Perturbation theory of the Stark effect in hydrogen to arbitrarily high order

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The solution of the Stark effect in hydrogen to arbitrarily high orders of perturbation theory is made feasible by the explicit formula for the Nth-order energy in terms of the separation constants through Nth order, derived here. The Nth-order separation constant $\beta_i^{(N)}$ is shown to be a polynomial of total degree N+1 in the parabolic quantum number n_i and the magnetic quantum number m. The polynomial coefficients have been tabulated through seventeenth order and are listed here through tenth order. Similarly, the Nth-order energy is a polynomial in the quantum numbers n_1 , n_2 , and m. The polynomial coefficients (which are more numerous than for $\beta_i^{(N)}$) have been tabulated through seventeenth order and are listed here through seventh order. Seventeenth order is high enough to permit a clear numerical demonstration of the asymptotic character of the perturbation series, and a "maximum useful field strength" is defined and illustrated. Energies calculated by perturbation theory for specific states are shown to be in excellent agreement with energies calculated nonperturbatively.

I. INTRODUCTION

The Stark effect in hydrogen was the first quantum-mechanical problem to be treated by Rayleigh-Schrödinger perturbation theory.¹ The perturbed energy is easy to obtain in first and second order. Third order² requires some effort, but fourth order^{3,4} is already so tedious that the correct formula was finally obtained only in 1973.⁵ Higher orders have not been published, except for the ground state,^{6,7} although there have been a number of nonperturbative calculations.⁸⁻¹⁶

Recently there has been experimental interest in high Rydberg states of atoms.¹⁷ Since detection techniques for these hydrogenic states have involved ionization of the atoms by an external electric field, it is of some interest to have accurate formulas for the Stark shifts and for the ionization lifetimes. The asymptotic formula^{18,19} for the lifetimes depends sensitively on the accuracy of the computed perturbed energy and separation constant.

The purpose of this paper is to discuss the solution of the Stark effect in hydrogen to arbitrarily high orders of perturbation theory.

In each order of perturbation theory, the separation constant and energy are polynomials in the parabolic and magnetic quantum numbers. We show that the total degree of the polynomials is only about half of that expected. We give the polynomial coefficients for the separation constant to tenth order and the energy coefficients to seventh order. Tables of both through seventeenth order are available from the Physics Auxiliary Publication Service (PAPS).²⁰

Although the relation between the values of the separation constants and the energy is very simple, the relation between the two perturbation series is not, because the expansion parameter for the separation constant series explicitly involves the energy. In the past it was unnecessary to go beyond fourth order, and a successive substitution method sufficed. Here the order may be arbitrarily high, so that we must derive the explicit relationship between the two series.

The Stark-effect perturbation series do not converge: they are asymptotic.²¹ For a given number of terms, it is possible to find a heuristic maximum useful field strength, which we illustrate through some simple numerical calculations.

II. PERTURBATION THEORY EQUATIONS FOR THE SEPARATION CONSTANTS

In this section the perturbation-theory equations for the separation constants are restated so that all the relevant quantities are polynomials in the quantum numbers.

The Schrödinger equation for hydrogen in a uniform electrostatic field,

$$(-\frac{1}{2}\nabla^2 - 1/r + Fz - E)\Psi = 0$$
 (1)

separates in scaled parabolic coordinates,^{1,22,23}

$$x = (-2E)^{-1/2} (\sigma \rho)^{1/2} \cos \phi , \qquad (2)$$

$$y = (-2E)^{-1/2} (\sigma \rho)^{1/2} \sin \phi, \qquad (3)$$

$$c = \frac{1}{2} (-2E)^{-1/2} (\sigma - \rho), \qquad (4)$$

$$\Psi = (\sigma \rho)^{-1/2} \Phi_1(\sigma) \Phi_2(\rho) e^{i m \phi}, \qquad (5)$$

$$\left[-\sigma\left(\frac{d}{d\sigma}\right)^2 + \frac{(m^2 - 1)}{(4\sigma)} + \frac{1}{4}\sigma + f\sigma^2 - \beta_1\right]\Phi_1(\sigma) = 0, \quad (6)$$

$$\left[-\rho\left(\frac{d}{d\rho}\right)^{2}+\frac{(m^{2}-1)}{(4\rho)}+\frac{1}{4}\rho-f\rho^{2}-\beta_{2}\right]\Phi_{2}(\rho)=0, \quad (7)$$

$$f = \frac{1}{4} \left(-2E \right)^{-3/2} F.$$
(8)

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The energy is determined from the separation constants by

$$E = -\frac{1}{2}(\beta_1 + \beta_2)^{-2}.$$
 (9)

Since the quantum number m appears only in the form $\frac{1}{4}(m^2-1)$, we denote the combination by M for convenience²⁴

$$M = \frac{1}{4}(m^2 - 1). \tag{10}$$

We also introduce the symbol k_i for the unperturbed separation constants,⁵ which play a major role in subsequent formulas

$$k_i = \beta_i^{(0)} \tag{11}$$

$$= n_i + \frac{1}{2} |m| + \frac{1}{2} (n_i = 0, 1, 2, ...).$$
 (12)

 k_i is a more convenient label than the parabolic quantum numbers n_1 and n_2 .

Since the essential difference between Eqs. (6) and (7) for β_1 and β_2 is the sign of f, it is sufficient to consider only one, say Eq. (6), and we write

$$\beta_1 = \beta(\boldsymbol{k}_1, \boldsymbol{M}, \boldsymbol{f}), \qquad (13)$$

$$\beta_2 = \beta(k_2, M, -f),$$
 (14)

$$\beta(k, M, f) = \sum_{n=0}^{\infty} \beta^{(N)}(k, M) f^{N}.$$
 (15)

When there is no danger of confusion, subscripts will be omitted. Subscripts will be indicated for the parabolic quantum numbers n_1 and n_2 , however, because we wish to reserve *n* (without a subscript) to represent the "principal quantum number."

To make all the important quantities in the calculation polynomials, we introduce an *un-normalized* basis. Let $|i\rangle$ denote the function.

$$|j\rangle = N_{j} [(n_{1} + j)! / (n_{1} + j + |m|)!^{3}]^{1/2}$$

$$\times \sigma^{(|m|+1)/2} e^{-\sigma/2} L_{n_{1}+j+|m|}^{|m|} (\sigma) , \qquad (16)$$

where $L_{n_1+j+|m|}^{|m|}$ denotes the usual generalized Laguerre polynomial,²⁵ and N_j is a constant that spoils the normalization, given by

$$N_{j} = \prod_{i=0}^{j-1} \left[(k+i)(k+i+1) - M \right]^{1/2} \quad (j \ge 1)$$

= 1 (j = 0) (17)
=
$$\prod_{i=1}^{|j|} \left[(k-i)(k-i+1) - M \right]^{-1/2} \quad (-1 \ge j \ge -n_{1}).$$

(18)

[Here k means $n_1 + \frac{1}{2}|m| + \frac{1}{2}$, as in Eq. (12).] The $|j\rangle$ are eigenfunctions of the unperturbed Hamiltonian $h^{(0)}$,

$$h^{(0)} \left| j \right\rangle = \left[-\sigma \left(\frac{d}{d\sigma} \right)^2 + \frac{M}{\sigma} + \frac{1}{4} \sigma \right] \left| j \right\rangle \tag{19}$$

$$= (k+j) | j \rangle .$$
 (20)

Note that the difference of any two unperturbed eigenvalues is an integer. The function $|0\rangle$, which is normalized to unity (with volume element $\sigma^{-1}d\sigma$), is the unperturbed eigenfunction belonging to the unperturbed separation constant k.

The matrix of the perturbation, which is defined by

$$\sigma^{2}|j\rangle = \sum_{i} v_{ij}(k, M)|i\rangle, \qquad (21)$$

is straightforward to evaluate from the well-known properties of the generalized Laguerre polynomials.²⁶ One finds that the nonzero v_{ij} are polynomials in k and M:

$$v_{ij}(k,M) = 1 \quad (i = j + 2)$$
 (22)

$$= -4(k+j+\frac{1}{2}) \quad (i=j+1) \tag{23}$$

$$= 6(k+j)^2 - 2M \quad (i=j)$$
(24)

$$= -4(k+j-\frac{1}{2})[(k+j-1)(k+j)-M]$$

$$(i=j-1)$$
 (25)
=[$(k+j-2)(k+j-1)-M$]

$$\times \left[(k+j-1)(k+j) - M \right] \quad (i=j-2) \quad (26)$$

$$= 0 \quad (|i-j| > 2) \,. \tag{27}$$

In the usual manner, with "intermediate normalization," we expand the perturbed eigenfunctions $\Phi^{(N)}$ on the basis $\{|j\rangle\}$ and obtain recursive relations for the expansion coefficients and for the separation constant:

$$\Phi = \sum_{N} \Phi^{(N)} f^{N}, \qquad (28)$$

$$\Phi^{(0)} = \left| 0 \right\rangle, \tag{29}$$

$$\Phi^{(N)} = \sum_{j \neq 0} c_j^{(N)}(k, M) | j \rangle \quad (N \ge 1) , \qquad (30)$$

$$c_{j}^{(0)} = \delta_{j0},$$
 (31)

$$c_{l}^{(N)} = -\frac{1}{l} \left(\sum_{j \neq 0} v_{lj} c_{j}^{(N-1)} - \sum_{\nu=1}^{N-1} \beta^{(N-\nu)} c_{j}^{(\nu)} \right) \quad (l \neq 0),$$

$$\beta^{(N)}(k,M) = \sum_{j \neq 0} v_{0j} c_j^{(N-1)}.$$
(33)

Because each nonzero v_{ij} is a polynomial in k and M of total degree 2 - i + j (where $|i-j| \leq 2$) in k and $M^{1/2}$, it is apparent from Eqs. (31)-(33) that the wave function coefficients $c_i^{(N)}$ are also polynomials in k and M of total degree 2N - l in k and $M^{1/2}$. (Note that l must satisfy $-2N \leq l \leq 2N$ for $c_i^{(N)}$ to be nonzero.) Similarly, the separation constants $\beta^{(N)}(k, M)$ are polynomials in k and $M^{1/2}$. Actually, we prove in Sec. III that the degree of

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 $\beta^{(N)}$ is only N+1. Since $\beta^{(N)}$ is known⁵ to have parity $(-1)^{N+1}$ with respect to k, we can write for $c_j^{(N)}$ and for $\beta^{(N)}$,

$$c_{j}^{(N)}(k,M) = \sum_{s=0}^{2N-j} \sum_{r=0}^{\lfloor (2N-j-s)/2 \rfloor} c_{jrs}^{(N)} M^{r} k^{s}, \qquad (34)$$

$$\beta^{(N)} = \sum_{r=0}^{\lfloor (N+1)/2 \rfloor } \sum_{t=0}^{\lfloor (N+1)/2 \rfloor - r} \beta^{(N)}_{r, N+1-2r-2t} \times M^r k^{N+1-2r-2t}, \qquad (35)$$

where $c_{irs}^{(N)}$ and $\beta_{rs}^{(N)}$ are constant coefficients, and where [i] denotes the greatest integer $\leq i$.

By programming the recursion formulas (31)-(33) with the polynomial formulas (34), (35), and

(22)-(27), one can generate with a computer the numerical values for the polynomial coefficients to as high an order as is desired. In Table I we have listed $\beta_{rs}^{(N)}$ for $N \leq 10$. Orders 1-4 agree with Alliluev and Malkin.⁵ The coefficients appear to be even integers for $N \leq 10$. Since the calculation has been carried out in double precision on a Digital Equipment Corp. System-10 computer, it is not possible to tell if the coefficients are integers beyond tenth order, as the numbers of significant digits in the $\beta_{rs}^{(N)}$ are too large. The number of nonzero $\beta_{rs}^{(2N)}$ in 2Nth order is $\frac{1}{2}(N+1)(N+2)$, and the number in (2N+1)st order is $\frac{1}{2}(N+1)(N+4)$. Tables of $\beta_{rs}^{(N)}$ for $N \leq 17$ are available from PAPS.²⁰

TABLE I. Separation constant, orders 1-10. $\beta^{(N)} = \sum_{r} \beta^{(N)}_{rs} M^r k^s$.

N	r	s	$\beta_{rs}^{(N)}$	N	r	s	$\beta_{rs}^{(N)}$	N	r	s	$\beta_{rs}^{(N)}$
1	0	2	6	7	0	2	484 419 360		1	. 6	-18 689 372 876 160
	1	0	-2		0	4	3997074984		1	8	-4741154003520
					0	6	6142374000		2	0	104008003840
2	0	1	-10		0	8	1 905 807 816		2	2	2121721190976
	0	3	-68		1	0	-60 441 600		2	4	4 891 069 777 440
	1	1	36		1	2	-1 663 495 776		2	6	2112491924256
					1	4	-4778464320		3	0	-25354142720
3	0	2	660		1	6	-2 379 35 2 416		3	2	-313679336640
	0	4	1 500		2	0	46 142 240		3	4	-313746377280
	1	0	-120		2	2	65 7 7 68 432		4	0	2 306 958 208
	1	2	-1032		2	4	728298480		4	2	12987661680
	2	0	44		3	0	-8 886 688		5	0	-66513664
					3	2	-53877984				
4	0	1	-2210		4	0	450 696	10	0	1	-6733922487500
	0	3	-37 900						0	3	-255756162331848
	0	5	-42756	8	0	1	-2564063050		0	5	-1 272 588 709 670 760
	1	1	12960		0	3	-83943936120		0	7	-1872704078101248
	1	3	35640		0	5	-331953386940		0	9	-947164090664440
	2	1	-3636		0	7	$-330\ 097\ 271\ 640$		0	11	-131843747886864
					0	9	-75783847700		1	1	59 428 632 96 1 0 80
5	0	2	381264		1	1	20 998 605 600		1	3	640629497858400
•	0	4	2 093 280		1	3	182925628704	1. N. A.	1	5	${\bf 1} {\bf 534} {\bf 645} {\bf 588} {\bf 890} {\bf 960}$
	0	6	1400784		1	5	302 909 140 800		1	7	1130767102209600
	1	0	-54240		1	7	104928884208		1	9	218 322 729 416 880
	1	2	-1 029 600		2	1	-12 837 436 200		2	1	-40501147719024
	1	4	-1364160		2	3	-60110940120		2	3	-271104353086320
	2	0	33712		2	5	-39436837272		2	5	-369115567782192
	2	2	229 200		3	1	2155524912		2	7	-112468417051680
	3	0	-3680		3	3	4336206000		3	1	$8\ 691\ 108\ 323\ 328$
					4	1	-98 489 844		3,	3	35612199123840
6	0	1	-1 656 500						3	5	21 237 767 907 552
	0	3	-43002592	9	Õ	2	1078065703680		4	1	$-717\ 606\ 581\ 400$
	0	5	-113822520		0	4	11206248868800		4	3	-1 343 525 553 360
	0	7	-50118384		0	6	25 598 194 895 280		5	1	19 093 706 352
	1	1	12051720		0	8	17695955613600				
	1	3	72439200		0	10	3118465869552				
	1	5	55719216		1	0	-123998092800				
	2	1	-5881416		1	2	-4217580898560				
	2	3	-13182480		1	4	$-17\ 545\ 111\ 106\ 880$				
	3	1	552912								

III. THEOREM ON TOTAL DEGREE OF $\beta^{(N)}$

In this section we prove that the total degree of $\beta^{(N)}$ with respect to k and $M^{1/2}$ is N+1.

Alliluev and Malkin⁵ observed extensive cancellation in calculating $\beta^{(2)}$, $\beta^{(3)}$, and $\beta^{(4)}$ that reduced the total degrees in k and $M^{1/2}$ to 3, 4, and 5 from 4, 6, and 8, respectively. To prove the cancellation for arbitrary order, the author finds it necessary first to convert to an orthonormal basis, and then to introduce a new coordinate representation not directly connected to the σ representation, but which has the same matrix representation. The new coordinate representation yields a tractable equation for the logarithm of the eigenfunction, which in turn gives the polynomial degree of $\beta^{(N)}$ directly, without the necessity of cancelling higher terms.

The orthonormal basis is obtained by removing the N_i from the definition of the functions $|j\rangle$:

 $|j\rangle' = N_j^{-1} |j\rangle$.

In the $|j\rangle'$ basis, the matrix of the perturbation becomes²⁷

$$v'_{ij}(k, M) = [(k+j+1)(k+j+2) - M]^{1/2} \times [(k+j)(k+j+1) - M]^{1/2} \quad (i=j+2),$$
(36)

$$= -4(k+j+\frac{1}{2}) [(k+j)(k+j+1) - M]^{1/2} \quad (i=j+1), \quad (37)$$

$$= 6(k+j)^2 - 2M \quad (i=j), \quad (38)$$

$$= -4(k+j-\frac{1}{2})[(k+j)(k+j-1) - M]^{1/2} \quad (i=j-1), \quad (39)$$

$$= [(k+j-2)(k+j-1) - M]^{1/2} \quad (i=j-2), \quad (42)$$

$$= 0 (|i-j| > 2).$$
 (41)

The new coordinate representation is obtained by the correspondence,

$$|\mu\rangle' = (2\pi)^{-1/2} e^{i\mu\phi} .$$
(42)

In the ϕ representation the unperturbed Hamiltonian is essentially the angular momentum operator \hat{l}_{x} , while v becomes a function of both ϕ and \hat{l}_{x}

$$h^{(0)} = \hat{l}_{z} + k$$

$$= -i \frac{d}{d\phi} + k ,$$
(43)
(44)

$$\hat{v} = e^{2i\phi} [(k+\hat{l}_{z})(k+\hat{l}_{z}+1) - M]^{1/2} [(k+\hat{l}_{z}+1)(k+\hat{l}_{z}+2) - M]^{1/2} - 4e^{i\phi}(k+\hat{l}_{z}+\frac{1}{2})[(k+\hat{l}_{z})(k+\hat{l}_{z}+1) - M]^{1/2} + 6(k+\hat{l}_{z})^{2} - 2M - 4e^{-i\phi}(k+\hat{l}_{z}-\frac{1}{2})[(k+\hat{l}_{z})(k+\hat{l}_{z}-1) - M]^{1/2} + e^{-2i\phi}[(k+\hat{l}_{z}-2)(k+\hat{l}_{z}-1) - M]^{1/2} \\ \times [(k+\hat{l}_{z}-1)(k+\hat{l}_{z}) - M]^{1/2} .$$
(45)

The unperturbed eigenfunction $\Phi^{(0)}$ is just the constant $(2\pi)^{-1/2}$.

We introduce what is essentially the logarithmic derivative of Φ by

$$\Phi(\phi) = (2\pi)^{-1/2} \exp\left[i \int^{\phi} \chi(\phi') \, d\phi'\right],\tag{46}$$

where $\chi \to 0$ when $f \to 0$. Since $\Phi(\phi)$ is single valued, we must have

$$\int_{0}^{2\pi} \chi(\phi) \, d\phi = 0 \,. \tag{47}$$

The eigenvalue equation for Φ then yields the equation satisfied by χ ,

$$\chi(\phi) + k - \beta = -\exp\left[-i\int^{\phi}\chi(\phi')\,d\phi'\right]f\,\hat{v}\,\exp\left[i\int^{\phi}\chi(\phi')\,d\phi'\right]\,.$$
(48)

For the purpose of obtaining the behavior of χ and β for large k and M, we expand \hat{v} in a power series in \hat{l}_{z} ,

$$\hat{v} = v_0(\phi) + v_1(\phi)\hat{l}_z + v_2(\phi)\hat{l}_z^2 + \cdots,$$
(49)

and χ in a power series in f,

$$\chi = f \chi^{(1)} + f^2 \chi^{(2)} + \cdots$$
 (50)

The v_{μ} depend on ϕ ; for instance,

$$v_{0}(\phi) = e^{2i\phi} [k(k+1) - M]^{1/2} [(k+1)(k+2) - M]^{1/2} - 4e^{i\phi}(k+\frac{1}{2}) [k(k+1) - M]^{1/2} + 6k^{2} - 2M$$

-4e^{-i\phi}(k-\frac{1}{2}) [k(k-1) - M]^{1/2} + e^{-2i\phi} [(k-1)(k-2) - M]^{1/2} [(k-1)k - M]^{1/2}. (51)

From Eqs. (45) and (49) we can characterize the behavior of v_{μ} for large k and M, or more precisely, for fixed but arbitrary k^2/M ,

$$v_{\mu}(\phi) = O[(ak + bM^{1/2})^{2-\mu}], \qquad (52)$$

where the constants a and b depend on the ratio k^2/M .

The system of equations obtained by substituting Eqs. (15), (49), and (50) for β , \hat{v} , and χ into Eq. (48) and collecting froms of the same order in f,

$$k - \beta^{(0)} = 0$$
, (53)

$$\chi^{(1)} - \beta^{(1)} = -v_0(\phi), \tag{54}$$

$$\chi^{(2)} - \beta^{(2)} = -\sum_{\mu=1}^{\infty} v_{\mu}(\phi) \left(-i \frac{\omega}{d\phi} \right) \chi^{(1)},$$
(55)
$$\sum_{\mu=1}^{\infty} \left(-d \right)^{\mu-1} \sum_{\nu=1}^{\infty} \left(-i \frac{\omega}{d\phi} \right) \chi^{(1)} ,$$

$$\chi^{(3)} - \beta^{(3)} = -\sum_{\mu=1}^{\infty} v_{\mu}(\phi) \left(i \frac{d}{d\phi} \right)^{\mu-1} \chi^{(2)} - \sum_{\mu=0}^{\infty} v_{\mu}(\phi) (-i)^{\mu-2} \sum_{\nu=1}^{\mu-1} \binom{\mu-1}{\nu} \left[\left(\frac{d}{d\phi} \right)^{\mu-1-\nu} \chi^{(1)} \right] \left[\left(\frac{d}{d\phi} \right)^{\nu-1} \chi^{(1)} \right], \quad (56)$$

and so forth, can be solved recursively. The $\beta^{(N)}$ are obtained via the single-valuedness condition (47)

$$\beta^{(1)} = (2\pi)^{-1} \int_0^{2\pi} v_0(\phi) \, d\phi \,, \tag{57}$$

$$\beta^{(2)} = (2\pi)^{-1} \int_0^{2\pi} \sum_{\mu=1}^\infty v_\mu(\phi) \left(-i \frac{d}{d\phi}\right)^{\mu-1} \chi^{(1)} d\phi , \qquad (58)$$

and so forth. It is not our purpose to solve Eqs. (53)-(56) for $\chi^{(N)}$ and $\beta^{(N)}$, but only to determine their behavior for large k and M. From Eqs. (51), (52), and (57), we see that $\chi^{(1)}$ and $\beta^{(1)}$ are $O[(ak+b M^{1/2})^2]$. From Eqs. (52) and (55), we see that $\chi^{(2)}$ and $\beta^{(2)}$ are $O[(ak+b M^{1/2})^3]$. It follows inductively from Eqs. (52) and (47) and from the system of equations whose first four members are Eqs. (53)-(56), that $\chi^{(N)}$ and $\beta^{(N)}$ are $O[(ak+b M^{1/2})^{N+1}]$. But since $\beta^{(N)}$ is a polynomial in k and M, to know its behavior for large k and $M^{1/2}$ is to know its total degree.

IV. PERTURBATION SERIES FOR E FROM THE SERIES FOR β_1 AND β_2

The usual procedure to obtain the series in F for E_*

$$E = E^{(0)} + E^{(1)}F + E^{(2)}F^2 + \cdots$$
(59)

from the series in f for $\beta_1 + \beta_2$ —expansion of $-\frac{1}{2}(\beta_1 + \beta_2)^{-2}$ in a series in f and successive ap-

proximation to eliminate E from $f = F(-2E)^{-3/2}$ is unwieldy after the first few terms, and also unnecessary: the "formula of Lagrange"²⁸ performs both steps at once.

Let us denote by γ the sum of the separation constants,

$$\gamma = \beta_1 + \beta_2 \tag{60}$$

$$=\gamma^{(0)} + \gamma^{(1)}f + \gamma^{(2)}f^{2} + \cdots, \qquad (61)$$

$$\gamma^{(N)} = \beta_1^{(N)} + \beta_2^{(N)} \tag{62}$$

$$=\beta^{(N)}(k_1, M) + (-1)^N \beta^{(N)}(k_2, M).$$
 (63)

 γ may be regarded as the "perturbed principal quantum number," since $\gamma^{(0)}$ is the usual unperturbed principal quantum number n

$$\gamma^{(0)} = k_1 + k_2 \tag{64}$$

$$=n_1+n_2+|m|+1$$
 (65)

(66)

The explicit formula that we derive for $E^{(N)}$ in terms of the sums of the separation constants $\gamma^{(N)}$ is

=n.

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$$E^{(N)} = 4^{-N} \sum_{\substack{i_1 \ge 0, i_2 \ge 0, \cdots, i_N \ge 0\\(i_1 + 2i_2 + 3i_3 + \cdots + Ni_N = N)}} \frac{(3N - 3)!}{(3N - 2 - \sum_j i_j)! i_1! i_2! \cdots i_N!} \gamma^{(0)_{3N-2} - \sum_j i_j} \gamma^{(1)i_1} \gamma^{(2)i_2} \cdots \gamma^{(N)i_N}.$$
(67)

For example,

j

$$E^{(1)} = \frac{1}{4} \gamma^{(1)} , \qquad (68)$$

$$E^{(2)} = 4^{-2} (\gamma^{(0)3} \gamma^{(2)} + \frac{3}{2} \gamma^{(0)2} \gamma^{(1)2}), \qquad (69)$$

$$\Sigma^{(3)} = 4^{-3} (\gamma^{(0)6} \gamma^{(3)} + 6 \gamma^{(0)5} \gamma^{(1)} \gamma^{(2)})$$

$$+5\gamma^{(0)4}\gamma^{(1)3}). (70)$$

The summation in Eq. (67), although complex in appearance, is just over the distinct partitions of N into sums of positive integers.²⁹

The derivation takes only a few lines. The relations (8), (9), (60), and (61) among E, $\beta_1 + \beta_2$, f, F, and γ translate into an implicit equation for γ :

$$\gamma = \gamma^{(0)} + \gamma^{(1)} (\frac{1}{4} \gamma^3 F) + \gamma^{(2)} (\frac{1}{4} \gamma^3 F)^2 + \cdots$$
 (71)

Langrange's formula gives any desired function of the solution of the above implicit equation, in particular, $-\frac{1}{2}\gamma^{-2}$:

$$E = -\frac{1}{2}\gamma^{-2}, \qquad (72)$$

$$= -\frac{1}{2}\gamma^{(0)-2} + \sum_{\mu=1}^{\infty} \frac{1}{\mu!} \left(\frac{d}{d\gamma^{(0)}}\right)^{\mu-1} \times \left\{ \left[\frac{d}{d\gamma^{(0)}} \left(-\frac{1}{2}\gamma^{(0)-2}\right)\right] \times \left[\sum_{\nu=1}^{\infty}\gamma^{(\nu)} \left(\frac{1}{4}\gamma^{(0)3}F\right)^{\nu}\right]^{\mu} \right\}. \qquad (73)$$

Collecting all terms multiplying F^N and evaluating the derivatives, one obtains Eq. (67).

V. POLYNOMIAL FORMS FOR $E^{(N)}$

 $E^{(N)}$ can be evaluated as a polynomial in k_1 , k_2 , and M via Eq. (67), which expresses $E^{(N)}$ in terms of the $\gamma^{(N)}$, which in turn are expressed as polynomials via Eq. (35). The resulting formula is slightly simplified, however, by first factoring $\gamma^{(0)} = k_1 + k_2 = n$ out of each $\gamma^{(N)}$

For even N, we find from Eqs. (63) and (35) that

$$\gamma^{(N)} = \sum_{r=0}^{N/2} \sum_{\mu=0}^{N/2-r} \beta^{(N)}_{r,2\mu+1} M^{r} (k_{1}^{2\mu+1} + k_{2}^{2\mu+1})$$
(N even), (74)

$$= (k_1 + k_2) \sum_{r=0}^{N/2} \sum_{\mu=0}^{N/2-r} \beta_{r,2\mu+1}^{(N)} M^r \times (k_1^{2\mu} - k_1^{2\mu-1}k_2 + k_1^{2\mu-2}k_2^2 - \cdots + k_2^{2\mu}) (N \text{ even}),$$
(75)

while for odd N,

$$\gamma^{(N)} = \sum_{r=0}^{(N+1)/2} \sum_{\mu=0}^{(N+1)/2-r} \beta_{r,2\mu}^{(N)} M^r (k_1^{2\mu} - k_2^{2\mu}) \quad (N \text{ odd}),$$

$$= (k_1 + k_2) \sum_{r=0}^{(N+1)/2} \sum_{r=0}^{(N+1)/2-r} \beta_{r,2\mu}^{(N)} M^r (k_1^{2\mu-1} - k_2^{2\mu-2} k_2 + \dots - k_2^{2\mu-1}) \quad (N \text{ odd}).$$
(76)

Thus the product of
$$\gamma$$
's in Eq. (67) has the natural factorization.

 $\gamma^{(0)_{3N-2}-\sum_{j}i_{j}}\gamma^{(1)i_{1}}\gamma^{(2)i_{2}}\cdots\gamma^{(N)i_{N}}$

$$= n^{3N-2} \left(\frac{\gamma^{(1)}}{\gamma^{(0)}}\right)^{i_1} \left(\frac{\gamma^{(2)}}{\gamma^{(0)}}\right)^{i_2} \cdots \left(\frac{\gamma^{(N)}}{\gamma^{(0)}}\right)^{i_N} .$$
(78)

Taking this factorization into account, we infer from Eq. (67) the formula for $E^{(N)}$:

$$E^{(N)} = 4^{-N} n^{3N-2} \sum_{\substack{r \ge 0 \ i \ge 0 \ j \ge 0 \ i \neq 0, \\ (i+j+Neven, \\ i+j+2r \le N}} \sum_{\substack{r \ge 0 \ i \ge 0 \ j \ge 0, \\ ri \neq k+2r \le N}} E^{(N)}_{rij} M^r k_1^i k_2^j.$$
(79)

When N is odd, it is clear from Eq. (76) that $\gamma^{(N)}$ is divisible by $k_1 - k_2$. Consequently $E^{(N)}$ [each term of which in Eq. (67) must contain at least one γ factor of odd order] is also divisible by $k_1 - k_2$, implying an alternative polynomial decom-

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$E_{rij}^{(N)}$	j	i	r	N	$E_{rij}^{(N)}$	j	i	r	N	$E_{rij}^{(N)}$	j	i	r	N	$E_{rij}^{(N)}$	j	i	r	N
-387 343 440	3	2	0		-18 397 680	3	1	0		31 920	0	5	0	5	6	0	1	Ó	1
-647 145 360	4	1	0		-8 482 320	4	0	0		95 088	1	4	0		-6	1	0	0	
-259 786 800	5	0	0		$-12\ 409\ 432$	0	2	0		62 160	2	3	0						
692 482 344	0	3	0		-10 958 528	1	1	0		-62160	3	2	0		1 4	0	2	0	2
537 695 544	1	2	0		-12 409 432	2	0	0		-95 088	4	1	0		-40	1	1	0	
-537 695 544	2	1	Ò		-1 360 000	0	0	0		-31920	5	0	0		-14	2	0	0	
-692 482 344	3	0	0		$909\ 216$	0	4	1		277 440	0	3	0		-10	0	0	0	
256 412 640	0	1	0		3445344	1	3	1		265440	1	2	0		36	0	0	1	
-256 412 640	1	0	0		5047776	2	2	1		-265 440	2	1	0						
-9 092 832	0	5	1		3445344	3	1	1		-277 440	3	0	0		132	0	3	0	3
-27 488 928	1	4	1		909 216	4	0	1		182544	0	1	0		156	1	2	0	
-19 002 912	2	3	1		9 907 200	0	2	1		-182544	1	0	0		-156	2	1	0	
19 002 912	3	2	1		18734400	1	1	1		11 90 4	0	3	1		-132	3	0	0	
27 488 928	4	1	1		9 907 200	2	0	1		1 344	1	2	1		300	0	1	0	
9 092 832	5	0	1		9 292 320	0	0	1		-1 344	2	1	1		-300	1	0	0	
-133 799 040	0	3	1		421 200	0	2	2		-11 904	3	0	1		264	0	1	1	
-138 516 480	1	2	1		-324000	1	1	2		27 360	0	1	1		-264	1	0	1	
138 516 480	2	1	1		421 200	2	0	2		-27 360	1	0	1						
133799040	3	0	1		301 584	0	0	2		34800	0	1	2		-1860	0	4	0	4
-154 141 344	0	1	1		222432	0	0	3		-34800	1	0	2		-5 340	1	3	0	
154141344	1	0	1												-7 548	2	2	0	
3864720	0	3	2		12865608	0	. 2	0	7						-5 340	3	1	0	
-2 482 320	1	2	2		55273752	1	6	0		-616 216	0	6	0	6	-1 860	4	0	0	
2 482 320	2	1	2		88743240	2	5	0		-2 327 664	1	5	0		-9100	0	2	0	
-3864720	3	0	2		46243224	3	.4	0		-3758904	2	4	0		-13580	1	1	0	
7 7 57 904	0	1	2	•	-46243224	4	3	0		-4 106 816	3	3	0		-9 100	2	0	0	
-7 757 904	1	0	2		-88743240	5	2	0		-3758904	4	2	0		-1760	0	0	0	
5 369 952	0	1	3		-55273752	6	1	0		-2 327 664	5	1	0		4536	0	2	1	
-5 369 952	1	0	3		-12 865 608	7	0	0		-616 216	6	0	0		4536	1	1	1	
					259 786 800	0	5	0		-8 482 320	0	4	0		4 5 3 6	2	0	1	
					647 145 360	1	4	0		-18 397 680	1	3	0		9 7 2 0	0	0	1	
					387 343 440	2	3	0		-19 846 320	2	2	0		2 196	0	0	2	

TABLE II. Perturbed energy, orders 1-7. $E^{(N)} = 4^{-N} n^{3N-2} \sum E^{(N)}_{rij} M^r k_1^i k_2^j$.

TABLE III. Odd-order energies with $k_1 - k_2$ factored out, orders 1–7.

						<i>E</i>	'='	4 -/*	$(k_1 - k_1)$	$-k_2$	$n^{3N-2}\sum E_{rij}^{(N)}M^{r}k_{1}^{i}k$	2.				
N	r	i	j	$E_{rij}^{(N)'}$			N	r	i	j	$E_{rij}^{(N)}$		N r	i	j	$E_{rij}^{(N)'}$
1 0	0	0	0	6				1	1	1	13 248		0	1	1	1 230 177 888
								1	0	2	11 904		0	0	2	692 482 344
3 (0	2	0	132				1	0	.0	27 360		. 0	0	0	256412640
C	0	1	1	288				2	0	0	34 800		1	4	0	-9092832
C	0	0	2	132									1	3	1	-36581760
C	0	0	0	300			7	0	6	0	12865608		1	2	2	-55584672
1	1	0	0	264				0	5	1	68 139 360		1	1	3	-36581760
								0	4	2	156882600		1	0	4	-9092832
5 0	0	4	0	31 920				0	3	3	203125824		1	2	0	-133799040
C	0	3	1	127 008				0	2	4	156882600		1	1	1	-272315520
0	0	2	2	189 168				0	1	5	68 139 360		1	0	2	-133799040
0	0	1	3	$127\ 008$				0	0	6	12865608		1	0	0	-154141344
0	0	0	4	31 920				0	4	0	259786800		2	2	0	3864720
0	0	2	0	$277\ 440$				0	3	1	906 932 160		2	1	1	1382400
· 0	0	1	1	542880				0	2	2	1294275600		2	0	2	3864720
0	0	0	2	$277\ 440$				0	1	3	906 932 160		2	0	0	7757904
0	0	0	0	182544				0	0	4	259786800		3	0	0	5369952
1	1	2	0	11 904				0	2	0	692 482 344					

position,

$$E^{(N)} = 4^{-N} (k_1 - k_2) n^{3N-2} \times \sum_{\substack{r \le 0 \ i \le 0}} \sum_{\substack{j \le 0 \ i \le 0 \ 2N-1}} E^{(N)'}_{rij} M^r k_1^i k_2^j \quad (N \text{ odd}) .$$
(80)

To find the numerical values of the polynomial coefficients $E_{rij}^{(N)}$, one must substitute Eqs. (75) and (77) for $\gamma^{(N)}/\gamma^{(0)}$ into Eq. (67) and collect terms proportional to $M^r k_1^i k_2^j$. To evaluate the $E_{rij}^{(N)}$; it is necessary to divide the polynomials $E^{(N)}$ by $k_1 - k_2$. It is straightforward to carry out these calculations on a computer, and we have listed the values of the $E_{rij}^{(N)}$ and the $E_{rij}^{(N)}$ for or-ders 1–7 in Tables II and III.

The number of nonzero $E^{(2N)}$ is $\frac{1}{3}(N+1)(N+\frac{3}{2})(N+2)$, which is also the number of nonzero $E_{rij}^{(2N+1)'}$. The number of nonzero $E_{rij}^{(2N+1)}$ is $\frac{1}{3}(N+1)(N+2)(N+3)$. As these numbers increase rapidly as N increases, we have not included tables for orders higher than N=7. Tables of $E_{rij}^{(N)}$ and $E_{rij}^{(N)'}$ for $N \leq 17$ are available from PAPS.²⁰

VI. CALCULATION OF PERTURBED ENERGIES FOR SPECIFIC STATES

The perturbed energy of a specific state can be calculated by putting the appropriate values of k_1 , k_2 , and M into Eq. (79) to get $E^{(N)}$, and then the desired value of F into Eq. (59) to get E. Given the asymptotic nature of the perturbation series, the question naturally arises how large a field strength can be used. For example, we list in Table IV the termwise contributions to the groundstate energy for fields of 0.02 and 0.10 a.u. At the lower field, the terms decrease in magnitude monotonically, the last term being 10^{-11} a.u. At the higher field, the terms decrease in magnitude until the sixth order, then *increase* monotonically so that the sixteenth-order term is almost 3.5 times the zeroth-order term. The behavior exhibited at the higher field is characteristic of an asymptotic series.

The terms of an asymptotic power series typically tend to decrease in magnitude until some critical term is reached, after which they tend to increase without limit.³⁰ In favorable cases the value of the (N+1)st term is a measure of the error in the Nth partial sum. In the higher-field example of Table IV, we would accordingly regard the fourth partial sum, -0.5281 a.u., as optimum, with an error of the order of 0.0049 a.u. The accurate value of the perturbed energy is reported¹⁰ to be -0.5275 a.u.

The important question from a practical point of view is how accurately the energy can be computed. That is, given a maximum error ϵ and a maximum number of terms available, what is the maximum value of F that can be used in the series, and how many terms should be included. Towards answering these questions, we fix ϵ and consider the sequence of numbers, $F_N(\epsilon)$:

$$F_N(\epsilon) = |\epsilon/E^{(N)}|^{1/N} \quad \text{(for } E^{(N)} \neq 0\text{)}. \tag{81}$$

 $F_N(\epsilon)$ is the value of F for which the Nth term has magnitude ϵ . Intuitively, we expect $F_n(\epsilon)$ to tend first to increase, then decrease as N increases. The maximum value of $F_N(\epsilon)$ can be taken as a "maximum useful field for error ϵ ," and the corresponding order of N (which also depends on ϵ) indicates how many terms should be included, since at this value of F, the magnitudes of the terms preceding and following the Nth are larger than the Nth, which has the magnitude ϵ . If $F_N(\epsilon)$ is still increasing at the last term available, then $F_N(\epsilon)$ may still be regarded as the "maximum useful field for error ϵ ," but with the qualification, "with terms of orders $\leq N$," since if more terms were available, it might be possible to go to higher fields without increasing the error.

TABLE IV. Termwise contributions to the ground-state energy at low and high fields, in a.u.

Order	F=0.	02 a.u.	F=0.1	10 a.u.
Ν	$F^{N}E^{(N)} \times 10^{12}$	Partial sum	$F^N E^{(N)} \times 10^4$	Partial sum
0	-500 000 000 000	-0.500 000 000 000	-5000	-0.5000
2	-900 000 000	-0.500 900 000 000	-225	-0.5225
4	-8 887 500	-0.500 908 887 500	-56	-0.5281
6	-314097	-0.500 909 201 597	-49	-0.5330
8	-20 332	-0.500 909 221 930	-79	-0.5409
10	-1992	-0.500 909 223 922	-195	-0.5604
12	-271	-0.500 909 224 193	-663	-0.6266
14	-49	-0.500909224242	-2992	-0.9259
16	-11	-0.500 909 224 254	-17347	-2.6606

		$F_N(E) = \epsilon/E^{(N)} ^{1/N}$							
			$\epsilon = 0$).5×	•				
Ν	E(W)	10-2	10-4	10-6	10-8				
0	-0.5								
2	-2.25	0.0471	0.00471	0.000471	0.0000471				
4	-55.546875	0.0974	0.0308	0.00974	0.00308				
6	-4907.771484375	0.100	0.0466	0.0216	0.0100				
8	-794 236.926 452 636 718	0.0944	0.0531	0.0298	0.0168				
10	-194 531 960.466 499 329	0.0873	0.0551	0.0348	0.0219				
12	-66 263 036 523.689 170 9	0.0806	0.0549	0.0374	0.0255				
14	-29 924 943 988 411.939 5	0.0747	0.0537	0.0387	0.0278				
16	-17 346 970 495 631 198.5	0.0694	0.0520	0.0390	0.0293				

TABLE V. Perturbed coefficients $E^{(N)}$ and maximum field parameters for the ground state for various error limits, in a.u.

To illustrate the calculation of the maximum useful field, we list in Table V the values of $E^{(N)}$ and $F_N^{(N)}(\epsilon)$ for $|\epsilon/E^{(0)}|=10^{-2}$, 10^{-4} , 10^{-6} , and 10^{-8} for the ground state. Note that the ground state $E^{(N)}$ for orders 2-10, when truncated to ten significant figures, agree exactly with Mendelsohn's.⁷ For an error of 10^{-2} relative to $E^{(0)}$, the maximum $F_N(\epsilon)$ occurs at N=6. Accordingly we would take E only through fourth order, and F not greater than 0.1 a.u. For a relative error of 10^{-4} , the maximum $F_N(\epsilon)$ occurs at N=12. Accordingly we would use the tenth-order partial sum with F not greater than 0.0551 a.u. For relative errors of 10^{-6} or 10^{-8} , we would use all terms available, and we would limit F by 0.0390 a.u. and by 0.0293 a.u., respectively.

In Table VI, we list $F_{17}(10^{-6}E^{(0)})$ or, in the cases for which $n_1 = n_2$, $F_{16}(10^{-6}E^{(0)})$ for a number of states. As with the ground state, $F_N(\epsilon)$ is still increasing at N=17 for relative error 10^{-6} (and 10^{-8}). For values of F less than those in Table VI, the energy through seventeenth order has a relative accuracy of 10^{-6} .

In Table VII, we compare some typical perturbation-theory energies with energies computed nonperturbatively. Not only is the agreement excellent, but we observe that for the ground state the error turns out to be less in magnitude than the imaginary part of the energy corresponding to purely-outgoing-wave boundary conditions,^{10,18,31} while for the excited state listed, the error is approximately one order of magnitude larger than

TABLE VI. Maximum useful field for accuracy 10^{-6} relative to $E^{(0)}$, for E calculated to 17th order for various states.

n	m	n_1	n_2	F(a.u.)	n	т	n_1	n_2	F(a.u.)	n	m	n_1	n_2	F(a.u.)
1	0	0	0	3.90×10 ⁻²		0	3	0	3.06×10^{-4}		1	3	0	1.38×10^{-4}
						1	0	2	3.21×10^{-4}		2	0	2	1.45×10^{-4}
2	0	0	1	$3.76 imes 10^{-3}$		1	1	1	$3.40 imes 10^{-4}$		2	1	1	1.51×10^{-4}
	0	1	0	3.76×10^{-3}		1	2	0	3.21×10^{-4}		2	2	0	1.45×10^{-4}
	1	0	0	$3.95 imes 10^{-3}$		2	0	1	$3.44 imes 10^{-4}$		3	0	1	1.55×10^{-4}
						2	1	0	$3.44 imes 10^{-4}$		3	1	0	1.55×10^{-4}
3	0 -	0	2	8.80×10^{-4}	1	3	0	0	3.55×10^{-4}		4	0	0	1.59×10^{-4}
	0	1	1	9.52×10^{-4}										
	0	2	0	8.80×10^{-4}	5	0	0	4	1.33×10^{-4}	10	0	9	0	$9.29 imes 10^{-6}$
	1	0	1	$9.43 imes 10^{-4}$		0	1	3	1.44×10^{-4}					
	1	1	0	9.43×10^{-4}		0	2	2	1.49×10^{-4}	14	2	0	11	2.56×10^{-6}
	2	0	0	$9.80 imes 10^{-4}$		0	3	1	1.44×10^{-4}					
						0	4	0	1.33×10^{-4}	15	0	14	0	1.89×10^{-6}
1	0	0	3	3.06×10^{-4}		1	0	3	1.38×10^{-4}					
	0	1	2	$3.40 imes 10^{-4}$		1	1	2	1.52×10^{-4}	20	0	19	0	6.07×10^{-7}
	0	2	1	$3.40 imes 10^{-4}$		1	2	1	1.52×10^{-4}	25	0	24	0	2.50×10^{-7}

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<i>F</i> (a.u.)	E(a.u.) calculated by perturbation theory	Order N	Next term ^a $ F^{N+1}E^{(N+1)} $	E(a.u.) calculated nonperturbatively
	C	Ground sta	te	· · · · · ·
0.001	-0.500 002 250 055 55	16	<10-15	-0.500 002 250 055 56 b
0.004	-0.500 036 014 240 15	16	<10-15	-0.5000360142400^{b}
0.005	-0.50005628479379	16	<10-15	-0.500 056 284 793 7 ^b
0.006	-0.50008107221907	16	<10-15	$-0.500\ 081\ 072\ 219\ 4^{b}$
0.007	-0.50011038395008	16	<10-15	-0.500 110 383 958 b
0.008	-0.50014422882008	16	<10-15	-0.500 144 228 821 b
0.009	-0.50018261708612	16	<10-15	-0.500 182 617 088 ^b
0.01	-0.50022556045796	16	<10-15	-0.500 225 560 459 ^b
0.02	-0.500909224254	16	<10-11	-0.500 909 224 258 b
0.03	-0.5020742636	16	<74×10 ⁻¹⁰	-0.502 074 273 b
0.04	-0.50377146	14	$75 imes 10^{-8}$	-0.5037715 ^b
0.05	-0.506 099	10	$16 imes 10^{-6}$	-0.506 1054 ^c
0.06	-0.509 18	8	$12\! imes\!10^{-5}$	-0.509 204 °
0.07	-0.512 94	6	46×10 ⁻⁵	-0.513 08 °
0.08	-0.5167	4	13×10^{-4}	-0.517 56°
0.09	-0.5219	4	26×10^{-4}	-0.5224°
0.10	-0.5281	4	$49 imes 10^{-4}$	-0.527 5°
0.11	-0.527 2	2	81×10^{-4}	-0.5325°
0.12	-0.532	2	$12\! imes\!10^{-3}$	-0.5374°
	Excited state wi	th $n=5$, n	$n = 1, n_1 = 3, n_2 = 0$)
1.556 $ imes 10^{-4}$	-0.01685516	17	$\sim \! 15 \! imes \! 10^{-8}$	-0.0168552372 ^d
1.9448×10^{-4}	-0.016 182 1	12	$57 imes10^{-7}$	-0.016 179 388 5 ^d
2.1393×10^{-4}	-0.015868	10	$18 imes 10^{-6}$	-0.015860468^{d}
$2.5282 imes 10^{-4}$	-0.015 294	6	$79 \! imes \! 10^{-6}$	-0.015269204^{d}
2.9172×10^{-4}	-0.01478	4	$20 imes 10^{-5}$	-0.014740243 ^d
3.3061×10^{-4}	-0.014 30	2	$49 imes 10^{-5}$	-0.014 242 49 ^d

TABLE VII. Energies calculated perturbatively compared with energies calculated nonperturbatively.

^a For the ground state, the next term is $F^{N+2}E^{(N+2)}$.

^bReference 8. We have taken the average of Alexander's upper and lower bounds.

^cReference 10.

^dReference 15.

the imaginary part of the energy.

Finally, in Table VIII we give the perturbed energy for a particular state of the n = 14 level to illustrate the accuracy obtainable (absolute, ~6 $\times 10^{-9}$ a.u.; relative, ~2×10⁻⁶) for a state experimentally observed in sodium, at field strengths at which ionization rates have been measured.¹⁷

TABLE VIII.	Energy	through	17th	order	for n	=14,
$m = 2, n_1 = 0, n_2$	=11.					

F (a.u.)	E (a.u.)	$F^{17}E^{(17)}$
2.4×10^{-6}	-0.0031570578	9 × 10 ⁻¹⁰
$2.5 imes 10^{-6}$	-0.0031854026	17×10^{-10}
$2.6 imes 10^{-6}$	-0.0032141327	33×10^{-10}
$2.7 imes 10^{-6}$	-0.0032432833	$63 imes 10^{-10}$

VII. CONCLUDING REMARKS

It has been shown how to generate the explicit formulas for the perturbed separation constant and energy in the Stark effect in hydrogen as polynomials in the quantum numbers, for arbitrarily high orders, and the usefulness of these formulas has been demonstrated by numerical examples. We conclude with some practical observations.

The number of nonzero polynomial coefficients in $\beta^{(N)}$ and $E^{(N)}$ increases as $\frac{1}{2}N^2$ and $\frac{1}{3}N^3$, respectively. The total number of $c_{irs}^{(N)}$ for orders up to and including N is roughly of order N⁴. As N increases, the large number of quantities becomes troublesome. The maximum value of N=17 that appears in this paper represents the largest calculation that fit in a computer memory of 256k words without extensive effort to conserve storage, without explicit overlaying, and without using disk files for intermediate results.

If one is interested in only a few states, to generate the $\beta^{(N)}$ and $E^{(N)}$ as polynomials for very high N would not appear to be as efficient as to generate the numerical values directly from Eqs. (32), (33), (63), and (67). Compared to the effort required to program the polynomial solution, it is rather easy to program Eqs. (32) and (33) to get numerical values of the $\beta^{(N)}$ for specific states. The general formula (67) then produces the $E^{(N)}$ from the $\beta^{(N)}$ with an even more modest effort. Such a computer program has been written to give $\beta^{(N)}$ and $E^{(N)}$ through N=25.

For some practical purposes, the energies computed by perturbation theory are probably of sufficiently high accuracy, since in some cases the errors inherent in the use of the asymptotic series comparable with the widths of the levels, which are nonzero because of field ionization. One is not normally interested in the energy of a state to an accuracy too much finer than its width.

Finally, we note that the perturbation series is much easier to use than any of the nonperturbative schemes.

Note added in proof. Precise measurements of intense-field Stark shifts in highly excited states of hydrogen have now been reported by Peter M. Koch [Phys. Rev. Lett. (to be published)]. In the two cases analyzed in detail, the discrepancy between perturbation theory and experiment is less than the experimental uncertainty for a large $n_2 - n_1$ state (n = 30, $n_1 = 0$, $n_2 = 29$, m = 0), but more than an order of magnitude larger than the uncertainty for a large $n_1 - n_2$ state (n = 25, $n_1 = 21$, $n_2 = 2$, |m| = 1). In the latter case, successive partial sums alternately bracket the experimental value from above and below. The author is grateful to Professor Koch for a preprint of his paper and for helpful discussions.

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- ¹E. Schrödinger, Ann. Phys. (Leipz.) <u>80</u>, 437 (1926). ²Y. Ishida and S. Hiyama, Sci. Pap. Inst. Phys. Chem.
- Res. (Jpn.) 9, 1 (1928).
- ³K. Basu, Bull. Calcutta Math. Soc. 26, 79 (1934).
- ⁴J. D. Bekenstein and J. B. Krieger, Phys. Rev. <u>188</u>, 130 (1969).
- ⁵S. P. Alliluev and I. A. Malkin, Zh. Eksp. Teor. Fiz. 66, 1283 (1974) [Sov. Phys. JETP 39, 627 (1974)].
- ⁶G. L. Sewell, Proc. Cambridge Philos. Soc. <u>45</u>, 678 (1949).
- ⁷L. B. Mendelsohn, Phys. Rev. 176, 90 (1968).
- ⁸M. A. Alexander, Phys. Rev. 178, 34 (1969).
- ⁹J. O. Hirschfelder and L. A. Curtiss, J. Chem. Phys. 55, 1395 (1971).
- ¹⁰M. Hehenberger, H. V. McIntosh, and E. Brändas, Phys. Rev. A <u>10</u>, 1494 (1974).
- ¹¹N. A. Gushina and V. K. Nikulin, Chem. Phys. <u>10</u>, 23 (1975).
- ¹²P. Froelich and E. Brändas, Phys. Rev. A <u>12</u>, 1 (1975).
- ¹³P. Froelich and E. Brändas, Int. J. Quantum Chem. Symp. 10, 353 (1976).
- ¹⁴W. P. Reinhardt, Int. J. Quantum Chem. Symp. <u>10</u>, 359 (1976).
- ¹⁵R. J. Damburg and V. V. Kolosov, J. Phys. B <u>9</u>, 3149 (1976).
- ¹⁶D. R. Herrick, J. Chem. Phys. 65, 3529 (1976).
- ¹⁷M. G. Littman, M. L. Zimmerman, and D. Kleppner, Phys. Rev. Lett. <u>37</u>, 486 (1976), and references cited therein.
- ¹⁸T. Yamabe, A. Tachibana, and H. J. Silverstone, Phys. Rev. A 16, 877 (1977).
- ¹⁹B. M. Smirnov and M. I. Chibisov, Zh. Eksp. Teor.

Fiz. <u>49</u>, 841 (1965) [Sov. Phys. JETP <u>22</u>, 585 (1966). ²⁰See, AIP document No. PAPS PLRAA-18-1853-25 for 12 pages containing tables of $\beta_{rs}^{(N)}$, $E_{rk}^{(N)}$, and $E_{rk}^{(N)}$, for N=1 through 17. Order by PAPS number and journal reference from American Institute of Physics, Physics Auxiliary Publication Service, 335 East 45th Street, New York, N.Y.10017. The price is \$1.50 for each microfiche (98 pages), or \$5 for photocopies of up to 30 pages with \$0.15 for each additional page over 30 pages. Airmail additional. Make checks payable to the American Institute of Physics. This material also appears in *Current Physics Microform*, the monthly microfilm edition of the complete set of journals published by AIP, on the frames immediately following this journal article.

- ²¹E. C. Titchmarsh, Eigenfunction ExpansionsAssociated with Second-Order Differential Equations, 2nd ed. (Oxford U.P., London, 1958), Chap. 20.
- ²²See, for instance, L. D. Landau and E.M. Lifshitz, *Quantum Mechanics* (Addison-Wesley, Reading, Mass., 1958).
- 23 The notation is that of Ref. 18.
- ²⁴The quantity $\frac{1}{4}(1-m^2)$ is the eigenvalue of the Casimir operator for the group O(2,1), whose relevance to the Stark effect in hydrogen was exploited by Alliluev and Malkin (Ref. 5).
- ²⁵There are unfortunately several different notations in use. We follow Ref. 21. The equation $L_n^m(x)$
- = $(d/dx)^m L_n(x)$ may aid in identifying the notation. In terms of the usage in Abramowitz and Stegun (Ref. 26) $L_{n+m}^m(x)$ (here) = $(-1)^n (n+m)! L_n^{(m)}(x)$ (Abramowitz and Stegun).
- ²⁶M. Abramowitz and I. A. Stegun, Handbook of Mathe-

matical Functions, U. S. Natl. Bur. Stand. Appl. Math. Ser. 55 (U.S. GPO, Washington, D.C., 1964).

²⁷Cf. Eqs. (26) of Ref. 5.

- ²⁸See, for instance, Eq. (3.6.7) of Ref. 26.
- ²⁹See, for instance, Sec. 24.2 and Table 24.2 of Ref. 26.

³⁰See, for instance, A. Erdélyi, Asymptotic Expansions (Dover, New York, 1956).

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