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Computer-generated formulas for three-center nuclear-attraction integrals (electrostatic potential) for Slater-type orbitals

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The computer-assisted C-matrix, Löwdin-alpha-function, single-center expansion method in spherical harmonics has been applied to the three-center nuclear-attraction integral (potential due to the product of separated Slater-type orbitals). Exact formulas are produced for 13 terms of an infinite series that permits evaluation to ten decimal digits of an example using 1s orbitals.

I. INTRODUCTION

Efforts continue to be made to facilitate the use of Slater-type orbitals (STO's), or, more generally speaking, exponential-type orbitals in problems of *ab initio* quantum chemistry and molecular physics.^{1,2} Here, we apply our method of computer-assisted single-center expansion in spherical harmonics to the difficult case of the three-center nuclear-attraction integral, which is the same as the problem of finding the potential due to a charge distribution that is given by the product of separated STO's. Earlier, partial results had been achieved for this integral by the use of elliptical coordinates³ and transform methods.⁴ We will make numerical comparisons with results recently obtained by Trivedi and Steinborn⁵ in which a transform method using the Steinborn *B* function is employed.

The initial work in single-center expansion methods was carried out by Coolidge.⁶ A suggestive formulation of this method with the use of coefficient matrices and the α -function notation was introduced by Löwdin.⁷ An essential simplification and the use of a "C matrix" was given to the closed form of the α function by Jones and Weatherford⁸ after its presentation by Sharma.⁹ Thus far, computer-generated formulas have been produced for two-center overlap,^{10,11} Coulomb,¹² hybrid,¹³ and exchange integrals;¹⁴ some three-center Coulomb-type formulas also have been generated.¹⁵

II. DERIVATION OF THE POTENTIAL

We place a STO, χ_a , at the origin of our coordinate system (r, θ, ϕ) to represent one of the factors of the expression for the electron density of electron 1:

$$\chi_a(1) = A_a r_1^{N_a - 1} e^{-\zeta_a r_1} Y_{L_a}^{M_a}(\theta_1, \phi_1) .$$
 (1)

 $A_a = (2\zeta_a)^{N_a + 1/2} [(2N_a)!]^{-1/2}$ is the normalization factor; N_a , L_a , and M_a are the quantum numbers of the orbital; and ζ_a is the screening constant. We place the other factor χ_b for the density of electron 1 in its local coordinate system (R, Θ, ϕ) that has been translated a distance α along the z axis:

$$\chi_{b}(1) = A_{b} R_{1}^{N_{b}-1} e^{-\zeta_{b} R_{1}} Y_{L_{b}}^{M_{b}}(\Theta_{1}, \phi_{1}) .$$
⁽²⁾

In terms of the original coordinate system, the expansion in spherical harmonics is as follows:⁸

$$\chi_{b}(1) = \frac{A_{b}}{\zeta_{b}^{N_{b}-1}} \left[\frac{(2L_{b}+1)(L_{b}+M_{b})!}{4\pi(L_{b}-M_{b})!} \right]^{1/2} \sum_{l=M_{b}}^{\infty} \left[\frac{4\pi(l+M_{b})!}{(2l+1)(l-M_{b})!} \right]^{1/2} (-1)^{M_{b}} \alpha_{l}^{N_{b}L_{b}M_{b}} (\zeta_{b}\alpha,\zeta_{b}r_{1}) Y_{l}^{M_{b}} (\theta_{1},\phi_{1}) ,$$

where

$$\alpha_{l}^{N_{b}L_{b}M_{b}}(\zeta_{b}\alpha,\zeta_{b}r_{1}) = \frac{(2l+1)(l-M_{b})!}{2(l+M_{b})!} \sum_{i=0}^{N_{b}+L_{b}+l} \sum_{j=0}^{N_{b}+l} C_{l}^{N_{b}L_{b}M_{b}}(i,j)H_{ij}(\zeta_{b}\alpha,\zeta_{b}r_{1})(\zeta_{b}\alpha)^{i-L_{b}-l-1}(\zeta_{b}r_{1})^{j-l-1}$$

and

$$H_{ij}(\zeta_b a, \zeta_b r_1) = \begin{cases} e^{-\zeta_b a} [(-1)^j e^{\zeta_b r_1} - e^{-\zeta_b r_1}], & r_1 < a \\ e^{-\zeta_b r_1} [(-1)^j e^{\zeta_b a} - e^{-\zeta_b a}], & r_1 > a \end{cases}$$

(3)

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$$\sum_{i=0}^{N+L+l} \sum_{j=0}^{N+l} C_{l}^{NLM}(i,j) a^{i} r^{j} = \sum_{P=0}^{\lfloor (L+M)/2 \rfloor} \sum_{q=0}^{L+M-2p} \sum_{v=0}^{L+M-2p-q} \sum_{p'=0}^{\lfloor (l-M)/2 \rfloor} \sum_{q'=0}^{l-M-2p'} \sum_{v'=0}^{L-M-2p'-q} \sum_{k=0}^{t} \sum_{k'=0}^{t-k} \frac{a^{x}}{4^{L+l-p-p'}} \times \frac{r^{y}(-1)^{v+q'+p+p'}(2L-2p)!(2l-2p')!}{(L-p)!p!p'!q!q'!v!v'!(L+M-2p-q-v)!} \frac{(N-L+2p+2q+2q')![(N-L-2p+2q+2q'-k-k')!]^{-1}}{(l-p')!(L-M-2p'-q'-v')!k!}, \quad (4)$$

where

$$x = N + L + 2l - 2p' - 2v' - 2v - k - k'$$

y = 2(p' + v + v') + k',

and

$$t = N - L + 2p + 2q + 2q'$$

The symbols [] mean reduce to integer. [By multiplying by $(-1)^{l+M}$ it is possible to have all positive elements.] We seek the potential at an arbitrary point (r_2, θ_2, ϕ_2) due to the charge distribution $\chi_a(1)\chi_b(1)$. Thus

$$V(\vec{\mathbf{r}}_2) = \int dv_1 \frac{\chi_a \chi_b}{r_{12}} .$$
⁽⁵⁾

The Laplace expansion of $1/r_{12}$ is

$$\frac{1}{r_{12}} = 4\pi \sum_{\lambda=0}^{\infty} \sum_{m=-\lambda}^{\lambda} (2\lambda+1)^{-1} r_{<}^{\lambda} / r_{>}^{\lambda+1} Y_{\lambda}^{m^{*}}(\theta_{1},\phi_{1}) Y_{\lambda}^{m}(\theta_{2},\phi_{2}) , \qquad (6)$$

where $r_{<}$ is to be replaced by r_1 or r_2 , depending on which is smaller, and $r_{>}$ is to be replaced by r_1 or r_2 , depending on which is larger. By making the appropriate substitutions in Eq. (5), the radial and angular variables separate and we get

$$V(\vec{r}_{2}) = k_{0} \sum_{\lambda} \sum_{m} Y_{\lambda}^{m}(\theta_{2}, \phi_{2}) \sum_{l} (2\lambda - 1)^{-1} \left[\frac{(l + M_{b})!}{(2l + 1)(l - M_{b})!} \right]^{1/2} \\ \times \int r_{1}^{2} dr_{1} r_{1}^{N_{a} - 1} e^{-\xi_{a} r_{1}} \alpha_{l}^{N_{b} L_{b} M_{b}} (\xi_{b} \alpha, \xi_{b} r_{1}) \frac{r_{<}^{\lambda}}{r_{>}^{\lambda + 1}} \langle \lambda, m | L_{a} M_{a} | l, M_{b} \rangle , \qquad (7)$$

where

$$k_0 = \frac{A_a A_b}{\zeta_b^{N_b - 1}} (-1)^{M_b} \left[\frac{(2L_b + 1)(L_b + M_b)!}{(L_b - M_b)!} \right]^{1/2},$$

and the angular brackets represent the angular integration over the product of three spherical harmonics, i.e., the Gaunt coefficients.¹⁶ Let us examine the case of 1s orbitals with screening constants of 1. Carrying out the algebra and using

$$Y_l^0(\theta,\phi) = \left(\frac{2l+1}{4\pi}\right)^{1/2} P_l(\cos\theta) , \qquad (8)$$

we get

$$V(r_{2},\theta_{2}) = \sum_{l=0}^{\infty} V_{l}(r_{2})P_{l}(\cos\theta_{2})$$
(9)

with

$$V_{l}(r_{2}) = 2 \sum_{i=0}^{l+1} \sum_{j=0}^{l+1} C_{l}(i,j) a^{-l-1+i} \int dr_{1} r_{1}^{2} e^{-r_{1}} \frac{r_{<}^{l}}{r_{>}^{l+1}} H_{ij}(a,r_{1}) r_{1}^{-l-1+j} .$$
(10)

The integral can be reduced to a summation by employing formulas (slightly modified) from a standard integral table:¹⁷

$$\int x^n e^{-wx} dx = -e^{-wx} \sum_{k=0}^n \frac{n!}{(n-k)!} \frac{x^{n-k}}{w^{k+1}}, \quad n \ge 0$$
(11)

$$\int \frac{e^{-wx}}{x^n} dx = -e^{-wx} \sum_{k=1}^{n-1} \frac{(n-k-1)!}{(n-1)!} \frac{(-w)^{k-1}}{x^{n-k}} + \frac{(-w)^{n-1}}{(n-1)!} \operatorname{Ei}(-wx) , \quad n > 0.$$
⁽¹²⁾

For formula generation we use a very simple form of computer algebra.¹⁸ For instance, the expression $4\alpha^2 r^3$ would be

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represented by the array element (2,3)=4. Hence, to multiply terms we multiply the coefficients and add the powers of the variables separately. This can be expeditiously done under iteration. We end up with obvious functions multiplied by computer-generated polynomial coefficients. The potentials for $r < \alpha$ and $r > \alpha$ must be determined separately. Here, we examine the $r < \alpha$ case. For every *l* value we anticipate a formula with ten kinds of functions, each multiplied by a polynomial in α and r. The ten functions are as follows: $e^{-\alpha}$, $e^{-\alpha}\text{Ei}(-2r)$, $e^{-\alpha}\gamma$, $e^{-\alpha}\ln\alpha$, $e^{-\alpha}\lnr$, $e^{-3\alpha}$, $e^{-\alpha}e^{-2r}$, $e^{-\alpha}\text{Ei}(-2\alpha)$, and $e^{\alpha}\text{Ei}(-2\alpha)$, where $\gamma=0.577\cdots$ (Euler's constant). Actually, only two functions appear with nonzero coefficients for all *l* values, namely, $e^{-\alpha}$ and $e^{-\alpha}e^{-2r}$. Only in the cases of l=1 and 2 must we consider more than two functions. To be explicit, we write the formulas for l=0 and 1, dropping the subscript 2 on *r*:

$$V_{0}(r) = a^{-1}(-1/r + 1.5 - r + r^{2}/3)e^{-a} + a^{-1}(1/r + 0.5)e^{-a-2r} + a^{0}(-0.5/r + 2 - r)e^{-a} + a^{0}(0.5/r)e^{-a-2r} + a^{1}(1.0)e^{-a},$$
(13)

$$V_{1}(r) = a^{-2}(3.75/r^{2} - 9 + 13r + 1.5r^{2})e^{-a} + a^{-2}(6r)e^{-a}\text{Ei}(-2r) + a^{-2}(6r)e^{-a}\ln a + a^{-2}(6r)e^{-a}\ln r + a^{-2}(-3.75/r^{2} - 7.5/r + 1.5)e^{-a-2r} + a^{-2}(6r)e^{a}\text{Ei}(-2a) + a^{-1}(3.75/r^{2} - 9 + r + 1.5r^{2})e^{-a} + a^{-1}(6r)e^{-a}\text{Ei}(-2r) + a^{-1}(6r)e^{-a}\ln a + a^{-1}(-6r)e^{-a}\ln r + a^{-1}(-3.75/r^{2} - 7.5/r + 1.5)e^{-a-2r} + a^{-1}(6r)e^{-a}\text{Ei}(-2a) + a^{0}(1/r^{2} - 3 - 4r^{2}/3)e^{-a} + a^{0}(2r)e^{-a}\text{Ei}(-2r) + a^{0}(2r)e^{-a}\ln a + a^{0}(-2r)e^{-a}\ln r + a^{0}(-1/r^{2} - 2/r + 1)e^{-a-2r} + a^{0}(-2r)e^{a}\text{Ei}(-2a).$$
(14)

The lowest powers of α and r in the polynomials appear to be -l-1, with the highest power of α being 1 and the highest power of r being l+2. For the values $\alpha=2$, r=0.5 we obtain

$$V_{0} = 0.388\,330\,440\,4\,, V_{1} = 0.053\,464\,134\,5\,,$$

$$V_{2} = 0.007\,004\,740\,9\,, V_{3} = 0.000\,977\,590\,4\,,$$

$$V_{4} = 0.000\,148\,745\,9\,, V_{5} = 0.000\,024\,463\,7\,,$$

$$V_{6} = 0.000\,004\,283\,0\,, V_{7} = 0.000\,000\,787\,3\,,$$

$$V_{8} = 0.000\,000\,150\,4\,, V_{9} = 0.000\,000\,029\,6\,,$$

$$V_{10} = 0.000\,000\,006\,0\,, V_{11} = 0.000\,000\,001\,2\,,$$

$$V_{12} = 0.000\,000\,000\,3\,.$$

Hence,

$$V(0.5,0^{\circ}) = \sum_{l=0}^{12} V_l = 0.449\,955\,373\,9\,,\tag{15}$$

and

$$V(0.5, 180^{\circ}) = \sum_{l=0}^{12} V_l (-1)^l$$

= 0.341 021 359 6. (16)

These two values given to ten decimal digits agree with the eight figures of Trivedi and Steinborn.⁵ (We note that the table given by Hirschfelder and Weygandt¹⁹ is only accurate to six figures.)

If we are satisfied with eight-decimal digit accuracy, we

may run our program in a purely numerical mode and only use 1.4 sec central processing unit (CPU) time (on the Florida State University CDC-Cyber 170-760). For higher accuracy we must generate exact formulas for V_l , which for l = 12 requires 44 sec CPU time, but of course these formulas need only be generated once and then put on magnetic tape. To avoid intolerable cancellation errors at high l values, the formulas must be expanded by machine in a Taylor series.¹¹

III. CONCLUSION

Although in some respects our method is formally equivalent to the single-center method of Harris and Michels²⁰ and that of Barnett and Coulson²¹ (at least for *s* orbitals), it is our insistence on algebra that is decisive. The fact that the *C*-matrix elements are integers permits exact formulas to be developed. The extensive use of algebra gives us a penetrating insight into numerical calculations that allows us to proceed with confidence. Our method gives us the ability to establish benchmark values; the use of high-speed mass storage devices makes it possible to efficiently store and retrieve formula coefficients for further developments.

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