

## Electron-Cs<sup>+</sup>-ion recombination in the presence of neutral helium atoms

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Three-body neutral stabilized recombination has been studied in a pulsed discharge afterglow apparatus. The rate coefficient of the He stabilized recombination reaction  $\text{Cs}^+ + e + \text{He} \rightarrow \text{Cs} + \text{He}$  at 625°K is deduced to be less than  $10^{-28} \text{ cm}^6 \text{ sec}^{-1}$ , with the most probable value being  $4 \times 10^{-29} \text{ cm}^6 \text{ sec}^{-1}$ . Since this value is much lower than that obtained for the recombination of  $\text{He}_2^+$  by the same mechanism, it suggests that the efficiency of the three-body neutral stabilized recombination process depends sensitively upon the atomic or molecular nature of the recombining ion.

### I. INTRODUCTION

The three-body (electron-ion-atom) recombination mechanism,



has been presumed for a long time to play an important role in the recombination of weakly ionized gases. Several calculations of an expression for the three-body rate coefficient  $k_N$  characteristic of reaction (1) have been proposed. The analytical formula of Massey and Burhop,<sup>1</sup> which is an extension of a simple calculation of the ion-ion recombination coefficient by Thompson,<sup>2</sup> and that of Pitaevski,<sup>3</sup> differs only by a numerical factor [ $(k_N)_{\text{Pit}} \sim 6(k_N)_{\text{Th}}$ ], and predicts that

$$k_N \sim \sigma/M T^{5/2}, \quad (2)$$

where  $M$  is the mass of atom  $B$ ,  $\sigma$  is the momentum-transfer cross section for  $e$ - $B$  collisions, and  $T$  is the temperature. Note that in this equation  $k_N$  is independent of the nature of the positive ion.

Considering process (1) in more detail for the case of  $\text{He}^+ - e - \text{He}$ , Bates and Khare<sup>4</sup> have carried out numerical calculations of  $k_N$  which are in good agreement with the Pitaevski formula at low temperatures (100°K), but which decrease faster with  $T$  than predicted by (2). For  $\text{H}^+ - e - \text{H}$ , Drawin<sup>5</sup> has derived  $k_N$  coefficients that are larger than those of Bates and Khare for  $\text{He}^+ - e - \text{He}$  by one to several orders of magnitude. Thus, the many discrepancies that exist among the theoretical determinations of  $k_N$  indicate that it is necessary to experimentally measure these constants. Note that all the theoretical investigations have been performed assuming that the gas temperature  $T_g$  is equal to the electron temperature  $T_e$ . Although it is certain that  $k_N$  is dependent on both  $T_e$  and  $T_g$ , this dependence is not precisely known.<sup>6</sup>

For an experimental study of process (1), however, a careful choice of the experimental conditions must be made in order to reduce as much as

possible the effects of other ionic loss channels. On the one hand, a ratio of the concentration of atoms  $B$  to that of the electrons is desired for which the contribution of the mechanism  $A^+ + e + e$  to the global electron-ion recombination is not predominant in comparison to process (1). On the other hand, the absolute values of the electron and atom densities are important. If the electron density  $n_e$  is too low, diffusion to the walls becomes the major loss process of the charged particles. A concentration of atoms  $B$ ,  $N_B$ , that is too high, however, would generally result in molecular ion formation, and if these ions recombine by a very efficient dissociative mechanism, this could become the dominant loss channel.

These types of difficulties contribute to the sparsity of experimental results relating to reaction (1). To our knowledge, the only published data concern  $\text{He}_2^+ + e + \text{He}$  recombination in an afterglow,<sup>7-9</sup> and we shall discuss later the relatively high value of the rate constant ( $k_N = 5 \cdot 10^{-27} \text{ cm}^6 \text{ sec}^{-1}$  for  $T = 300 \text{ }^\circ\text{K}$ ) that was obtained.<sup>9</sup> In this paper we present the results of our afterglow investigations of the three-body recombination process in a cesium-helium mixture:



We chose these particular reactants for the following reasons:

(a) Cs has an ionization potential (3.9 eV) much lower than the first excitation potential of He (19.8 eV); it is thus possible to create a plasma in which only cesium atoms are excited or ionized, while the helium atoms remain in their ground state, even when the ratio of the densities  $N_{\text{Cs}}/N_{\text{He}}$  is small.

(b) According to Eq. (2), the coefficient  $k_N$  is larger when the third body is a light atom.

(c) The molecular ion  $\text{He-Cs}^+$  is most likely weakly bound (Mason and Schamp<sup>10</sup> estimate 0.014 eV); so the concentration of such ions would remain small.

The coefficient  $k_N$  for reaction (3) was obtained by fitting the measured electron-density decay frequencies with calculated curves representing the contribution of the various processes that contribute to this decay.

## II. EXPERIMENT AND MEASUREMENT TECHNIQUES

The experimental setup is quite similar to ones previously used.<sup>11,12</sup> The experimental cell consists of a cylindrical glass discharge tube 8 cm in diameter. A pulsed discharge (25 msec long with a repetition frequency between 3 and 20 Hz) is applied between two heated electrodes. The cell is placed in an oven, and the cesium pressure is controlled by regulating the temperature of the coolest point of the cell. An excellent agreement has been observed between the saturated vapor pressure corresponding to the temperature of the coolest point of the cell and that deduced from the cesium density as measured by absorption of the 8521-Å resonance line.

The electron density, averaged over the tube diameter, is deduced from the phase shift of a free-space electromagnetic wave (4-mm wavelength). The radial distribution of  $n_e$ , and consequently the electron density along the tube axis ( $n_{e0}$ ), is obtained from optical measurements.<sup>12</sup> The actual electron temperature  $T_e$  is deduced by observing the intensity of the continuous spectrum emitted during recombination to the 6P level.<sup>10,12</sup>

Typical experimental conditions are  $N_{\text{Cs}} = 10^{15} \text{ cm}^{-3}$ , initial electron density  $(1-2) \times 10^{13} \text{ cm}^{-3}$ , and  $T_{\text{oven}} = 625^\circ \text{K}$ . The working pressure of helium ( $p_{\text{He}}$ ) was limited to less than 60 torr, since at higher helium pressures we have observed that during the discharge period cataphoresis effects tend to make the plasma unstable.

## III. EXPERIMENTAL RESULTS AND INTERPRETATION

Because of its influence on the recombination coefficient  $k_N$  [see Eq. (2)], we shall consider first the electron temperature and its decay in the afterglow. Then, we shall look at the late afterglow and the useful information we can obtain concerning various loss processes for the electrons. Finally, we shall consider the early afterglow ( $n_e > 10^{11} \text{ cm}^{-3}$ ) and the various recombination mechanisms involved there.

The evolution of the electron temperature during the afterglow is shown in Fig. 1. After decaying quickly during the first 100  $\mu\text{sec}$ ,  $T_e$  then decreases very slowly (from 1100 to 800°K in 9 msec). The shape of the  $T_e$  decay curve is very similar to that observed in pure cesium afterglows and may be interpreted in the same way.<sup>11</sup> The initial fast decay of  $T_e$  corresponds to the

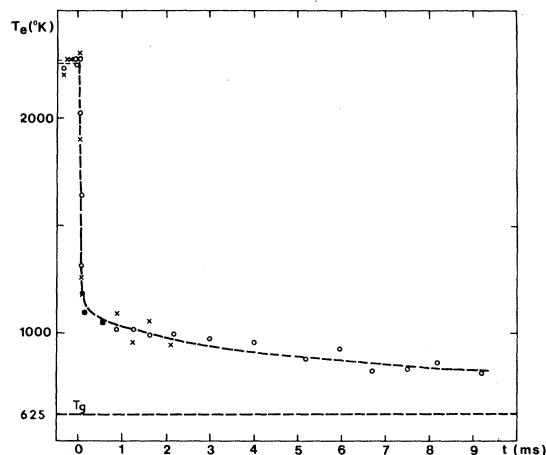


FIG. 1. Electron temperature as a function of time.  $p_{\text{He}} = 60$  torr. Repetition rate of the discharge: 16.7 Hz ( $\times$ ), 6.7 Hz ( $\circ$ ).

cooling of the electrons by collisions with heavy particles (ions and neutral atoms) which are at a lower temperature  $T_g$ . A quasiequilibrium situation is then reached where the electron density and electron temperature decays are coupled together. In this situation, the three-body electron-stabilized recombination supplies the free electrons with energy and compensates for the inelastic losses with heavy particles.<sup>13</sup> This quasiequilibrium holds as long as recombination is the main loss process.

In Fig. 1 we have plotted the electron-temperature decay for two repetition rates of the discharge pulses. From the equality of the quasiequilibrium temperature measured under both conditions, we can infer that no noticeable heating effect of the heavy particles is induced during the discharge pulse.

Figure 2 presents the time decay of the electron density on the axis of the tube ( $n_{e0}$ ). For the later afterglow, the decay becomes exponential. This behavior might seem to be characteristic of a decay controlled by diffusion, but if we compare the measured time constant to the diffusion time constant calculated from the reduced mobilities of  $\text{Cs}^+$  in helium<sup>14,15</sup> and in cesium,<sup>16</sup> we find that the experimental value (18 msec) is much smaller than the calculated one (68 msec). This large difference cannot be explained either by the uncertainties related to the values of the reduced mobilities, or by the presence of a molecular ion created by the attachment of a  $\text{Cs}^+$  ion with one or several neutral atoms, since such a heavy ion would not diffuse sufficiently fast to explain the measured decay time constant of the electron density. An explanation of the above can be found, however. It is known (see, for example, Ref. 14) that an expo-

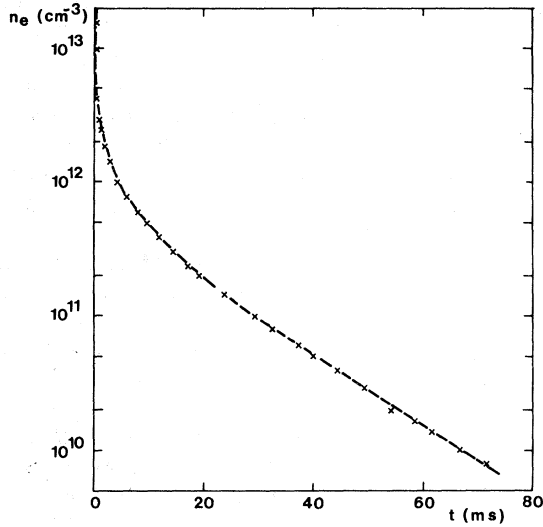
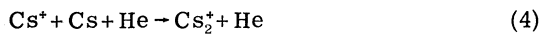


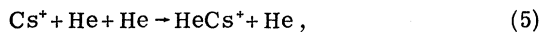
FIG. 2. Decay of electron density along the axis of the discharge tube.  $p_{\text{He}} = 60$  torr. Repetition rate: 6.7 Hz.

nential decay of the electron density is observed in an afterglow when, besides being lost by ambipolar diffusion, electrons recombine rapidly with molecular ions that have been created by a slow three-body conversion process from atomic ions. Under these circumstances, the atomic ions remain the majority ions and the decay rate ( $1/\tau_{\text{exp}}$ ) of the electron density is equal to the sum of the ambipolar diffusion and the three-body ion conversion frequencies,  $\nu_D$  and  $\nu_C$ , respectively.

If we adopt this interpretation for our work, the three-body ion conversion process being either



or



with  $\text{Cs}_2^+$  or  $\text{HeCs}^+$  recombining rapidly, we obtain at  $p_{\text{He}} = 60$  torr for  $\nu_C$  a value of  $40 \text{ sec}^{-1}$ , since  $1/\tau_{\text{exp}} = 55 \text{ sec}^{-1}$  and  $\nu_D = 15 \text{ sec}^{-1}$  for  $\text{Cs}^+$  ions and electrons. At  $p_{\text{He}} = 30$  torr, uncertainties in the measured  $\tau_{\text{exp}}$  and the computed  $\nu_D$  prevent us from precisely determining the value of  $\nu_C$ . From the comparison of  $\nu_C$  at 60 and 30 torr, the dependence of  $\nu_C$  on helium pressure cannot be accurately obtained, so that the exact nature of the conversion process, (4) or (5), cannot be elucidated. We suggest, however, that it is process (4) (see Ref. 16), since analogous processes have been evoked to explain the late afterglow decay of discharges in mercury mixed with helium,<sup>17</sup> or in pure cesium at high pressures.<sup>18</sup>

We will now try to interpret the behavior of the first part of the afterglow at  $p_{\text{He}} = 60$  torr ( $n_{e0} \geq 10^{11}$

$\text{cm}^{-3}$ ,  $t < 25$  msec). The experimental decay rate of the electron density along the axis of the tube,

$$\nu_{\text{exp}} = \frac{1}{\tau_{\text{exp}}} = - \frac{1}{n_{e0}} \frac{dn_{e0}}{dt}, \quad (6)$$

can be reasonably assumed to contain contributions from (i) the two processes which have been already found responsible for the decay of  $n_{e0}$  in the later afterglow, i.e., ambipolar diffusion (frequency  $\nu_D$ ) and three-body ion conversion (frequency  $\nu_C$ ), (ii) three-body electron stabilized recombination



(frequency  $\nu_R$ ), and finally, (iii) a recombination process corresponding to reaction (3) with a frequency  $\nu_N$ . In order to calculate these four frequencies from the measured parameters ( $p_{\text{He}}$ ,  $n_{\text{Cs}}$ ,  $n_{e0}$ ,  $T_e$ ), we need to make the following assumptions:

(a) The diffusion frequency can be expressed as

$$\nu_D = - \left( \frac{1}{n_{e0}} \frac{dn_{e0}}{dr} \right)_{\text{diff}} = \frac{D_a \nabla^2 n_{e0}}{n_{e0}} \approx \frac{D_i}{\Lambda^2} \left( 1 + \frac{T_e}{T_i} \right), \quad (8)$$

where  $D_a$  and  $D_i$  are the ambipolar and  $\text{Cs}^+$  ion diffusion coefficients, respectively, and  $T_i$  is the ionic temperature (assumed equal to the oven temperature). Relation (8) is strictly true only if diffusion is the dominant loss process. In that case  $\Lambda$  is the diffusion length characterizing the spatial distribution of charged particles in the cell according to the fundamental mode. If diffusion is not the major loss process, the spatial distribution is different and the last equality of (8) is no longer valid. A good approximation can be obtained, however, if  $\Lambda$  is written as  $\Lambda_0 \xi$ , where  $\Lambda_0$  is the classical diffusion length and where  $\xi$  takes into account the modification of the spatial distribution (see Ref. 19).

(b) As already mentioned, the exact nature of the conversion process of  $\text{Cs}^+$  ions into molecular ions which then induce an electron loss is not known, but whatever it is, it will be dependent on neither  $n_e$  nor  $T_e$ . Consequently, the same value of  $\nu_C$  deduced from the late afterglow decay ( $40 \text{ sec}^{-1}$  at  $p_{\text{He}} = 60$  torr) has been assumed to be present during the whole afterglow.

(c) The decay frequency  $\nu_R$  characterizing three-body electron-stabilized recombination (7) can be calculated according to the relation deduced from results obtained in pure cesium for the same range of  $n_e$  and  $T_e$ <sup>11</sup>:

$$\nu_R \equiv \alpha n_{e0} = 2.4 \times 10^{-10} n_{e0}^2 T_e^{-4.2},$$

where  $\alpha$  is the recombination coefficient. Note that the variation of  $\nu_R$  with  $n_{e0}$  will appear less

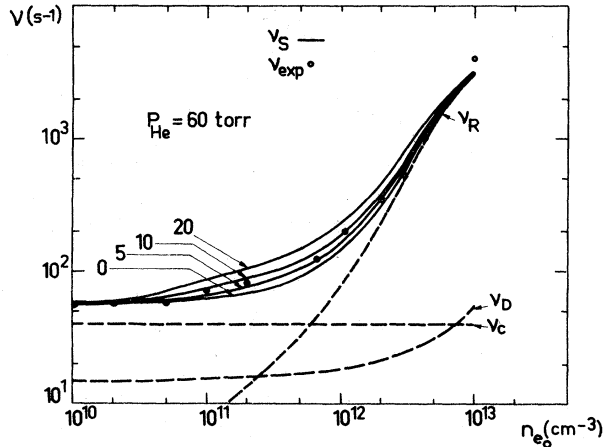


FIG. 3. Comparison of the inverse of the measured decay time constant of  $n_{e0}$  with the calculated sum  $\nu_s = \nu_R + \nu_D + \nu_C + \nu_N$  for the values of  $k_N^{(0)} = 0, 5, 10,$  and  $20 \times 10^{-29} \text{ cm}^6 \text{ sec}^{-1}$ . For electron densities lower than  $5 \times 10^{11} \text{ cm}^{-3}$ , the electron temperatures are extrapolated.

than quadratic because  $T_e$  also increases with  $n_{e0}$  (see Figs. 1 and 2).

(d) Process (3) is characterized by a frequency  $\nu_N = k_N(T_e, T_g)n_{e0}N_{\text{He}}$ . The variation of  $k_N$  as a function of  $T_e$  for  $T_e \neq T_g$  is not known, as already mentioned. We have chosen to write

$$k_N(T_e, T_g) = k_N^{(0)}(T_g) \times \left( \frac{T_e}{T_g} \right)^{-2.5}, \quad (9)$$

where  $k_N^{(0)}(T_g)$  is the three-body recombination coefficient when  $T_e = T_g$ . By taking this form of  $k_N$ , we certainly overestimate its dependence upon  $T_e$ .<sup>6</sup>

The frequencies  $\nu_D$ ,  $\nu_C$ , and  $\nu_R$  are plotted in Fig. 3 as a function of  $n_{e0}$ . Also plotted is the sum  $\nu_s = \nu_D + \nu_C + \nu_R + \nu_N$  as calculated for four values of  $k_N^{(0)}$  (with  $k_N^{(0)} = 0, 5, 10,$  and  $20 \times 10^{-29} \text{ cm}^6 \text{ sec}^{-1}$ ). These values of  $\nu_s$  are compared with

the measured  $\nu_{\text{exp}}$  in Fig. 3. It appears that  $\nu_D$  and  $\nu_C$  are the major terms contributing to the sum  $\nu_s$  for  $n_{e0} < 2 \times 10^{11} \text{ cm}^{-3}$ , while  $\nu_R$  becomes preponderant for  $n_{e0} > 2 \times 10^{12} \text{ cm}^{-3}$ . Between these limits is the region where the decay of  $n_{e0}$  is the most sensitive to the value of  $\nu_N$ . In this region the best agreement between experimental results and calculated curves is obtained for  $k_N^{(0)}(625^\circ\text{K}) = 5 \times 10^{-29} \text{ cm}^6 \text{ sec}^{-1}$ .

Note that if a weaker dependence of  $k_N(T_e, T_g)$  as a function of  $T_e$  had been chosen in Eq. (9), it would lead to a smaller value of  $k_N^{(0)}$ . For example, if we take the exponent in Eq. (9) to be  $-1.5$  (which appears to be an upper limit as suggested by Ref. 6), we obtain  $k_N^{(0)}(625^\circ\text{K}) = 3.5 \times 10^{-29} \text{ cm}^6 \text{ sec}^{-1}$ . Therefore, taking into account the dispersion of the experimental values of  $\nu_{\text{exp}}$  and the uncertainties introduced in the calculation of  $\nu_s$ , we think it reasonable to attribute to  $k_N^{(0)}$  an upper limit of  $10^{-28}$  and a probable value around  $4 \times 10^{-29} \text{ cm}^6 \text{ sec}^{-1}$ .

#### IV. DISCUSSION

The values of the  $k_N^{(0)}$  coefficients associated with the three-body recombination mechanism (1) and calculated from the theoretical models<sup>1-4</sup> are compared in Table I with the available experimental results. The experimental conditions for the  $\text{He}_2^+ + e^- + \text{He}$  study<sup>8</sup> were  $10^9 \leq n_e \leq 10^{11} \text{ cm}^{-3}$  and  $10 \leq p_{\text{He}} \leq 100 \text{ torr}$ . Although the value obtained for  $k_N(625^\circ\text{K})$  in the present work is lower than that calculated according to Pitaevski,<sup>3</sup> it is not in contradiction with those deduced by Thompson<sup>1</sup> or by Bates and Khare.<sup>4</sup> However, it is two orders of magnitude lower than the experimental value of  $k_N$  obtained for the recombination of the  $\text{He}_2^+$  ion stabilized by an He atom.<sup>9</sup> Such a variation can be explained neither by the difference between the gas temperatures of the two experiments nor by the difference between the mass of the ions involved, as is suggested by Ref. 4. It is possible

TABLE I. Theoretical and experimental three-body coefficients  $k_N^{(0)}$  (in  $\text{cm}^6 \text{ sec}^{-1}$ ) corresponding to recombination process (1). The notation  $A(B)$  means  $A \times 10^B$ .

	Thompson <sup>a</sup>	Pitaevski <sup>b</sup>	Bates and Khare <sup>c</sup>	Experimental results	References
$\text{Cs}^+ - \text{He}$ (625°K)	6.1(-29)	3.8(-28)	~3.5(-29)	<1.(-28) Probable value 4.(-29)	This work
$\text{He}_2^+ - \text{He}$ (300°K)	3.9(-28)	2.4(-27)	~6(-28)	5 ± 1(-27)	9

<sup>a</sup>References 1 and 2.

<sup>b</sup>Reference 3.

<sup>c</sup>The values of  $k_N^{(0)}$  in this column have been extrapolated from those given in Ref. 4, assuming the ionic mass dependency suggested by Bates and Khare.

that the molecular nature of the  $\text{He}_2^+$  ion is responsible for the large value of  $k_N$ . One can note that the results obtained in the study of a helium afterglow suggest that the three-body recombination stabilized by a helium atom would be less efficient for  $\text{He}^+$  than for  $\text{He}_2^+$ .<sup>20</sup> Further experimental work is needed to compare the different behavior of atomic and molecular ions with respect to process (1).

The  $\text{Cs}^+ + e + e \rightarrow \text{Cs} + e$  recombination reaction has a coefficient which nearly equals  $4 \times 10^{-22} \text{ cm}^6 \text{ sec}^{-1}$  for  $T_e = 625 \text{ }^\circ\text{K}$ .<sup>11</sup> In view of the value found for  $k_N^{(0)}$  we can see that the stabilization of the  $\text{Cs}^+ + e$  recombination by an atom may compete with the stabilization by an electron only if  $n_e/n_{\text{He}} \leq 10^{-7}$ .

### V. CONCLUSION

By using a static afterglow system, we have examined the role played by the three-body neutral stabilized recombination process [process (3)] in the recombination of  $\text{Cs}^+$  ions. For our

experimental conditions ( $n_e/n_{\text{He}} \approx 10^{-6}$ ), this process does not appear to play an important role by comparison to diffusion, ion conversion, or the  $\text{Cs}^+ + e + e$  recombination mechanism. A value of the rate coefficient for reaction (3) is found to be approximately  $k_N^{(0)} = 4 \times 10^{-29} \text{ cm}^6 \text{ sec}^{-1}$ , at  $T_e = T_g = 625 \text{ }^\circ\text{K}$ , which agrees with theoretical estimates of Thompson<sup>1</sup> and Bates and Khare,<sup>4</sup> but not with Pitaevski.<sup>3</sup> The value of  $k_N$  obtained here appears to be two orders of magnitude lower than the one measured for  $\text{He}_2^+ + e + \text{He}$ ,<sup>9</sup> suggesting that the molecular nature of that ion plays an important role in the recombination mechanism.

### ACKNOWLEDGMENTS

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<sup>12</sup>B. Sayer, J. C. Jeannet, and J. Berlande, *J. Phys. (Paris)* **33**, 993 (1972).

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<sup>14</sup>L. B. Loeb, *Basic Processes of Gaseous Electronics* (University of California, Berkeley, 1961).

<sup>15</sup>R. S. Bergman and L. M. Chanin, *Phys. Rev. A* **8**, 1076 (1973).

<sup>16</sup>Two ionic conversion processes can be envisaged:

$\text{Cs}^+ + 2\text{He} \rightarrow \text{HeCs}^+ + \text{He}$  or  $\text{Cs} + \text{Cs}^+ + \text{He} \rightarrow \text{Cs}_2^+ + \text{He}$ , which would, respectively, have a frequency  $\nu_C$  proportional to  $p_{\text{He}}^2$  or  $p_{\text{He}}$ . Because of the probable weak binding energy of  $\text{HeCs}^+$  (see Ref. 9), we think that the correct conversion process is the one leading to  $\text{Cs}_2^+$ . Using this hypothesis, the value of  $\nu_C$  deduced from the experiment at  $p_{\text{He}} = 60$  torr would lead to a three-body coefficient for the conversion process close to  $4 \times 10^{-32} \text{ cm}^6 \text{ sec}^{-1}$ . This value is about the same order of magnitude as the one found by Biondi (Ref. 17) for  $\text{Hg}^+ + \text{Hg} + \text{He} \rightarrow \text{Hg}_2^+ + \text{He}$  ( $1.7 \times 10^{-31} \text{ cm}^6 \text{ sec}^{-1}$ ).

<sup>17</sup>M. A. Biondi, *Phys. Rev.* **90**, 730 (1953).

<sup>18</sup>B. Sayer, J. C. Jeannet, G. Saint-Hilaire, J. Lozingot, and J. Berlande, *J. Phys. B* **8**, 503 (1975).

<sup>19</sup> $\Lambda$  has been taken equal to  $\Lambda_0 \xi$ , where  $\Lambda_0 = R/2.405$ , is the classical diffusion length of a long cylinder having a radius  $R$ .  $\xi(t)$  is a correction factor which is introduced for taking into account the evolution of the spatial distribution of the charged particles with time.  $\xi(t)$  has been taken as the ratio between the measured halfwidth of the  $n_e$  radial distribution and the Bessel distribution function  $J_0$ . During the initial discharge period, the plasma is "pinched" and thus does not fill in radially all the volume of the tube; it then diffuses during the afterglow. Consequently,  $\xi$  is smaller than 1 at the beginning of the afterglow, and then, in our experiment, approaches unity after about 2 msec.

<sup>20</sup>L. C. Pitchford and R. Deloche (private communication).