## Accurate measurement of relative oscillator strengths by phase-matched nonlinear optics

J. J. Wynne and R. Beigang\*

IBM Thomas J. Watson Research Center, P.O. Box 218, Yorktown Heights, New York 10598 (Received 5 December 1980)

A new method for accurately measuring relative oscillator strengths is presented. The method is based on measurements of the frequencies at which a nonlinear-optical sum-mixing process is phase matched. By applying this technique to Ca I, the relative strengths of the  $4s^{21}S_0$ -4snp  ${}^1P_1^0$  transitions for n = 10-13 are measured to an accuracy of  $\pm 2\%$ , more than six times better than previous measurements.

The two general types of atomic spectroscopic data used by "consumers"—atomic theorists, plasma physicists, astrophysicists, etc.—are (1) atomic line positions (including shapes) and (2) atomic transition strengths or oscillator strengths (f values). Advances in tunable dye laser technology and laser spectroscopic techniques now permit measurements<sup>1</sup> of line positions to accuracies of greater than 1 part in  $10^{11}$ . In contrast, f values are generally not known<sup>2</sup> to accuracies of better than 10%.

In this paper, we report on a method of accurately measuring *relative f* values using a nonlinear optical mixing technique. Key features of this method are (1) the relative f values are determined by frequency measurements, (2) a measurement of the light intensity is unnecessary, and (3) sample density and length, while affecting the measurement, need not be measured to get accurate f values. We have applied our technique to Ca I and measured the f values of the  $4s^{21}S_0-4snp$   ${}^{1}P_1^{0}$  transitions for n = 10, 11, 12, 13, and 14, relative to the f value of the main resonance line,  $4s^{21}S_0-4s4p$   ${}^{1}P_1^{0}$ . Accuracies of  $\pm 2\%$  are obtained for n = 10-13, a significant improvement over the best previous measurements<sup>3</sup> having accuracies of  $\geq \pm 12\%$ .

Our method is based on the dependence of nonlinear-optical sum mixing on phase matching.<sup>4</sup> The signal generated in a nonlinear-optical medium at a new frequency will, in general, maximize when the driven nonlinear polarization wave stays in phase with the wave it radiates; i.e., the process is phase matched. Beginning with Young et al.,<sup>5</sup> various studies of nonlinear-optical mixing in atomic vapors have been carried out where a buffer gas was added or the laser frequencies were tuned to achieve phase matching.<sup>6,7</sup> The aim of these studies was to maximize the signal. In contrast, Puell and Vidal,<sup>8</sup> using Xe to phase-match third-harmonic generation (THG) in Rb, determined the f value of the 5s-5ptransition in Rb1 relative to the refractive index contribution of Xe. They also pointed out that relative fvalues for an atomic species could be obtained by using tunable lasers to phase match a nonlinear-optical sum-mixing process. It is in this spirit that the present work has been undertaken.

Our experiments were carried out using two N<sub>2</sub>laser-pumped dye lasers, at frequencies  $v_1$  and  $v_2$ , respectively. Typical laser intensities for the experiment were  $\sim 100 \text{ kW/cm}^2$  at  $\nu_2$  and  $\sim 30 \text{ kW/cm}^2$  at  $v_1$ . Laser linewidths<sup>9</sup> were typically  $\sim 0.1$  cm<sup>-1</sup>. The copropagating laser beams passed through a heated pipe containing Ca and  $\sim 10$ -torr Ne buffer gas. The heated zone was  $\sim 20$  cm long, and a typical temperature for the measurements was 850 °C (measured at the center of the pipe), corresponding to a Ca pressure of  $\sim 2$  torr. In the Ca vapor the visible laser beams generated a uv signal at  $v_3 = 2v_1 + v_2$ , and this signal was recorded as  $v_2$  was tuned, with  $v_1$  held fixed. For calibration purposes and for resonant enhancement of the nonlinear susceptibility  $\chi^{(3)}(\nu_3 = 2\nu_1 + \nu_2), \nu_1$  was set at half the frequency of the  $4s^{2} S_0 - 4s 5s S_0$  two-photon transition of Ca1 at 33317.3 cm<sup>-1</sup>. This setting was ensured by blocking the beam at  $v_2$ , tuning  $v_1$ , and (1) looking for a maximum in the THG process due to resonance enhancement of  $\chi^{(3)}(3\nu_1)$ , and (2) simultaneously observing the maximum in the three-photon photoionization process. The uv was detected with a solar blind photomultiplier, and the ionization was monitored via space-charge-limited thermionic diode detection.<sup>10</sup> With  $v_1$  appropriately set, light at  $v_2$  was added. For the transitions studied,  $v_3$  took on values between ~47650 and 48570 cm<sup>-1</sup>. The  $3v_1$  background at 49 975.9 cm<sup>-1</sup> was blocked using a filter consisting of  $\sim$  1-cm thickness of 0.4-*M* KCl in H<sub>2</sub>O. During the data runs, a small fraction of the  $\nu_2$  light passed through an air-spaced, Fabry-Perot etalon to monitor accurately the tuning of  $\nu_2$ , accomplished either by rotating the laser's diffraction grating or by placing it in a vacuum chamber and slowly admitting N<sub>2</sub> or Freon 14 ( $CF_4$ ).<sup>11</sup> A typical data run is shown in Fig. 1. These data clearly show that the uv signal maximizes for  $v_3 \sim 4 \text{ cm}^{-1}$  below resonance with the  $4s^{21}S_0$ - $4s10p^{1}P_1^0$  transition at 47 662.1 cm<sup>-1</sup>, marked

<u>23</u>

2736



FIG. 1. Typical data with  $\nu_3$  in vicinity of  $4s^{21}S_0$ -4s 10p<sup>1</sup>P<sup>0</sup> transition at 47 662.1 cm<sup>-1</sup>. The insert on the left-hand side shows the relationship between the laser frequencies and the energy levels of Ca I.

by the maximum in the ionization signal.

In order to extract relative *f*-value information from this data, we take the well-known formulation<sup>8</sup> for the dependence of the signal intensity,  $I(v_3)$  on  $I(v_1)$ ,  $I(v_2)$ ,  $\chi^{(3)}(v_3 = 2v_1 + v_2)$ , and phase matching, namely,

$$I(\nu_3) \propto I^2(\nu_1) I(\nu_2) |\chi^{(3)}|^2 (\sin \Delta k L/2)^2 / (\Delta k)^2 \qquad (1)$$

Here,  $\Delta k$ , the phase mismatch, is defined as  $\Delta k = 2k_1 + k_2 - k_3$ , and the vapor sample is assumed to have constant density, N, over a length L. Equation (1) is valid in the limit of plane, collinear waves, negligible linear absorption, and small signals at  $v_3$ . Over a small tuning range of  $v_2$  so that  $I(v_2)$  is essentially constant,  $I(v_3)$  depends on  $v_2$  through the dependences of  $\chi^{(3)}$  and  $\Delta k$  on  $v_2$ . For  $v_3$  near  $v_{n'}$ we may write  $\chi^{(3)}(v_3 = 2v_1 + v_2) \propto 1/(v_{n'} - v_3)$ . The connection between  $I(v_3)$  and f values comes from the dependence of  $k_i$  on the refractive indices  $n(v_i)$ ,  $k_i = 2\pi n(v_i) v_{i}$ , and

$$n(\nu_{i}) = 1 + \frac{Ne^{2}}{2\pi mc^{2}} \left\{ \sum_{n} \frac{f_{n}}{\nu_{n}^{2} - \nu_{i}^{2}} + \int \frac{df}{d\nu_{c}} \frac{1}{\nu_{c}^{2} - \nu_{i}^{2}} d\nu_{c} \right\} .$$
 (2)

In Eq. (2), e is the electronic charge, m the electron mass, and c the speed of light. Equation (2) is valid when only the atomic ground state is populated and  $v_i$  is far enough removed from any resonance line frequency,  $v_n$ , so that linewidth factors are negligible. The sum is over all discrete transitions, and the integral is over all continuum transitions from the ground state,  $df/dv_c$  being the continuum density of oscillator strength. Note that even with an inhomogeneous vapor density N(z), since  $\chi^{(3)}$  and  $\Delta k$  are both proportional to N(z), it may be shown<sup>7</sup> that the above formulation is still valid provided that the product *NL* is replaced by  $\int N(z) dz$ , the integration being over the entire sample length. The proportionality between  $\Delta k$  and N(z) is contingent upon a negligible contribution to  $\Delta k$  from other gas atoms in the vapor, a condition readily satisfied with the use of ~10-torr Ne buffer gas.

In general, the uv signal strength as a function of  $v_3$  depends on all  $f_n$  and  $df/dv_c$ . However, by a careful choice of frequencies, certain selected transitions may be chosen to most strongly affect  $n(v_i)$ , and the uv generation measurement will give accurate information about these transitions. For Cal, with  $v_1$  and  $v_2 \sim 14\,000 - 17\,000 \text{ cm}^{-1}$ ,  $n(v_1)$  and  $n(v_2)$  are dominated by the main resonance line,  $4s^{21}S_0 - 4s4p^{-1}P_1^0$ , with  $\nu_R = 23652.3 \text{ cm}^{-1}$  and  $f_R = 1.75^{12}$  In order to measure the f value,  $f_{n'}$ , of a higher transition  $4s^{21}S_0$ -4sn'p <sup>1</sup> $P_1^0$ , relative to  $f_R$ , we tune  $\nu_3$  below and near to  $\nu_{\mu'}$ . When the contribution of this transition to  $n(\nu_3)$ , governed by the frequency offset,  $\nu_{n'}$ - $\nu_3$ , is sufficient, phase matching may be achieved, with  $I(\nu_3)$  showing a maximum as a function of  $v_3$ . The position of the maximum is determined primarily by the ratio  $f_{\mu'}/f_R$ . However, the resonant enhancement of  $\chi^{(3)}$  pulls the signal maximum from exact phase matching  $(\Delta k = 0)$  toward exact resonance  $(\nu_3 = \nu_{n'})$ .

The idealized situation just described, where only  $f_R$  and  $f_n$ , need be considered, is not completely realized. The accuracy of the determination of  $f_n/f_R$  from the  $I(\nu_3)$  vs  $\nu_3$  data depends on how accurately the other  $f_n$  and  $df/d\nu_c$  are known. For the present case, using values for  $f_n$  up to n = 17 in the discrete sum and  $df/d\nu_c$  values for higher frequencies in the continuum integral of Eq. (2),<sup>13</sup> we find that the calculated phase-matching peak position shifts closer to exact resonance by 7% (of its offset from resonance) as compared to neglecting all f except  $f_n$ , and  $f_R$ .

Examples of the uv signal, calculated from Eq. (1), for  $\nu_3$  near the  $4s^{21}S_0$ - $4s10p^{1}P_1^0$  transition, and with all relative f values taken from Geiger,<sup>13</sup> are shown in Fig. 2(a). One sees that the phase-matching peak occurs  $\sim 4 \text{ cm}^{-1}$  below resonance and that the peak narrows and becomes more symmetric, with the maximum moving farther from resonance as NL increases. In the limit  $NL \rightarrow \infty$ , the calculated curve would peak exactly at  $\Delta k = 0$ .

A nonlinear least-squares fit of the experimental data to Eq. (1) was carried out in order to extract relative f values for various n. The fitted parameters were  $f_n/f_R$  and NL. A typical fit is shown in Fig. 2(b) where the central maximum is seen to be fitted quite accurately. Data were taken at several temperatures so that N varied by as much as a factor  $\sim 10$ .



FIG. 2. uv signal as a function of  $\nu_3$  in the vicinity of the  $4s^2 - 4s 10p$  transition at 47 662.1 cm<sup>-1</sup>. (a) Calculated signal for three different values of *NL*. (b) Calculated signal (---) resulting from nonlinear least-squares fit of Eq. (1) to experimental data (---).

The values for  $f_n/f_R$  extracted from fits of data taken for a given n' showed no systematic dependence on temperature. Data at several values of Ne buffer gas pressure, from 10-200 torr, also showed no systematic dependence, as expected.

Our results are collected in Table I. The  $\pm 5\%$  uncertainty listed for n = 14 is the standard deviation of the mean for three data runs. However, the  $\pm 2\%$  uncertainty listed for n = 10-13 is larger than the statistical variation of several data runs for each *n*, reflecting the uncertainty of the other  $f_n$  and  $df/d\nu_c$  values used in extracting each  $f_n/f_R$  from the data fits.<sup>13</sup> For comparison purposes, Table I also lists the previously most accurate relative *f* values, derived

TABLE I. Cal: relative f values  $f(4s^{2}s_{0}-4snp^{1}P_{1}^{0})/f(4s^{2}s_{0}-4s4p^{1}P_{1}^{0}).$ 

| 11 | Phase-matched nonlinear optics | Hook <sup>a</sup><br>method    |
|----|--------------------------------|--------------------------------|
| 10 | $1.73 \times 10^{-3} \pm 2\%$  | $1.85 \times 10^{-3} + 12\%$   |
| 11 | $1.01 \times 10^{-3} \pm 2\%$  | $1.17 \times 10^{-3} \pm 12\%$ |
| 12 | $6.93 \times 10^{-4} \pm 2\%$  | $7.89 \times 10^{-4} \pm 12\%$ |
| 13 | $4.60 \times 10^{-4} \pm 2\%$  | $4.98 \times 10^{-4} \pm 15\%$ |
| 14 | $3.7 \times 10^{-4} \pm 5\%$   | $3.95 \times 10^{-4} \pm 23\%$ |

<sup>a</sup>Parkinson et al., Ref. 3.

from Hook measurements.<sup>3</sup> In each case, our measurement falls within the error range of the values given by Parkinson *et al.*,<sup>3</sup> but always toward the lower end of the range. At this time, we have no explanation for the  $\sim 10\%$  systematic difference between results of the two types of measurements.

In summary, we have demonstrated a new method of determining relative f values in atomic vapors, which, when applied to Cal, yields values at least six times more accurate than previous measurements. The precision is limited by inaccuracies in the experimental values of other resonance line transitions. By applying this technique to more members of the  $4s^2$ -4*snp* series, a self-consistent set of *f* values accurate to better than  $\pm 1\%$  should be obtainable. Such precison would be welcomed by solar physicists trying to determine the abundance of Ca in the photosphere of the sun and theoreticians calculating f values for Ca. A further advantage of this method is its capability of measuring f values for uv transitions using only visible and near-infrared laser light. Applications to transitions in atoms other than Ca1, including ions, and to excited-state-excited-state transitions are under consideration.

## ACKNOWLEDGMENTS

We wish to thank L. H. Manganaro for technical assistance, R. T. Hodgson for his original suggestion to make this type of measurement, G. C. Bjorklund for useful and informative discussions, and J. A. Armstrong and D. R. Grischkowsky for critical readings of this manuscript.

- \*Present address: Institut F. Atom-und Festkorperphysik, FU Berlin FB20, Boltzmannstr. 3, 1000 Berlin 33, West Germany.
- <sup>1</sup>J. C. Bergquist, R. L. Barger, and D. J. Glaze, in *Laser Spectroscopy IV*, edited by H. Walther and K. W. Rothe (Springer-Verlag, Berlin, 1979), p. 120.
- <sup>2</sup>W. L. Wiese and G. A. Martin, in CRC Handbook of Chemistry and Physics, edited by R. C. Weast (CRC Press, Inc., Boca Raton, Florida, 1980), p. E-349.
- <sup>3</sup>W. H. Parkinson, E. M. Reeves and F. S. Tomkins, J. Phys. B <u>9</u>, 157 (1976).
- <sup>4</sup>J. A. Giordmaine, Phys. Rev. Lett. <u>8</u>, 19 (1962); P. D. Maker, R. W. Terhune, M. Nisenoff, and C. M. Savage, *ibid.* <u>8</u>, 21 (1962).
- <sup>5</sup>J. F. Young, G. C. Bjorklund, A. H. Kung, R. B. Miles, and S. E. Harris, Phys. Rev. Lett. <u>27</u>, 1551 (1971).
- <sup>6</sup>P. P. Sorokin, J. J. Wynne, and J. R. Lankard, Appl. Phys. Lett. <u>22</u>, 342 (1973); R. T. Hodgson, P. P. Sorokin, and

- J. J. Wynne, Phys. Rev. Lett. <u>32</u>, 343 (1974).
- <sup>7</sup>G. C. Bjorklund, J. E. Bjorkholm, P. F. Liao, and R. H. Storz, Appl. Phys. Lett. <u>29</u>, 729 (1976).
- <sup>8</sup>H. Puell and C. R. Vidal, Opt. Commun. <u>19</u>, 279 (1976).
- <sup>9</sup>Explicit values for frequencies  $\nu$  are given in units of cm<sup>-1</sup>. <sup>10</sup>K. H. Kingdon, Phys. Rev. <u>21</u>, 408 (1923); J. A.
- Armstrong, P. Esherick, and J. J. Wynne, Phys. Rev. A <u>15</u>, 180 (1977).
- <sup>11</sup>**R**. Wallenstein and T. W. Hänsch, Appl. Opt. <u>13</u>, 1625 (1974).
- <sup>12</sup>W. H. Smith and H. S. Liszt, J. Opt. Soc. Am. <u>61</u>, 938 (1971), measured  $f_R$  to an accuracy of  $\pm 10\%$ .
- <sup>13</sup>J. Geiger, J. Phys. B <u>12</u>, 2277 (1979), from which we take  $f_n$  values from column 4 of Table 3, and  $df/d\nu_c$  from Eq. (11) and Table I, setting  $I_0$  to match the discrete f values to  $df/d\nu_c$  through Eq. (12). We estimate an overall uncertainty of  $\pm 20\%$  for these f values.