Maxwell fish eye for polarized light

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We consider the propagation of polarized light in a medium with Maxwell fish eye refraction index profile. We show that polarization violates the additional symmetries of the medium so that ray trajectories no longer remain closed. Then we suggest a modified, polarization-dependent Maxwell fish eye refraction index which restores all symmetries of the initial profile and yields closed trajectories of polarized light. Explicit expressions for the polarization-dependent integrals of motion and the solutions of corresponding ray trajectories are presented.

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I. INTRODUCTION

It is well known that the minimal action principle came into physics from geometric optics. Initially, it was invented for the description of the propagation of light and is presently known as the Fermat principle:

$$S_{\text{Fermat}} = \frac{1}{\lambda_0} \int d\tilde{l}, \quad d\tilde{l} := n(\mathbf{r}) |d\mathbf{r}/d\tau| d\tau, \qquad (1)$$

where $n(\mathbf{r})$ is the refraction index and λ_0 is the wavelength in vacuum. This action could be interpreted as the action of a system on a three-dimensional curved space equipped with the "optical metrics" or the Fermat metrics (see [1]) of Euclidean signature

$$d\tilde{l}^2 = n^2(\mathbf{r})d\mathbf{r} \cdot d\mathbf{r}.$$
 (2)

Thus, the symmetries of the system which describe the propagation of light in a particular medium are coming from the symmetries of the optical metrics of that particular medium. In systems with a maximal number of functionally independent integrals of motion (2N-1) integrals for an N-dimensional system), all the trajectories of the system become closed. The closeness of the trajectories makes such profiles highly relevant in the study of cloaking and perfect imaging phenomena. The most well-known profile of this sort is the so-called Maxwell fish eye profile which is defined by the metrics of a (three-dimensional) sphere or pseudosphere [by pseudosphere we mean the upper (or lower) sheet of a two-sheet hyper-boloid]:

$$n_{Mfe}(\mathbf{r}) = \frac{n_0}{|1 + \kappa \mathbf{r}^2|}, \quad \kappa = \pm \frac{1}{4r_0^2},$$
 (3)

where the plus (minus) sign in the expression for κ corresponds to a sphere (pseudosphere) with the radius r_0 , and

 $n_0 > 0$. Apart from applications in cloaking and perfect imaging phenomena [2–4], the Maxwell fish eye is a common profile in quantum optics with single atoms and photons [5], optical resonators [6], discrete spectrum radiation [7], etc. Moreover, there are many experimental implementations of Maxwell fish eye lenses [8–10].

In most of the listed studies, scalar wave approximation was used and polarization of electromagnetic waves was not taken into account. In these systems, the ray trajectories belong to the plane which is orthogonal to the angular momentum. Introduction of spin (polarization) results in the rotation of this plane by a constant angle proportional to spin; moreover, it breaks the nonrotational symmetries of optical systems with the Maxwell fish eye profile, so that photon trajectories no longer remain closed [11]. Thus, the key property of the Maxwell fish eye profile which makes it relevant in cloaking and perfect imaging phenomena is violated.

In the present paper, we continue our study of a polarized light passing through the Maxwell fish eye profile within the geometrical optics approximation. The key point of our study is that we introduce a polarization-dependent deformation of the Maxwell fish eye profile:

$$n_{Mfe}^{s}(\mathbf{r}) = \frac{n_{Mfe}(\mathbf{r})}{2} \left(1 + \sqrt{1 - \frac{4\kappa s^2 \lambda_0^2}{n_0} \frac{1}{n_{Mfe}(\mathbf{r})}} \right), \quad (4)$$

where $n_{Mfe}(\mathbf{r})$ is the original Maxwell fish eye profile given by (3), and s is the light polarization. For linearly and circularly polarized light we have s=0 and $s=\pm 1$ respectively. The proposed deformation restores all the symmetries of the optical Hamiltonian, with Maxwell fish eye profile, which were broken after the inclusion of polarization. It also ensures the closeness of the trajectories for polarized photons and can be used for cloaking and perfect imaging of polarized photons. It is seen that the spin-induced term is proportional to the dimensionless parameter $s^2 \lambda_0^2 / r_0^2$, where r_0 is the characteristic length of the profile defined in (3). This means that spin will play a significant role only in the vicinity of the

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wave and geometrical optics border $s\lambda_0/r_0 \sim 1$, since we are working in the framework of a geometrical optics approximation, $\lambda_0 \ll r_0$. Below we will investigate the influence of spin (polarization) on the ray trajectories in the deformed Maxwell fish eye profile given by (4).

The paper is organized as follows. In Sec. II, we describe the Hamiltonian formulation of the geometric optical system given by the action (1). We also present some other text-book facts on the duality between Coulomb and free-particle systems on a (pseudo)sphere which allow us to relate the Maxwell fish eye and Coulomb profiles and will be used in our further consideration.

In Sec. III, we present the Hamiltonian formalism for the polarized light propagating in an optical medium and propose the general scheme of the deformation of an isotropic refraction index profile which allows us to restore the initial symmetries after the inclusion of polarization.

In Sec. IV, we use the proposed scheme for the construction of a "polarized Maxwell fish eye" profile (4) which inherits all the symmetries of the original profile (3) when light polarization is taken into account. We present the explicit expressions for the symmetry generators of the corresponding Hamiltonian system and find the expressions of the Casimirs of their symmetry algebra.

In Sec. V the explicit expressions for the trajectories of polarized light are presented. It is shown that these trajectories are no longer orthogonal to the angular momentum but turn to a fixed angle relative to it. Despite deviations from circles, these trajectories remain closed.

Through the text we will use the notation $r := |\mathbf{r}|$, $\mathbf{r} := (x_1, x_2, x_3)$, $\mathbf{p} := (p_1, p_2, p_3)$, $p := |\mathbf{p}|$, and so on.

II. SCALAR WAVES

Due to reparametrization invariance of the action (1), the Hamiltonian constructed by the standard Legendre transformation is identically zero. However, the constraint between momenta and coordinates appears there:

$$\Phi := \frac{\mathbf{p}^2}{n^2(\mathbf{r})} - \lambda_0^{-2} = 0.$$
 (5)

Hence, in accordance with the Dirac's constraint theory [12] the respective Hamiltonian system is defined by the canonical Poisson brackets

$$\{x_i, p_i\} = \delta_{ii}, \quad \{p_i, p_i\} = \{x_i, x_i\} = 0$$
 (6)

and by the Hamiltonian

$$\mathcal{H}_0 = \alpha(\mathbf{p}, \mathbf{r})\Phi = \alpha(\mathbf{p}, \mathbf{r}) \left(\frac{p^2}{n^2(\mathbf{r})} - \lambda_0^{-2} \right) \approx 0.$$
 (7)

Here α is the Lagrangian multiplier which could be an arbitrary function of coordinates and momenta, and i, j = 1, 2, 3. The notation "weak zero," $\mathcal{H}_0 \approx 0$, means that when writing down the Hamiltonian equations of motion we should take into account the constraint (5) only after the differentiation:

$$\frac{df(\mathbf{r}, \mathbf{p})}{d\tau} = \{f, \mathcal{H}_0\} = \{f, \alpha\}\Phi + \alpha\{f, \Phi\} \approx \alpha\{f, \Phi\}. \quad (8)$$

The arbitrariness in the choice of the function α reflects the reparametrization invariance of (1). For the description of the

equations of motion in terms of arc length of the original Euclidian space one should choose (see, e.g., [13])

$$\alpha = \frac{n^2(\mathbf{r})}{p + \lambda_0^{-1} n(\mathbf{r})} \quad \Rightarrow \quad \mathcal{H}_{\text{Opt}} = p - \lambda_0^{-1} n(\mathbf{r}). \tag{9}$$

With this choice, the equations of motion take the conventional form [14]

$$\frac{d\mathbf{p}}{dl} = \lambda_0^{-1} \nabla n(\mathbf{r}), \quad \frac{d\mathbf{r}}{dl} = \frac{\mathbf{p}}{p}, \tag{10}$$

where $dl := \alpha(\mathbf{r}, \mathbf{p})d\tau$ is the element of arc length. These equations describe the motion of a wave package with center coordinate \mathbf{r} and momentum \mathbf{p} in a medium with refraction index $n(\mathbf{r})$.

Assume we have a Hamiltonian system given by the Poisson bracket (6) and by the Hamiltonian

$$H = \frac{p^2}{2g(\mathbf{r})} + V(\mathbf{r}). \tag{11}$$

In accordance with the Maupertuis principle, after fixing the energy surface H = E, we can relate its trajectories with the optical Hamiltonian (7) with the refraction index

$$n(\mathbf{r}) = \lambda_0 \sqrt{2g(\mathbf{r})[E - V(\mathbf{r})]}.$$
 (12)

Clearly, the optical Hamiltonian (7) [as well as the Hamiltonian (9)] with the refraction index (12) inherits all the symmetries and constants of motion of the Hamiltonian (11).

Canonical transformations preserve the symmetries of the Hamiltonians and their level surfaces. Hence, we are able to construct the physically nonequivalent optical Hamiltonians (and refraction indices) with the identical symmetry algebra. The simplest illustration is the well-known relation between the Coulomb Hamiltonian which defines the so-called Coulomb refraction index profile and the free-particle Hamiltonian on the three-dimensional sphere, which defines the "Maxwell fish eye" refraction index (see, e.g., [15]). First, we fix the energy surface of the Coulomb Hamiltonian and get the respective refraction index:

$$H_{\text{Coul}} - E := \frac{p^2}{2} - \frac{\gamma}{r} - E = 0$$

 $\Rightarrow n_{\text{Coul}} = \lambda_0 \sqrt{2(E + \gamma/r)}, \text{ where } \gamma > 0.$ (13)

The constants of motion of the Coulomb problem (and of the respective optical Hamiltonian) are given by the rotational momentum and by the Runge-Lenz vector

$$\mathbf{L} = \mathbf{r} \times \mathbf{p}, \quad \mathbf{A} = \mathbf{L} \times \mathbf{p} + \gamma \frac{\mathbf{r}}{r}$$
 (14)

which form the algebra

$$\{A_i, A_j\} = -2\varepsilon_{ijk}H_{\text{Coul}}L_k, \quad \{A_i, L_j\} = \varepsilon_{ijk}A_k,$$

$$\{L_i, L_j\} = \varepsilon_{ijk}L_k.$$
 (15)

Now, let us perform a simple canonical transformation,

$$(\mathbf{p}, \mathbf{r}) \to (-\mathbf{r}, \mathbf{p}). \tag{16}$$

As a result, the first equation in (13) reads

$$r^2 - \frac{2\gamma}{p} - 2E = 0 \implies p - \frac{2\gamma}{r^2 - 2E} = 0.$$
 (17)

Interpreting the second equation as an optical Hamiltonian, we get the refraction index profile known as the "Maxwell fish eye" (3) with the parameters κ and n_0 defined as

$$\kappa := -\frac{1}{2E}, \quad \frac{n_0}{\lambda_0} := 2\epsilon \kappa \gamma, \tag{18}$$

where $\epsilon = -\operatorname{sgn}(r^2 + 1/\kappa)$.

The integrals of motion (14) result in the symmetry generators of the optical Hamiltonian with the Maxwell fish eye refraction index:

$$L \to L, \quad A \to \frac{T}{2\kappa},$$

$$\mathbf{T} = (1 - \kappa r^2)\mathbf{p} + 2\kappa(\mathbf{r}\mathbf{p})\mathbf{r} = \left(2 - \frac{n_0}{n_{Mfe}(\mathbf{r})}\right)\mathbf{p} + 2\kappa(\mathbf{r}\mathbf{p})\mathbf{r}.$$
(19)

These integrals form the so(4) algebra for $\kappa > 0$, and so(1.3) algebra for $\kappa < 0$:

$$\{L_i, L_j\} = \varepsilon_{ijk}L_k, \quad \{T_i, L_j\} = \varepsilon_{ijk}T_k, \quad \{T_i, T_j\} = 4\kappa \varepsilon_{ijk}L_k.$$
 (20)

In the next sections we will use the above described duality for the construction of the Maxwell fish eye profile for polarized light.

III. INCLUSION OF POLARIZATION

Let us briefly discuss the inclusion of polarization.

To this end we should add to the scalar Lagrangian $L_0 = \mathbf{p}\dot{\mathbf{r}} - p + \lambda_0^{-1}n$ the additional term $L_1 = -s\mathbf{A}(\mathbf{p})\dot{\mathbf{p}}$, where s is the spin of the photon and \mathbf{A} is the the vector-potential of the "Berry monopole" [i.e., the potential of the magnetic (Dirac) monopole located at the origin of momentum space] [13]:

$$\mathbf{F} := \frac{\partial}{\partial \mathbf{p}} \times \mathbf{A}(\mathbf{p}) = \frac{\mathbf{p}}{p^3}.$$
 (21)

From the Hamiltonian viewpoint this means to preserve the form of the Hamiltonian (7) and replace the canonical Poisson brackets (6) by the twisted ones

$$\{x_i, p_j\} = \delta_{ij}, \quad \{x_i, x_j\} = s\varepsilon_{ijk}F_k(\mathbf{p}), \quad \{p_i, p_j\} = 0, \quad (22)$$

where i, j, k = 1, 2, 3, and F_k are the components of the Berry monopole (21). On this phase space the rotation generators take the form

$$\mathbf{J} = \mathbf{r} \times \mathbf{p} + s \frac{\mathbf{p}}{p} \tag{23}$$

while the equations of motion read

$$\frac{d\mathbf{p}}{dl} = \lambda_0^{-1} \nabla n(\mathbf{r}), \quad \frac{d\mathbf{r}}{dl} = \frac{\mathbf{p}}{p} - \frac{s}{\lambda_0} \mathbf{F} \times \nabla n(\mathbf{r}). \tag{24}$$

However, the above procedure, i.e., twisting the Poisson bracket with preservation of the Hamiltonian, violates the nonkinematical (hidden) symmetry of the system. To get the profiles admitting the symmetries in the presence of polarization, we use the following observation [16] (see [17] for its quantum counterpart). Assume we have the three-dimensional rotationally invariant system

$$\mathcal{H}_0 = \frac{p^2}{2g(r)} + V(r), \quad \{x_i, p_j\} = \delta_{ij},$$

$${p_i, p_j} = {x_i, x_j} = 0.$$
 (25)

For the inclusion of interaction with a magnetic monopole, we should transition from the canonical Poisson brackets to the twisted ones:

$$\{x_i, p_j\} = \delta_{ij}, \quad \{p_i, p_j\} = s\varepsilon_{ijk}\frac{x_k}{r^3}, \quad \{x_i, x_j\} = 0.$$
 (26)

The rotation generators then read

$$\mathbf{J} = \mathbf{r} \times \mathbf{p} + s \frac{\mathbf{r}}{r} : \quad \{J_i, J_j\} = \varepsilon_{ijk} J_k. \tag{27}$$

By modifying the initial Hamiltonian to

$$\mathcal{H}_s = \frac{p^2}{2g(r)} + \frac{s^2}{2g(r)r^2} + V(r), \tag{28}$$

we find that trajectories of the system preserve their form, but the plane to which they belong fails to be orthogonal to the the axis **J**. Instead, it turns to the constant angle

$$\cos \theta_0 = \frac{s}{|\mathbf{J}|}.\tag{29}$$

For systems with hidden symmetries, one can find the appropriate modifications of the hidden symmetry generators respecting the inclusion of the monopole field.

To apply this observation to systems with polarized light, we should choose the appropriate integrable system with a magnetic monopole, and then perform the canonical transformation (16) which yields the Poisson brackets for polarized light (26). Afterwards we need to solve the equation

$$r^2 + \frac{s^2}{p^2} - 2g(p)[E - V(p)] = 0 \implies p = \frac{n_{inv}^s(r)}{\lambda_0}.$$
 (30)

For example, to get the "polarized Coulomb profile" we have to start from the free-particle Hamiltonian on a three-dimensional sphere or hyperboloid interacting with the Dirac monopole:

$$H_s = \frac{(1 + \kappa r^2)^2}{2} \left(p^2 + \frac{s^2}{r^2} \right). \tag{31}$$

Then, after fixing the energy surface $H_s = E$ and performing canonical transformation (16) we arrive at the third-order (with respect to p^2) algebraic equation

$$(1 + \kappa p^2)^2 \left(r^2 + \frac{s^2}{p^2}\right) = 2E(>0) \Leftrightarrow y^3 u - y^2 (u - \kappa s^2)$$
$$-Ey + E = 0, \tag{32}$$

with $y := 1 + \kappa p^2$, $u := r^2$.

This equation has either one real and two complex solutions or three real solutions, which describe the "polarized Coulomb profiles."

Conversely, when we start from the Coulomb problem with a Dirac monopole we will arrive at the "polarized Maxwell fish eye," i.e. the deformation of the "Maxwell fish eye" which preserves, in the presence of polarized light, all symmetries of initial scalar system. The latter is considered in detail in the next section.

IV. POLARIZED MAXWELL FISH EYE

Let us consider the Coulomb system with a Dirac monopole, which is known as the "MICZ-Kepler system" [18]. It is defined by the twisted Poisson brackets (26) and by the Hamiltonian

$$H_{MICZ} = \frac{p^2}{2} + \frac{s^2}{2r^2} - \frac{\gamma}{r}.$$
 (33)

Besides the conserved angular momentum (27), this system has the conserved Runge-Lenz vector

$$\mathbf{A}_{s} = \mathbf{J} \times \mathbf{p} + \gamma \frac{\mathbf{r}}{r},\tag{34}$$

which forms the symmetry algebra of Coulomb problem (20) [with the replacement $(\mathbf{L}, \mathbf{A}) \rightarrow (\mathbf{J}, \mathbf{A}_s)$]. After performing canonical transformation (16), we get

$$H_{MICZ} = E \Leftrightarrow r^2 + \frac{s^2}{p^2} - \frac{2\gamma}{p} - 2E = 0.$$
 (35)

Solving this quadratic equation for p, we get the refraction index given by the expression (4), where the notation (18) is used.

The rotation generator (27) transforms to (23), and the Runge-Lenz vector (34) transforms to \mathbf{T}_s/κ , where

$$\mathbf{T}_{s} = \left(2 - \frac{n_0}{n_{Mfe}^{s}(\mathbf{r})}\right)\mathbf{p} + 2\kappa(\mathbf{r}\mathbf{p})\mathbf{r} + \frac{2\kappa s}{n_{Mfe}^{s}(\mathbf{r})}\mathbf{J}.$$
 (36)

Along with (23), these generators form the symmetry algebra of the original Maxwell fish eye profile (20) [where the pair (\mathbf{L} , \mathbf{T}) is replaced by (\mathbf{J} , \mathbf{T}_s)]. The Casimirs of the symmetry algebra are given by the expressions

$$\mathbf{T}_{s}^{2} + 4\kappa(\mathbf{J}^{2} - s^{2}) = \frac{n_{0}^{2}}{\lambda_{0}^{2}}, \quad \mathbf{T}_{s} \cdot \mathbf{J} = \frac{sn_{0}}{\lambda_{0}}.$$
 (37)

Hence, for $\kappa > 0$ the vectors $\sqrt{4\kappa} \mathbf{J}$ and \mathbf{T}_s form a parallelogram with the fixed lengths of diagonals

$$|\mathbf{T}_s \pm \sqrt{4\kappa} \mathbf{J}| = \left| \frac{n_0}{\tilde{\chi}_0} \pm \sqrt{4\kappa} s \right|. \tag{38}$$

This immediately leads to the conclusion that for $\kappa > 0$ the generators \mathbf{T}_s and \mathbf{J} reach the lower or upper bounds being parallel to each other:

$$\left(|\mathbf{J}|_{\min} = s, |\mathbf{T}_s|_{\max} = \frac{n_0}{\lambda_0}\right),$$

$$\times \left(|\mathbf{J}|_{\max} = \frac{n_0}{\lambda_0 \sqrt{4\kappa}}, |\mathbf{T}_s|_{\min} = \sqrt{4\kappa s}\right). \tag{39}$$

Notice also that for $\kappa > 0$ we get a restriction of rays in the finite domain

$$\kappa > 0 : r \leqslant \sqrt{\frac{n_0^2}{4s^2\lambda_0^2\kappa^2} - \frac{1}{\kappa}}.$$
(40)

One can also note that spin appears in the expression for the refraction index (4) along with the factor $\kappa \lambda_0^2 = (\lambda_0/2r_0)^2$. In order to stay within the bounds of the geometrical optics approximation, this factor must be reasonably small. Therefore, the influence of the spin will be far more notable within a certain range of distance from the core of the fish eye.

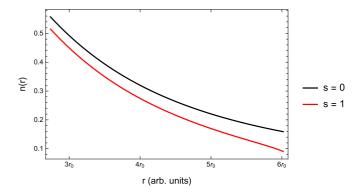


FIG. 1. Maxwell fish eye refraction index profile for s = 0 and s = 1 when $n_0 = 1.5$, $\lambda_0 = 1$, $r_0 = 5$.

The latter happens when the condition $4\kappa s^2 \hat{\tau}_0^2 / n_0 \approx n_{Mfe}(\mathbf{r})$ holds. At these distances the refraction index in the presence of spin can be much smaller compared to the refraction index with zero spin (see Fig. 1).

V. TRAJECTORIES

Let us study the ray trajectories of the polarized light propagating in the medium with the above constructed profile (4). One can see that

$$\mathbf{r} \cdot \mathbf{J} = s \frac{\mathbf{r} \mathbf{p}}{p}, \quad \mathbf{r} \cdot \mathbf{T}_s = \frac{n_0}{\lambda_0} \frac{\mathbf{r} \mathbf{p}}{p} \quad \Rightarrow \quad \mathbf{r} \cdot \left(\mathbf{J} - \frac{s \lambda_0}{n_0} \mathbf{T}_s \right) = 0.$$
 (41)

Hence, ray trajectories are orthogonal to the axis

$$\mathbf{E}_3 = \mathbf{J} - \frac{s\dot{\mathbf{x}}_0}{n_0} \mathbf{T}_s,\tag{42}$$

and, therefore, the trajectories belong to the plane spanned by the following vectors:

$$\mathbf{E}_{1} = \mathbf{T}_{s} \times \mathbf{J},$$

$$\mathbf{E}_{2} = \mathbf{E}_{3} \times \mathbf{E}_{1} = (\mathbf{J}^{2} - s^{2}) \left(\mathbf{T}_{s} - \frac{4s\lambda_{0}\kappa}{n_{0}} \mathbf{J} \right) :$$

$$\mathbf{E}_{3} \cdot \mathbf{E}_{2} = \mathbf{E}_{3} \cdot \mathbf{E}_{1} = 0.$$
(43)

Then, from the expression $J \cdot (r \times T_s)$ we immediately obtain the solution for the ray trajectories:

$$\mathbf{r} \cdot (\mathbf{T_s} \times \mathbf{J}) = (J^2 - s^2) \left(2 - \frac{n_0}{n_{mfe}^s} \right). \tag{44}$$

This prompts us to introduce the following orthogonal frame:

$$\mathbf{e}_i = \frac{\mathbf{E}_i}{|\mathbf{E}_i|} : \quad \mathbf{e}_i \cdot \mathbf{e}_j = \delta_{ij}, \tag{45}$$

where

$$|\mathbf{E}_1|^2 = (\mathbf{J}^2 - s^2) \left(\frac{n_0^2}{\lambda_0^2} - 4\kappa \mathbf{J}^2 \right),$$

$$|\mathbf{E}_3|^2 = (\mathbf{J}^2 - s^2) \left(1 - \frac{4s^2 \lambda_0^2 \kappa}{n_0^2} \right), \quad |\mathbf{E}_2|^2 = |\mathbf{E}_1|^2 |\mathbf{E}_3|^2.$$
 (46)

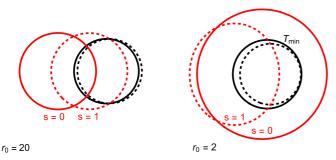


FIG. 2. Deformations of the ray trajectories for different values of r_0 when $n_0 = 1.5$, $\lambda_0 = 1$. The black curves correspond to the basic trajectories where $T = T_{min} = s/r_0$. The red (light gray) curves correspond to trajectories corresponding to the same value of T. Dashed curves are the trajectories corresponding to the same value of T but for circularly polarized light (s = 1). The first figure ($r_0 = 20$) corresponds to the case when the deformations of the profile only result in the shift of the centers of the trajectories, without affecting their shapes. Conversely, in the second figure ($r_0 = 2$), the deformations of the Maxwell fish eye result in highly deformed trajectories.

Decomposing ${\bf r}$ over this frame, we introduce the polar coordinates

$$\mathbf{r} = x_1 \mathbf{e}_1 + x_2 \mathbf{e}_2, \quad x_1 = r \cos \varphi, \quad x_2 = r \sin \varphi. \tag{47}$$

Then, having in mind Eqs. (43) and (46), we can immediately rewrite the equation (44) in polar coordinates:

$$1 - |\kappa| |\mathbf{a}_s| r \cos \varphi = \frac{1 + \kappa r^2}{1 + \sqrt{1 - \frac{4\kappa s^2 \lambda_0^2}{n_0^2} (1 + \kappa r^2)}},$$
 (48)

where

$$|\mathbf{a}_s|^2 := R_s^2 - \frac{1}{\kappa}, \quad R_s^2 := \frac{n_0^2 - 4\kappa s^2 \lambda_0^2}{4\lambda_0^2 \kappa^2 (J^2 - s^2)}.$$
 (49)

So the trajectories of polarized light are not circles anymore, in contrast to the case of scalar waves. However, they can be described by the parameters $\mathbf{a_s}$ and R_s (49) which, in the limit $s \to 0$, become the center coordinate and the radius of the circle, respectively. Indeed, for s = 0 the solution (44) results in the the equation for a circle with the center located at the $\mathbf{e_1}$ axis:

$$s = 0$$
: $\mathbf{r} \cdot (\mathbf{T} \times \mathbf{L}) = L^2 (1 - \kappa r^2) \Rightarrow (\mathbf{r} - \mathbf{a}_0)^2 = R_0^2,$ (50)

where

$$\mathbf{a}_0 := \frac{\mathbf{T} \times \mathbf{L}}{2\kappa L^2}, \quad |\mathbf{a}_0|^2 = R_0^2 - \frac{1}{\kappa}, \quad R_0 := \frac{n_0}{2|\kappa|\lambda_0 L}.$$
 (51)

Given $s\lambda_0 \ll n_0/\kappa$, the deformation of the circle is negligible, which is not the case for $s\lambda_0 \sim n_0/\sqrt{\kappa}$.

As we can see from Fig. 2, for $r_0 = 20$ the only notable manifestation of the polarization is the shift of the center of the trajectory. However, it is worth noting that since the dashed lines are not circles anymore when talking about the center of the trajectory for s = 1, we refer to the point which becomes the center of the circular trajectory when we pass from s = 1 to s = 0. The second picture illustrates the circular trajectories

and their deformations for $s\lambda_0 \sim n_0/\sqrt{\kappa}$. In this case, the original profile (3) and the deformed one (4) differ drastically. The circular trajectories are notably deformed.

Detailed knowledge of trajectory parameters (49) can be used in different applications. For example, in the conformal mapping scheme the cloaking area is the outer space of closed trajectories [3]. Therefore as it follows from (49) there is no cloaking for polarized photons when $J \rightarrow s$.

VI. CONCLUDING REMARKS

The standard Maxwell fisheye profile does not ensure closed ray trajectories for polarized photons (the only exceptions are linearly polarized photons corresponding to the s=0 spin statistic), while the closeness of trajectories is the main property that is used in perfect imaging and cloaking phenomena. In this paper, we suggested the deformation of the Maxwell fisheye profile, which ensures the closeness of the trajectories of the ray trajectories for the polarized photons. We examined the properties of the deformed profile and have shown that the main difference between the cases of polarized and nonpolarized photons is observed at the vicinity of wave and geometrical optics border $s\lambda_0/r_0 \sim 1$.

The proposed modification scheme is applicable for any isotropic refraction index n(r). Namely, to preserve the qualitative properties of scalar wave trajectories for the propagating polarized light, we should replace it with the modified index $n^s(r)$ which is the solution (with respect to p) of the following equation:

$$p = \frac{1}{\hbar_0} n \left(\sqrt{r^2 + \frac{s^2}{p^2}} \right) \quad \Rightarrow \quad p = n^s(r), \tag{52}$$

where *s* is polarization of light. The proposed deformation preserves the additional symmetries of the system (if any), and thus guarantees the closeness of trajectories of polarized light.

Seemingly, the suggested scheme could be extended to some nonisotropic but integrable profiles as well. On the other hand, nonisotropic integrable profiles are not common objects in the present study, though they obviously can be constructed by the use of existing integrable models. For example, choosing a textbook integrable system, the two-center Coulomb problem [1], and performing trivial canonical transformation (16) we can construct (taking into account the expressions for constants of motion; see, e.g., [19]) an anisotropic profile which could be interpreted as a superposition of two "Maxwell fish eye" profiles. Furthermore, using the proposed scheme, we can construct a "polarized Maxwell double fish eye" profile as well, starting from the "two-center MICZ-Kepler problem" [20], i.e., from the two-center Coulomb problem specified by the presence of magnetic monopoles located at the attraction centers. We hope to consider this problem elsewhere.

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