Continuous real-time sensing with a nitrogen-vacancy center via coherent population trapping

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We propose and theoretically analyze the use of coherent population trapping of a single diamond nitrogenvacancy (NV) center for continuous real-time sensing. The formation of the dark state in coherent population trapping prevents optical emissions from the NV center. Fluctuating magnetic fields, however, can kick the NV center out of the dark state, leading to a sequence of single-photon emissions. A time series of the photon counts detected can be used for magnetic field estimations, even when the average photon count per update time interval is much smaller than 1. For a theoretical demonstration, the nuclear spin bath in a diamond lattice is used as a model fluctuating magnetic environment. For fluctuations with known statistical properties, such as an Ornstein-Uhlenbeck process, Bayesian inference-based estimators can lead to an estimation variance that approaches the classical Cramer-Rao lower bound and can update dynamical information in real time with the detection of just a single photon. Real-time sensing using coherent population trapping adds a powerful tool to the emerging technology of quantum sensing.

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I. INTRODUCTION

Quantum sensing exploits the sensitivity of a simple quantum system, such as solid-state spins, cold atoms, or superconducting circuits, to a given physical quantity to derive an estimate for the physical quantity. The near-term prospect for emerging quantum technologies and the potential for discoveries in wide ranging research areas have stimulated intense research efforts in quantum sensing [1,2]. One of the most promising systems for quantum sensing is a negatively charged nitrogen-vacancy (NV) center in diamond, which can enable sensing of magnetic field, electric field, temperature, and strain with nanometer spatial resolution [3-6].

The most widely used quantum sensing approach has been Ramsey interferometry or Ramsey fringes, which probe the coherent time evolution of a single or a collection of two-level systems [3,7]. Incorporation of Bayesian phase estimations, adaptive measurements, and machine learning in Ramsey interferometry can further improve the sensitivity, increase the dynamical range, and reach the Heisenberg limited scaling of the underlying sensing process [8–14]. Ramsey interferometry is especially suitable for sensing of static as well as periodic signals. Repeated Ramsey interferometric measurements also allow the tracking of time-dependent fields [13,15,16]. Special Bayesian phase estimation and wave-form reconstruction techniques have been developed for the time-dependent sensing [12,15,17]. Nevertheless, since each Ramsey interferometric measurement consists of three separate stages, initialization, coherent time evolution, and read-out, Ramsey interferometry cannot provide continuous real-time sensing.

Here we propose and theoretically analyze the use of coherent population trapping (CPT) in a NV center for continuous real-time sensing. In a CPT process, a Λ -type three-level system is trapped in a dark state, i.e., a special superposition of the two lower states, which is decoupled from the upper state due to destructive quantum interference [18]. The formation of the dark state prevents optical emissions from the NV center. A fluctuating magnetic environment, however, can kick the NV center out of the dark state, leading to a sequence of single-photon emissions from the NV center. These single-photon emissions carry the information on the fluctuating magnetic field and can in principle provide continuous real-time sensing of the magnetic field, as illustrated schematically in Fig. 1. Nevertheless, the overall collection or detection efficiency for optical emissions from a NV center is only a few percent under typical experimental conditions. For a realistic implementation, a primary challenge is to make continuous real-time estimations of the magnetic field with the few photons detected.

We show that this challenge can be overcome with the use of Bayesian inference. For the theoretical analysis, we have used the nuclear spin bath in diamond as a model fluctuating magnetic environment. We have employed the stochastic Schrodinger equation (SSE) to simulate single-photon emissions from a NV center in a CPT setting [19] and have treated the nuclear spin bath as an Ornstein-Uhlenbeck (OU) process [20-22]. Estimations of fluctuating magnetic fields are obtained from a time series of photon counts, for which the average number of photons per update time interval is much smaller than 1. By taking advantage of the known statistical properties of the fluctuating environment, we demonstrate that the Bayesian estimator can provide dynamical information on a time scale that is comparable to the inverse of the average photon counting rate and specifically can update dynamical information in real time with the detection of just a single photon. Additional theoretical analysis also shows that the Bayesian estimator can approach the classical Cramer-Rao lower bound (CRLB).

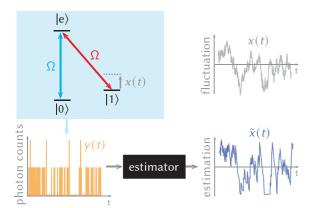


FIG. 1. Schematic of CPT-based continuous real-time sensing of the fluctuating magnetic environment at a NV center. The fluctuating magnetic field, which induces single-photon emissions from the NV center, can be estimated from the time series of the photon counts detected.

Note that continuous sensing using resonant fluorescence of a two-level system has been proposed and analyzed in earlier studies, for which the distribution of waiting times between detected photon counts is used for Bayesian estimations of parameters such as Rabi frequency or laser detuning [23-26]. This scheme has also been extended to a Λ -type three-level system, for which two-channel photon counting is used, though the role of CPT was not examined. The Bayesian estimations in these studies use the entire data set including data pre- and post-time of interest, yielding information on the past quantum state [27]. As such, these estimations do not provide information in real time.

II. PHYSICAL MODEL

A. Coherent population trapping

For a NV center in diamond, several energy-level schemes have been used for the realization of CPT [28–32]. Without losing generality, here we consider two ground spin states, $m_s = 0$ and $m_s = 1$, coupling to an excited state, $|e\rangle$, through two dipole optical transitions with frequencies v_0 and v_1 , respectively. Two external optical fields with frequencies ω_0 and ω_1 couple to the two respective optical transitions with equal Rabi frequency Ω , as shown schematically in Fig. 1. With $\omega_0 \approx v_0, \omega_1 \approx v_1$, and $\rho_{ee} \ll 1$, the steady-state excited-state population is given by [31]

$$\rho_{ee} = \frac{\Omega^2}{2\Gamma\kappa} \left[1 - \frac{\Omega^2}{2\kappa} \frac{\gamma_s + \Omega^2/2\kappa}{(\delta - \omega_B)^2 + (\gamma_s + \Omega^2/2\kappa)^2} \right], \quad (1)$$

where γ_s and κ are the decay rates for the spin coherence and optical dipole coherence, respectively, Γ is the spontaneous emission rate of the excited state, $\delta = \omega_0 - \omega_1$ is the laser detuning, and $\omega_B = \nu_0 - \nu_1$ is the frequency separation between the two spin states. CPT, which corresponds to the formation of a dark state for the two lower spin states, occurs near the Raman resonance with $\Delta = \delta - \omega_B = 0$. Note that the external optical fields lead to a power broadening described by the term $\Omega^2/2\kappa$ in Eq. (1). The CPT process can be characterized by a dimensionless cooperativity, defined as $C = \Omega^2/2\kappa \gamma_s$. CPT-based sensing will be carried out in the regime of $C \gg 1$.

B. Nuclear spin bath

We simulate the magnetic field fluctuations induced by the nuclear spin bath as an OU process characterized by a memory time τ_N [33], with

$$dx = -\frac{1}{\tau_N} x dt + \sqrt{\frac{2\sigma^2}{\tau_N}} dW_t, \qquad (2)$$

where $x(t) = \omega_B(t) - \langle \omega_B(t) \rangle$ represents the bath-induced fluctuation in ω_B , dW_t denotes a Wiener increment that has a Gaussian distribution with mean zero and variance dt, and the autocorrelation function for x(t) is characterized by

$$R(t) = \langle x(t_0)x(t_0+t) \rangle = \sigma^2 e^{-|t|/\tau_N}.$$
 (3)

For the two bath parameters, σ can be derived directly from the decay of the Ramsey fringes and τ_N can be obtained from additional experiments of spin echoes [22]. For the numerical simulations in this paper, we take $\tau_N = 1$ ms and $\sigma/2\pi = 0.13$ MHz (which corresponds to a dephasing time $T_2^* = 1.7 \,\mu$ s [34]) for a ¹³C nuclear spin bath in diamond. Figure 2(a) shows an example of the simulated x(t) and the autocorrelation function averaged over 1000 runs.

C. Time series of photon counts

Magnetic field fluctuations induced by the nuclear spin bath of the diamond carbon lattice lead to corresponding fluctuations in ω_B , which can kick a NV center out of or into the dark state in a CPT setting. This in turn induces fluctuations in the excited state population and leads to a sequence of singlephoton emissions. Note that the nuclear spin bath fluctuates at a time scale many orders of magnitude slower than the NV radiative lifetime (12 ns [35]). It takes only a few spontaneous emission events for the NV center to reach the steady state, as confirmed in an earlier experimental study [36]. In this CPT setting, the single-photon emissions from the excited state carry the information on the fluctuating nuclear spin bath.

We have used the SSE to simulate single-photon emissions from a Λ -type three-level system in a CPT setting (see Appendix A), for which we take $\Gamma/2\pi = 13$ MHz [35]. A Rabi frequency of $\Omega/2\pi = 2.8$ MHz and a Raman detuning or bias of $\Delta_0 = \delta - \langle \omega_B \rangle = 2\pi \times 0.25$ MHz are also used in the numerical simulation. Figure 2(b) shows the time series of the photons emitted by the three-level system, along with the corresponding excited state population. Only a small fraction of these photons are detected in a realistic sensing experiment. Figure 2(c) plots the time series of the photons detected, for which an overall collection or detection efficiency, $\eta = 1.6\%$, is used. Under these conditions, the average photon count rate is approximately 10000 per second. For comparison, the underlying fluctuations in ω_B are also plotted in Fig. 2(c). Because of the very low photon counting rate, it is difficult to directly discern the pattern of the fluctuations from the time series of the detected photon counts.

Alternatively, we can simulate single-photon emissions using the population of the excited state with the assumption that the system has a Poisson distributed probability to emit a pho-

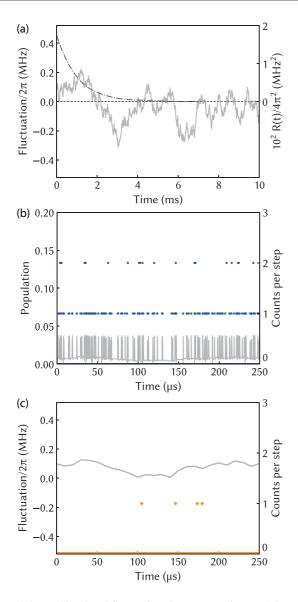


FIG. 2. (a) Simulated fluctuations in ω_B (grey line) and the autocorrelation R(t) (dot-dashed line). (b) The excited state population (grey lines) as a function of time and the corresponding time series of photon counts emitted by the NV center (blue dots). (c) The time series of the photon counts detected (orange diamond) and the underlying fluctuation in ω_B . The CPT parameters used are $(\Omega, \Gamma, \Delta_0)/2\pi = (2.8, 13, 0.25)$ MHz. The step size used for the photon counting is 0.1 μ s.

ton with the average detection rate given by $\eta \Gamma \rho_{ee}$. For this simulation, we make the approximation that ρ_{ee} adiabatically follows the magnetic field of the bath, since the time scale of the bath fluctuation is orders of magnitude longer than the NV spontaneous emission lifetime. As shown in Appendix A, magnetic field estimations using the single-photon emissions obtained with the steady-state ρ_{ee} exhibit the same behavior as those obtained with the SSE. For numerical calculations that require more than a few hundred runs, we have thus simulated single-photon emissions using the steady-state ρ_{ee} .

Note that for the simulation of single-photon emissions, we have set $\gamma_s = 0$ since in the limit of $C \gg 1$, the CPT spectral

linewidth is completely dominated by power broadening, as shown in Eq. (1). In this limit, intrinsic decoherence of the NV center does not affect CPT-based real-time sensing, as will be discussed in more detail in Appendix A.

III. ESTIMATORS

In this section, we discuss estimators, which we have used to output a time series of estimated frequencies, $\{\tilde{\mathbf{x}}_n\} = \{\tilde{x}_1, \tilde{x}_2, \dots, \tilde{x}_n, \dots\}$, from a time series of photon counts, $\{\mathbf{y}_n\} = \{y_1, y_2, \dots, y_n, \dots\}$, where y_n is the number of photons detected during the *n*th time interval and the duration of the time interval τ is small compared with τ_N . We will also discuss how close these estimations can be to the actual frequencies, $\{\mathbf{x}_n\} = \{x_1, x_2, \dots, x_n, \dots\}$.

A. Bayesian estimator

We have used the Bayesian inference described by the Bayes update rule,

$$p(x_n|y_n, y_{n-1}, \dots, y_1) \propto p_{\bar{y}_n}(y_n|x_n) \times p'(x_n|y_{n-1}, \dots, y_1).$$

(4)

to estimate x(t) from the time series of photon counts, where $p'(x_n|y_{n-1}, \ldots, y_1)$ is the prior probability distribution based on the previous time series of photon counts, $p(x_n|y_n, y_{n-1}, \ldots, y_1)$ is the *a posteriori* probability distribution, and $p_{\bar{y}_n}(y_n|x_n)$ is the likelihood of detecting y_n photons in the *n*th time interval given x_n . The likelihood follows a Poisson distribution,

$$p_{\bar{y}_n}(y_n|x_n) = \frac{\bar{y}_n^{y_n} e^{-\bar{y}_n}}{y_n!},$$
(5)

where $\bar{y}_n = \eta \tau \Gamma \rho_{ee}(x_n)$ is the average number of detected photons expected for the given time interval and the update interval τ is sufficiently short such that \bar{y}_n is small compared with 1. The estimation of x(t) as a function of time is then given by

$$\tilde{x}(t) = \int p(x,t)xdx.$$
(6)

The prior probability distribution for x_n can be improved if the statistical properties of the fluctuating field are known. For the OU process, the probability of finding x_n at $t + \tau$ given x_{n-1} at *t* is given by a normal distribution,

$$p_{OU}(x_n, t + \tau | x_{n-1}, t) = N\{x_{n-1}e^{-\tau/\tau_N}, \sigma^2[1 - \exp(-2\tau/\tau_N)]\}$$
(7)

where the normal distribution has a mean of $x_{n-1}e^{-\tau/\tau_N}$ and a variance of $\sigma^2[1 - \exp(-2\tau/\tau_N)]$. In this case, the prior probability distribution for x_n can be improved as

$$p'(x_{n} = x|y_{n-1}, \dots, y_{1})$$

$$= \int d\omega p(x_{n-1} = x - \omega|y_{n-1}, \dots, y_{1})$$

$$\times p_{OU}(x_{n} = x_{n-1} + \omega, t + \tau |x_{n-1}, t)$$
(8)

We will refer to this improved Bayesian estimation process as the OU Bayesian estimator. Note that the OU Bayesian estimator has also been used in an earlier study on the estimations of magnetic field fluctuations in an atomic ensemble [37].

B. Average count estimator

For comparison, we have also used the photon counts detected in a relatively long duration, τ_a , to directly estimate x(t). Specifically, we carry out the estimation by using

$$y_n^{(a)} = \eta \Gamma \tau_a \rho_{ee}(x_n), \tag{9}$$

where $y_n^{(a)}$ is the number of photons detected between time $n\tau - \tau_a$ and $n\tau$. For achieving acceptable photon counts (of order 10), we took $\tau_a = 100\tau$ for the results presented in Sec. IV. This average count approach is not expected to be able to track bath fluctuations that occur in a relatively short time scale.

Note that electromagnetically induced transparency (EIT) of an ensemble of NV centers, which is closely related to CPT, has been used for static sensing [38]. CPT of an ensemble of NV centers, which can feature orders of magnitude greater photon counting rates than those of a single NV, can also be exploited for time-dependent sensing. A single quantum sensor, such as a single NV, however, is necessary for sensing microscopic fluctuations or for sensing at nanometer resolution.

C. Cramer-Rao lower bound

How close an estimation is to the actual quantity is characterized by the average estimation variance, defined as $\operatorname{Var}[\tilde{x}_n(\mathbf{y})] = \langle [\tilde{x}_n(\mathbf{y}) - x_n(\mathbf{y})]^2 \rangle$. Theoretically, the classical CRLB sets a lower bound on the variance. As shown in Appendix **B**, the CRLB for the OU-Bayesian estimation is given by

$$\operatorname{Var}[\tilde{x}_{n}(\mathbf{y})] \geqslant \frac{\sigma^{2}}{\sqrt{1 + 2\tau_{N}\eta\Gamma\sigma^{2}g(\sigma)}},$$
(10)

where $g(\sigma) = \langle (\partial \rho_{ee}/\partial x)^2 / \rho_{ee}(x) \rangle_x$.

For real-time sensing, only the historical data can be used. Data taken after the estimation event cannot be used for the estimation. However, the CRLB in Eq. (10) assumes that the entire data set can be used for the estimation. As shown in Appendix B, the CRLB when only the historical data can be used is revised as

$$\operatorname{Var}[\tilde{x}_{n}(\mathbf{y})] \geq \frac{\sigma^{2}}{\sqrt{1 + 2\tau_{N}\eta\Gamma\sigma^{2}g(\sigma)}} \times \frac{2}{1 + 4/\sqrt{1 + 32\tau_{N}\eta\Gamma\sigma^{2}g(\sigma)}}, \quad (11)$$

where we have assumed that $\tau_N \eta \Gamma \sigma^2 g(\sigma) > 2$ (which is satisfied in most cases). In the limit that $\tau_N \eta \Gamma \sigma^2 g(\sigma) \gg 1$, the CRLB for estimations using the historical data set is twice that using the entire data set.

It should be noted that a necessary condition for achieving CRLB is that the posterior is a Gaussian distribution [39]. For our system, the posterior, $p(x_n|y_n, y_{n-1}, ..., y_1)$, is approximately Gaussian only when ρ_{ee} depends linearly or quadratically on *x*.

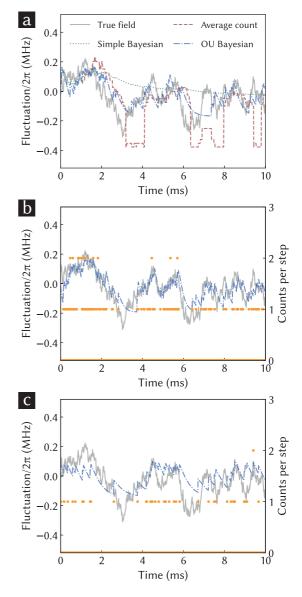


FIG. 3. (a) Estimations obtained with the average photon account estimator, the simple Bayesian estimator, and the OU Bayesian estimator. The CPT and bath parameters used are the same as those for Fig. 2, with an average count per time interval of 0.1. (b), (c) Comparison of the detected photon counts (orange dots) with the corresponding estimations obtained with the OU Bayesian estimator. The average photon count per time interval is 0.2 and 0.05 for (b) and (c), respectively, with other conditions the same as those used in (a). The solid grey lines in (a)–(c) plot the actual fluctuations.

IV. RESULTS AND DISCUSSION

We have used the three estimators discussed in Sec. III for estimations of ω_B from a sequence of single-photon emissions. The estimations as a function of time shown in Fig. 3(a) are based on the simulated fluctuations and the corresponding time series of photon counts shown in Appendix A. The update time interval used is $\tau = 10 \,\mu$ s and the average photon count per time interval is 0.1. As shown in Fig. 3(a), the estimations obtained with the average count estimator exhibit considerable deviations from the true frequency due to the low photon count rate. The estimations also feature significant delays with respect to the fluctuations in the actual frequency. The delays are due to the relatively long photon counting period τ_a needed to avoid excessive fluctuations in the photon counts used for the estimation. In comparison, the estimations obtained with the simple Bayesian estimator provides a good estimate for the relatively long time behavior of the frequency fluctuations. However, the photon counts are too low for the estimator to provide dynamical information on a relatively short time scale.

Much improved estimations are obtained with the OU Bayesian estimator, as shown in Fig. 3(a). By taking advantage of the known statistical properties of the OU process, the Bayesian estimator is able to yield dynamical information on a time scale much shorter than that achieved with the simple Bayesian estimator. The OU Bayesian estimator provides continuous real-time sensing of the fluctuating magnetic field, even when the time scale of the fluctuation is comparable to the inverse of the average photon count rate.

Figures 3(b) and 3(c) compare directly the detected photon counts with the corresponding estimations obtained with the OU Bayesian estimator, for which we varied the detection efficiency to achieve an average photon count per time interval of 0.2 (with $\eta = 3.2\%$) and 0.05 (with $\eta = 0.8\%$), respectively, with other conditions the same as those used for Fig. 3(a). As shown in Figs. 3(b) and 3(c), the Bayesian estimator can update dynamical information of the bath in real time with the detection of just a single photon. In comparison, an earlier study has used the complete CPT spectrum of a single NV for the sensing of the nuclear spin bath in diamond [29], for which it takes about 100 detected photons (or about 5 ms) to obtain a single estimation.

For a quantitative analysis of the effectiveness of the estimators, we have numerically calculated the estimation variances and have compared these variances with those expected from the CRLB. The variances shown in Fig. 4 are the averaged results of 100 runs. Each run covers the time duration from t = 2 ms to t = 10 ms and uses sequences of single photon emissions generated from the steady-state ρ_{ee} , as discussed in Sec. II C. The parameters used for the bath and the CPT process are the same as those used for Fig. 2 unless otherwise specified.

Figure 4(a) shows the variances obtained with the three different estimators as a function of the memory time τ_N of the nuclear spin bath. For comparison, Fig. 4(a) also plots the theoretically expected CRLB for the OU Bayesian estimator and the variance σ^2 of the actual frequency fluctuations. As expected, the variances obtained with the average count estimator are far above the CRLB. These variances also exceed σ^2 . The variances obtained with the simple Bayesian estimator fall below σ^2 but are still far above the CRLB. In comparison, the variances obtained with the OU Bayesian estimator nearly approach the CRLB. Both the variances and the CRLB improve gradually with increasing τ_N . By taking advantage of the known statistical properties of the fluctuating fields, the OU Bayesian estimator from the detected single photons.

For the OU Bayesian estimator, the estimations are expected to depend on deviations of the statistical parameters used in the estimator from the true parameters. Figure 4(b) shows the ratio of the estimation variances over the corre-

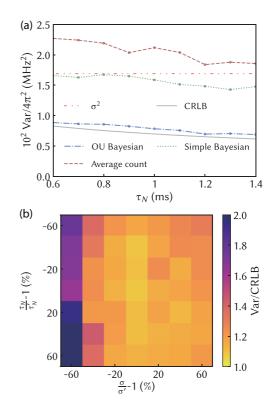


FIG. 4. (a) The average variances for estimations obtained with the average count estimator, simple Bayesian estimator, and OU Bayesian estimator as a function of the bath memory time τ_N . The estimation variances are also compared with the CRLB of the OU Bayesian estimator and with σ^2 , the variance of the fluctuating ω_B . (b) The ratio of the estimation variance over the CRLB as τ_N and σ used in the OU Bayesian estimator deviate from the true parameters, $\tau'_N = 1 \text{ ms and } \sigma'/2\pi = 0.13 \text{ MHz}.$

sponding CRLB as τ_N and σ used in the estimator move away from the true parameters. Variances close to the CRLB can still be obtained when the deviations of τ_N and σ from the true parameters are within 20%. The estimation variances, however, can exceed those obtained from the simple Bayesian estimator when τ_N and σ deviate from the true parameters by more than 50%. As discussed earlier, τ_N and σ can be deduced from experiments such as Ramsey fringes and spin echoes. For the experimental implementation of the OU Bayesian estimator, feedback control, which uses the estimations to keep the NV in the dark state, can also be employed for the further optimization of the statistical parameters used in the estimator. In addition, the feedback control can also serve as a verification for the estimations.

The choice of the CPT parameters, including Rabi frequency Ω and bias $\Delta_0 = \delta - \langle \omega_B \rangle$, also strongly affects the effectiveness of the estimations. For example, little information on the magnetic field fluctuations can be obtained from a sequence of single-photon emissions when Δ_0 is set to near 0. Figure 5(a) shows the variances for the OU Bayesian estimator as a function of Ω and Δ_0 , with all other parameters the same as those in Fig. 2. In this case, the optimal estimation is achieved with $\Omega/2\pi = 2.5$ MHz and $\Delta_0/2\pi = 0.2$ MHz. In comparison, the corresponding CRLBs shown in Fig. 5(b) decrease with decreasing Ω and Δ_0 , which is expected from

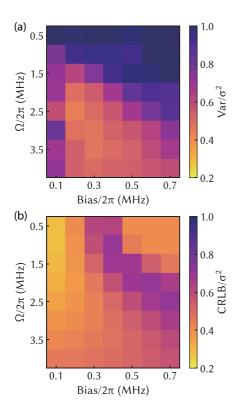


FIG. 5. (a) Estimation variances vs Ω and bias (i.e., Δ_0). The color bar corresponds to Var/ σ^2 . (b) CRLBs for the OU Bayesian estimator vs Ω and bias. The color bar corresponds to CRLB/ σ^2 .

Eq. (11). As can be seen from Eq. (11), the CRLB is minimized when $g(\sigma) = \langle (\partial \rho_{ee}/\partial x)^2 / \rho_{ee}(x) \rangle_x$ reaches maximum.

The optimal CPT parameters for the OU Bayesian estimator shown in Fig. 5(a) reflect a tradeoff between a steep $\partial \rho_{ee}/\partial x$ and the need to avoid the zero bias region. As discussed in more detail in Appendix C, this also corresponds to the tradeoff between the working range of the sensing process and how sensitive the sensing process can be. Note that as discussed in Sec. III C, the CRLB is only achievable when the posterior is a Gaussian distribution [39], which occurs when ρ_{ee} depends either linearly or quadratically on x. Near zero bias, the CRLB can no longer be achieved.

V. CONCLUSION

We have shown theoretically that CPT can enable continuous real-time sensing using a single quantum sensor such as a NV center in diamond. The use of Bayesian inference for the real-time sensing allows effective estimation of magnetic fields from a time series of detected photon counts, for which the average photon count per time interval is much smaller than 1, and can lead to estimation variances that approach the CRLB and to a time resolution that is comparable to the inverse of the average photon counting rate. Combining CPT with Bayesian inference that can take advantage of the known statistical properties of the bath makes it possible to update dynamical information of the bath in real time with the detection of just a single photon.

While a NV center along with the surrounding nuclear spin bath has been used as a model system for the theoretical analysis, other spin systems such as defect centers in SiC can also serve as CPT-based quantum sensors [40]. In addition to the nuclear spin bath, the real-time sensing can be applied to studies of time-varying magnetic field changes or fluctuations in a variety of systems, for example, two-dimensional materials on a diamond surface [41,42]. Real-time sensing of a single nuclear spin is also possible, provided that the nuclear spin in question features a magnetic field that exceeds considerably that induced by the nuclear spin bath [43,44]. Together with feedback control techniques, the real-time sensing can also be used for the protection of a spin qubit from the surrounding magnetic fluctuations [16]. Overall, CPT-based real-time sensing can add a powerful tool to the emerging technology of quantum sensing.

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APPENDIX A: SIMULATION OF TIME SERIES OF PHOTON COUNTS

For the Λ -type three-level system in Fig. 1 and with $\omega_0 = \nu_0$, the Hamiltonian in the rotating frame is given by

$$H(t) = \frac{\hbar\Omega}{2} (|e\rangle\langle 0| + |e\rangle\langle 1| + |0\rangle\langle e| + |1\rangle\langle e|) + \hbar[\Delta_0 + x(t)]|1\rangle\langle 1|.$$
(A1)

For CPT-based real-time sensing, we operate the CPT process with $C \gg 1$. In this regime, the power broadening [see Eq. (1)] far exceeds the intrinsic decoherence rate ($\gamma_s/2\pi \approx 0.1$ kHz [45]). We have thus set $\gamma_s = 0$ in our simulation unless otherwise specified. The master equation is then given by

$$\dot{\rho}(t) = -\frac{i}{\hbar} [H(t), \rho(t)] + \frac{\Gamma}{2} D(|0\rangle\langle e|)\rho(t) + \frac{\Gamma}{2} D(|1\rangle\langle e|)\rho(t),$$
(A2)

where $D(\hat{O})\rho(t) = \hat{O}\rho(t)\hat{O}^+ - \{\hat{O}^+\hat{O}, \rho(t)\}/2$ and we have also assumed $\kappa = \Gamma/2$.

We have used the SSE to simulate single-photon emissions from the three-level system. The SSE, which tracks every collapse of the system, is given by (with $\hbar = 1$)

$$d|\psi\rangle = -i\left(H - \frac{i\Gamma}{2}|e\rangle\langle e|\right)|\psi\rangle dt + \frac{\Gamma}{2}\langle |e\rangle\langle e|\rangle|\psi\rangle dt + \left(\frac{|0\rangle\langle e|}{\sqrt{\langle |e\rangle\langle e|\rangle}} - 1\right)|\psi\rangle dN_0 + \left(\frac{|1\rangle\langle e|}{\sqrt{\langle |e\rangle\langle e|\rangle}} - 1\right)|\psi\rangle dN_1,$$
(A3)

where N_0 and N_1 are the accumulated photon counts from the $|e\rangle \leftrightarrow |0\rangle$ and $|e\rangle \leftrightarrow |1\rangle$ transitions, respectively. Without any optical emission, the system evolves under an effective Hamiltonian, $H-i\Gamma(|e\rangle\langle e|-\langle |e\rangle\langle e|\rangle)/2$. When a photon is emitted via the $|e\rangle \leftrightarrow |0\rangle$ (or $|e\rangle \leftrightarrow |1\rangle$) transition, the system collapses to the $|0\rangle$ (or $|1\rangle$) state.

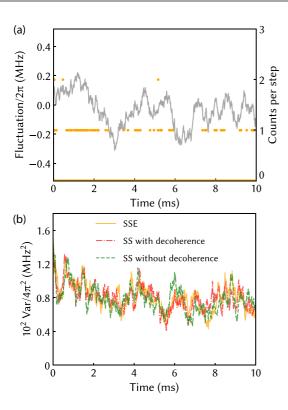


FIG. 6. (a) Time series of photon counts generated by the SSE (orange dots), along with the underlying fluctuations in ω_B (grey lines). The step size for photon counting is 10 μ s. (b) Estimation variances obtained from time series of photon counts generated by the SSE (solid orange lines) and those by the steady-state (SS) population of the excited state (dashed green line with $\gamma_s = 0$ and dot-dashed red line with $\gamma_s/2\pi \approx 0.1$ kHz). The results shown are averages over 100 runs. All other parameters used are the same as for Fig. 3(a).

Figure 6(a) shows, along with the underlying fluctuations in ω_B , the calculated time series of photons detected with $\eta =$ 1.6% and with the collection and detection loss modeled as a random process. The parameters used are the same as those for Fig. 2(c), except that the step size used for the photon counting is 10 μ s for Fig. 6(a) and is 0.1 μ s for Fig. 2(c). The data set in Fig. 6(a) are used for the estimations shown in Fig. 3(a).

We have also used the steady-state population of the excited state to model the single photon emission process. In this case, the steady-state solution to Eq. (A2) is obtained for a given x_n . The A-type three-level system in the excited state has a Poisson distributed probability to emit photons, with a mean detected photon count, $\bar{y}_n = \eta \tau \Gamma \rho_{ee}(x_n)$, for a given time interval τ . Figure 6(b) compares the estimation variances derived from the time series of photon counts obtained with the SSE and those obtained with the steadystate population, for which we have used the OU Bayesian estimator and the same parameters as for Fig. 6(a). As can be seen from Fig. 6(b), estimations obtained with these two simulation approaches exhibit essentially the same behaviors. In addition, we have also included in Fig. 6(b) results obtained with the steady-state population and with $\gamma_s/2\pi \approx 0.1$ kHz, which confirms that with $C \gg 1$, the intrinsic decoherence of the NV center does not affect the CPT-based real-time sensing process.

APPENDIX B: DERIVATION OF THE CRAMER-RAO LOWER BOUND

For the estimation of a time series of parameters, there exists a theoretical lower bound for the estimation variance (i.e., the CRLB), $\operatorname{Var}[\tilde{x}_i(\mathbf{y})] \ge (\mathbf{F}^{-1})_{ii}$, where $(\mathbf{F}^{-1})_{ii}$ is the diagonal component of the inverse Fisher information matrix. The Fisher matrix consists of two parts, $\mathbf{F}_{ij} = \mathbf{F}_{ij}^M + \mathbf{F}_{ij}^B$, with

$$\mathbf{F}_{ij}^{M} = \left\langle \frac{\partial \ln p(\mathbf{y}|\mathbf{x})}{\partial x_{i}} \frac{\partial \ln p(\mathbf{y}|\mathbf{x})}{\partial x_{j}} \right\rangle, \tag{B1}$$

$$\mathbf{F}_{ij}^{B} = \left\langle \frac{\partial \ln p(\mathbf{x})}{\partial x_{i}} \frac{\partial \ln p(\mathbf{x})}{\partial x_{j}} \right\rangle. \tag{B2}$$

For the OU-Bayesian estimations, \mathbf{F}_{ij}^{M} represents the information provided by the time series of the photon counts, whereas \mathbf{F}_{ij}^{B} contains the information provided by the correlations inherent in the fluctuations [39].

We can evaluate the Fisher matrix analytically, with $\mathbf{F}_{ij}^{M} = \Gamma \eta \tau g(\sigma) \delta_{ij}$ [46]. In the limit that the update interval $\tau \to 0$, we have

$$F^{M}(t_{i},t_{j}) = \lim_{\tau \to 0} \frac{\mathbf{F}_{ij}^{M}}{\tau^{2}} = \Gamma \eta g(\sigma) \delta(t_{i} - t_{j}).$$
(B3)

In this limit, the Fourier transform of \mathbf{F}_{ij}^{B} is given by [46]

$$\mathcal{F}\{F^B(t_i, t_j)\} = \mathcal{F}\left\{\lim_{\tau \to 0} \frac{\mathbf{F}^B_{ij}}{\tau^2}\right\} = \frac{1}{2\pi \mathcal{F}\{\sigma^2 e^{-|t_i - t_j|/\tau_N}\}}.$$
 (B4)

Using the above results, we have derived the lower bound for the estimation variance,

$$\operatorname{Var}[\tilde{x}_{t}(\mathbf{y})] \ge F^{-1}(0) = \frac{\sigma^{2}}{\sqrt{1 + 2\Gamma\eta\sigma^{2}\tau_{N}g(\sigma)}}.$$
 (B5)

The CRLB in Eq. (B5) is for the entire data set and occurs near the middle of the time series. However, only the historical information can be used in real-time sensing. In this case, the estimation effectively takes place at the end of the available data set. The CRLB for the real-time sensing can thus be obtained via the ratio between $\mathbf{F}_{T,T}^{-1}$, the diagonal element of the inverse Fisher matrix at the end of the data set, and $\mathbf{F}_{T/2,T/2}^{-1}$, the diagonal element of the inverse Fisher matrix at the middle of the data set.

It can be shown that $\mathbf{F}_{T,T}^{-1}/\mathbf{F}_{T/2,T/2}^{-1} = \det(\mathbf{F}_T')/\det(\mathbf{F}_{T/2}')^2$, where

$$\mathbf{F}_{T}' = \begin{pmatrix} 1 - \tau_{0}^{4} & \tau_{0}^{3} - \tau_{0} & & \\ \tau_{0}^{3} - \tau_{0} & 1 - \tau_{0}^{4} & \ddots & & \\ & \ddots & \ddots & \ddots & \\ & & \ddots & 1 - \tau_{0}^{4} & \tau_{0}^{3} - \tau_{0} \\ & & & & \tau_{0}^{3} - \tau_{0} & 1 - \tau_{0}^{2} \end{pmatrix}$$

is a matrix of order $n = T/\tau$, with $g_0 = \Gamma \eta g(\sigma) \tau \sigma_0^4 / \sigma^2$, $\tau_0 = e^{-\tau/\tau_N}$, and $\sigma_0^2 = \sigma^2 (1-\tau_0^2)$. Using recurrence relations, we obtain

$$\frac{\mathbf{F}_{T,T}^{-1}}{\mathbf{F}_{T/2,T/2}^{-1}} = \frac{\left(Ar_1^n + Br_2^n\right)}{\left(Ar_1^{n/2} + Br_2^{n/2}\right)^2},\tag{B7}$$

where *A*, *B*, r_1 , r_2 , which are determined by the boundary conditions, are functions of τ_0 , g_0 . In the limit that $n \gg 1$, we have

$$\frac{\mathbf{F}_{T,T}^{-1}}{\mathbf{F}_{T/2,T/2}^{-1}} \approx \frac{Ae^{TC_{+}} + Be^{TC_{-}}}{A^{2}e^{TC_{+}} + B^{2}e^{TC_{-}} + 2ABe^{T(C_{+}+C_{-})/2}}, \quad (B8)$$

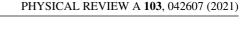
where $C_{\pm} = (-8 \pm \sqrt{1 + 32\Gamma \eta g(\sigma)\sigma^2 \tau_N})/4\tau_N$. For our choice of parameters, $\Gamma \eta g(\sigma)\sigma^2 \tau_N > 2$. The CRLB when only historical information can be used is then given by

$$\mathbf{F}_{T,T}^{-1} \ge F^{-1}(0) \cdot \lim_{T \to \infty} \frac{Ae^{TC_{+}} + Be^{TC_{-}}}{A^{2}e^{TC_{+}} + B^{2}e^{TC_{-}} + 2ABe^{T(C_{+}+C_{-})/2}} = \frac{2F^{-1}(0)}{1 + 4/\sqrt{1 + 32\Gamma\eta\sigma^{2}\tau_{N}g(\sigma)}}.$$
(B9)

APPENDIX C: TRADEOFF BETWEEN WORKING RANGE AND SENSITIVITY

For sensing of a constant parameter, the minimum detectable signal, which scales inversely with the total detection time, is defined as the signal amplitude that yields unit signalto-noise ratio and the sensitivity is defined as the minimum detectable signal per unit time [1]. However, this definition of sensitivity is no longer applicable for sensing a time-varying

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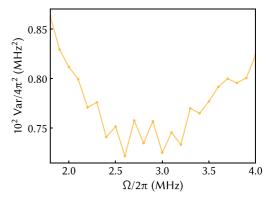


FIG. 7. Estimation variances vs Ω , for which the optimal Δ_0 is used for each Ω , with other parameters the same as for Fig. 5(a).

parameter with a given distribution, for which we would like to know the estimation performance averaged over the entire distribution rather than at a single point. In this case, we can choose the estimation variance $Var[\tilde{x}_n(\mathbf{Y})]$ as an effective measure of how good the sensing process is [17]. Note that $\sqrt{Var[\tilde{x}_n(\mathbf{Y})]}$ is not the minimum detectable signal.

Although it is difficult to define a sensitivity for the sensing of a time-varying parameter, for our case it is apparent that the sharper the CPT dip is [more precisely the greater $(\partial \rho_{ee}/\partial x)^2 / \rho_{ee}(x)$ is], the more sensitive the sensing process becomes. In comparison, the working range of the sensing process, which is the maximum range of the distinguishable signal, scales with the width of the CPT dip. There is thus a tradeoff between the working range and how sensitive the sensing process is. Since the power broadening of the CPT dip is given by Ω^2/Γ , the optimal choice of Ω for the sensing process reflects this tradeoff.

To illustrate this tradeoff, Fig. 7 shows the calculated estimation variance as a function of Ω , for which the optimal bias Δ_0 is used for each Ω . When Ω is too small, the working range is small compared with the actual range of the frequency fluctuation, which is a few times of σ . Whereas when Ω is too large, the sensing process becomes relatively insensitive. Both cases result in relatively large Var[$\tilde{x}_n(\mathbf{Y})$], as shown in Fig. 7.

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