

Shock-wave generation and propagation in dissipative and nonlocal nonlinear Rydberg mediaLu Qin,^{1,2,3} Chao Hang⁴, Guoxiang Huang,⁴ and Weibin Li^{2,3}¹*Department of Physics, Henan Normal University, Xinxiang 453007, China*²*School of Physics and Astronomy, University of Nottingham, Nottingham, NG7 2RD, United Kingdom*³*Centre for the Mathematics and Theoretical Physics of Quantum Non-Equilibrium Systems, University of Nottingham, Nottingham NG7 2RD, United Kingdom*⁴*State Key Laboratory of Precision Spectroscopy, East China Normal University, Shanghai 200062, China*

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We investigate the generation of optical shock waves in strongly interacting Rydberg atomic gases with a spatially homogeneous dissipative potential. The Rydberg atom interaction induces an optical nonlocal nonlinearity. We focus on local nonlinear ($R_b \ll R_0$) and nonlocal nonlinear ($R_b \sim R_0$) regimes, where R_b and R_0 are the characteristic length of the Rydberg nonlinearity and beam width, respectively. In the local regime, we show spatial width and contrast of the shock wave change monotonically when increasing strength of the dissipative potential and optical intensity. In the nonlocal regime, the characteristic quantity of the shock wave depends on R_b/R_0 and dissipative potential nontrivially and on the intensity monotonically. We find that formation of shock waves dominantly takes place when R_b is smaller than R_0 , while the propagation dynamics is largely linear when R_b is comparable to or larger than R_0 . Our results reveal nontrivial roles played by dissipation and nonlocality in the generation of shock waves, and provide a route to manipulate their profiles and stability. Our study furthermore opens new avenues to explore non-Hermitian physics and nonlinear wave generation and propagation by controlling dissipation and nonlocality in the Rydberg media.

DOI: [10.1103/PhysRevA.110.013703](https://doi.org/10.1103/PhysRevA.110.013703)**I. INTRODUCTION**

Nonlinear hydrodynamic flows are found in different media [1,2], such as plasma [3–5], acoustics systems [6], polymerized ionic liquid [7], quantum-mechanical piston [8,9], ultracold quantum gases [10–15], and optical media [16–27]. In these systems, the competition of nonlinearity, dispersion, and dissipation gives rise to nonlinear wave phenomena, such as solitons [28–30], and rogue wave [31,32]. In defocusing nonlinear media, an initially smooth wave steepens when propagation, eventually reaching a point of gradient catastrophe [33] that lead to the formation of shock waves [8,33–46]. Profiles of shock waves depend on the dissipation of the medium. In dissipation-free media, the formed dispersive shock waves (DSW) show a strong oscillatory structure due to the interplay between the nonlinearity and dispersion [16,18,38,47–52]. This steepening can also be mediated by dissipation, where the nonlinear wave acquires a monotonic shock front without any oscillations. In such cases, a dissipative shock wave, sometimes also called viscous shock wave (VSW), emerges [7,8,47,53].

Recently, it has been shown that cold atomic gases interacting with laser fields provide a fertile ground for studying shock waves [10–15]. When additionally coupling the light to

highly excited Rydberg states [54,55], strong and long-range interactions between Rydberg atoms can be mapped to light fields through electromagnetically induced transparency (EIT) [56], generating strong nonlocal nonlinearities [57,58]. The characteristic length of the nonlocal nonlinearity, determined by the blockade radius of the Rydberg gas, is in the order of micrometers, which is comparable to typical beam width. Using the strong Rydberg nonlocal nonlinearity (NNL), it has been shown that DSWs can be generated and manipulated in Rydberg atom gases [38]. Dissipation plays an important roles in the study of Rydberg systems [59–61]. In cold atom gases, dissipation, on the other hand, can be induced and controlled [28,30]. This opens new opportunities for exploring shock waves in the interplay between the nonlocal nonlinearity and controllable dissipation that is otherwise difficult to achieve in other systems.

In this work, we study the generation and propagation of shock waves within a cold Rydberg atomic gas setting, incorporating an engineered, homogeneous dissipative potential that can be changed from loss to gain. This change is controlled by employing an incoherent pumping [30] and controlling the laser detuning [28]. A nonlocal optical nonlinear interaction is induced by coupling low-lying electronic states to Rydberg S state via EIT [56,57]. Depending on the blockade radius R_b of the NNL and beam width R_0 , the system is in a local regime when $R_b \ll R_0$, and nonlocal regime when $R_b \sim R_0$. In the local regime, the nonlinear Schrödinger (NLS) equation governing the light propagation is cast into coupled Riemann equations. Formation of shock waves is signified by wave breaking. Excluding dissipation, wave breaking points

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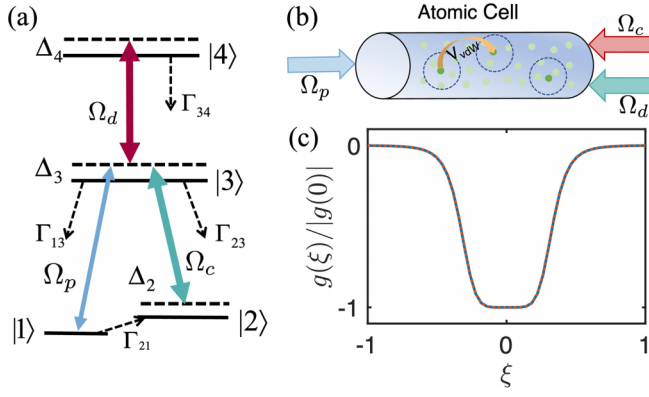


FIG. 1. System. (a) Inverted Y-type level diagram. A weak probe laser field (half-Rabi frequency Ω_p) couples transition $|1\rangle \leftrightarrow |3\rangle$. A strong control (dressed) laser field [half-Rabi frequency Ω_c (Ω_d)] couples transition $|2\rangle \leftrightarrow |3\rangle$ ($|3\rangle \leftrightarrow |4\rangle$). In Rydberg state $|4\rangle$ atoms interact strongly through the van der Waals interaction. Here Δ_α are detuning and $\Gamma_{\alpha\beta}$ ($\alpha < \beta$) are spontaneous emission decay rates. The emission from state $|3\rangle$ generates an effective loss potential. To achieve the gain, we apply an incoherent pumping Γ_{21} between states $|1\rangle$ and $|2\rangle$. (b) The blue, red, and cyan arrows indicate the propagating direction of the probe, control, and dressed fields. The strong Rydberg atom interaction gives rise to an blockade radius (dashed circle) around Rydberg atoms. The blockade radius can be tuned by the excitation laser. (c) Response function $g(\xi)$. The numerically obtained $g(\xi)$ (blue solid line) agrees with the analytical approximation Eq. (4) (dashed red line). Details of the response function can be found in Sec. II B.

are obtained analytically through the Riemann equations. Oscillation contrast [39,62] and spatial width of the shock wave depends on the strength of the dissipation and optical intensity monotonically in the local regime. In the nonlocal regime, the nonlocal degree of the optical nonlinearity modifies amplitudes and width of shock waves. Importantly properties of the shock wave exhibit complicated dependence on the nonlocality. We show that shock wave generation and propagation are important when $R_b < R_0$. When R_b and R_0 are comparable, the medium is effectively linear, where the NNL becomes a homogeneous dispersive potential approximately.

The paper is arranged as follows. In Sec. II, we present the physical model that can lead to the dissipative and nonlocal nonlinear potential. The NLS equation that describes the propagation of the probe laser field is derived. In Sec. III, light propagation in the local regime is discussed. The impact of the dissipative potential on the oscillation contrast, width, and shock width is investigated. In Sec. IV, we explore the influence of the nonlocality on the generation and propagation of shock waves. Finally, conclusions are given in Sec. V.

II. MODEL AND LIGHT PROPAGATION EQUATIONS

A. Physical model

We consider a gas of cold atoms with an inverted Y-type four-level configuration [see Fig. 1(a)], where a weak probe laser field with half-Rabi frequency Ω_p couples the transitions $|1\rangle \leftrightarrow |3\rangle$. A strong control and a dressed laser fields with half-Rabi frequencies Ω_c and Ω_d couple the transition

$|2\rangle \leftrightarrow |3\rangle$ and $|3\rangle \leftrightarrow |4\rangle$, correspondingly. Detuning Δ_α ($\alpha = 2, 3, 4$) gives difference between laser frequency and atomic transition. And $\Gamma_{\alpha\beta}$ are spontaneous emission decay rates from $|\beta\rangle$ to $|\alpha\rangle$ ($\alpha < \beta$). Here the incoherent decay from state $|3\rangle$ causes loss of the probe field. To generate gain, an incoherent pumping (with pumping rate Γ_{21}) is used to pump atoms from $|1\rangle$ to $|2\rangle$. Driven by the control laser Ω_c a small number of atoms are populated in state $|3\rangle$, which provides a gain effect [30,63].

In this setting, state $|4\rangle$ is a high-lying Rydberg state. The interaction between two Rydberg atoms, respectively, at positions \mathbf{r} and \mathbf{r}' is described by van der Waals potential $V_{\text{vdW}} \equiv \hbar V(\mathbf{r}' - \mathbf{r}) = -\hbar C_6/|\mathbf{r}' - \mathbf{r}|^6$ [64]. When the light propagates in the medium [see Fig. 1(b)], Rydberg excitation in the vicinity of a Rydberg atom is strongly suppressed, due to the long-range Rydberg-Rydberg interaction. Such spatial dependent Rydberg blockade leads to nonlocal nonlinear optical interactions [65]. Note that in the inverted Y-shaped excitation scheme shown in Fig. 1(a), the transition $|1\rangle \rightarrow |2\rangle \rightarrow |3\rangle$ forms a Λ -type EIT, while $|1\rangle \rightarrow |3\rangle \rightarrow |4\rangle$ forms a ladder-type Rydberg EIT. The interplay of the two paths can be controlled by the external lasers, giving rise to dissipative, nonlocal nonlinear interactions.

Under electric-dipole and rotating-wave approximations, the Hamiltonian of the system is $\hat{H} = \mathcal{N}_a \int d^3r \hat{\mathcal{H}}$ with Hamiltonian density $\hat{\mathcal{H}}$ ($\hbar \equiv 1$),

$$\hat{\mathcal{H}} = - \sum_{\alpha=1}^4 \Delta_\alpha \hat{S}_{\alpha\alpha}(\mathbf{r}) - [\Omega_p \hat{S}_{13}(\mathbf{r}) + \Omega_c \hat{S}_{23}(\mathbf{r}) + \Omega_d \hat{S}_{34}(\mathbf{r}) + \text{H.c.}] + \mathcal{N}_a \int d^3r' \hat{S}_{44}(\mathbf{r}') V(\mathbf{r}' - \mathbf{r}) \hat{S}_{44}(\mathbf{r}),$$

where \mathcal{N}_a is atomic density, and $\hat{S}_{\alpha\beta}(\mathbf{r}) = |\beta\rangle\langle\alpha| \exp[i(\mathbf{k}_\beta - \mathbf{k}_\alpha) \cdot \mathbf{r} - (\omega_\beta - \omega_\alpha + \Delta_\beta - \Delta_\alpha)t]$ is the atomic transition operator between states $|\alpha\rangle$ and $|\beta\rangle$. For weak excitation, $\hat{S}_{\alpha\beta}(\mathbf{r})$ are approximated by bosonic operators [66]. Taking into account of decay, dynamics of the density matrix (matrix elements $\rho_{\alpha\beta} \equiv \langle \hat{S}_{\alpha\beta} \rangle$) is described by the Bloch equation $\partial \hat{\rho} / \partial t = -i[\hat{H}, \hat{\rho}] / \hbar - \Gamma[\hat{\rho}]$, where Γ is the relaxation matrix describing the spontaneous emission and dephasing (see Appendix A). Propagation of the semiclassical probe field is governed by the Maxwell equation,

$$i \left(\frac{\partial}{\partial z} + \frac{1}{c} \frac{\partial}{\partial t} \right) \Omega_p + \frac{c}{2\omega_p} \left(\frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} \right) \Omega_p + \kappa_{13} \rho_{31} = 0,$$

where $\kappa_{13} = \mathcal{N}_a \omega_p |\mathbf{p}_{13}|^2 / (2\epsilon_0 c \hbar)$, with ω_p the weak probe laser field of center frequency, \mathbf{p}_{13} the electric dipole matrix element associated with the transition $|1\rangle \leftrightarrow |3\rangle$, ϵ_0 the vacuum dielectric constant, and c the speed of light in vacuum. In deriving the propagation equation, the paraxial and slowly varying envelope approximations have been applied.

B. Nonlinear envelope equation

For weak probe field, the Maxwell and Bloch equations can be solved perturbatively. The Rydberg-Rydberg interaction, on the other hand, is treated beyond the simple mean-field approximation [28,67]. We then solve the Bloch equation up to the third order of Ω_p . This allows us to derive a

(2 + 1)-dimensional [(2 + 1)D] nonlocal nonlinear Schrödinger (NNLS) equation of the probe field [30,63,68],

$$i\frac{\partial}{\partial z}\Omega_p + \frac{c}{2\omega_p}\nabla_{\perp}^2\Omega_p - V_1\Omega_p + W|\Omega_p|^2 + \int d^2r'_{\perp}G(\mathbf{r}_{\perp}, \mathbf{r}'_{\perp})|\Omega_p(\mathbf{r}'_{\perp}, z)|^2\Omega_p(\mathbf{r}) = 0, \quad (1)$$

with $\mathbf{r}_{\perp} = (x, y)$. Here V_1 is the dissipative potential that gives homogeneous gain or loss in the medium controlled by the laser parameters [30]. This is achieved by employing jointly the incoherent pumping and spontaneous decay of the excited states. Details on the control of the dissipative potential can be found in Appendix B. Nonlinear coefficient W characterizes the local Kerr nonlinearity (contributed by the weak, short-range interactions between photons and atoms [29]). $G(\mathbf{r})$ is a nonlocal nonlinear response function characterizing respectively NNL [28,67,69].

Using ^{87}Rb as an example with electronic states $|1\rangle = |5S_{1/2}, F=1, m_F=-1\rangle$, $|2\rangle = |5S_{1/2}, F=1, m_F=1\rangle$, $|3\rangle = |5P_{3/2}, F=1, m_F=0\rangle$, and $|4\rangle = |nS_{1/2}\rangle$. We consider the principal quantum number $n=30$. The corresponding dispersion coefficient is $C_6 \approx -2\pi \times 68 \text{ MHz } \mu\text{m}^6$ [70]. As $C_6 < 0$, the Rydberg-Rydberg interaction is repulsive. Other typical parameters $\Delta_2 = 2\pi \times 3.18 \text{ MHz}$, $\Delta_3 = -2\pi \times 31.8 \text{ MHz}$, $\Delta_4 = 2\pi \times 1.59 \text{ MHz}$, $\Gamma_3 = 2\pi \times 6.1 \text{ MHz}$, $\Gamma_4 = 2\pi \times 2.02 \text{ kHz}$, $\Omega_c = 2\pi \times 6.37 \text{ MHz}$, $\Omega_d = 2\pi \times 1.59 \text{ MHz}$, and $\mathcal{N}_a = 2.3 \times 10^{10} \text{ cm}^{-3}$.

With approximation and focusing on the diffraction along the x axis, we convert the propagation equation in a dimensionless form,

$$i\frac{\partial u}{\partial \zeta} + \frac{1}{2}\frac{\partial^2 u}{\partial \xi^2} - \mathcal{V}u + g_0 \int d\xi' g(\xi', \xi)|u(\xi', \zeta)|^2 u = 0, \quad (2)$$

where we have defined dimensionless quantities $u = \Omega_p/U_0$, $\xi = x/R_0$, $\zeta = z/L_{\text{diff}}$, $\mathcal{V} = 2L_{\text{diff}}V_1$, $g = 2L_{\text{diff}}U_0^2R_0^2 \int G(\mathbf{r}_{\perp}, \mathbf{r}'_{\perp})d\mathbf{y}'$, and $g_0 = 1/|\int g(\xi', \xi)d\xi'|$. The optical field and spatial coordinate have been scaled with respect to the maximal Rabi frequency U_0 and beam radius R_0 . We have scaled z with respect to the characteristic diffraction length $L_{\text{diff}} = \omega_p R_0^2/c$, with c being the speed of light in vacuum and $\omega_p \approx 2\pi \times 3.85 \times 10^{14} \text{ Hz}$ the probe light frequency. Note that, compared to the Rydberg induced nonlinearity, the conventional Kerr nonlinearity is marginal and has been neglected in Eq. (2). To show this, we find that the dimensionless Kerr nonlinearity is $\mathcal{W} = 2L_{\text{diff}}U_0^2W$. Considering the probe beam radius $R_0 = 5 \mu\text{m}$ and $U_0 = 2\pi \times 1.59 \text{ MHz}$, we obtain $L_{\text{diff}} \approx 0.2 \text{ mm}$, and the dimensionless Kerr nonlinearity $\mathcal{W} \approx 0.01$, i.e., much smaller than the strength (in the order of 1) of the NNL. In our setting, we have assumed the strong control field is a plane wave field, resulting in a homogeneous complex potential \mathcal{V} . The real part of the complex potential is associated with the refractive index. The imaginary part $V_I = \text{Imag}[\mathcal{V}]$ characterizes the dissipation, i.e., the potential is gain (loss) when $V_I > 0$ ($V_I < 0$).

Due to the Rydberg blockade, the response function has a soft-core shape [see Fig. 1(c)]. However the expression of $g(\xi', \xi)$ is typically lengthy and complicated (see Appendix B

for discussions). It can be approximated by an analytical form [30,38]

$$g(\xi', \xi) \approx -\frac{1}{b_1 + (b_2/\sigma^6)|\xi' - \xi|^6}, \quad (3)$$

where $\sigma = R_b/R_0$ characterizes the nonlocal degree of the nonlinearity. Here $R_b = |C_6/\delta_{\text{EIT}}|^{1/6}$ is the radius of the blockade sphere, with $\delta_{\text{EIT}} \approx |\Omega_c|^2/|\Delta_3|$ the linewidth of EIT transition spectrum (i.e., the width of EIT transparency window) [57,65]. Numerically the blockade radius is $R_b \approx 1.94 \mu\text{m}$ when $n=30$. From Eq. (3), one can observe that the response function has a soft-core profile, with depth $1/b_1$ and soft-core radius $\sigma(b_1/b_2)^{1/6}$. When fixing b_1 and b_2 , one can change the landscape of the response function by changing σ , where the depth will not be affected. Coefficients b_1 and b_2 depend on the laser parameters. Their values are determined through solving the Bloch equation perturbatively (see Appendix B), and can be modified by varying, e.g., the detuning and control laser Rabi frequency without affecting the probing field. The response function can be cast into a different form,

$$g(\xi', \xi) \approx -\frac{B_1}{B_2\sigma^6 + |\xi' - \xi|^6}, \quad (4)$$

The relation between $B_{1,2}$ and $b_{1,2}$ are $B_1 = \sigma^6/b_2$ and $B_2 = b_1/b_2$, i.e., the latter is not affected by σ directly. When fixing B_1 and B_2 , the profile of $g(\xi', \xi)$ becomes wider and its strength weaker when increasing σ . This form thus provides a different way to examine and understand the nonlocal effect.

With the laser and atomic parameters given previously, the dimensionless coefficients in the response function can be obtained, $B_1 = 0.001$, $B_2 = 0.38$, $b_1 = 1.0$, and $b_2 = 2.6$. We plot the approximate response function in Fig. 1(c), which agrees with the numerical one derived from the Bloch equation. Both the analytical and numerical response function capture the soft-core shape. Note that this system has a defocusing nonlinearity as $g(\xi', \xi) < 0$, which is crucial for the generation of shock waves.

When R_b is comparable to R_0 , the nonlinear interaction is nonlocal (i.e., σ is finite). In the opposite regime when $R_b \ll R_0$, we have a local regime as $\sigma \sim 0$. The nonlocality parameter σ can be varied by varying R_b or R_0 . The blockade radius can be tuned by changing detuning, laser intensities, or choosing different Rydberg states as $C_6 \propto n^{11}$ with n to be the principal quantum number. When $\sigma \sim 0$, we can make a local field approximation, i.e., $\int d\xi' g(\xi', \xi)|u(\xi', \zeta)|^2 \approx |u(\xi, \zeta)|^2 \int d\xi' g(\xi', \xi)$. Carrying out the spatial integration, we arrive at a propagation equation with local nonlinear interactions,

$$i\frac{\partial u}{\partial \zeta} + \frac{1}{2}\frac{\partial^2 u}{\partial \xi^2} - \mathcal{V}u + \bar{g}_0|u|^2 u = 0, \quad (5)$$

where $\bar{g}_0 = -g_0\pi\sigma(b_1/b_2)^{1/6}/(3b_1)$ is the effective interaction strength. The local interaction is similar to the conventional Kerr nonlinearity, though the nonlinearity is much stronger. In the following, we will discuss the local and nonlocal regime separately.

III. LOCAL RYDBERG NONLINEARITY REGIME

A. Euler-like fluid equation

In the local regime, we start to investigate the generation of shock waves with the hydrodynamic approach. By treating the light field as a classical fluid, the hydrodynamic equation can be obtained by using the Madelung transformation $u(\xi, \zeta) = \sqrt{\rho(\xi, \zeta)}e^{i\phi(\xi, \zeta)}$, Eq. (5) can be transformed into two Euler-like fluid equations,

$$\frac{\partial \rho}{\partial \zeta} + \frac{\partial}{\partial \xi}(\rho v) = 2\rho V_I, \quad (6a)$$

$$\frac{\partial v}{\partial \zeta} + \frac{\partial}{\partial \xi} \left[\frac{1}{2}v^2 + \bar{g}_0\rho + Q \right] = 0, \quad (6b)$$

where $Q = -\frac{1}{2\sqrt{\rho}}\frac{\partial^2\sqrt{\rho}}{\partial \xi^2}$ is quantum pressure, and $v = \partial\phi/\partial\xi$ is the flow velocity of the light fluid. Neglecting Q for the moment, Eq. (6b) becomes

$$\frac{\partial v}{\partial \zeta} + \frac{\partial}{\partial \xi} \left(\frac{1}{2}v^2 + \bar{g}_0\rho \right) = 0. \quad (7)$$

Equations (6a) and (7) can be cast into the diagonal Riemann form

$$\frac{\partial r_1}{\partial \zeta} + c_1 \frac{\partial r_1}{\partial \xi} = d_1, \quad (8a)$$

$$\frac{\partial r_2}{\partial \zeta} + c_2 \frac{\partial r_2}{\partial \xi} = d_2, \quad (8b)$$

where the Riemann invariant and hyperbolic speeds are defined by $r_i = v/2 \pm \sqrt{\bar{g}_0\rho}$ and $c_i = v \pm \sqrt{\bar{g}_0\rho}$, with $c_1 = (3r_1 + r_2)/2$, and $c_2 = (r_1 + 3r_2)/2$. $d_{1,2} = \pm(r_1 - r_2)V_I/2$. The light fluid intensity and the flow velocity are given by $\rho = (r_1 - r_2)^2/(4\bar{g}_0)$ and $v = r_1 + r_2$.

We numerically solve the Riemann equation without dissipative potential with the initial condition

$$\rho(\xi, 0) = \rho_b + \rho_h e^{-\xi^2/\xi_0^2}, \quad \text{and} \quad v(\xi, 0) = 0. \quad (9)$$

Here ρ_b and ρ_h are the background and hump intensity, and ξ_0 is the width of the hump. Figures 2(a) and 2(b) are the propagation of the right- and left-moving Riemann waves, respectively. The wave is stable before the shock onsets. After a critical distance (marked by stars on the figure) the shock forms. A distinctive feature is that the wave becomes steepening at the critical distance. These points are the shock wave breaking points (see more discussions in Sec. III B). These points are symmetric for the left- and right-moving solutions, as depicted in Fig. 2(c). In other words the left- and right-moving components will form shock waves after propagating equal distances.

B. Breaking point

In the local regime and without dissipation, the breaking point can be found analytically. We first linearize Eq. (8) by means of the hodograph transform (see, Refs. [33,39,71]), which treats ξ and ζ as functions of r_1 and r_2 . The transformation yields,

$$\frac{\partial \xi}{\partial r_1} - c_2 \frac{\partial \zeta}{\partial r_1} = 0, \quad \frac{\partial \xi}{\partial r_2} - c_1 \frac{\partial \zeta}{\partial r_2} = 0. \quad (10)$$

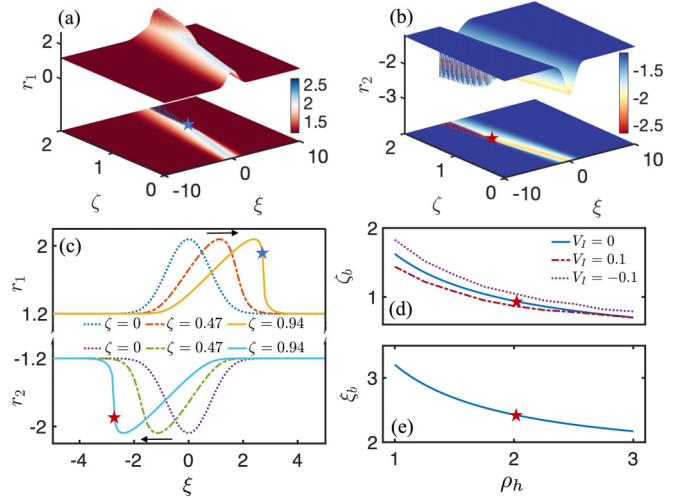


FIG. 2. (a) Right-moving and (b) left-moving Riemann waves. The stars represent the breaking position of the wave, after which the shock wave forms. (c) Riemann waves r_1 and r_2 as function as ξ for $\zeta = 0, 0.47, 0.94$. The stars represent the breaking position. Top panel: The right-moving part r_1 . Bottom panel: The left-moving part r_2 . (d) Breaking point in the ζ direction (ζ_b) as function of the hump peak intensity ρ_h . The star marks the breaking point corresponding to (a) and (b). (e) The same as (d), but for breaking point in the ξ direction (ξ_b). Other parameters are $\rho_b = 1$, $\rho_h = 2$, $\xi_0 = 1$, $\bar{g}_0 = -1.44$, and $V_I = 0$. In (d), we additionally show data with dissipative potentials.

We then introduce two functions $w_1(r_1, r_2)$ and $w_2(r_1, r_2)$ such that

$$\xi - c_1\zeta = w_1, \quad \xi - c_2\zeta = w_2. \quad (11)$$

Using the initial condition, w_1 and w_2 can be obtained,

$$w_{1,2} = \begin{cases} \xi_0 \sqrt{\ln \rho_h - \ln [r_{1,2}^2/\bar{g}_0 - \rho_b]}, & \xi > 0, \\ -\xi_0 \sqrt{\ln \rho_h - \ln [r_{1,2}^2/\bar{g}_0 - \rho_b]}, & \xi < 0. \end{cases} \quad (12)$$

Wave breaking corresponds to the occurrence of a gradient catastrophe for which $|\partial r_{1,2}/\partial \xi| \rightarrow \infty$. As the right- and left-moving wave are symmetric with respect to the ξ , we determine the breaking point of the right-moving branch explicitly. From Eq. (11), wave breaking occurs at a distance ζ such that [39]

$$\zeta = \frac{2}{3} \left| \frac{\partial w_1}{\partial r_1} \right|_{r_1=r_1^*}. \quad (13)$$

One can evaluate ζ_b approximately when the point of largest gradient in $r_1(\xi)$ lies in a region where $r_2 \approx \sqrt{\bar{g}_0\rho_b}$. In this case, Eq. (13) becomes

$$\zeta_b \approx \frac{2\xi_0\sqrt{\rho^*}}{3\sqrt{\bar{g}_0(\rho^* - \rho_b)}\sqrt{\ln[\rho_h/(\rho^* - \rho_b)]}}. \quad (14)$$

The shortest of distance ζ is reached close to the point ξ^* for which $|\partial\rho/\partial\xi|$ is maximal. We denote ξ^* the coordinate of this point in the ξ direction and $\rho^* = \rho(\xi^*)$. According to $\partial\xi/\partial\rho = \partial^2\xi/\partial\rho^2 = 0$ at $\rho = \rho^*$, it is readily to find $\ln[\rho_h/(\rho^* - \rho_b)] = \rho^*/(\rho^* + \rho_b)$. This leads to the

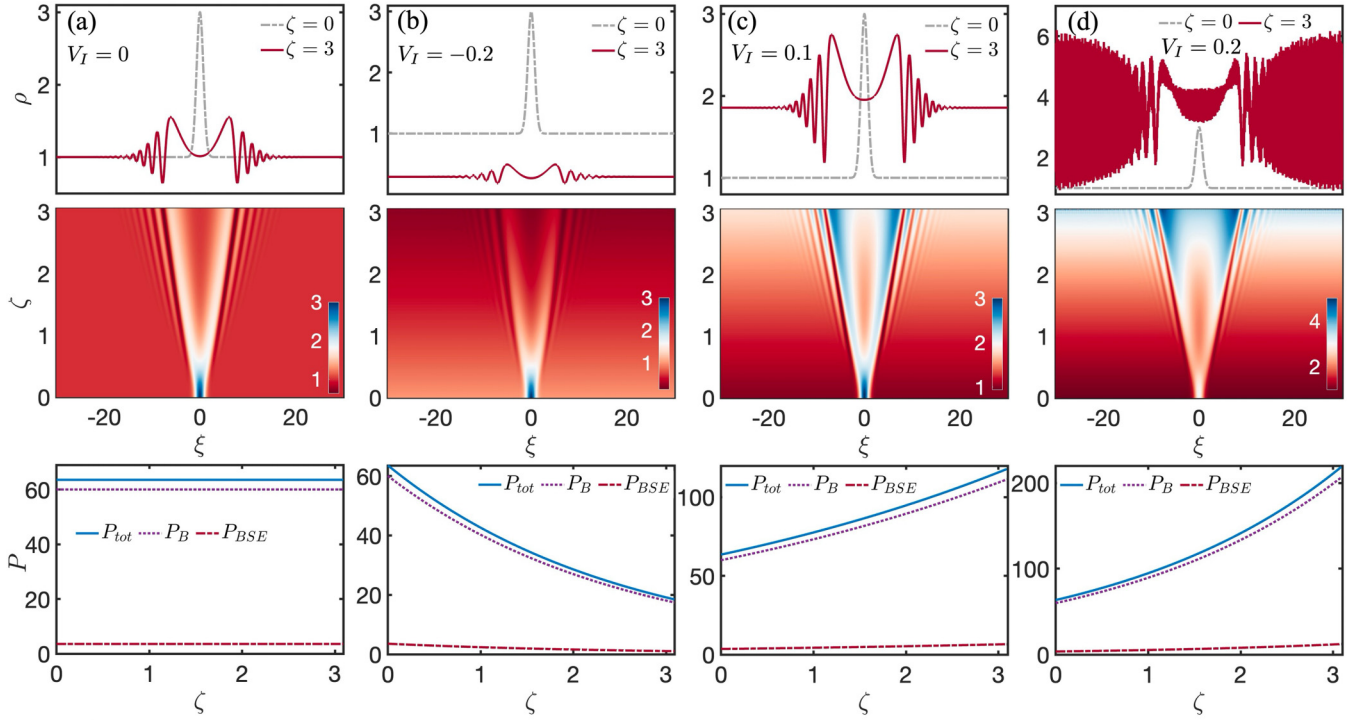


FIG. 3. (a)–(d) show the propagation of shock waves with $V_I = 0$, $V_I = -0.2$ (loss), $V_I = 0.1$, and $V_I = 0.2$ (gain). Top panel: The probe field intensity $\rho = |u|^2$ at $\zeta = 0$ and 3 (dotted-dashed gray and solid red lines). Middle panel: Wave propagation from $\zeta = 0$ to $\zeta = 3$. Bottom panel: The total power of the probe field (P_{tot} , blue solid line), the background (P_B , dotted purple line) and exchange (P_{BSE} , dotted-dashed red line) power. In all panels, the initial conditions are $\rho_b = 1$, $\rho_h = 2$, and $\xi_0 = 1$.

approximate relation,

$$\zeta_b \approx \frac{2\xi_0}{3(\rho^* - \rho_b)} \sqrt{\frac{\rho^* + \rho_b}{\bar{g}_0}}. \quad (15)$$

The breaking point as a function of ρ_h is shown in the Fig. 2(d), which matches the numerical calculation well. The breaking point ζ_b is reduced when increasing the hump intensity. Such relation is useful in controlling the generation of shock waves. For example, increasing the intensity of the hump peak allows for a shorter distance and faster visibility of the shock wave. Moreover, the breaking point ξ_b along the ξ axis when the wave breaks along the ζ direction can be obtained [40],

$$\xi_b \approx c_s(\rho^*)\zeta_b + \xi_0 \sqrt{\ln \rho_h - \ln[\rho^* - \rho_b]}. \quad (16)$$

Here $c_s = \sqrt{\bar{g}_0 \rho}$ is the local sound speed. The results are shown in Fig. 2(e), which agrees with the numerical calculation well.

Including dissipation in the Riemann function, analytical solutions are in general not possible. Instead, we find the breaking point numerically. Breaking points for $V_I = 0.1$ and -0.1 are shown in Fig. 2(d). It is found that the breaking point ζ_b decreases as the dissipative potential changes from loss to gain. These results highlight the importance of dissipative potential on the generation of shock waves. For example, the gain potential (i.e., $V_I > 0$) accelerates the generation of shock waves, as the wave breaks earlier. The loss potential ($V_I < 0$) slows down their generation.

On the other hand, the nonlocality can also affect the breaking point. To be specific, the breaking point ζ_b increases with σ . Therefore, a strong nonlocality will postpone the occurrence of shock waves [20].

C. Wave propagation

We now turn to investigate the propagation of shock waves by numerically solving Eq. (5), where the quantum pressure is taken into account explicitly. In Fig. 3 propagation of shock waves without external potential [Fig. 3(a)], in loss potential [Fig. 3(b)], and in gain potential [Figs. 3(c) and 3(d)] are shown. Without dissipation, the initial hump splits into two density peaks first. When the shock wave forms, the wave front oscillates rapidly in the ξ direction, as depicted in the top and middle panel of Fig. 3(a). The end of oscillations edge corresponds to small-amplitude edge of the shock wave [38,72]. Before the shock wave reaches the boundary, the background field is not perturbed.

In a loss potential ($V_I = -0.2$), the intensity of the background wave decays exponentially, as shown in Fig. 3(b). Shock waves form in the central region, characterized by the rapid oscillation at the shock edge. Amplitudes of the oscillation edge become smaller than that of the $V_I = 0$ case. Results of shock waves in a gain potential with $V_I = 0.1$ are shown in Fig. 3(c). From the top panel of Fig. 3(c), the intensity of the shock wave and background field, and the small-amplitude edge are all larger than the cases $V_I = 0$ and $V_I = -0.2$. This is a direct manifestation of the gain effect. On the other hand, further increasing the strength of the gain potential, the shock

wave and background field quickly become unstable, causing catastrophic collapse [38].

In the presence of the dissipative potential, power of the field will decay or grow exponentially with the propagation distance ζ . The total, background, and exchange power between shock wave and background fields are obtained, $P_{\text{tot}} = \int |u(\xi, \zeta)|^2 d\xi \approx e^{2V_I \zeta} \int |u(\xi, \zeta = 0)|^2 d\xi$, $P_B = \int |u_b(\xi, \zeta)|^2 d\xi \approx e^{2V_I \zeta} \int \rho_b d\xi$, and $P_{\text{BSE}} = P_{\text{tot}} - P_B \approx e^{2V_I \zeta} \int \rho_h \exp(-\xi^2/\xi_0^2) d\xi$. Here $u_b(\xi, \zeta) = \sqrt{\rho_b} \exp(V_I \zeta)$ is the boundary intensity at (ξ_L, ζ) . When $V_I = 0$, the total, background as well as the exchange power is conserved as a function of ζ [bottom panel of Fig. 3(a)]. Their values are determined by the initial values. On the other hand, the power will grow (decay) exponentially when $V_I > 0$ ($V_I < 0$) when propagating in the medium, as shown in the bottom panel of Figs. 3(b) and 3(c).

D. Contrast and shock width

As shown in Fig. 3, profiles of the shock wave exhibit a nontrivial dependence on the dissipative potential and initial state. Once the shock wave forms, rapid oscillations are found along the ξ axis. Including the quantum pressure, the gradient divergence of u is not available, which makes it impossible to calculate the breaking point. As shock waves oscillate rapidly, the maximal and minimal values of the oscillations provide a way to characterize the amplitude of the shock wave. Therefore we calculate the visibility of the oscillations near the soliton edge (i.e., the start of oscillation edge) of the shock wave by measuring the contrast [39,62]

$$C = \frac{\rho_{\text{max}} - \rho_{\text{min}}}{\rho_{\text{max}} + \rho_{\text{min}}}, \quad (17)$$

where ρ_{max} and ρ_{min} are the maximum and minimum values of ρ , as depicted in the left lower insert of Fig. 4(a). At a fixed propagation distance ζ , C as a function V_I is shown in Fig. 4(a). As V_I increases (from loss to gain), the contrast of the formed shock wave decreases. In Fig. 4(b), we show contrast C when varying ρ_h . For a given V_I , the contrast of the shock wave increases with increasing ρ_h . Changing V_I , such trend remains the same. These results show that we could enhance the contrast by using larger ρ_h .

We also calculate the shock wave width L_1 (measured from the center to the end of the intensity oscillation) [18,73], and the oscillation width L_2 (measured from the start to the end of the oscillation), as indicated in the insert of Fig. 4(a). We find L_1 increases with V_I , as shown in Fig. 4(c). The reason is that the small-amplitude edge has a slight increase when the potential changes from loss to gain. However, the width of oscillation L_2 (orange solid line) decreases due to the soliton edge increases.

In Fig. 4(d) width L_1 and L_2 as function of the hump peak intensity ρ_h are shown. The larger the hump density ρ_h , the wider the width L_1 and L_2 . We can understand these distances by examining the local sound speed $c_s \propto \sqrt{\rho}$. c_s is a function of local density ρ that consists of both the background and hump density. Increasing the hump density will increase the local sound speed. This means L_1 will be larger with higher ρ_h , after propagating certain ζ . The inner region defined by

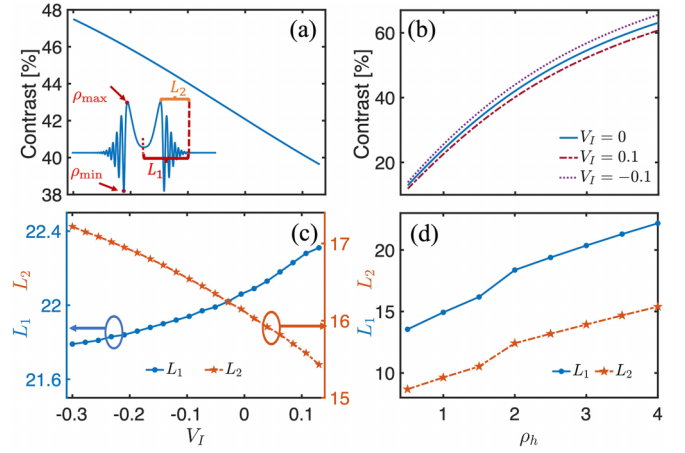


FIG. 4. (a) Oscillation contrast versus V_I . The insert illustrate the maximum and minimum values of ρ , i.e., ρ_{max} and ρ_{min} , shock width L_1 , and oscillation width L_2 . The parameters $\rho_b = 1$, $\rho_h = 2$, $\xi_0 = 1$, $\bar{g}_0 = -1.44$, and $\zeta = 3$. (b) Contrast versus ρ_h with $V_I = 0, 0.1$, and -0.1 , respectively. (c) Shock width L_1 , measured from the center to the end of oscillations, with respect to the intensity of imaginary potential V_I (blue solid line). The oscillation width L_2 , measured from the start to the end of oscillations (orange dotted-dashed line). (d) L_1 and L_2 by varying the hump peak intensity ρ_h with $V_I = 0$. The other parameters same as (a).

L_2 , on the other hand, describes propagation of solitons [38]. Its front travels at the sound speed approximately. Hence L_2 increases when c_s (ρ_h) is larger.

IV. NONLOCAL RYDBERG NONLINEARITY REGIME

We will consider the full soft-core potential using Eq. (4). To understand the role played by the nonlocality, we will solve Eq. (2) numerically by taking into the full soft-core potential Eq. (4). There are two different ways to change the soft-core potential Eq. (4). One can vary σ by fixing B_1 and B_2 , which requires us to change parameters b_1 and b_2 correspondingly. The depth of the potential [see Fig. 1(c)] is $B_1/(B_2\sigma^6) = 1/b_1$. As $b_1 = B_2\sigma^6/B_1$, the depth of the potential increases rapidly as σ decreases, which eventually makes the numerical calculation unstable. Hence σ can not be too small in practice. To avoid this divergence, one can alternatively change σ while keeping the potential depth (hence b_1) constant.

A. Fix B_1 and B_2

To avoid numerical instabilities, we have chosen the smallest $\sigma = 0.2$ in the numerical simulation. When σ is small, the formation of shock waves are featured by conspicuous oscillations, as illustrated in Fig. 5(a). The behavior is similar to that of the local nonlinearity regime. Increasing σ , the oscillation frequency decreases. As a result, the contrast increases rapidly with increasing σ , and reaches its maximum value $C \approx 0.47$ around $\sigma \approx 0.33$, as shown in Fig. 5(b). It subsequently decreases and saturates to a constant value. The width L_1 and L_2 depend on σ sensitively when σ is small. They decrease rapidly with increasing σ , as shown in Figs. 5(c) and 5(d). This dependence can be understood by analyzing the sound speed. When B_1 and B_2 are fixed, we

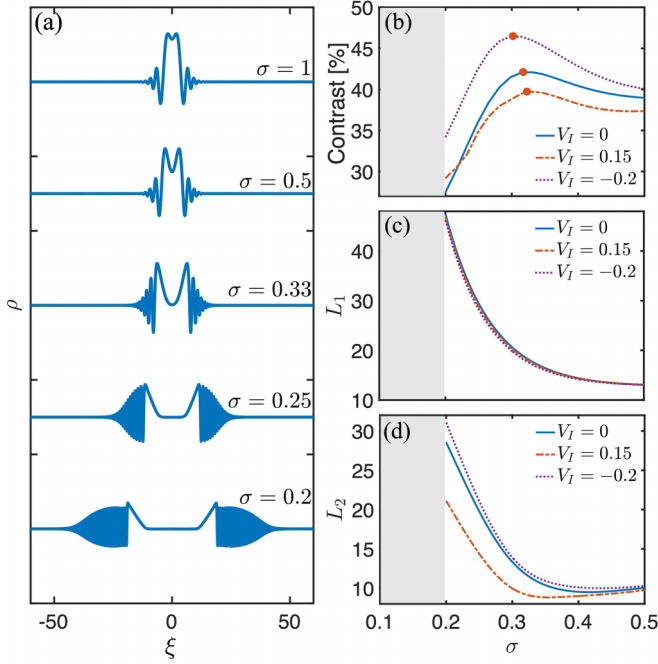


FIG. 5. (a) Shock waves for different nonlocality degree $\sigma = 0.2, 0.25, 0.33, 0.5$, and 1 (from bottom to top) with $B_1 = 0.001$, $B_2 = 0.38$, $V_I = 0$, and $\zeta = 3$. (b) The oscillation contrast as function of the nonlocality degree σ with $V_I = 0, 0.15$, and -0.2 at $\zeta = 3$. The red dot represents the maximal contrast. (c) The width L_1 and (d) oscillation width L_2 as function of the nonlocality degree σ with $V_I = 0, 0.15$, and -0.2 at $\zeta = 3$. The initial condition $\rho_h = 2$, $\rho_b = 1$, and $\xi_0 = 1$.

obtain $c_s = \sqrt{\pi g_0 \rho B_1 / (3B_2^{5/6} \sigma^5)}$. The sound speed decreases rapidly ($c_s \propto 1/\sqrt{\sigma^5}$) when σ increases. Therefore both L_1 and L_2 reduce when σ is large.

When increasing σ , we in fact drive the response from highly nonlinear to a linear regime. In other words, when σ is small, the wave dynamics is strongly nonlinear, which promotes the generation of shock waves. By increasing σ , however, the response of the Rydberg medium becomes effectively linear. When σ (R_b) is large, the soft-core potential is nearly a constant compared to the typical wavelength of the excitation. Assuming the nonlocal potential is a constant, one can carry out the integration in Eq. (2) and obtain a linear potential [73], i.e., $\int d\xi' g(\xi', \xi) |u(\xi')|^2 u(\xi) \approx g(0, 0) P_{\text{tot}} u(\xi)$. As a result, the resulting wave will propagate linearly, i.e., behaves like phonons (see Appendix B). We will focus on shock wave generation and propagation in the nonlocal regime. In practice, this requires roughly $\sigma < 0.5$, i.e., the blockade radius is half of R_0 . When $\sigma > 0.5$, the generated wave is linear and show similar propagation dynamics.

When increasing ρ_h , the contrast increases monotonically, as shown in Fig. 6(a). Both L_1 and L_2 become larger for higher ρ_h , as $L_1, L_2 \propto c_s \propto \sqrt{\rho_h}$. Similar dependence is also found in the local nonlinear case [see Figs. 4(b) and 4(d)]. The difference is that both L_1 and L_2 are slightly larger in the nonlocal Rydberg medium than that of the local medium (for given ρ_h), mainly due to that the strength of the nonlocal interaction is different in the two figures.

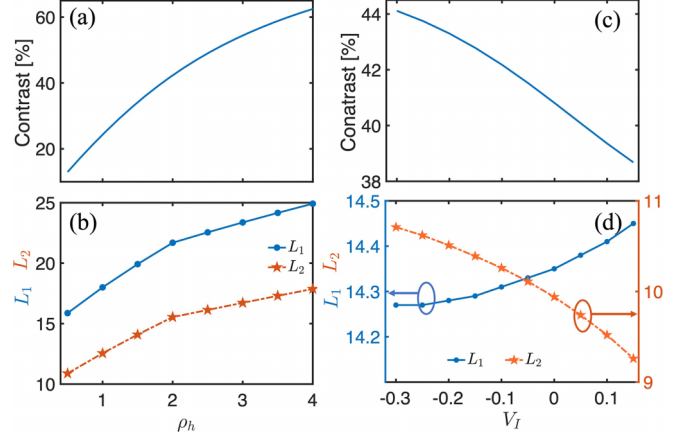


FIG. 6. (a) Oscillation contrast and (b) width L_1 and L_2 versus ρ_h with $\sigma = 0.33$ and $V_I = 0$. (c) Oscillation contrast and (d) width L_1 and L_2 versus V_I with $\sigma = 0.39$, $\rho_h = 2$ at $\zeta = 3$. The other parameters are the same as Fig. 5.

We then examine the characteristic quantities as a function of V_I numerically. The results are shown in Figs. 6(c) and 6(d). When varying V_I , the contrast, L_1 and L_2 exhibit similar dependence on V_I as found in the local nonlinearity case. Changing V_I from -0.3 to 0.15 , the contrast decreases slowly, as shown in Fig. 6(c), akin to the finding in the local regime, as shown in Fig. 4(a). Compared to the local case, shock width L_1 is barely changed when increasing V_I , as shown in Fig. 6(d). The weak dependence comes from the fact that the small-amplitude edge only has a slight change. Oscillation width L_2 decreases apparently, due to the soliton edge increases, similar to the local case shown in Fig. 4(c). The numerical data show that the contrast, L_1 and L_2 are all smaller than that of the local nonlinear case.

When varying σ , one has to change b_1 and b_2 simultaneously in order to keep B_1 and B_2 constant, as $B_1 = \sigma^6/b_2$ and $B_2 = b_1/b_2$. In practice, it would be complicated to realize such a scheme. Moreover the response function Eq. (4) decreases rapidly when increasing σ , where the nonlocal and nonlinear effect is diminished, too.

B. Fix the potential depth

To avoid the complication mentioned above, we will vary σ while keeping a fixed depth of the response function. This can be achieved by using response function Eq. (3), where b_1 and b_2 are fixed. The nonlocality degree σ can be adjusted by changing R_0 or R_b (through principal quantum number n). In the following calculation, we will assume R_0 is varied while other parameters are given previously. In this case, the sound speed $c_s \propto \sqrt{\sigma \rho}$, which means the larger σ , the more separation between the left- and right-moving shock wave, as shown in Fig. 7(a). When $\sigma = 0.5$ and 1 , we find high density peaks in the inner region, which result from that the system is effectively linear. They also affect the contrast, shown in Fig. 7(b). It decreases gradually, arrives at a minimum, and then increases again with increasing σ . For large σ , the rising contrast is purely caused by the inner peaks and the background density, where the amplitude of the shock wave is marginal. Moreover, both L_1 and L_2 increase monotonically

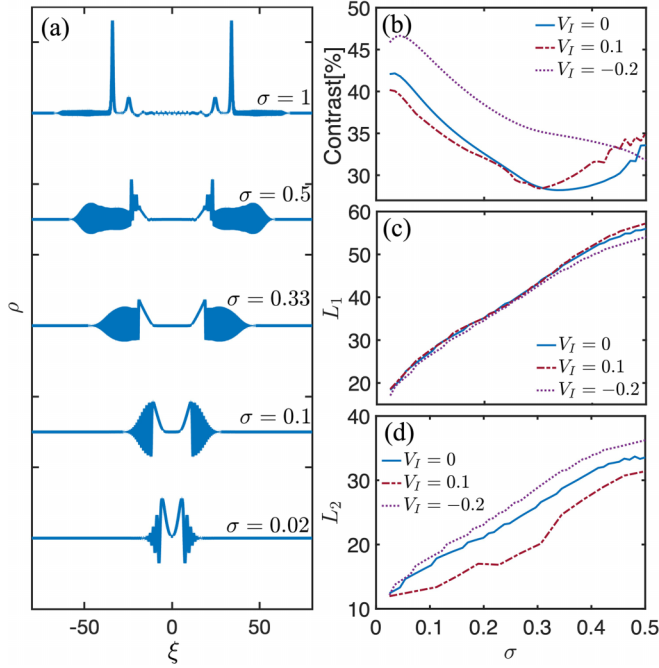


FIG. 7. (a) Shock waves for different nonlocality degree $\sigma = 0.02, 0.1, 0.33, 0.5$, and 1 , with $b_1 = 0.022, b_2 = 0.5, V_I = 0$, and $\zeta = 3$. (b) The oscillation contrast of the shock wave as function of the nonlocality degree σ with $V_I = 0, 0.1$, and -0.2 at $\zeta = 3$. (c) The shock width L_1 and (d) oscillation width L_2 as a function of σ with $V_I = 0, 0.1$, and -0.2 at $\zeta = 3$. The initial condition $\rho_h = 2, \rho_b = 1$, and $\xi_0 = 1$.

as we increase σ , due to $c_s \propto \sqrt{\sigma\rho}$. Thus after propagating distance ζ , the left- and right-moving shock waves separately significantly.

Dissipation, on the other hand, leads to a nearly global shifts to the contrast. As shown in Fig. 7(b), the contrast becomes larger when V_I is negative (loss potential). When V_I is positive, the contrast is only shifted lower slightly when $\sigma < 0.3$. The dissipation barely modifies L_1 as we increase σ [Fig. 7(c)], which shows the robustness of the shock wave propagation. For different V_I , non-negligible changes to L_2 are found when varying σ [Fig. 7(d)]. This results from the fact that the speed of the soliton is modified apparently. We want to point out that values of the contrast, L_1 and L_2 approach to those of the local regime shown in Fig. 4 (when other parameters are identical), when $\sigma \rightarrow 0$. This indicates that the local approximation is consistent with results from the general, nonlocal response function.

When increasing ρ_h , the contrast, L_1 and L_2 all increase [see Figs. 8(a) and 8(b)]. This is because not only the sound speed, but also amplitudes of the oscillation increases with larger ρ_h . Similar trends are also found in the previous case shown in Figs. 6(a) and 6(b). However, the contrast depends on V_I nontrivially in the current case. By increasing V_I , we find contrast has a minimal value C_{\min} [see Fig. 8(c)], which is different from the situation shown in Fig. 6(c). In the latter case, the contrast declines monotonically with increasing V_I in the given parameter range. C_{\min} depends on not only V_I , but also σ . We numerically obtain C_{\min} and the respective parameter V_I and σ . In the inset of Fig. 8(c), the corresponding

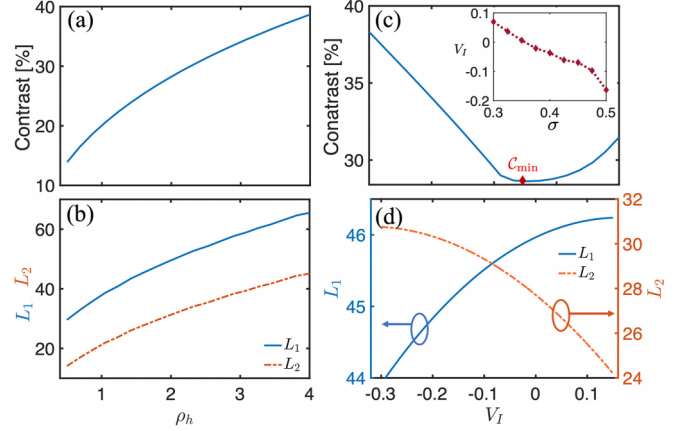


FIG. 8. (a) Oscillation contrast and (b) Shock width L_1 and oscillation width L_2 versus ρ_h with $\sigma = 0.33$ and $V_I = 0$ at $\zeta = 3$. (c) Oscillation contrast and (d) Shock width L_1 and oscillation width L_2 versus V_I with $\sigma = 0.33, \rho_h = 2$ at $\zeta = 3$. The other parameters same as Fig. 7. In (c), the inset illustrate the minimal contrast value C_{\min} versus σ and V_I .

V_I and σ are plotted. It shows that when σ increases, one has to decrease V_I in order to find C_{\min} . Finally, L_1 and L_2 as a function of V_I are shown in Fig. 8(d). The trend is similar to that of the previous case [Fig. 6(d)]. Their values are, however, larger in general. This results from the fact that the sound speed has different dependence on parameters in the two cases.

V. CONCLUSION

In this paper, we have elaborated a scheme that enables the generation and propagation of shock waves within an atomic gas involving a homogeneous dissipative potential and long-range Rydberg interaction under the condition of EIT. We have demonstrated that the homogeneous gain or loss potential significantly alters the power in the local nonlinearity regime. Both the oscillation contrast of shock waves and the oscillation width change monotonically when increasing strength of the dissipative potential and optical intensity. Different from the local regime, we have shown that in the NNL regime, the contrast of the shock wave changes nonmonotonically when increasing strength of the dissipative potential and nonlocal degree of the nonlinearity. Furthermore, the nonlocal degree of the nonlinearity modifies the oscillation amplitude and width of shock waves. Hence, the nonlocal nonlinearity can be used in controlling properties of the shock wave. Additionally, we have illustrated the hump intensity of the initial state can enhance the visibility of shock waves in both local and NNL regimes. Our results reveal the nontrivial roles of dissipation and nonlocality in the generation of shock waves, providing new routes to manipulate their profiles and stability. Our study opens new avenues for exploring non-Hermitian dynamics [74–77], and nonlinear wave dynamics [78–80] modulated by the interplay between the NNL, and local and nonlocal [59,81] dissipation in highly controllable Rydberg gases.

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APPENDIX A: EXPLICIT EXPRESSION OF THE OPTICAL BLOCH EQUATION

The dynamics of the atomic motion is governed by the optical Bloch equation

$$\frac{\partial \hat{\rho}}{\partial t} = -\frac{i}{\hbar}[\hat{H}, \hat{\rho}] - \Gamma[\hat{\rho}]. \quad (\text{A1})$$

Here $\hat{\rho}$ is the density matrix (DM) describing the atomic population and coherence, with the DM elements defined by $\rho_{\alpha\beta} \equiv \langle \hat{S}_{\alpha\beta} \rangle$; Γ is the relaxation matrix that characterizes the spontaneous emission and dephasing, whose Lindblad form is

$$\begin{aligned} \Gamma[\hat{\rho}] &= -\frac{1}{2} \sum_m (C_m^\dagger C_m \rho + \rho C_m^\dagger C_m) + \sum_m C_m \rho C_m^\dagger \\ &= \begin{pmatrix} (\Gamma_{13} - \Gamma_{21})\rho_{11} & -\frac{1}{2}\Gamma_{21}\rho_{12} & -\frac{1}{2}\Gamma_3\rho_{13} & -\frac{1}{2}(\Gamma_{34} + \Gamma_{21})\rho_{14} \\ -\frac{1}{2}\Gamma_{21}\rho_{21} & (\Gamma_{23} + \Gamma_{21})\rho_{22} & -\frac{1}{2}\Gamma_3\rho_{23} & -\frac{1}{2}\Gamma_{34}\rho_{24} \\ -\frac{1}{2}\Gamma_3\rho_{31} & -\frac{1}{2}\Gamma_3\rho_{32} & (\Gamma_{34} - \Gamma_3)\rho_{33} & -\frac{1}{2}\Gamma_3\rho_{34} \\ -\frac{1}{2}(\Gamma_{34} + \Gamma_{21})\rho_{41} & -\frac{1}{2}\Gamma_{34}\rho_{42} & -\frac{1}{2}\Gamma_3\rho_{43} & -\Gamma_{34}\rho_{44} \end{pmatrix}, \end{aligned} \quad (\text{A2})$$

here $C_{21} = \sqrt{\Gamma_{21}}|2\rangle\langle 1|$, $C_{13} = \sqrt{\Gamma_{13}}|1\rangle\langle 3|$, $C_{23} = \sqrt{\Gamma_{23}}|2\rangle\langle 3|$, and $C_{34} = \sqrt{\Gamma_{34}}|3\rangle\langle 4|$.

Based on the Hamiltonian \hat{H} given in the main text, we obtain the explicit expression of the optical Bloch equation with the following form:

$$i\frac{\partial}{\partial t}\rho_{11} - i\Gamma_{13}\rho_{33} + i\Gamma_{21}\rho_{11} + \Omega_p^*\rho_{31} - \Omega_p\rho_{13} = 0 \quad (\text{A3a})$$

$$i\frac{\partial}{\partial t}\rho_{22} - i\Gamma_{23}\rho_{33} - i\Gamma_{21}\rho_{11} + \Omega_c^*\rho_{32} - \Omega_c\rho_{23} = 0, \quad (\text{A3b})$$

$$\begin{aligned} i\frac{\partial}{\partial t}\rho_{33} - i\Gamma_{34}\rho_{44} + i\Gamma_3\rho_{33} - \Omega_p^*\rho_{31} + \Omega_p\rho_{13} - \Omega_c^*\rho_{32} \\ + \Omega_c\rho_{23} + \Omega_d^*\rho_{43} - \Omega_d\rho_{34} = 0, \end{aligned} \quad (\text{A3c})$$

$$i\frac{\partial}{\partial t}\rho_{44} + i\Gamma_{34}\rho_{44} - \Omega_d^*\rho_{43} + \Omega_d\rho_{34} = 0, \quad (\text{A3d})$$

for the diagonal elements, and

$$\left(i\frac{\partial}{\partial t} + d_{21}\right)\rho_{21} + \Omega_c^*\rho_{31} - \Omega_p\rho_{23} = 0, \quad (\text{A4a})$$

$$\left(i\frac{\partial}{\partial t} + d_{31}\right)\rho_{31} + \Omega_d^*\rho_{41} + \Omega_p(\rho_{11} - \rho_{33}) + \Omega_c\rho_{21} = 0, \quad (\text{A4b})$$

$$\left(i\frac{\partial}{\partial t} + d_{32}\right)\rho_{32} + \Omega_d^*\rho_{42} + \Omega_p\rho_{12} + \Omega_c(\rho_{22} - \rho_{33}) = 0, \quad (\text{A4c})$$

$$\begin{aligned} \left(i\frac{\partial}{\partial t} + d_{41}\right)\rho_{41} + \Omega_d\rho_{31} - \Omega_p\rho_{43} \\ - \mathcal{N}_a \int d^3\mathbf{r}' V(\mathbf{r}' - \mathbf{r})\rho_{44,41}(\mathbf{r}', \mathbf{r}, t) = 0, \end{aligned} \quad (\text{A4d})$$

$$\begin{aligned} \left(i\frac{\partial}{\partial t} + d_{42}\right)\rho_{42} + \Omega_d\rho_{32} - \Omega_c\rho_{43} \\ - \mathcal{N}_a \int d^3\mathbf{r}' V(\mathbf{r}' - \mathbf{r})\rho_{44,42}(\mathbf{r}', \mathbf{r}, t) = 0, \end{aligned} \quad (\text{A4e})$$

$$\begin{aligned} \left(i\frac{\partial}{\partial t} + d_{43}\right)\rho_{43} - \Omega_p^*\rho_{41} - \Omega_c^*\rho_{42} + \Omega_d(\rho_{33} - \rho_{44}) \\ - \mathcal{N}_a \int d^3\mathbf{r}' V(\mathbf{r}' - \mathbf{r})\rho_{44,43}(\mathbf{r}', \mathbf{r}, t) = 0, \end{aligned} \quad (\text{A4f})$$

for the nondiagonal elements. Here $d_{\alpha\beta} = \Delta_\alpha - \Delta_\beta + i\gamma_{\alpha\beta}$, $\gamma_{\alpha\beta} = (\Gamma_\alpha + \Gamma_\beta)/2$, $\Gamma_\beta = \sum_{\alpha < \beta} \Gamma_{\alpha\beta}$, with $\Gamma_{\alpha\beta}$ the spontaneous emission decay rate from $|\beta\rangle$ to $|\alpha\rangle$; Γ_{21} is the rate of population exchange between $|1\rangle$ and $|2\rangle$; $\rho_{44,4\alpha}(\mathbf{r}', \mathbf{r}, t) = \langle \hat{S}_{44}(\mathbf{r}', t)\hat{S}_{4\alpha}(\mathbf{r}, t) \rangle$ are two-body DM elements; the interaction between two Rydberg atoms, respectively, at positions \mathbf{r} and \mathbf{r}' is described by the potential $V(\mathbf{r}' - \mathbf{r}) = -\hbar C_6/|\mathbf{r}' - \mathbf{r}|^6$, with C_6 the dispersion parameter.

APPENDIX B: SOLUTION OF THE MB EQUATIONS

1. Solutions of one-body density-matrix elements

Since the probe field is much weaker than the control and dressed fields, we can take Ω_p as an expansion parameter and the perturbation expansion $\rho_{\alpha\alpha} = \rho_{\alpha\alpha}^{(0)} + \varepsilon\rho_{\alpha\alpha}^{(1)} + \varepsilon^2\rho_{\alpha\alpha}^{(2)} + \dots$ ($\alpha = 1, 2, 3, 4$), and $\rho_{\alpha\beta} = \varepsilon\rho_{\alpha\beta}^{(1)} + \varepsilon^2\rho_{\alpha\beta}^{(2)} + \dots$ ($\alpha = 2, 3, 4; \beta = 1, 2, 3; \alpha > \beta$). Substituting the above expansions into the Eqs. (A3) and (A4), we obtain a set of linear but inhomogeneous equations, which can be solved order by order [82].

At the zeroth order, the solutions read

$$\rho_{11}^{(0)} = 2\Gamma_{13}(2Z^2 + 2XY + X\Gamma_{34})/M, \quad (\text{B1a})$$

$$\rho_{22}^{(0)} = \Gamma_{21}[4(Z^2 + XY) + 2(X + Z)\Gamma_{34} - (\Gamma_{13} + \Gamma_{23})(2Y + \Gamma_{34})]/M, \quad (\text{B1b})$$

$$\rho_{33}^{(0)} = 2\Gamma_{21}[2(Z^2 + XY) + X\Gamma_{34}]/M, \quad (\text{B1c})$$

$$\rho_{44}^{(0)} = 2\Gamma_{21}[2(Z^2 + XY) - Z(\Gamma_{13} + \Gamma_{23})]/M, \quad (\text{B1d})$$

$$\rho_{32}^{(0)} = [(|\Omega_c|^2 - d_{42}d_{43})\rho_{22}^{(0)} - |\Omega_d|^2\rho_{44}^{(0)} + (|\Omega_d|^2 - |\Omega_c|^2 + d_{42}d_{43})\rho_{33}^{(0)}]\Omega_c/D_1, \quad (\text{B1e})$$

$$\rho_{42}^{(0)} = [d_{43}\rho_{22}^{(0)} - (d_{32} + d_{43})\rho_{33}^{(0)} + d_{32}\rho_{44}^{(0)}]\Omega_c\Omega_d/D_1, \quad (\text{B1f})$$

$$\rho_{43}^{(0)} = [|\Omega_c|^2\rho_{22}^{(0)} + (|\Omega_d|^2 - |\Omega_c|^2 + d_{42}d_{32})\rho_{33}^{(0)} - (|\Omega_d|^2 - d_{42}d_{32})\rho_{44}^{(0)}]/D_1, \quad (\text{B1g})$$

where $M = \Gamma_{21}[12(XY + Z^2) + 2\Gamma_{34}(2X + Z)] + \Gamma_{13}[2\Gamma_{34}X + 4XY - \Gamma_{21}(\Gamma_{34} + 2Y + 2Z) + 4Z^2] - \Gamma_{21}\Gamma_{23}[\Gamma_{34} + 2(Y + Z)]$, $X = 2\text{Im}[(d_{42}d_{43} - |\Omega_c|^2)/D_1]|\Omega_c|^2$, $Y = 2\text{Im}[(|\Omega_d|^2 - d_{42}d_{32})/D_1]|\Omega_d|^2$, $Z = 2\text{Im}[1/D_1]|\Omega_d|^2|\Omega_c|^2$, and $D_1 = d_{32}d_{42}d_{43} - d_{32}|\Omega_c|^2 - d_{43}|\Omega_d|^2$.

From the zeroth-order solution, we find that the incoherent population pumping rate Γ_{21} is a key parameter in the

zeroth-order solution. If $\Gamma_{21} = 0$, all populations are in the ground state, i.e., $\rho_{11}^{(0)} = 1$, and other state population and coherence are both zero, $\rho_{\alpha\beta}^{(0)} = 0$. However, when $\Gamma_{21} \neq 0$, we have $\rho_{33}^{(0)} \neq 0$, and hence a gain to the probe field will be realized when the probe field is coupled to the states $|1\rangle$ and $|3\rangle$.

First-order solutions. At the first order, the solutions of $\rho_{21}^{(1)}$, $\rho_{31}^{(1)}$, and $\rho_{41}^{(1)}$ are given by

$$\rho_{21}^{(1)} = [(d_{31}d_{41} - |\Omega_d|^2)\rho_{32}^{*(0)} - d_{41}\Omega_c(\rho_{33}^{(0)} - \rho_{11}^{(0)}) + \Omega_d\Omega_c\rho_{43}^{(0)}]/D_2\Omega_p \equiv a_{21}^{(1)}\Omega_p, \quad (\text{B2a})$$

$$\rho_{31}^{(1)} = [-\Omega_c d_{41}\rho_{32}^{*(0)} + d_{21}d_{41}(\rho_{33}^{(0)} - \rho_{11}^{(0)}) - d_{21}\Omega_d\rho_{43}^{(0)}]/D_2\Omega_p \equiv a_{31}^{(1)}\Omega_p, \quad (\text{B2b})$$

$$\rho_{41}^{(1)} = [\Omega_c\Omega_d\rho_{32}^{*(0)} - d_{21}\Omega_d(\rho_{33}^{(0)} - \rho_{11}^{(0)}) + (d_{21}d_{31} - |\Omega_c|^2)\rho_{43}^{(0)}]/D_2\Omega_p \equiv a_{41}^{(1)}\Omega_p, \quad (\text{B2c})$$

where $D_2 = d_{31}|\Omega_c|^2 + d_{21}|\Omega_d|^2 - d_{21}d_{31}d_{41}$. Other $\rho_{\alpha\beta}^{(1)}$ are zero.

Second-order solutions. At the second order, the matrix elements can be solved by the equation

$$\begin{pmatrix} i\Gamma_{21} & 0 & -i\Gamma_{13} & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\ -i\Gamma_{21} & 0 & -i\Gamma_{23} & 0 & 0 & 0 & 0 & 0 & \Omega_c^* & -\Omega_c \\ 0 & 0 & 0 & i\Gamma_{34} & -\Omega_d^* & \Omega_d & 0 & 0 & 0 & 0 \\ 0 & \Omega_c & -\Omega_c & 0 & 0 & 0 & \Omega_d^* & 0 & d_{32} & 0 \\ 0 & 0 & 0 & 0 & -\Omega_c & 0 & d_{42} & 0 & \Omega_d & 0 \\ 0 & 0 & \Omega_d & -\Omega_d & d_{43} & 0 & -\Omega_c^* & 0 & 0 & 0 \\ 0 & \Omega_c^* & -\Omega_c^* & 0 & 0 & 0 & 0 & \Omega_d & 0 & d_{32}^* \\ 0 & 0 & 0 & 0 & 0 & -\Omega_c^* & 0 & d_{42}^* & 0 & \Omega_d^* \\ 0 & 0 & \Omega_d^* & -\Omega_d^* & 0 & d_{43}^* & 0 & -\Omega_c & 0 & 0 \\ 1 & 1 & 1 & 1 & 0 & 0 & 0 & 0 & 0 & 0 \end{pmatrix} \begin{pmatrix} \rho_{11}^{(2)} \\ \rho_{22}^{(2)} \\ \rho_{33}^{(2)} \\ \rho_{44}^{(2)} \\ \rho_{43}^{(2)} \\ \rho_{34}^{(2)} \\ \rho_{42}^{(2)} \\ \rho_{24}^{(2)} \\ \rho_{32}^{(2)} \\ \rho_{23}^{(2)} \end{pmatrix} = \begin{pmatrix} 2i\text{Im}[\Omega_p\rho_{13}^{(1)}] \\ 0 \\ 0 \\ -\Omega_p\rho_{12}^{(1)} \\ 0 \\ \Omega_p^*\rho_{41}^{(1)} \\ -\Omega_p^*\rho_{21}^{(1)} \\ 0 \\ \Omega_p\rho_{14}^{(1)} \\ 0 \end{pmatrix}. \quad (\text{B3})$$

Solving the matrix equations above yields $\rho_{\alpha\beta}^{(2)} = a_{\alpha\beta}^{(2)}|\Omega_p|^2$ with the coefficients $a_{\alpha\beta}$ being the function detuning Δ_α , spontaneous emission decay rate $\Gamma_{\alpha\beta}$, and half-Rabi frequencies Ω_d , Ω_c .

Third-order solutions. At the third order, the solutions of $\rho_{21}^{(3)}$, $\rho_{31}^{(3)}$, and $\rho_{41}^{(3)}$ can be obtained from the equations

$$\begin{pmatrix} d_{21} & \Omega_c^* & 0 \\ \Omega_c & d_{31} & \Omega_d^* \\ 0 & \Omega_d & d_{41} \end{pmatrix} \begin{pmatrix} \rho_{21}^{(3)} \\ \rho_{31}^{(3)} \\ \rho_{41}^{(3)} \end{pmatrix} = \begin{pmatrix} a_{23}^{(2)} \\ a_{33}^{(3)} - a_{11}^{(2)} \\ a_{43}^{(2)} \end{pmatrix} |\Omega_p|^2\Omega_p + (0 \quad 0 \quad \mathbb{A})^T, \quad (\text{B4})$$

where $\mathbb{A} = \mathcal{N}_a \int d\mathbf{r}' V(\mathbf{r}' - \mathbf{r}) a_{44,41} |\Omega_p(\mathbf{r}')|^2 \Omega_p$. Expressions of $\rho_{31}^{(3)}$ at the third order is obtained from Eq. (B4)

$$\rho_{31}^{(3)} = a_{31}^{(3)} |\Omega_p|^2 \Omega_p + \mathcal{N}_a \int d^3\mathbf{r}' V(\mathbf{r}' - \mathbf{r}) b_{31}^{(3)} |\Omega_p(\mathbf{r}')|^2 \Omega_p, \quad (\text{B5})$$

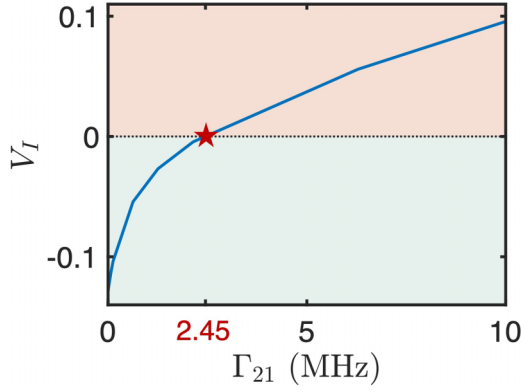


FIG. 9. (a) The imaginary potential as function as incoherent pumping Γ_{21} . When $\Gamma_{21} < 2.45$ MHz, the potential is a loss. Otherwise, the potential is a gain. The red star represents $V_I = 0$ when $\Gamma_{21} = 2.45$ MHz.

where $a_{31}^{(3)} = [d_{21}d_{41}(a_{33}^{(2)} - a_{11}^{(2)}) - d_{41}\Omega_c a_{32}^{*(2)} - d_{21}\Omega_d^* a_{43}^{(2)}] / D_2$ and $b_{31}^{(3)} = d_{21}\Omega_d^* a_{44,41}^{(3)} / D_2$. Combining first three order solutions of ρ_{31} with Maxwell equation, we obtain the nonlocal nonlinear Schrödinger equation

$$i\frac{\partial\Omega_p}{\partial z} + \frac{c}{2\omega_p}\nabla_{\perp}^2\Omega_p - V_I\Omega_p + W|\Omega_p|^2\Omega_p + \int d\mathbf{r}' G(\mathbf{r}', \mathbf{r})|\Omega_p(\mathbf{r}')|^2\Omega_p = 0, \quad (\text{B6})$$

$$\begin{pmatrix} 2d_{41} - V & 2\Omega_d & 0 & 0 & 0 & 0 \\ \Omega_d^* & d_{41} + d_{31} & \Omega_d & \Omega_c & 0 & 0 \\ 0 & \Omega_d^* & d_{31} & 0 & \Omega_c & 0 \\ 0 & \Omega_c^* & 0 & d_{41} + d_{21} & \Omega_d & 0 \\ 0 & 0 & \Omega_c^* & \Omega_d^* & d_{31} + d_{21} & \Omega_c \\ 0 & 0 & 0 & 0 & \Omega_c^* & d_{21} \end{pmatrix} \begin{pmatrix} \rho_{41,41}^{(2)} \\ \rho_{41,31}^{(2)} \\ \rho_{31,31}^{(2)} \\ \rho_{41,21}^{(2)} \\ \rho_{31,21}^{(2)} \\ \rho_{21,21}^{(2)} \end{pmatrix} = \begin{pmatrix} 2a_{43}^{(0)}a_{41}^{(1)} \\ (a_{33}^{(0)} - a_{11}^{(0)})a_{41}^{(1)} + a_{43}^{(0)}a_{31}^{(1)} \\ (a_{33}^{(0)} - a_{11}^{(0)})a_{31}^{(1)} \\ a_{32}^{*(0)}a_{41}^{(1)} + a_{43}^{(0)}a_{21}^{(1)} \\ (a_{33}^{(0)} - a_{11}^{(0)})a_{21}^{(1)} + a_{23}^{(0)}a_{31}^{(1)} \\ a_{32}^{*(0)}a_{21}^{(1)} \end{pmatrix} \Omega_p^2. \quad (\text{B8})$$

The solution for $\rho_{41,41}^{(2)} = a_{41,41}^{(2)}(\mathbf{r}' - \mathbf{r})\Omega_p^2(\mathbf{r}')$ with

$$a_{41,41}^{(2)}(\mathbf{r}' - \mathbf{r}) = \frac{P_0}{P_1 + P_2V(\mathbf{r}' - \mathbf{r})}, \quad (\text{B9})$$

where P_0 , P_1 , and P_2 are the functions of Ω_d , Ω_c , Δ_{α} , and $\Gamma_{\alpha\beta}$.

The two-body equations for $\rho_{44,41}^{(3)}$ are 27 order linear equations, which are very lengthy and hence are omitted here, and solution has the form

$$\rho_{44,41}^{(3)}(\mathbf{r}' - \mathbf{r}) = \frac{\sum_{n=0}^2 P_n V^n(\mathbf{r}' - \mathbf{r})}{\sum_{n=0}^3 Q_n V^n(\mathbf{r}' - \mathbf{r})} |\Omega_p(\mathbf{r}')|^2 \Omega_p(\mathbf{r}) \approx \rho_{41,41}^{(2)} \rho_{14}^{(1)} \equiv a_{44,41}^{(3)} |\Omega_p|^2 \Omega_p, \quad (\text{B10})$$

where P_n , Q_n are the functions of Ω_d , Ω_c , Δ_{α} , and $\Gamma_{\alpha\beta}$.

Note that the response function in dimensionless NNLS equation Eq. (2) can be obtained from the two-body correlators $\rho_{44,41}^{(3)}$. The expression reads

where $V_I = -\kappa_{13}a_{31}^{(1)}$ is linear potential, $W = \kappa_{13} [d_{21}d_{41}(a_{33}^{(2)} - a_{11}^{(2)}) - d_{41}\Omega_c a_{32}^{*(2)} - d_{21}\Omega_d^* a_{43}^{(2)}] / D_2$ is the local nonlinear coefficient, and the nonlocal response function $G(\mathbf{r}', \mathbf{r}) = (\kappa_{13}d_{21}\Omega_d^* \mathcal{N}_a / D_2) V(\mathbf{r}' - \mathbf{r}) a_{44,41}^{(3)}(\mathbf{r}' - \mathbf{r})$. Note that the two-body equations $a_{44,41}^{(3)}$ should be solved first in order to solve the nonlocal nonlinear Schrödinger equation above.

In dimensionless NNLS equation Eq. (2), the linear potential reads

$$\mathcal{V} = -2\kappa_{13}L_{\text{diff}} \left[-\Omega_c d_{41} \rho_{32}^{*(0)} + d_{21} d_{41} (\rho_{33}^{(0)} - \rho_{11}^{(0)}) - d_{21} \Omega_d \rho_{43}^{(0)} \right] / D_2 \equiv V_R + iV_I. \quad (\text{B7})$$

Here, V_R and V_I are the real and imaginary parts of linear potential, respectively. V_R is a constant and independent on the intensity of the shock wave, and is indeed negligible. $V_I < 0$ show a loss to the probe field due to the dissipation of system when $\rho_{33}^{(0)} = 0$ in the zeroth order for $\Gamma_{21} = 0$. With increasing of incoherent pumping Γ_{21} , $\rho_{33}^{(0)} \neq 0$ and a gain to the probe field will be realized. As results show in Fig. 9, a loss or gain potential V_I can be realized by adjusting the Γ_{21} . When $\Gamma_{21} < 2.45$ MHz, the potential is a loss. Otherwise, the potential is a gain.

2. Solutions of two-body density-matrix elements

We solve the second-order solution of the correlators of $\rho_{41,41}^{(2)}$, which can be obtained from

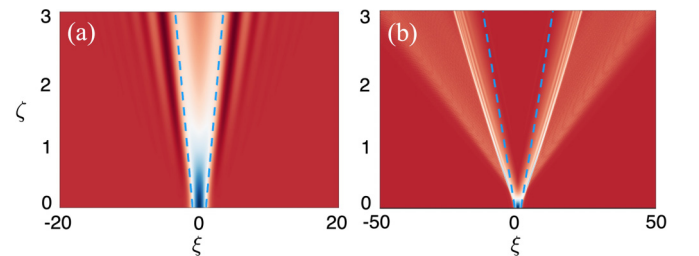


FIG. 10. (a) Wave propagation when $\sigma = 0.5$. The dashed line marks the trajectory of the sound wave. It is clear that the wave propagates at the speed of sound c_s , indicating the system is in the linear regime. Other parameters are same as in Fig. 5(a), corresponding to the linear propagation discussed in Sec. IV A. (b) The same as (a). Other parameters are same as in Fig. 7(a), corresponding to the linear propagation discussed in Sec. IV B.

$g = 2L_{\text{diff}}U_0^2R_0^2\kappa_{13}d_{21}\Omega_d^*\mathcal{N}_a/D_2 \int V(\mathbf{r}' - \mathbf{r}) a_{44,41}^{(3)} dy'$. It can be approximated by analytical form [30,38]

$$g(\xi', \xi) \approx - \int \frac{1}{b_1 + \frac{b_2}{\sigma^6}[(\xi' - \xi)^2 + (y'/R_0)^2]^3} dy' \\ \approx - \frac{B_1}{B_2\sigma^6 + |\xi' - \xi|^6}, \quad (\text{B11})$$

where $\sigma = R_b/R_0$ characterizes the nonlocality degree of the nonlinearity. b_1 and b_2 are the coefficients determined by laser parameters (i.e., Ω_d , Ω_c , Δ_α , and $\Gamma_{\alpha\beta}$). The relation between $B_{1,2}$ and $b_{1,2}$ are $B_1 = \sigma^6/b_2$ and $B_2 = b_1/b_2$.

As discussed in the main text, the response becomes linear when R_b is comparable to the beam radius R_0 , where the soft-core potential is nearly a constant compared to the wavelength of the excitation. Hence the wave behaves like phonons and propagates linearly, as depicted in Fig. 10.

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