#### Oscillator strengths for the magnesium isoelectronic sequence<sup>\*</sup>

P. Shorer, C. D. Lin,<sup>†</sup> and W. R. Johnson<sup>‡</sup>

Center for Astrophysics, Harvard College Observatory and Smithsonian Astrophysical Observatory, Cambridge, Massachusetts 02138 (Received 4 April 1977)

The relativistic random-phase approximation is used to calculate the excitation energies and absorption oscillator strengths for the  $3s^2$   ${}^{1}S_{0}$ -3s3p  ${}^{1}P_{1}$ , -3s4p  ${}^{1}P_{1}$ , -3s5p  ${}^{1}P_{1}$ , and -3s3p  ${}^{3}P_{1}$  transitions of selected elements of the magnesium isoelectronic sequence through nuclear charge Z = 92. The results are compared with previous theoretical and experimental values. The data plotted as a function of  $Z^{-1}$  enable us to study

deviations from the predictions of the nonrelativistic Z-expansion theory.

# I. INTRODUCTION

Accurate absorption oscillator strengths for highly stripped atoms are needed in astrophysics for the determination of solar abundances, temperatures, and densities and in the study of laboratory plasmas both for diagnostic purposes and for the determination of the effects of impurities on controlled thermonuclear fusion.

In order to calculate reliable values for transition rates and oscillator strengths, the effects of interelectron correlation must be taken into account. The nonrelativistic random-phase approximation<sup>1</sup> (RPA) has been very successful in including the most important correlation effects for systems with small nuclear charges (Z), but as the nuclear charge increases along the isoelectronic sequence, relativistic effects become important also.<sup>2,3</sup> A random-phase approximation treating relativistic effects nonperturbatively (RRPA) has been applied to the calculation of oscillator strengths of the resonance transitions of elements of the helium<sup>4,5</sup> and beryllium<sup>6</sup> isoelectronic sequences, as well as to transitions between excited states of heliumlike ions.<sup>7</sup> A relativistic calculation of oscillator strengths enables the study of systematic trends to be continued into the high Zend of the isoelectronic sequence where nonrelativistic calculations break down. For low Z elements, where relativistic effects are expected to be unimportant, a relativistic calculation has the virtue that forbidden transitions are obtained in the same manner as allowed transitions.

In this paper, we calculate the excitation energies and absorption oscillator strengths for the allowed resonance transitions  $3s^2 \, {}^{1}S_0 - 3s3p^{1}P_1$ ,  $-3s4p^{1}P_1$ ,  $-3s5p^{1}P_1$ , and for the intercombination transition  $3s^2 \, {}^{1}S_0 - 3s3p^3P_1$  of elements of the magnesium isoelectronic sequence. We use the RRPA in a frozen core approximation which neglects correlations with the 1s, 2s, and 2p inner-shell electrons.

# **II. THEORY**

The equations of the RRPA are derived using the time-dependent Hartree-Fock (TDHF) method.<sup>8</sup> The ground state of an N-electron closed-shell atom is described by a single determinant constructed of the N Dirac-Hartree-Fock (DHF) orbitals  $u_i(\mathbf{r})$ . An external field  $A_+e^{-i\omega t}+A_-e^{i\omega t}$  of frequency  $\omega$  induces perturbations  $w_{i\pm}$  in the orbitals  $u_i$  which then take the form

$$u_{i}(\mathbf{r}) + w_{i+}(\mathbf{r})e^{-i\omega t} + w_{i-}(\mathbf{r})e^{i\omega t}.$$
(1)

Using the variational principle

 $\left\{\delta\Psi(t), (H-i\partial/\partial t)\Psi(t)\right\} = 0,$ 

where

$$H = \sum_{i=1}^{N} \vec{\alpha}_i \cdot \vec{p}_i + \beta_i m - \frac{Ze^2}{r_i} + \sum_{i>j} \frac{e^2}{|\vec{r}_i - \vec{r}_j|}$$

is the Dirac Hamiltonian in natural units for an N electron atom, we obtain the zero-order equations, which are just the usual DHF equations, by equating to zero the terms containing neither the external field nor the perturbations. The firstorder equations, obtained by linearizing in the perturbations and the external field, and separately equating the coefficients of  $\exp(\pm i\omega t)$  equal to zero, are the RRPA equations

$$(h_0 + V - \epsilon_i \mp \omega) w_{i\pm} = (A_{\pm} - V_{\pm}^{(1)}) u_i, \quad i = 1, 2, \dots, N,$$
(2)

where

$$\begin{split} h_{0} &= \vec{\alpha} \cdot \vec{p} + \beta m - e^{2}Z/r , \\ Vw_{i} &= \sum_{j=1}^{N} e^{2} \int \frac{d^{3}r'}{|\vec{r} - \vec{r}'|} \left[ (u_{j}^{*}u_{j})'w_{i} - (u_{j}^{*}w_{i})'u_{j} \right] , \\ V_{\pm}^{(1)}u_{i} &= \sum_{j=1}^{N} e^{2} \int \frac{d^{3}r'}{|\vec{r} - \vec{r}'|} \left[ (u_{j}^{*}w_{j\pm})'u_{i} + (w_{j\pm}^{*}u_{j})'u_{i} - (w_{j\pm}^{*}u_{i})'u_{j\pm} - (w_{j\pm}^{*}u_{j})'u_{j\pm} \right] \end{split}$$

with  $A_{\pm}$  the amplitude of the external perturbation

16

1109

of frequency  $\pm \omega$  and  $\epsilon_i$  the orbital eigenvalue of the *i*th DHF equation.

An eigenvalue problem for the atomic excitation energies  $\omega_k$  and the eigenfunctions  $w_{ik\pm}$  is obtained from the set of homogeneous equations which results from setting the perturbing external field,  $A_{\pm}$ , equal to zero in Eqs. (2). Normalization of the eigenfunctions is such that

$$\sum_{i=1}^{N} \left( \langle w_{ik+} | w_{ik+} \rangle - \langle w_{ik-} | w_{ik-} \rangle \right) = 1.$$
 (3)

The transition rate from an excited state k to the ground state 0 by emission of a single photon, for electric multipole radiation, is given by

$$A_{k \to 0} = 8\pi \alpha \omega_{k} \left| \sum_{i=1}^{N} \left( \langle u_{i} | \vec{\mathbf{a}}_{JM}^{(1)*} \cdot \vec{\alpha} | w_{ik+} \rangle + \langle u_{i} | \vec{\alpha} \cdot \vec{\mathbf{a}}_{JM}^{(1)} | w_{ik-} \rangle \right) \right|^{2}, \quad (4)$$

where J and M are the photon angular momentum quantum numbers and where

$$\vec{\mathbf{a}}_{JM}^{(1)} = \left[ j_{J}'(\omega r) + j_{J}(\omega r) / \omega r \right] \vec{\mathbf{Y}}_{JM}^{(+1)}(\hat{r}) + \left[ J(J+1) \right]^{1/2} j_{J}(\omega r) \vec{\mathbf{Y}}_{JM}^{(-1)}(\hat{r}) / \omega r$$
(5)

is the vector potential of the multipole field in the Coulomb gauge with  $j_J$  a spherical Bessel function and  $Y_{JM}^{(\pm 1)}$  vector spherical harmonics. For electric multipole radiation, a gauge transformation can be performed resulting in the replacement of  $\vec{\alpha} \cdot \vec{a}_{JM}^{(1)}$  in Eq. (4) by  $b_{JM} + \vec{\alpha} \cdot \vec{c}_{JM}$ , where

$$b_{JM} = i \left(\frac{J+1}{J}\right)^{1/2} \left(j_J(\omega r) - \frac{\omega r j_{J+1}(\omega r)}{J+1}\right) Y_{JM}(\hat{r})$$

$$\vec{c}_{JM} = \frac{\omega r j_J(\omega r)}{[J(J+1)]^{1/2}} \hat{r} Y_{JM}(\hat{r}).$$

With the former choice of gauge, the matrix element of Eq. (4) reduces in the nonrelativistic limit to the velocity form of the dipole matrix element, while with the latter choice, the matrix element reduces to the length form.<sup>9</sup> As in the nonrelativistic random-phase approximation, <sup>1(b)</sup> both forms are equivalent in the RRPA.<sup>5</sup>

By adopting the frozen core approximation, we reduce the problem from one involving eleven channels to one involving only the two virtual excitation channels  $3s_{1/2}$ - $np_{1/2}$  and  $3s_{1/2}$ - $np_{3/2}$ . Such a simplifying procedure has been successful in retaining the important interelectron correlation effects while reducing the numerical complexity in previous calculations.<sup>6,10</sup> For all transitions described in this work, we calculate the oscillator strengths with both choices of gauge. Since the inner-shell correlation effects are neglected, the length form should be the more reliable because of the greater sensitivity of the velocity form to

conditions close to the nucleus.<sup>6</sup> The difference between the length and velocity results is an indication of the importance of inner-shell correlation effects.

The RRPA equations are solved iteratively beginning with the zero-order approximation for the energies and oscillator strengths that represents an intermediate coupling Dirac-Hartree-Fock solution. The nonrelativistic designations  ${}^{1}P_{1}$  and  ${}^{3}P_{1}$ are used to label the intermediate coupling solutions. As in Ref. 6, the Breit interaction is neglected. The reduction of the RRPA equations and a procedure for their numerical solution have been described in previous papers.<sup>5,6</sup>

### **III. RESULTS AND DISCUSSION**

Tables I, II, and III present the RRPA results for the  $3s^2 {}^{1}S_0 - 3s 3p {}^{1}P_1$ ,  $-3s 4p {}^{1}P_1$ , and  $-3s 5p {}^{1}P_1$ transitions. They contain also the calculated Dirac-Hartree-Fock intermediate coupling results as well as values taken from the literature. A more extensive tabulation of comparison values for the neutral member is presented in Table IV. Table V compares the RRPA oscillator strengths for the resonance transition with the results of other relativistic calculations. The RRPA excitation energies for the intercombination transition  $3s^2 {}^{1}S_0 - 3s3p {}^{3}P_1$  are of low accuracy (the error ranging from 22% for Cl vI to 13% for Fe XV) and the differences between the length and velocity forms of the oscillator strengths are sizable. Because of the uncertain accuracy of the oscillator strengths for these particular transitions, we present only a sample of our  $3^{3}P_{1}$  results in Table VI, along with semiempirical oscillator strengths<sup>5,6</sup> obtained by correcting for the difference between the theoretical and experimental singlet-triplet energy separation.

For low values of the nuclear charge, where relativistic effects are small, the transition energy can be written<sup>27,28</sup>

$$\Delta E = \epsilon_0 Z^2 + \epsilon_1 Z + \epsilon_2 + \epsilon_3 Z^{-1} + \cdots$$

with  $\epsilon_0 = 0$  if  $\Delta n = 0$ , where *n* is the principal quantum number of the active electron. It is instructive to plot  $\Delta E/Z$  as a function of  $Z^{-1}$  for transitions in which the principal quantum number is unchanged and to plot  $\Delta E/Z^2$  when there is a change in *n* (cf. Fig. 1).

The  $3^{1}P_{1}$  curve shows a departure from linearity as Z increases above 33. Both the  $4^{1}P_{1}$  and  $5^{1}P_{1}$ curves are slightly concave upward throughout the low and moderate Z elements, with a sharper increase in slope as the nuclear charge increases above Z = 50. These increases in slope at high Z are attributable to relativistic effects. Because of

16

TABLE I.  $3s^{2} {}^{1}S_{0} - 3s^{3}p^{1}P_{1}$ . RRPA excitation energies (in a.u.) and oscillator strengths (length form) compared with DHF intermediate coupling results and other values. A (B) denotes  $A \times 10^{B}$ .

		Excitatio	n energy			Oscillato	r strength	
Z	RRPA	DHF	(RPA)	EXP	RRPA	DHF	(RPA)	Other
12	0.1496	0.1314	0.1492 <sup>a</sup>	0.1597 <sup>b</sup>	1.67	1.91	1.665 <sup>a</sup>	$1.81 \pm 0.18^{d}$
13	0.2667	0.2433	0.2653	0.2727	1.85	2.10	1.877	$1.84 \pm 0.18$
14	0.3745	0.3525	0.3716	0.3776°	1.73	2.04	1.739	$1.70 \pm 0.17$
15	0.4782	0.4591	0.4734	0.4793 <sup>b</sup>	1.59	1.91	1.591	$1.60 \pm 0.16$
16	0.5796	0.5637	0.5725	0.5793	1.47	1.79	1.457	$1.46 \pm 0.15$
<b>17</b>	0.6799	0.6670	0.6697	0.6787	1.36	1.66	1.342	$1.28 \pm 0.13$
18	0.7798	0.7696	0.7658	0.7779	1.27	1.55	1.248	$1.21 \pm 0.12$
20	0.9798	0.9746		0.9773	1.11	1.37		$1.09 \pm 0.11$
22	1.1825	1.1816		1.1801°	9.95(-1)	1.22		
<b>24</b>	1.3903	1.3930		1.3879	9.02(-1)	1.10		
26	1.6052	1.6112	1.5111	1.6035 <sup>f</sup>	8.27(-1)	1.01	7.83(-1)	$8.3(-1)^{f}$
28	1.8297	1.8385		1.8285 <sup>g</sup>	7.67(-1)	9.30(-1)		
29	1.9463	1.9564			7.41(-1)	8.96(-1)		
30	2.0663	2.0777			7.17(-1)	8.65(-1)		7.08(-1) <sup>h</sup>
36	2.8795	2.8965			6.11(-1)	7.25(-1)		6.03(-1)
42	3.9273	3.9474		1	5.49(-1)	6.38(-1)		5.40(-1)
54	7.2596	7.2795			5.09(-1)	5.68(-1)		5.02(-1)
74	20.2087	20.2185			6.16(-1)	6.52(-1)		6.12(-1)
79	25.9631	25.9697			6.68(-1)	7.01(-1)		6.84(-1) <sup>i</sup>
92	49.1440	49.1420			8.43(-1)	8.69(-1)		$8.42(-1)^{h}$

<sup>a</sup>Stewart, Ref. 10.

<sup>b</sup>Moore, Ref. 11(a).

<sup>c</sup> Moore, Ref. 11(b).

<sup>d</sup>Wiese, Smith, and Miles, Ref. 12. <sup>e</sup>Fawcett and Peacock, Ref. 13.

these relativistic effects, the leading term of the Z expansion of the transition energy takes the form<sup>29</sup>

$$\Delta E/Z = (\epsilon_1 + Z^3 \alpha^2 \epsilon_r) + \cdots, \quad \Delta n = 0,$$
  
$$\Delta E/Z^2 = (\epsilon_0 + Z^2 \alpha^2 \epsilon_r) + \cdots, \quad \Delta n \neq 0.$$

For all transitions considered, the difference between the RRPA and DHF intermediate coupling energies becomes smaller as the nuclear charge increases. The difference of 12% for MgI in the  $3^{1}P_{1}$  transition is reduced to 1% by Ar VII. For the  $4^{1}P_{1}$  transition the difference is already 1% for the second member of the sequence while for the  $5^{1}P_{1}$ transition it is 1% for the neutral member.

The time-dependent Hartree-Fock (RPA) calculations of Stewart<sup>10</sup> on the  $3^{1}P_{1}$  and  $4^{1}P_{1}$  transitions provide a direct comparison with the RRPA calculations when relativistic effects can be neglected. The RRPA excitation energies are within 2% of the nonrelativistic values for all elements calculated by Stewart for both transitions with the exception of the  $3^{1}P_{1}$  transition of Fe XV for which our value is 6% higher. Our Fe XV value agrees closely with the measured value of Cowan and Widing<sup>14</sup> as

<sup>f</sup> Cowan and Widing, Ref. 14. <sup>g</sup> Fawcett and Hayes, Ref. 15.

<sup>h</sup>Cheng and Johnson, Ref. 16.

<sup>i</sup>Cowan, Ref. 17.

well as with their value calculated taking into account relativistic effects and configuration interaction. The RRPA energies for the allowed transitions are in good agreement with the experimental values for all but the most weakly ionized elements of the isoelectronic sequence.

Figure 2 presents the calculated oscillator strengths in graphical form as a function of  $Z^{-1}$ , as suggested by the expression for the oscillator strength in the nonrelativistic Z expansion method

$$f = f_0 + f_1 Z^{-1} + f_2 Z^{-2} + \cdots$$

with  $f_{\rm 0}$  involving radial integrals over hydrogenic wave functions.

In the region where relativistic effects are expected to be small, say Z < 30, the curves display the typical behavior as a function of  $Z^{-1,30}$  Proceeding from high to low  $Z^{-1}$ , the RRPA curve for the  $3^{1}S_{0}$ - $3^{1}P_{1}$  transition attains a maximum at Al II, after which it decreases steadily toward the nonrelativistic hydrogenic value of zero. The  $3^{1}S_{0}$ - $4^{1}P_{1}$  transition oscillator strengths pass through a minimum at Si III and the  $3^{1}S_{0}-5^{1}P_{1}$  transition oscillator strengths pass through a minimum at Al II.

TABLE II.  $3s^2 {}^1S_0 - 3s4p^1P_1$ . RRPA excitation energies (in a.u.) and oscillator strengths (length form) compared with DHF intermediate coupling results and other values. A(B) denotes  $A \times 10^B$ .

	E	xcitation en	ergy		Oscillator strength					
Z	RRPA	DHF	(RPA)	EXP	RRPA	DHF	(RPA)	Other		
12	0.2023	0.1968	0.2021 <sup>a</sup>	0.2248 <sup>b</sup>	1.64(-1)	1.29(-1)	1.83(-1) <sup>a</sup>	1.25(-1) <sup>d</sup>		
13	0.4611	0.4547	0.4599	0.4872	9.67(-3)	3.06(-2)	2.5(-3)	2.7(-3)		
14	0.7774	0.7713	0.7748	$0.8041{^{\rm c}}$	6.36(-3)	1.08(-3)	1.7(-2)	3.11(-2)		
15	1.1471	1.1414	1.1427	1.1733 <sup>b</sup>	3.92(-2)	6.00(-3)	5.9(-2)	8.44(-2)		
16	1.5688	1.5634	1.5606		8.10(-2)	2.74(-2)	1.01(-1)	1.37(-1)		
17	2.0415	2.0365	2.0260		1.24(-1)	5.62(-2)	1.41(-1)	9.16(-2)		
<b>18</b>	2.5650	2.5603	2.5457	2.5805	1.64(-1)	8.80(-2)	1.81(-1)			
20	3.7641	3.7597		3.7923 <sup>e</sup>	2.35(-1)	1.51(-1)				
22	5.1662	5.1621		5.1939	2.94(-1)	2.08(-1)				
<b>24</b>	6.7728	6.7688		6.8022 <sup>b</sup>	3.38(-1)	2.54(-1)				
26	8.5858	8.5819	8.4718	$8.602^{f}$	3.70(-1)	2.91(-1)	4.22(-1)	$5.0(-1)^{f}$		
28	10.6077	10.6038		10.632 <sup>g</sup>	3.92(-1)	3.18(-1)				
29	11.6978	11.6939		11.720	3.99(-1)	3.28(-1)				
30	12.8412	12.8373			4.05(-1)	3.37(-1)				
36	20.8451	20.8413			4.17(-1)	3.63(-1)				
42	30.8879	30.8840			4.10(-1)	3.66(-1)				
54	57.6591	57.6552			3.79(-1)	3.48(-1)				
74	126.2469	126.2427			2.91(-1)	2.73(-1)				
79	149.2703	149.2660			2.62(-1)	2.46(-1)				
92	223.9457	223.9410			1.79(-1)	1.68(-1)				
aS bN cN dV	tewart, Ref. Ioore, Ref. Ioore, Ref.	ughlin Be	f 18	<sup>e</sup> Ekberg, Re <sup>f</sup> Cowan and <sup>g</sup> Feldman, F	ef. 19. Widing, Re Katz, Behri	f. 14. ng, and Coh	en, Ref. 20.			

TABLE III.  $3s^2 {}^1S_0 - 3s5 p^1P_1$ . RRPA excitation energies (in a.u.) and oscillator strengths (length form) compared with DHF intermediate coupling results and other values. A(B) denotes  $A \times 10^{B}$ .

	E	xcitation energ	gy	Oscillator strength			
Z	RRPA	DHF	EXP	RRPA	DHF	Other	
12	0.2238	0.2216	0.2492ª	4.16(-2)	3.68(-2)	2.86(-2)°	
13	0.5412	0.5386	0.5735	6.33(-5)	4.86(-3)	8 (_4)	
14	0.9440	0.9414	0.9775 <sup>b</sup>	8.09(-3)	2.72(-4)	2.45(-2)	
15	1.4258	1.4233	1.4583 <sup>a</sup>	2.30(-2)	6.40(-3)		
16	1.9838	1.9814		3.80(-2)	1.66(-2)		
17	2.6167	2.6144		5.17(-2)	2.79(-2)		
18	3.3239	3.3217		6.41(-2)	3.91(-2)		
20	4.9599	4.9579	5.0009 <sup>d</sup>	8.42(-2)	5.94(-2)		
22	6.8912	6.8893	6.9666	9.89(-2)	7.55(-2)		
24	9.1187	9.1168		1.09(-1)	8.76(-2)		
26	11.6442	11.6423		1.15(-1)	9.59(-2)		
28	14.4699	14.4681		1.19(-1)	1.01(-1)		
29	15.9963	15.9944		1.20(-1)	1.03(-1)		
30	17.5988	17.5970		1.20(-1)	1.05(-1)		
36	28.8383	28.8365		1.20(-1)	1.09(-1)		
42	42.9454	42.9436		1.18(-1)	1.09(-1)		
54	80.3659	80.3641		1.11(-1)	1.04(-1)		
74	174.4096	174.4077		9.14(-2)	8.71(-2)		
79	205.3614	205.3594		8.47(-2)	8.08(-2)		
92	303.7250	303.7229		6.36(-2)	6.07(-2)		

<sup>a</sup>Moore, Ref. 11(a).

<sup>b</sup>Moore, Ref. 11(b).

<sup>c</sup>Victor, Stewart, and Laughlin, Ref. 18.

<sup>d</sup>Ekberg, Ref. 19.

TABLE IV.	RRPA	oscillator	strengths	for	Mgı	compared	with	values	obtained	by	other	methods.	A(B)	denotes
$A \times 10^{B}$ .														

$n^1P_1$	RRPA	DHF	TDHF	(RPA)	Model potential <sup>c</sup>	"Frozen-cores" <sup>d</sup>	Electron-beam phase-shift <sup>e</sup>	Beam-foil <sup>f</sup> Hook <sup>g</sup>
3	1.67	1.91	1.67 <sup>a</sup>	1.66 <sup>b</sup>	1.72	1.76	$1.86 \pm 0.29$	$1.67 \pm 0.15^{f}$
4	1.64(-1)	1.29(-1)	1.83(-1)	1.50(-1)	1.25(1)	1.19(-1)	$1.8 \pm 0.4(-1)$	$1.07 \pm 0.02(-1)^{g}$
5	4.16(_2)	3.68(-2)	•••	3.54(-2)	2.86(-2)	2.73(-2)	$5.5 \pm 1(-2)$	2.27 ± 0.12(-2)

<sup>a</sup>Stewart, Ref. 10.

<sup>b</sup>Amusia, Cherepkov, Pavlin, Radojević, and Živanović, Ref. 21.

<sup>c</sup>Victor, Stewart, and Laughlin, Ref. 18.

<sup>d</sup>Saraph, Ref. 22.

<sup>e</sup>Smith and Liszt, Ref. 23.

<sup>f</sup>Andersen, Desesquelles, Jessen, and Sørensen, Ref. 24.

<sup>g</sup> Mitchell, Ref. 25 (obtained from relative f values based on f = 1.72 for  $3^{1}P_{1}$ ).

TABLE V. RRPA oscillator strengths for the resonance transition compared with the results of other relativistic calculations. L denotes length form; V denotes velocity form. A(B) denotes  $A \times 10^{B}$ .

		RRPA		Multiconi Hartree	iguration e-Fock <sup>a</sup>	Paran poten	netric tial <sup>b</sup>	"HFR" °	Superposition of configurations <sup>d</sup>	
	Ion	$f_L$	$f_{V}$	$f_L$	$f_{V}$	$f_L$	$f_{V}$	f	$\int f$	
	Mg	1.67	1.64	1.73	1.79	1.64	2.02			
	Ar <sup>+6</sup>	1.27	1.28	1.25	1.24	1.17	1.07	1.309		
	Fe <sup>+14</sup>	8.27(-1)	8.51(-1)	8.18(-1)	8.25(-1)	7.9(-1)	6.9(-1)	8.50(-1)	8.18(-1)	
•	Kr <sup>+24</sup>	6.11(-1)	6.32(-1)	6.03(-1)	6.11(-1)			6.19(-1)	6.05(-1)	
	Mo <sup>+30</sup>	5.49(-1)	5.66(-1)	5.40(-1)	5.47(-1)			5.50(-1)	5.43(-1)	
	$Xe^{+42}$	5.09(-1)	5.21(-1)	5.02(-1)	5.07(-1)			5.07(-1)	5.08(-1)	
	W *62	6.16(_1)	6.21(_1)	6.12(_1)	6.14(_1)			6.25(_1)	6.14(_1)	

<sup>a</sup>Cheng and Johnson, Ref. 16.

<sup>b</sup>Aymar and Luc-Koenig, Ref. 3(c).

<sup>c</sup>Cowan, Ref. 17.

<sup>d</sup>Weiss, Ref. 26.

TABLE VI.  $3s^{2} \, {}^{1}S_{0} - 3s 3p \, {}^{3}P_{1}$ . RRPA excitation energies (in a.u.) and oscillator strengths (length form) for magnesiumlike chlorine, titanium, and iron compared with other values. Semiempirical oscillator strengths (EMP) are obtained by multiplying the RRPA values by  $E(\text{EXP})\Delta E^{2}(\text{RRPA})/E(\text{RRPA})\Delta E^{2}(\text{EXP})$  where E and  $\Delta E$  are the triplet energies and singlet-triplet energy differences, respectively. A(B) denotes  $A \times 10^{B}$ .

 			-				
 Z	E(RRPA)	$\Delta E$ (RRPA)	E(EXP)	$\Delta E$ (EXP)	f (RRPA)	<i>f</i> (EMP)	f (other)
17	0.3658	0.3141	0.4463 <sup>a</sup>	0.2324 a,b	8.61(-5)	1.92(_4)	1.8(-4)°
22	0.6946	0.4879	0.7920 <sup>a</sup>	0.3881 <sup>a,d</sup>	6.97(-4)	1.26(-3)	
26	0.9698	0.6354	1.0920°	0.5115°	2.17(-3)	3.77(_3)	3.3(_3)°

<sup>a</sup> Ekberg, Ref. 19.

<sup>b</sup> Moore, Ref. 11(a).

<sup>c</sup> Cheng and Johnson, Ref. 16.

<sup>d</sup>Fawcett and Peacock, Ref. 13.

<sup>e</sup>Cowan and Widing, Ref. 14.

The curves then increase with decreasing  $Z^{-1}$ . Relativistic effects cause departures from the non-relativistic predictions at higher Z.

For all three allowed resonance transitions, the RRPA length and velocity results are in close agreement, except near the minima in the  $4^{1}P_{1}$  and  $5^{1}P_{1}$  curves. They are in harmony with the nonrel-ativistic TDHF values of Stewart for the  $3^{1}S_{0}-3^{1}P_{1}$  transition for small Z. The oscillator strengths given by Wiese, Smith, and Miles<sup>12</sup> for  $Z \leq 20$  are estimated to have an uncertainty of less than 10%;

the RRPA results are in good agreement with these values.

Uncertainties in the theoretical oscillator strengths are expected to be large for those transitions for which the oscillator strength curves are near minima due to cancellation in the radial matrix elements and configuration interaction effects.<sup>3(C), 1B, 30(b)</sup> Thus, agreement among the values calculated by different authors for the low Z elements of the  $3^{1}S_{0}-4^{1}P_{1}$  transition is less satisfactory than for the  $3^{1}S_{0}-3^{1}P_{1}$  transition. For the



FIG. 1. (a)  $3s^2 {}^{1}S_0 - 3s 3p {}^{1}P_1$ . RRPA excitation energy in a.u., divided by the nuclear charge as a function of  $Z^{-1}$ . (b)  $3s^2 {}^{1}S_0 - 3s 4p {}^{1}P_1$ . RRPA excitation energy in a.u., divided by the square of the nuclear charge as a function of  $Z^{-1}$ . (c)  $3s^2 {}^{1}S_0 - 3s 5p {}^{1}P_1$ . RRPA excitation energy in a.u., divided by the square of the nuclear charge as a function of  $Z^{-1}$ . (c)  $3s^2 {}^{1}S_0 - 3s 5p {}^{1}P_1$ . RRPA excitation energy in a.u., divided by the square of the nuclear charge as a function of  $Z^{-1}$ . On the scale to which the graphs have been drawn, the corresponding DHF results coincide with the RRPA data points.





 $4^{1}P_{1}$  transition, the difference between the RRPA and TDHF values of Stewart is of the order of 10% for MgI and Ar VII, while for elements closer to the minimum [Fig. 2(b)] the difference is a factor of 3 or 4. Whereas the RRPA value for MgI for the  $3^{1}P_{1}$  transition is in excellent agreement with the nonrelativistic RPA calculations of Amusia, Cherepkov, Pavlin, Radojević, and Živanović<sup>21</sup> in which correlations with the *L*-shell electrons are taken into account, the RRPA values for the 4 and  $5^{1}P_{1}$  transitions are respectively 9% and 15% greater than the corresponding RPA values. The differences may be largely attributed to the sensitivity of the transition matrix element to small changes in the wave function for elements for which the oscillator strength versus  $Z^{-1}$  curve is near a minimum. The RRPA values for MgI for the  $4^{1}P_{1}$  and  $5^{1}P_{1}$  transitions lie between the experimentally determined values of Smith and Liszt<sup>23</sup> and of Mitchell.<sup>25</sup>

The line strength S is related to the oscillator strength by

$$f=\frac{2}{3}(\Delta E/g)S$$

where g is the degeneracy of the initial state and  $\Delta E$  is the transition energy. In the Z-expansion

$$S = S_0 Z^{-2} + S_1 Z^{-3} + S_2 Z^{-4} + \cdots$$

As given by the nonrelativistic charge expansion calculations of Crossley and Dalgarno,<sup>31</sup> S<sub>0</sub> equals 271.65 and S<sub>1</sub> equals 3915.1 for the  $3s^2 {}^{1}S_0 - 3s3p {}^{1}P_1$ transition. For  $U^{+80}$ , the RRPA value of  $Z^2S$  for this transition is 218 as compared to the nonrelativistic value of 320, indicating that the high-Zdeparture of the resonance transition oscillator

- †Permanent address: Department of Physics, Kansas State University, Manhattan, Kans. 66506.
- ‡ Permanent address: Department of Physics, University of Notre Dame, Notre Dame, Ind. 46556.
- \*W. R. J. supported in part by NSF Grant No. GP-42738 and P. S. and C. D. L. supported in part by ERDA Contract No. EY-76-S-02-2887.
- <sup>1</sup>(a) R. F. Stewart, D. K. Watson, and A. Dalgarno, J. Chem. Phys. 63, 3222 (1975); (b) M. Ya. Amusia and N. A. Cherepkov, Case Stud. At. Phys. 5, 47 (1975); (c) C. D. Lin, Phys. Rev. A 9, 171 (1974); 9, 181 (1974).
- <sup>2</sup>A. W. Weiss, in *Beam-Foil Spectroscopy*, edited by I. A. Sellin and D. J. Pegg (Plenum, New York, 1976), Vol. 1, p. 51.
- <sup>3</sup>(a) L. Armstrong, Jr., W. R. Fielder, and D. L. Lin, Phys. Rev. A 14, 1114 (1976); (b) Y. K. Kim and J. P. Desclaux, Phys. Rev. Lett. 36, 139 (1976); (c) A. Aymar and E. Luc-Koenig, Phys. Rev. A 15, 821 (1977).
- <sup>4</sup>W. R. Johnson, C. D. Lin, and A. Dalgarno, J. Phys. B 9, L303 (1976).
- <sup>5</sup>W. R. Johnson and C. D. Lin, Phys. Rev. A <u>14</u>, 565 (1976).
- <sup>6</sup>C. D. Lin and W. R. Johnson, Phys. Rev. A <u>15</u>, 1046 (1977).
- <sup>7</sup>C. D. Lin, W. R. Johnson, and A. Dalgarno, Phys. Rev. A 15, 154 (1977).
- <sup>8</sup>A. Dalgarno and G. A. Victor, Proc. R. Soc. A 291, 291 (1966).
- <sup>9</sup>I. P. Grant, J. Phys. B 7, 1458 (1974).
- <sup>10</sup>R. F. Stewart, Mol. Phys. <u>30</u>, 745 (1975).
- <sup>11</sup>(a) C. E. Moore, Atomic Energy Levels, Natl. Bur. Stand. Ref. Data Ser. No. 35 (U.S. GPO, Washington, D. C., 1971), Vols. I and II; (b) C. E. Moore, Selected Tables of Atomic Spectra, ibid, 1965, No. 3, Sec. 1.
- <sup>12</sup>W. L. Wiese, M. W. Smith, and B. M. Miles, Atomic Transition Probabilities, Nat. Bur. Stand. Ref. Data Ser. No. 22 (U.S. GPO, Washington, D. C., 1969), Vol. II.
- <sup>13</sup>B. C. Fawcett and N. J. Peacock, Proc. Phys. Soc.

strength from the nonrelativistic prediction is attributable to the relativistic corrections to the excitation energy rather than to the transition integral.

#### ACKNOWLEDGMENT

The authors thank Professor A. Dalgarno for many useful discussions. This work was supported in part by the Energy Research and Development Administration under Contract No. EY-76-S-02-2887.

91, 973 (1967).

- <sup>14</sup>R. D. Cowan and K. G. Widing, Astrophys. J. 180, 285 (1973).
- <sup>15</sup>B. C. Fawcett and R. W. Hayes, J. Phys. B 5, 366 (1972).
- <sup>16</sup>K. T. Cheng and W. R. Johnson (to be published).
- <sup>17</sup>R. D. Cowan, Los Alamos Scientific Lab. Rept. No. LA-6679-MS (unpublished).
- <sup>18</sup>G. A. Victor, R. F. Stewart, and C. Laughlin, Astrophys. J. Suppl. Ser. 31, 237 (1976).
- <sup>19</sup>J. O. Ekberg, Phys. Scr. 4, 101 (1971).
- <sup>20</sup>U. Feldman, L. Katz, W. Behring, and L. Cohen, J. Opt. Soc. Am. 61, 91 (1971).
- <sup>21</sup>M. Ya. Amusia, N. A. Cherepkov, I. Pavlin, V. Radojević, and Dj. Živanović in Fourth International Conference on Atomic Physics Abstracts of Contributed Papers, edited by J. Kowalski and H. G. Weber (Heidelberg University, Heidelberg, 1974), pp. 328-331. The results for Mg1 appear in Ref. 1 (b).
- <sup>22</sup>H. E. Saraph, J. Phys. B 9, 2379 (1976).
- <sup>23</sup>W. H. Smith and H. S. Liszt, J. Opt. Soc. Am. <u>61</u>, 938 (1971).
- <sup>24</sup>T. Andersen, J. Desesquelles, K. A. Jessen, and G. Sørensen, J. Quant. Spectrosc. Radiat. Transfer 10, 1143 (1970).
- <sup>25</sup>C. J. Mitchell, J. Phys. B <u>8</u>, 25 (1975).
- <sup>26</sup>A. W. Weiss (unpublished). Values taken from At. Data Fusion 3, 27 (1977).
- <sup>27</sup>(a) D. Layzer, Ann. Phys. (N.Y.) 8, 271 (1959); (b)
- D. Layzer, Z. Horak, M. N. Lewis, and D. P. Thompson, Ann. Phys. (N.Y.) 29, 101 (1964).
- <sup>28</sup>R. J. S. Crossley, in Advances in Atomic and Molecular Physics, edited by D. R. Bates and I. Estermann (Academic, New York, 1969), Vol. 5, p. 237. <sup>29</sup>H. T. Doyle, in Ref. 28, p. 337.
- <sup>30</sup>(a) W. L. Wiese and A. W. Weiss, Phys. Rev. 175, 50 (1968); (b) W. L. Wiese, Appl. Opt. 7, 2361 (1968). <sup>31</sup>R. J. S. Crossley and A. Dalgarno, Proc. R. Soc.
- A 286, 510 (1965).