

Use of a scaling relationship for synchrotron radiation

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A single table is presented which can be used with a scaling relationship to predict the intensity of synchrotron radiation for any electron energy, wavelength, and observation angle. The accuracy of the scaling relationship and the table has been tested to a level of 1 part in 10^5 using current values for the fundamental constants.

The radiation emitted in electron storage rings and synchrotrons has become increasingly important for applications in many fields of physics, chemistry, biology, and material sciences.¹⁻³ The broad range of wavelengths (from the infrared down to the x-ray region), the high degree of polarization, and high intensity provide a number of favorable characteristics.

Recently, we have been concerned with using synchrotron radiation as an intensity source for the absolute calibration of spectrometers in the vacuum-ultraviolet region, cf. Ref. 4. As is well known, the absolute intensity of radiation emitted from an electron storage ring can be expressed in terms of the electron energy and current, the local bending radius of the electron orbit, the wavelength of the observed radiation, and the observation angle. The spectral characteristics can be conveniently parametrized in terms of a critical wavelength near the peak of the radiation distribution in the orbital plane.¹⁻⁷ In this article we present a method for determining the intensity using a scaling relationship which allows one to determine the intensity of radiation for any electron energy, wavelength, and observation angle by using just one single table which gives the wavelength in terms of the critical wavelength and using a universal function for the intensity.

Analytic expressions for the intensity of synchrotron radiation exist as functions of the electron energy, bending radius, observation angle, and wavelength⁸⁻¹⁰; however, the equations contain terms involving Bessel functions for which accurate numerical tables or computational algorithms are not widely known. Thus, it is difficult to readily determine the intensity of synchrotron radiation for different operating conditions. Furthermore, we know of no reports in which the basic equations have been tested numerically to an accuracy of better than $1:10^3$. Such tests are important for calibration purposes if there are no experimental checks of the theory. The theory itself contains a number of approximations which are not necessarily accurate at the 1 part in 10^3 level.

The angular distribution of the intensity of synchrotron radiation emitted from a beam of relativistic electrons of kinetic energy E in a circular orbit can be obtained from Schott's equation^{8,11} for discrete wavelengths based on harmonics of the orbital frequency

$$\frac{dN_E(\lambda, \Psi)}{d\Omega} = i \frac{\Delta\lambda}{\lambda} I_E(\lambda, \Psi), \quad (1)$$

where the intensity function is given as

$$I_E(\lambda, \Psi) = \frac{9e}{8\epsilon_0 hc} \frac{\gamma^6}{(\lambda/\lambda_c)^2} [\beta^4 J_\nu'^2(\beta\nu \cos\Psi) + \beta^2 \tan^2\Psi J_\nu^2(\beta\nu \cos\Psi)], \quad (2)$$

with

$$\beta = (1 - \gamma^{-2})^{1/2}, \quad (3)$$

$$\gamma = \frac{E}{mc^2}, \quad (4)$$

$$\nu = \frac{2\pi R}{\lambda}, \quad (5)$$

and

$$\lambda_c = \frac{4\pi R}{3\gamma^3}. \quad (6)$$

The derivative of J_ν is taken with respect to the complete argument $\beta\nu \cos\Psi$. The angular distribution $dN_E(\lambda, \Psi)/d\Omega$ is given in units of photons per second per steradian, where Ψ is the angle between the line of observation and its projection on the orbital plane which is tangent to the orbital ring, λ is the wavelength, $\Delta\lambda$ is the bandwidth, R is the bending radius of the electron orbit, and i is the current. In the derivation of Eq. (2) it is assumed that quantum effects can be neglected, the electrons emit incoherently, the diameter of the electron beam is infinitesimal, and the observation distance is large compared to the bending radius. In generating the table the values for the fundamental constants e , m , h , c , and ϵ_0 were taken from Cohen and Taylor.¹²

In general, if one wishes to calculate the intensity $dN_E(\lambda, \Psi)/d\Omega$ for a particular set of values of E , λ , and Ψ , one needs to evaluate the Bessel function $J_\nu(\nu z)$ and its derivative $J_\nu'(\nu z)$ for large

orders of ν . Since these functions change rapidly with z , our calculations were performed using 16 significant digits.

The widely known approximate formula for synchrotron radiation derived independently by Schwinger⁹ and Ivanenko and Sokolov¹⁰ in which asymptotic representations of the Bessel functions of high order are used is given as

$$I_E(\lambda, \Psi) \approx \frac{3e}{8\pi^2 \epsilon_0 hc} \frac{\gamma^2}{(\lambda/\lambda_c)^2} (1+X^2)^2 \times \left(K_{2/3}^2(\xi) + \frac{X^2}{1+X^2} K_{1/3}^2(\xi) \right), \quad (7)$$

where

$$X = \gamma \Psi, \quad (8)$$

$$\xi = \frac{1}{2(\lambda/\lambda_c)} (1+X^2)^{3/2}, \quad (9)$$

and $K_{1/3}$ and $K_{2/3}$ are modified Bessel functions of the second kind. Inspection of Eq. (7) shows that by using the dimensionless parameters $\bar{\lambda} = \lambda/\lambda_c$ and $\bar{\Psi} = \bar{E}\Psi$ one can write a scaling relationship

$$I_E(\lambda, \Psi) = \bar{E}^2 \bar{I}(\bar{\lambda}, \bar{\Psi}), \quad (10)$$

where \bar{E} is the energy in GeV and $\bar{\Psi}$ is in rad. Thus, if one table of the universal intensity function $\bar{I}(\bar{\lambda}, \bar{\Psi})$ is available for a large range of $\bar{\lambda}$ and $\bar{\Psi}$ one can scale the results and produce a table for the intensity function for any electron energy.

Equation (10) is similar in form to parametric equations derived by Green,⁷ some of which were reprinted in Ref. 1; however, our equation is more compact and incorporates the angular dependence of the intensity directly. The results of Green presented in Ref. 1 for $\Psi = 0$ only are in graphical form and are accurate to only 10–20%. Reference 7, containing more extensive graphs, is unpublished and not widely available. Our results are accurate to better than 1 part in 10^5 .

For practical calculations the critical wavelength λ_c can be expressed as

$$\lambda_c(\text{\AA}) = 5.5893 \frac{R}{\bar{E}^3}, \quad (11)$$

where R is in meters.

The polarization fraction of the intensity is given by

$$P = \frac{I_{\parallel} - I_{\perp}}{I_{\parallel} + I_{\perp}}, \quad (12)$$

where I_{\parallel} and I_{\perp} are the parallel and perpendicular components, respectively, with respect to the orbital plane containing the electron acceleration vector.

Table I lists the universal intensity function

$\bar{I}(\bar{\lambda}, \bar{\Psi})$ and the polarization fraction. Figure 1 shows $\bar{I}(\bar{\lambda}, \bar{\Psi})$ vs $\bar{\lambda}$ for several observation angles $\bar{\Psi}$.

Precise numerical analysis shows that the scaling relationship, Eq. (10), and $I_E(\lambda, \Psi)$ found from Eq. (7) agree with the more accurate formula Eq. (2) to within less than (a) $1:10^5$ for $\bar{E} \geq 0.5$ and $\bar{\lambda} \leq 10^4 \bar{E}^3$, and (b) $1:10^3$ for $\bar{E} \geq 0.05$ and $\bar{\lambda} \leq 10^6 \bar{E}^3$. Uncertainties in the fundamental constants e and m limit the accuracy of the calculations to five significant figures in general and to three significant figures for values of $\bar{\lambda}$ which are less than 0.1, see Table I.

A complete table for a particular storage ring can be obtained from the universal Table I by replacing the $\bar{\lambda}$ and $\bar{\Psi}$ scales by actual λ and Ψ scales. The wavelengths λ are found by multiplying the $\bar{\lambda}$ values in Table I by the λ_c for the storage ring considered. The angles Ψ are found by dividing the $\bar{\Psi}$ values in Table I by \bar{E} .

As an example of the use of Table I we consider calculating the intensity $dN_E/d\Omega$ for the Stanford Synchrotron Radiation Laboratory¹ (SSRL) operating at 2.5 GeV. To determine the intensity of 9 Å radiation emitted at $\Psi = 0.2$ mrad, we let

$$E = 2.5 \text{ GeV } (\bar{E} = 2.5),$$

$$R = 12.7 \text{ m (radius of curvature),}$$

$$i = 10 \text{ mA (current of electrons),}$$

$$\Delta\lambda/\lambda = 0.01 \text{ (1\% bandwidth averaging linearly).}$$

The intensity from Eqs. (1) and (10) is

$$\frac{dN_E}{d\Omega} = i \frac{\Delta\lambda}{\lambda} I_E(\lambda, \Psi),$$

where

$$I_E(\lambda, \Psi) = \bar{E}^2 \bar{I}(\bar{\lambda}, \bar{\Psi}).$$

We note that $\bar{\lambda} = \lambda/\lambda_c \approx 2.0$ since $\lambda_c = 4.54$ Å from Eq. (11), and

$$\bar{\Psi} = \bar{E}\Psi = 5 \times 10^{-4} \text{ rad.}$$

From Table I $\bar{I}(\bar{\lambda}, \bar{\Psi}) = 1.2082 \times 10^{22}$ photons/sec Å sr from which $dN_E/d\Omega \approx 7.55 \times 10^{18}$ photons/sec sr.

Practical calculations are concerned with the flux of radiation passing through an aperture. An expression for the flux through a rectangular aperture centered in the orbital plane and subtending an acceptance angle $\Delta\phi$ in the orbital plane and an opening angle $2\Psi_0$ perpendicular to the orbital plane is given by

$$N_E(\lambda, \Psi_0) = i \frac{\Delta\lambda}{\lambda} \Delta\phi F_E(\lambda, \Psi_0), \quad (13)$$

where

TABLE I. (Continued.)

λ	0.0	0	5.00	-4	1.00	-3	1.50	-3	2.00	-3	2.50	-3	3.00	-3	3.50	-3	4.00	-3	4.50	-3	5.00	-3	
2.0	0	1.7973	22	1.2082	22	4.1351	20	4.1141	16	7.923	8	1.870	-4										
		(1.00000)		(0.47531)		(0.16830)		(0.07578)		(0.04142)		(0.02573)											
3.0	0	1.5441	22	1.3142	22	1.6291	21	3.960	18	3.1239	13	1.2843	5	4.325	-8								
		(1.00000)		(0.51412)		(0.18988)		(0.08538)		(0.04612)		(0.02830)		(0.01897)									
4.0	0	1.3470	22	1.2744	22	2.9805	21	3.5739	19	5.7003	15	3.0916	9	1.4301	0	1.4014	-13						
		(1.00000)		(0.54490)		(0.20926)		(0.09450)		(0.05070)		(0.03084)		(0.01452)									
5.0	0	1.1977	22	1.1995	22	4.0821	21	1.2729	20	1.2326	17	1.2510	12	4.416	4	1.7836	-6	2.6981	-20				
		(1.00000)		(0.57027)		(0.22689)		(0.10320)		(0.05517)		(0.03334)		(0.01550)		(0.01145)							
6.0	0	1.0820	22	1.1207	22	4.8782	21	2.8729	20	9.254	17	6.618	13	4.2037	7	9.4027	-2	2.9241	-13				
		(1.00000)		(0.59173)		(0.24307)		(0.11152)		(0.05952)		(0.03580)		(0.01647)		(0.01211)							
8.0	0	9.1445	21	9.8131	21	5.7665	21	7.4984	20	1.0842	19	8.900	15	2.0941	11	7.0760	4	1.7130	-4	1.4785	-15		
		(1.00000)		(0.62649)		(0.27195)		(0.12716)		(0.06794)		(0.04063)		(0.02647)		(0.01840)							
1.0	1	7.9855	21	8.7077	21	6.0873	21	1.2698	21	4.5161	19	1.6025	17	3.2929	13	2.2585	8	2.9682	1	4.2809	-8	3.8748	-19
		(1.00000)		(0.65378)		(0.29718)		(0.14166)		(0.07600)		(0.04533)		(0.02939)		(0.02031)							
1.5	1	6.1943	21	6.8343	21	5.9396	21	2.3117	21	2.7228	20	6.7963	18	2.5104	16	9.5317	12	2.5748	8	3.4173	2	1.5370	-5
		(1.00000)		(0.70285)		(0.34904)		(0.17396)		(0.09480)		(0.05659)		(0.03647)		(0.02499)							
2.0	1	5.1512	21	5.6806	21	5.4462	21	2.8785	21	6.1548	20	4.0704	19	6.3716	17	1.7995	15	6.9677	11	2.8051	7	8.892	1
		(1.00000)		(0.73638)		(0.39001)		(0.20190)		(0.11196)		(0.06722)		(0.04329)		(0.02955)							
3.0	1	3.9576	21	4.3313	21	4.5177	21	3.2134	21	1.2420	21	2.1720	20	1.4394	19	3.0224	17	1.6771	15	2.0476	12	4.5751	8
		(1.00000)		(0.78056)		(0.45220)		(0.24866)		(0.14251)		(0.08688)		(0.05623)		(0.03834)							
4.0	1	3.2768	21	3.5580	21	3.8354	21	3.1430	21	1.6273	21	4.6193	20	6.2931	19	3.6011	18	7.5620	16	5.0837	14	9.5348	11
		(1.00000)		(0.80917)		(0.49832)		(0.28691)		(0.16917)		(0.10478)		(0.06835)		(0.04674)							
5.0	1	2.8286	21	3.0507	21	3.3378	21	2.9650	21	1.8250	21	6.9187	20	1.4514	20	1.5151	19	7.0688	17	1.3223	16	8.883	13
		(1.00000)		(0.82961)		(0.53454)		(0.31924)		(0.19287)		(0.12126)		(0.07977)		(0.05478)							
6.0	1	2.5075	21	2.6892	21	2.9629	21	2.7696	21	1.9096	21	8.7706	20	2.4521	20	3.8204	19	3.0344	18	1.1228	17	1.7674	15
		(1.00000)		(0.84514)		(0.56407)		(0.34719)		(0.21422)		(0.13654)		(0.09058)		(0.06250)							
8.0	1	2.0725	21	2.2034	21	2.4372	21	2.4161	21	1.9133	21	1.1147	21	4.4577	20	1.1450	20	1.7677	19	1.5341	18	6.9937	16
		(1.00000)		(0.86746)		(0.60993)		(0.39359)		(0.25147)		(0.16416)		(0.11063)		(0.07709)							
1.0	2	1.7872	21	1.8879	21	2.0858	21	2.1334	21	1.8303	21	1.2276	21	6.0781	20	2.1058	20	4.8423	19	7.0065	18	6.0450	17
		(1.00000)		(0.88297)		(0.64440)		(0.43106)		(0.28321)		(0.18864)		(0.12893)		(0.09070)							
1.5	2	1.3650	21	1.4207	21	1.5639	21	1.6543	21	1.5688	21	1.2638	21	8.2953	20	4.2757	20	1.6704	20	4.7749	19	9.6389	18
		(1.00000)		(0.90733)		(0.70339)		(0.50081)		(0.34641)		(0.23998)		(0.16885)		(0.12127)							

TABLE I. (Continued.)

λ	ψ	0.0	2.00	4.00	6.00	8.00	1.00	1.20	1.40	1.60	1.80	2.00
2.0	2	1.1272 21 (1.00000)	1.3502 21 (0.39466)	2.8575 20 (0.11012)	1.3767 18 (0.04026)	2.7058 13 (0.01849)	3.7519 5 (0.00994)	6.2091 -7 (0.00596)				
3.0	2	8.6050 20 (1.00000)	1.0536 21 (0.46548)	4.3675 20 (0.14684)	1.4083 19 (0.05619)	1.1252 16 (0.02617)	6.9875 10 (0.01410)	1.0384 3 (0.00844)	1.1263 -8 (0.00545)			
4.0	2	7.1043 20 (1.00000)	8.6896 20 (0.51630)	4.9838 20 (0.17825)	4.1443 19 (0.07094)	2.1092 17 (0.03355)	2.7709 13 (0.01817)	3.9013 7 (0.01087)	2.4228 -1 (0.00701)	2.7192 -12 (0.00479)		
5.0	2	6.1228 20 (1.00000)	7.4387 20 (0.55529)	5.1474 20 (0.20572)	7.5392 19 (0.08470)	1.1645 18 (0.04065)	9.5454 14 (0.02214)	2.0618 10 (0.01328)	5.7810 3 (0.00857)	1.0319 -5 (0.00584)	5.7370 -17 (0.00417)	
6.0	2	5.4223 20 (1.00000)	6.5347 20 (0.58652)	5.1003 20 (0.23014)	1.0876 20 (0.09759)	3.5194 18 (0.04750)	9.7746 15 (0.02603)	1.3035 12 (0.01565)	4.6337 6 (0.01010)	2.4279 -1 (0.00689)	1.0341 -10 (0.00491)	
8.0	2	4.4762 20 (1.00000)	5.3109 20 (0.63407)	4.7775 20 (0.27208)	1.6234 20 (0.12125)	1.3224 19 (0.06052)	1.6876 17 (0.03358)	2.1895 14 (0.02030)	1.8619 10 (0.01314)	6.6650 4 (0.00897)	6.4347 -3 (0.00639)	1.0720 -11 (0.00471)
1.0	3	3.8576 20 (1.00000)	4.5159 20 (0.66910)	4.3926 20 (0.30721)	1.9675 20 (0.14256)	2.7850 19 (0.07274)	8.8689 17 (0.04083)	4.5049 15 (0.02484)	2.5735 12 (0.01612)	1.1616 8 (0.01102)	2.9040 2 (0.00785)	2.8147 -5 (0.00579)
1.5	3	2.9440 20 (1.00000)	3.5614 20 (0.72771)	3.5774 20 (0.37567)	2.5971 20 (0.18822)	6.7705 19 (0.10046)	7.2867 18 (0.05787)	2.2823 17 (0.03572)	1.6520 15 (0.02337)	2.1880 12 (0.01604)	4.1899 8 (0.01146)	9.1545 3 (0.00846)
2.0	3	2.4303 20 (1.00000)	2.7281 20 (0.76491)	3.0056 20 (0.42672)	2.2948 20 (0.22603)	9.7285 19 (0.12497)	1.9219 19 (0.07356)	1.4937 18 (0.04601)	3.8472 16 (0.03034)	2.7596 14 (0.02093)	4.6244 11 (0.01499)	1.5168 8 (0.01108)
3.0	3	1.8547 20 (1.00000)	2.0370 20 (0.81088)	2.2940 20 (0.49992)	2.0617 20 (0.28641)	1.2502 20 (0.16699)	4.5217 19 (0.10177)	8.7081 18 (0.06510)	7.9747 17 (0.04356)	3.0966 16 (0.03033)	4.5397 14 (0.02186)	2.2351 12 (0.01622)
4.0	3	1.5310 20 (1.00000)	1.6587 20 (0.83900)	1.8736 20 (0.55129)	1.8135 20 (0.33354)	1.3094 20 (0.20227)	6.3920 19 (0.12667)	1.9353 19 (0.08254)	3.3395 18 (0.05594)	3.0157 17 (0.03928)	1.3073 16 (0.02848)	2.4933 14 (0.02122)
5.0	3	1.3194 20 (1.00000)	1.4159 20 (0.85836)	1.5958 20 (0.59009)	1.6079 20 (0.37199)	1.2853 20 (0.23268)	7.4989 19 (0.14900)	2.9757 19 (0.09864)	7.5051 18 (0.06759)	1.1242 18 (0.04784)	9.3386 16 (0.03487)	4.0131 15 (0.02608)
6.0	3	1.1684 20 (1.00000)	1.2450 20 (0.87267)	1.3978 20 (0.62079)	1.4426 20 (0.40430)	1.2317 20 (0.25941)	8.0826 19 (0.16927)	3.8382 19 (0.11360)	1.2463 19 (0.07861)	2.6153 18 (0.05604)	3.3508 17 (0.04105)	2.4746 16 (0.03082)
8.0	3	9.6447 19 (1.00000)	1.0176 20 (0.89271)	1.1328 20 (0.66693)	1.1991 20 (0.45623)	1.1075 20 (0.30466)	8.3975 19 (0.20496)	4.9838 19 (0.14072)	2.2170 19 (0.09904)	7.0867 18 (0.07149)	1.5602 18 (0.05286)	2.2666 17 (0.03996)
1.0	4	8.3115 19 (1.00000)	8.7103 19 (0.90626)	9.6208 19 (0.70046)	1.0303 20 (0.49673)	9.9434 19 (0.34199)	8.2043 19 (0.23568)	5.5581 19 (0.16485)	2.9838 19 (0.11768)	1.2270 19 (0.08586)	3.7365 18 (0.06400)	8.1446 17 (0.04868)
1.5	4	6.3428 19 (1.00000)	6.5788 19 (0.92698)	7.1576 19 (0.75573)	7.7310 19 (0.56907)	7.8593 19 (0.41338)	7.2210 19 (0.29771)	5.8153 19 (0.21568)	4.0017 19 (0.15827)	2.2988 19 (0.11800)	1.0778 19 (0.08945)	4.0322 18 (0.06894)

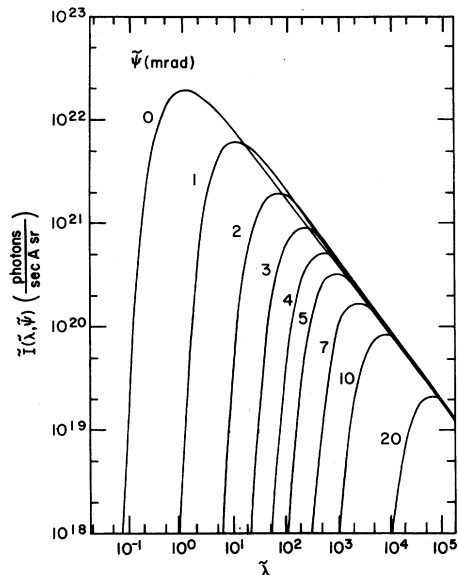


FIG. 1. Spectral distribution for several angles $\tilde{\Psi}$ of the universal intensity function $\tilde{I}(\tilde{\lambda}, \tilde{\Psi})$ as a function of $\tilde{\lambda} = \lambda/\lambda_c$, where λ_c is the critical wavelength. For an electron energy E , the intensity $dN_E(\lambda, \Psi)/d\Omega$ in photons/sec sr per bandwidth $(\Delta\lambda/\lambda)$ is found by multiplying $\tilde{I}(\tilde{\lambda}, \tilde{\Psi})$ by $i(\Delta\lambda/\lambda)\tilde{E}^2$, where \tilde{E} is the electron energy in GeV. The angle $\tilde{\Psi}$ is found by multiplying Ψ by \tilde{E} .

$$F_E(\lambda, \Psi_0) = \int_{\Psi_0}^{\Psi_0} I_E(\lambda, \Psi) \cos\Psi d\Psi. \quad (14)$$

A scaling relationship also holds to a good approximation for the flux function $F_E(\lambda, \Psi_0)$ for those cases where the radiation pattern is confined within a small angle to the orbital plane. This scaling law is linear in \tilde{E} , that is

$$F_E(\lambda, \Psi_0) = \tilde{E}\tilde{F}(\tilde{\lambda}, \tilde{\Psi}_0), \quad (15)$$

and agrees with Eq. (14) to a few parts in 10^5 for $\tilde{\lambda} \leq 5 \times 10^5 \tilde{E}^3$.

A table for the universal flux function $\tilde{F}(\tilde{\lambda}, \tilde{\Psi}_0)$ and a more extensive one for $\tilde{I}(\tilde{\lambda}, \tilde{\Psi})$ accurate to five significant figures will be published elsewhere.¹³

It should be mentioned that in practice a main source of error will be due to the limited accuracy to which the electron motion in a storage ring can be parametrized. The high accuracy of our tables, however, will assist in checking computational algorithms in addition to providing an overview of the spectral characteristics.

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