Nonlinearities of King's plot and their dependence on nuclear radii

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Investigations of isotope shifts of atomic spectral lines provide insights into nuclear properties. Deviations from the linear dependence of the isotope shifts of two atomic transitions on nuclear parameters, leading to a nonlinearity of the so-called King plot, are actively studied as a possible way of searching for new physics. In the present Letter we calculate the King-plot nonlinearities originating from the standard-model atomic theory. The calculation is performed both analytically, for a model example applicable for an arbitrary atom, and numerically, for one-electron ions. It is demonstrated that the standard-model predictions of the King-plot nonlinearities are hypersensitive to experimental errors of nuclear charge radii. This effect significantly complicates identifications of possible King-plot nonlinearities originating from the new physics.

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Introduction. Introduced by King in 1963 [1], the King plot has been widely used as an extremely useful tool for interpreting the results of isotope-shift atomic experiments. After measuring the energies of two transitions for four or more isotopes of the same element, one can arrange the results as a King plot, which should be-to a very high accuracy-linear. In this way the experimental results can be cross-checked without any further theoretical input. Moreover, the two parameters of the linear plot can be interpreted in terms of the mass and field shifts and separately compared with theoretical calculations [2].

Progress in quantum logic techniques and collinear spectroscopy achieved during the last years resulted in a dramatic increase of experimental accuracy. In particular, measurements of optical-clock transitions were demonstrated on a few-Hz precision level [3,4]. An even higher accuracy can be achieved by using entangled states [5] and coherent highresolution optical spectroscopy [6]. Naturally, a question arises if the King plot is going to stay linear at this new level of experimental precision. Recent experiments in strontium and ytterbium ions [7,8] gave first indications that the linearity is actually broken on the level of several standard deviations.

It was recently demonstrated [9-11] that the linearity of the King plot can be used in the search for new physics, specifically, to constrain the coupling strength of hypothetical new-physics boson fields to electrons and neutrons. It remained unclear, however, whether any actually observed nonlinearity can be clearly interpreted as a manifestation of new physics, since the standard-model theory can also produce effects that slightly bend the King plot. Such nonlinear effects have been the subject of several recent studies [12–14].

The King plot is defined so that its construction does not require any knowledge about the nuclear charge radii; only the nuclear masses are involved. Any prediction of King-plot nonlinearities, however, requires an experimental input in the form of nuclear radii. One might assumed that in view of the extreme smallness of the nonlinearities, our limited knowledge of the nuclear radii should not cause any problems. In the present Letter we will demonstrate that it is exactly the opposite. Even very small errors in experimental nuclear radii lead to greatly amplified uncertainties of the King-plot nonlinearities, which bring about problems in asserting even the order of the magnitude of the effect. This makes the Kingplot nonlinearities a very sensitive tool for studying nuclear radii, but diminishes their valuableness in the search for new physics.

Arbitrary atom. Let us consider the isotope shift of the energy of the transition a between isotopes with mass numbers A_i and A_0 , which will be denoted as E_a^i (the reference isotope A_0 will be implicit everywhere and suppressed in the notations for brevity). It is convenient to introduce the so-called reduced isotope-shift energies n_a^i as

$$n_a^i = \frac{E_a^i}{\frac{m}{M_i} - \frac{m}{M_0}},\tag{1}$$

where *m* is the electron mass and M_k is the nuclear mass of the isotope A_k . Within the standard formulation, the reduced isotope-shift energy is represented as a sum of the mass-shift and the field-shift contributions,

$$n_a^i = K_a + F_a r_i^{(2)}, (2)$$

where K_a and F_a are the mass-shift and the field-shift con-stants, respectively, $r_i^{(2)} \equiv \left[\frac{R_i^2}{k_c^2} - \frac{R_0^2}{k_c^2}\right] / \left[\frac{m}{M_i} - \frac{m}{M_0}\right]$, R_k is the root-mean-square (rms) nuclear charge radius of the isotope A_k , and $\lambda_C = \frac{\lambda_C}{2\pi}$ is the reduced Compton wavelength.

It is usually assumed that the isotope-shift constants K_a and F_a depend only on the transition but not on the isotope. In this case, by considering two transitions, *a* and *b*, and three pairs of isotopes (A_i , A_0) with i = 1, 2, and 3, one can form the so-called King plot [1]. It is easy to show that three points (x_i , y_i) = (n_a^1 , n_b^1), (n_a^2 , n_b^2), (n_a^3 , n_b^3) lie on a straight line,

$$y = \left(K_b - \frac{F_b}{F_a}K_a\right) + \frac{F_b}{F_a}x.$$
 (3)

It is important that the linear dependence of (n_a^i, n_b^i) does not rely on a theoretical knowledge of the isotope-shift constants; the only theoretical input is the representation of the isotope shifts in the form of Eq. (2). This representation is remarkably accurate, but at the level of present-day experimental precision this accuracy may not be sufficient.

Let us now consider a generalization of Eq. (2) that takes into account that the field shift, in addition to R^2 , depends also on a higher power of *R*. Specifically, we write

$$n_a^i = K_a + F_a r_i^{(2)} + F_a w_a r_i^{(3)}, \tag{4}$$

where $r_i^{(3)} \equiv \left[\frac{R_i^3}{\lambda_c^3} - \frac{R_0^3}{\lambda_c^3}\right] / \left[\frac{m}{M_i} - \frac{m}{M_0}\right]$ and w_a is an additional isotope-shift constant, which depends on the transition but not on the isotope. We keep in mind that $r_i^{(3)} \approx \left(\frac{R_0}{\lambda_c}\right) r_i^{(2)} \ll r_i^{(2)}$ since $R_0/\lambda_c \approx 0.01$.

Obviously, the three points (n_a^i, n_b^i) of Eq. (4) no longer lie on a straight line but rather on a parabola, $y = a + bx + cx^2$. The nonlinearity of this function is connected with the coefficient $c = \frac{1}{2}y''(x)$. Evaluating the second derivative of the exact fit of the three points (n_a^i, n_b^i) and neglecting $r_i^{(3)}$ on the background of $r_i^{(2)}$, we arrive at the following expression,

$$y'' = 2\frac{F_b}{F_a^2}(w_a - w_b)\mathcal{P}_{\text{nucl}},$$
(5)

where

$$\mathcal{P}_{\text{nucl}} = \frac{r_3^{(2)} (r_1^{(3)} - r_2^{(3)}) + r_1^{(2)} (r_2^{(3)} - r_3^{(3)}) + r_2^{(2)} (r_3^{(3)} - r_1^{(3)})}{(r_1^{(2)} - r_2^{(2)}) (r_1^{(2)} - r_3^{(2)}) (r_2^{(2)} - r_3^{(2)})}.$$
(6)

It is now clear that if we would like to predict the nonlinearity of King's plot in the approximation of Eq. (4), we have to calculate the field-shift constants $F_{a,b}$ and $w_{a,b}$ and multiply them by the nuclear factor \mathcal{P}_{nucl} , which is determined by the experimental nuclear parameters (charge radii and masses) of the isotopes.

Let us now address the following question: To which accuracy can we determine the nuclear factor \mathcal{P}_{nucl} basing on the available experimental data? We consider an example of four isotopes of tin (Z = 50) with $A_i = (118, 120, 122, 124)$. The nuclear radii are taken [15] as $R_{118} = 4.6393$ (1) fm, $R_{120} = 4.6519$ fm, $R_{122} = 4.6634$ (1) fm, and $R_{124} = 4.6735$ (1) fm. Note that we keep only the *relative* uncertainty of the radii with respect to the A = 120 isotope, omitting the common systematic uncertainty of 0.0020 fm, which will be ignored in the present context. The nuclear masses are taken from Ref. [16]; their uncertainties do not play any role here.

We now evaluate \mathcal{P}_{nucl} numerically, varying each of R_i within their uncertainties. The result is surprising: Although the nuclear radii are supposed to be known with a five-digit accuracy, the results for \mathcal{P}_{nucl} may vary by several orders of magnitude. Specifically, we obtain $|\mathcal{P}_{nucl,min}| = 5.6 \times 10^{-5}$

and $|\mathcal{P}_{\text{nucl,max}}| = 1.3 \times 10^{-3}$. The reason for such striking behavior is that both the numerator and the denominator in Eq. (6) can nearly vanish for some combinations of R_i , leading to very small or very large values of $\mathcal{P}_{\text{nucl}}$.

We have to conclude that regardless of the accuracy of the theoretical calculations of the isotope-shift constants, any predictions for the nonlinearities of the King plot are problematic because of their extreme sensitivity to errors of the experimental nuclear radii. A way to circumvent this problem might be to consider the ratio of the nonlinearities for two pairs of transitions. We note that in Eq. (5) the nuclear part is factorized out and, therefore, the ratio of the second derivatives y'' for two pairs of transitions is free from nuclear uncertainties. However, the exact factorization holds only for one-parameter extensions of the standard formula such as Eq. (4).

The realistic situation is more complicated. In particular, the field-shift constant contains terms with $\ln R$ in its expansion and is modified by the nuclear-polarization effects which often demonstrate irregular dependence on the isotope mass number. The general expression for the reduced isotope-shift energy can be written as

$$n_a^i = K_a^i + F_a^i r_i^{(2)}, (7)$$

where the isotope-shift "constants" K_a^i and F_a^i depend on the isotope parameters; this dependence is indicated by the superscript *i*. For a general atom, the theoretical predictions of the isotope dependence of these constants are rather complicated. In the present work we will address this problem for the simplest case of H-like ions.

H-like ions. We now turn to examining the case of oneelectron ions. From now on, we will adopt the relativistic units with $c = \hbar = 1$, which greatly simplifies the following formulas. We start with the mass shift. The main isotope dependence of the mass-shift constant comes from the quadratic part $\sim (\frac{m}{M})^2$ of the nuclear recoil effect. Specifically, we write

$$K_a^i = K_a^{(1)} + \frac{m}{M_i} K_a^{(2)},\tag{8}$$

where the first-order mass-shift constant $K^{(1)}$ is well known (see, e.g., the review [17]), and $K^{(2)}$ is the second-order massshift constant calculated in the present work. For a reference state with quantum numbers (njl), the second-order massshift constant is expressed as a second-order perturbation correction induced by the relativistic recoil operator H_{rec} [18],

$$K_{njl}^{(2)} = \sum_{k \neq (njl)} \frac{\langle njl | H_{\rm rec} | k \rangle \langle k | H_{\rm rec} | njl \rangle}{\varepsilon_{njl} - \varepsilon_k},\tag{9}$$

where the summation over k is performed over the complete Dirac spectrum and

$$H_{\rm rec} = \frac{1}{2} \left[\boldsymbol{p}^2 - \frac{Z\alpha}{r} \left(\boldsymbol{\alpha} + \frac{(\boldsymbol{\alpha} \cdot \boldsymbol{r})\boldsymbol{r}}{r^2} \right) \cdot \boldsymbol{p} \right].$$
(10)

Here, p is the electron momentum and α is the vector of Dirac matrices. We calculate $K_{njl}^{(2)}$ in two ways: analytically, performing the summation over the Dirac spectrum with the help of generalized virial relations [19], and numerically, with the finite basis set *B*-spline method [20]. The results obtained by two different methods agree with each other. For light ions, we also observe good agreement with known results for the

first terms of the $Z\alpha$ expansion [17]

$$K_{njl}^{(2)} = -\frac{(Z\alpha)^2}{2n^2} + \frac{(Z\alpha)^4}{2n^3} \left(\frac{3}{2n} - \frac{2}{2l+1}\right) + \cdots .$$
(11)

The isotope dependence of the field-shift constant arises from two main sources: (i) the deviation of the *R* dependence of the finite nuclear size (fns) correction from the R^2 form and (ii) the effect of the nuclear polarization. The fns correction to the energy levels ΔE_{fns} is calculated numerically, by solving the Dirac equation with the potential of an extended nuclear charge (see Ref. [21] for details). The obtained results are represented, factorizing out the leading *R* dependence [22], as

$$\Delta E_{\rm fns} = R^{2\gamma} G_{\rm fns}(Z\alpha, R), \qquad (12)$$

where $\gamma = \sqrt{(j + \frac{1}{2})^2 - (Z\alpha)^2}$ and *R* is the nuclear rms radius. The field-shift constant is obtained from the fns correction to the transition energy as $F_a^i = \frac{\Delta E_{\text{fms},a}}{R^2}$. We note that the deviation of the $R^{2\gamma}$ term in Eq. (12) from the standard R^2 factor does not cause any nonlinearity of King's plot, because it leads to a multiplication of *F* in Eq. (2) by a factor that does not depend on the transition. So, the fns contribution to the nonlinearity comes only from the *R* dependence of the function G_{fns} . We note that this dependence is very weak; in order to reliably calculate it for medium-*Z* ions, we had to solve the Dirac equation in the extended-precision arithmetics.

The nuclear polarization correction to the energy of a state with quantum numbers (njl) is expressed as [23,24]

$$\Delta E_{\rm np} = -\alpha \sum_{LM} B(EL) \sum_{k} \frac{|\langle njl | F_L Y_{LM} | k \rangle|^2}{\varepsilon_k - \varepsilon_{njl} + \operatorname{sgn}(\varepsilon_k)\omega_L}, \quad (13)$$

where $B(EL) = B(EL; L \rightarrow 0)$ are the reduced probabilities of nuclear transitions from the excited ("L") to the ground ("0") level, ω_L are the nuclear excitation energies with respect to the ground state, F_L are radial functions given by Eq. (4) of Ref. [24], Y_{LM} are the spherical harmonics, and the summation over k is performed over the complete spectrum of electronic states. Multipole contributions in Eq. (13) are usually separated into two classes: contributions from the giant resonance transitions and those from the lowest-lying nuclear rotational transitions. Among the latter, the electrical quadrupole transition between the ground 0^+ the lowest-lying 2^+ state is the dominant channel for most of the even-even nuclides. Our calculation of the nuclear polarization correction includes the dominant E2 nuclear rotational transition and the giant resonance transitions with $L \leq 3$. Experimental results for the nuclear quadrupole transition probabilities B(E2) and excitation energies ω_L were taken from Ref. [25]. The summation over the Dirac spectrum was performed with help of the finite basis set B-spline method [20]. We find good agreement with the numerical results of Ref. [24]. A similar calculation was recently presented in Ref. [26]; the difference is that in that work empirical approximate formulas were used for B(E2)and only the dominant L = 1 giant resonance was included.

The numerical results of our calculation of the nuclear polarization correction are conveniently parametrized in terms of the dimensionless function g_{np} , defined as a multiplicative



FIG. 1. The isotope dependence of the finite nuclear size and nuclear polarization effects, with the leading $R^{2\gamma}$ dependence factorized out, for the 1s state of H-like barium (Z = 56). Black diamonds show the finite nuclear size correction, in terms of the function G_{fns} defined by Eq. (12). The red circles present results for the combined nuclear size and nuclear polarization effect, in terms of the function $G_{\text{nucl}} = G_{\text{fns}}(1 - \frac{1}{1000}g_{\text{np}})$. Both functions $G_{\text{fns}}(A)$ and $G_{\text{nucl}}(A)$ are normalized to their values for the isotope $A_0 = 134$.

factor to the fns correction,

$$\Delta E_{\rm np} = -\frac{g_{\rm np}(Z,A)}{1000} \Delta E_{\rm fns}.$$
 (14)

We studied the nuclear polarization effect for the 1s, 2s, and $2p_{1/2}$ states; for the $2p_{3/2}$ state the corresponding correction is very small and thus neglected. We find that g_{np} varies from 0.3 to 1.0 for all nuclei encountered in the present work, $Z \in (20, 92)$. It is instructive to compare the isotope dependence of the fns and the nuclear polarization effects. Such a comparison is presented in Fig. 1 for a chain of isotopes of Ba⁵⁵⁺. We observe that both effects yield significant contributions, but the nuclear polarization is larger and, more importantly, its isotope dependence is nonmonotonic. This is connected with the behavior of the nuclear transition probabilities B(E2), which tend to grow as the nuclide mass number moves away from the stability region. We note that the kink of the plot for G_{fns} is due to the unevenness of the dependence of R on the mass number A.

We are now in a position to calculate the King-plot nonlinearities for H-like ions. We adopt the definition [12,14] of the nonlinearity of a three-point curve as a shift of the ordinate of the third point from the straight line defined by the first two points,

$$\delta\epsilon_{ab}^{3} = \left(\frac{m}{M_{3}} - \frac{m}{M_{0}}\right) \left[n_{b}^{3} - n_{b}^{1} - \frac{n_{b}^{2} - n_{b}^{1}}{n_{a}^{2} - n_{a}^{1}} \left(n_{a}^{3} - n_{a}^{1}\right)\right].$$
 (15)

The nonlinearity of the King plot for the transition pair (a, b) is defined by symmetrizing the above expression with respect of a and b,

$$\Delta_{\rm NL}(ab) = \frac{1}{2} \left(\left| \delta \epsilon_{ab}^3 \right| + \left| \delta \epsilon_{ba}^3 \right| \right). \tag{16}$$

The nonlinearity Δ_{NL} has a dimension of energy and approximately indicates the experimental accuracy needed in



FIG. 2. King-plot nonlinearities of H-like ions for two transition pairs, (a, b) (top) and (b, c) (bottom), where $a = 1s \rightarrow 2p_{1/2}$, $b = 2s \rightarrow 2p_{1/2}$, and $c = 2s \rightarrow 2p_{3/2}$. The calculated points correspond to the isotope chains $(A_0, A_0 + 2, A_0 + 4, A_0 + 6)$, with the reference (A_0) isotopes: ⁴⁰Ca, ⁶⁴Zn, ⁸²Kr, ⁹²Mo, ¹¹⁸Sn, ¹³²Ba, ¹⁵⁴Gd, ¹⁶⁶Yb, ¹⁹⁶Hg. The dashed lines represent the variation of the nonlinearities when the nuclear radii are varied within their experimental uncertainties.

order to detect it. We note that the numerical evaluation of Δ_{NL} involves large numerical cancellations (due to multiple subtractions of similar numbers) and needs to be performed in an extended-precision arithmetics.

The results of our numerical calculations of the Kingplot nonlinearities for different H-like ions are presented in Fig. 2. We studied two pairs of transitions, (a, b) and (b, c), with $a = 1s \rightarrow 2p_{1/2}$, $b = 2s \rightarrow 2p_{1/2}$, and $c = 2s \rightarrow 2p_{3/2}$. From Fig. 2 we make several conclusions which could have been anticipated from our earlier analytical considerations. First, we find that the calculated nonlinearities behave irregularly as a function of Z, because they crucially depend on the differences of charge radii of the nuclides. Second, we find that the experimental errors of the nuclear radii cause tremendously amplified uncertainties of the resulting nonlinearities. Generally, only the upper bounds of the nonlinearities can be predicted and these bounds depend crucially on the assumed uncertainties of the nuclear radii. We would like to stress that for our analysis we selected the nuclides with the best-known charge radii. Specifically, the relative uncertainties of the rms radii for the studied nuclides with Z > 40 are about 0.0001 fm [15]. For less-known isotopes, the uncertainties of nonlinearities are larger by orders of magnitude.

Although our numerical calculations were performed for the specific choice of H-like ions, we expect that the situation remains qualitatively the same for many-electron atoms, including the case of singly charged ions which are most relevant from an experimental point of view. This expectation is supported by the analytical example discussed in the first part of the present Letter. In order to obtain quantitative results for the upper bounds for the standard-model King-plot nonlinearities for singly charged ions, dedicated calculations are needed for each particular element.

The observed hypersensitivity of the King-plot nonlinearities on the nuclear radii can be used for improving our knowledge of the isotope differences of the nuclear charge radii. For example, if one of the three differences of the nuclear radii involved in the King plot is known much more poorly than the other two, it should be possible to improve its value basing on the measured King-plot nonlinearity.

The dependence of the predicted King-plot nonlinearities on the nuclear radii can be significantly reduced if we study the ratio of the nonlinearities for two transition pairs, e.g., $\Delta_{\rm NL}(ab)/\Delta_{\rm NL}(bc)$. This can be seen from the analytical example of Eq. (5) and also from the fact that the plots for $\Delta_{\rm NL}(ab)$ and $\Delta_{\rm NL}(bc)$ in Fig. 2 look very similar. However, the factorization of the nuclear degrees of freedom, exact in the case of Eq. (5), becomes only approximate in the complete calculation. As a consequence, we still encounter combinations of nuclear radii leading to strong variations of the ratio of the nonlinearities.

Turning to perspectives of using the King-plot nonlinearities in the search for new physics beyond the standard model, we conclude that the findings of the present Letter significantly complicate such searches. More specifically, when an experiment detects a nonlinearity (as, e.g., in Refs. [7,8]), it should be very difficult to distinguish whether this is caused by new physics or the standard atomic theory, because of our limited knowledge of the nuclear charge radii. Therefore, the upper bounds on the new physics derived from such an observation will be defined by the observed nonlinearity and any further progress in experimental precision will not improve these bounds. An alternative approach would be to use independent constraints to eliminate the possibility of new physics [27–30] and interpret the observed nonlinearities in terms of constraints on the nuclear radii.

Summarizing, we calculated nonlinearities of the King plot originating from the standard-model atomic theory. The calculations were performed first analytically for a model problem applicable for an arbitrary atom. Next, we performed a detailed numerical calculation for hydrogenlike ions, taking into account the quadratic nuclear recoil, the higher-order finite nuclear size, and the nuclear polarization effects. We found that the standard-model predictions of the King-plot nonlinearities are hypersensitive to experimental errors of nuclear charge radii, often leading to uncertainties even in the order of magnitude of the effect. Our restricted knowledge of the nuclear radii leads to large ambiguities in the standard-model predictions of the nonlinearities, thus greatly complicating an identification of possible effects originating from new physics.

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