# Radiative collision-induced photoionization

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Semiclassical expressions of two-photon ionization of atoms induced by radiative collisions are derived. The dependence of the ionization yield on the atomic forces, field intensity, and energy gap is derived. Although absorption tends to decrease as the field intensity rises due to stimulated emission at the second crossing, the two-photon ionization yield can be nearly saturated at the first crossing, thus enhancing the absorption. Both regions, ionization in single collisions and ionization between collisions, are treated. In the latter we find that saturation of the ionization can be achieved at much reduced intensities. This process promises an extremely sensitive method for studying radiative collisions, especially when absorption or fluorescence becomes extremely weak.

### I. INTRODUCTION

Atomic collisions in the presence of electromagnetic fields are continuing to receive attention theoretically and experimentally<sup>1-24</sup> because of their relevance to laser fusion and isotope separation and to numerous other interesting applications which include development of radiative-collision lasers,<sup>11</sup> up conversion of long-wavelength radiation, increasing the reaction rates of selected gas-phase chemical reactions,<sup>18</sup> extracting potential curves of the interacting system,<sup>16</sup> and studying spin and charge exchange.<sup>15,18</sup> Recently we extended the optical-collision studies by studying two-photon ionization of the colliding atoms.<sup>19</sup> Because complete conversion of every absorption event to an ion pair is possible,<sup>23</sup> the ionization channel promises extreme sensitivity since it is easier to detect a small number of electrons than a small number of photons. This allows measurement at the extreme wing where absorption or fluorescence becomes vanishingly small.<sup>24</sup> Even single absorption events can be detected with ease by gas proportional counters which are sensitive to single electrons.23

In this work, we study some aspects of atomatom (-ion) collisions in the presence of electromagnetic fields by studying two-photon ionization of atoms assisted by radiative collision. The process involves an atom in an excited state colliding with another atom (ion). We assume that, in absence of the collision, the two-photon ionization has a vanishingly small cross section because of the absence of resonant intermediate states or because of a high ionization potential. In the presence of the collision and when the photon energy plus the energy of the excited state of the atom matches an energy level of the atom, resonant collisional transfer takes place, thus providing a real resonant intermediate level for the laser field interaction with the system. These systems can be chosen such that one more photon absorbed by the system can further ionize it. Ionizing the atom-atom system, therefore, can be sensitive in monitoring near resonant collisional transfer, especially when the yield becomes vanishingly small.

We derive the ionization-yield dependence on the interaction forces between the atoms, as well as on the excitation intensity and the energy gap. Although the absorption of the system might normally decrease as the intensity of excitation increases due to stimulated emission at the second crossing, the two-photon ionization yield can be nearly saturated at the first crossing, thus enhancing the absorption. When the absorption is weak and the ionization takes place between collisions, saturation of the ionization can be achieved at moderate excitation intensities. The far-wing wavelength dependence of the ionization yield induced by collisional transfer falls off with the square root of the detuning, while the corresponding effect in charge exchange drops faster and is more complicated since it is due to overlap.

This paper consists of four sections. In Sec. II, we give the formulation and the solution for the absorption probabilities. Section III treats the ionization yield, and, in Sec. IV, we give some concluding remarks.

### **II. ABSORPTION PROBABILITY**

We consider an atom-atom (-ion) collision in the presence of an electromagnetic field. One atom A is assumed to be in an excited state, while the other atom (ion) B is in its ground state. In the case where the process can occur at long range with no overlap taking place, an atomic basis of the process can be used. However, in the case

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where overlap is an integral part of the interaction, the process should be treated from the viewpoint of quasimolecular absorption. In any case, when radiative collisions are considered, where one of the colliding partners is in an excited state, the effective interaction Hamiltonian depends on both the collisional coupling and the electromagnetic field coupling. This effective Hamiltonian can be best derived taking an atomic viewpoint of the interaction. Therefore, we will consider the atomic viewpoint first.

To solve for the two-photon ionization of the system, we start with the Schrödinger equation in the dipole classical-field approximation:

$$i\hbar \frac{\partial}{\partial t}\psi = H\psi, \qquad (1)$$

$$H = H_A + H_B + H_{AB} - \vec{\mu}_A \cdot \vec{\epsilon}, \qquad (2)$$

where  $H_A$  and  $H_B$  are the electronic Hamiltonians of the isolated atoms;  $H_{AB}$  is the atom-atom interaction;  $\mu_A$  and  $\mu_B$  are the dipole moments of atoms A and B, respectively; and  $\epsilon$  is the classical electromagnetic field  $\vec{E} \cos \omega t$  with frequency  $\omega$ . In writing  $H_{AB}$ , the impact-parameter method will be used where the internuclear coordinates are treated classically and without acceleration in the motion. That is,  $\vec{R} = \vec{b} + \vec{V}_t$  and  $\vec{b} \cdot \vec{V} = 0$ , where  $\vec{b}$  is the impact parameter and  $\vec{V}$  the relative velocity of the nuclei. However, deviation from straight-line trajectories will be considered in the case of charge transfer at slow velocities.

We assume that the dipole-moment matrix elements between  $|2a\rangle$  and  $|0a\rangle$  and between  $|1b\rangle$  and



FIG. 1. Energy level diagram of the colliding system.

 $|0b\rangle$  are oriented along the lines joining the two atoms (rotating-atom approximation). The motive for using the rotating-atom approximation is that it eliminates the magnetic sublevels and thus simplifies the calculation. However, it is only good to a factor of 2. This approximation has been previously used in dealing with collisions in line broadening<sup>25</sup> (50% error), in Penning ionization by atoms in resonance states<sup>26</sup> (30% error), and in several other types of collisions.<sup>27</sup> Therefore, the system is represented by field-free electronic states, and we consider only three product discrete states and an infinite set of continuum states (Fig. 1). The wave function of the system can then be written as follows:

$$\begin{split} \psi &= a_0 \exp\left[-i\left(\omega_{A1} + \omega_{B0}\right)t\right] \left|1a\right\rangle \left|0b\right\rangle + a_1 \exp\left[-i\left(\omega_{A2} + \omega_{B0}\right)t\right] \left|2a\right\rangle \left|0b\right\rangle \\ &+ a_2 \exp\left[-i\left(\omega_{A0} + \omega_{B1}\right)t\right] \left|0a\right\rangle \left|1b\right\rangle + \int a_c \exp\left[-i\left(\omega_{A0} + \omega_{Bc}\right)t\right] \left|0a\right\rangle \left|cb\right\rangle d\hbar\omega_{Bc}, \end{split}$$

where  $\omega_{xi}$  is the absolute frequency of level  $|ix\rangle$ with i=0,1,2, or c; and x=A or B. The product state  $|1a\rangle |0b\rangle$  is the ground state of the system where atom A is an initial excited state  $|1a\rangle$  while the atom *B* is in its ground state  $|0b\rangle$ . State  $|2a\rangle |0b\rangle$ is reached after the system absorbs one photon, and thus atom A is excited to  $|2a\rangle$  state while atom B is still in its ground state  $|0b\rangle$ . After the collision takes place and the transfer occurs, the system evolves to the product state  $|0a\rangle |1b\rangle$  where atom A is in the ground state and atom B is in an excited state  $|1b\rangle$ . The product  $|0a\rangle |cb\rangle$  describes the photoionization of the system after the transfer process has occurred. State  $\ket{cb}$  is a continuum state of the atom B. An infinite set of continuum states has been included by performing the integration over the energy of the photoelectron. Timedependent amplitudes of the discrete states and continuum states are represented by  $a_m$  and  $a_c$ , respectively.

To calculate the two-photon ionization amplitude, one has to solve the equations of motion (1)-(3)with the initial condition  $a_0(-\infty)=1$ ,  $a_1(-\infty)=a_2(-\infty)$  $=a_c(-\infty)=0$ . Substituting Eq. (3) in Eq. (1) and using the rotating-wave approximation, one gets the following set of equations for the amplitudes  $a_m$  and  $a_c$  when the field is linearly polarized:

$$\frac{d}{dt}a_0 = i\mu_A E e^{-i\Delta_1 t} a_1, \tag{4}$$

$$\frac{d}{dt}a_1 = i\mu_A^* E e^{i\Delta_1 t} a_0 + ive^{i\Delta_2 t} a_2, \tag{5}$$

(3)

$$\frac{d}{dt}a_{2} = iv^{*}e^{-i\Delta_{2}t}a_{1} + i\int\mu_{B}Ee^{-\Delta_{c}t}a_{c}d\hbar\omega_{Bc}, \qquad (6)$$

$$\frac{d}{dt}a_{c} = i\mu_{B}^{*}Ee^{i\Delta_{c}t}a_{2}, \qquad (7)$$

where  $\mu_A$  is the matrix element in units of  $\hbar$ , of the atomic dipole operator between states  $|1a\rangle$  and  $|2a\rangle$ ; v is the matrix element, in units of  $\hbar$ , of the interaction Hamiltonian  $H_{AB}$  between states  $|2a\rangle$ and  $|1b\rangle$ ;  $\Delta_c = \omega_c - \omega$ ;  $\Delta_1 = \omega_0 - \omega$ ;  $\Delta_2 = \omega_1 - \omega_2$ ;  $\mu_B$  is the matrix element, in units of  $\hbar$ , of the dipole operator between the  $|1b\rangle$  state and the continuum state  $|cb\rangle$ , where  $\omega_i$  are given in Fig. 1. In writing Eqs. (4)-(7), we have neglected scattering between the continuum states by neglecting the coupling between them in Eq. (7).

In order to simplify the calculation, we separate the continuum states into two sets; the first set is close to the resonance condition  $|\Delta_c| \leq |\Delta_c^0|$ , where  $|\Delta_c^0|$  is a small quantity, and the second set is that which is sufficiently far away from resonance  $|\Delta_c| \geq |\Delta_c^0|$ . When  $\Delta_c$  is large, energy is not conserved as a result of nonresonant continuum states; consequently, intensity-dependent frequency shifts are introduced in level  $|1b\rangle$ . The near-resonant continuum states, however, result in intensitydependent sink, or decay to level  $|1b\rangle$ . When  $\Delta_c$  is large, the right side of Eq. (7) oscillates swiftly; and if the field is a slowly varying function of time, one can solve Eq. (7) for  $a_c$  by integrating over time by parts:

$$a_c = \left(\mu_B^* E / \Delta_c\right) e^{i\Delta_c t} a_2. \tag{8}$$

However, when near-resonant states are considered, i.e.,  $\Delta_c$  is small,  $a_c$  can only be represented as a full integral:

$$a_c = i \,\mu_B^* \int E e^{i\Delta_c t} a_2(t) \,dt \,. \tag{9}$$

Substituting expression (8), which is valid for  $|\Delta_c| \ge |\Delta_c^0|$ , in Eq. (6) results in an expression for the shift as an integral over  $\Delta_c$ . Substituting expression (9), which is valid for  $|\Delta_c| \le |\Delta_c^0|$ , in Eq. (6) and carrying out the integration over  $\Delta_c$  first with the assumption that  $\mu_B$  is fairly constant for  $|\Delta_c| \le |\Delta_c^0|$  enables one to easily carry out the integration over time:

$$\frac{da_0}{dt} = i\,\mu_A E e^{-i\Delta_1 t} a_1,\tag{10}$$

$$\frac{da_1}{dt} = i\,\mu_A^* E e^{i\,\Delta_1 t} a_0 + iv e^{i\,\Delta_2 t} a_2,\tag{11}$$

$$\frac{da_2}{dt} + (\Gamma + is)a_2 = iv * e^{-i\Delta_2 t}a_1, \qquad (12)$$

where

$$\Gamma = \gamma_1(t) + \gamma_0, \tag{13}$$

 $\gamma_0$  is a spontaneous-emission decay rate included phenomenologically,

$$s = -E^2 P \int \frac{\mu_B^2}{\Delta_c} dh \omega_{Bc} , \qquad (14)$$

$$\gamma_1 = h\pi E^2 \mu_B^2, \tag{15}$$

and *P* stands for principal value. The above scheme of eliminating the continuum states was also used in photoionization of isolated<sup>28</sup> and colliding atoms.<sup>19</sup> We consider cases where the laser is tuned to near resonance, i.e.,  $\Delta_1 \sim \Delta_2$ , and write this as  $\Delta_1 = \Delta_2 + \delta'$ ,  $\Delta_2 = \Delta$ , where  $\delta' \ll \Delta_2$ . Also, let us define  $a_0 = \overline{A}_0$ ,  $a_1 = \overline{A}_1 \exp[i(\Delta_2 + \delta')t]$ , and  $a_2 = \overline{A}_2 e^{i\delta't}$ . Equations (10)–(12) reduce to

$$\frac{d\overline{A}_{0}}{dt} = i\,\mu_{A}E\overline{A}_{1},\tag{16}$$

$$\frac{d\overline{A}_{1}}{dt} + i(\Delta + \delta')\overline{A}_{1} = i\mu_{A}^{*}E\overline{A}_{0} + iv\overline{A}_{2}, \qquad (17)$$

$$\frac{d\overline{A}_2}{dt} + i(\delta' + s - i\Gamma)\overline{A}_2 = iv *\overline{A}_1.$$
(18)

The interesting situation is when  $\hbar \Delta \gg kT$  so that thermal collisional transfer does not take place. Also, to prevent thermal transfer,  $\mu_A E$  has to be much smaller than  $\hbar \Delta$  so that power broadening will not bridge the gap. Therefore, in this situation  $\overline{A}_1$  is sufficiently far from resonance that  $d\overline{A}_1/dt$  can be neglected relative to  $\Delta \overline{A}_1$  in Eq. (17), thus, one can solve for  $\overline{A}_1$  in terms of  $\overline{A}_0$ and  $\overline{A}_2$ . This procedure of eliminating a nonresonant intermediate state results in frequency shift of the effective collisionally induced photoabsorption process

$$\overline{A}_{1} = \frac{\mu_{A}^{*}E}{\Delta + \delta'}\overline{A}_{0} + \frac{v}{\Delta + \delta'}\overline{A}_{2}.$$
(19)

Substituting Eq. (19) in Eqs. (16) and (18), we get

$$\left[\frac{d}{dt} - \frac{i\,\mu_A^2 E^2}{\Delta + \,\delta'}\right] \overline{A}_0 = \frac{i\,\mu_A E v_2}{\Delta + \,\delta'} \overline{A}_2,\tag{20}$$

$$\left[\frac{d}{dt}+i\left(\delta'+s-i\Gamma-\frac{v_1^2}{\Delta+\delta'}\right)\right]\overline{A}_2 = \frac{i\mu_A^*Ev_2}{\Delta+\delta'}\overline{A}_0.$$
 (21)

Note that we labeled the matrix elements v by  $v_1$ and  $v_2$  in order to allow for the possibility of overlap in the interaction, as we will explain in more detail below. Equations (4)–(7), which describe the two-photon ionization process assisted by collisional transfer, are reduced to two-state problems with frequency shifts caused by both the electric field and by the collisional coupling. The intermediate nonresonant state  $|2a\rangle|0b\rangle$  causes a shift in the ground state  $|1a\rangle|0b\rangle$ . The continuum state causes a complex shift  $\gamma_1 + is$  in the collisionally populated state  $|0a\rangle|1b\rangle$ . The real part  $\gamma_1$  is an intensity-induced decay due to photoionization, and the imaginary part s is an intensity-induced frequency shift due to the virtual transitions with the continuum states. The two-state system is coupled by an effective interaction potential which is the product of the laser field and the collision interaction potential, and scaled down by the frequency gap  $\Delta$ . With a phase transformation

$$\overline{A}_{0} = \exp\left[\frac{i\mu_{A}^{2}}{\Delta + \delta'}\int E^{2} dt\right]A_{0},$$

$$\overline{A}_{2} = \exp\left[\frac{i\mu_{A}^{2}}{\Delta + \delta'}\int E^{2} dt\right]A_{2},$$
(22)

Eqs. (20) and (21) become

$$\frac{d}{dt}A_0 = i\,\alpha\,A_2,\tag{23}$$

$$\frac{d}{dt}A_2 + (i\delta + \Gamma)A_2 = i\alpha^*A_0, \qquad (24)$$

where

$$\delta(t) = \left(\delta' + s + \frac{\mu_A^2 E^2}{\Delta + \delta'} - \frac{v_1^2}{\Delta + \delta'}\right),\tag{25}$$

$$\alpha = \frac{\mu_A E v_2}{\Delta + \delta'}.$$
 (26)

Note that when the process occurs at long range, where there is no overlap,  $v_1 = v_2$ . This arises in atom-atom collisions. However, when overlap is an integral part of the interaction, as is the case in charge-exchange collisions between an ion and an atom,  $v_1 \neq v_2$ . Below we discuss both cases. The classical Hamiltonian for the interaction of a singly charged ion and a neutral atom is given by  $H' = e^2 r/R^2$ , where r is the local coordinate of the electron and R is the internuclear separation. To calculate the energy shift, one uses an atomic basis set which for a two-level atom and to second order gives an interaction shift  $v_1^2 = -c_4/R^4$ . The classical interaction Hamiltonian between two neutral atoms when dipole-dipole or Van der Waals interaction is dominant is  $e^2 x_1 x_2 / R^3$ , where  $x_1$  and  $x_2$  are the local coordinates of the electrons of each atom and R is the internuclear separation. To second order, this Hamiltonian yields a shift  $v_1^2 = -c_6/R^6$ , if the two atoms are different, as here, where  $c_6$  is the Van der Waals interaction constant.

Let us consider the form of the interaction coupling  $v_2$ . For the case of the interaction of a hydrogen atom and a proton, the interaction potential can be explicitly evaluated. The long-range interaction takes the form

$$v_2 = -\frac{e^2}{\hbar a_0} \frac{R}{a_0} e^{-R/a_0},$$
 (27)

where e is the electronic charge,  $a_0$  is the Bohr radius, and R is the internuclear separation. For a single electron charge-exchange collision between different partners, Olson *et al.*<sup>29</sup> showed that the interaction potential can be approximated by a similar formula with an effective internuclear separation  $R_0$  and strength

$$v_{2} = -\frac{G}{\hbar} \frac{R_{0}}{a_{0}} \exp\left(\frac{-0.86R_{0}}{a_{0}}\right).$$
(28)

The above form is expected to hold when the colliding particles are sufficiently far apart and when the wave functions are nearly hydrogenic. The strength G depends on the ionization potential of the respective atoms:

$$G = I_{H} (I_{1} I_{2} / I_{H}^{2})^{1/2} \left[ (I_{1} / I_{H})^{1/2} + (I_{2} / I_{H})^{1/2} \right],$$
(29)

where  $I_H = e^2/2a_0$  is the ionization potential of hydrogen and  $I_1$  and  $I_2$  are the ionization potentials of the colliding atoms.

When the charge transfer takes place in an intense field, the form of the interaction Hamiltonian can be derived as discussed above, using atomic basis. The result is the product  $\mu_A E v_2/(\Delta + \delta')$ ; therefore, this product, when  $v_2$  is substituted, becomes

$$H_{\rm in} = \left(\frac{\mu_A E}{\Delta + \delta'}\right) \frac{G}{\hbar} \frac{R_0}{a_0} \exp\left(\frac{-0.86R_0}{a_0}\right),\tag{30}$$

where  $R_0$  is the internuclear separation where absorption takes place.

Equations (20) and (21) with  $v_1 \neq v_2$  describe many physical problems. These include dipole-dipole, dipole-quadrupole, quadrupole-quadrupole, charge transfer, S-S transition, and spin change. In the case where overlap is important (as, for example, charge exchange), one can use the quasimolecular absorption viewpoint, as discussed above. For completeness, we describe briefly the two-photoionization process induced by the charge-transfer analog of radiative collision, using the quasimolecular picture.

Let us consider the time-dependent, electronicfield Schrödinger equation:

$$i\frac{\hbar\partial}{\partial t}\psi(\vec{\mathbf{x}},\vec{\mathbf{R}}) = (H_{e1} + H_{in})\psi(\vec{\mathbf{x}},\vec{\mathbf{R}}), \qquad (31)$$

$$H_{e1} = T_x + V_e(\vec{\mathbf{x}}, \vec{\mathbf{R}}), \qquad (32)$$

where  $\vec{\mathbf{x}}$  and  $\vec{\mathbf{R}}$  stand for electronic and nuclear coordinates, respectively;  $T_x$  is the electronic kinetic energy operator;  $v_e(\vec{\mathbf{x}}, \vec{\mathbf{R}})$  is the electrostatic interaction among electrons and nuclei; and  $H_{in}$  is

the classical interaction potential between the molecular system and the laser field.

Since the effective interaction potential  $H_{in}$  is derived, one can proceed to treat the two-photon process from the quasimolecular viewpoint. We represent the system by two field-free adiabatic discrete electronic states  $\phi_i(\vec{\mathbf{x}}, \vec{\mathbf{R}})$  and a continuum  $\phi_c$ . We assume  $\langle \phi_i | \phi_j \rangle = \delta_{ij}$ , where i, j = 1, 2, or c. We will not attempt to include the magnetic sublevels in this study in order to simplify the derivation. Calculations of line broadening where anisotropy of the system was included indicated that the two-level calculation is only 50% off from the complete calculation.<sup>22-24</sup> Therefore,

$$\psi(\mathbf{\tilde{x}}, \mathbf{\tilde{R}}) = a_1(t)\phi_1(\mathbf{\tilde{x}}, \mathbf{\tilde{R}}) \exp\left(-\frac{i}{\hbar} \int_{t_1}^t w_1(\mathbf{\tilde{R}}) dt\right) + a_2(t)\phi_2(\mathbf{\tilde{x}}, \mathbf{\tilde{R}}) \exp\left(-\frac{i}{\hbar} \int_{t_1}^t w_2(\mathbf{\tilde{R}}) dt\right) + \int a_c(w_c, t)\phi_c(\mathbf{\tilde{x}}, \mathbf{\tilde{R}}) \exp\left(-\frac{i}{\hbar} \int_{t_1}^t w_c(\mathbf{\tilde{R}}) dt\right) dw_c,$$
(33)

where  $\phi_i$  are the field-free adiabatic discrete electronic states,  $w_i$  are the corresponding fieldfree adiabatic potential energy surfaces,  $\phi_c$  are continuum states with corresponding  $w_c$  energies,  $t_1$  is some initial time, and  $a_i$  and  $a_c$  are time-dependent amplitudes. Explicit forms for  $a_1$ ,  $a_2$ , and  $a_c$  are found by solving the equation of motion (26) given the initial conditions  $a_1(-\infty) = 1$  and  $a_2(-\infty)$  $= a_c(-\infty) = 0$ .

In the case where the system is electronically adiabatic in the absence of the field, we drop terms containing the nonadiabatic coupling  $\langle \phi_1 | \dot{\phi}_2 \rangle$ ,  $\langle \phi_2 | \dot{\phi}_1 \rangle$ ,  $\langle \phi_2 | \dot{\phi}_o \rangle$ , and  $\langle \phi_c | \dot{\phi}_2 \rangle$ ; and in the rotatingwave approximation Eqs. (31)–(33) reduce to the following set of equations:

$$\begin{aligned} \frac{da_1}{dt} &= i \,\alpha \exp\left(-i \int_{t_1}^t \tilde{\delta}(t') \,dt'\right) a_2, \end{aligned} \tag{34} \\ \frac{da_2}{dt} &+ \gamma_0 a_2 = i \,\alpha^* \exp\left(i \int_{t_1}^t \tilde{\delta}(t') \,dt'\right) a_1 \\ &+ i E \int \mu_B \exp\left(-i \int_{t_1}^t \delta_c(t') \,dt'\right) a_c \,dw_{Bc}, \end{aligned} \tag{35}$$

$$\frac{da_c}{dt} = i\,\mu_B^* E\,\exp\left(i\int_{t_1}^t \delta_c(t\,\prime)\,dt\,\prime\right)a_2 \ , \tag{36}$$

where

$$\tilde{\delta}(t) = (1/\hbar)(w_2 - w_1 - \hbar\omega), \qquad (37)$$

$$\delta_c(t) = (1/\hbar)(w_c - w_2 - \hbar\omega), \qquad (38)$$

$$\alpha = \langle \mathbf{1} | H_{\mathbf{in}} | 2 \rangle = \frac{\mu_A E}{(\Delta + \delta')} \frac{G}{\hbar} \frac{R^*}{a_0} e^{-\mathbf{0} \cdot \mathbf{3} \mathbf{6} R^* / a_0}.$$
(39)

Note that we used the matrix elements  $\mu_A$  and  $\mu_B$  which we used in the atomic viewpoint.

We eliminate the continuum states using the procedure described above in the atom-atom collision case. Equations (34)-(36) reduce to *cut* 

 $\frac{da_1}{dt} = L_1 a_2,$ 

 $\frac{da_2}{dt} + \tilde{\Gamma}(t)a_2 = L_2 a_1,$ 

where

$$\tilde{\Gamma}(t) = \gamma_0 + is + \gamma_1, \tag{42}$$

$$\Gamma(t) = \gamma_0 + \gamma_1$$

$$L_{1} = i \alpha \exp\left(-i \int \tilde{\delta}(t') dt'\right), \qquad (43)$$

$$L_2 = i \alpha^* \exp\left(i \int \tilde{\delta}(t') dt'\right), \qquad (44)$$

$$s = -E^2 P \int \frac{\mu_B^2}{\delta_c} dw_c \tag{45}$$

$$\gamma_1 = \hbar \pi \,\mu_B^2 E^2,\tag{46}$$

and P stands for principal value.

Note that with a phase transformation of the amplitude  $a_2 = A_2 \exp(i \int \delta dt)$ , Eqs. (40) and (41) become identical in form to Eqs. (23) and (24) derived using the atomic viewpoint. The two sets become identical if  $\delta + s$  in Eqs. (40) and (41) is the same as  $\delta$  in Eq. (25). This implies that the difference frequency with all the interactions taken into account is

$$s + \mu_A^2 E^2/(\Delta + \delta') - v^2/(\Delta + \delta').$$

This includes Stark shifts due to nonresonant continuum and discrete states.

The above system of two coupled equations with time-varying amplitudes arises in many physical situations such as collision phenomena, laser interactions with isolated systems, and collisions in the presence of laser fields.

We now proceed to solve Eqs. (40) and (41) for  $a_2$ , from which expressions for the photoionization can be calculated. Since changes in the amplitudes can only occur at the resonance condition

(40)

(41)

 $\delta(t_0) = 0$ , we consider times near  $t_0$  and define the quantity  $A_2$ :

$$A_{2} = \exp\left(\int_{0}^{\tau} \frac{\tilde{\Gamma}(t)}{2} dt\right)$$
$$\times \exp\left(-\frac{1}{2}\int_{0}^{\tau} \frac{d}{dt} \ln L_{2} dt\right) a_{2}, \qquad (47)$$

where  $\tau = t - t_0$  and  $t_0$  is the time where resonance occurs,  $\delta(t_0) = 0$ . Substituting Eq. (47) in Eqs. (40) and (41), we obtain

$$\left(\frac{d^2}{d\tau^2} + f(\tau)\right)A_2 = 0, \qquad (48)$$

where

$$f(t) = -L_1 L_2 - \frac{1}{4} \left( \tilde{\Gamma} + \frac{d}{dt} \ln L_2 \right)^2 + \frac{1}{2} \frac{d}{dt} \tilde{\Gamma} + \frac{1}{2} \frac{d^2}{dt^2} \ln L_2.$$
(49)

We proceed to solve Eq. (48) by the Landau-Zener method near the resonance time  $t_0$ . We first write  $f(\tau)$  explicitly in terms of the field intensity, detuning, and decay:

$$f(t) = \alpha^{2} + \frac{1}{4} \left( \delta - i \frac{d}{d\tau} \ln \alpha - i\Gamma \right)^{2} + \frac{1}{2} \frac{d\Gamma}{d\tau} + \frac{1}{2} \left( \frac{d^{2}}{d\tau^{2}} \ln \alpha + i \frac{d\delta}{d\tau} \right),$$
(50)

where we have defined  $\delta = \tilde{\delta} + s$ . Close to  $t = t_0$  or  $\tau = 0$ , one can expand  $\delta(\tau)$  as  $\tau d\delta(0)/d\tau$  and write

$$f(t) = g_0 + \frac{1}{2}ig_1 + \frac{g_1^2}{4} \left(\tau - \frac{i}{g_1}\frac{d}{d\tau}\ln\alpha - \frac{i\Gamma}{g_1}\right)^2, \quad (51)$$

$$g_0 = \alpha^2 + \frac{1}{2} \frac{d}{d\tau} \Gamma + \frac{1}{2} \frac{d^2}{d\tau^2} \ln \alpha, \qquad (52)$$

$$g_1 = \frac{d\delta(0)}{d\tau} \,. \tag{53}$$

Substituting Eqs. (51)-(53) in Eq. (48), and with the definitions of z and n,

$$z = \left(\tau - \frac{i}{g_1} \frac{d}{d\tau} \ln \alpha - \frac{i\Gamma}{g_1}\right) g_1^{1/2} e^{-i\pi/4}, \tag{54}$$

$$n = \frac{ig_0}{g_1} = g + ip , \qquad (55)$$

where q and p are real, Eq. (48) reduces to the following equation:

$$\left(\frac{d^2}{dz^2} + n + \frac{1}{2} - \frac{z^2}{4}\right) A_2 = 0.$$
 (56)

Equation (56) reduces to Weber equation if n is time independent. For constant pulse amplitude and slowly moving particles, this condition can be assumed; therefore, one can easily show that the asymptotic probability

$$|a_2|^2 = (1 - e^{-2\pi p}) e^{-\Gamma \tau}, \tag{57}$$

where

$$p = \frac{\alpha^2}{d\delta(0)/d\tau}.$$
(58)

Note that the absorption probability  $1 - e^{-2\pi p} de_{-}$ pends exponentially on the function p. Since p=  $\alpha^2 (d\delta(0)/d\tau)$ , the absorption depends on the square of the coupling  $\alpha$  and on the time rate of change of the detuning from the isolated atom level evaluated at the point of crossing. In radiative-type collisions, as the processes we are discussing, the coupling  $\alpha$  of the reduced two-level system depends on the product of the collisional and electromagnetic field couplings. This results in the dependence of p on the parameters of the collisional interaction through  $\alpha^2$  and  $d\delta/d\tau$ . This is in contrast to optical-type collisions where the electromagnetic field is tuned near the resonance in one atom. In these collisions, the only dependence on the collisional parameters enter through  $\delta$ .

The absorption probability decays exponentially due to spontaneous decay rate  $\gamma_0$  and the induced ionization rate  $\gamma_1$ , which depends on the intensity of the field and on the oscillator strength of the transition between the excited state and the continuum state.

## **III. IONIZATION YIELD**

Equation (57) describes the absorption at one crossing point where the system is initially in the ground state. In a single collision, there are two times where the system comes into resonance with the photon field: when the two atoms, separated by a distance R, are approaching to, and departing from each other. Since, during collision, not much decay or ionization takes place unless those rates are of order  $10^{12} \text{ sec}^{-1}$ , one can show that after two symmetric resonance times  $|a_2|^2$  is

$$|a_2|^2 = 2(1 - e^{-2\pi p}) e^{-2\pi p}, \tag{59}$$

where we have taken the probabilities induced at the first crossing as initial conditions at the second crossing.

We discuss now some limiting cases of Eq. (59). The above calculated probability is valid for both the weak- and strong-field cases. For strong fields where  $2\pi p \gg 1$ ,  $|a_2|^2$  goes to zero. Note that after one crossing [Eq. (57)]  $|a_2|^2$  reduces to 1 for strong fields, which indicates that the first crossing induces complete inversion, while the second crossing induces complete stimulated

emission which leaves  $a_2$  unpopulated. Since in the limit of very weak field  $|a_2|^2$  goes to zero, there exists an intermediate field where  $|a_2|^2$ reaches a maximum. However, when  $2\pi p \ll 1$ , then  $|a_2|^2$  becomes linear in the field intensity:

 $|a_2|^2 = 4\pi p.$  (60)

Far away in the wing-e.g., when the atoms get closer to each other, resulting in more distortion of the resonance lines—we expect the slope of  $\delta$  to be large. Therefore, in this regime the condition  $2\pi p \ll 1$  holds; and the population of the excited state is linear in the field intensity. This picture is consistent with the fact that the inverse of the slope of the potential-curve difference is a measure of the time the system stays in resonance. We consider now the ionization yield. We first consider the case where the ionization rate is of the order  $10^{12}$ ; there, appreciable ionization is achieved before other collisions take place and before the second crossing. This condition requires power densities of the order  $10^{11}$  W/cm<sup>2</sup> for a photoionization cross section of the order  $10^{-17}$  cm<sup>2</sup>. Such high power will induce power broadening of the order of  $5 \times 10^{13}$  sec<sup>-1</sup> which will tend to wash out the crossing. However, very far away on the wing, i.e., when the atoms get very close to each other, the potential curves become very steep and a slight change in R will induce a detuning effect larger than the power broadening. Therefore, on the extreme far wing and for nearly complete ionization and inversion during the collision, one needs to consider only one crossing as given in Eq. (57).

We next calculate the ionization yield. Since the continuum states were eliminated from the coupled equations by introducing a shift and a decay to the excited state, one can use this scheme to calculate the total ionization probability. From Eq. (41) and from the result that the decay of the excited state into the continuum is given by  $\gamma_1$  [Eq. (46)], one writes the total probability of ionization r as

$$\boldsymbol{\gamma} = \int_{-\infty}^{\infty} \gamma_1 \left| a_2 \right|^2 dt \,. \tag{61}$$

Using the asymptotic expressions calculated for  $|a_2|^2$ , one finds that

$$r = \frac{\gamma_1(1 - e^{-2\pi\rho})}{\gamma_0 + \gamma_1} \left\{ 1 - \exp\left[-(\gamma_0 + \gamma_1)T\right] \right\}, \tag{62}$$

where *T* is the pulse width. Saturation of the ionization can be achieved if  $\gamma_1 \gg \gamma_0$  and  $(\gamma_0 + \gamma_1)T \gg 1$ . In this case *r* reduces to

$$r = 1 - e^{-2\pi p}.$$
 (63)

When  $2\pi p \gg 1$ ,  $\gamma_1 \gg \gamma_0$ , and  $\gamma_1 T \gg 1$ , both states are

saturated and the ionization probability becomes 1.

If the absorption is weak while the ionization is nearly saturated, one has to consider both crossings; and the ionization yield will be twice the yield in Eq. (63) in the limit of  $2\pi p \ll 1$ .

Let us consider a situation where the ionization rate is much smaller than the reciprocal of the time of resonance. This case typically arises with power densities of the order  $10^8 \text{ W/cm}^2$  and ionization rates of the order  $10^{10}$  sec<sup>-1</sup>. After the absorption in a single collision takes place, the system evolves according to the slower processes between collisions. For a typical coupling and a distance of 10 Å, the frequency detunings can reach  $10^{12}$  sec<sup>-1</sup> which results in complete dephasing of the interaction. Therefore, the rate equation treatment is adequate for the processes between collisions. In the rate equation treatment, one calculates an absorption cross section first by integrating over the impact parameters. Far on the wing where  $2\pi p \ll 1$ ,  $|a_2|^2$  reduces to  $4\pi p$ , as given in Eq. (60); therefore, the cross section  $\sigma$  is

$$\sigma = 4\pi \int_{0}^{R_{0}} 2\pi b \, db \, p \,, \tag{64}$$

where  $R_0$  is the internuclear separation where absorption takes place.

Substituting for p in Eq. (64) from (58) and carrying out the integration, we get

$$\sigma = \frac{4\pi^2}{3} \frac{a_0^2}{V} \frac{dR_0^3}{d\delta},\tag{65}$$

where  $\alpha_0$  is  $\alpha$  evaluated at the crossing, and V is the relative speed.

The velocity average can now be carried out. The result is a thermal rate constant k:

$$k = N(8KT/\pi M)^{1/2}\overline{\sigma}, \qquad (66)$$

where

$$\overline{\sigma} = 2 \int_0^\infty x^3 e^{-x^2} \sigma(x) \, dx \,, \tag{67}$$

*k* is the rate of absorption, *M* is the reduced mass of the system, *T* is the temperature, *K* is Boltzmann's constant,  $x = (M/2KT)^{1/2}V$ , and *N* is the foreign-gas density. This results in

$$k = \frac{8N\pi^2}{3} \alpha_0^2 \frac{dR_0^3}{d\delta}.$$
 (68)

When the collision between atoms is considered,  $v_1 = v_2 = c_6^{1/2}/R^3$  and  $\alpha_0^2 = \mu_A^2 E^2 v^2 / (\Delta + \delta')^2$ . Therefore, k becomes

$$k = \frac{4N\pi^2}{3} \left( \hbar c_6^{1/2} \right) \frac{\mu_A^2 E^2}{(\Delta + \delta')^2} \frac{1}{\delta^{1/2}}$$
(69)

which predicts that the absorption rate on the far wing goes like  $\delta^{-1/2}$ . This drop-off is much slower

than that of optical collisions where the excitation frequency is tuned close to a transition in one of the atoms.<sup>19</sup> In such line-broadening cases the far-wing absorption rate goes like  $\delta^{-3/2}$ .

In the case of charge exchange, the wavelength dependence is much more complicated due to the overlap interaction energies. Substituting for  $\alpha_0^2$  from Eq. (38) and for  $\hbar^\delta$  the quantity  $c_4/R_0^4$ , we get for k:

$$k = N\pi^{2} \frac{\mu_{A}^{2} E^{2}}{(\Delta + \delta')^{2}} \frac{G^{2}}{\hbar^{2}} \left(\frac{c_{4}}{\hbar}\right)^{5/4} \frac{1}{a_{0}\delta^{9/4}} \\ \times \exp\left[-\frac{1.72}{a_{0}\delta^{1/4}} \left(\frac{c_{4}}{\hbar}\right)^{1/4}\right].$$
(70)

Note that for large  $\delta^{1/4}$  the exponential becomes negligible and the rate of absorption becomes proportional to  $\delta^{2,25}$ . The reason for this fast decay, as compared to (69), is that the distortion in the energy levels is less steep than that in atom-atom collision and the interaction energies at the crossing go like  $1/R^4$  for small *R* compared to  $1/R^6$  in the atom-atom collision.

Note that k can be written as  $I\overline{\sigma}$ , where I is the fluence of the light source and  $\overline{\sigma}$  is an average cross section. Note also that the temperature dependence of the rate is absent. This result agrees with results that are derived from statistical theories. In these theories, the statistical distribution of frequencies is defined by regarding the intensity at frequency  $\nu$  as proportional to the time interval during which  $\nu$  is irradiated. This time interval in turn is proportional to the relative volume of configuration space in which the frequency perturbation is  $\nu$ . Since our results [Eqs. (69) and (70) were derived on the assumption that  $2\pi p \ll 1$  in Eq. (60), we predict the statistical arguments will fail as the intensity of light increases. This can be reached at frequencies close to the line center. However, one has to be careful since our calculation requires sharp crossing, as was discussed after Eq. (60). This, in turn, requires somewhat steep potential curves. Therefore, higher powers of  $2\pi p$  will be important; therefore, higher powers of the configuration space become important, resulting in temperature-dependent rate constants.

The assumption of straight-line trajectories in the case of interaction between an ion and a neutral is not valid at slow velocities. The attractive force resulting from the polarization of the neutral by the incident ion leads to curved or spiraling orbits. In the classical model, all particles incident at an impact parameter less than a certain value will, in fact, get trapped and spiral inward to the origin. Although some particles might approach at an impact parameter greater than that of the curve crossing, they may still make it to the crossing, but with a velocity that differs from that of the initial approach. Due to this orbiting effect, we consider a modification to the Landau-Zener result [Eq. (70)].

The total energy of the reduced system E is

$$E = \frac{1}{2}M \left(\frac{dR}{dt}\right)^2 + \frac{L^2}{2MR^2} - \frac{c_4}{R^4},$$
 (71)

where *L* is the angular momentum of the system and *R* and *M* are the internuclear separation and the reduced mass, respectively. Due to conservation of energy and angular momentum, the total energy and angular momentum can be written in terms of the initial velocity  $V_0$  and impact parameter  $\rho_0$  as  $E = \frac{1}{2}MV_0^2$  and  $L = MV_0\rho_0$ . The classical capture radius  $\rho_c$  is defined as that at which *E* is equal to the maximum of the effective potential  $V_e$ :

$$\frac{v_e = M V_0^2 \rho_0^2}{2R^2} - \frac{c_4}{R^4}.$$
 (72)

This results in

$$\rho_c = (4c_4/E_0)^{1/4}. \tag{73}$$

For  $\rho$  greater than  $\rho_c$ , the maximum impact parameter  $\rho_{\text{max}}$  such that a curve crossing at  $R = R_0$  can be reached can be calculated from Eq. (71) by putting  $dR_c/dt = 0$ . This results in

$$\rho_{\max} = R_0 (1 + \rho_c^4 / 4R_0^4)^{1/2}. \tag{74}$$

Also, from Eq. (71) one can find the minimum curve crossing  $R_{o_{\min}}$  which may be reached for a given impact parameter:

$$R_{0_{\min}} = \left\{ \frac{1}{2} \left[ \rho^2 - (\rho^4 - \rho_c^4)^{1/2} \right] \right\}^{1/2}.$$
(75)

The velocity at the curve crossing can also be derived from Eq. (71):

$$\frac{d}{dt}R_0 = v_0 \left(1 - \frac{\rho^2}{R_0^2} + \frac{\rho_c^4}{4R_0^4}\right)^{1/2}.$$
(76)

Due to the orbiting phenomena which result in the above equations (73)-(76), the absorption cross section

$$\sigma = 4\pi \int_{0}^{\rho_{\max}} e^{-2\pi p} (1 - e^{-2\pi p}) b \, db \tag{77}$$

gets modified. Note that the above equations imply that for a crossing radius  $R_0$  in the range  $0 < R_0 \le \rho_c / \sqrt{2}$ ,  $\rho_{\text{max}}$  is equal to  $\rho_c$ ; while for the range  $\rho_c / \sqrt{2} < R_0 < \infty$ ,  $\rho_{\text{max}}$  is given by Eq. (74).

Using the above equations and with x and the dimensionless quantities  $\lambda$  and  $\eta$  defined below, one can show that Eq. (77) becomes

$$\sigma = 4\pi R_0^2 (1 + \rho_c^4 / 4R_0^4) \int_0^{x_m} e^{-\eta x} (1 - e^{-\eta x}) x^{-3} \, dx \,, \qquad (78)$$

$$x_{m} = \frac{(1 + \rho_{c}^{4}/4R_{0}^{4})^{1/2}}{(\rho_{c}^{2}/2R_{0}^{2}) - 1} \quad \text{for } 0 \le R_{0} \le \rho_{c}/\sqrt{2},$$
(79)

$$x_m = \infty \text{ for } \rho_c / \sqrt{2} \le R_0 \le \infty, \tag{80}$$

where

$$\lambda = \rho_c^4 / 4R_0^4, \tag{81}$$

$$\eta = \frac{2\pi \,\alpha^2}{(dR/d\delta) V_0 (1+\lambda)^{1/2}} , \qquad (82)$$

$$x = (1+\lambda)^{1/2} (1+\lambda - \rho^2 / R_0^2)^{-1/2}, \qquad (83)$$

and

$$2\pi p = \eta x. \tag{84}$$

Let us consider the case of weak fields such that  $2\pi p \ll 1$  or  $\eta$  is very small, and the capture radius  $\rho_c$  is much larger than the crossing radius  $R_0$ . Then Eq. (78) gives

$$\sigma = 4\pi \rho_c^2 \; \frac{2\pi p}{V_0} \; \frac{\partial R_0}{\partial \tau}.$$
(85)

Result (85) shows that the effect of orbiting is to replace the crossing distance  $R_0$  by the classical capture radius  $\rho_c$ . Note that  $\rho_c$  depends on the energy of the system as given in Eq. (73); therefore, the velocity dependence of the cross section becomes  $1/v^2$  rather than 1/v as in the case of no orbiting. Substituting Eq. (73) in Eq. (85), we get

$$\sigma = 16\pi^2 \left(\frac{2c_4}{M}\right)^{1/2} p \frac{\partial R_0}{\partial t} \frac{1}{V_0^2} . \tag{86}$$

Note that the product  $p(\partial R_0/\partial t)$  is a velocity-independent quantity. The thermal energy of the cross section yields an absorption rate k as given in Eqs. (66) and (67). The result is

$$\overline{\sigma} = \frac{8\pi^2}{KT} (2c_4 M)^{1/2} p \, \frac{\partial R_0}{\partial t} \,. \tag{87}$$

The absorption-rate k dependence on the detuning  $\delta = c_4 / \hbar R^4$  can now be written easily. Using Eqs. (39), (58), (66), and (87), we get

$$k = 32N\pi^{3/2} \left(\frac{c_4}{KT}\right)^{1/2} \alpha_0^2 \frac{\partial R_0}{\partial \delta} .$$
 (88)

Comparing Eq. (88) with Eq. (68) shows the reduced sensitivity to the crossing radius when orbiting effects are taken into consideration. Substituting  $c_4/R^4$  for  $\hbar\delta$  reduces Eq. (88) to

$$k = 8N\pi^{3/2} \left(\frac{c_4^2}{\hbar KT}\right)^{1/2} \frac{\mu_A^2 E^2}{(\Delta + \delta')^2} \frac{G^2}{a_0 \hbar^2} \frac{1}{\delta^{3/2}} \\ \times \exp\left[-\frac{1.72}{a_0 \delta^{1/4}} \left(\frac{c_4}{\hbar}\right)^{1/4}\right].$$
(89)

This absorption rate falls more slowly with  $\delta$  than that of the case of straight trajectory.

With the absorption rates derived, one can write the rate equations describing the two-photon ionization.<sup>19</sup> Let  $n_0$ ,  $n_1$ , and  $n_e$  be the populations of the ground state, excited state, and the number of atoms ionized, respectively, and  $N_0$  be the initial population of the ground state. Then one writes

$$\frac{dn_0}{dt} = -\sigma_a I n_0 + \sigma_a I \frac{g_1}{g_0} n_1 + \gamma_0 n_1, \qquad (90)$$

$$\frac{dn_{1}}{dt} + \gamma_{0}n_{1} = \sigma_{a}In_{0} - \sigma_{a}I\frac{g_{1}}{g_{0}}n_{1} - \sigma_{i}In_{1}, \qquad (91)$$

$$\frac{dn_e}{dt} = \sigma_i I n_1, \tag{92}$$

where  $\sigma_a$  and  $\sigma_i$  are the absorption and ionization cross sections, respectively; *I* is the photon fluence;  $g_0$  and  $g_1$  are the *g* factors of the ground and excited states, respectively; and  $\gamma_0$  is a spontaneous decay of the excited state. Or one may write

$$\begin{bmatrix} \frac{d^2}{dt^2} + \left(\sigma_a I + \sigma_a I \frac{g_1}{g_0} + \sigma_i I + \gamma_0\right) \frac{d}{dt} + \sigma_a I \left(\gamma_0 + \sigma_a I \frac{g_1}{g_2} + \sigma_i I\right) \\ - \sigma_a^2 I^2 \frac{g_1}{g_0} - \sigma_a I \gamma_0 \end{bmatrix} n_1 = 0.$$
(93)

Note that  $\sigma_i I$  is not equal to the isolated-pair ionization rate  $\gamma_1$  [Eq. (46)] because here ionization takes place during collisions. During collision, quantum substates are mixed. The cross section  $\sigma_i$  includes the degeneracy of the excited state. Equation (93) can easily be shown to have the following solution:

$$n_1 = \frac{N_0 \sigma_a I}{f_2 - f_1} (e^{-f_1 t} - e^{-f_2 t}).$$
(94)

where

$$f_1 = \alpha - \beta, \quad f_2 = \alpha + \beta, \tag{95}$$

and

$$\boldsymbol{\alpha} = \frac{1}{2} \left[ \sigma_{\mathbf{z}} I \left( 1 + \frac{g_{1}}{g_{0}} \right) + \sigma_{\mathbf{z}} I + \gamma_{0} \right], \qquad (96)$$

$$\beta = \left[\frac{1}{4} \left(\gamma_0 + \sigma_a I \frac{g_1}{g_2} + \sigma_i I - \sigma_a I\right)^2 + \sigma_a^2 I^2 \frac{g_1}{g_0} + \sigma_a I \gamma_0\right]^{1/2}.$$
(97)

Therefore, the ionization yield can be derived from Eq. (92):

$$N_e = \sigma_i I \int_0^\tau n_1 dt , \qquad (98)$$

where  $\tau$  is the pulse length. Substituting Eq. (94) in Eq. (98), we get

$$N_e = \frac{N_0 \sigma_a I}{f_2 - f_1} \sigma_i I \left( \frac{1 - e^{-f_1 \tau}}{f_1} - \frac{1 - e^{-f_2 \tau}}{f_2} \right).$$
(99)

Let us again consider the case where the ionization is saturated while the laser is tuned on the far wing. This case is of interest because the ionization line shape reduces to the absorption line shape; therefore, the analysis becomes simpler. When  $\gamma_0 \ll \sigma_i I \gg \sigma_a I$ , one can show that  $f_1$  reduces to  $\sigma_a I$  and  $f_2$  reduces to  $\sigma_i I$ . Thus

$$N_e = N_0 \left( 1 - \frac{\sigma_i e^{-\sigma_a I\tau} - \sigma_a e^{-\sigma_i I\tau}}{\sigma_i - \sigma_a} \right), \tag{100}$$

which reduces when  $\sigma_i \gg \sigma_a$  and  $\sigma_i I \tau \gg 1$  to

$$N_e = N_0 \left( 1 - e^{-\sigma_a t \tau} \right). \tag{101}$$

Expressions (99) and (101) describe the ionization yield of the gas mixture for arbitrary degree of saturation of the ionization and for complete ionization, respectively. The latter expression is the absorption of the system since saturation of the ionization is achieved. On the far wing and when  $\sigma_a I \tau \ll 1$ , the absorption yield is

$$N_e/N_0 = \sigma_a I \tau. \tag{102}$$

Equation (102) shows linear dependence on buffergas pressure and on the light intensity. The dependence on the detuning enters in  $\sigma_a$ . Substituting the velocity-averaged cross section in the case of atom-atom scattering for  $\sigma_a$ , we get

$$\frac{N_e}{N_0} = \frac{4N\pi^2}{3} (c_6)^{1/2} \frac{\mu_A^2 E^2}{(\Delta + \delta')^2} \,\delta^{-1/2}.$$
(103)

At a particular intensity,  $\ln N_e/N_0$  vs  $\ln \delta$  is a straight line with slope  $-\frac{1}{2}$  and intercept that depends on  $c_6$  and the other parameters; therefore, an accurate measurement of the slope and intercept determines the interaction force. Similarly, substituting the velocity-averaged cross section for the case of charge-transfer collision for  $\sigma_a$  in the case of straight-line trajectory, we get

$$\frac{N_{a}}{N_{0}} = N\pi^{2} \frac{\mu_{A}^{2}E^{2}}{(\Delta + \delta')^{2}} \frac{G^{2}}{\hbar^{2}a_{0}} \left(\frac{c_{4}}{\hbar}\right)^{5/4} \delta^{-9/4}$$

$$\times \exp\left[-\frac{1.72}{a_{0}} \left(\frac{c_{4}}{\hbar\delta}\right)^{1/4}\right]. \tag{104}$$

At a particular electromagnetic field intensity,  ${\rm ln}N_e/N_0$  is

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$$\ln(N_e/N_0) = -\frac{9}{4}\ln\delta - c_1\delta^{-1/4} + c_2 , \qquad (105)$$

which indicates that  $\ln(N_e/N_0)$  vs  $\ln\delta$  is not a straight line. The inclusion of orbiting in the analysis results in [Eqs. (89) and (102)]

$$\frac{N_e}{N_0} = 8N\pi^{3/2} \left(\frac{c_4^2}{\hbar KT}\right)^{1/2} \frac{\mu_A^2 E^2}{(\Delta + \delta')^2} \frac{G^2}{\hbar^2} \, \delta^{-3/2} \\ \times \exp\left[-\frac{1.72}{\delta^{1/4} a_0} \left(\frac{c_4}{\hbar}\right)^{1/4}\right].$$
(106)

which falls off as  $\delta^{-3/2}$  for large  $\delta$ .

In the wing, and although the absorption cross section falls off rapidly with detuning,  $\sigma_d I \tau$  [Eq. (101)] can still be of the order of unity even for moderate electromagnetic fields. This can be achieved by increasing the buffer-gas pressure. With one atmosphere of buffer gas and with power densities of the order of MW/cm<sup>2</sup>, nonlinear absorption takes place even as far out as 70 Å detunings. Saturation of the absorption with such power densities can be easily achieved at 10–15 Å detuning.

## **IV. CONCLUSION**

We have derived expressions for two-photon ionization of atoms induced by radiative-type collisions or its charge-exchange analog. Both regimes where appreciable ionization takes place during single collisions and between collisions are treated. The dependence of the ionization yield on the interatomic forces at long range is derived, thus making these results of importance to experimental studies. This process promises an extremely sensitive method for studying radiative collisions, especially when absorption or fluorescence becomes extremely weak. This is made possible because of the ability with present lasers to ionize every absorption event.

Although absorption tends to decrease as the field intensity rises due to stimulated emission at the second crossing, the two-photon ionization yield can be nearly saturated at the first crossing, thus enhancing the absorption.

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