapolated values for the nonrelativistic ionization energies $T_{nr}({}^3D)$ and $T_{n,r}({}^{1}D)$ are systematically too large. Our results indicate that their total ${}^{3}D$ ionization energies include approximate cancellations of the errors of their $T_{nr}({}^3D)$ and Δ_r (³D) values, whereas the errors of their T_{nr} ⁽¹D) and $\Delta_r({}^1D)$ values are in the same direction. The $2{}^3S_1$ ionization energy derived previously⁹ by combining experimental data with Kono and Hattori's $T(4^3D)$ and $T(5³D)$ total ionization energies is not significantly changed by use of "adjusted" theoretical $T(n^3D)$ and $T(n⁻¹D)$ energies obtained as above.

The difference between our $\Delta_{r,st}$ (1s3d) value in Table IX and Kono and Hattori's value⁷ is fairly consistent with their uncertainty estimates of about 1% for the Δ_r (3³D) and Δ_r (3¹D) contributions. The discrepancy of 2.8 MHz between our Δ_{rst} (1s7d) value and their value again indicates errors of about 1% in their Δ_r (7³D) and Δ_r (7¹D) values, significantly larger than their uncertainty estimates of 0.2% for these quantities. Evaluations of the Δ , contributions with more accurate wave functions are clearly needed.

C. $1sn f$ level structures and ionization energies

We have used our nonrelativistic ionization energies for $n³F$ and $n¹F$ terms together with $1snf$ energy matrix elements calculated by Cok and Lundeen²⁵ to obtain predicted ionization energies for the levels including relativistic and other contributions. The three magnetic interactions h_{so} (spin-orbit), h_{off} (off-diagonal element), and h_{ss} (spin-spin), as well as the Δ_r , and $-\epsilon_M$ energies in Table X, were calculated in an "extended adiabatic" approximation.²⁵ The numbers in parentheses following the three magnetic-interaction values are the extendedadiabatic corrections to the corresponding hydrogenicapproximation values; we take these corrections as roughly indicative of the uncertainties.

The results of diagonalizing the $1snf$ energy matrices with values of the exchange energies from Table VII and the matrix elements from Ref. 25 are given under "Calculation" in Table XI. The quantity $\Delta T({}^{1}F)$ is the calculated increase in the ${}^{1}F$ ionization energy with respect to the nonrelativistic (variational) ${}^{1}F$ position. The uncertainties given for the $\Delta T({}^{1}F)$ displacement and for our calculated values of the three $1snf$ fine-structure separations include only the exchange and magnetic-interactions uncertainties; we have not attempted to estimate the errors of the Δ , and ϵ_M energies from Ref. 25. The Lamb shift were assumed to be negligible compared to other approximations in these calculations.

The experimental $1s5f$ and $1s6f$ fine structures in Table XI are from measurements of $n F-n G$ transitions by Farley et $al.^{29}$ Our estimates of the errors, which are very rough, are in general larger than would be obtained from the standard deviations of the individual determinations, in accordance with the discussion in Ref. 29. A similar comment applies to the errors given with the experimental $1s7f$ fine structure separations, which are from ⁷ D-7 F measurements by Farley, MacAdam, and Wing.³⁰ We averaged data from Cok and Lundeen's $8 F-8 G$ determinations³¹ and from the $8 D-8 F$ measure

TABLE X. Values for the magnetic (h_{so}, h_{off}, h_{ss}), relativistic (Δ), and mass-polarization ($-\epsilon_M$) energies for lsnf configurations $(n = 4-8)$. All values from Ref. 25. Units are MHz.

1snf	n_{so}	n _{off}	n_{ss}	$\Delta.(^1F)^a$	$\Delta (^3F)^a$	$-\epsilon_M(^1F)^a$	$-\epsilon_M({}^3F)^{\rm a}$
1s4f	$-32.509(55)$	$-97.841(48)$	32.598(50)	268.97	268.99	9.53	9.25
1s5f	$-16.639(37)$	$-50.081(38)$	16.679(36)	190.27	190.29	5.49	5.25
1s6f	$-9.625(26)$	$-28.976(28)$	9.649(24)	130.38	130.40	3.38	3.20
1s7f	$-6.060(18)$	$-18.245(20)$	6.074(17)	91.23	91.25	2.20	2.08
1s8f	$-4.059(13)$	$-12.222(14)$	4.070(11)	65.70	65.70	l.50	1.42

^aThe quantities Δ , and $-\epsilon_M$ are tabulated here for consistency with other parts of this paper. The absolute values of these energies are from the values of the H_{rel} and H_{mp} matrix elements in Table V of Ref. 25.

comparison. Units are MHz.

^aSeparations derived from measurements of $n F$ -n G transitions $(n = 5, 6)$ given in Ref. 29.

^bSeparations derived from measurements of 7 D-7 F transitions given in Ref, 30.

 e Separations derived from measurements of 8 D-8 F transitions given in Ref. 30 and 8 $F-8$ G transitions in Ref. 31.

ments of Ref. 30 to obtain the experimental $1s8f$ finestructure separations, the two sets of measurements agreeing within about 0.2 MHz.

Most of our calculated $1snf$ fine-structure separations agree with the experimental values within the combined estimated errors. It is interesting to note in connection with these results that only for $l = 3$ is the 1snl exchange energy roughly comparable in order of magnitude with the corresponding magnetic and relativistic energies. Cok and Lundeen²⁵ calculated the fine structures for several 1snf configurations using their values for all interaction energies except the exchange energy and evaluating the latter quantity parametrically by fitting the experimental structure. The effect of any possibly significant diagonal interaction omitted from the matrix would have been absorbed by the exchange-energy value. Our results show that any such omitted contribution is at most at the level of the uncertainties.

Farley³² obtained the $1snf$ fine-structure separations given under "Series formula" in Table XI by fitting experimental measurements²⁹⁻³¹ to the formul

$$
T(n^{3}F_{J}) - T(n^{1}F_{3}) = An^{-3} + Bn^{-5}
$$
 (8)

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and evaluating the constants A and B for each of the three series of intervals [three different J values for n^3F_J in (8)]. Our calculated separations agree with the seriesformula values within the errors in only about half the cases in Table XI, but the standard deviations quoted for the series-formula values are known to be typically smaller than the actual errors.²⁹ The series-formula predic tion³² for the 4³ F_3 -4¹ F_3 separation is apparently low by about 26 MHz, which is about 16 times the standard deviation obtained from the fit of the data for higher n values. A good fit of these separations down to $n = 4$ would probably require at least a third term on the righthand side of (8).

In Table XII we compare $1snf¹F$ ionization energies based on two very different types of calculations: the "variational" $T_r({}^1F)$ term values in the fourth column were obtained by addition of the $\Delta T({}^{1}F)$ corrections from Table XI to our calculated $T_{nr}^{1}(F)$ values from Table IV, whereas the $T_r({}^1F)$ values for $n=5-8$ under Ref. 9 were derived from Drachman's²⁸ core-polarization theoretical term values for $1s_ng$ or $1s_nh$ configurations combined with microwave-spectroscopic data. We note again that the errors for our $\Delta T({}^{1}F)$ corrections include no allowances for errors of the Δ , relativistic contributions, and thus the errors given for the variational $T_{\nu}({}^{1}F)$ energies may be unrealistically small. The difference between the variational $T_r({}^1F)$ energy for each $n = 5-8$ and the corresponding $T_r({}^1F)$ value based on the core polarization theoretical $n \nG$ or n H term value is given in the last column. The variational $T_r({}^1F)$ values are systematically larger than the core-polarization/microwave-data values by about 3 MHz $(n = 5)$ to 1 MHz $(n = 7, 8)$. These differences are comparable to the estimated errors of the $T_r({}^1F)$ values from Ref. 9; in addition, the actual errors of our $\Delta T({}^{1}F)$ values could be of the order of a few MHz.

The $1s4f$ levels have not yet been determined experimentally with high accuracy. The trend of the systematic differences between our $T_r({}^1F)$ values and those based on theoretical n G and n H energies for $n = 5-8$ indicates a corresponding difference of about 0.0002 cm^{-1} (6 MHz for $n = 4$; this extrapolated difference can be taken as a minimum uncertainty estimate for our predicted $4¹F$ ionization energy. The most accurate previous prediction for the $4^{T}F$ position was obtained in Ref. 9 by use of the nonrelativistic energy from Ref. 13 and an "adjusted" theoretical value of 77.5 MHz for the exchange energy K . As noted in Ref. 9, the net calculated displacement ΔT (4 ¹F) is relatively small because the 4 ¹F term-value increase due to the relativistic Δ , contribution is almost canceled by the decrease due to the triplet-singlet repulsion. The $T_r({}^1F)$ energy from Ref. 9 agrees with our new value within 3 MHz (Table XII), and the separations of the $1s4f$ levels in Ref. 9 agree with our predictions in Table XI within a maximum deviation of about 2 MHz.

V. CONCLUSION

Combined CI-Hy techniques have proven effective in obtaining new values of the nonrelativistic energies of the

TABLE XII. Ionization energies (term values) for $1snf'F$ levels ($n = 4-8$). The nonrelativistic term values $T_{\text{nr}}({}^1F)$ in the second column are from the present calculations (Table IV). These values were corrected by addition of the calculated $\Delta T({}^{1}F)$ values from Table IX to obtain the $T_{r}({}^{1}F)$ values in the fourth column. The $T_r({}^1F)$ values for $n = 5-8$ from Ref. 9 (fifth column) were obtained by a combination of core-polarization theoretical 1sng or 1snh term values (Ref. 28) and microwave-spectroscopic data. The difference of the $T_r({}^1F)$ values in the fourth and fifth columns is given in the final column. All term values are based on a value of 109 722.273 50(7) cm^{-1} for the ⁴He Rydberg constant (Ref. 27). Units are cm^{-1} .

$\ln f$	$T_{\text{nr}}({}^{1}F)$ Variational	$\Delta T({}^1F)$	$T.(^1F)$ Variational	$T_r({}^1F)$ Ref. 9	Difference
1s4f	6858.771 002(11)	0.000723(6)	6858.771725(13)	6858.7718(6) ^a	$-0.0001(6)$
ls5f	4389.535 479(11)	0.002723(5)	4389.538 202(12)	4389.538 10(7) ^b	0.00010(7)
1s6f	3048.234 816(11)	0.002437(4)	3048.237 253(12)	3048.237 19(5) ^c	0.00006(5)
ls7f	2239.486042(11)	0.001906(3)	2239.487948(12)	2239.487 918(37) ^d	0.000030(39)
				2239.487922(18) ^e	0.000026(22)
1s8f	1714.585 443(6)	0.001460(2)	1714.586 903(6)	1714.586 879(30) ^f	0.000024(31)
				1714.586 866(19) ⁸	0.000037(21)

'Based on a previous variational calculation (Ref. 13).

 b Based on 5 G theoretical term value.

 c Based on 6 G theoretical term value.

 d Based on 7 G theoretical term value.

 e Based on 7 H theoretical term value.

 f Based on 8 G theoretical term value.

 B Based on 8 H theoretical term value.

 $n^{-1}D$ and $n^{-3}D$ terms $(n=3-8)$ and the $n^{-1}F$ and $n^{-3}F$ terms $(n = 4-8)$ of neutral helium. These calculations are superior to all previous results reported in the literature. We suggest that the exact nonrelativistic energies differ from our extrapolated energies by no more than 3×10^{-10} a.u. in the worst case and 10^{-11} a.u. in the best case.

Combining these results with available calculated values for relativistic and other smail contributions, we have obtained new predictions for the ionization energies of the $n D$ and $n F$ energy levels. The predicted $n D$ and n F ionization energies for $n = 5-8$ are consistent with core-polarization theoretical $n G$ and $n H$ ionization energies and connecting data within uncertainties of the order of ¹ to several MHz. These results and comparisons with experimentally determined n^3D-n^1D separations $(n=3-8)$ lead to suggested semiempirical values for the $n³D$ and ¹D relativistic contributions that should be checked and improved by more accurate calculations. Our calculated $n \nightharpoondown F$ fine-structure separations agree with

the experimental values for $n = 5-8$ within deviations of the order of the experimental uncertainties (-1) to 0.1 MHz), but more accurate evaluations of the relativistic, magnetic, and mass-polarization contributions would also allow improved tests of the theory for these configurations.

After this paper was submitted for publication, a paper by Schilling et aI .³³ appeared giving measurements of the ls4f level separations with estimated uncertainties of ¹ to 2 MHz. Their experimental values for the three separations agree with our predicted values in Table XI within a maximum deviation of 0.22 MHz.

ACKNOWLEDGMENTS

The authors would like to express their gratitude to Dr. Bruce R. Miller and Dr. John R. Rumble for helpful discussions about this work.

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from Refs. 7 and 9 have been corrected to this $R(^{4}He)$ value, which was obtained by applying the usual reduced-mass correction, $R({}^{4}\text{He})=R_{\infty}[1+(m_e/M_{\alpha})]^{-1}$, to the value $R_{\infty} = 109737.31569(7)$ cm⁻¹ [P. Zhao, W. Lichten, H. P. Layer, and J. C. Bergquist, Phys. Rev. A 34, 5138 (1986)]. The ratio of the electron mass to the ⁴He nuclear mass m_e/M_a was taken as 1.370 933 54(3) $\times 10^{-4}$ [E.R. Cohen and B. N. Taylor, CODATA Bull. 63, ¹ (1986)].

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