


Nuclear mass and size corrections to the magnetic shielding

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We derive finite nuclear mass and finite nuclear size corrections to the magnetic shielding in light ions. These corrections are important for the accurate determination of nuclear magnetic moments. We correct several previous formulas for the nuclear mass corrections and present improved results for the magnetic shielding in ^1H , $^3\text{He}^+$, and ^3He . Finally, we obtain an ^3He atomic magnetic moment, which serves as an accurate probe to measure magnetic fields.

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I. INTRODUCTION

The nuclear magnetic moment in atoms is partially shielded by atomic electrons. This effect is not very significant: about 10^{-5} for light elements. Nevertheless, because nuclear magnetic moments are determined from the Zeeman shift in atomic systems, the calculation of the magnetic shielding is necessary for their accurate determination. For example, the recent measurement of the magnetic moment of the $^3\text{He}^+$ ion [1] together with the calculation of magnetic shielding [2,3] allowed for the most accurate determination so far of the helion magnetic moment. A similar measurement is planned for $^9\text{Be}^{3+}$, which will result in an improved determination of the ^9Be nuclear magnetic moment [4]. Moreover, accurate values for nuclear magnetic moments are important for the determination of atomic hyperfine splitting (HFS), testing quantum electrodynamics (QED), and the nuclear structure theory [5]. This is because HFS is very sensitive to the distribution of the magnetic moment within the nucleus.

In this work, we point out two interesting effects that are frequently overlooked in calculations of nuclear magnetic shielding [6], namely, those due to the finite nuclear mass and the finite nuclear size. Nuclear mass corrections are as large as relativistic corrections for light atomic systems, while finite nuclear size effects are much smaller, but they are expected to be significant for heavier elements. These finite nuclear mass effects have already been the subject of several works [7,8]. Here, we rederive them thoroughly, correct some mistakes, and update numerical values for the most relevant cases of the H, $^3\text{He}^+$, and ^3He elements. The finite nuclear size effects have been studied only numerically and only for hydrogen-like systems [9,10]. Here, we derive a compact analytic formula in terms of the charge, magnetic, and effective Zemach nuclear radii, which accounts also for nuclear inelastic effects.

II. BREIT-PAULI HAMILTONIAN WITH THE HOMOGENOUS MAGNETIC FIELD

To account for finite nuclear mass effects, we have to treat nuclei on an equal footing with all electrons. Therefore, we

consider a system of charged particles, each having its own mass m_a , charge e_a , spin s_a , and the so-called g-factor g_a , which is related to the magnetic moment by

$$\vec{\mu}_a = \frac{g_a e_a}{2 m_a} \vec{s}_a. \quad (1)$$

These particles are electrons with spin 1/2 and nuclei with an arbitrary spin. Our derivation employs a Breit-Pauli Hamiltonian with homogenous magnetic field and with separation of center of mass motion. It closely follows the lines of Ref. [7]. Let us therefore introduce the total mass M

$$M = \sum_a m_a, \quad (2)$$

center of mass variables

$$\vec{R} = \sum_a \frac{m_a}{M} \vec{r}_a, \quad (3)$$

$$\vec{P} = \sum_a \vec{p}_a, \quad (4)$$

and relative coordinates

$$\vec{x}_a = \vec{r}_a - \vec{R}, \quad (5)$$

$$\vec{q}_a = \vec{p}_a - \frac{m_a}{M} \vec{P}, \quad (6)$$

such that

$$[x_a^i, q_b^j] = i \delta^{ij} \left(\delta_{ab} - \frac{m_b}{M} \right), \quad (7)$$

$$[R^i, P^j] = i \delta^{ij}, \quad (8)$$

$$[x_a^i, P^j] = [R^i, q_a^j] = 0. \quad (9)$$

The Hamiltonian of a bound system of charged particles in an external magnetic field including leading relativistic corrections and with the separated-out center of mass motion is [7]

$$\begin{aligned}
H_{\text{in}} = & \sum_a \left\{ \frac{\vec{\pi}_a^2}{2m_a} - \frac{e_a}{2m_a} g_a \vec{s}_a \cdot \vec{B} - \frac{\vec{\pi}_a^4}{8m_a^3} + \frac{e_a}{8m_a^3} \left[4\vec{\pi}_a^2 \vec{s}_a \cdot \vec{B} + (g_a - 2) \{ \vec{\pi}_a \cdot \vec{B}, \vec{\pi}_a \cdot \vec{s}_a \} \right] - \frac{e_a^2}{2} \chi_a \vec{B}^2 \right\} \\
& + \sum_{a>b,b} \frac{e_a e_b}{4\pi} \left\{ \frac{1}{r_{ab}} - \frac{1}{2m_a m_b} \pi_a^i \left(\frac{\delta^{ij}}{r_{ab}} + \frac{r_{ab}^i r_{ab}^j}{r_{ab}^3} \right) \pi_b^j - \frac{2\pi}{3} \langle r_{Ea}^2 + r_{Eb}^2 \rangle \delta^3(r_{ab}) - \frac{2\pi g_a g_b}{3m_a m_b} \vec{s}_a \cdot \vec{s}_b \delta^3(r_{ab}) \right. \\
& \left. + \frac{g_a g_b}{4m_a m_b} \frac{s_a^i s_b^j}{r_{ab}^3} \left(\delta^{ij} - 3 \frac{r_{ab}^i r_{ab}^j}{r_{ab}^2} \right) \right\} + \sum_{a,b} \frac{e_a e_b}{4\pi} \frac{1}{2r_{ab}^3} \left[\frac{g_a}{m_a m_b} \vec{s}_a \cdot \vec{r}_{ab} \times \vec{\pi}_b - \frac{(g_a - 1)}{m_a^2} \vec{s}_a \cdot \vec{r}_{ab} \times \vec{\pi}_a \right], \quad (10)
\end{aligned}$$

where $r_{ab} = |\vec{r}_a - \vec{r}_b|$, and

$$\vec{\pi}_a = \vec{q}_a + \frac{1}{2} \vec{D}_a \times \vec{B}, \quad (11)$$

$$\vec{D}_a = e_a \vec{x}_a + \frac{m_a}{M} \vec{D}, \quad (12)$$

$$\vec{D} = \sum_a e_a \vec{x}_a. \quad (13)$$

For a point spin $s = 1/2$ particle, $g = 2$, $\langle r_E^2 \rangle = 3/(4m^2)$, and $\chi = 1/(4m^3)$. For a finite size particle $\langle r_E^2 \rangle$ includes the mean square charge radius. An equivalent Hamiltonian for a system of spin $1/2$ point particles was originally obtained by Hegstrom in Ref. [11]. Our Hamiltonian in Eq. (10), however, is valid for arbitrary spin particles and has a more compact form.

The magnetic interaction resulting from H_{in} neglecting the terms quadratic in \vec{B} is

$$\begin{aligned}
\delta H = & - \sum_a \frac{e_a}{2m_a} (\vec{x}_a \times \vec{q}_a + g_a \vec{s}_a) \cdot \vec{B} \\
& + \sum_a \frac{1}{4m_a^3} \left[q_a^2 \vec{D}_a \times \vec{q}_a \cdot \vec{B} + 2e_a q_a^2 \vec{s}_a \cdot \vec{B} \right. \\
& \left. + e_a (g_a - 2) \vec{q}_a \cdot \vec{s}_a \vec{q}_a \cdot \vec{B} \right] + \sum_{a \neq b, b} \frac{e_a e_b}{4\pi} \\
& \times \left[- \frac{1}{4m_a m_b} q_a^i \left(\frac{\delta^{ij}}{r_{ab}} + \frac{r_{ab}^i r_{ab}^j}{r_{ab}^3} \right) (\vec{D}_b \times \vec{B})^j \right. \\
& \left. + \frac{1}{4r_{ab}^3} \frac{g_a}{m_a m_b} (\vec{s}_a \times \vec{r}_{ab}) \cdot (\vec{D}_b \times \vec{B}) \right. \\
& \left. - \frac{1}{4r_{ab}^3} \frac{(g_a - 1)}{m_a^2} (\vec{s}_a \times \vec{r}_{ab}) \cdot (\vec{D}_a \times \vec{B}) \right]. \quad (14)
\end{aligned}$$

This is a general interaction Hamiltonian, which is valid for an arbitrary set of particles. In particular, one can obtain the bound electron g-factor or the magnetic shielding in atomic and molecular systems. In the next section we derive the atomic magnetic shielding with full account of the nuclear mass.

III. FINITE NUCLEAR MASS CORRECTIONS

We will derive the magnetic shielding constant for arbitrary ions with the vanishing orbital angular momentum \vec{L} . The interaction of the nuclear spin with the magnetic field is

obtained from Eq. (14) as

$$\begin{aligned}
\delta H = & - \frac{e_N}{2m_N} g_N \vec{s}_N \cdot \vec{B} \\
& + \frac{e_N}{4m_N^3} \left[2q_N^2 \vec{s}_N \cdot \vec{B} + (g_N - 2) \vec{q}_N \cdot \vec{s}_N \vec{q}_N \cdot \vec{B} \right] \\
& + \sum_b' \frac{e_N e}{4\pi} \frac{(\vec{s}_N \times \vec{r}_{Nb})}{4r_{Nb}^3} \\
& \cdot \left[\frac{g_N}{m_N m_e} \vec{D}_b \times \vec{B} - \frac{(g_N - 1)}{m_N^2} \vec{D}_N \times \vec{B} \right], \quad (15)
\end{aligned}$$

where we assumed that for nucleus $a = N$, all other particles are electrons, and \sum' denotes summation over electrons only. For an electronic state with a spherical symmetry

$$(\vec{s}_N \times \vec{X}) \cdot (\vec{Y} \times \vec{B}) = -\frac{2}{3} \vec{s}_N \cdot \vec{B} \vec{X} \cdot \vec{Y}, \quad (16)$$

one introduces the scalar shielding constant σ

$$\delta H = -\frac{g_N e_N}{2m_N} \vec{s}_N \cdot \vec{B} (1 - \sigma). \quad (17)$$

This σ is conveniently split into two parts, consisting of the first- and the second-order matrix elements

$$\sigma = \sigma_1 + \sigma_2. \quad (18)$$

σ_1 results from the first-order matrix element of δH in Eq. (15),

$$\begin{aligned}
\sigma_1 = & \frac{1}{2g_N m_N^2} \left[2 + \frac{(g_N - 2)}{3} \right] \langle q_N^2 \rangle \\
& + \frac{1}{3} \frac{e_e}{4\pi} \left\langle \sum_b' \frac{\vec{r}_{bN}}{r_{bN}^3} \cdot \left[\frac{1}{m_e} \vec{D}_b - \frac{(g_N - 1)}{g_N m_N} \vec{D}_N \right] \right\rangle. \quad (19)
\end{aligned}$$

Because $\sum_a m_a \vec{x}_a = 0$, the position \vec{x}_N of the nucleus with respect to mass center and the dipole operator \vec{D} can be expressed in terms of the electron coordinates only,

$$\vec{x}_N = -\frac{m_e}{M} \sum_a' \vec{r}_{aN}, \quad (20)$$

$$\vec{D} = e_e \sum_a' \vec{r}_{aN} \left(1 + (Z - N_e) \frac{m_e}{M} \right), \quad (21)$$

where $M = m_N + N_e m_e$, N_e is the number of electrons, and Z is the nuclear charge in units of the elementary charge.

Consequently, the shielding constant σ_1 takes the form

$$\begin{aligned} \sigma_1 &= \frac{\alpha}{3m_e} \left\langle \sum_a' \frac{1}{r_a} \right\rangle + \frac{(4+g_N) \langle p_N^2 \rangle}{6g_N m_N^2} \\ &+ \frac{\alpha}{3} \left\langle \sum_b \frac{\vec{r}_{bN}}{r_{bN}^3} \cdot \sum_a' \vec{r}_{aN} \right\rangle \frac{1}{g_N} \frac{m_e}{M} \\ &\times \left[\frac{(Z-N_e)}{M} + (1-g_N) \left(\frac{1}{m_e} + \frac{Z}{m_N} \right) \right], \end{aligned} \quad (22)$$

where $\vec{p}_N = -\sum_a' \vec{p}_a$, and we used $\vec{P}|\phi\rangle = 0$. Consider now the following matrix element:

$$\begin{aligned} \left\langle \sum_a' \frac{\vec{r}_{aN}}{r_{aN}^3} \cdot \sum_b \vec{r}_{bN} \right\rangle &= \frac{1}{iZ\alpha} \left\langle [\vec{p}_N, H - E] \sum_b \vec{r}_{bN} \right\rangle \\ &= \sum_b' \frac{1}{iZ\alpha} \langle \vec{p}_N [H - E, \vec{r}_{bN}] \rangle \\ &= \langle p_N^2 \rangle \frac{1}{Z\alpha} \frac{M}{m_N m_e}, \end{aligned} \quad (23)$$

which is used to simplify σ_1 ,

$$\begin{aligned} \sigma_1 &= \frac{\alpha}{3m_e} \left\langle \sum_a' \frac{1}{r_a} \right\rangle + \frac{\langle p_N^2 \rangle}{3g_N m_N^2} \left[3 - \frac{g_N}{2} \right. \\ &\left. + \frac{m_N}{M} \left(1 - \frac{N_e}{Z} \right) + (1-g_N) \frac{m_N}{Z m_e} \right]. \end{aligned} \quad (24)$$

The σ_2 part is given by the second-order interaction coming from the Hamiltonian

$$\begin{aligned} \delta H &= -\sum_a \frac{e_a}{2m_a} \vec{x}_a \times \vec{q}_a \cdot \vec{B} - \frac{e_N e_e}{4\pi} \frac{\vec{s}_N}{2m_N} \\ &\cdot \sum_b \frac{\vec{r}_{bN}}{r_{bN}^3} \times \left[g_N \frac{\vec{q}_b}{m_e} - (g_N - 1) \frac{\vec{q}_N}{m_N} \right], \end{aligned} \quad (25)$$

namely,

$$\begin{aligned} \sigma_2 &= \frac{2}{3} \left\langle \sum_a \frac{e_a}{2m_a} \vec{x}_a \times \vec{q}_a \frac{1}{(E-H)} \sum_b' \frac{e_e}{4\pi} \frac{\vec{r}_{bN}}{r_{bN}^3} \right. \\ &\left. \times \left[\frac{\vec{p}_b}{m_b} - \frac{(g_N-1)}{g_N} \frac{\vec{p}_N}{m_N} \right] \right\rangle. \end{aligned} \quad (26)$$

Using

$$\begin{aligned} \sum_a \frac{e_a}{2m_a} \vec{x}_a \times \vec{q}_a &= \frac{e_e}{2m_e} \vec{L} + \left(\frac{e_e}{2m_e} + \frac{Z e_e}{2m_N} \right) \\ &\times \frac{m_e}{M} \sum_a' \vec{r}_{aN} \times \vec{p}_N, \end{aligned} \quad (27)$$

we arrive at

$$\begin{aligned} \sigma_2 &= \frac{\alpha}{3M} \left(1 + \frac{Z m_e}{m_N} \right) \left\langle \sum_a' \vec{r}_{aN} \times \vec{p}_N \frac{1}{(E-H)} \right. \\ &\left. \times \sum_b \frac{\vec{r}_{bN}}{r_{bN}^3} \left[\frac{\vec{p}_b}{m_e} - \frac{(g_N-1)}{g_N} \frac{\vec{p}_N}{m_N} \right] \right\rangle. \end{aligned} \quad (28)$$

The total shielding constant is $\sigma = \sigma_1 + \sigma_2$, where σ_1 is given in Eq. (24) and σ_2 in Eq. (28). For the numerical cal-

TABLE I. Contributions to the shielding constant $10^6\sigma$, for ^1H , $^3\text{He}^+$, and ^3He using Refs. [3,8]. The new results are $\sigma^{(2,1)}(\text{He})$, $\sigma^{(2,2)}$, $\sigma^{(6)}$, and σ_{fs} . Because the direct numerical calculation of QED corrections to $\sigma^{(6)}$ is not sufficiently accurate for low Z [9,10], we estimate uncertainty from QED corrections at this order by assuming that it does not exceed the known relativistic contribution to σ^6 . σ_{fs} was calculated using $r_C(p) = r_M(p) = 0.84$ fm [12], $\tilde{r}_Z(p) = 0.87$ fm [5], $r_C(h) = r_M(h) = 1.97$ fm [13], and $\tilde{r}_Z(h) = 2.60$ fm, and other physical constants are from Ref. [14].

	^1H	$^3\text{He}^+$	^3He
$\sigma^{(2,0)}$	17.750 451 5	35.500 903 0	59.936 771 0
$\sigma^{(2,1)}$	-0.017 603 7	-0.013 933 4	-0.023 020 1
$\sigma^{(2,2)}$	0.000 014 1	0.000 001 4	0.000 002 1(7)
$\sigma^{(4,0)}$	0.002 546 9	0.020 375 1	0.052 663 1
$\sigma^{(4,1)}$	0.000 000 0(28)	0.000 000 0(74)	0.000 000 0(192)
$\sigma^{(5,0)}$	0.000 018 4	0.000 082 0	0.000 096 3
$\sigma^{(6,0)}$	0.000 000 2(2)	0.000 006 5(65)	0.000 012 9(129)
σ_{fs}	-0.000 000 1	-0.000 006 7	-0.000 013 5(67)
$10^6\sigma$	17.735 427(3)	35.507 427(10)	59.966 512(24)
Previous	17.735 436(3)	35.507 434(9)	59.967 029(23)

culations, it is convenient to apply the expansion in the mass ratio, which takes the form (in a.u.)

$$\sigma = \sigma^{(2,0)} + \sigma^{(2,1)} + \dots, \quad (29)$$

$$\sigma^{(2,0)} = \frac{\alpha^2}{3} \left\langle \sum_a' \frac{1}{r_a} \right\rangle, \quad (30)$$

$$\begin{aligned} \sigma^{(2,1)} &= \frac{\alpha^2}{3} \frac{m_e}{m_N} \left[\left\langle \sum_a' \frac{1}{r_a} \frac{1}{(E-H)} p_N^2 \right\rangle + \langle p_N^2 \rangle \frac{(1-g_N)}{Z g_N} \right. \\ &\left. + \left\langle \sum_a' \vec{r}_a \times \vec{p}_N \frac{1}{(E-H)} \sum_b' \frac{\vec{r}_b \times \vec{p}_b}{r_b^3} \right\rangle \right], \end{aligned} \quad (31)$$

where all matrix elements in the above are assumed with infinite nuclear mass, and $\sigma^{(i,j)}$ denotes the expansion term of order $\alpha^i (m_e/m_N)^j$. The last term in the above differs in sign from that derived previously in Refs. [7,8], see Table I for the updated numerical values.

For the hydrogenic ion in nS state the nonrelativistic shielding constant, using Eq. (24), is

$$\begin{aligned} \sigma &= \frac{(Z\alpha)^2}{3n^2} \frac{m_N}{m_N + m_e} + \frac{(Z\alpha)^2 m_e^2}{3n^2 g_N (m_N + m_e)^2} \\ &\times \left[3 - \frac{g_N}{2} + \frac{m_N}{m_N + m_e} \left(1 - \frac{1}{Z} \right) - (g_N - 1) \frac{m_N}{Z m_e} \right] \\ &= \frac{(Z\alpha)^2}{3n^2 g_N (1+x)^2} \left[\left(3 - \frac{g_N}{2} + \frac{x}{1+x} \right) \right. \\ &\left. + \frac{x^2}{Z} \left(\frac{1}{1+x} + g_N \right) \right], \end{aligned} \quad (32)$$

where $x = m_N/m_e$. It is convenient to define $\delta g_N = -g_N \sigma$

$$\begin{aligned} \delta g_N &= \frac{(Z\alpha)^2}{3n^2 (1+x)^2} \left[\frac{g_N}{2} - 3 - \frac{x}{1+x} \right. \\ &\left. - \frac{x^2}{Z} \left(\frac{1}{1+x} + g_N \right) \right], \end{aligned} \quad (33)$$

which is in agreement with the known formula for the electron g -factor in the S-state of the hydrogenic ion [7, 15],

$$\delta g_e = \frac{(Z\alpha)^2}{3n^2(1+x)^2} \left[\left(\frac{g_e}{2} - 3 - \frac{x}{1+x} \right) - Zx^2 \left(\frac{1}{1+x} + g_e \right) \right], \quad (34)$$

with $x = m_e/m_N$, which verifies the new formula for the magnetic shielding in hydrogen-like ions. Its small electron mass expansion takes the following form:

$$\sigma = \frac{Z\alpha^2}{3n^2} \left[1 + \frac{m_e}{m_N} \left(\frac{1}{g_N} - 2 \right) + \frac{m_e^2}{m_N^2} \left(\frac{4Z-3}{g_N} - \frac{Z}{2} + 3 \right) + \dots \right], \quad (35)$$

where the quadratic in the mass ratio term differs from that derived previously in Ref. [[3], Eq. (64)] due to the computational mistake. As seen from Table I, the largest uncertainty for light ions comes from the relativistic recoil correction $\sigma^{(4,1)}$, but this has not yet been studied in the literature.

IV. FINITE NUCLEAR SIZE CORRECTIONS

Let us pass now to another nuclear correction, which is due to the finite distribution of the charge and the magnetic moment within the nucleus. We will study this correction for hydrogenic ions only, but generalization for an arbitrary ion is straightforward. For light hydrogenic ions this effect is given by

$$\sigma_{\text{fs}} = -\frac{Z\alpha^2}{3} [2(Z\alpha)^2 m^2 (r_C^2 + r_M^2) + 8(Z\alpha)^3 m \tilde{r}_Z], \quad (36)$$

where $m = m_e$, r_C is the charge radius, r_M the magnetic radius, and \tilde{r}_Z the effective Zemach radius of the nucleus. This formula is proved as follows.

The shift of nonrelativistic hydrogenic levels due to r_C is given by

$$\begin{aligned} \delta H &= eA^0 - \frac{e}{6} r_C^2 \bar{\nabla} \bar{E} \\ &= -Z\alpha \left(\frac{1}{r} - \frac{2\pi}{3} \delta^3(r) r_C^2 \right), \end{aligned} \quad (37)$$

where $e = e_e$. The finite nuclear size affects the nonrelativistic wave function, which in turn affects the matrix elements for the nuclear magnetic shielding,

$$\begin{aligned} \sigma_C &= 2 \frac{\alpha}{3m} \left\langle \frac{1}{r} \frac{1}{(E-H)} \frac{2\pi}{3} Z\alpha r_C^2 \delta^3(r) \right\rangle \\ &= -2(Z\alpha)^2 m^2 r_C^2 \frac{Z\alpha^2}{3}. \end{aligned} \quad (38)$$

To derive the contribution from the magnetic radius of the nucleus, let us rederive the leading shielding that comes from the $e^2 \bar{A}^2/(2m)$ term in the kinetic energy of the electron

$$\begin{aligned} \delta E &= \frac{\alpha}{2m} \left\langle (\bar{\mathbf{B}} \times \vec{r}) \cdot \left(\vec{\mu} \times \frac{\vec{r}}{r^3} \right) \right\rangle \\ &= \vec{\mu} \cdot \bar{\mathbf{B}} \frac{\alpha}{3m} \left\langle \vec{r} \cdot \frac{\vec{r}}{r^3} \right\rangle. \end{aligned} \quad (39)$$

The shielding σ is thus given by

$$\sigma = -\frac{\alpha}{3m} \left\langle \vec{r} \cdot \bar{\nabla} \left(\frac{1}{r} \right) \right\rangle. \quad (40)$$

The magnetic radius r_M enters the magnetic interaction similarly to r_C in Eq. (37); therefore, the shift due to the magnetic radius is

$$\begin{aligned} \sigma_M &= \frac{\alpha}{3m} \frac{2\pi}{3} r_M^2 \langle \vec{r} \cdot \bar{\nabla} (\delta^3(r)) \rangle \\ &= \frac{\alpha}{3m} \frac{2\pi}{3} r_M^2 (-3) \langle \delta^3(r) \rangle \\ &= -2(Z\alpha)^2 m^2 r_M^2 \frac{Z\alpha^2}{3}. \end{aligned} \quad (41)$$

The calculation of the shift due to the Zemach radius \tilde{r}_Z is more complicated. \tilde{r}_Z represents the hyperfine anomaly, namely, $(E_{\text{hfs}}^{\text{exp}} - E_{\text{hfs}}^{\text{point}})/E_F = -2Z\alpha m \tilde{r}_Z$, see Eq. (48). If we assume that it comes exclusively from the charge and magnetic moment distribution, it becomes r_Z given by Eq. (49), which can only be derived from the Dirac equation. Let us thus start derivation from the relativistic hyperfine splitting

$$E_{\text{hfs}} = -e \langle \psi^\dagger | \vec{\alpha} \cdot \bar{\mathbf{A}}_I | \psi \rangle, \quad (42)$$

where, for a point nucleus,

$$\bar{\mathbf{A}}_I = \frac{1}{4\pi} \vec{\mu}_I \times \frac{\vec{r}}{r^3}. \quad (43)$$

In the nonrelativistic limit E_{hfs} is given by the Fermi formula

$$\begin{aligned} E_F &= -\frac{e}{2m} \langle \phi | \{ \vec{\sigma} \cdot \vec{p}, \vec{\sigma} \cdot \bar{\mathbf{A}}_I \} | \phi \rangle \\ &= -\langle \phi | \vec{\mu}_e \cdot \bar{\mathbf{B}}_I | \phi \rangle \\ &= -\frac{2}{3} \langle \phi | \vec{\mu}_e \cdot \vec{\mu}_I \delta^3(r) | \phi \rangle. \end{aligned} \quad (44)$$

We are now ready to consider the leading finite nuclear size correction E_Z to the hyperfine splitting

$$\begin{aligned} E_Z &= 2 \left\langle \begin{array}{c} \phi^\dagger(0) \\ 0 \end{array} \middle| (-e) \vec{\gamma} \cdot \bar{\mathbf{A}}_I \frac{1}{\not{p} - m} e \gamma^0 A^0 \middle| \begin{array}{c} \phi(0) \\ 0 \end{array} \right\rangle \\ &= 2e^2 \int \frac{d^3p}{(2\pi)^3} \frac{1}{\bar{p}^2} \\ &\quad \times \left\langle \begin{array}{c} \phi^\dagger(0) \\ 0 \end{array} \middle| \vec{\gamma} \cdot \bar{\mathbf{A}}_I(-\vec{p}) (\not{p} + m) \gamma^0 A^0(\vec{p}) \middle| \begin{array}{c} \phi(0) \\ 0 \end{array} \right\rangle, \end{aligned} \quad (45)$$

where $p^0 = m$ and

$$A^0(\vec{p}) = -\frac{Ze}{\bar{p}^2} G_E(\vec{p}^2), \quad (46)$$

$$\bar{\mathbf{A}}_I(\vec{p}) = -i \vec{\mu}_I \times \frac{\vec{p}}{p^2} G_M(\vec{p}^2), \quad (47)$$

with normalization $G_E(0) = G_M(0) = 1$. E_Z can be simplified to

$$\begin{aligned} E_Z &= \frac{2Z\alpha m}{\pi^2} \int \frac{d^3p}{p^4} [G_E(p^2) G_M(p^2) - 1] E_F \\ &= -2Z\alpha m r_Z E_F, \end{aligned} \quad (48)$$

where

$$r_Z = \int d^3 r_1 \int d^3 r_2 \rho_E(r_1) \rho_M(r_2) |\bar{r}_1 - \bar{r}_2|, \quad (49)$$

and where ρ_E and ρ_M are the Fourier transforms of G_E and G_M . If we are about to represent complete hyperfine anomaly, then r_Z becomes \tilde{r}_Z in Eq. (48), because it may include the nuclear inelastic contribution.

Let us now combine perturbation due to r_Z and the homogenous magnetic field

$$\delta E = 2 \langle \bar{\psi}_M | (-e) \vec{\gamma} \cdot \vec{A}_1 \frac{1}{\not{p} - e \not{A} - m} e \gamma^0 A^0 | \psi_M \rangle, \quad (50)$$

where

$$| \psi_M \rangle = \left(I - \frac{1}{2m} \vec{\gamma} (\vec{p} - e \vec{A}) + \frac{e}{8m^2} \vec{\sigma} \vec{B} \right) \begin{pmatrix} \phi_M(0) \\ 0 \end{pmatrix}, \quad (51)$$

and where ϕ_M is an eigenstate of

$$H_M = \frac{p^2}{2m} - \frac{Z\alpha}{r} - \frac{e}{2m} \vec{\sigma} \vec{B} \left(1 - \frac{p^2}{2m^2} + \frac{Z\alpha}{6mr} \right). \quad (52)$$

We claim that the $e\vec{A}$ terms in the propagator and in the wave function can be neglected, because they lead to an additional p^2 in the denominator and their contribution thus goes with the nuclear radius to the third power. Therefore, we have only two corrections due to the last terms in Eqs. (51) and (52), namely,

$$\begin{aligned} \delta E = & 2 \left\langle \phi \left| H_Z \frac{e}{8m^2} \vec{\sigma} \vec{B} \right. \right\rangle \\ & - 2 \langle \phi | H_Z \frac{1}{(E - H_Y)} \frac{e}{2m} \vec{\sigma} \vec{B} \left(-\frac{p^2}{2m^2} + \frac{Z\alpha}{6mr} \right) | \phi \rangle, \end{aligned} \quad (53)$$

where

$$H_Z = \frac{2}{3} \vec{\mu}_e \cdot \vec{\mu}_1 \delta^3(r) (2Z\alpha m r_Z). \quad (54)$$

Therefore,

$$\delta E = \vec{\mu}_1 \cdot \vec{B} \alpha (Z\alpha)^3 m r_Z \frac{8}{3} \left(\frac{1}{4} - \frac{X}{m^2 (Z\alpha)^3} \right), \quad (55)$$

where

$$\begin{aligned} X = & \langle \phi | \pi \delta^3(r) \frac{1}{(E - H_Y)} \left(-\frac{p^2}{2m^2} + \frac{Z\alpha}{6mr} \right) | \phi \rangle \\ = & \frac{5}{4} (Z\alpha)^3 m^2. \end{aligned} \quad (56)$$

Thus, with $\delta E = \vec{\mu}_1 \cdot \vec{B} \sigma_Z$,

$$\sigma_Z = -\frac{8}{3} \alpha (Z\alpha)^4 m r_Z, \quad (57)$$

which proves Eq. (36). In addition, Yerokhin [16] verified this equation by numerically calculating the magnetic shielding with Dirac wave functions for various Z , charge, and magnetic radii of the nucleus. The advantage of Eq. (36) over the direct numerical calculation is the presence of \tilde{r}_Z instead of r_Z , which represents the sum of elastic and inelastic contributions to HFS, and thus can be determined from the HFS anomaly.

V. SUMMARY

The total magnetic shielding for hydrogen-like ions including contributions up to order α^6 is (cf. Eq. (25) of Ref. [2])

$$\begin{aligned} \sigma = & \frac{Z\alpha^2}{3} + \frac{97}{108} Z^3 \alpha^4 + \frac{289}{216} Z^5 \alpha^6 + \frac{8\alpha^2}{9\pi} (Z\alpha)^3 \\ & \times \left[\ln(Z\alpha)^{-2} + 2 \ln k_0 - 3 \ln k_3 - \frac{221}{64} + \frac{3}{5} \right] \\ & + \frac{Z\alpha^2}{3} \left[\left(\frac{1}{g_N} - 2 \right) \frac{m}{m_N} + \left(\frac{4Z-3}{g_N} - \frac{Z}{2} + 3 \right) \frac{m^2}{m_N^2} \right] \\ & - \frac{Z\alpha^2}{3} [2(Z\alpha)^2 m^2 (r_C^2 + r_M^2) + 8(Z\alpha)^3 m \tilde{r}_Z], \end{aligned} \quad (58)$$

where [17]

$$\ln k_0 = 2.984\,128\,556, \quad (59)$$

$$\ln k_3 = 3.272\,806\,545. \quad (60)$$

Numerical results for all these known contributions to the magnetic shielding in H, He⁺, and He are presented in Table I. The updated values are $\sigma^{(2,1)}$ for He, where we corrected the sign error in the last term in Eq. (31). This leading recoil correction to the magnetic shielding is about 0.02×10^{-6} , which is the relative 2×10^{-8} correction in the determination of nuclear magnetic moments. The higher-order recoil correction, the last term in Eq. (35), which is also corrected in this work, is much smaller and thus is negligible at present accuracy of measurements. The same holds for nuclear finite size effects, described by Eq. (36); they are negligible for light elements and can safely be neglected. However, the nuclear finite size effects can be significant for heavy elements, where they strongly affect binding energies and hyperfine splitting.

Finally, our recommended values for the nuclear magnetic shieldings are in the penultimate row, and they are compared to previous recommendations from Ref. [3] in the last row. The largest change of 0.5×10^{-9} is for the He atom; changes to H and He⁺ ion are negligible.

We can now use these new shieldings to recalculate the helium magnetic moment from He⁺ measurement

$$\mu(^3\text{He}^+) = -4.255\,099\,606\,9(30)(17) \times \frac{\mu_N}{2}, \quad (61)$$

namely, it is

$$\begin{aligned} \mu(^3\text{He}^{++}) = & \frac{\mu(^3\text{He}^+)}{1 - \sigma(^3\text{He}^+)} \\ = & -2.127\,625\,350\,0(17) \mu_N, \end{aligned} \quad (62)$$

which differs slightly from that in Ref. [1], while our recommended value for the atomic ³He magnetic moment is

$$\begin{aligned} \mu(^3\text{He}) = & \mu(^3\text{He}^+) \frac{1 - \sigma(^3\text{He})}{1 - \sigma(^3\text{He}^+)} \\ = & -2.127\,497\,763\,7(17) \mu_N, \end{aligned} \quad (63)$$

which can serve as a reference in gaseous NMR measurements [18] because it is the most accurately known atomic magnetic moment.

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- [1] A. Schneider, B. Sikora, S. Dickopf, M. Müller, N. S. Oreshkina, A. Rischka, I. A. Valuev, S. Ulmer, J. Walz, Z. Harman, C. H. Keitel, A. Mooser, and K. Blaum, *Nature (London)* **606**, 878 (2022).
- [2] D. Wehrli, A. Spyszkiewicz-Kaczmarek, M. Puchalski, and K. Pachucki, *Phys. Rev. Lett.* **127**, 263001 (2021).
- [3] D. Wehrli, M. Puchalski, and K. Pachucki, *Phys. Rev. A* **105**, 032808 (2022).
- [4] A. Mooser (private communication).
- [5] S. G. Karshenboim, *Phys. Rep.* **422**, 1 (2005).
- [6] M. Jaszunski, A. Antušek, P. Garbacz, K. Jackowski, W. Makulski, and M. Wilczek, *Prog. Nucl. Magn. Reson. Spectrosc.* **67**, 49 (2012).
- [7] K. Pachucki, *Phys. Rev. A* **78**, 012504 (2008).
- [8] A. Rudziński, M. Puchalski, and K. Pachucki, *J. Chem. Phys.* **130**, 244102 (2009).
- [9] V. A. Yerokhin, K. Pachucki, Z. Harman, and C. H. Keitel, *Phys. Rev. Lett.* **107**, 043004 (2011).
- [10] V. A. Yerokhin, K. Pachucki, Z. Harman, and C. H. Keitel, *Phys. Rev. A* **85**, 022512 (2012).
- [11] R. A. Hegstrom, *Phys. Rev. A* **7**, 451 (1973).
- [12] R. Pohl, A. Antognini, F. Nez, F. D. Amaro, F. Biraben, J. a. M. R. Cardoso, D. S. Covita, A. Dax, S. Dhawan, L. M. P. Fernandes, A. Giesen, T. Graf, T. W. Hänsch, P. Indelicato, L. Julien, C.-Y. Kao, P. Knowles, E.-O. Le Bigot, Y.-W. Liu, J. A. M. Lopes *et al.*, *Nature (London)* **466**, 213 (2010).
- [13] K. Schuhmann, L. M. P. Fernandes, F. Nez, M. A. Ahmed, F. D. Amaro, P. Amaro, F. Biraben, T.-L. Chen, D. S. Covita, A. J. Dax, M. Diepold, B. Franke, S. Galtier, A. L. Gouvea, J. Götzfried, T. Graf, T. W. Hänsch, M. Hildebrandt, P. Indelicato, L. Julien *et al.* (The CREMA Collaboration), [arXiv:2305.11679](https://arxiv.org/abs/2305.11679) [physics.atom-ph].
- [14] E. Tiesinga, P. J. Mohr, D. B. Newell, and B. N. Taylor, *Rev. Mod. Phys.* **93**, 025010 (2021).
- [15] H. Grotch and R. A. Hegstrom, *Phys. Rev. A* **4**, 59 (1971).
- [16] V. Yerokhin (private communication).
- [17] K. Pachucki, A. Czarnecki, U. D. Jentschura, and V. A. Yerokhin, *Phys. Rev. A* **72**, 022108 (2005).
- [18] T. R. Gentile, P. J. Nacher, B. Saam, and T. G. Walker, *Rev. Mod. Phys.* **89**, 045004 (2017).