# Improved equation of state for the classical one-component plasma 

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#### Abstract

We compute the internal energy of the classical one-component plasma, for values of the Coulomb coupling parameter $\Gamma$ between 1 and 300 , by using converged Monte Carlo chains and a more accurate potential approximation than that used by Hansen. The liquid data are fitted to a simple, very accurate fourparameter formula from which the Helmholtz free energy is derived. The solid data are likewise fitted to a one-parameter formula. The intersection of the two free-energy curves gives an estimate of the fluid-solid transition at $\Gamma=168 \pm 4$.


## I. INTRODUCTION

The classical one-component plasma (OCP) is an idealized system of ions immersed in a uniform sea of electrons such that the whole system is electrically neutral. This system is characterized by the Coulomb coupling parameter $\Gamma$, where $\Gamma=(Z e)^{2} / \overline{\mathrm{r}} k T$ with the ion sphere radius $\bar{r}=(4 \pi N / 3 V)^{-1 / 3}, N$ the number of ions and $V$ is the volume of the system.

Following the pioneering Monte Carlo work of Brush, Sahlin, and Teller ${ }^{1}$ (hereafter BST) on the OCP, Hansen ${ }^{2}$ dramatically improved the accuracy of the equation of state by improving the approximation of the interparticle potential and by using long Monte Carlo runs. However, DeWitt ${ }^{3}$ has questioned the accuracy of Hansen's data for large values of the Coulomb coupling parameter ( $40<\Gamma<160$ ). DeWitt thought that the reduced precision for large $\Gamma$ may have been due to Monte Carlo noise or to possible systematic errors. In this paper we report the results of calculations similar to but more accurate than the results of Hansen and Brush, Sahlin and Teller. Systematic deviations from Hansen's results are found and our estimation for the fluid-solid transition is $\Gamma=$ $168 \pm 4$.

For the potential approximation used here the error in evaluating the Madelung sum is of order 1 part in $10^{6}$. This compared with an error of order 1 part in $10^{3}$ for Hansen's approximation. We also extend the range of the pair distribution functions by including in the averages, particles in the volume ( $V=L^{3}$ ) inside the basic periodic cube but outside the inscribed sphere with radius $L / 2$.

In Sec. II we discuss the various approximations to the Ewald potential (Hansen, ${ }^{1}$ and DeWitt and Hubbard ${ }^{4}$ ). It is shown that our approximation to the potential is at least two orders of magnitude
more accurate than Hansen's. The fit to the potential which we used (cubic harmonics specialized for a vector computer) is described in detail in Appendix A. The data and a simple very accurate four-parameter fit to the liquid data, and a oneparameter fit to the solid data are discussed in Sec. III. Appendix B gives analytic formulas for normalizing the extended pair distribution function.

## II. INTERPARTICLE POTENTIAL APPROXIMATION

Following BST, the total internal energy of a system $U$ may be written

$$
\begin{equation*}
\frac{U}{N k T}=\frac{U_{0}}{N k T}+\frac{1}{2} \sum_{i \neq j=1}^{N} \frac{\psi}{N k T}\left(\left|\overrightarrow{\mathrm{r}}_{i}-\overrightarrow{\mathrm{r}}_{j}\right|\right), \tag{1}
\end{equation*}
$$

where

$$
\begin{equation*}
\frac{\psi}{N k T}=\frac{\Gamma}{N L}\left[U_{1}(\gamma)+U_{2}(\overrightarrow{\mathrm{r}})\right], \tag{2}
\end{equation*}
$$

and

$$
\begin{align*}
U_{1}(r)= & (1 / r) \operatorname{erfc}\left(\pi^{1 / 2} r\right)-1,  \tag{3}\\
U_{2}(\overrightarrow{\mathrm{r}})= & \sum_{\mathrm{n}}^{\prime} \frac{1}{|\overrightarrow{\mathrm{n}}-\overrightarrow{\mathrm{r}}|} \operatorname{erfc}\left(\pi^{1 / 2}|\overrightarrow{\mathrm{n}}-\overrightarrow{\mathrm{r}}|\right) \\
& +\sum_{\mathrm{n}}^{\prime} \frac{1}{\pi n^{2}} \exp \left(-\pi n^{2}\right)\{\exp [i(2 \pi \overrightarrow{\mathrm{n}} \cdot \overrightarrow{\mathrm{r}})]\} . \tag{4}
\end{align*}
$$

Here the interparticle separation $\overrightarrow{\mathrm{r}}=\overrightarrow{\mathrm{r}}_{i}-\overrightarrow{\mathrm{r}}_{j}$ is in units of $L$ where $L=(4 \pi N / 3)^{1 / 3}$ in units of the ion sphere radius $\bar{r}$ with $N$ the number of particles.

In these units the minimum and maximum values for any component of $r$ are 0 and 1 , respectively, inside the basic periodic cube. $\overrightarrow{\mathrm{n}}$ is a vector with integer components and the prime on the summations indicates that the zero $\overrightarrow{\mathrm{n}}$ vector ( $0,0,0$ ) is to be excluded. In terms of the above,

$$
\frac{U_{0}}{N k T}=\frac{\Gamma}{2 L}\left[U_{2}(0)-3\right]=-\frac{\Gamma}{2 L} E_{m},
$$

TABLE I. Errors in potential approximations (all errors are relative).

| Error description | Polynomial <br> used here | DeWitt and <br> Hubbard (Ref. 4) | Hansen (Ref. 2) |
| :---: | :---: | :---: | :---: |
| rms <br> average | $6.5 \times 10^{-8}$ | $4.2 \times 10^{-7}$ | $1.4 \times 10^{-5}$ |
| (indicates amount <br> of cancellation for <br> a uniform distribution) <br> maximum | $2.1 \times 10^{-4}$ | $-2.0 \times 10^{-5}$ | $-4.0 \times 10^{-4}$ |

or $E_{m}=3-U_{2}(0)$ where $E_{m}=2.837297479$ is the Madelung energy of a simple cubic crystal with lattice spacing $L$. Since Eq. (4) is difficult to compute rapidly it is customary to approximate it in some other form. BST used a Taylor expansion, Hansen ${ }^{2}$ used what he called an "optimized Kubic harmonic" expansion, and DeWitt and Hubbard ${ }^{4}$ used a 3D table look-up and interpolation. Since we are using a vector computer which evaluates polynomials rapidly we shall use ordinary cubic harmonics such as those used by Slater and DeCicco. ${ }^{5}$ We describe this approach in detail in Appendix A.

We now show that our potential approximation is several orders of magnitude more accurate than previous calculations. To evaluate the errors in the various potential approximations we computed the potential at 64000 random points to within a few parts in $10^{10}$ using Eq. (4). This value was then compared with the values of Hansen, DeWitt, and Hubbard and our polynomial values. The results are shown in Table I for 64000 random points. As noted by Hansen, ordinary cubic harmonics converge poorly for $r>L / 2$ but this error may be decreased by taking enough terms. The cubic harmonic approximation places the largest relative error where the potential is the smallest. This is in contrast with Hansen's optimized Kubic harmonics which place the largest error where the potential contributes much more to the total sum. The errors in the DeWitt and Hubbard scheme are distributed more uniformly since it is a table look-up and interpolation procedure. Their method could be made more accurate but table look-up is a relatively slow procedure on a vector machine. To estimate the systematic errors of the various potentials we placed the particles in a bcc ( 128 ions) or fcc ( 108 ions) lattice and evaluated the static internal energy. The resulting static energy may be compared directly with the Madelung energy of that lattice. The results show that for the potential used here the systematic errors are of order 1 part in $10^{6}$ whereas Hansen's potential has systematic errors of order 1 part in $10^{3}$. With the polynomial potential the errors de-
crease with increasing numbers of ions, for example, bcc lattices with 432 and 686 ions have systematic errors of 4 parts and 1 part in $10^{7}$, respectively.

By rearranging the Ewald sum we were able to write a computer code that was about three times faster than the polynomial code, but it was not used because it was one-half order of magnitude less accurate.

## III. RESULTS AND DISCUSSION

Table II gives the internal energies ( $U / N k T$ ), corrected for center-of-mass motion, for our calculations on the one-component plasma. The number of configurations in the Monte Carlo (MC) chains varied from $10^{6}$ (with $10^{5}$ previously discarded) for the lower $\Gamma$ on up to $8.6 \times 10^{6}$ (with $0.2 \times 10^{6}$ to $1.5 \times 10^{6}$ previously discarded) for the higher $\Gamma$. Most of the calculations were done using 128 particles; a few $54,108,250$, and 432 particle calculations were also done. The "lattice" after some of the entries indicates that a bcc or fcc lattice start was used. The statistical errors $\sigma$ are the standard deviations of the means. These errors were computed after ensuring that there was negligible slope in the internal energy data.

Comparison of the data for the 128 ions and 250 ions shows a dependence on the numbers of ions ( $N$ dependence). As expected this $N$ dependence is not nearly as strong when comparing the 250 data with the 432 data. Apparently a few hundred particles are necessary for the $N$-dependence errors to be negligible. Consequently the results in this paper which are all based upon the $N=128$ runs are subject to a small systematic error. For $\Gamma=$ 180 the random start was seen to solidify after $1.6 \times 10^{6}$ configurations. Figure 1 shows the pair distribution functions for the liquid (dashed curve) and solid (solid curve). All lattice starts with $\Gamma<160$ melted and all random starts with $\Gamma \geqslant 170$ solidified. For calculation of structure factors, we give various pair distribution functions in Table III.

The systematic deviations of Hansen's ${ }^{2}$ data from

TABLE II. Internal energies. $U / N k T$ is the excess internal energy per ion divided by $k T$. We give our results and those of Hansen (Ref. 2) and Pollock and Hansen (Ref. 10).

| $\Gamma$ | $U / N k T$ | Error $( \pm \sigma)$ | Number of ions $N$ | Type of start | $U / N k T$ (Refs. 2 and 10) |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 1.0 | -0.573 | 0.000 | 108 | $1 i q{ }^{\text {a }}$ |  |
| 1.0 | -0.573 | 0.000 | 128 | liq | -0.580 |
| 2.0 | -1.320 | 0.001 | 128 | liq | -1.318 |
| 3.0 | -2.112 | 0.001 | 128 | liq | -2.111 |
| 4.0 | -2.927 | 0.001 | 128 | liq | -2.926 |
| 6.0 | -4.592 | 0.001 | 108 | liq |  |
| 6.0 | -4.593 | 0.001 | 128 | liq | -4.590 |
| 10.0 | -7.993 | 0.001 | 108 | liq |  |
| 10.0 | -7.992 | 0.002 | 128 | liq | -7.996 |
| 15.0 | -12.310 | 0.002 | 108 | liq |  |
| 15.0 | -12.309 | 0.003 | 128 | liq | -12.313 |
| 20.0 | -16.668 | 0.004 | 128 | liq | -16.667 |
| 30.0 | -25.426 | 0.004 | 108 | liq |  |
| 30.0 | -25.431 | 0.004 | 128 | liq | -25.429 |
| 40.0 | -34.251 | 0.004 | 108 | liq |  |
| 40.0 | -34.248 | 0.002 | 128 | liq | -34.232 |
| 60.0 | -51.941 | 0.006 | 54 | liq |  |
| 60.0 | -51.961 | 0.006 | 108 | liq |  |
| 60.0 | -51.961 | 0.007 | 128 | liq | -51.936 |
| 80.0 | -69.714 | 0.006 | 54 | liq |  |
| 80.0 | -69.727 | 0.005 | 108 | liq |  |
| 80.0 | -69.715 | 0.005 | 128 | liq | -69.690 |
| 100.0 | -87.511 | 0.007 | 54 | liq |  |
| 100.0 | -87.534 | 0.004 | 108 | liq |  |
| 100.0 | -87.500 | 0.005 | 128 | liq | -87.480 |
| 100.0 | -87.498 | 0.005 | 250 | liq |  |
| 110.0 | -96.411 | 0.005 | 128 | liq | -96.360 |
| 125.0 | -109.825 | 0.007 | 108 | liq |  |
| 125.0 | -109.780 | 0.004 | 128 | liq | -109.732 |
| 125.0 | -109.793 | 0.005 | 250 | liq |  |
| 125.0 | -109.799 | 0.004 | 432 | liq |  |
| 140.0 | -123.239 | 0.010 | 108 | liq |  |
| 140.0 | -123.160 | 0.006 | 128 | liq | -123.095 |
| 140.0 | -123.165 | 0.007 | 250 | liq |  |
| 150.0 | -132.078 | 0.007 | 128 | liq |  |
| 160.0 | -141.098 | 0.009 | 108 | liq |  |
| 160.0 | -141.000 | 0.006 | 128 | liq | -140.890 |
| 160.0 | -141.051 | 0.007 | 432 | liq |  |
| 160.0 | -141.690 | 0.010 | 108 | lat ${ }^{\text {b }}$ |  |
| 160.0 | -141.729 | 0.006 | 128 | lat | -141.717 |
| 160.0 | -141.716 | 0.009 | 250 | lat |  |
| 170.0 | -150.703 | 0.005 | 128 | lat | -150.696 |
| 180.0 | -159.654 | 0.007 | 108 | lat |  |
| 180.0 | -159.675 | 0.004 | 128 | lat | -159.662 |
| 180.0 | -159.669 | 0.005 | 250 | lat |  |
| 180.0 | -159.667 | 0.004 | 432 | lat |  |
| 200.0 | -177.574 | 0.005 | 108 | lat |  |
| 200.0 | -177.619 | 0.004 | 128 | lat | -177.604 |
| 220.0 | -195.551 | 0.006 | 128 | lat | -195.538 |
| 240.0 | -213.469 | 0.003 | 128 | lat | -213.470 |
| 300.0 | -267.205 | 0.008 | 108 | lat |  |
| 300.0 | -266.675 | 0.006 | 125 | lat |  |
| 300.0 | -267.243 | 0.004 | 128 | lat | -267.242 |

[^0]

FIG. 1. Pair distribution functions for the liquid (dashed curve) and the solid (solid curve) for $\Gamma=180$.
our data can be seen more clearly in Fig. 2 where we have plotted the thermal fraction of the internal energy, $\left(U-U_{0}\right) / N k T$, for our data (solid curve) and Hansen's ${ }^{2}$ data, and the data of Pollock and Hansen ${ }^{10}$ (dashed curve). Since Hansen's statistical errors are larger than ours by a factor of order 2 , the differences can mostly be attributed to the fact that our potential is two to three orders of magnitude more accurate than Hansen's.

The liquid $U / N k T$ data ( $1 \leqslant \Gamma \leqslant 160$ ) for 128 -particle runs were fitted to the form

$$
\begin{equation*}
U / N k T=a \Gamma+b \Gamma^{1 / 4}+c \Gamma^{-1 / 4}+d, \tag{5}
\end{equation*}
$$

by minimizing the sum of the squares of the relative errors using $a, b, c$, and $d$ as parameters. The fitted $U / N k T$ are shown in Table IV along with

TABLE III. Pair distributions as a function of $\Gamma$ and radius (ion sphere units).

| $\Gamma$ | 80 | 110 | 125 | 140 | 150 | 160 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Radius |  |  |  |  |  |  |
| 0.0406 | 0.00000 | 0.00000 | 0.00000 | 0.00000 | 0.00000 | 0.00000 |
| to |  |  |  |  |  |  |
| 0.9343 | 0.00000 | 0.00000 | 0.00000 | 0.00000 | 0.00000 | 0.00000 |
| 1.0155 | 0.00009 | 0.00000 | 0.00000 | 0.00000 | 0.00000 | 0.00000 |
| 1.0967 | 0.00122 | 0.00007 | 0.00000 | 0.00000 | 0.00000 | 0.00000 |
| 1.1780 | 0.01407 | 0.00276 | 0.00122 | 0.00048 | 0.00029 | 0.00017 |
| 1.2592 | 0.08493 | 0.02962 | 0.01589 | 0.00957 | 0.00645 | 0.00463 |
| 1.3404 | 0.30161 | 0.16145 | 0.11679 | 0.08586 | 0.06637 | 0.05393 |
| 1.4217 | 0.71033 | 0.52691 | 0.44756 | 0.37714 | 0.32813 | 0.29571 |
| 1.5029 | 1.23254 | 1.11559 | 1.05375 | 0.99015 | 0.92628 | 0.89207 |
| 1.5842 | 1.67427 | 1.71958 | 1.71491 | 1.71490 | 1.70176 | 1.68981 |
| 1.6654 | 1.89391 | 2.06666 | 2.13504 | 2.19421 | 2.24500 | 2.26708 |
| 1.7466 | 1.88092 | 2.08405 | 2.18090 | 2.25957 | 2.33161 | 2.37848 |
| 1.8279 | 1.70898 | 1.87124 | 1.94683 | 2.01267 | 2.07637 | 2.11711 |
| 1.9091 | 1.47963 | 1.56849 | 1.61510 | 1.65084 | 1.68386 | 1.70876 |
| 1.9904 | 1.25013 | 1.27867 | 1.29415 | 1.30714 | 1.31253 | 1.32301 |
| 2.0716 | 1.06053 | 1.04632 | 1.04028 | 1.03500 | 1.02550 | 1.02030 |
| 2.1528 | 0.91731 | 0.87993 | 0.85878 | 0.84111 | 0.82694 | 0.81629 |
| 2.2341 | 0.81472 | 0.76561 | 0.73594 | 0.71828 | 0.69934 | 0.68262 |
| 2.3153 | 0.75079 | 0.69499 | 0.66714 | 0.64750 | 0.62569 | 0.60664 |
| 2.3966 | 0.72072 | 0.66416 | 0.63859 | 0.61436 | 0.59326 | 0.57968 |
| 2.4778 | 0.71586 | 0.66407 | 0.63973 | 0.61745 | 0.59613 | 0.58618 |
| 2.5590 | 0.73845 | 0.68853 | 0.66643 | 0.64785 | 0.62929 | 0.62430 |
| 2.6403 | 0.77804 | 0.73426 | 0.71705 | 0.70168 | 0.69128 | 0.68719 |
| 2.7215 | 0.83578 | 0.79972 | 0.78997 | 0.77747 | 0.77272 | 0.76965 |
| 2.8028 | 0.90310 | 0.88072 | 0.87504 | 0.87249 | 0.86949 | 0.86840 |
| 2.8840 | 0.97407 | 0.96790 | 0.97042 | 0.97282 | 0.97732 | 0.97820 |
| 2.9652 | 1.04191 | 1.05575 | 1.06519 | 1.07133 | 1.08096 | 1.08186 |
| 3.0465 | 1.09864 | 1.13105 | 1.14571 | 1.15708 | 1.16836 | 1.17405 |
| 3.1277 | 1.14003 | 1.18322 | 1.20401 | 1.21622 | 1.23281 | 1.24284 |
| 3.2090 | 1.15827 | 1.20902 | 1.23033 | 1.24529 | 1.26377 | 1.27603 |
| 3.2902 | 1.15440 | 1.20539 | 1.22477 | 1.24232 | 1.25771 | 1.27080 |
| 3.3714 | 1.13313 | 1.17749 | 1.19225 | 1.20771 | 1.22021 | 1.22803 |
| 3.4527 | 1.10064 | 1.12940 | 1.13957 | 1.15272 | 1.16164 | 1.16177 |
| 3.5339 | 1.06233 | 1.07523 | 1.07702 | 1.08672 | 1.08819 | 1.08720 |
| 3.6151 | 1.02069 | 1.01889 | 1.01634 | 1.01822 | 1.01324 | 1.01043 |
| 3.6964 | 0.98498 | 0.96926 | 0.96196 | 0.95449 | 0.94743 | 0.94320 |
| 3.7776 | 0.95336 | 0.92939 | 0.91811 | 0.90699 | 0.89554 | 0.88775 |
| 3.8589 | 0.93300 | 0.90315 | 0.89061 | 0.87442 | 0.86122 | 0.85172 |
|  |  |  |  |  |  |  |

TABLE III. (Continued.)

| $\Gamma$ | 80 | 110 | 125 | 140 | 150 | 160 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 3.9401 | 0.92172 | 0.88849 | 0.87514 | 0.85987 | 0.84550 | 0.83906 |
| 4.0213 | 0.92156 | 0.88860 | 0.87510 | 0.86095 | 0.84777 | 0.84254 |
| 4.1026 | 0.92909 | 0.90225 | 0.89113 | 0.88078 | 0.86819 | 0.86544 |
| 4.1838 | 0.94701 | 0.93150 | 0.92284 | 0.91830 | 0.90834 | 0.90729 |
| 4.2651 | 0.97148 | 0.96287 | 0.96127 | 0.96276 | 0.95693 | 0.96139 |
| 4.3463 | 0.99444 | 0.99560 | 1.00124 | 1.00638 | 1.00868 | 1.01579 |
| 4.4275 | 1.01348 | 1.02535 | 1.03683 | $1.04162^{-}$ | 1.05201 | 1.05945 |
| 4.5088 | 1.02866 | 1.04716 | 1.06153 | 1.06801 | 1.08286 | 1.08969 |
| 4.5900 | 1.03864 | 1.06118 | 1.07460 | 1.08154 | 1.09998 | 1.10124 |
| 4.6713 | 1.04234 | 1.06787 | 1.07664 | 1.08395 | 1.10209 | 1.09711 |
| 4.7525 | 1.03951 | 1.06632 | 1.07161 | 1.07730 | 1.09063 | 1.08581 |
| 4.8337 | 1.03471 | 1.05518 | 1.05898 | 1.06395 | 1.07395 | 1.06993 |
| 4.9150 | 1.02951 | 1.04228 | 1.03882 | 1.04756 | 1.05094 | 1.04596 |
| 4.9962 | 1.02024 | 1.02684 | 1.01902 | 1.02795 | 1.02811 | 1.02492 |
| 5.0775 | 1.00637 | 1.00897 | 0.99966 | 1.00804 | 1.00118 | 1.00193 |
| 5.1587 | 0.99649 | 0.99044 | 0.98240 | 0.98823 | 0.97869 | 0.98028 |
| 5.2399 | 0.98668 | 0.97348 | 0.97150 | 0.97131 | 0.96004 | 0.96734 |
| 5.3212 | 0.98057 | 0.96139 | 0.96258 | 0.95833 | 0.94687 | 0.95261 |
| 5.4024 | 0.97663 | 0.95364 | 0.95760 | 0.94657 | 0.93801 | 0.94744 |
| 5.4837 | 0.97474 | 0.95245 | 0.95565 | 0.94511 | 0.93571 | 0.94412 |
| 5.5649 | 0.97538 | 0.95626 | 0.96067 | 0.94579 | 0.93638 | 0.95007 |
| 5.6461 | 0.97765 | 0.96684 | 0.97064 | 0.95296 | 0.94712 | 0.96140 |
| 5.7274 | 0.98490 | 0.98221 | 0.98581 | 0.96798 | 0.96790 | 0.96850 |
| 5.8086 | 0.99247 | 1.00121 | 1.00371 | 0.98883 | 0.99213 | 0.98187 |
| 5.8898 | 1.00704 | 1.02113 | 1.01820 | 1.00836 | 1.01700 | 0.99651 |
| 5.9711 | 1.01656 | 1.03188 | 1.03084 | 1.03135 | 1.04001 | 1.02203 |
| 6.0523 | 1.02249 | 1.04542 | 1.03697 | 1.05050 | 1.06827 | 1.04014 |
| 6.1336 | 1.03081 | 1.04675 | 1.03569 | 1.06142 | 1.08433 | 1.05042 |
| 6.2148 | 1.02964 | 1.03956 | 1.02464 | 1.06399 | 1.08333 | 1.05093 |
| 6.2960 | 1.02101 | 1.01902 | 1.01457 | 1.04607 | 1.07168 | 1.04902 |
| 6.3773 | 1.00736 | 1.01406 | 1.00906 | 1.02164 | 1.04800 | 1.03531 |
| 6.4585 | $0.99548{ }^{\circ}$ | 0.99930 | 1.00686 | 0.98763 | 1.00706 | 1.01399 |
| 6.5398 | 0.98887 | 0.99361 | 0.99349 | 0.96939 | 0.96762 | 0.99714 |
| 6.6210 | 0.97859 | 0.98444 | 0.98951 | 0.95913 | 0.92446 | 0.98583 |



FIG. 2. Thermal fractions of internal energy (compared with a bcc lattice $U_{0}=-0.895929256 \Gamma$ ). Solidcurve data from this paper, dashed-curve data from Hansen (Ref. 2) and Pollock and Hansen (Ref. 10).

TABLE IV. Fitted liquid data for $1 \leqslant \Gamma \leqslant 160 . U / N k T$ $=a \Gamma+b \Gamma^{1 / 4}+c \Gamma^{-1 / 4}+d$, where $a=-0.89752, b=0.94544$, $c=0.17954$, and $d=-0.80049$.

|  | $U / N k T$ |  |  |  |
| ---: | ---: | ---: | ---: | ---: |
| $\Gamma$ | MC | $U / N k T$ <br> fitted | Relative <br> error | Absolute <br> error |
| 1.0 | -0.573 | -0.573 | 0.00003 | 0.000 |
| 2.0 | -1.320 | -1.320 | -0.00012 | 0.000 |
| 3.0 | -2.112 | -2.112 | -0.00014 | 0.000 |
| 4.0 | -2.927 | -2.927 | 0.00022 | -0.001 |
| 6.0 | -4.593 | -4.591 | 0.00036 | -0.002 |
| 10.0 | -7.992 | -7.994 | -0.00013 | 0.001 |
| 15.0 | -12.309 | -12.311 | -0.00022 | 0.003 |
| 20.0 | -16.668 | -16.667 | 0.00007 | -0.001 |
| 30.0 | -25.431 | -25.437 | -0.00021 | 0.005 |
| 40.0 | -34.248 | -34.252 | -0.00013 | 0.005 |
| 60.0 | -51.961 | -51.956 | 0.00010 | -0.005 |
| 80.0 | -69.715 | -69.715 | 0.00001 | 0.000 |
| 100.0 | -87.500 | -87.506 | -0.00007 | 0.006 |
| 110.0 | -96.411 | -96.410 | 0.00001 | -0.001 |
| 125.0 | -109.780 | -109.775 | 0.00004 | -0.005 |
| 140.0 | -123.157 | -123.149 | 0.00006 | -0.008 |
| 150.0 | -132.078 | -132.068 | 0.00007 | -0.010 |
| 160.0 | -141.000 | -140.991 | 0.00007 | -0.009 |

TABLE V. $U / N k T$ for $\Gamma \leqslant 1$.

| Calculation <br> $\Gamma$ | Monte Carlo (MC) <br> calculated here | HNC <br> (Ref. 9) | BSC <br> (Ref. 8) | Abe <br> (Refs. 6 and 2) |
| :---: | :---: | :---: | :---: | :---: |
| 0.1 |  | -0.02580 | -0.0258 | $-0.02590^{\mathrm{b}}$ |
| 0.2 | -0.1225 | $-0.06903^{\text {b }}$ | -0.0688 | -0.0718 |
| 0.3 | -0.1784 | $-0.1195^{\mathrm{b}}$ |  |  |
| 0.4 | -0.2383 | $-0.1773^{\mathrm{b}}$ |  |  |
| 0.5 | $-0.3018^{\mathrm{b}}$ | $-0.2360^{\mathrm{b}}$ |  |  |
| 0.6 | $-0.3668^{\mathrm{b}}$ | -0.3015 |  |  |
| 0.7 | $-0.4337^{\mathrm{b}}$ | -0.3644 |  |  |
| 0.8 | $-0.5035^{\mathrm{b}}$ | -0.4316 |  |  |
| 0.9 | $-0.573^{\mathrm{b}}$ | -0.5004 | -1.968 | -0.580 |
| 1.0 |  |  |  |  |

${ }^{a}$ All values are the result of $10^{6}$ configurations.
${ }^{\mathrm{b}}$ Values used to calculate $F\left(\Gamma_{1}\right) / N k T$ numerically. The Abe analytic cluster expansion was used for $0 \leqslant \Gamma \leqslant 0.1$.
the relative and absolute errors of the fit and the values of the parameters. The fit to this form is extremely good; the relative rms error is $3 \times 10^{-5}$. A suggestive argument is presented by DeWitt and Rosenfeld ${ }^{6}$ for the form of Eq. (5).
To find the liquid equation of state we calculate the Helmholtz free energy from the formula

$$
\frac{F(\Gamma)}{N k T}=\int_{0}^{\Gamma} \frac{U}{N k T} \frac{d \Gamma^{\prime}}{\Gamma^{\prime}} .
$$

Substituting for $U / N k T$ from Eq. (5) we find

$$
\begin{align*}
\frac{F(\Gamma)}{N k T}= & a \Gamma+4\left(b \Gamma^{1 / 4}-\frac{C}{\Gamma^{1 / 4}}\right)+d \ln \Gamma \\
& -[a+4(b-c)]+\frac{F\left(\Gamma_{1}\right)}{N k T}, \tag{6}
\end{align*}
$$

where $F\left(\Gamma_{1}\right) / N k T$ is the free energy at $\Gamma=1$. $F\left(\Gamma_{1}\right) / N k T$ was calculated by summing the results from a Simpson's rule integration of the starred values of $U / N k T$ given in Table $\mathrm{V}(0.1 \leqslant \Gamma \leqslant 1.0)$ and an analytic integration of the Abe cluster ${ }^{7}$ expansion formula [Eq. (19), Ref. 2] for $U / N k T$ $(0.0 \leqslant \Gamma \leqslant 0.1)$. The Abe formula was chosen for the lower end because it agrees well with the $U / N k T$ values given by Broyles, Sahlin, and Carley ${ }^{8}$ and by the hypernetted chain (HNC) approximation. ${ }^{9}$ The HNC values were used in the integration for $\Gamma=0.1,0.2,0.3,0.4$, and 0.5 . At $\Gamma=0.6$ the HNC $U / N k T$ agrees very well with the MC $U / N k T$ consequently the MC values were used in the integration for $0.6 \leqslant \Gamma \leqslant 1.0$. The integration gives the result $F\left(\Gamma_{1}\right) / N k T=-0.420$. We consider this result to be more accurate than the corresponding Hansen ${ }^{2}$ value [ $F_{\text {Hansen }}\left(\Gamma_{1}\right) / N k T=$ -0.445 ] since a much denser grid was used for the integration and the $U / N k T$ values are more accurate. [Note that $F\left(\Gamma_{1}\right) / N k T$ in Hansen's paper includes the perfect gas contribution.]

After including the perfect gas contribution [Eq.
(23), Ref. 2] Eq. (6) becomes

$$
\begin{align*}
F(\Gamma) / N k T= & a \Gamma+4\left(b \Gamma^{1 / 4}-c / \Gamma^{1 / 4}\right) \\
& +(d+3) \ln \Gamma-[a+4(b-c)+1.135] . \tag{7}
\end{align*}
$$

This formula for the liquid free energy may be used to calculate other thermodynamic variables of interest. We will use it shortly, after we have calculated the solid free energy, to find $\Gamma$ at the liquid-solid transition ( $\Gamma_{m}$ ).

Following Pollock and Hansen ${ }^{10}$ we fit our 128-ion bcc lattice data to the form

$$
\begin{equation*}
\frac{U}{N k T}=\frac{U_{0}}{N k T}+\frac{3}{2}+\frac{h}{\Gamma^{2}}, \tag{8}
\end{equation*}
$$

by minimizing the squares of the relative errors. For the bcc lattice, $U_{0} / N k T=-0.895929 \Gamma$. The fitted $U / N k T$ are shown in Table VI along with the relative and absolute errors of the fit and the value of the parameter $h$. To obtain the Helmholtz free energy we use the harmonic approximation results from Ref. 10. This leads to
$\frac{F}{N k T}=-0.895929256 \Gamma+\frac{9}{2} \ln \Gamma-1.8856-\frac{h}{2 \Gamma^{2}}$.
Since we now have formulas for the free energy

TABLE VI. Fitted solid data for $160 \leqslant \Gamma 300 . U / N k T$ $=1.5+U_{0}+h / \Gamma^{2}$ where $U_{0}=-0.895929 \Gamma$ and $h=2980$.

| $\Gamma$ | $U / N k T$ <br> MC | $U / N k T$ <br> fitted | Relative <br> error | Absolute <br> error |
| :---: | :---: | :---: | ---: | ---: |
| 160.0 | -141.729 | -141.732 | -0.00002 | 0.003 |
| 170.0 | -150.703 | -150.705 | -0.00001 | 0.002 |
| 180.0 | -159.675 | -159.675 | 0.00000 | 0.000 |
| 200.0 | -177.619 | -177.611 | 0.00004 | -0.008 |
| 220.0 | -195.551 | -195.543 | 0.00004 | -0.008 |
| 240.0 | -213.469 | -213.471 | -0.00001 | 0.002 |
| 300.0 | -267.243 | -267.246 | -0.00001 | 0.003 |

of both the liquid [Eq. (7)] and solid [Eq. (9)] we are now in a position to find the intersection of the two curves which gives the fluid-solid transition. We find that $\Gamma_{m}=168 \pm 4$. The error in $\Gamma_{m}$ was calculated by systematically biasing the $U / N k T$ data (Table II) by $\pm \sigma_{i}$. The error of $\pm 4$ corresponds the maximum and minimum intersection points of the three liquid free-energy curves (obtained by fitting $U_{i} / N k T-\sigma_{i}, U_{i} / N k T$, and $U_{i} / N k T+\sigma_{i}$ ) with the corresponding three solid free-energy curves.
Our results for $\Gamma_{m}=168 \pm 4$ may be compared with the value of Pollock and Hansen, ${ }^{10} \Gamma_{m}=155 \pm 10$. Again the difference is attributed to the relatively large systematic error in Hansen's ${ }^{2}$ liquid data (Fig. 2). We note here that Van Horn's ${ }^{11}$ estimate of $\Gamma_{m}=170 \pm 10$ (based on the Lindeman melting criterion, harmonic theory, and empirical data) is in agreement with our more exact result.

## IV. CONCLUSION

Using a very accurate representation of the Ewald potential and a Monte Carlo technique we have evaluated the internal energy of the classical one-component plasma ( $1 \leqslant \Gamma \leqslant 300$ ). The liquid internal-energy data were fitted to a simple fourparameter equation which was then integrated to obtain the Helmholtz free energy of the system. After fitting the lattice internal-energy data to a one-parameter formula, and obtaining the solid free energy, we find that the fluid-solid transition is at $\Gamma=168 \pm 4$.
Note added in proof. Subsequent calculations with $N=250,432$, and 686 have shown a dependence of $U / N k T$ on $N$ for values of $\Gamma$ just below melting. The most recent $\Gamma_{m}$ using the new results is $\Gamma_{m}=171 \pm 3$.

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## APPENDIX A: CUBIC HARMONIC POTENTIAL APPROXIMATION

In this Appendix we discuss the form of the potential approximation following Slater and DeCicco. ${ }^{5}$ For a system of point charges surrounded by a uniform background charge, we expect the potential energy of a point charge to be given by
the sum of a solution to Poisson's equation for the uniform background [Eq. (A1)] and another solution for the point charges [Eq. (A2)].

First, for the potential of the uniform negative background, we need the solution to Poisson's equation $\nabla^{2} \phi=-4 \pi \rho$ in spherical coordinates. In this case the charge density $\rho$ is $-Z e$. This leads to the solution $\phi=2 \pi r^{2} Z e$ and thus the potential energy per NKT may be written as

$$
\begin{equation*}
\frac{W_{1}}{N k T}=\frac{(Z e)^{2}}{N k T} \phi=\frac{\Gamma}{N L} \frac{2}{3} \pi r^{2}, \tag{A1}
\end{equation*}
$$

since $\bar{r} L \equiv 1$. A term of this form did not appear in Hansen's approximation.
Next, we anticipate that the solution to Poisson's equation with cubic symmetry for the point charges is given by the expression

$$
\begin{aligned}
\frac{W_{2}}{N k T}= & \lim _{r \rightarrow 0} \frac{\Gamma}{N L}\left[u_{1}(r)+1+u_{2}(r)\right. \\
& +\frac{\Gamma}{N L}(\text { fitted cubic harmonics })
\end{aligned}
$$

which in the limit goes to

$$
\begin{align*}
\frac{W_{2}}{N k T}= & \frac{\Gamma}{N L}\left(\frac{1}{r}+1-E_{m}\right) \\
& +\frac{\Gamma}{N L}(\text { fitted cubic harmonics }) \tag{A2}
\end{align*}
$$

The fitting parameters in the cubic harmonics remain to be determined.
We digress for a few comments about the cubic harmonics. Explicitly we used cubic harmonic polynomials in the coordinates $l, m$, and $n$ where $l=x / r, m=y / r$, and $n=z / r$, and $x, y$, and $z$ are the usual Cartesian coordinates.

Examples of the first 3 normalized functions follow:

$$
\begin{aligned}
K_{2}= & P_{2}(l)+P_{2}(m)+P_{2}(n)=0, \\
K_{4}= & V_{1}^{\prime} r^{4}\left[P_{4}(l)+P_{4}(m)+P_{4}(n)\right] \\
= & V_{1} r^{4}\left(l^{4}+m^{4}+n^{4}-\frac{3}{5}\right), \\
K_{6}= & V_{2}^{\prime} r^{6}\left[P_{6}(l)+P_{6}(m)+P_{6}(n)\right] \\
= & V_{2} r^{6}\left[l^{6}+m^{6}+n^{6}\right. \\
& \left.\quad-\frac{15}{11}\left(l^{4}+m^{4}+n^{4}\right)+\frac{30}{77}\right],
\end{aligned}
$$

where $P_{i}(l)$ are the usual Legendre polynomials and the $V_{i}$ 's are the fitted constants (the $V_{i}^{\prime}$ 's differ by a constant factor).

To obtain the required accuracy we used terms through $K_{22}$. Terms of degree 12, 16, 18, 20, and 22 are degenerate and consequently each needs an orthogonal function of the same degree to complete the basis set. We used only the 12th degree function,

$$
\bar{K}_{12}=V_{11} \operatorname{Re}\left[(l+i m)^{12}+(m+i n)^{12}+(n+i l)^{12}\right],
$$

where the $V_{11}$ is fitted along with $V_{1}$ (for $K_{4}$ ) through $V_{10}$ (for $K_{22}$ ).
Adding Eqs. (A1) and (A2) we obtain for the potential between two particles with separating distance $r=\left(x^{2}+y^{2}+z^{2}\right)^{1 / 2}$,

$$
\begin{aligned}
\frac{\psi}{N K T}=\frac{\Gamma}{N L} & \left(\frac{1}{r}+C_{0}+C_{2} r^{2}+V_{1} r^{4}\left(S_{4}+C_{4,1}\right)\right. \\
& +V_{2} r^{6}\left(S_{6}+C_{6,1} S_{4}+C_{6,2}\right) \\
& \cdot \\
& \cdot \\
& \cdot \\
& +V_{10} r^{22}\left(S_{22}+C_{22,1} S_{20}+C_{22,2} S_{18}+\cdots\right. \\
& \left.\left.+C_{22,9} S_{4}+C_{22,10}\right)+V_{11} \bar{K}_{12}\right)
\end{aligned}
$$

where $C_{0}=1-E_{m}, C_{2}=2 \pi / 3, S_{j}=\left(x^{j}+y^{j}+z^{j}\right) / r^{j}$, and the $C_{j i}$ are the collected Legendre coefficients

For computational convenience this expression is rearranged by collecting like powers of $r^{2}$. Letting $T_{j}=S_{j} r^{j}=x^{j}+y^{j}+z^{j}, U_{j}=V_{(j-2) / 2}$, and $U_{2}=C_{2}$ we have
$\frac{\psi}{N K T}=\frac{\Gamma}{N L}\left(\frac{1}{r}+C_{0}\right.$

$$
\begin{align*}
& +T_{2}\left(U_{2}+C_{4,1} U_{4} r^{2}+C_{6,2} U_{6} r^{4}+C_{8,3} U_{8} r^{6}\right. \\
& \left.\quad+\cdots+C_{22,10} U_{22} r^{20}\right) \\
& +T_{4}\left(U_{4}+C_{6,1} U_{6} r^{2}+C_{8,2} U_{8} r^{4}+C_{10,3} U_{10} r^{6}\right. \\
& \left.\quad+\cdots+C_{22,9} U_{11} r^{18}\right) \\
& \text { - } \\
& \text { - } \\
& \text { - } \\
& +T_{20}\left(U_{20}+C_{22,1} U_{22} r^{2}\right) \\
& +T_{22} U_{22}  \tag{A3}\\
& \left.+U_{24} \bar{K}_{12}\right)
\end{align*}
$$

The coefficients for the $r^{2}$ terms (e.g., $C_{4,1} U_{4}$ and $C_{22,1} U_{22}$ ) are calculated once and placed in a matrix $\left(A_{j i}\right)$ which is subsequently used in calculating the coefficients for the $T_{j}$ 's.
To calculate the potential given the separation $x, y$, and $z$ we first compute the $T_{j}$ coefficient as polynomials in $r^{2}$ using $A_{j i}$. Since the $T_{j}$ 's are simply polynomials in $x^{2}, y^{2}$, and $z^{2}$ we next calculate each $x^{2}, y^{2}$, and $z^{2}$ polynomial using the $T_{j}$ coefficients. These results are then added to the left-over terms ( $1 / r+C_{0}+U_{24} \bar{K}_{12}$ ) for the final result.

A table of 8000 values of the function $\psi / N k T$
[Eqs. (2), (3), and (4) were generated. These values were approximated by the function $\psi / N k T$ in Eq. (A3) by adjusting the parameters $V_{1}, V_{2}, \ldots$ so as to minimize the mean-square error. Note that Eq. (A3) does not require evaluation of the error function complement.

## APPENDIX B: EXTENSION OF THE RANGE OF THE PAIR DISTRIBUTION FUNCTIONS FOR SPHERICALLY SYMMETRIC DISTRIBUTIONS

In a Monte Carlo calculation, for each configuration the pair separations are computed between a particle under consideration and all others whose $x, y$, and $z$ coordinates are within half of the periodic distance. The number of particles within an interval at a certain distance, say $a \pm d$ where $d$ is small radial increment, are accumulated into a bin. At the end of the calculation the number in the bin is divided by the volume of the bin at that distance times the number of configurations to get the pair distribution function for that bin. Although all pair separations are calculated within a box of side $L$ centered on a particle, pair distributions are usually only tabulated for pairs in a sphere of diameter $L$ centered on the particle. About half of the pairs in the box are outside the sphere. We show here a method for using the pairs which are usually neglected. For spherically sy imetric distributions, this method will extend the jair distribution function from $\frac{1}{2}$ to $\sqrt{\frac{3}{2}}$ where the length of the edge of the basic cube is taken as 1 . For distributions without spherical symmetry, the following procedure provides a method for determining the deviation from spherical symmetry.

Consider a sphere of radius a centered in a cube with edge length 1. For a spherical shell with radial increment $d$, the volume is given by the expression $V_{1}(a, d)=V_{V}(a)-S_{V}(a-d)$ where $S_{V}(a)=$


FIG. 3. Normalizing volume for the pair distribution functions.
$\frac{4}{3} \pi a^{3}$. As the shell grows outwards, the outer sphere becomes tangent to the six faces of the cube at $a=\frac{1}{2}$. For $a>\frac{1}{2}$ and before the outer sphere is tangent to the edges of the cube, at $a=\sqrt{2} / 2$, the volume of the shells must be diminished by the volume of the "caps" protruding beyond the faces of the cube. The volume of the shells in this region are given by the formula: $V_{2}(a, d)=V_{1}(a, d)$ $-\left[C_{V}(a)-C_{V}(a-d)\right]$, where $C_{V}(a)=2 \pi\left(a-\frac{1}{2}\right)^{2}\left(2 a+\frac{1}{2}\right)$. For $\sqrt{2} / 2<a<\sqrt{3} / 2$ the caps overlap at the edges of the cube and the shell volume is diminished unnecessarily by the volume in the 12 overlaps. The volume of the shells in this overlap region may be expressed as

$$
V_{3}(a, d)=V_{2}(a, d)+[O(a)-O(a-d)]
$$

where

$$
\begin{aligned}
O(a)= & \left(12 a^{2}-1\right)\left\{\sin ^{-1 \frac{1}{2}} /\left[\left(a^{2}-\frac{1}{4}\right)^{1 / 2}\right]-\frac{1}{2} \pi\right\} \\
& +16 a^{3} \sin ^{-1}\left\{\left(a^{2}-\frac{1}{2}\right)^{1 / 2} /\left[2\left(a^{2}-\frac{1}{4}\right)\right]^{1 / 2}\right\} \\
& +2\left(a^{2}-\frac{1}{2}\right)^{1 / 2} .
\end{aligned}
$$

Figure 3 shows the relative volumes of the shells out to $a=\sqrt{3} / 2$. The smaller volumes near $\sqrt{3} / 2$ (where the spherical shells are near the vertices of the cube) indicate decreasing accuracy in the statistics for the pair distribution functions in this region.
${ }^{1}$ S. G. Brush, H. L. Sahlin, and E. Teller, J. Chem. Phys. 45, 2102 (1966).
${ }^{2}$ J. P. Hansen, Phys. Rev. A 8, 3096 (1973).
${ }^{3}$ H. E. DeWitt, Phys. Rev. A 14, 1290 (1976).
${ }^{4}$ H. E. DeWitt and W. B. Hubbard, Astrophys. J. 205, 295 (1976).
${ }^{5}$ J. C. Slater and P. DeCicco, MIT Solid State and Molecular Theory Group Quarterly Progress Report No. 50, 46 (1963).
${ }^{6}$ H. E. DeWitt and Y. Rosenfeld, Phys. Lett. (in press).
${ }^{7}$ R. Abe, Prog. Theor. Phys. 22, 213 (1959).
${ }^{8}$ A. A. Broyles, H. L. Sahlin, and D. D. Carley, Phys. Rev. Lett. 10, 319 (1963).
${ }^{9}$ Calculated by the method of K-C. Ng, J. Chem. Phys. 61, 2680 (1974).
${ }^{10}$ E. L. Pollock and J. P. Hansen, Phys. Rev. A 8, 3110 (1973).
${ }^{11}$ H. M. Van Horn, Phys. Lett. 28A, 707 (1969).


[^0]:    ${ }^{\mathrm{a}}$ liq denotes liquid.
    ${ }^{\mathrm{b}}$ lat denotes lattice.

